CY121

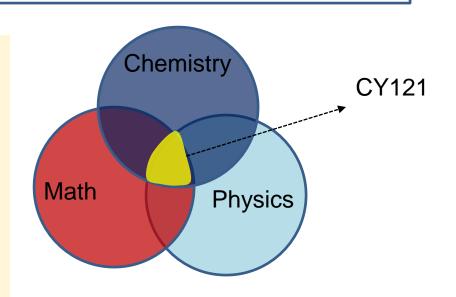
Inorganic & Physical Chemistry

Instructor:

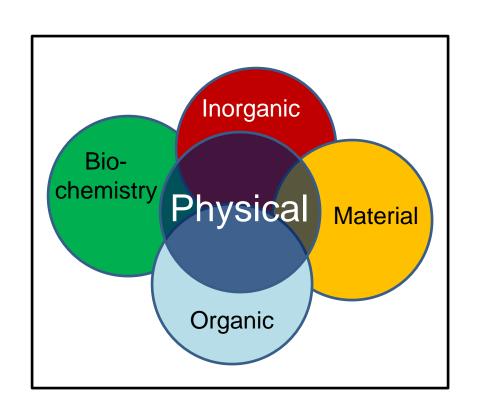
Dr. Debashis PandaAssociate Professor

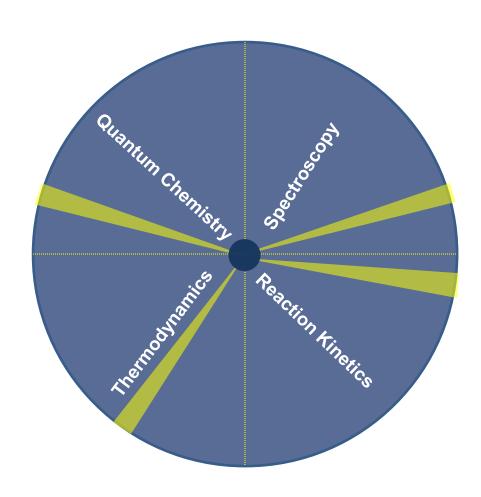
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■ sites.google.com/site/dpanda258



What are we going to cover?





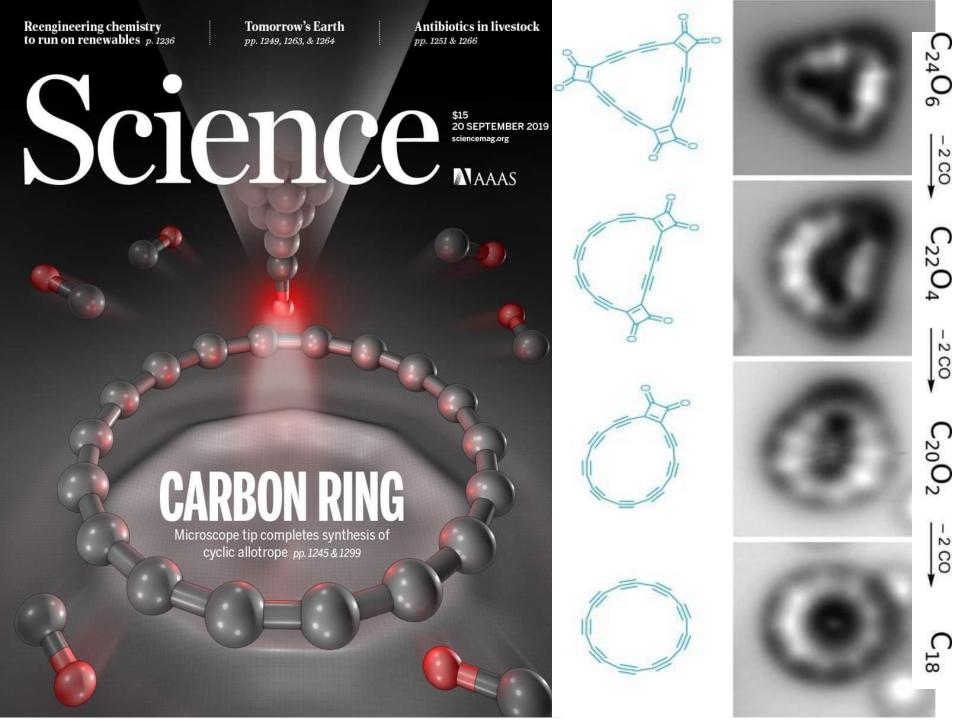
Is there anything "New" to learn?

Unit IV.

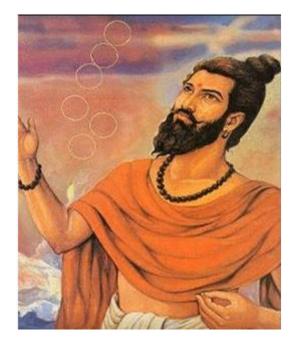
- Wave functions, Probability density, Operator, Eigen function & Eigen Value
- Schrodinger equations,
- Particle in a Box/Ring problem,
- Hydrogen atom & Atomic orbitals

Quantum Chemistry

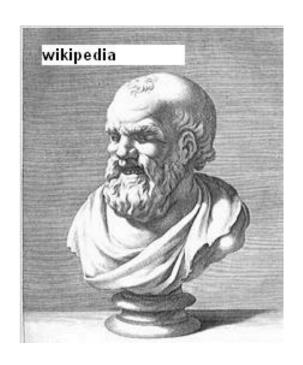
It describes atoms, molecules, their stability and the chemical transformations "ACCURATELY".



History of Atomic Structures



"But in reality there are atoms and space."



Acharya Kanad

~ 600 BC

Kan in Sanskrit means "smallest particle"

Real Name: Kashyap

Democritus

~ 400 BC

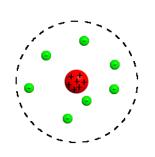
Debate

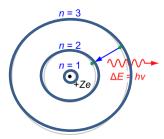
Speculation vs. Authenticity

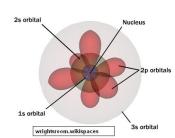
History of Atomic Structures

Atoms can be divided into smaller Particles

















Schrodinger



1803









1897

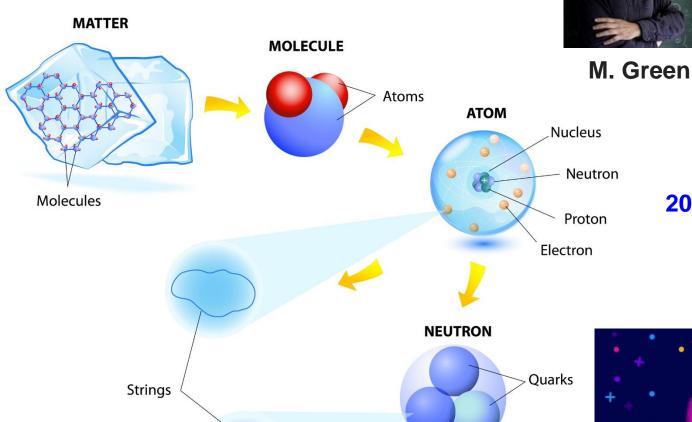
1911

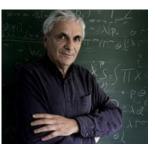
1922

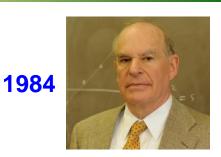
1930

Where we are now on Atomic Structure !!!

STRING THEORY







J. H. Schwarz



Michael R. Douglas



2003

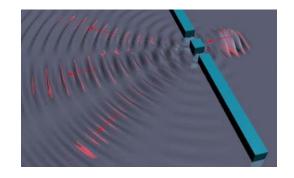
History of the Electron

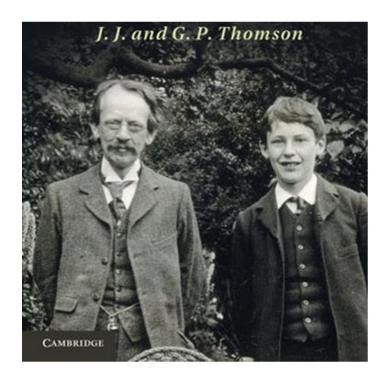
Thompson:

Electrons are well known "particles" with a negative charge

Davisson and Germer Expts:

Electron show interference patterns





Their achievements have been summarized as JJ showing that the electron is a particle and GP showing that it is not, or rather that the electron can also behave as a wave.

