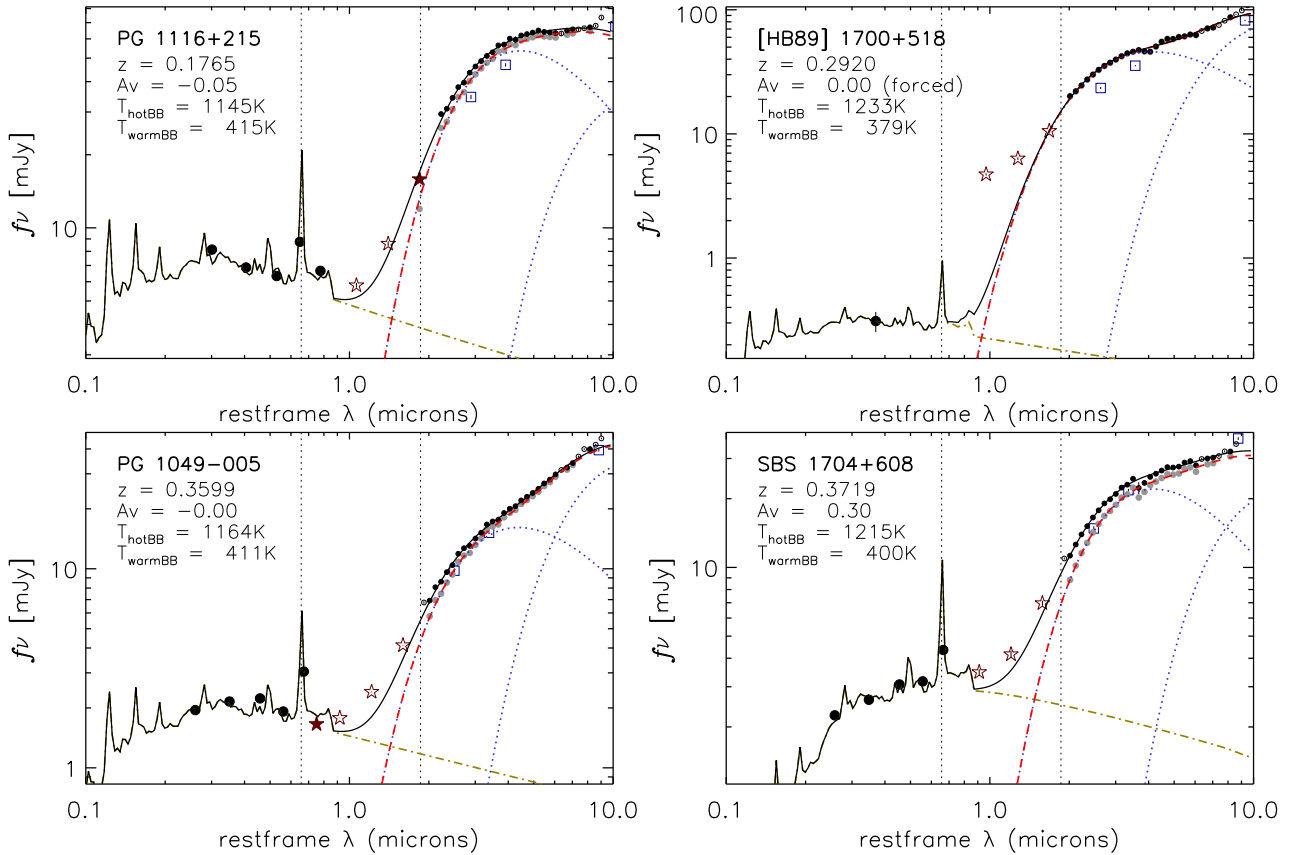


# The near-to-mid infrared spectrum of quasars

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## APPENDIX A: SEDS AND MODEL FITS FOR INDIVIDUAL SOURCES



**Figure A1.** Spectral energy distributions of the sources in the sample. Circles, stars, and squares represent broadband photometry in the observed-frame optical (from SDSS), NIR (2MASS/UKIDSS/VHS), and MIR (WISE), respectively. Filled symbols indicate bands used to fit the disk component (yellow dot-dashed line). The (AKARI+)-IRS spectra resampled at  $\lambda/\Delta\lambda = 20$  are shown with small black dots. Grey dots below the (AKARI+)-IRS spectra represent the dust spectrum obtained after subtraction of the disk component. The model for the dust spectrum (red dashed line) is the linear combination of two black-bodies at adjustable temperatures (blue dotted lines). The combined disk+dust model is represented by the black solid line. The vertical dotted lines mark the restframe wavelength of the  $H\alpha$  and  $Pa\alpha$  recombination lines.

