

# A L<sup>A</sup>T<sub>E</sub>X template and user guide

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## Abstract

L<sup>A</sup>T<sub>E</sub>X [1] is free typesetting software used to create technical documents in the sciences and mathematics. Most scientific publishers nowadays prefer—and may even require—documents to be prepared in L<sup>A</sup>T<sub>E</sub>X. What makes L<sup>A</sup>T<sub>E</sub>X better than other word processing programs is its ability to handle copious amounts of equations, tables, figures, citations, and cross-references with ease. It allows the author to focus on the content of the document, rather than wasting time on the minute formatting details. This document serves two purposes: it is both a guide to the most commonly used L<sup>A</sup>T<sub>E</sub>X commands, as well as a L<sup>A</sup>T<sub>E</sub>X document template, having itself been prepared in L<sup>A</sup>T<sub>E</sub>X.

## 1 Installing L<sup>A</sup>T<sub>E</sub>X

L<sup>A</sup>T<sub>E</sub>X is available to download free-of-charge. See the official website [1] for specific installation instructions for your operating system. The present document was created with TeXworks [2] operating on Windows. Additionally, a reference manager such as JabRef [3] is recommended for assembling lists of references. See Section 2 below for more details.

## 2 Citing references

An important part of any technical publication is citing previous work. And anyone who has ever created citations manually knows that it is nothing less than a *nightmare*. Fortunately, L<sup>A</sup>T<sub>E</sub>X takes all of the hard work out of it. Start by installing a free reference manager such as JabRef [3]. In JabRef, create (and save) a new library, and add entries for all of the sources you want to cite in your paper. For example, if you wanted to cite Newton’s *Principia* [4], you would create a new Book entry with the following field values:

- Title: Philosophiae Naturalis Principia Mathematica
- Publisher: Royal Society of London
- Year: 1687
- Author: Isaac Newton

You will also need to specify a “Bibtexkey,” which is a unique identifier that you will use to cite this work. For Newton’s *Principia*, we might choose Newton1687 as the Bibtexkey. Once you have all of your sources saved in your library, you are ready to go. To create a list of references at the end of your paper, use the `\bibliography{}` command. For example, if you have saved your JabRef library as references.bib, put `\bibliography{references}` at the end of your document where you want the list of references to appear. Then, wherever you want to cite a reference, use the `\cite{}` command along with the Bibtexkey. If we want to cite Newton’s *Principia*, we would type `\cite{Newton1687}`. You can also specify the style of the list of references with a style file. The present document uses the style file asmems4.bst, so we include the command `\bibliographystyle{asmems4}` on the line above `\bibliography{references}`.

### 3 Equations

One of L<sup>A</sup>T<sub>E</sub>X's greatest strengths is its ability to render really nice-looking equations. Numbered equations are created using the `equation` environment, enclosed between the commands `\begin{equation}` and `\end{equation}`. In-line equations are placed between two dollar signs `$$`. For example, Newton's second law of motion states that

$$\mathbf{F} = \frac{d\mathbf{p}}{dt}, \quad (1)$$

where  $\mathbf{F}$  is the net force on a particle,  $\mathbf{p}$  is the particle's linear momentum, and  $t$  is time [4].

Maxwell's equations of electromagnetism are as follows:

$$\text{Gauss's first law:} \quad \oiint_{\partial V} \mathbf{E} \cdot \hat{\mathbf{n}} dA = q/\epsilon_0 \quad \nabla \cdot \mathbf{E} = \rho/\epsilon_0 \quad (2)$$

$$\text{Gauss's second law:} \quad \oiint_{\partial V} \mathbf{B} \cdot \hat{\mathbf{n}} dA = 0 \quad \nabla \cdot \mathbf{B} = 0 \quad (3)$$

$$\text{Faraday's law:} \quad \oint_{\partial S} \mathbf{E} \cdot d\mathbf{r} = -\frac{d}{dt} \iint_S \mathbf{B} \cdot \hat{\mathbf{n}} dA \quad \nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \quad (4)$$

$$\text{Ampere's law:} \quad \oint_{\partial S} \mathbf{B} \cdot d\mathbf{r} = \mu_0 j + \mu_0 \epsilon_0 \frac{d}{dt} \iint_S \mathbf{E} \cdot \hat{\mathbf{n}} dA \quad \nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \epsilon_0 \frac{\partial \mathbf{E}}{\partial t} \quad (5)$$