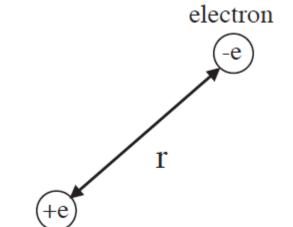
FAILURE OF CLASSICAL MECHANICS

THE CLASSICAL DESCRIPTION OF AN ATOM

How is an e-bound to the nucleus?



$$F(r) = \frac{-e^2}{4\pi \varepsilon_0 r^2}$$



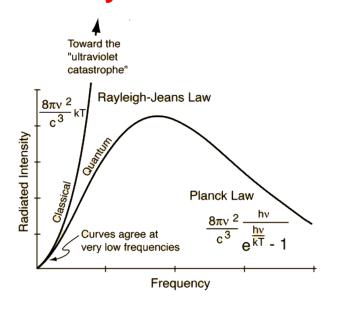
H nucleus

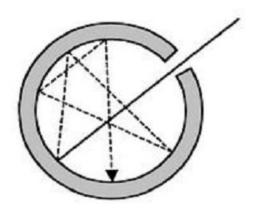
The closer the electron is to the nucleus, the larger the attractive force between the two charges.

This predicts that the electron should plummet into the nucleus.

FAILURES OF CLASSICAL MECHANICS

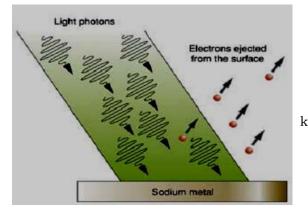
1. Black-Body Radiation

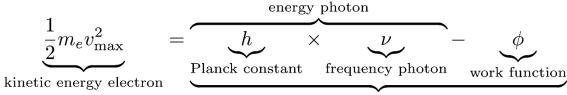




Sun, stars, Heated iron rods...

2. Photo-Electric effect





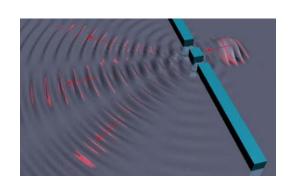
If this is larger than zero, i.e. > 0, then the electron flies off, or else it remains undisturbed.

Wave-Particle Duality: de Broglie Hypothesis

Thompson:

Electrons are well known "particles" with a negative charge

Davisson and Germer Expts: Electron show interference patterns





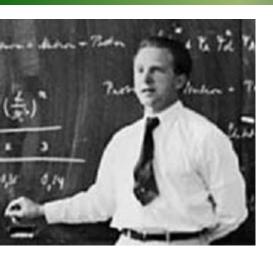
de Broglie

$$\lambda = \frac{h}{p} = \frac{h}{mv}$$

Electron-wave moving @ 106 m/s

$$\lambda = \frac{h}{mv} = \frac{6.6 \times 10^{-34} \text{ J/s}}{9.1 \times 10^{-31} \text{ Kg} \times 1 \times 10^{6} \text{ m/s}} = 7 \times 10^{-10} m$$

Heisenberg's Uncertainty Principle

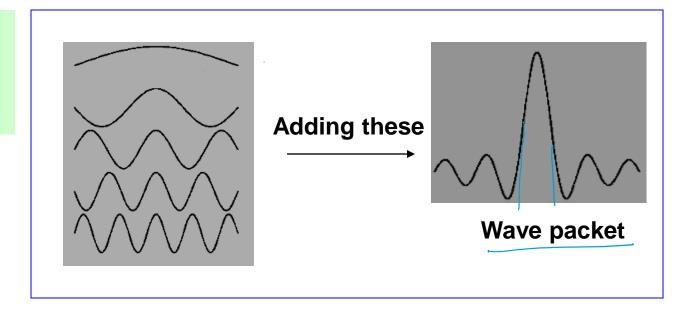


If only one wave with known k (momentum), the position becomes completely uncertain, since 1 wave spreads indefinitely in space.

Specifying position accurately would require many waves, each having a fixed value of momentum (k),

→ P becomes completely uncertain.

$$\Delta x.\Delta p_x \ge \frac{h}{4\pi}$$

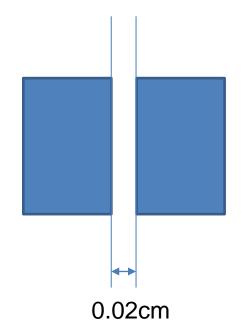


Can we rationalize Uncertainty Principle?

A Simple Experiment

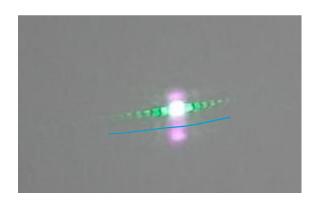
What do we need?

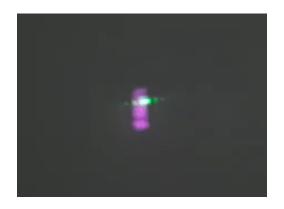
- 1. Two Visiting Cards
- 2. One Laser Pointer (red/green)
- 3. Cello tape
- 4. Camera/Cellphone









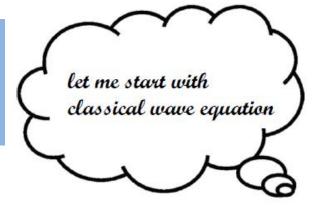


Schrodinger's Philosophy



PARTICLES can be WAVES & WAVES can be PARTICLES

Wave-like Equation for describing electrons, atoms or molecules.



Schrödinger

A concoction of

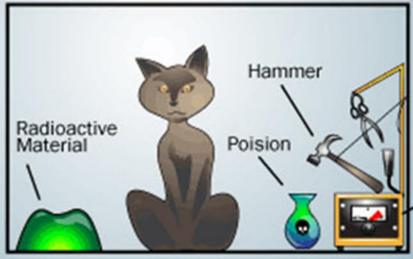
$$E = T + V = \frac{1}{2}mv^2 + V = \frac{p^2}{2m} + V$$

$$E = hv = \hbar \omega$$
 Wave is Particle

$$\lambda = \frac{h}{p} = \hbar k$$
 Particle is Wave



Schrödinger's Cat



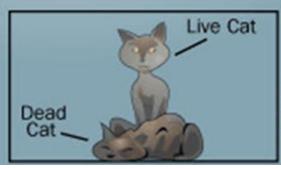
Geiger Counter



The material does not decay; the cat lives.



The material has decayed; the cat has been killed by the poision.



According to the Copenhagen interpretation, the cat is both alive and dead. It exists in a state of "superposition."

POSTULATE-1

Any bound state of a dynamical system of n particles is described as completely as possible by an acceptable, square-integrable function $\Psi(q_1, q_2, \dots, q_{3n}, \omega_1, \omega_2, \dots, \omega_n, t)$, where the q's are spatial coordinates, ω 's are spin coordinates, and t is the time coordinate. $\Psi^*\Psi d\tau$ is the probability that the space-spin coordinates lie in the volume element $d\tau (\equiv d\tau_1 d\tau_2 \cdots d\tau_n)$ at time t, if Ψ is normalized.

The state of a quantum mechanical system is completely specified by the function $\Psi(\mathbf{r},t)$ that depends on the coordinates of the particle, \mathbf{r} and the time, t. This function is called the wavefunction or state function and has the property that $\Psi^*(\mathbf{r},t)\Psi(\mathbf{r},t)d\tau$ is the probability that the particle lies in the volume element $d\tau$ located at \mathbf{r} and time t.

This is the *probabalistic* interpretation of the wavefunction. As a result the wavefunction must satisfy the condition that finding the particle *somewhere* in space is 1 and this gives us the normalisation condition,

$$\int_{-\infty}^{+\infty} \Psi^*(\mathbf{r}, t) \Psi(\mathbf{r}, t) d\tau = 1$$

The other conditions on the wavefunction that arise from the probabilistic interpretation are that it must be single-valued, continuous and finite. We normally write wavefunctions with a normalisation constant included.

POSTULATE- 2

To every observable dynamical variable M there can be assigned a linear hermitian operator $\widehat{\mathbf{M}}$ in quantum mechanics.

Observable	Classical Symbol	Quantum Operator	Operation
position	${f r}$	\hat{r}	multiply by ${f r}$
momentum	p	\hat{p}	$-i\hbar(\hat{i}\frac{\partial}{\partial x}+\hat{j}\frac{\partial}{\partial y}+\hat{k}\frac{\partial}{\partial z})$
kinetic energy	T	\hat{T}	$\frac{-\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right)$
potential energy	$V({f r})$	$\hat{V}(\mathbf{r})$	multiply by $V(\mathbf{r})$
total energy	E	${\cal H}$	$\frac{-\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) + V(\mathbf{r})$
angular momentum	l_x	\hat{l}_x	$-i\hbar(yrac{\partial}{\partial z}-zrac{\partial}{\partial y})$
	l_y	\hat{l}_y	$-i\hbar(zrac{\partial}{\partial x}-xrac{\partial}{\partial z})$
	l_z	\hat{l}_z	$-i\hbar(xrac{\partial}{\partial y}-yrac{\partial}{\partial x})$

Hermitian Operators: A physical variable must have real expectation values (and eigenvalues).

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	l_z	\hat{l}_z	$-i\hbar(xrac{\partial}{\partial y}-yrac{\partial}{\partial x})$

Hermitian Operators: A physical variable must have real expectation values (and eigenvalues).

