

PROBLEM 1

A $\epsilon = -\frac{\mu}{2a}$

$a_{ISS} = 6378.137 + 422 \text{ km}$

$a_{GPS} = 6378.137 + 20200 \text{ km}$

$\epsilon_{ISS} < \epsilon_{GPS} \rightarrow \boxed{\text{FALSE}}$

$a \uparrow \rightarrow \epsilon \uparrow$

B $P = 2\pi \sqrt{\frac{a^3}{\mu}}$

$a \uparrow \rightarrow P \uparrow$

$P_{ISS} < P_{GPS} \rightarrow \boxed{\text{FALSE}}$

C NEWTON \rightarrow UNIVERSAL LAW OF GRAVITATION

D KEPLER \rightarrow FIRST LAW

E PTOLEMY / COPERNICUS

F 1. THE ORBIT OF EACH PLANET IS AN ELLIPSE
WITH THE SUN AT A FOCUS

2. THE LINE JOINING THE PLANET TO THE SUN
SWEEPS OUT EQUAL AREAS IN EQUAL TIMES

3. THE SQUARE OF THE PERIOD OF A PLANET
IS PROPORTIONAL TO THE CUBE OF ITS MEAN
DISTANCE FROM THE SUN.

$$\ddot{r} + \frac{\mu}{r^3} r = 0$$

$$r = \frac{a(1-e^2)}{1+e \cos r}$$

C $r_a = 25000 \text{ km}$
 $r_p = 8000 \text{ km}$

$$v_a = 2.3 \text{ km/sec}$$

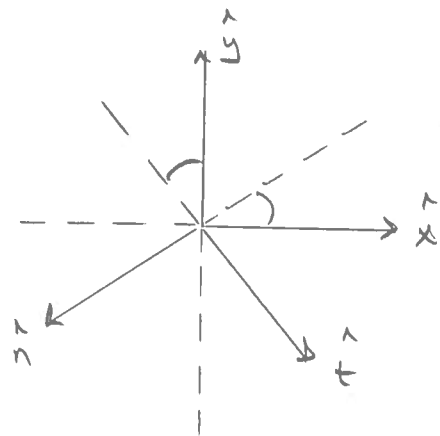
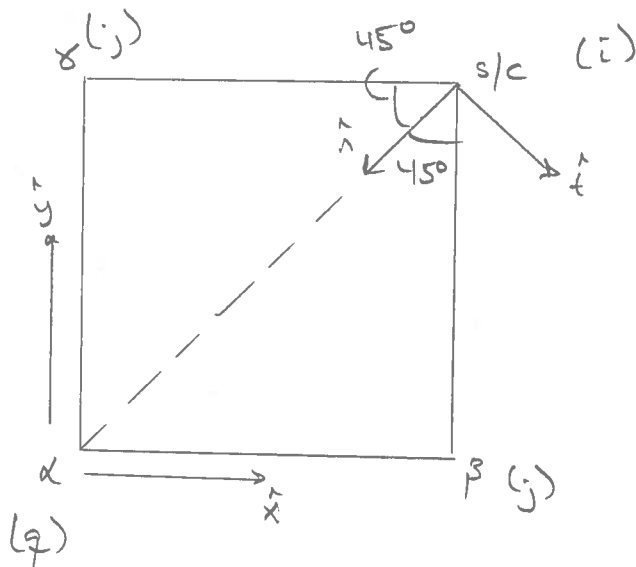
$$v_p = 8.7 \text{ km/sec}$$

$$E = -\frac{\mu}{2a} = \frac{v^2}{2} - \frac{\mu}{r}$$

$$v = \sqrt{2\left(\frac{\mu}{r} - \frac{\mu}{2a}\right)}$$

$$v_p > v_a \rightarrow \boxed{\text{FALSE}}$$

2017 MIDTERM SOLUTION



REFERENCE FRAME

$$\hat{r}_{\alpha s} = -\hat{n}$$

$$\hat{r}_{s\beta} = -\hat{y}$$

$$\hat{r}_{\alpha\beta} = \hat{x}$$

$$\hat{r}_{s\gamma} = -\hat{x}$$

$$\hat{r}_{\alpha\gamma} = \hat{y}$$

A RELATIVE N-BODY EOMS.

