

CONIC SECTION

$$r = \frac{h^2/\mu}{1 + e \cos \theta} = \frac{p}{1 + e \cos \theta}$$

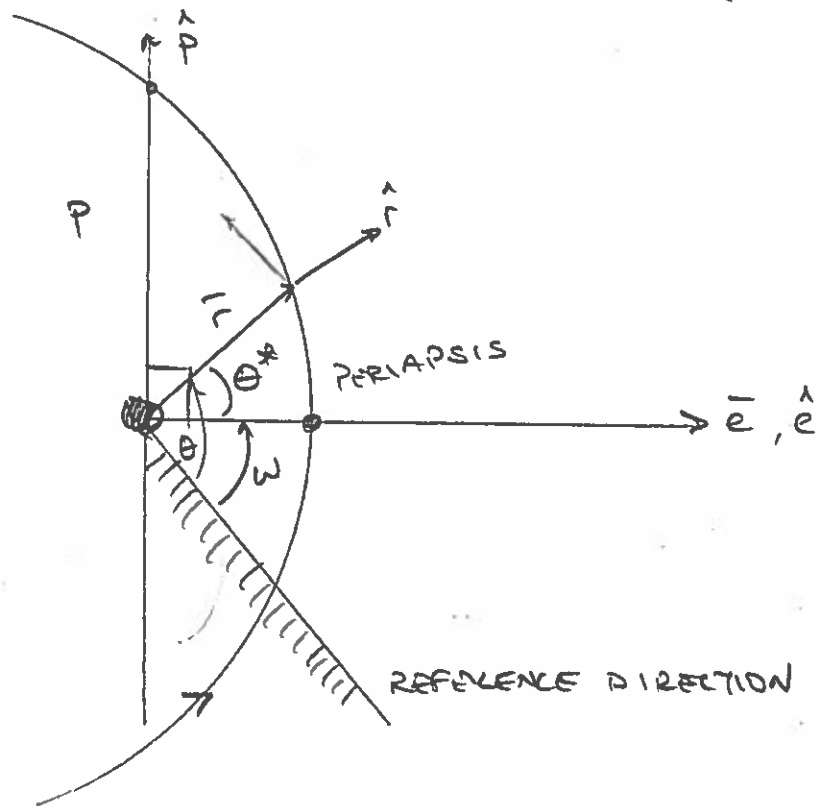
STANDARD POLAR EQ. OF
A CONIC SECTION

$$p = \frac{h^2}{\mu}$$

$$e = \sqrt{1 + \frac{2\epsilon h^2}{\mu}}$$

SEMI-LATUS RECTUM

ECCENTRICITY



$2p = \text{LATJUS RECTUM}$

θ^* - TRUE ANOMALY (P, V)

$$\theta - \omega = \theta^* = \nu$$

$$h = \sqrt{\mu p}$$

$$\epsilon = \frac{-\mu^2}{2h^2} (1 - e^2)$$

$$\text{DEFINE } a = \frac{p}{1 - e^2}$$

SEMI MAJOR AXIS

$$\epsilon = \left[-\frac{\mu}{2a} = \frac{v^2}{2} - \frac{\mu}{r} \right]$$

JULIAN DATE - CONTINUOUS COUNT OF DAYS SINCE 1 JAN 4713 B

JDO = NOON 1 JAN 4713 BC

JD2451545 = 12:00 UTC JAN 1 2000

$$\text{MJD} = \text{JD} - 2451545$$

USE ASTRO LIBRARY TO CONVERT JD.