POSTULATE- 3

In any measurement of the observable associated with operator \hat{A} , the only values that will ever be observed are the eigenvalues, a, that satisfy the eigenvalue equation,

$$\hat{A}\Psi = a\Psi$$

POSTULATE- 4

If a system is in a state described by the normalised wavefunction, Ψ , then the average value of the observable corresponding to \hat{A} is given by

$$<\hat{A}> = \int_{-\infty}^{+\infty} \Psi^* \hat{A} \Psi d\tau$$

Eigen values

Operator operating on a function results in re-generating the same function multiplied by a number

$$\hat{A} \cdot f(x) = a \cdot f(x)$$
 Eigen Value Equation

The function f(x) is eigenfunction of operator \hat{A} and a its eigenvalue

$$f(x) = Sin(\alpha x)$$
$$\frac{d}{dx}f(x) = \alpha \cdot Cos(\alpha x)$$

$$Sin(\alpha x)$$
 is an eigenfunction of operator $\frac{d^2}{dx^2}$ and $-\alpha^2$ is its eigenvalue

$$\frac{d^2}{dx^2}f(x) = \frac{d}{dx}[\alpha.\cos(\alpha x)] = -\alpha^2 \cdot \sin(\alpha x) = -\alpha^2 \cdot f(x)$$

Operators

- A symbol that tells you to do something to whatever follows it
- Operators can be real or complex,
- Operators can also be represented as matrices

Classical Variable

Position, x

Momentum, $p_x = mv$

Kinetic Energy,
$$T_x = \frac{p_x^2}{2m}$$

Kinetic Energy, $T = \frac{p_x^2}{2m} + \frac{p_y^2}{2m} + \frac{p_z^2}{2m}$ $\hat{T} = \frac{-\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial u^2} + \frac{\partial^2}{\partial z^2} \right)$

Potential Energy, V(x)

QM Operator

$$\hat{p}_x = \frac{\hbar}{i} \frac{d}{dx} = -i\hbar \frac{d}{dx}$$

$$\widehat{T}_x = \frac{-\hbar^2}{2m} \frac{d^2}{dx^2}$$

$$\widehat{T} = \frac{-\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right)$$

$$\widehat{V}(x)$$

Operators

Linear momentum

$$p_x = mv_x$$

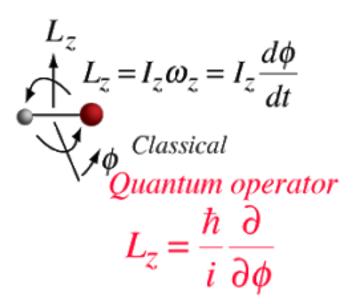
$$\stackrel{m}{\bullet} \stackrel{v_{\chi}}{\longrightarrow}$$

Classical

Quantum operator

$$p_{x} = \frac{\hbar}{i} \frac{\partial}{\partial x}$$

Angular momentum



Rationalizing the Momentum Operator

$$\Psi(x,t) = e^{ikx}e^{i\omega t}$$

$$\Psi(x) = Ae^{ikx}$$

$$\frac{\partial \Psi}{\partial x} = ik \, \Psi = i \, \frac{p}{\hbar} \Psi$$

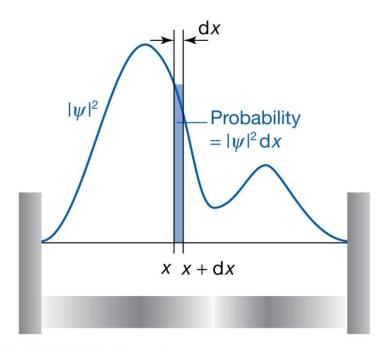
$$\frac{\hbar}{i} \frac{\partial \Psi}{\partial x} = p \Psi$$

$$k = \frac{2\pi}{\lambda} = \frac{2\pi}{h/p} = \frac{p}{\hbar}$$

$$\hat{p} = \frac{\hbar}{i} \frac{\partial}{\partial x}$$

Wave function (Ψ)

For every dynamical system, there exists a wave function Ψ , from which all possible predictions about the physical properties of the system can be obtained.



If the wavefunction of a particle has the value ψ at some point x, then the probability of finding the particle between x and x + dx is proportional to $|\psi|^2 dx$.

Wave function

A wave function consists of all information about the given quantum mechanical system.

Conditions for an Acceptable Wave function:

- 1) The wave function must be continuous. All its partial derivatives must also be continuous.
- 2) It must be quadratically integrable.
- 3) It must be single valued.
- 4) The wave function must satisfy the boundary conditions of the quantum mechanical system it represents.

Born Interpretation: Probability density



The state of a quantum mechanical system is completely specified by a wavefunction $\Psi(x,t)$, which can be complex

All possible information can be derived from $\Psi(x,t)$

From the analogy of classical wave equation, Intensity is replaced by Probability. The probability is proportional to the square of the of the wavefunction $|\Psi(x,t)|^2$, known as probability density P(x)

Born Interpretation

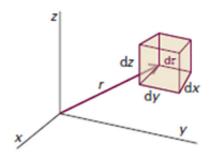


Probability density

$$P(x) = |\Psi(x,t)|^2 = \Psi^*(x,t) \cdot \Psi(x,t)$$

Probability

$$P(x_a \le x \le x_a + dx) = |\Psi(x,t)|^2 dx = \Psi^*(x_a,t) \cdot \Psi(x_a,t) dx$$



Probability in 3-dimensions

$$P(x_{a} \le x \le x_{a} + dx, y_{a} \le y \le y_{a} + dy, z_{a} \le z \le z_{a} + dz)$$

$$= \Psi^{*}(x_{a}, y_{a}, z_{a}, t').\Psi(x_{a}, y_{a}, z_{a}, t')dxdydz$$

$$= |\Psi(x_{a}, y_{a}, z_{a}, t')|^{2} d\tau$$

Normalization of Wave function

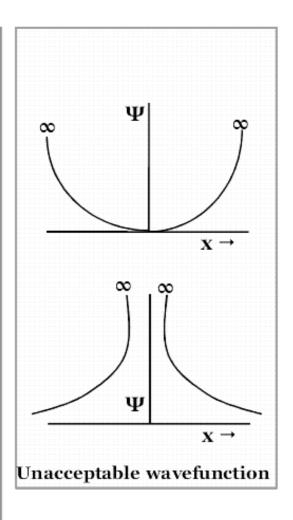
Since $\Psi^*\Psi d\tau$ is the probability, the total probability of finding the particle somewhere in space has to be unity

$$\iiint_{all \ space} \Psi^*(x, y, z).\Psi(x, y, z) dxdydz$$

$$= \int_{\mathbb{R}} \Psi^* \Psi d\tau - /\Psi |\Psi\rangle -$$

$$= \int_{all \ space} \Psi^* \Psi d\tau = \langle \Psi | \Psi \rangle = 1$$

If integration diverges, i.e. $\rightarrow \infty$: Ψ can not be normalized, and therefore is NOT an acceptable wave function. However, a constant value C \neq 1 is perfectly acceptable.



 Ψ must vanish at $\pm \infty$, or more appropriately at the boundaries and Ψ must be finite

Laws of Quantum Mechanics: Average Value

The mathematical description of quantum mechanics is built upon the concept of an operator

The values which come up as result of an experiment are the eigenvalues of the self-adjoint linear operator.

The average value of the observable corresponding to operator is

 $\langle a \rangle = \int \Psi * \hat{A} \Psi dv$

The state of a system is completely specified by the wavefunction $\Psi(x,y,z,t)$ which evolves according to time-dependent Schrodinger equation

Time-Independent Schrödinger equation

TIME-DEPENDENT Schrödinger equation :

$$\widehat{H} \cdot \Psi(x, y, z, t) = i\hbar \frac{\partial}{\partial t} \Psi(x, y, z, t)$$

$$\Psi(x,y,z,t) = \psi(x,y,z) \cdot \phi(t)$$

$$\Psi(x,y,z,t) = \psi(x,y,z) \cdot \phi(t) \qquad \widehat{H} = \frac{-\hbar^2}{2m} \nabla^2 + V(x,y,z)$$

$$\widehat{H}(\psi \cdot \phi) = i\hbar \frac{\partial}{\partial t} (\psi \cdot \phi)$$

 \widehat{H} operates only on ψ and $\frac{\partial}{\partial t}$ operates only on ϕ

$$\phi \cdot \widehat{H} \psi = \psi \left(i\hbar \frac{\partial}{\partial t} \phi \right) \implies \frac{\widehat{H} \psi}{\psi} = \frac{1}{\phi} \left(i\hbar \frac{\partial}{\partial t} \phi \right)$$

Time-Independent Schrödinger equation

In classical mechanics \hat{H} represents total energy We can therefore write

$$\widehat{H}\psi = W\psi$$
 as $\widehat{H}\psi = E\psi$

$$\widehat{H}\psi = E\psi$$

$$\left[-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x) \right] \psi(x) = E \cdot \psi(x)$$

