

Constants

$$\begin{aligned}\omega_{\oplus} &= 15^{\circ} \text{ sidereal hr}^{-1} \\ \omega_{\oplus} &= 0.000\,072\,921\,151\,467 \text{ rad solar sec}^{-1} \\ \mu_{\oplus} &= 398\,600.5 \text{ km}^3 \text{ s}^{-2} \\ R_{\oplus} &= 6378.137 \text{ km} \\ e_{\oplus} &= 0.081\,819\,190\,842\,6 \\ J_2 &= 1.082\,63 \times 10^{-3} \\ \rho_o &= 1.225 \text{ kg m}^{-3}\end{aligned}$$

Satellites - Basics

$$\begin{aligned}0 &= \ddot{\bar{r}} + \frac{\mu}{r^3} \bar{r} \quad r = \frac{a(1-e^2)}{1+e\cos\nu} \\ h &= rv\cos\gamma = \sqrt{\mu a(1-e^2)} \quad \bar{h} = \bar{r} \times \bar{v} \\ a &= \frac{r_a + r_p}{2} \quad e = \frac{2c}{2a} = \frac{r_a - r_p}{r_a + r_p} \\ r_p &= a(1-e) \quad r_a = a(1+e) \\ p &= a(1-e^2) = \frac{h^2}{\mu} \quad \bar{e} = \frac{\bar{v} \times \bar{h}}{\mu} - \frac{\bar{r}}{r} \\ P &= 2\pi\sqrt{\frac{a^3}{\mu}} \quad \varepsilon = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a}\end{aligned}$$

Time

$$\begin{aligned}\theta_g &= \theta_{g_0} + 1.00273790935 \times 2\pi * D \\ D &= \text{Day Num} - 1 + \frac{\text{HR}}{24} + \frac{\text{MIN}}{1440} \\ &+ \frac{\text{SEC}}{86400} \quad \text{where D is elapsed days} \\ \text{LST} &= \theta_g + \lambda_E \\ \text{EST} &= \text{UT} - 5 \quad \text{EDT} = \text{UT} - 4\end{aligned}$$

Preliminary Orbit Determination

$$\begin{aligned}x &= \left| \frac{a_e}{\sqrt{1-e^2\sin^2 L}} + H \right| \cos L \\ z &= \left| \frac{a_e(1-e^2)}{\sqrt{1-e^2\sin^2 L}} + H \right| \sin L \\ \bar{r}_s &= x \cos \text{LST} \hat{i} + x \sin \text{LST} \hat{j} + z \hat{k} \\ \rho_s &= -\rho \cos \alpha \cos \beta \\ \rho_e &= \rho \sin \alpha \cos \beta \\ \rho_z &= \rho \sin \beta \\ \dot{\rho}_s &= -\dot{\rho} \cos \alpha \cos \beta + \rho \dot{\alpha} \sin \alpha \cos \beta + \rho \dot{\beta} \cos \alpha \sin \beta \\ \dot{\rho}_e &= \dot{\rho} \sin \alpha \cos \beta + \rho \dot{\alpha} \cos \alpha \cos \beta - \rho \dot{\beta} \sin \alpha \sin \beta \\ \dot{\rho}_z &= \dot{\rho} \sin \beta + \rho \dot{\beta} \cos \beta \\ [\text{IJK}] &= \text{ROT}_3(\text{LST}) \text{ROT}_2(\text{COLAT}) [\text{SEZ}]\end{aligned}$$

$$\begin{aligned}\bar{r} &= \bar{\rho} + \bar{r}_s \\ \bar{v} &= \dot{\bar{\rho}} + \bar{\omega}_{\oplus} \times \bar{r} \\ \bar{v}_2 &= -\Delta t_{32} \left(\frac{1}{\Delta t_{21} \Delta t_{31}} + \frac{\mu}{12r_1^3} \right) \bar{r}_1 \\ &+ (\Delta t_{32} - \Delta t_{21}) \left(\frac{1}{\Delta t_{21} \Delta t_{32}} + \frac{\mu}{12r_2^3} \right) \bar{r}_2 \\ &+ \Delta t_{21} \left(\frac{1}{\Delta t_{32} \Delta t_{31}} + \frac{\mu}{12r_3^3} \right) \bar{r}_3\end{aligned}$$