CONIC SECTIONS DEFINED BY TYPES

$0 \leq e < 1$	$e < 0$	ELLIPSE
$e = 1$	$e = 0$	PARABOLA
$e > 1$	$e > 0$	HYPERBOLA

γ - FLIGHT PATH ANGLE

$$0 < \theta < 180 \rightarrow \gamma > 0$$

$$180 < \theta < 360 \rightarrow \gamma < 0$$

CIRCLE SPECIAL CASE $e=0$

$$a=r=p \quad \boxed{\epsilon = -\frac{\mu}{2a} = \frac{v^2}{2} - \frac{\mu}{r}} \rightarrow v_c = \sqrt{\frac{\mu}{r}} \quad \text{CIRCULAR ORBIT SPEED AT } r$$

GENERAL ELLIPSE

$$\frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a}$$

$$v^2 = \frac{2\mu}{r} - \frac{\mu}{a} = 2v_c^2 - \frac{\mu}{a}$$

VELOCITY OF ELLIPSE IS ALWAYS LESS THAN $\sqrt{2}$ X CIRCULAR VELOCITY

$$\boxed{v < \sqrt{2} v_c} \quad \text{FIGURE OUT IF ORBIT IS CLOSE}$$

PERIOD

$$\frac{dA}{dt} = \frac{h}{2} \rightarrow dt = \frac{2}{h} dA \quad \text{AREA VELOCITY}$$

$$P = \frac{2}{h} (\pi ab) \quad \leftarrow \text{AREA OF ELLIPSE}$$

$$\begin{cases} b = a(1-e^2)^{1/2} \\ p = a(1-e^2) \\ h = \sqrt{\mu p} \end{cases}$$

$$\boxed{P = 2\pi \sqrt{\frac{a^3}{\mu}}}$$

PERIOD

BIG P - PERIOD

LITTLE p = SEMI-MAJOR AXIS

ENERGY

$$\epsilon = -\frac{\mu}{2a} = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{1}{2} \left(\frac{\mu}{h} \right)^2 (1-e^2)$$

$$e \uparrow \Rightarrow e \uparrow$$

CONIC SECTIONS

RADIUS / DISTANCE - $r = \frac{h^2 / \mu}{1 + e \cos \sqrt{}}$

CONIC EQUATION

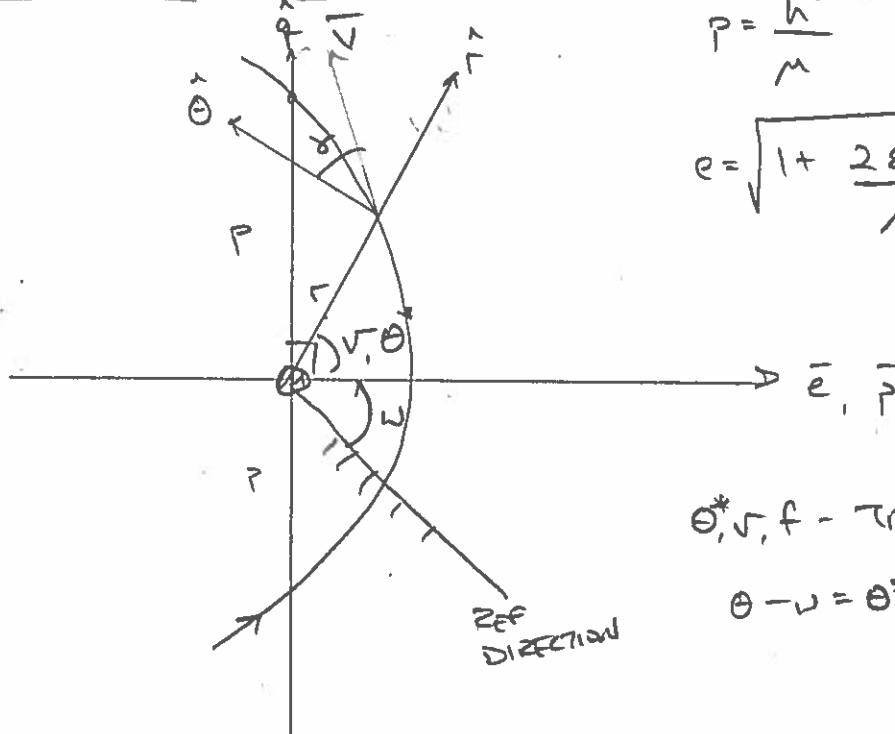
$$\vec{r} = r \hat{r}$$

$$\vec{v} = \dot{r} \hat{r} + r \dot{\theta} \hat{\theta} = \frac{\mu}{h} e \sin \theta \hat{r} + \frac{\mu}{h} (1 + e \cos \theta) \hat{\theta}$$

