



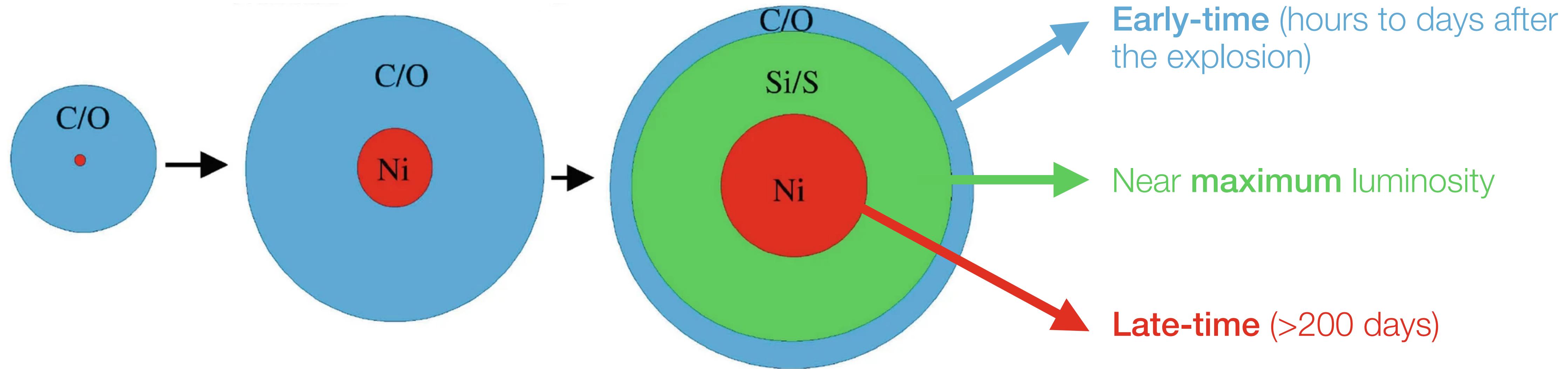
From the First to the Last Photons

Unveiling the progenitor demographics of Type Ia supernovae

Chang Liu, Adam A. Miller
CIERA/Northwestern