Schrodinger equation is an eigen-value equation

There can be many solutions $\psi_n(x)$ each corresponding to different energy E_n

Time-Independent Schrödinger equation

In 3-dimensions the Schrodinger equation is

$$\left[-\frac{\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) + V(x, y, z) \right] \psi(x, y, z) = E \cdot \psi(x, y, z)$$

For 'n' particle system the Schrodinger equation in 3dimensions is

$$\sum_{i=1}^{n} \left[-\frac{\hbar^{2}}{2m} \left(\frac{\partial^{2}}{\partial x_{i}^{2}} + \frac{\partial^{2}}{\partial y_{i}^{2}} + \frac{\partial^{2}}{\partial z_{i}^{2}} \right) + V(x,y,z) \right] \psi = E \cdot \psi$$

$$\psi \Leftarrow \psi(x_1, x_2, x_3, ...x_{n-1}, x_n, y_1, y_2, y_3, ...y_{n-1}, y_n, z_1, z_2, z_3, ...z_{n-1}, z_n)$$

Exactly Solvable QM System: Free Particle

Time-independent Schrodinger equation

$$\widehat{H}\psi = E\psi$$

$$\left[-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x) \right] \psi(x) = E \cdot \psi(x)$$

For a free particle V(x)=0There are no external forces acting

$$-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial x^2}\psi(x) = E \cdot \psi(x)$$

Free Particle

Let us assume

$$\psi(x) = A\sin kx + B\cos kx$$

Trial Solution

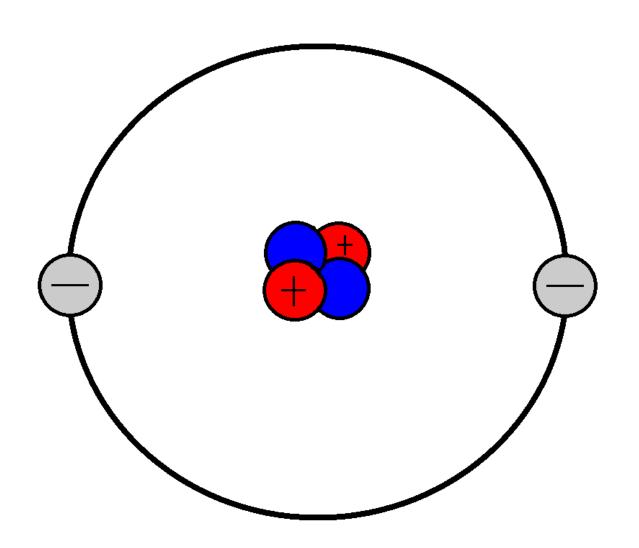
$$\psi(x) = A\sin kx + B\cos kx$$

$$\frac{\partial}{\partial x}\psi(x) = \frac{\partial}{\partial x}(A\sin kx + B\cos kx) = k(A\cos kx - B\sin kx)$$

$$\frac{\partial^2}{\partial x^2}\psi(x) = -k^2 \left(A \sin kx + B \cos kx \right) = -k^2 \psi(x)$$

$$-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial x^2}\psi(x) = E \cdot \psi(x)$$

$$\frac{\hbar^2}{2m}k^2\psi(x) = E \cdot \psi(x) \quad \Rightarrow \quad \left(E = \frac{\hbar^2 k^2}{2m}\right) \Rightarrow \quad k = \pm \frac{\sqrt{2mE}}{\hbar}$$



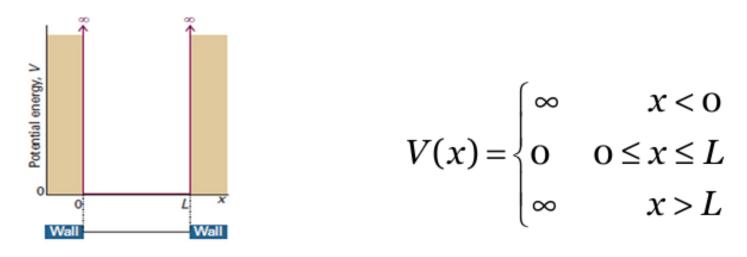
Free Particle

$$E = \frac{\hbar^2 k^2}{2m}$$
 There are no restrictions on k
 E can have any value
Energies of free particles are continuous

$$\psi(x) = A\sin\frac{\sqrt{2mE}}{\hbar}x + B\cos\frac{\sqrt{2mE}}{\hbar}x$$

No Quantization

All energies are allowed



$$V(x) = \begin{cases} \infty & x < 0 \\ 0 & 0 \le x \le L \\ \infty & x > L \end{cases}$$

$$-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial x^2}\psi(x)+V(x)\psi(x)=E\cdot\psi(x)$$

For regions in the space x < 0 and $x > L \Rightarrow V = \infty$

$$\frac{\partial^2}{\partial x^2} \psi(x) = \frac{2m}{\hbar^2} (V - E) \cdot \psi(x) = \infty \cdot \psi(x)$$

Boundary Condition $x = 0 \implies \psi(x) = 0$

$$\psi(x) = A \sin kx + B \cos kx \implies \psi(x) = A \sin kx$$

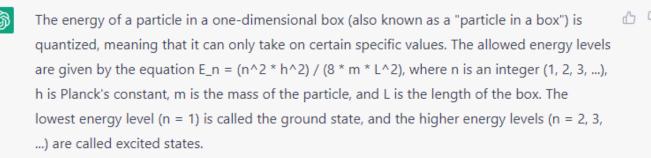
Boundary Condition $x = L \implies \psi(L) = 0$

$$\sin kL = 0 \implies kL = n\pi \quad n=1,2,3,4...$$

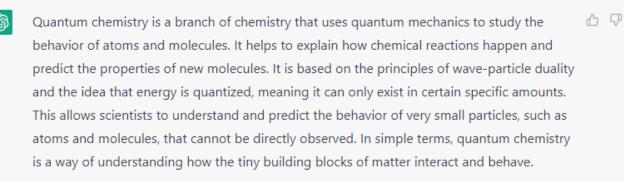
$$E = \frac{\hbar^2 k^2}{2m}$$
 and $k = \frac{n\pi}{L}$

$$E_n = \frac{n^2 \pi^2 \hbar^2}{2mL^2} = \frac{n^2 h^2}{8mL^2}$$
 n=1,2,3,4...





DE Explain quantum chemistry in simple terms



write a poem on particle in a 1D box



Trapped in a box, a particle so small,

No room to move, no space at all.

A prison of length, with no way to escape,

A fate it must bear, with no hope to shape.

It bounces and bounces, in its tiny cell,
With energy high, and a tale to tell.
A wave of probability, in the confined space,
A quantum dance, with a steady pace.

It cannot be seen, but it's always there,
A prisoner of the box, with no room to spare.
But its energy states, they form a range,
A spectrum of light, with a hidden change.

So let us not pity, the particle's plight,

For it is a marvel, in the world of quantum light.

It may be confined, but it's still alive,

A spark in the darkness, that will always thrive.



write a poem on RGIPT Jais Amethi

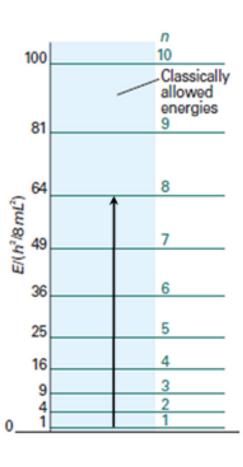


In the land of Amethi, so green and fair, RGIPT stands tall, a beacon of care. Where knowledge and technology meet, Its halls are bustling with students so sweet.

With a faculty of great renown,
RGIPT's reputation will never drown.
Research and innovation lead the way,
As students learn and grow each day.

In Jais, a town of peace and pride,
RGIPT's future shines bright and wide.
With a vision to excel and reach new heights,
RGIPT's story will be one of great lights.

So let us sing of RGIPT's glory,
In Amethi's land, so fair and hoary.
May its name be known far and wide,
A leader in education, forever abide.