Transfers

$$\begin{aligned}\Delta V_{OTB}^2 &= V_1^2 + V_2^2 - 2V_1 V_2 \cos \Delta\gamma \\ \tan \gamma &= \frac{e \sin \nu}{1 + e \cos \nu} \quad \Delta V_s = 2V_i \sin \frac{\theta}{2} \\ \Delta V_{COMB}^2 &= V_1^2 + V_2^2 - 2V_1 V_2 \cos \Delta i\end{aligned}$$

Rendezvous

$$\begin{aligned}\text{TOF} &= \pi \sqrt{\frac{a^3}{\mu}} \quad \omega = \sqrt{\frac{\mu}{r_{\text{circ}}^3}} \\ \alpha_{\text{lead}} &= \omega_t \times \text{TOF} \quad \phi_f = \pi - \alpha_{\text{lead}} \\ \text{Wait time}_{\text{coplanar}} &= \frac{\phi_f - \phi_i \pm 2\pi n}{\omega_t - \omega_i} \\ \text{Wait time}_{\text{noncoplanar}} &= \frac{\phi_f - \phi_i \pm 2\pi n}{\omega_t} = \frac{\alpha_i - \alpha_f + 2\pi n}{\omega_t}\end{aligned}$$

Keplers Problem

$$\begin{aligned}n &= \sqrt{\frac{\mu}{a^3}} \quad M_f = M_i + n \times \text{TOF} - 2k\pi \\ M &= E - e \sin E \\ \cos E &= \frac{e + \cos \nu}{1 + e \cos \nu} \quad \cos \nu = \frac{\cos E - e}{1 - e \cos E} \\ E_{n+1} &= E_n + \frac{M - E_n + e \sin E_n}{1 - e \cos E_n}\end{aligned}$$

COEs to RV

$$\begin{aligned}\bar{r} &= r [\cos \nu \hat{p} + \sin \nu \hat{q}] \\ \bar{v} &= \sqrt{\frac{\mu}{p}} [-\sin \nu \hat{p} + (e + \cos \nu) \hat{q}] \\ [\text{IJK}] &= \text{ROT}_3(\Omega) \text{ROT}_1(i) \text{ROT}_3(\omega) [PQW]\end{aligned}$$

Perturbations

$$\bar{n} = n_0 \left[1 + \frac{3}{2} J_2 \left(\frac{R_{\oplus}}{p_0} \right)^2 \sqrt{1 - e_0^2} \left(1 - \frac{3}{2} \sin^2 i_0 \right) \right]$$

$$n = n_0 + \dot{n}_0 \Delta t$$

$$\dot{e}_{drag} = \frac{-2(1 - e_0)\dot{n}_0}{3\bar{n}}$$

$$e = e_0 + \dot{e}_{drag} \Delta t$$

$$\dot{\Omega}_{J_2} = \left[-\frac{3}{2} J_2 \left(\frac{R_{\oplus}}{p_0} \right)^2 \cos i_0 \right] \bar{n}$$

$$\Omega = \Omega_0 + \dot{\Omega}_{J_2} \Delta t$$

$$\dot{\omega}_{J_2} = \left[\frac{3}{2} J_2 \left(\frac{R_{\oplus}}{p_0} \right)^2 \left(2 - \frac{2}{5} \sin^2 i_0 \right) \right] \bar{n}$$

$$\omega = \omega_0 + \dot{\omega}_{J_2} \Delta t$$

$$M = M_0 + n_0 \Delta t + \frac{\dot{n}_0}{2} \Delta t^2$$

$$\bar{a}_{drag} = -\frac{1}{2} \rho \frac{C_D A}{m} v \bar{v}$$

Entry - $p_0 = 1.225 \text{ kg m}^{-3}$

$$\dot{r} = v \sin \phi_E$$

$$\dot{v} = \frac{\rho_0}{2\Delta} v^2 \exp(-\beta h) \quad \beta = \frac{1}{7.315 \text{ km}}$$

$$\Delta = \frac{m}{C_D A}$$

$$v = v_e \exp \left(\frac{\rho_0}{2\Delta\beta \sin \phi_E} \exp(-\beta h) \right)$$

$$h_{max_g} = \frac{1}{\beta} \ln \left(\frac{-\rho_0}{\Delta\beta \sin \phi_E} \right)$$

$$max_g = \frac{-\beta v_E^2 \sin \phi_E}{2g_0 \exp(1)}$$

$$v_{max_g} = v_E \exp(-0.5) \approx 0.61 v_E$$

Proximity Operations

$$x(t) = \frac{\dot{x}_0}{\omega} \sin \omega t - \left(3x_0 + \frac{2\dot{y}_0}{\omega} \right) \cos \omega t + \left(4x_0 + \frac{2\dot{y}_0}{\omega} \right)$$