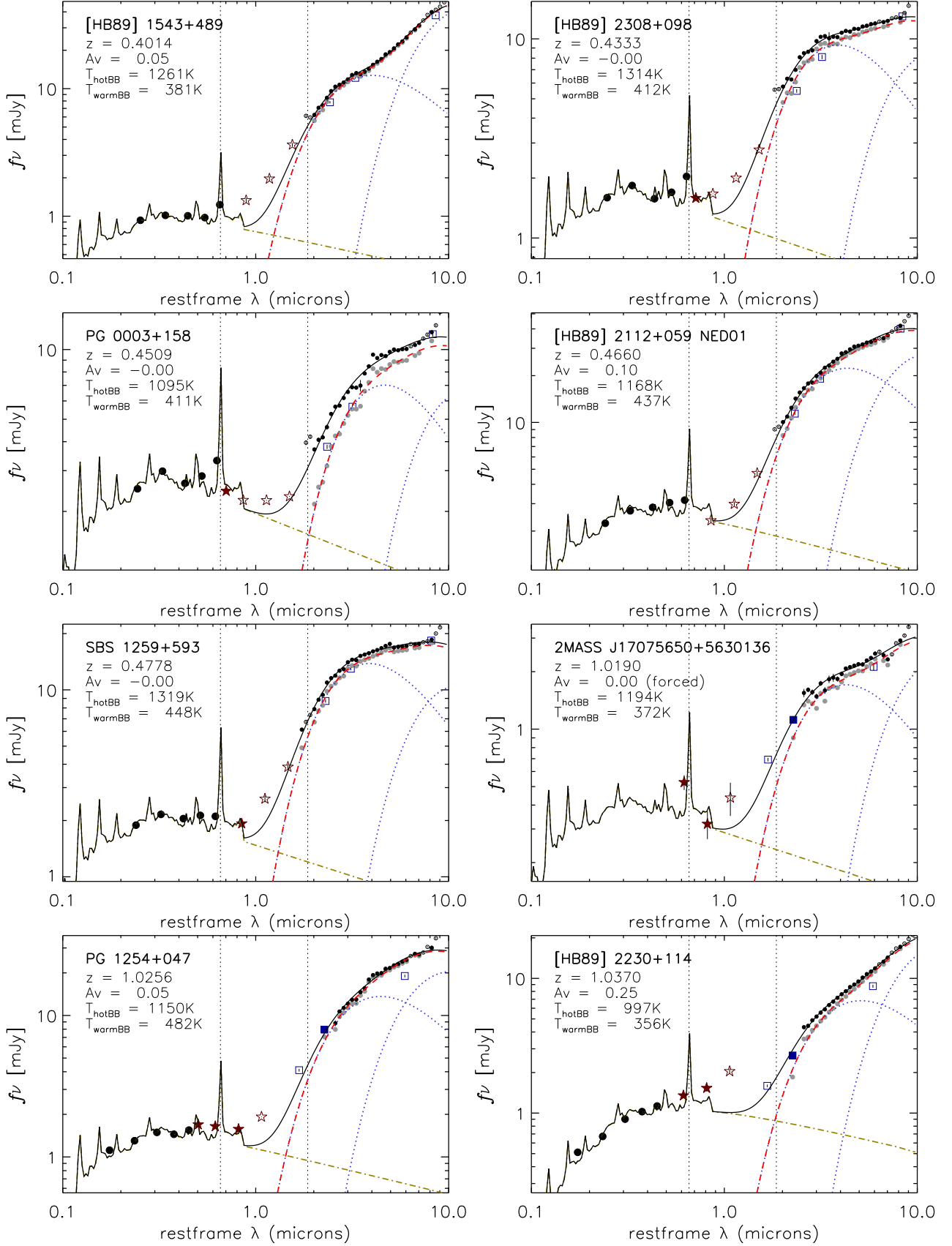


Figure A1. continued

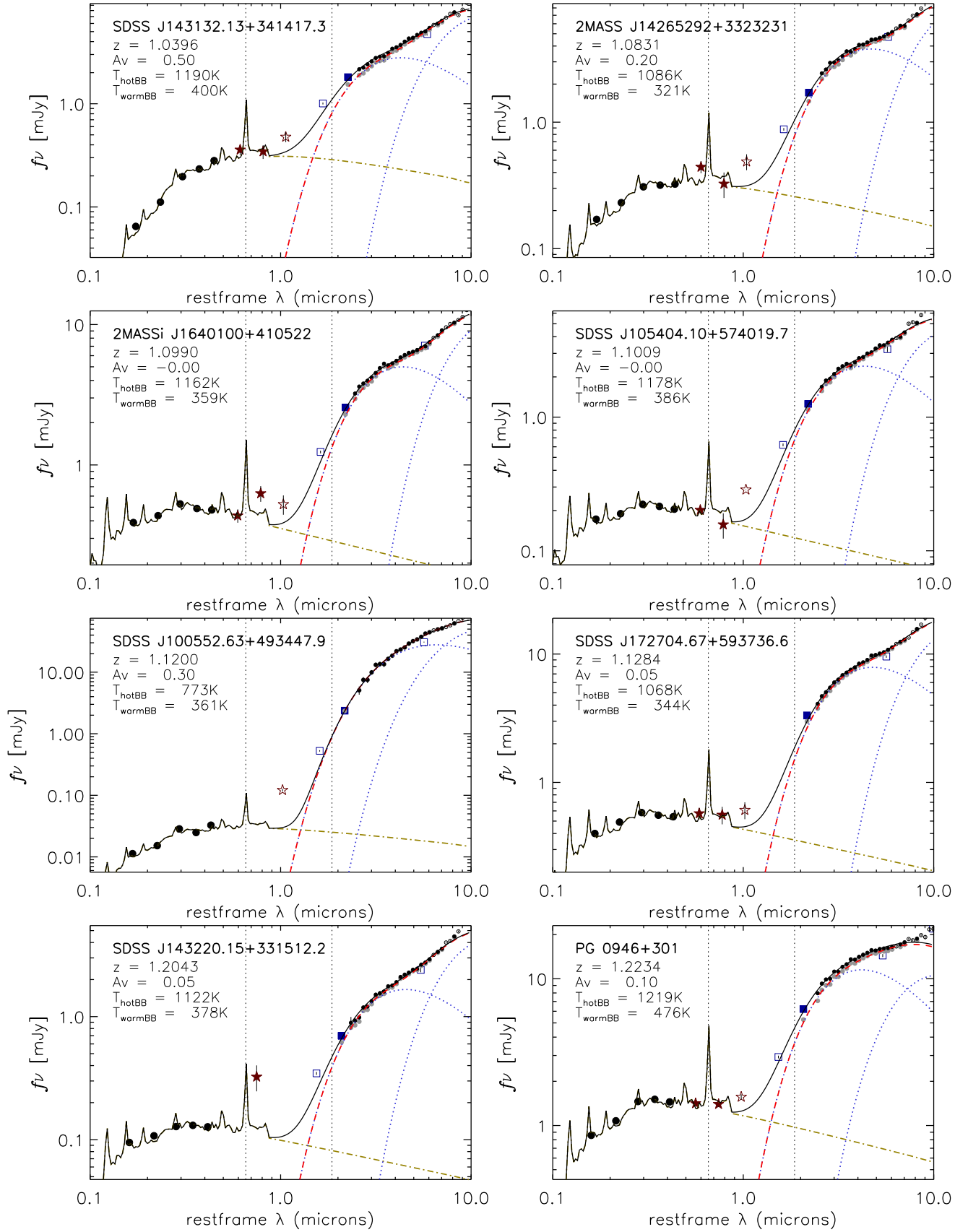