$$\ddot{\vec{r}}_{qi} + \frac{m_i + m_q}{r_{qi}^2} \hat{r}_{qi} = \sum_{j=1, j \neq i}^n m_j \left(\frac{\hat{r}_{ij}}{r_{ij}^2} - \frac{\hat{r}_{ji}}{r_{ji}^2} \right)$$

MOTION OF S/C WRT ALPHA.

$$\ddot{\vec{r}}_{\alpha s} = - \underbrace{\frac{(m_s + m_\alpha)}{r_{\alpha s}^2}}_{\text{DOMINANT}} \hat{r}_{\alpha s} + m_\beta \underbrace{\left(\frac{\hat{r}_{s\beta}}{r_{s\beta}^2} - \frac{\hat{r}_{\alpha\beta}}{r_{\alpha\beta}^2} \right)}_{\text{DIRECT} \quad \text{INDIRECT}} + m_\gamma \underbrace{\left(\frac{\hat{r}_{s\gamma}}{r_{s\gamma}^2} - \frac{\hat{r}_{\alpha\gamma}}{r_{\alpha\gamma}^2} \right)}_{\text{DIRECT} \quad \text{INDIRECT}}$$

3 DOMINANT ACCEL : $-\frac{2\mu}{(\sqrt{2}d)^2} (-\hat{n}) = \frac{\mu}{d^2} \hat{n} = \bar{A}_D$

DIRECT : $\frac{\mu}{d^2} (-\hat{y}) + \frac{\mu}{d^2} (-\hat{x}) = \bar{A}_{\text{DIRECT}} \left\{ \begin{array}{l} \text{MAG: } \frac{\sqrt{2}\mu}{d^2} \\ \text{DIRECTION: } +\hat{n} \end{array} \right.$

$$\text{INDIRECT ACCEL: } -\frac{\mu}{d^2} (\hat{x}) - \frac{\mu}{d^2} (\hat{y}) = \bar{A}_{\text{INDIRECT}} \left\{ \begin{array}{l} \text{MAG: } \frac{\sqrt{2} \mu}{d^2} \\ \text{DIR: } + \hat{n} \end{array} \right.$$

C TOTAL ACCELERATION:

$$\begin{aligned} & \frac{\mu}{d^2} \hat{n} - \frac{2\mu}{d^2} \hat{x} - \frac{2\mu}{d^2} \hat{y} \\ &= \left(\underbrace{\frac{\mu}{d^2}}_{\text{DOMINANT}} + \underbrace{2\frac{\sqrt{2}\mu}{d^2}}_{\text{PERTURBING}} \right) \hat{n} \quad \left\{ \begin{array}{l} \text{MAG: } 3.828 \frac{\mu}{d^2} \\ \text{DIRECTION: } + \hat{n} \end{array} \right. \end{aligned}$$

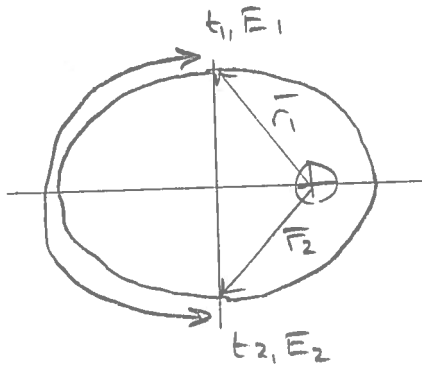
D NET PERTURBING ACCEL IN $+\hat{n}$ DIRECTION
 \Rightarrow TOWARDS ALPHA

PERTURBING ACCEL $>$ DOMINANT ACCEL.

$$\frac{P_{\text{EXT}}}{D_{\text{OM}}} = 3.828 !$$

A MODEL COMPRISED OF ONLY TWO BODIES (ALPHA + S/C)
 IS NOT REASONABLE. NEGLECTING THE PERTURBING
 ACCELERATION IS NOT A VALID ASSUMPTION.