Here  $\mathbf{E}$  and  $\mathbf{B}$  are the electric and magnetic fields, respectively;  $\mu_0$  and  $\epsilon_0$  are the permeability and permittivity of free space, respectively;  $\partial V$  is the closed surface bounding an arbitrary volume  $V$ ;  $q$  is the total charge enclosed within  $V$ ;  $\rho$  is the charge density over  $V$ ;  $\partial S$  is the closed curve bounding an arbitrary surface  $S$ ;  $j$  is the total current piercing  $S$ ;  $\mathbf{J}$  is the current density over  $S$ ;  $\nabla$  is the gradient operator; and  $t$  is time [5]. Note that L<sup>A</sup>T<sub>E</sub>X automatically numbers equations sequentially, without any work required on your part.

The Einstein field equations, which govern how space and time curve in the presence of matter, state that

$$R_{\mu\nu} + \left(\Lambda - \frac{1}{2}R\right) g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}, \quad (6)$$

where  $R_{\mu\nu}$  are the components of the Ricci curvature tensor,  $\Lambda$  is the cosmological constant,  $R$  is the scalar curvature,  $g_{\mu\nu}$  are the components of the metric tensor,  $G$  is Newton's gravitational constant,  $c$  is the speed of light in a vacuum, and  $T_{\mu\nu}$  are the components of the stress-energy tensor [6].

The Schrödinger equation of quantum mechanics states that

$$-\frac{\hbar^2}{2m} \nabla^2 \Psi + V\Psi = i\hbar \frac{\partial \Psi}{\partial t}, \quad (7)$$

where  $\Psi(\mathbf{r}, t)$  is a particle's wave function,  $m$  is the particle's mass,  $V(\mathbf{r}, t)$  is the particle's potential energy,  $i$  is the imaginary unit,  $\hbar$  is the reduced Planck constant,  $\nabla$  is the gradient operator, and  $t$  is time [7, 8].

Some common mathematical symbols can be found in Appendix A. If you can't find what you're looking for, try Detexify [9]. You can draw the symbol you are looking for and it will generate several suggestions.

## 4 Floating environments: tables and figures

Tables and figures are examples of *floating environments*, which tend to “float” around as you edit a document.

### 4.1 Tables

Tables are created using the `table` and `tabular` environments. The location of the table is specified by a string at the beginning of the `table` environment: `t` stands for ‘top,’ `b` stands for ‘bottom,’ `h` stands for ‘here,’ and `p` stands for new ‘page.’ If you include more than one option,  $\text{\LaTeX}$  will automatically select the one it deems best. You can also force it to use a particular option by including an exclamation point `!` at the beginning of the location string. The text justification in each column is specified by a string at the beginning of the `tabular` environment: `l` stands for ‘left-justified,’ `c` stands for ‘centered,’ and `r` stands for ‘right-justified.’ Columns are separated by the ampersand `&` character, and rows are separated by `\\`. In technical writing, table captions go *above* the corresponding tables. For example, see Table 1 for the Greek alphabet with the corresponding  $\text{\LaTeX}$  syntax.

Table 1. The Greek alphabet. Note that the table caption goes *above* the table.