2025.8.22 One Hundred Years of Supernova Science

Near-M_{Ch} WD



Sub-M_{Ch} WD in a double detonation

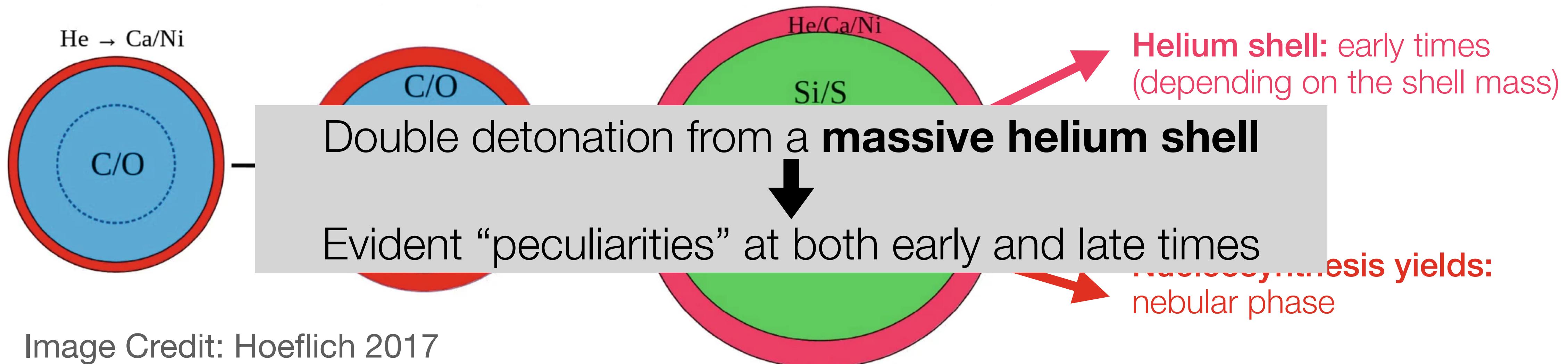
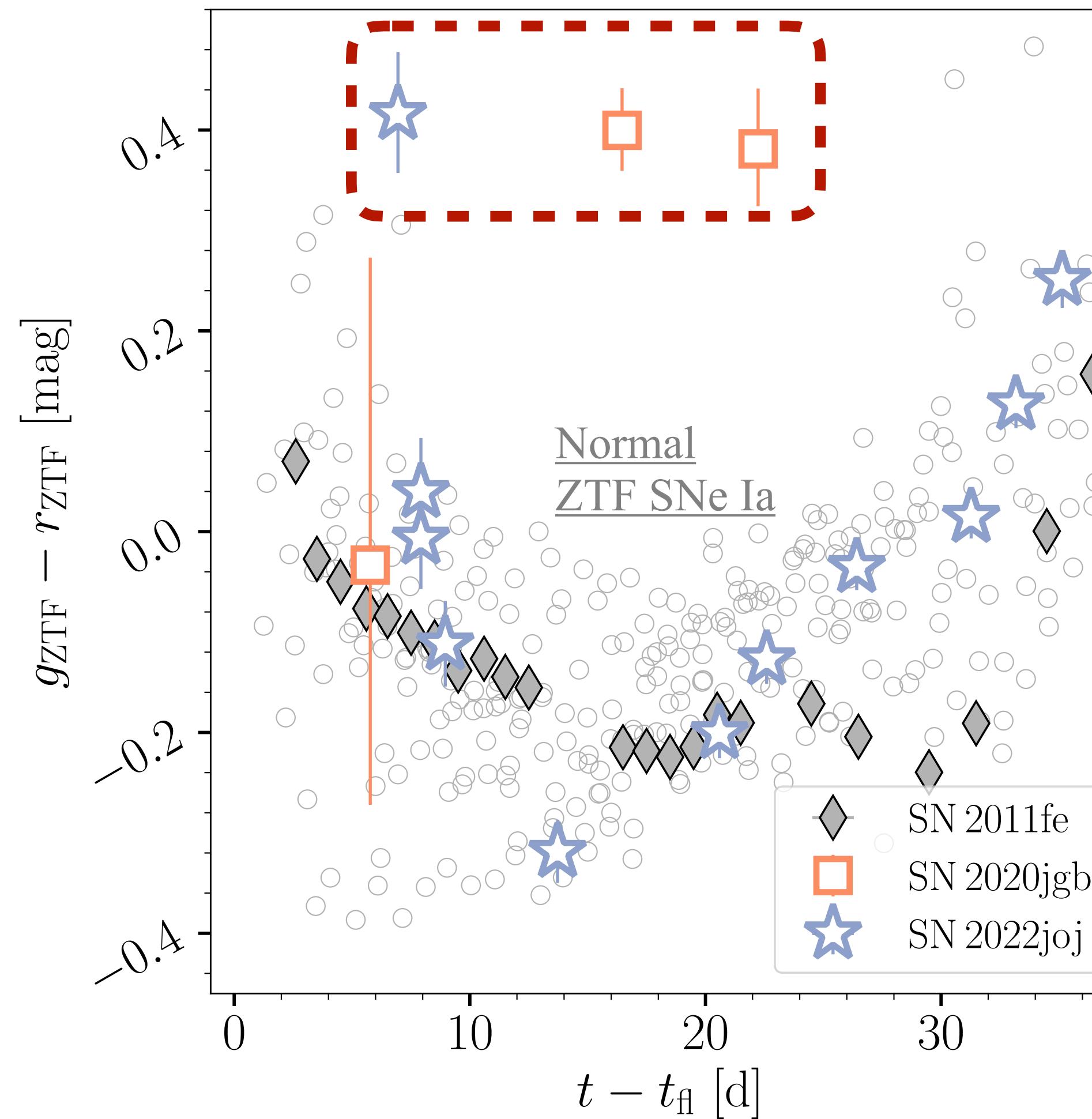


