Statement of Purpose: MSU

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About Me_

As an undergraduate student in astrophysics, I have a strong urge of pursuing a Ph.D degree to continue my adventure on mystery outside our tiny planet.

Though having been fascinated by astrophysics for long, the strong faith of being an astronomer, I believe, dates back in my freshman year in biology major. Though I got the **highest GPA (3.87/4)** among 120 freshmen in biology, I was finally tired of cumbersome taxonomy and test-tube-washing. Fortunate enough, after taking several fundamental courses on physics & astronomy, I realized that for me, the subtle balance of astonishing physical pictures and beautiful mathematical structures is so well attained in astrophysics, when I decided to switch to astronomy major. The training in School of Physics at Peking University has then lain a firm foundation of *mathematics* (calculus, linear algebra, PDE, and statistics) and *physics* (analytical mechanics, statistical mechanics, electrodynamics, and quantum mechanics) for me. More advanced astrophysics courses (spectroscopy, cosmology, general relativity, gravitational-wave astrophysics, etc.) have equipped me for conducting research. Since my sophomore year, I have kept in **the first place in GPA (3.84/4)** among 28 students in astronomy major.

One of my greatest pursuits is to combine state-of-the-art observations with powerful computational methods to understand complex astrophysical environment. This sweet 'temptation' has driven me to explore observational and computational astrophysics from various binary systems to astrochemistry & astrobiology within interstellar medium.

Research Experience

My first scientific project started at the end of my second year under the supervision of **Prof. Xian Chen** in *Kavli Institute for Astronomy and Astrophysics at Peking University* and **Prof. Fujun Du** in *Purple Mountain Observatory, Chinese Academy of Sciences*. Starting from a 'crazy' implication of recent observations that the supermassive black hole in the Milky Way was active about 2-8 Myr ago, I worked as an 'archaeologist' digging out the astrochemical history of our galaxy with numerical methods. During such an ancient AGN event, tremendous hard X-ray radiation should be able to efficiently penetrate the dusty Galactic disk and lead to the synthesis of complex molecules related to the origin of life. To run the long term astrochemical simulation, I built our chemical network based on a classic gas-phase model (osu_01_2007). For self-consistency, I included necessary surface processes important for molecule formation. Under a detailed AGN and galactic absorption model, I carefully embedded X-ray ionization at various distances from the galactic center in the network. The simulation was executed with the KROME package, a widely-used library-like code for astrophysical simulation with chemistry included. We are extremely excited to see that several typical prebiotic species show observable changes in abundance distribution in Galactic disk given former AGN events. Our paper is under final revision and will be submitted to ApJ soon.

My research was not limited to one single subfield. In the summer of 2019, I went to *Caltech* as part of the Summer Undergraduate Research Fellowship (SURF) program and touched the field of transients for the first time under the mentorship of **Prof. Shrinivas Kulkarni**. With light curves from Zwicky Transient Facility (ZTF), a state-of-the-art optical time-domain survey with unprecedented field of view and survey efficiency, I conducted a systematic search for periodic white dwarf binaries. Starting with a population of 486,641 WD candidates identified in *Gaia*, I performed a cross match between the so-far most complete catalog and ZTF data to select a subset of 90,000 sources with over 100 observations in ZTF. My carefully designed periodogram based on Lomb-Scargle method was then applied to the extracted light curves. A sample of 81 periodic white dwarfs (WDs) with periods between 1 and 3 hr to our surprise stood out. With combined analysis of both shapes of light curves and color information from Gaia and PanSTARRS, I classified several sources of interest including

a contact binary candidate with possibly the shortest period known and an unusual, strongly ellipsoidal-modulated double white dwarfs system with an extremely low-mass (ELM) component. A catalog of the 81 sources with our preliminary classification has been made. Our paper will be updated with continuing follow-ups before submission to ApJ.

After an attempt to reveal the nature of various stellar systems with degenerate components in observational way, I grow intensely interested in exploring more on the intrinsic physics of compact binaries. At present I am a visiting undergraduate student at *UC Santa Cruz* working on my thesis on the mass transfer in compact binaries with **Prof. Enrico Ramirez-Ruiz**. We focus on the so-called AM CVn binary systems, of which the components are white dwarfs so close to each other that the mass transfer flow from the donor directly strikes the accretor. They are believed to be important progenitors of Type Ia supernovae. Understanding the nature of these direct impact systems, dominant sources of GWs for space-based interferometers, will be of prime importance to next-generation detectors like *LISA*. To get an intuition of mass transfer, I independently built a 3-body integrator with FORTRAN to calculate the ballistic trajectory of a particle in Roche lobe overflow in a binary system, and successfully reproduced the trajectories and angular momenta transition processes in former literature. Then I have been trying to conduct hydrodynamical simulations with the radiation MHD simulation code **FLASH** to test the long-term stability of direct impact mass transfer. With Python package yt, I got to visualize the torque density during mass transfer to analyze how the orbital angular momentum evolves. In the long run, we will try to work out how light curves and even gravitational radiation properties for AM CVn systems are constrained by model parameters with simulated data.

Conducting research in these most advanced facilities around the world enables me to I take a bite of various cutting-edge astronomical subfields, from astrochemistry to compact objects, and developed general skills of data analysis, visualization, and simulation on supercomputers. My proficiency in English communication and academic writing has risen to a new level after working overseas for months. More importantly, I gradually get to learn how to stay out of frustration, anxiety and loneliness, so that I am able to persist in a long-term project.

Why MSU? - Academic Interests _

Closely cooperating with experimental & theoretical physicists, astronomy research at MSU is well-known for cutting edge study on high energy astrophysical events. Members the stellar & nuclear group, with strong computational background, use state-of-the-art codes including Enzo and FLASH to construct epic models for compact objects, core-collapse supernovae, and accreting processes. Graduate students will receive training in observation in world-class facilities including SOAR as well as computational skills as part of the optional graduate courses and have access to world-leading supercomputers, such as Blue Waters. MSU also encourages graduate students to start research early in their first year and pairs beginning graduate students with senior graduate mentors, and senior graduate as part of the mentoring program, which will definitely help me get started and hone my academic abilities.

I am not too surprised to find that MSU meets so perfectly with my research experience and interests. **Prof. Stephen Zepf** has performed observational work on multiwavelength property for X-ray sources in globular clusters with SOAR, HST, and Chandra. **Prof. Jay Strader** also has great contribution in searching for black holes in globular clusters, and is highly involved in time domain surveys such as ASSAS-SN. I also appreciate **Prof. Laura Chomiuk**'s multiwavelength work, from radio to gamma-ray band, on interacting compact binaries to study their evolution as Type Ia supernovae progenitor. My experience in the summer with high energy transients and data analysis in time-domain surveys enables me to conduct further research on these compact objects. And again, out of an intense interest to combine observation with computational methods, I do admire **Prof. Sean Couch**'s simulation work on core-collapse supernovae with FLASH, for which I have plenty of coding experience.

I do appreciate it if you could consider my application. I am looking forward to meeting with you for a more detailed talk soon.