Letter	Capital	Lowercase	$\text{\LaTeX}$ syntax
Alpha	A	$\alpha$	A, <code>\alpha</code>
Beta	B	$\beta$	B, <code>\beta</code>
Gamma	$\Gamma$	$\gamma$	<code>\Gamma</code> , <code>\gamma</code>
Delta	$\Delta$	$\delta$	<code>\Delta</code> , <code>\delta</code>
Epsilon	E	$\varepsilon$	E, <code>\varepsilon</code>
Zeta	Z	$\zeta$	Z, <code>\zeta</code>
Eta	H	$\eta$	H, <code>\eta</code>
Theta	$\Theta$	$\theta$	<code>\Theta</code> , <code>\theta</code>
Iota	I	$\iota$	I, <code>\iota</code>
Kappa	K	$\kappa$	K, <code>\kappa</code>
Lambda	$\Lambda$	$\lambda$	<code>\Lambda</code> , <code>\lambda</code>
Mu	M	$\mu$	M, <code>\mu</code>
Nu	N	$\nu$	N, <code>\nu</code>
Xi	$\Xi$	$\xi$	<code>\Xi</code> , <code>\xi</code>
Omicron	O	$\omicron$	O, <code>\omicron</code>
Pi	$\Pi$	$\pi$	<code>\Pi</code> , <code>\pi</code>
Rho	P	$\rho$	P, <code>\rho</code>
Sigma	$\Sigma$	$\sigma$	<code>\Sigma</code> , <code>\sigma</code>
Tau	T	$\tau$	T, <code>\tau</code>
Upsilon	Y	$\upsilon$	Y, <code>\upsilon</code>
Phi	$\Phi$	$\phi$	<code>\Phi</code> , <code>\phi</code>
Chi	X	$\chi$	X, <code>\chi</code>
Psi	$\Psi$	$\psi$	<code>\Psi</code> , <code>\psi</code>
Omega	$\Omega$	$\omega$	<code>\Omega</code> , <code>\omega</code>

## 4.2 Figures

Figures are created using the `figure` environment along with the `\includegraphics[]{}` command. The location of the figure is specified by a string at the beginning of the `figure` environment: `t` stands for ‘top,’ `b` stands for ‘bottom,’ `h` stands for ‘here,’ and `p` stands for new ‘page.’ If you include more than one option, L<sup>A</sup>T<sub>E</sub>X will automatically select the one it deems best. You can also force it to use a particular option by including an exclamation point `!` at the beginning of the location string. In technical writing, figure captions go *below* the corresponding figures. For example, see Figure 1 for a portrait of Issac Newton.

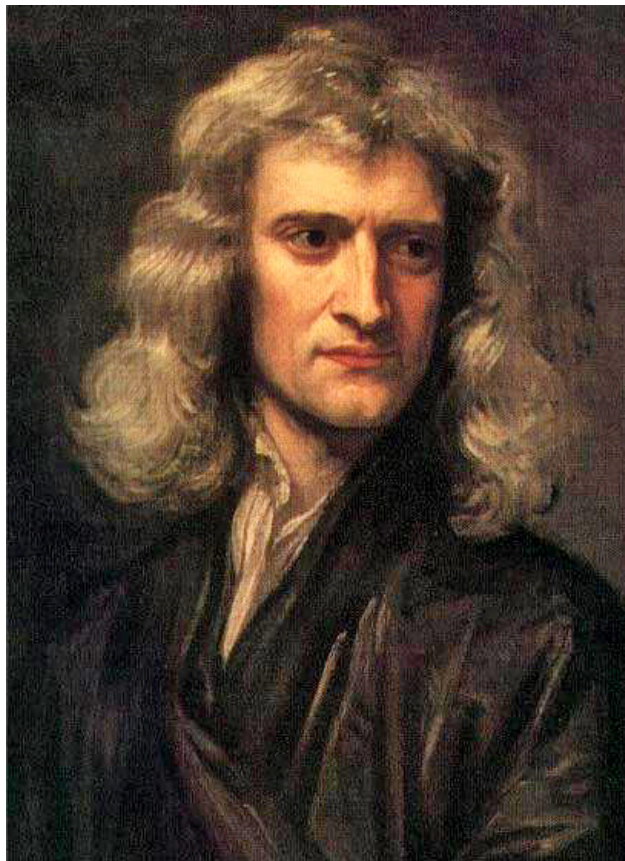
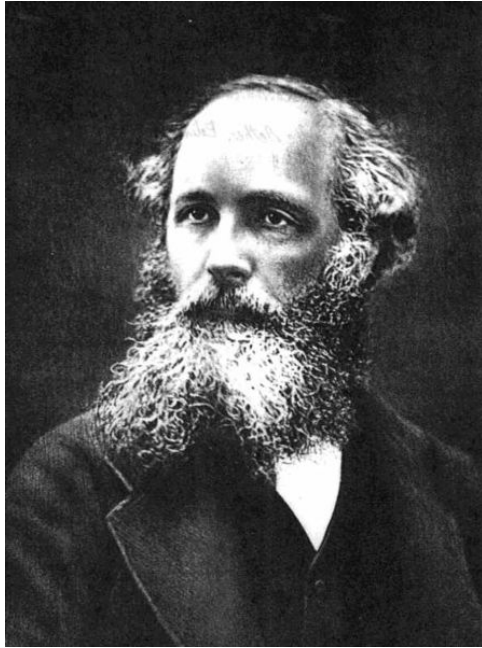
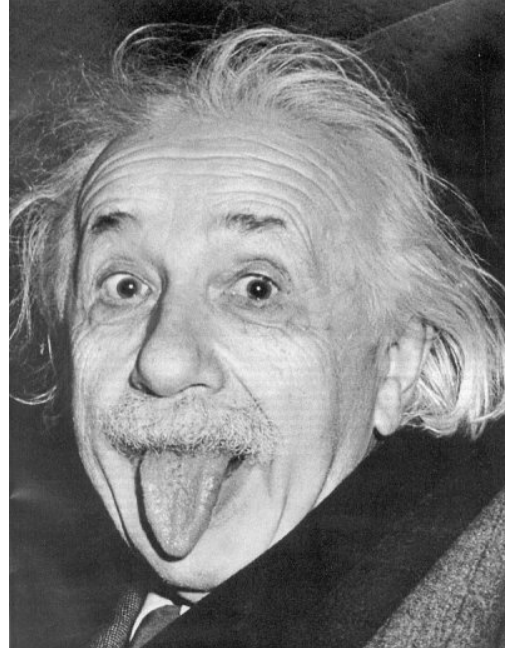