$$\dot{x}(t) = \dot{x}_0 \cos \omega t + (3\omega x_0 + 2\dot{y}_0) \sin \omega t$$

$$y(t) = \left(6x_0 + \frac{4\dot{y}_0}{\omega} \right) \sin \omega t + \frac{2\dot{x}_0}{\omega} \cos \omega t - (6\omega x_0 + 3\dot{y}_0) t + \left(y_0 - \frac{2\dot{x}_0}{\omega} \right)$$

$$\dot{y}(t) = (6\omega x_0 + 4\dot{y}_0) \cos \omega t - 2\dot{x}_0 \sin \omega t - (6\omega x_0 + 3\dot{y}_0)$$

$$z(t) = z_0 \cos \omega t + \frac{\dot{z}_0}{\omega} \sin \omega t$$

$$\dot{z}(t) = -z_0 \omega \sin \omega t + \dot{z}_0 \cos \omega t$$

Attitude Kinematics

$$ROT_1 = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \alpha & -\sin \alpha \\ 0 & \sin \alpha & \cos \alpha \end{bmatrix}$$

$$ROT_2 = \begin{bmatrix} \cos \beta & 0 & \sin \beta \\ 0 & 1 & 0 \\ -\sin \beta & 0 & \cos \beta \end{bmatrix}$$

$$ROT_3 = \begin{bmatrix} \cos \gamma & -\sin \gamma & 0 \\ \sin \gamma & \cos \gamma & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Note: Assuming column vectors.

Summary of Orbital Properties

Any Type of Orbit

$$\begin{aligned}
 \vec{h} &= \vec{r} \times \vec{v}, & \vec{r} &= r\hat{u}_r, & \vec{v} &= v_r\hat{u}_r + v_\theta\hat{u}_\theta, \\
 h &= rv \cos \gamma = rv_\theta = r^2\dot{\theta}, & r &= \frac{h^2/\mu}{1 + e \cos \theta}, & v_r &= \frac{\mu}{h}e \sin \theta = \dot{r}, \\
 \vec{e} &= \frac{\vec{v} \times \vec{h}}{\mu} - \frac{\vec{r}}{r}, & r_p &= \frac{h^2/\mu}{1 + e}, & v_\theta &= \frac{\mu}{h}(1 + e \cos \theta) = r\dot{\theta}, \\
 \mathcal{E} &= \frac{1}{2}v^2 - \frac{\mu}{r} = -\frac{1}{2}\frac{\mu^2}{h^2}(1 - e^2), & \tan \gamma &= \frac{v_r}{v_\theta} = \frac{e \sin \theta}{1 + e \cos \theta}.
 \end{aligned}$$

Circular Orbits: ($e = 0$)

$$v = \sqrt{\frac{\mu}{r}}, \quad \mathcal{E} = -\frac{\mu}{2r}, \quad T = \frac{2\pi}{\sqrt{\mu}}r^{3/2}.$$

Elliptic Orbits: ($0 < e < 1$)

$$\begin{aligned}
 r &= \frac{a(1 - e^2)}{1 + e \cos \theta}, & a &= \frac{h^2/\mu}{1 - e^2} = \frac{1}{2}(r_a + r_p), & T &= \frac{2\pi}{\sqrt{\mu}}a^{3/2}, \\
 r_p &= \frac{h^2/\mu}{1 + e} = a(1 - e), & b &= a\sqrt{1 - e^2}, & e &= \frac{r_a - r_p}{r_a + r_p}, \\
 r_a &= \frac{h^2/\mu}{1 - e} = a(1 + e), & \mathcal{E} &= -\frac{\mu}{2a}, & h &= \sqrt{\mu a(1 - e^2)}.
 \end{aligned}$$

Parabolic Orbits: ($e = 1$)

$$v = \sqrt{\frac{2\mu}{r}}, \quad \mathcal{E} = 0.$$

Hyperbolic Orbits: ($e > 1$)

$$\begin{aligned}
 r &= \frac{a(e^2 - 1)}{1 + e \cos \theta}, & a &= \frac{h^2/\mu}{e^2 - 1}, & \theta_\infty &= \cos^{-1}(-1/e), \\
 r_p &= a(e - 1), & b &= a\sqrt{e^2 - 1}, & \beta &= \cos^{-1}(1/e), \\
 \mathcal{E} &= \frac{\mu}{2a}, & h &= \sqrt{\mu a(e^2 - 1)}.
 \end{aligned}$$