A



$$E_1 = 90^\circ \quad E_2 = 270^\circ$$

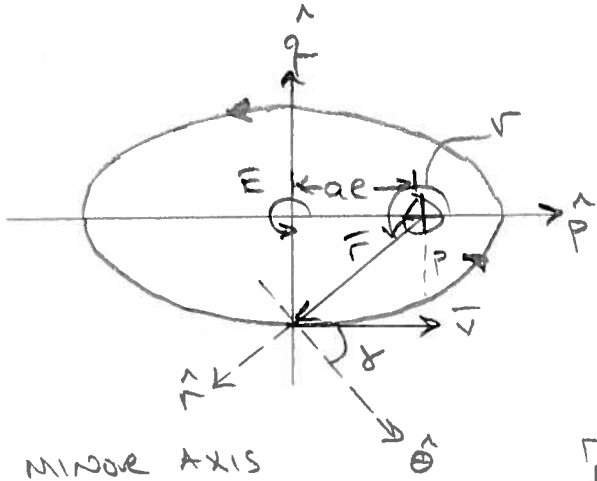
$$r(t-T) = E - e \sin E$$

$$\left. \begin{aligned} t_1 - T &= \frac{1}{n} \left(\frac{\pi}{2} - e \right) \\ t_2 - T &= \frac{1}{n} \left(\frac{3\pi}{2} + e \right) \end{aligned} \right\} (t_2 - t_1) = \frac{1}{n} (\pi + 2e)$$

$$P = \frac{2\pi}{n} \Rightarrow \boxed{\frac{t_2 - t_1}{P} = \frac{\pi + 2e}{2\pi}}$$

B $e = 0.75 \Rightarrow \frac{t_2 - t_1}{P} = 0.7387 \Rightarrow \boxed{73.9\%}$

2017 MIDTERM SOLUTIONS



MINOR AXIS

$$\boxed{\Gamma = a = BR} \quad \frac{a}{p} = 8$$

$$V = \sqrt{\frac{\mu}{R}}$$

$$\vec{r} = -4R \hat{p} - 4\sqrt{3} \hat{q} \rightarrow r = 8R$$

$$|\vec{v}| = 3 \frac{\text{rad}}{\text{sec}} \quad \vec{v} = v \hat{r}$$

$$\uparrow E = 270^\circ = -90^\circ$$

$$q_e = 4R \Rightarrow \boxed{e = 1/2}$$

$$\boxed{b = 4\sqrt{3} R} \quad \frac{b}{R} = 4\sqrt{3}$$

$$p = a(1 - e^2) = \boxed{6R = P} \quad \frac{P}{R} = 6$$

$$\Sigma = \frac{V^2}{2} - \frac{\mu}{r} \Rightarrow \Sigma = -\frac{V^2}{2} \Rightarrow \boxed{\Sigma = -4.5 \frac{\text{km}^2}{\text{sec}^2}}$$

$$h^2 = \mu p \quad \text{BUT} \quad v^2 = \frac{\mu}{r} = \frac{\mu}{8R} \Rightarrow \underline{h^2 = 432 R^2}$$

$$\frac{h}{R} = 12\sqrt{3} \quad \frac{\text{km}}{\text{sec}^2}$$

$$V = 180^\circ + 60^\circ = 240^\circ$$

$$\tan \delta = \frac{e}{\sqrt{1-e^2}} \quad \text{① MINOR AXIS}$$

$$\delta = -30^\circ \text{ BELOW LOCAL HORIZON}$$