The lowest allowed energy level is for n=1

$$E_1 = \frac{\pi^2 \hbar^2}{2mL^2}$$
 has a non zero value \Rightarrow Zero Point Energy

$$h\nu = \Delta E = E_f - E_i = \frac{n_f^2 h^2}{8mL^2} - \frac{n_i^2 h^2}{8mL^2} = \left(n_f^2 - n_i^2\right) \frac{h^2}{8mL^2}$$

Wavefunction
$$\psi(x) = A \sin kx = A \sin \frac{n\pi}{L}x$$

Normalization
$$\int_0^L \psi^*(x) \cdot \psi(x) \cdot dx = A^2 \int_0^L \sin^2 \frac{n\pi}{L} x \cdot dx = 1$$

$$A = \sqrt{\frac{2}{L}} \qquad \psi(x) = \sqrt{\frac{2}{L}} \sin \frac{n\pi}{L} x$$

Wavefunction
$$\psi(x) = \sqrt{\frac{2}{L}} \sin \frac{n\pi}{L} x$$

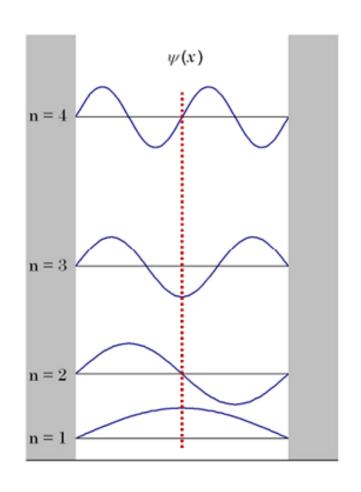
Wave function

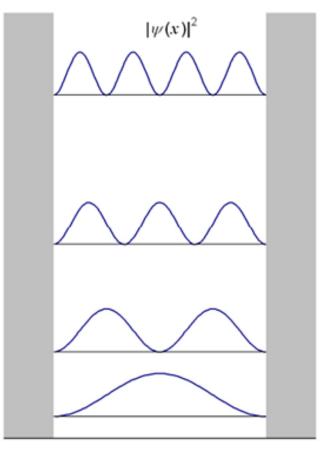
Probability density

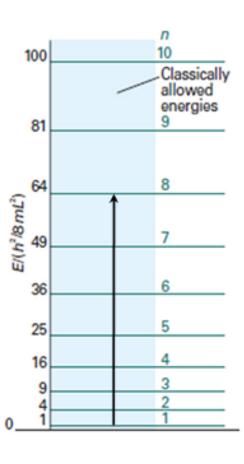
n=1,3.. (odd) Symmetric (even function)

n=2,4.. (even) Anti-Symmetric (odd function)

Number of Nodes (zero crossings) = n-1







The lowest allowed energy level is for n=1

$$E_1 = \frac{\pi^2 \hbar^2}{2mL^2}$$
 has a non zero value \Rightarrow Zero Point Energy

Wavefunction
$$\psi(x) = A \sin kx = A \sin \frac{n\pi}{L}x$$

Normalization
$$\int_0^L \psi^*(x) \cdot \psi(x) \cdot dx = A^2 \int_0^L \sin^2 \frac{n\pi}{L} x \cdot dx = 1$$

$$A = \sqrt{\frac{2}{L}} \qquad \psi(x) = \sqrt{\frac{2}{L}} \sin \frac{n\pi}{L} x$$

Expectation Value for position

$\langle x \rangle = \int \psi^* \cdot x \cdot \psi \cdot dx$ $= \int_0^L \sqrt{\frac{2}{L}} \sin \frac{n\pi}{L} x \cdot x \cdot \sqrt{\frac{2}{L}} \sin \frac{n\pi}{L} x \cdot dx$ $= \frac{2}{L} \int_0^L x \cdot \sin^2 \frac{n\pi}{L} x \cdot dx$

Expectation Value for momentum

$$\begin{aligned} \left\langle p_{x}\right\rangle &= \int \psi^{*} \cdot \left(-i\hbar \frac{\partial}{\partial x}\right) \cdot \psi \cdot dx \\ &= -i\hbar \int_{0}^{L} \sqrt{\frac{2}{L}} \sin \frac{n\pi}{L} x \cdot \frac{\partial}{\partial x} \cdot \sqrt{\frac{2}{L}} \sin \frac{n\pi}{L} x \cdot dx \\ &= \frac{-2i\hbar n\pi}{L^{2}} \int_{0}^{L} \sin \frac{n\pi}{L} x \cdot \cos \frac{n\pi}{L} x \cdot dx \\ &= 0 \end{aligned}$$

Building Up the Periodic Table: The Basis

1. Various quantum numbers

2. Hund's Rule:

When more than one orbital has the same energy (e.g. px, py, pz), electron occupy separate orbitals and do so with parallel spins.

3. Pauli (Exclusion) Principle

No more than two electrons shall occupy a single orbital and, if two do occupy a single orbital, then their spins must be paired. or

"no two electrons can have the same four quantum numbers"

4. The order of orbitals for a given quantum number depends on

Shielding Effects (Z*)
Penetration of orbitals

Electronic Configuration

H 1s¹
He 1s²
Li 1s²2s¹
.....
F 1s²2s²2p⁵
Ne 1s²2s²2p⁶

3s < 3p < 3d

Four Quantum Numbers

The principal quantum number, n = 1, 2, ..., determines the energy E of the electron in the atom.

The angular momentum quantum number, $\ell = 0, 1, 2, \dots, n-1$,

determines the orbital angular momentum of the

electron.

$$\ell = 0$$
 s "sharp"
 $\ell = 1$ p "principal"
 $\ell = 2$ d "diffuse"
 $\ell = 3$ f "fundamental or fine"

The magnetic quantum number, $m = 0, \pm 1, ..., \pm \ell$

determines the amount of angular momentum of the electron around a particular axis.

Orbit and **Orbital**

An **Orbit** is the path of an electron around the nucleus of an atom.

When a planet moves around the sun, you can plot a definite path for it which is called an orbit.

The impossibility of drawing orbits for electrons

The *Heisenberg Uncertainty Principle* says - that you can't know with certainty both where an electron is and where it's going next.

An **Orbital** is the three dimensional space where the probability of finding an electron around the nucleus is maximum. In other words, 'Orbital' refers to an electron's probable location.

Schrodinger equation for Hydrogen-like Atom

$$\left[-\frac{\hbar^2}{2mr^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) - \frac{\hbar^2}{2mr^2} \left(\frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right) + V(r) \right] \Psi(r, \theta, \phi) \ = \ E \Psi(r, \theta, \phi)$$

$$\psi(r,\theta,\phi) = R(r) P(\theta) F(\phi)$$

$$Y_{l,m_l}(\theta,\phi) \longrightarrow \text{Angular part}$$
Radial part

$$R(r)$$
 Solution exists iff $n=1,\,2,\,3,\,....$ $\ell=0,\,1,\,2,....,\,n-1$ $F(\phi)$ Solution exists iff $m_i=\ell,\,\ell+1,\,....,\,+\ell$

In three dimensions the Schrodinger equation

$$\left(-\frac{\hbar^2}{2m}\nabla^2 + V\right)\psi = E\psi$$

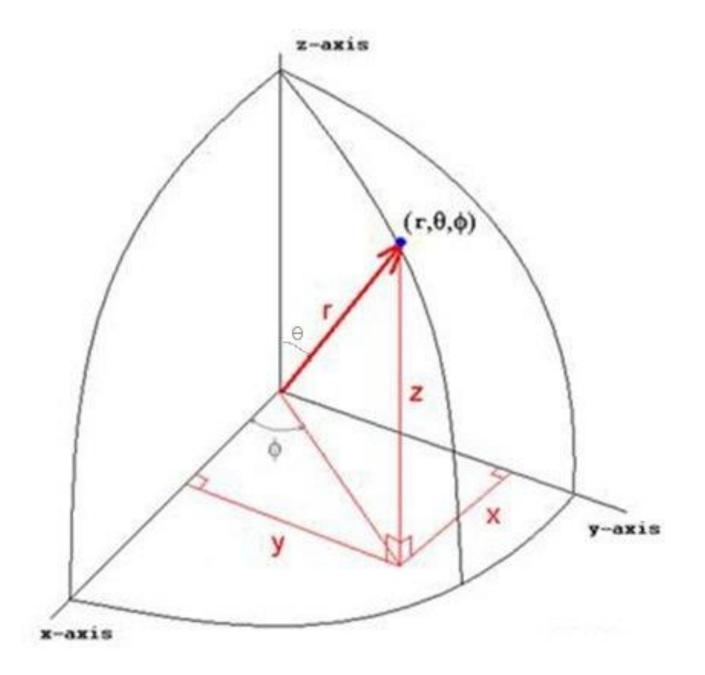
where ∇^2 is the Laplacian operator

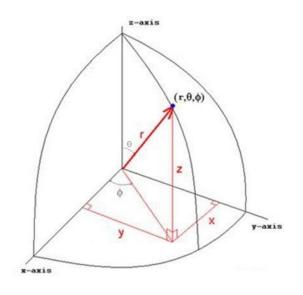
Laplacian in spherical coordinates

$$\nabla^2 = \frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r} + \frac{1}{r^2} \left(\frac{\partial^2}{\partial \theta^2} + \frac{1}{\tan \theta} \frac{\partial}{\partial \theta} + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right)$$