Transformation between Orbital Elements and (\vec{r}, \vec{v})

Given (\vec{r}, \vec{v}) , find the orbital elements $(h, e, \theta, \Omega, i, \omega)$

$$\begin{aligned}
 r &= |\vec{r}|, \\
 \vec{h} &= \vec{r} \times \vec{v}, \quad h = |\vec{h}|, \\
 \vec{e} &= \frac{1}{\mu} \vec{v} \times \vec{h} - \frac{\vec{r}}{r}, \quad e = |\vec{e}|, \\
 \vec{N} &= \hat{z} \times \vec{h}, \\
 i &= \cos^{-1} \left(\frac{\vec{h} \cdot \hat{z}}{h} \right) \quad (0 \leq i \leq \pi), \\
 \Omega &= \tan^{-1} \left(\frac{\hat{y} \cdot \vec{N}}{\hat{x} \cdot \vec{N}} \right) = \text{numpy.arctan2}(\hat{y} \cdot \vec{N}, \hat{x} \cdot \vec{N}), \\
 \omega &= \tan^{-1} \left(\frac{\vec{h} \cdot (\vec{N} \times \vec{e})}{h(\vec{N} \cdot \vec{e})} \right) = \text{numpy.arctan2}(\vec{h} \cdot (\vec{N} \times \vec{e}), h(\vec{N} \cdot \vec{e})), \\
 \theta &= \tan^{-1} \left(\frac{\vec{h} \cdot (\vec{e} \times \vec{r})}{h(\vec{e} \cdot \vec{r})} \right) = \text{numpy.arctan2}(\vec{h} \cdot (\vec{e} \times \vec{r}), h(\vec{e} \cdot \vec{r})).
 \end{aligned}$$

(Use the Numpy `numpy.arctan2` function to compute \tan^{-1} , i.e. $\tan^{-1}(y/x) = \text{numpy.arctan2}(y, x)$).

Given the orbital elements $(h, e, \theta, \Omega, i, \omega)$, find (\vec{r}, \vec{v})

$$\begin{aligned}
 \hat{N} &= \cos \Omega \hat{x} + \sin \Omega \hat{y}, \\
 \hat{h} &= \sin i \sin \Omega \hat{x} - \sin i \cos \Omega \hat{y} + \cos i \hat{z}, \\
 \hat{N}_t &= -\sin \Omega \cos i \hat{x} + \cos \Omega \cos i \hat{y} + \sin i \hat{z}, \\
 \hat{u}_r &= \cos(\theta + \omega) \hat{N} + \sin(\theta + \omega) \hat{N}_t, \\
 \hat{u}_\theta &= -\sin(\theta + \omega) \hat{N} + \cos(\theta + \omega) \hat{N}_t, \\
 r &= \frac{h^2}{\mu} \frac{1}{1 + e \cos \theta}, \\
 \mathcal{E} &= -\frac{1}{2} \frac{\mu^2}{h^2} (1 - e^2), \\
 v &= \sqrt{2 \left(\mathcal{E} + \frac{\mu}{r} \right)}, \\
 \gamma &= \tan^{-1} \left(\frac{e \sin \theta}{1 + e \cos \theta} \right), \\
 \vec{r} &= r \hat{u}_r, \\
 \vec{v} &= v \cos \gamma \hat{u}_\theta + v \sin \gamma \hat{u}_r.
 \end{aligned}$$

Orbital Position as a Function of Time

Time since periapsis passage: t (for any type of orbit)

$$\frac{\mu^2}{h^3} t = \int_0^\theta \frac{d\theta}{(1 + e \cos \theta)^2}. \quad (1)$$

Circular orbit: $\theta \sim t$

$$t = \frac{\theta}{2\pi} T, \quad (2)$$

where $T = \frac{2\pi}{\sqrt{\mu}} r^{3/2}$.

Elliptic orbit: $\theta \sim E \sim M_e \sim t$

$$\tan \frac{E}{2} = \sqrt{\frac{1-e}{1+e}} \tan \frac{\theta}{2}, \quad (3)$$

$$M_e = E - e \sin E, \quad (4)$$

$$t = \frac{M_e}{2\pi} T, \quad (5)$$

where $T = \frac{2\pi}{\sqrt{\mu}} a^{3/2}$.