$$\tan \gamma = \frac{e \sin \theta}{1 + e \cos \theta}$$

$$p = \frac{h^2}{\mu} \quad \text{SEMI-LATUS RECTUM}$$

$$e = \sqrt{1 + \frac{2 \epsilon h^2}{\mu}} \quad \text{ECCENTRICITY}$$



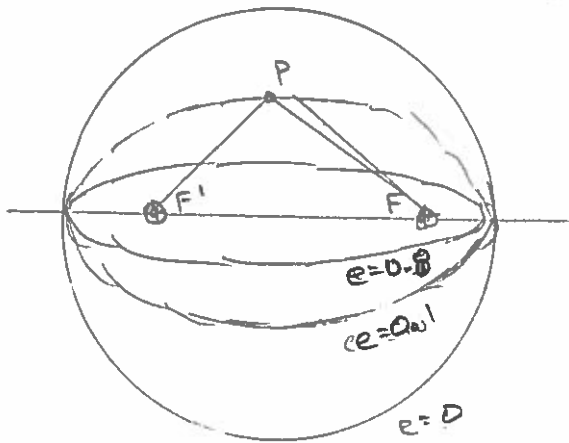
θ^*, ν, f - TRUE ANOMALY

$$\theta - \nu = \theta^* = \nu$$



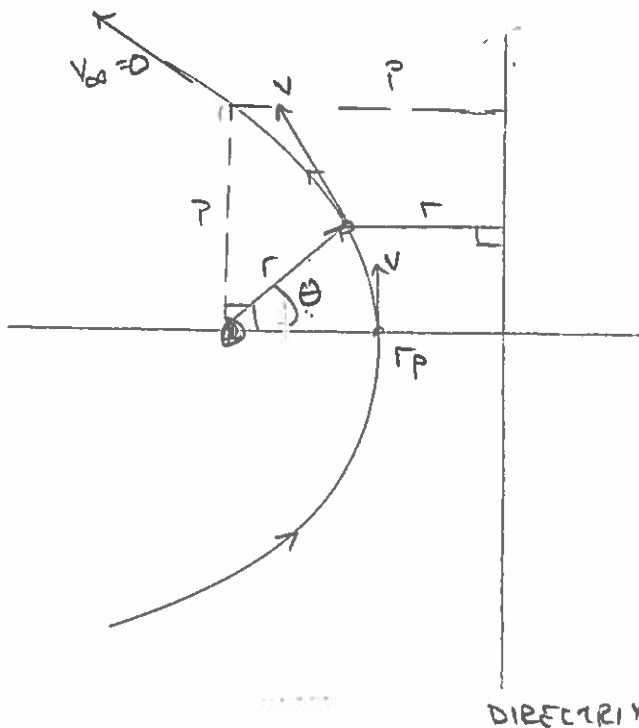
EX ELLIPSES

$$\overline{PF} + \overline{PF'} = 2a \quad \text{CONSTANT}$$



PARABOLA $e=1$ $a=\infty$ $\epsilon=0$ SPECIFIC ENERGY NOT CLOSED

IDEALIZED ORBIT - NOT REALLY FEASIBLE IN REAL LIFE



MINIMUM ENERGY TO ESCAPE CENTRAL BODY

$$-\frac{\mu}{2a} = \epsilon = \frac{v^2}{2} - \frac{\mu}{r} = 0 \rightarrow v^2 = \frac{2\mu}{r} \Rightarrow \boxed{v_e = \sqrt{2} v_c} \quad v_c = \sqrt{\frac{\mu}{r}}$$