Using the Laplacian in spherical coordinates, the Schrodinger equation becomes

$$-\frac{\hbar^2}{2m} \left[\frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right] \psi + V(r) \psi = E \psi$$





$$z = r \cos \theta$$

$$r_{projection} = r \sin \theta$$

$$x = r_{projection}\cos\varphi = r\cos\varphi\sin\theta$$

$$y = r_{projection} \sin \varphi = r \sin \varphi \sin \theta$$

Seperation of Variables: Solution to SE equation of the form

$$\psi(r, \theta, \phi) = R(r) Y(\theta, \phi)$$

$$\frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) R(r) Y(\theta, \phi) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) R(r) Y(\theta, \phi)$$

$$+\frac{1}{r^2\sin^2\theta}\frac{\partial^2}{\partial\phi^2}R(r)Y(\theta,\phi) - \frac{2m}{\hbar^2}\Big[V(r) - E\Big]R(r)Y(\theta,\phi) = 0$$



$$Y(\theta,\phi)\frac{1}{r^2}\frac{\partial}{\partial r}\left(r^2\frac{\partial}{\partial r}\right)R(r) + R(r)\frac{1}{r^2\sin\theta}\frac{\partial}{\partial \theta}\left(\sin\theta\frac{\partial}{\partial \theta}\right)Y(\theta,\phi)$$
$$+R(r)\frac{1}{r^2\sin^2\theta}\frac{\partial^2}{\partial \phi^2}Y(\theta,\phi) - \frac{2m}{\hbar^2}\Big[V(r) - E\Big]R(r)Y(\theta,\phi) = 0.$$

$$Y(\theta,\phi)\frac{1}{r^2}\frac{\partial}{\partial r}\left(r^2\frac{\partial}{\partial r}\right)R(r) + R(r)\frac{1}{r^2\sin\theta}\frac{\partial}{\partial \theta}\left(\sin\theta\frac{\partial}{\partial \theta}\right)Y(\theta,\phi)$$
$$+R(r)\frac{1}{r^2\sin^2\theta}\frac{\partial^2}{\partial \phi^2}Y(\theta,\phi) - \frac{2m}{\hbar^2}\Big[V(r) - E\Big]R(r)Y(\theta,\phi) = 0.$$

Dividing the equation by R(r) Y (θ , ϕ), multiplying by r^2 , and rearranging

$$\left\{ \frac{1}{R(r)} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) R(r) - \frac{2mr^2}{\hbar^2} \left[V(r) - E \right] \right\}
+ \left[\frac{1}{Y(\theta, \phi) \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) Y(\theta, \phi) + \frac{1}{Y(\theta, \phi) \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} Y(\theta, \phi) \right] = 0.$$

$$\left\{\frac{1}{R(r)}\frac{\partial}{\partial r}\left(r^2\frac{\partial}{\partial r}\right)R(r)-\frac{2mr^2}{\hbar^2}\Big[V(r)-E\Big]\right\}$$

$$+ \left[\frac{1}{Y(\theta, \phi) \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) Y(\theta, \phi) + \frac{1}{Y(\theta, \phi) \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} Y(\theta, \phi) \right] = 0.$$

The only way the sum of the groups can be zero, **if each group is equal to the same constant.** The constant chosen is known as the separation constant **(K)**.

Some of the physicists and mathematicians of the previous 300 years predicted the separation constant,

the enlightened choice of $\ell(\ell + 1)$ as the separation constant (K).

The radial equation

$$\frac{1}{R(r)}\frac{d}{dr}\left(r^2\frac{d}{dr}\right)R(r) - \frac{2mr^2}{\hbar^2}\left[V(r) - E\right] = l(l+1)$$

The angular equation

$$\frac{1}{Y(\theta,\phi)\sin\theta} \frac{\partial}{\partial\theta} \left(\sin\theta \frac{\partial}{\partial\theta} \right) Y(\theta,\phi) + \frac{1}{Y(\theta,\phi)\sin^2\theta} \frac{\partial^2}{\partial\phi^2} Y(\theta,\phi) = -l(l+1)$$

The signs on the right side are opposite so they do, in fact, sum to zero.

The angular equation

$$\frac{1}{Y(\theta,\phi)\sin\theta} \frac{\partial}{\partial \theta} \left(\sin\theta \frac{\partial}{\partial \theta} \right) Y(\theta,\phi) + \frac{1}{Y(\theta,\phi)\sin^2\theta} \frac{\partial^2}{\partial \phi^2} Y(\theta,\phi) = -l(l+1)$$

The spherical harmonics are also separable

$$Y(\theta, \phi) = f(\theta) g(\phi)$$

$$\frac{1}{f(\theta) g(\phi) \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) f(\theta) g(\phi) + \frac{1}{f(\theta) g(\phi) \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} f(\theta) g(\phi) = -l(l+1)$$

$$\Rightarrow \frac{1}{f(\theta)\sin\theta} \frac{\partial}{\partial \theta} \left(\sin\theta \frac{\partial}{\partial \theta} \right) f(\theta) + \frac{1}{g(\phi)\sin^2\theta} \frac{\partial^2}{\partial \phi^2} g(\phi) = -l(l+1).$$

$$\frac{1}{f(\theta)\,\sin\theta}\frac{\partial}{\partial\theta}\left(\sin\theta\frac{\partial}{\partial\theta}\right)f(\theta) + \frac{1}{g(\phi)\,\sin^2\theta}\frac{\partial^2}{\partial\phi^2}g(\phi) = -l(l+1).$$

Multiplying the equation by sin²θ and rearranging

$$\frac{\sin \theta}{f(\theta)} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) f(\theta) + l(l+1) \sin^2 \theta + \frac{1}{g(\phi)} \frac{\partial^2}{\partial \phi^2} g(\phi) = 0.$$

$$\frac{\sin \theta}{f(\theta)} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) f(\theta) + l(l+1) \sin^2 \theta + \frac{1}{g(\phi)} \frac{\partial^2}{\partial \phi^2} g(\phi) = 0.$$

Again, the only non-trivial solution such that the sum is zero is if the groups of terms each dependent on a single variable is equal to the same constant. Again using an enlightened choice, we pick m² as the separation constant

$$\frac{\sin\theta}{f(\theta)}\frac{d}{d\theta}\left(\sin\theta\frac{d}{d\theta}\right)f(\theta) + l(l+1)\sin^2\theta = m^2$$

$$\frac{1}{g(\phi)}\frac{d^2}{d\phi^2}g(\phi) = -m^2$$

that is how the magnetic moment quantum number is introduced

The radial equation

$$\frac{1}{R(r)}\frac{d}{dr}\left(r^2\frac{d}{dr}\right)R(r) - \frac{2mr^2}{\hbar^2}\left[V(r) - E\right] = l(l+1)$$

The Polar Angle Equation

$$\frac{\sin\theta}{f(\theta)}\frac{d}{d\theta}\left(\sin\theta\frac{d}{d\theta}\right)f(\theta) + l(l+1)\sin^2\theta = m^2$$

The azimuthal angle equation

$$\frac{1}{q(\phi)}\frac{d^2}{d\phi^2}g(\phi) = -m^2$$

The solution to the azimuthal angle equation

$$\frac{1}{g(\phi)}\frac{d^2}{d\phi^2}g(\phi) = -m^2$$

$$g(\phi) = e^{im\phi}$$

$$\Rightarrow g_m(\phi) = e^{im\phi}$$

there are as many solutions as there are allowed values of m

Hint

$$\frac{d^2}{d\phi^2}g_m(\phi) = \frac{d^2}{d\phi^2}e^{im\phi} = \frac{d}{d\phi}(im)e^{im\phi} = (im)^2e^{im\phi} = -m^2g_m(\phi).$$

$$\frac{1}{g(\phi)}\frac{d^2}{d\phi^2}g(\phi) = -m^2$$

The solution to the Polar Angle Equation

$$\frac{\sin\theta}{f(\theta)}\frac{d}{d\theta}\left(\sin\theta\frac{d}{d\theta}\right)f(\theta) + l(l+1)\sin^2\theta = m^2$$

$$\sin\theta \, \frac{d}{d\theta} \left(\sin\theta \, \frac{d}{d\theta} \right) f(\theta) + l(l+1) \, \sin^2\theta \, f(\theta) - m^2 \, f(\theta) = 0.$$

$$\sin\theta \, \frac{d}{d\theta} \left(\sin\theta \, \frac{d}{d\theta} \right) f(\theta) = \sin\theta \, \frac{d}{d\theta} \left(\sin\theta \, \frac{df(\theta)}{d\theta} \right)$$

$$= \sin\theta \, \left(\cos\theta \, \frac{df(\theta)}{d\theta} + \sin\theta \, \frac{d^2f(\theta)}{d\theta^2} \right)$$

$$= \sin^2\theta \, \frac{d^2f(\theta)}{d\theta^2} + \sin\theta \, \cos\theta \, \frac{df(\theta)}{d\theta}.$$

$$\sin^2 \theta \, \frac{d^2 f(\theta)}{d\theta^2} + \sin \theta \, \cos \theta \, \frac{d f(\theta)}{d\theta} + l(l+1) \, \sin^2 \theta \, f(\theta) - m^2 f(\theta) = 0.$$

$$\sin^2\theta \, \frac{d^2 f(\theta)}{d\theta^2} + \sin\theta \, \cos\theta \, \frac{d f(\theta)}{d\theta} + l(l+1) \, \sin^2\theta \, f(\theta) - m^2 f(\theta) = 0.$$

This is a kind of the associated Legendre equation.