Parabolic orbit: $\theta \sim M_p \sim t$

$$M_p = \frac{1}{2} \tan \frac{\theta}{2} + \frac{1}{6} \tan^3 \frac{\theta}{2}, \quad (6)$$

$$\tan \frac{\theta}{2} = \left[3M_p + \sqrt{(3M_p)^2 + 1} \right]^{1/3} - \left[3M_p + \sqrt{(3M_p)^2 + 1} \right]^{-1/3}, \quad (7)$$

$$t = M_p / \left(\frac{\mu^2}{h^3} \right). \quad (8)$$

Hyperbolic orbit: $\theta \sim F \sim M_h \sim t$

$$\tanh \frac{F}{2} = \sqrt{\frac{e-1}{e+1}} \tan \frac{\theta}{2}, \quad (9)$$

$$M_h = e \sinh F - F, \quad (10)$$

$$t = M_h / \left(\frac{\mu^2}{h^3} (e^2 - 1)^{3/2} \right). \quad (11)$$

	Axial Rotational Period (Rev/Day)	Mean Equatorial Radius (km)	Gravitational Parameter $\mu = Gm$ (km ³ /sec ²)	Semi-major Axis of Orbit (km)	Orbital Period (sec)	Eccentricity of Orbit	Inclination of Orbit to Ecliptic (deg)
☉ Sun	0.0394011 (not rigid)	695990.00	1.32712000E+11	————	————	————	————
☾ Moon*	0.0366004	1737.50	4.902800E+03	3.84400000E+05 (around Earth)	2360592 27.322 Earth days	0.05540000	5.16000
☿ Mercury	0.0170515	2.439000E+03	2.203210E+04	5.79092000E+07	7600568.601 87.97 Earth Days	0.205631	7.00487
♀ Venus	0.0041149 (retrograde)	6051.80	3.24859000E+05	1.08209000E+08	19414191.77 224.7 Earth Days	0.006773	3.39471
⊕ Earth	1.0027576	6378.14	3.98600000E+05	1.49589800E+08	31555647.16 365.23 Earth Days	0.01671020	4.98816000E-05
♂ Mars	0.9746985	3397.00	4.282840E+04	2.27937000E+08	59353583.28 686.96 Earth Days	0.09341230	1.85061
♃ Jupiter	2.4181458	71492.00	1.26687000E+08	7.7841200E+08	374396573 11.87 yr	0.04839270	1.30530
♄ Saturn	2.2522523	60330.00	3.79313000E+07	1.42673000E+09	929341659.8 29.47 yr	0.05415060	2.48446
♅ Uranus	1.3921178 (retrograde)	26200.00	5.79397000E+06	2.87097000E+09	2653128427 84.13 yr	0.04716770	0.76986
♆ Neptune	1.4897579	25225.00	6.835110E+06	4.49825000E+09	5203301252 165 yr	0.00858587	1.76917
♇ Pluto	0.1565631	1195.00	8.737670E+02	5.906638E+09	7829522968	0.24880800	17.14180

	Axial Rotational Period (Rev/Day)	Equatorial Radius (km)	Gravitational Parameter $\mu = Gm$ (km ³ /sec ²)	Semi-major Axis of Orbit (km)	Orbital Period	Eccentricity of Orbit	Inclination of Orbit to Ecliptic (deg)	
♀	Charon	0.1562500	593	108	19600 (about Pluto)	6.38725 days	0.00	96.16
	Nix	?	23→68	?	48675 (about Pluto)	24.8562 d	0.0023	0.1
	Hydra	?	32→84	?	64780 (about Pluto)	38.2065 d	0.0052	0.25
	Ganymede	0.1397711	2631.2	9887.834000	1070000.0 (about Jupiter)	7.154553 d	0.002	0.195
	Callisto	0.592059	2410.30	7179.29	1883000.0 (about Jupiter)	16.68902 d	0.007	0.281
	Titan	0.6271393	2575.50	8978.190000	1221830.0 (about Saturn)	15.94542 d	0.0292	0.33
	Titania	assumed synchronous	788.9	235.544000	435910.0 (about Uranus)	8.706 d	0.00220000	0.14000
	Ceres	2.6448030	474.00 km	63.200000	413906175	1680.982 d	0.07990478	10.58674
	Phobos	synchronous	11.10	0.000629	9377 (about Mars)	.31891 d	0.01510000	26.04000
Triton	synchronous	1350.00	1427.598000	354759 (about Neptune)	.5877 d (retrograde)	0.00001600	129.18200	