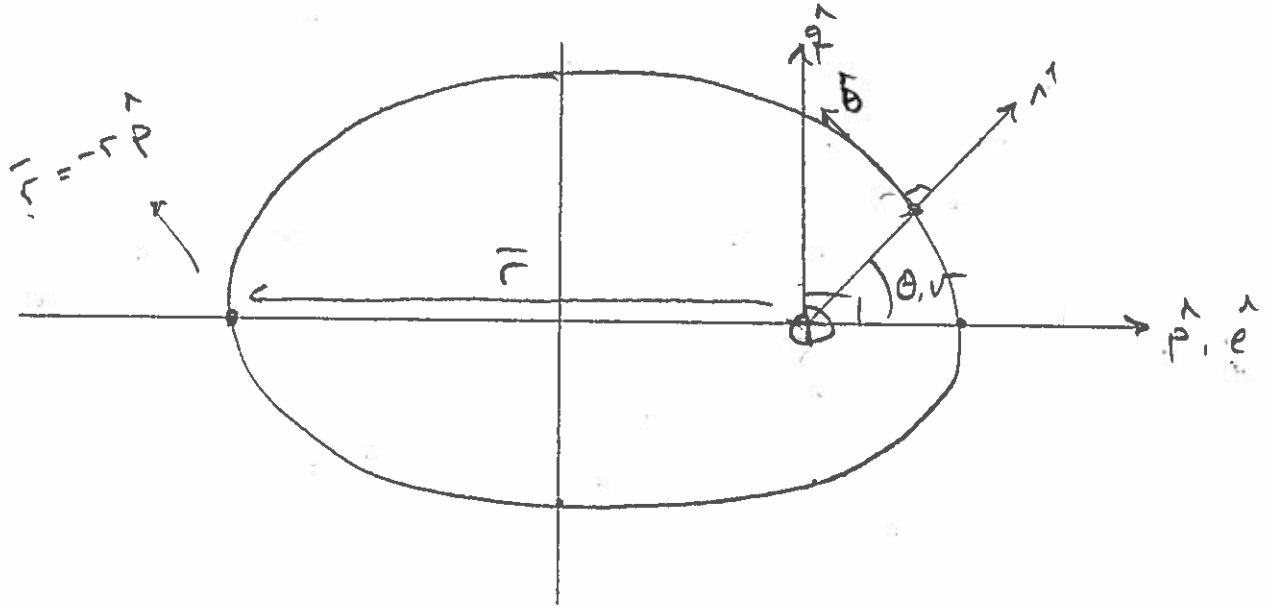
ESCAPE SPEED AT DISTANCE r

$$\frac{v_\infty^2}{2} - \frac{\mu}{r_\infty} = 0 \rightarrow v_\infty = 0$$

CAN JUST BARELY ESCAPE GRAVITY INFLUENCE

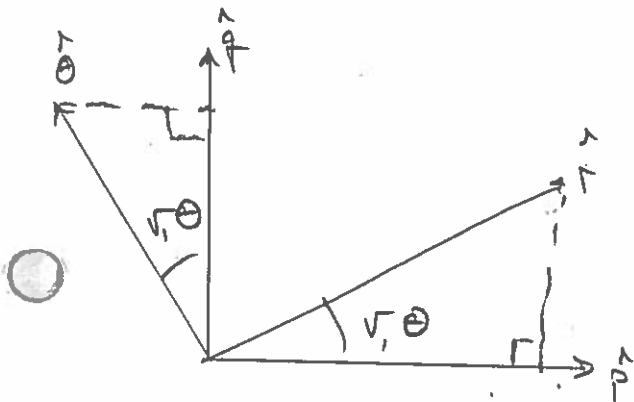
(2) PERIFOCAL REFERENCE FRAME ($\hat{p}, \hat{q}, \hat{w}$) $\hat{w} = \hat{h}$

- REFERENCE FRAME THAT IS FIXED TO THE ORBIT
- ORIGIN AT ATTRACTING FOCUS OF ORBIT
- \hat{p} - POINTS TOWARDS PERIAPSIS
- \hat{q} - POINTS AT $v=90^\circ$, ALONG SEMI-MAJOR AXIS DIRECTION
- \hat{w} - ALONG ANGULAR MOMENTUM VECTOR