Figure A1. continued

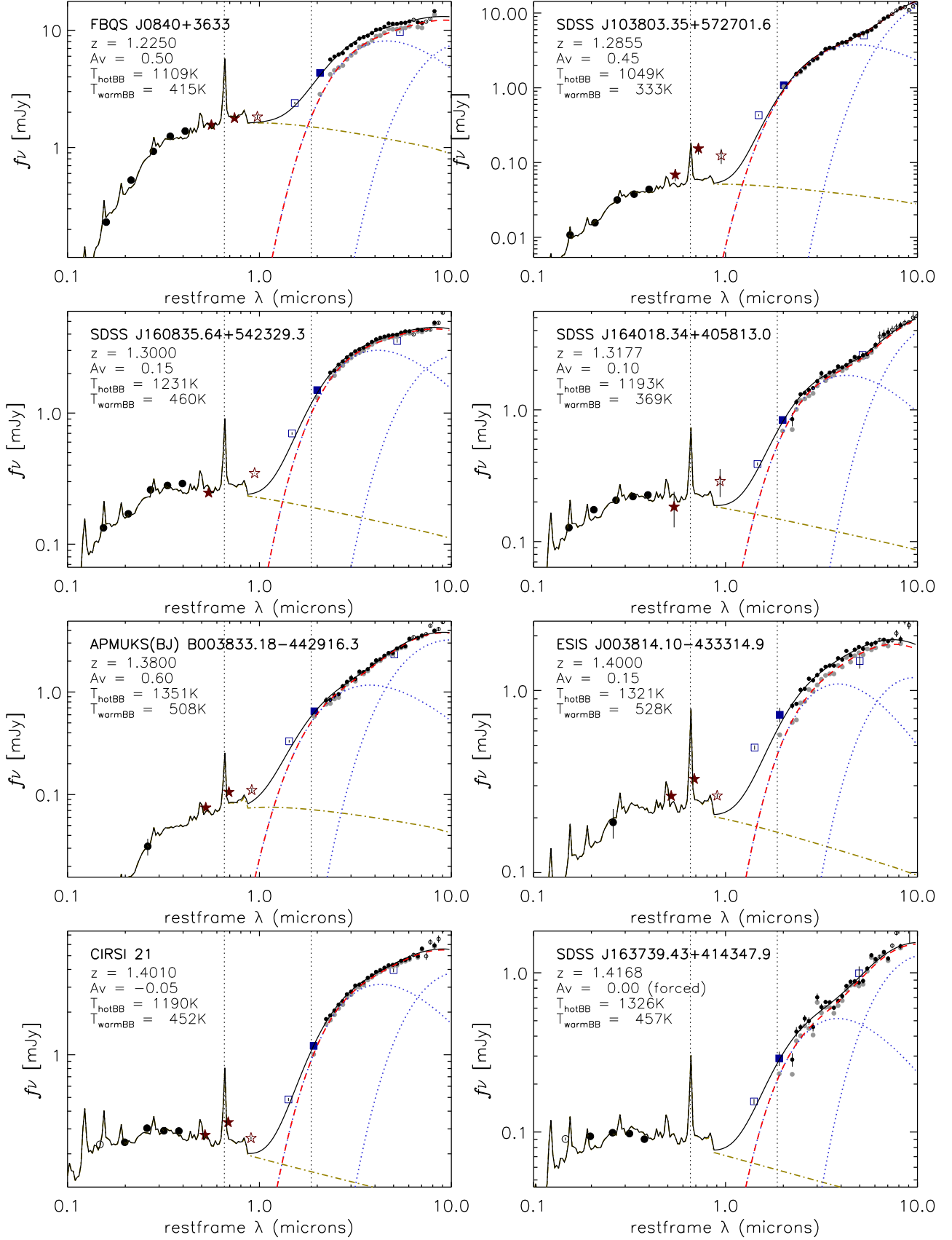


Figure A1. continued

