On the Monte Carlo simulation of anisotropic Newtonian gravitation

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Abstract

This paper contains a short introduction to anisotropic Newtonian gravitation. The main focus is on some C++ code.

1 Introduction

$$\alpha = \frac{\beta(R+\epsilon) - \beta(R)}{\epsilon}.\tag{1}$$

$$g = \frac{-\alpha}{r^2}. (2)$$

$$a_N = \sqrt{\frac{nGc\hbar \log 2}{4k\pi R^4}},\tag{3}$$

$$v_N = \sqrt{a_N R}. (4)$$

$$v_{\text{flat}} = 2v_N, \tag{5}$$

$$a_{\text{flat}} = \frac{v_{\text{flat}}^2}{R}.$$
 (6)

$$g_N = \frac{a_N k 2\pi M}{Rc\hbar \log 2}. (7)$$

$$a_N \propto g.$$
 (8)

$$a_{\rm ratio} = \frac{a_{\rm flat}}{a_N}. (9)$$

$$g_{\text{ratio}} = \frac{g}{g_N}.\tag{10}$$

D is where $g_{\text{ratio}} \geq a_{\text{ratio}}$.

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Figure 1: Where D=3, as viewed from the side. The field lines are isotropic, spherical.



Figure 2: Where D = 2.1, as viewed from the side. The field lines are increasingly anisotropic.

Figure 3: Where D=2.001, as viewed from the side. The field lines are anisotropic, disk-like.

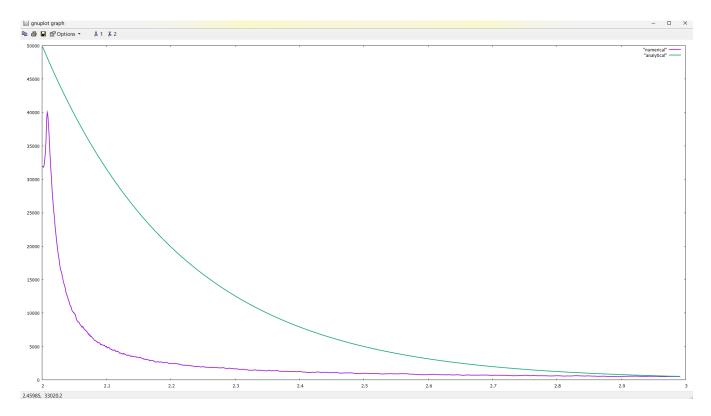


Figure 4: $R = 100, r = 1, n = 10^8, \epsilon = 1.$

References

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