The solutions to the associated Legendre equation are the associated Legendre polynomials.

$$P_{l,m}(x) = (-1)^m \sqrt{(1-x^2)^m} \frac{d^m}{dx^m} P_l(x)$$
$$P_l(x) = \frac{(-1)^l}{2^l l!} \frac{d^l}{dx^l} (1-x^2)^l.$$

The solutions to the polar angle equation are the associated Legendre polynomials, which are different for each choice of orbital angular momentum and magnetic moment quantum number.

Angular Wave Functions

$$Y_{l,m}(\theta,\phi) = (-1)^m \sqrt{\frac{(2l+1)(l-m)!}{4\pi(l+m)!}} P_{l,m}(\cos\theta) e^{im\phi} \qquad m \ge 0.$$

$$Y_{l,-m}(\theta,\phi) = Y_{l,m}^*(\theta,\phi), \qquad m < 0$$

the $P_{l,m}(\cos \theta)$ are associated Legendre polynomials

The solution to the Angular Equation

$$l = 0$$
 $Y_{0,0} = \frac{1}{\sqrt{4\pi}}$

$$l = 1 Y_{1,\pm 1} = \sqrt{\frac{3}{8\pi}} \sin \theta e^{\pm i\varphi}$$

$$l = 2 Y_{2,\pm 2} = \sqrt{\frac{15}{32\pi}} \sin^2 \theta e^{\pm 2i\varphi}$$

$$l = 3 Y_{3,\pm 3} = \sqrt{\frac{35}{64\pi}} \sin^3 \theta e^{\pm 3i\varphi}$$

The solution to the Radial Equation

$$\frac{1}{R(r)}\frac{d}{dr}\left(r^2\frac{d}{dr}\right)R(r) - \frac{2mr^2}{\hbar^2}\left[V(r) - E\right] = l(l+1)$$

$$\frac{1}{R(r)}\frac{d}{dr}\left(r^2\frac{d}{dr}\right)R(r) - \frac{2\mu r^2}{\hbar^2}\left[-\frac{e^2}{r} - E\right] - l(l+1) = 0$$

$$\Rightarrow \frac{d}{dr} \left(r^2 \frac{d}{dr} \right) R(r) - \frac{2\mu r^2}{\hbar^2} \left[-\frac{e^2}{r} - E \right] R(r) - l(l+1) R(r) = 0$$

$$\Rightarrow \frac{d}{dr}\left(r^2\frac{d}{dr}\right)R(r) + \left[\frac{2\mu r^2}{\hbar^2}\frac{e^2}{r} + \frac{2\mu r^2}{\hbar^2}E - l(l+1)\right]R(r) = 0.$$

The solution to the Radial Equation

$$y_j^k(x) = e^{-x/2} x^{(k+1)/2} L_j^k(x)$$

$$k = 2l + 1,$$

$$(k + 1)/2 = l + 1,$$

$$i + l + 1 = n$$

Radial Wave Functions

$$R_{n,l}(r) = \sqrt{\left(\frac{2}{na_0}\right)^3 \frac{(n-l-1)!}{2n[(n+l)!]^3}} e^{-r/na_0} \left(\frac{2r}{na_0}\right)^l L_{n-l-1}^{2l+1} \left(\frac{2r}{na_0}\right)^l$$

The solution to the Radial Equation

$$R_{10} = 2\left(\frac{Z}{a_0}\right)^{\frac{3}{2}} e^{-Zr/a_0}$$

$$R_{21} = \frac{1}{\sqrt{3}} \left(\frac{Z}{2a_0}\right)^{\frac{3}{2}} \left(\frac{Zr}{a_0}\right) e^{-Zr/2a_0}$$

$$R_{20} = 2\left(\frac{Z}{2a_0}\right)^{\frac{3}{2}} \left(1 - \frac{Zr}{2a_0}\right) e^{-Zr/2a_0}$$

$$R_{32} = \frac{2\sqrt{2}}{27\sqrt{5}} \left(\frac{Z}{3a_0}\right)^{\frac{3}{2}} \left(\frac{Zr}{a_0}\right)^2 e^{-Zr/3a_0}$$

$$R_{31} = \frac{4\sqrt{2}}{3} \left(\frac{Z}{3a_0}\right)^{\frac{3}{2}} \left(\frac{Zr}{a_0}\right) \left(1 - \frac{Zr}{6a_0}\right) e^{-Zr/3a_0}$$

$$R_{30} = 2\left(\frac{Z}{3a_0}\right)^{\frac{3}{2}} \left(1 - \frac{2Zr}{3a_0} + \frac{2(Zr)^2}{27a_0^2}\right) e^{-Zr/3a_0}$$

The First Seven Laguerre Polynomials

Laguerre polynomials are solutions to the Laguerre equation

$$L_0(x) = 1$$

$$L_1(x) = -x + 1$$

$$L_2(x) = x^2 - 4x + 2$$

$$L_3(x) = -x^3 + 9x^2 - 18x + 6$$

$$L_4(x) = x^4 - 16x^3 + 72x^2 - 96x + 24$$

$$L_5(x) = -x^5 + 25x^4 - 200x^3 + 600x^2 - 600x + 120$$

$$L_6(x) = x^6 - 36x^5 + 450x^4 - 2400x^3 + 5400x^2 - 4320x + 720$$

Associated Laguerre polynomials can be calculated from Laguerre polynomials using the generating function

$$L_j^k(x) = \left(-1\right)^k \frac{d^k}{dx^k} L_{j+k}(x)$$

Some Associated Laguerre Polynomials.

Associated Laguerre polynomials can be calculated from Laguerre polynomials using the generating function

$$L_j^k(x) = \left(-1\right)^k \frac{d^k}{dx^k} L_{j+k}(x)$$

$$\begin{array}{lll} L_0^0(x) = L_0(x) & L_0^2(x) = 2 \\ L_1^0(x) = L_1(x) & L_3^0(x) = L_3(x) \\ L_1^1(x) = -2x + 4 & L_3^1(x) = -4x^3 + 48x^2 - 144x + 96 \\ L_0^1(x) = 1 & L_2^3(x) = 60x^2 - 600x + 1200 \\ L_2^0(x) = L_2(x) & L_3^3(x) = -120x^3 + 2160x^2 - 10800x + 14400 \\ L_2^1(x) = 3x^2 - 18x + 18 & L_3^2(x) = -20x^3 + 300x^2 - 1200x + 1200 \\ L_2^2(x) = 12x^2 - 96x + 144 & L_1^3(x) = -24x + 96 \\ L_1^2(x) = -6x + 18 & L_0^3(x) = 6 \end{array}$$

Schrodinger equation for Hydrogen-like Atom

n	l	m _e	F(φ)	Ρ(θ)	R(r)
1	0	0	$\frac{1}{\sqrt{2\pi}}$	$\frac{1}{\sqrt{2}}$	$\frac{2}{a_0^{3/2}}e^{-r/a_0}$
2	0	0	$\frac{1}{\sqrt{2\pi}}$	$\frac{1}{\sqrt{2}}$	$\frac{1}{2\sqrt{2} \mathbf{a}_0^{3/2}} \left[2 - \frac{\mathbf{r}}{\mathbf{a}_0} \right] \mathbf{e}^{-\mathbf{r}/2\mathbf{a}_0}$
2	1	0	$\frac{1}{\sqrt{2\pi}}$	√6 cos θ	$\frac{1}{2\sqrt{6}}\frac{\mathbf{r}}{\mathbf{a}_0^{3/2}}\frac{\mathbf{r}}{\mathbf{a}_0} e^{-\mathbf{r}/2\mathbf{a}_0}$
2	1	±1	$\frac{1}{\sqrt{2\pi}}e^{\pm i\phi}$	√3 sin θ	$\frac{1}{2\sqrt{6}}\frac{\mathbf{r}}{\mathbf{a}_0^{3/2}}\frac{\mathbf{r}}{\mathbf{a}_0}\mathbf{e}^{-\mathbf{r}/2\mathbf{a}_0}$