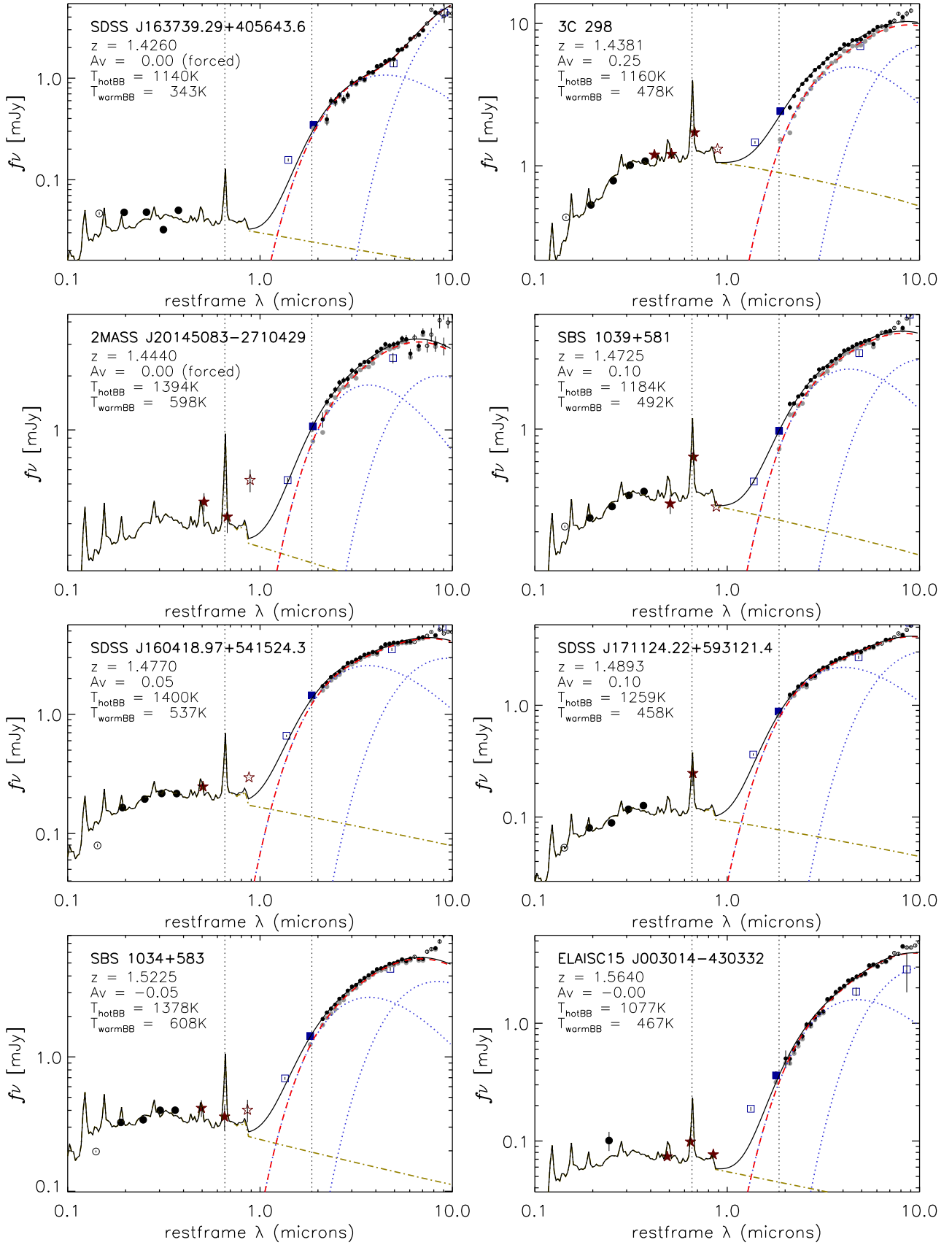


Figure A1. continued

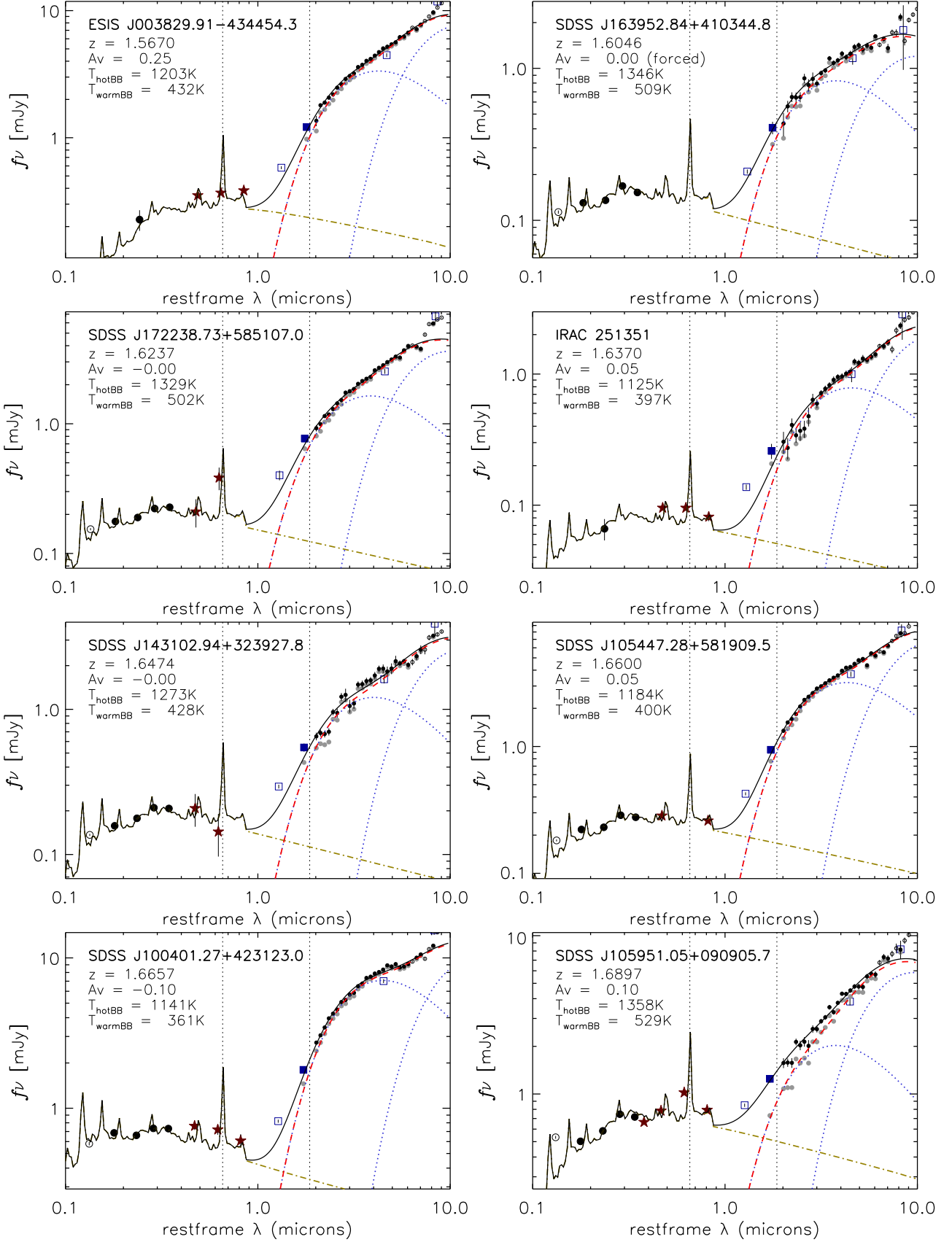