Figure 1. Portrait of Sir Isaac Newton [10]. Note that the figure caption goes *below* the figure.

You can make figures with multiple subfigures using the `subfigure` environment. For example, see Figure 2 for photographs of James Clerk Maxwell, Albert Einstein, and Erwin Schrödinger. Each subfigure can have its own caption and can be referenced independently. For more details about cross-referencing subfigures, see Section 5.



(a)



(b)



(c)

Figure 2. Photographs of (a) James Clerk Maxwell [11], (b) Albert Einstein [12], and (c) Erwin Schrödinger [13].

## 5 Cross-references

Cross-referencing is the act of referring to a numbered object, such as a figure, table, equation, or section. For example, you are currently reading Section 5 of this document, you probably just finished looking at Figure 2, and way back in Section 3, you saw Equations (1)–(7). As with citations, creating cross-references manually is a recipe for disaster. Imagine that you have just finished writing a report with over one hundred equations, when you decide that the last equation should have been first. If you were doing the cross-references manually, you would have to go through the entire document and change the number of each equation, not only for the equations themselves, but also everywhere you cross-referenced an equation!

Fortunately,  $\text{\LaTeX}$  makes cross-references a breeze. Anything you want to cross-reference can be assigned a label using the `\label{ }` command. The label is similar to the BibTeXkey of a reference: it is a unique identifier used to refer to the corresponding figure, table, equation, section, or whatever. You can then refer to the object using the `\eqref{ }` command in the case of equations, and the `\ref{ }` command in the case of everything else. For example, you might label Equation (1) `\label{eq:Newton2}`, and refer to it as `\eqref{eq:Newton2}`. Similarly, you might label Figure 1 `\label{fig:Newton}`, and refer to it as `Figure~\ref{fig:Newton}`. By assigning objects labels rather than specific numbers,  $\text{\LaTeX}$  can automatically renumber all the objects when their order is changed, or when a new object is inserted in the middle. In the case of figures with multiple subfigures, you can label and subsequently cross-reference the entire figure as a whole—e.g., Figure 2—or each subfigure individually—e.g., Figure 2(b).