(1) LOCAL VERTICAL / LOCAL HORIZONTAL FRAME

- ROTATING FRAME ATTACHED TO SATELLITE
- \hat{r} - ALONG THE POSITION VECTOR - RADIALLY OUTWARD
- $\hat{\theta}$ - IN ORBIT PLANE - \perp TO \hat{r} - LOCAL HORIZONTAL PLANE
- \hat{h} - ALONG ANGULAR MOMENTUM VECTOR



$$R_{LVLH \rightarrow PQW} = \begin{bmatrix} \hat{r} & \hat{\theta} & \hat{h} \end{bmatrix} \rightarrow \begin{bmatrix} \hat{p} & \hat{q} & \hat{w} \end{bmatrix}$$

$$\begin{bmatrix} \hat{p} \\ \hat{q} \\ \hat{w} \end{bmatrix} = \begin{bmatrix} c\sqrt{e} & -s\sqrt{e} & 0 \\ s\sqrt{e} & c\sqrt{e} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \hat{r} \\ \hat{\theta} \\ \hat{h} \end{bmatrix}$$

— WE ALREADY HAVE POSITION/VELOCITY IN $\hat{r} \hat{\theta} \hat{h}$
FROM TRAJECTORY EQ

$$r = \frac{h^2/m}{1 + e \cos \psi}$$

— POSITION $\vec{r} = r \hat{r} + 0 \hat{h}$

VELOCITY $\vec{v} = \frac{M}{h} e \sin \theta \hat{r} + \frac{M}{h} (1 + e \cos \theta) \hat{\theta} + 0 \hat{h}$

TRANSFORM TO $\hat{p} \hat{q} \hat{w}$

$$\vec{r}_{pqw} = R_{LVLT2} \cdot \vec{r}_{LVLT} = \boxed{r \cos \theta \hat{p} + r \sin \theta \hat{q} = \vec{r}_{pqw}}$$

$$\vec{v}_{pqw} = R_{LVLT2} \cdot \vec{v}_{LVLT} =$$

$$= \left[\frac{M}{h} e \sin \theta \cdot \cos \theta + \frac{M}{h} (1 + e \cos \theta) \cdot -\sin \theta \right] \hat{p}$$

$$+ \left[\frac{M}{h} e \sin \theta \cdot \sin \theta + \frac{M}{h} (1 + e \cos \theta) \cdot \cos \theta \right] \hat{q}$$

$$+ 0 \hat{h}$$

$$\boxed{\vec{v}_{pqw} = -\frac{M}{h} \sin \theta \hat{p} + \frac{M}{h} (e + \cos \theta) \hat{q}}$$

EXAMPLE FCN - KEPLER MODULE

SPECIFIC MECHANICAL ENERGY

$$\mathcal{E} = \frac{V^2}{2} - \frac{\mu}{r} \quad \mu, V, r \rightarrow \text{FCN} \rightarrow \mathcal{E}$$

$$\text{TESTS: } V=0 \quad r=1 \rightarrow \mathcal{E} = -\mu$$

$$V=1 \quad r=1 \rightarrow \mathcal{E} = \frac{1}{2} - \mu$$

PROBLEM 3

$$V = 8.215$$

$$r = 12172$$

$$\mathcal{E} = 1 \quad \mu = \mu_{\oplus}$$

RADIUS OF PERAPSIS + APOAPSIS - ELLIPTICAL ORBITS

$$r_p = a(1-e)$$

$$a, e \rightarrow \text{FCN} \rightarrow r_p, r_a$$

$$r_a = a(1+e)$$

$$\text{TESTS: } a = 6378 \quad e = 0 \rightarrow r_p = r_a = a$$

$$a = 8000 \quad e = 0.5 \rightarrow r_p = 4000, r_a = 12000$$

PROBLEM 2

$$a = 38500, e = 0.8181$$

$$r_p = 7000 \text{ km} \quad r_a = 70000 \text{ km}$$