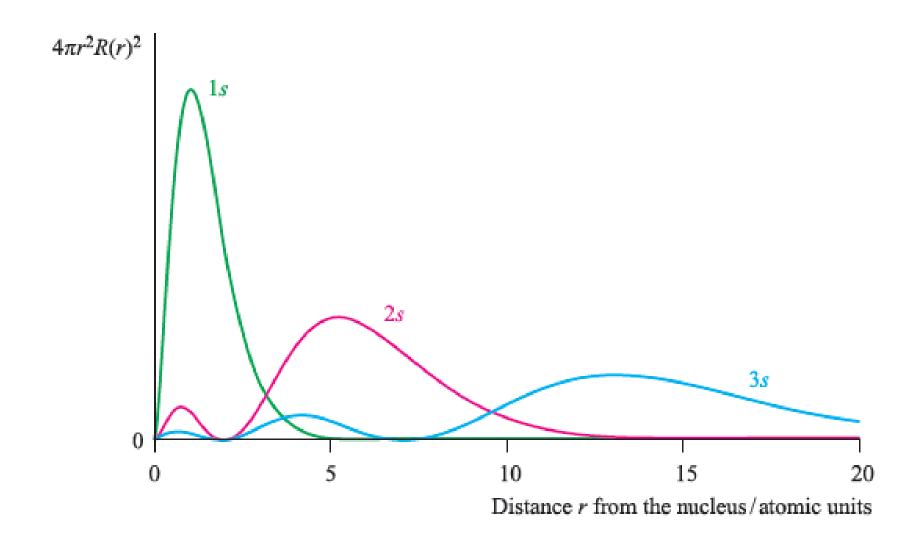
Schrodinger equation for Hydrogen-like

n	l	m_{ℓ}	$Ψ_n \ell_{m_\ell}(r, \theta, \phi)$
1	0	0	1s √π a₀3/2 e-r/a ₀
2	0	0	2s $\frac{1}{4\sqrt{2\pi}\mathbf{a}_0^{3/2}} \left[2 - \frac{\mathbf{r}}{\mathbf{a}_0} \right] \mathbf{e}^{-\mathbf{r}/2} \mathbf{a}_0$
2	1	0	2p $\frac{1}{4\sqrt{2\pi}} \frac{\mathbf{r}}{\mathbf{a}_0^{3/2}} \frac{\mathbf{e}^{-\mathbf{r}/2} \mathbf{a}_0}{\mathbf{a}_0} \cos \theta$
2	1	±1	2p $\frac{1}{8\sqrt{\pi} a_0^{3/2}} \frac{\mathbf{r}}{a_0} \mathbf{e}^{-\mathbf{r}/2\mathbf{a}_0} \sin \theta \mathbf{e}^{\pm \mathbf{i} \phi}$
			$\mathbf{a}_0 = \frac{\hbar^2}{\text{me}^2} = .0529 \text{ nm} = \text{first Bohr radius}$

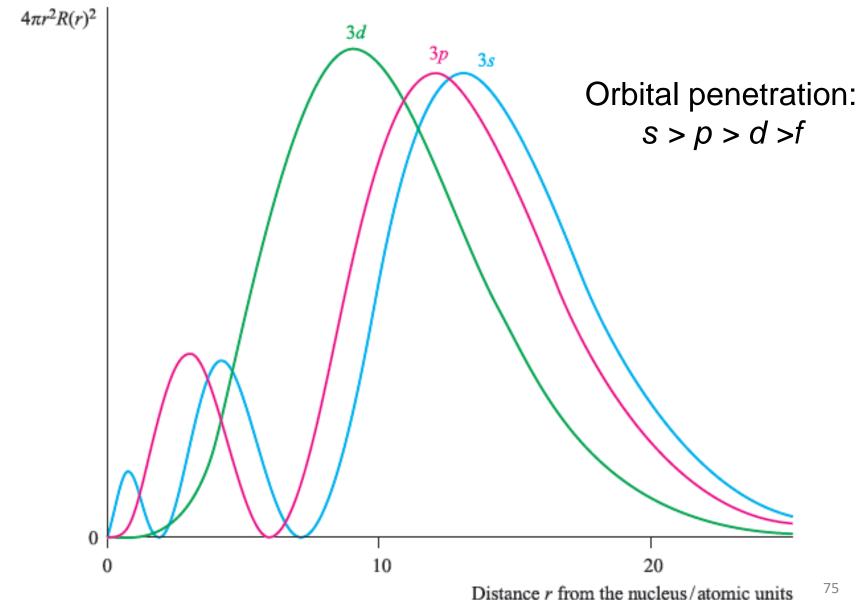
Schrodinger equation for Hydrogen-like Atom

Schrodinger equation for Hydrogen-like Atom

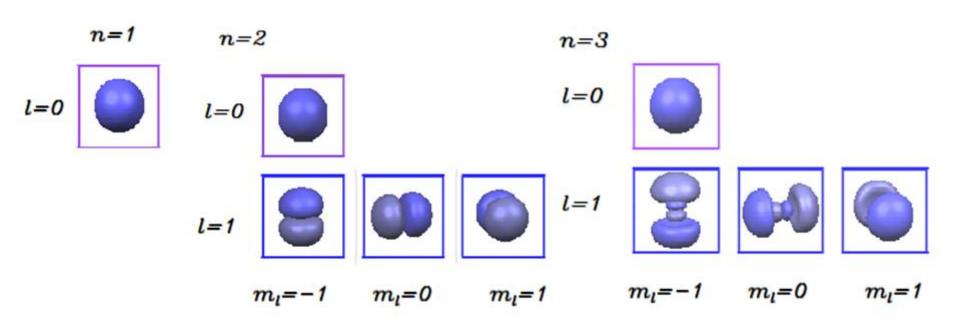
Distribution of Radial Probability density



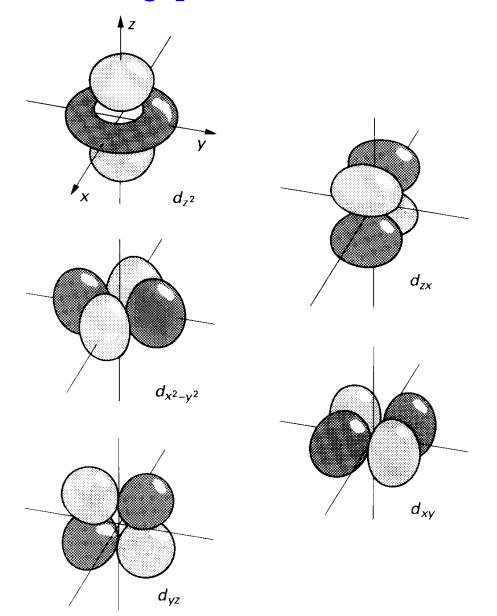
Distribution of Radial Probability density



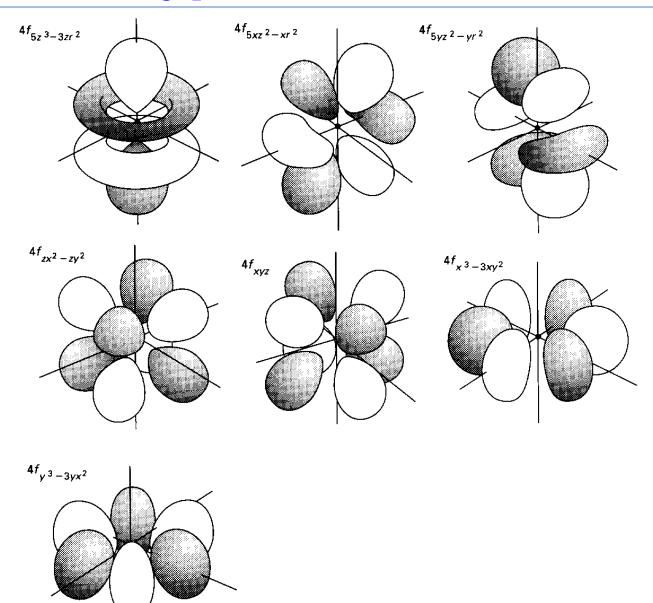
s & p-type Orbitals



d -type Orbitals



f -type Orbitals



Electronics Engineering

- 1. What are the conditions for which a wave function will be acceptable?
- 2. Using Planck distribution of energy spectral density for blackbody radiation, show that at room temperature the proportion of blue light (λ = 450 nm) in comparison to red light (λ = 700 nm) is insignificant. $\rho(\lambda,T) = \frac{8\pi hc}{\lambda^5 (e^{hc}/\lambda kT 1)}$
- 3. Evaluate the kinetic energy of the particle with wavefunction, where χ (chi) is a parameter. $\psi = (\cos \chi) e^{ikx} + (\sin \chi) e^{-ikx}$
- 4. Determine the probability of finding particle in an 1D box from 0.25L to 0.5L for n=2, L= 710 pm. Predict the probability from Classical Physics Standpoint.

Information Technology

- 1. State the postulates of quantum chemistry.
- 2. Calculate the wavelength of a neutron with a translational kinetic energy equal to kT at 300 K.
- 3. What is the uncertainty in position of a projectile (1.0 g) moving with a speed 1 μ ms⁻¹.
- 4. Write down the Hamiltonian operator of He₂ molecule.
- 5. Calculate kinetic energy for a particle in an 2D box for $n_x=n_y=1$, $L_x=L_y=288$ pm. Draw the wavefunction for the particle confined in this square well (2D box).