## References

- [1] The  $\text{\LaTeX}$  Project. <https://www.latex-project.org/>.
- [2] Texworks. <http://www.tug.org/texworks/>.
- [3] Jabref. <http://www.jabref.org/>.
- [4] Newton, I., 1687. *Philosophiae Naturalis Principia Mathematica*. Royal Society of London.
- [5] Maxwell, J. C., 1865. “A dynamical theory of the electromagnetic field”. *Philosophical Transactions of the Royal Society of London*, **155**, pp. 459–512.
- [6] Einstein, A., 1916. “Die Grundlage der allgemeinen Relativitätstheorie (The basis of the general theory of relativity)”. *Annalen der Physik*, **354**(7), pp. 769–822.
- [7] Schrödinger, E., 1926. “Quantisierung als Eigenwertproblem (Quantization as eigenvalue problem)”. *Annalen der Physik*, **79**, pp. 361–376, 489–527, 734–756.
- [8] Schrödinger, E., 1926. “Quantisierung als Eigenwertproblem (Quantization as eigenvalue problem)”. *Annalen der Physik*, **81**, p. 109.
- [9] Detexify. <http://detexify.kirelabs.org/classify.html>.
- [10] Isaac Newton. <https://commons.wikimedia.org/wiki/File:GodfreyKneller-IsaacNewton-1689.jpg>.
- [11] James Clerk Maxwell. [https://commons.wikimedia.org/wiki/File:James\\_clerk\\_maxwell.jpg](https://commons.wikimedia.org/wiki/File:James_clerk_maxwell.jpg).
- [12] Albert Einstein. <https://www.flickr.com/photos/alg0angelical/3307637331>.
- [13] Erwin Schrödinger. [https://commons.wikimedia.org/wiki/File:Erwin\\_Schr%C3%B6dinger\\_\(1933\).jpg](https://commons.wikimedia.org/wiki/File:Erwin_Schr%C3%B6dinger_(1933).jpg).

## Appendix A. Common mathematical symbols

Table 2 shows some common mathematical symbols and their  $\text{\LaTeX}$  syntax. Remember: if all else fails, use Detexify [9]!

Table 2. Some common mathematical symbols.

Description	Symbol(s)	$\text{\LaTeX}$ syntax
Greater than or equal to	$\geq$	<code>\geq</code>
Less than or equal to	$\leq$	<code>\leq</code>
Much greater than	$\gg$	<code>\gg</code>
Much less than	$\ll$	<code>\ll</code>
Identically equal to	$\equiv$	<code>\equiv</code>
Approximately equal to	$\approx$	<code>\approx</code>
Element of	$\in$	<code>\in</code>
Set union	$\cup$	<code>\bigcup</code>
Set intersection	$\cap$	<code>\bigcap</code>
Subset of	$\subset$	<code>\subset</code>
Similar to	$\sim$	<code>\sim</code>
Proportional to	$\propto$	<code>\propto</code>
Maps to	$\mapsto$	<code>\mapsto</code>
Right arrow	$\rightarrow$	<code>\rightarrow</code>
Implies	$\Rightarrow$	<code>\Rightarrow</code>
Square root	$\sqrt{x}$	<code>\sqrt{x}</code>
$n$ th root	$\sqrt[n]{x}$	<code>\sqrt[n]{x}</code>
Infinity	$\infty$	<code>\infty</code>
Summation	$\sum_{n=1}^{\infty}$	<code>\sum_{n=1}^{\infty}</code>
Product	$\prod_{n=1}^{\infty}$	<code>\prod_{n=1}^{\infty}</code>
Time derivative (Newton)	$\dot{x}$	<code>\dot{x}</code>
Partial differential	$\partial$	<code>\partial</code>
Integral	$\int_a^b$	<code>\int_a^b</code>
Double integral	$\iint_S$	<code>\iint_S</code>
Triple integral	$\iiint_V$	<code>\iiint_V</code>
Closed line integral	$\oint_{\Gamma}$	<code>\oint_{\Gamma}</code>
Closed surface integral	$\oiint_S$	<code>\oiint_S</code>
Cartesian basis vectors	$\hat{i}, \hat{j}, \hat{k}$	<code>\hat{i}, \hat{j}, \hat{k}</code>
Gradient	$\nabla$	<code>\nabla</code>
Dot product	$\cdot$	<code>\cdot</code>
Cross product	$\times$	<code>\times</code>
Tensor product	$\otimes$	<code>\otimes</code>