Figure A1. continued

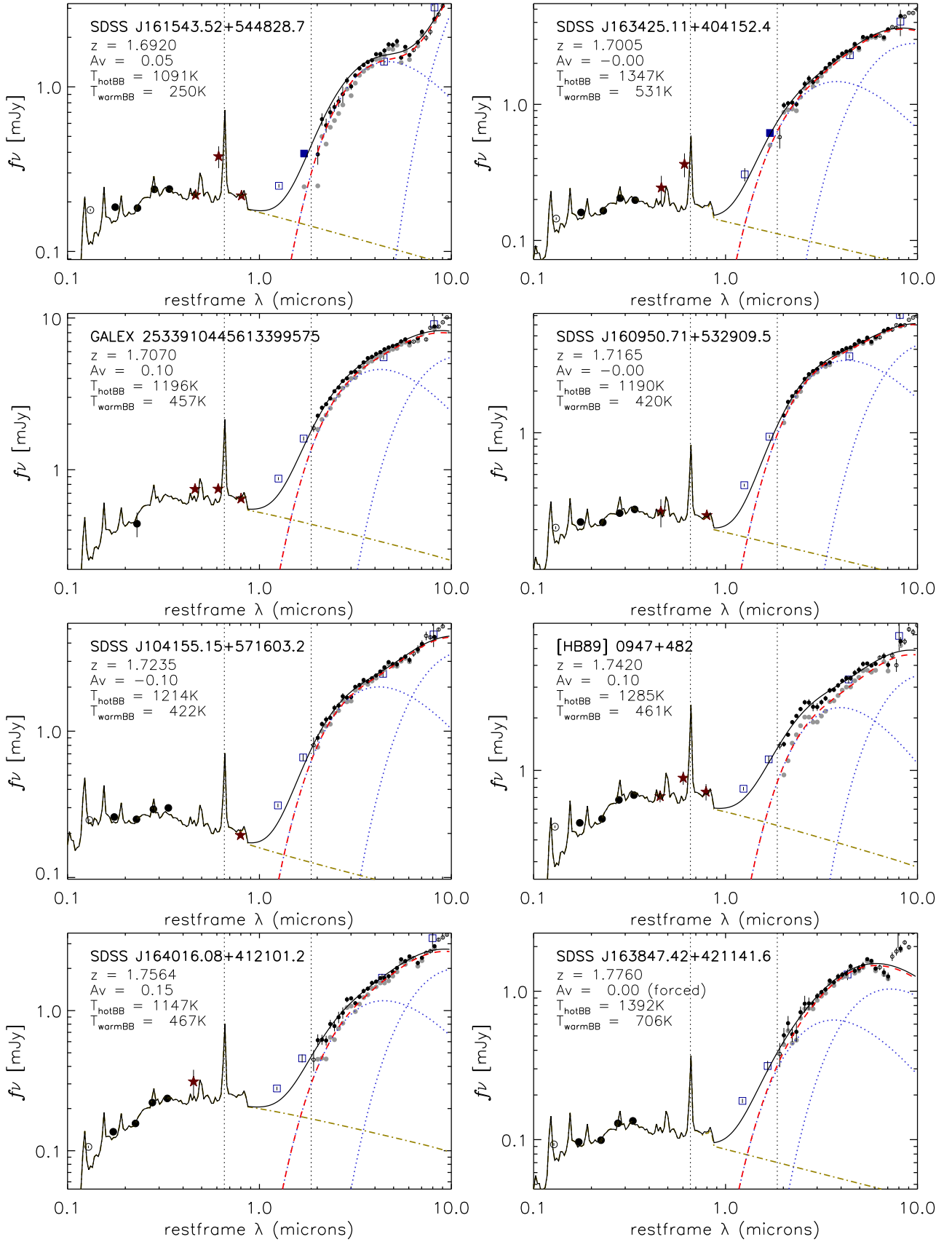


Figure A1. continued

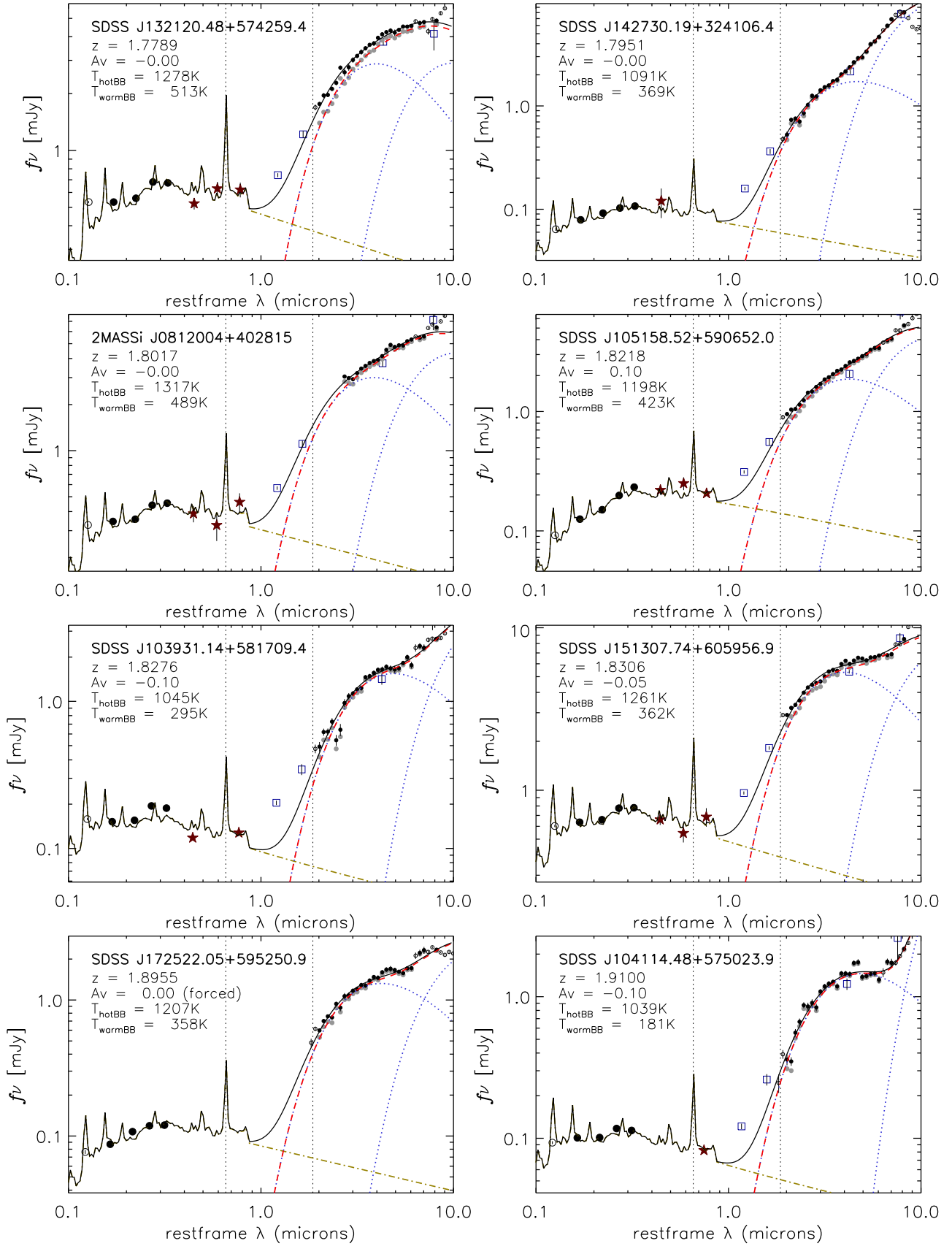


Figure A1. continued



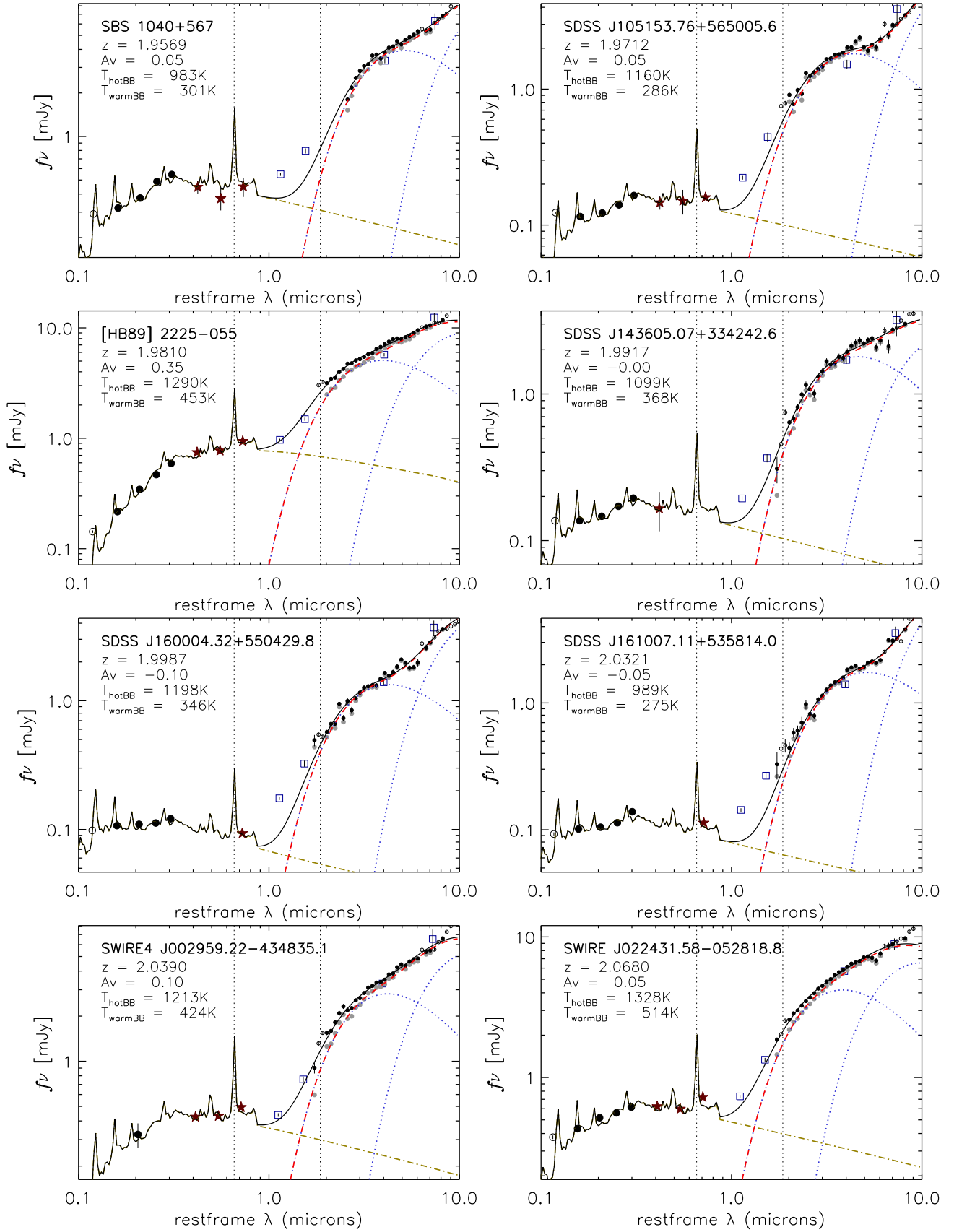


Figure A1. continued

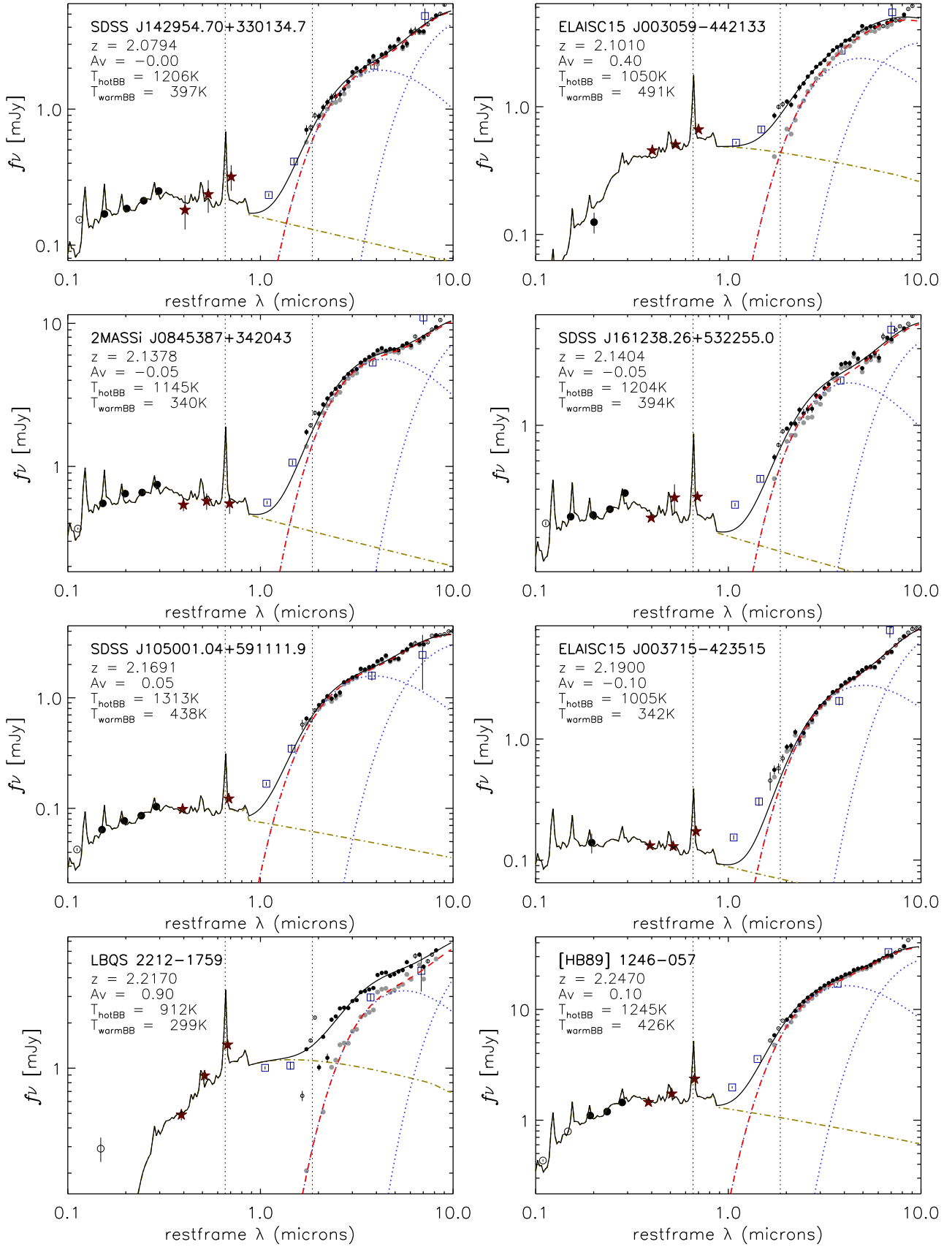


Figure A1. continued

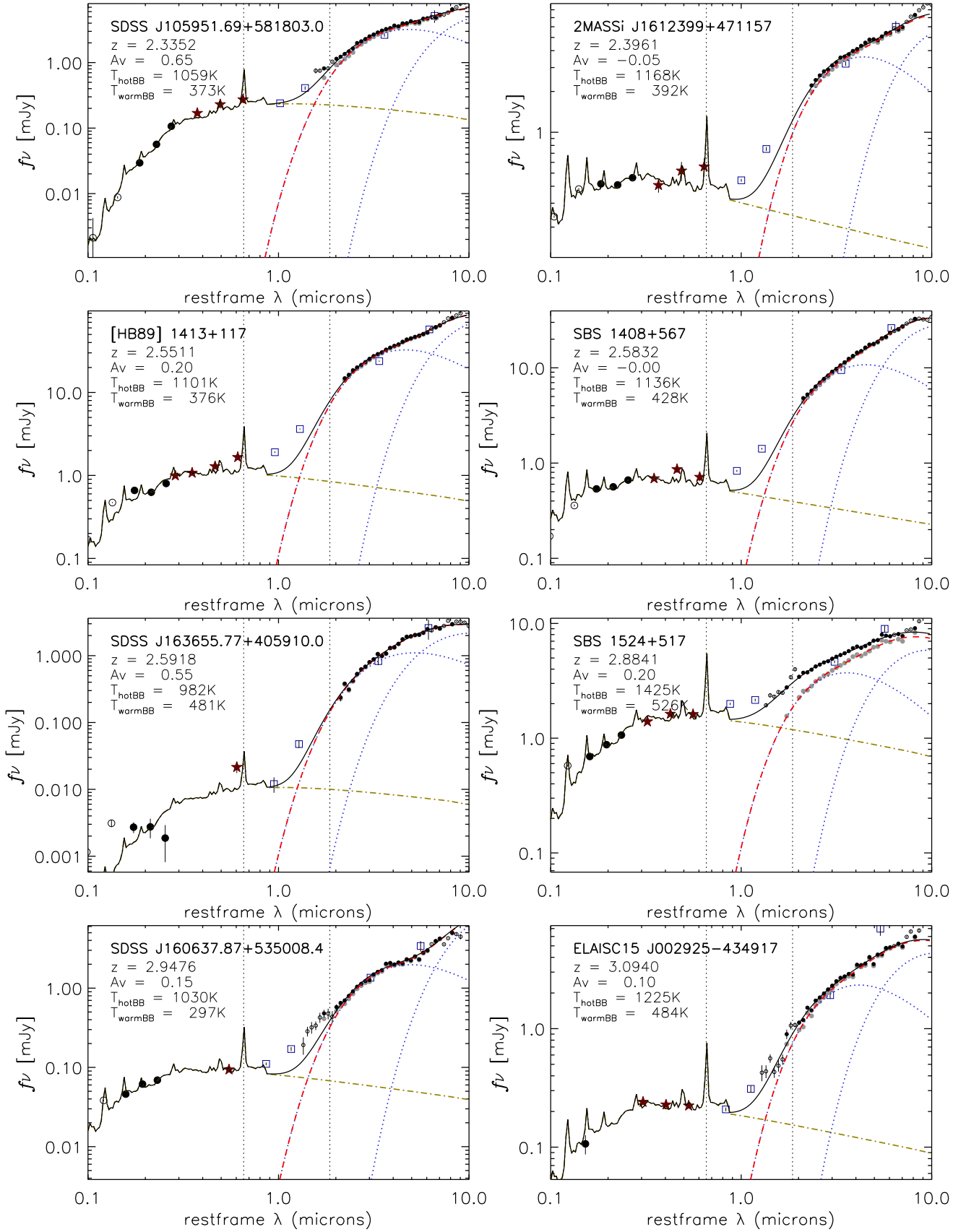
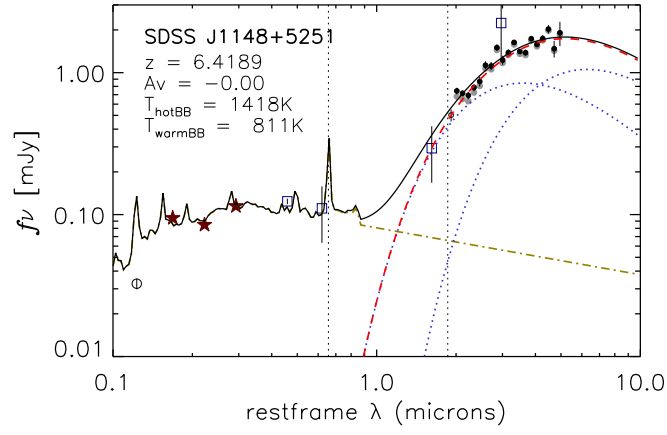


Figure A1. continued



**Figure A1.** continued