Image Credit: Höflich 2017

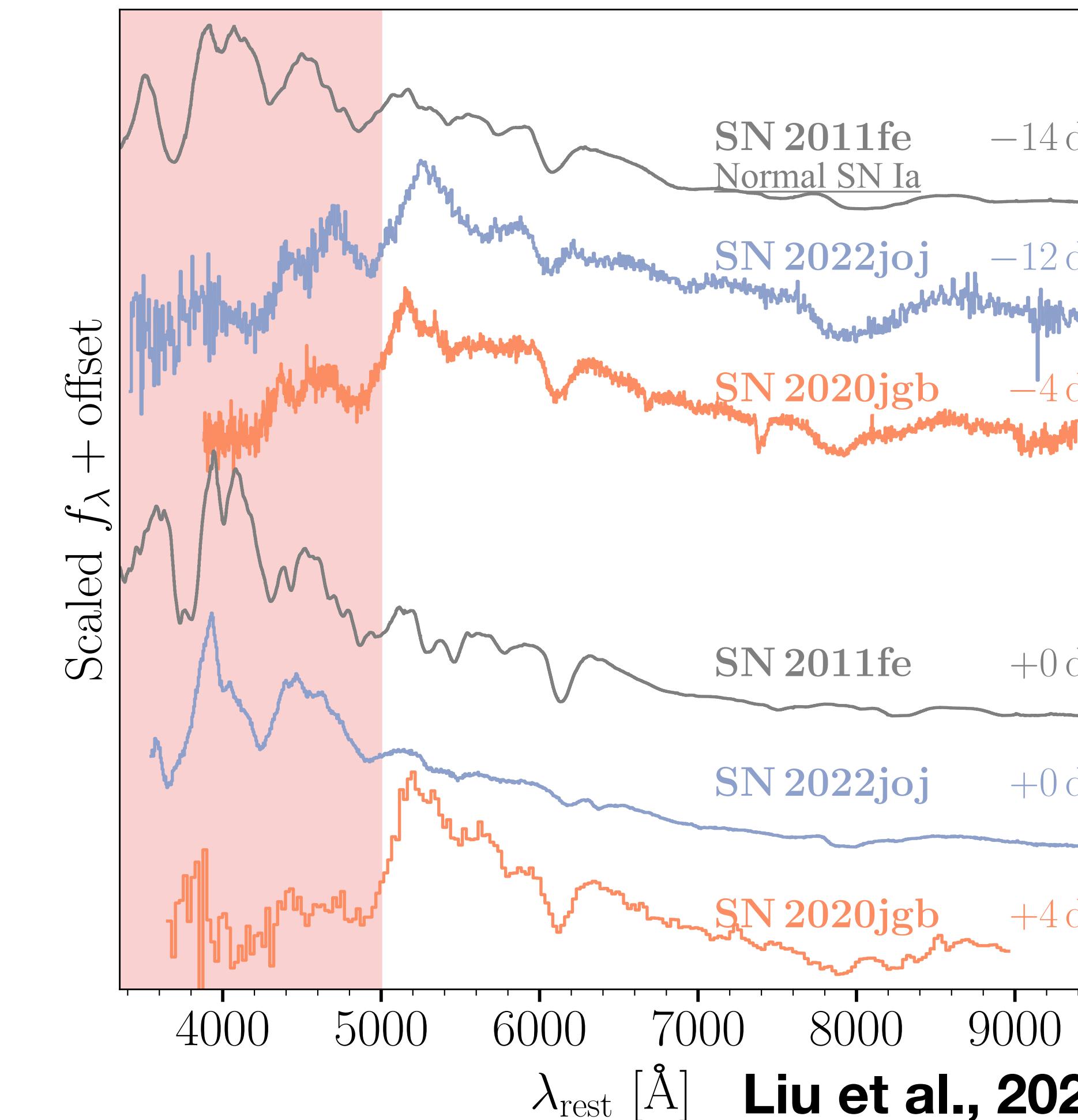
Massive He-shell Double Detonation

From days after explosion to maximum luminosity

Unusual red colors



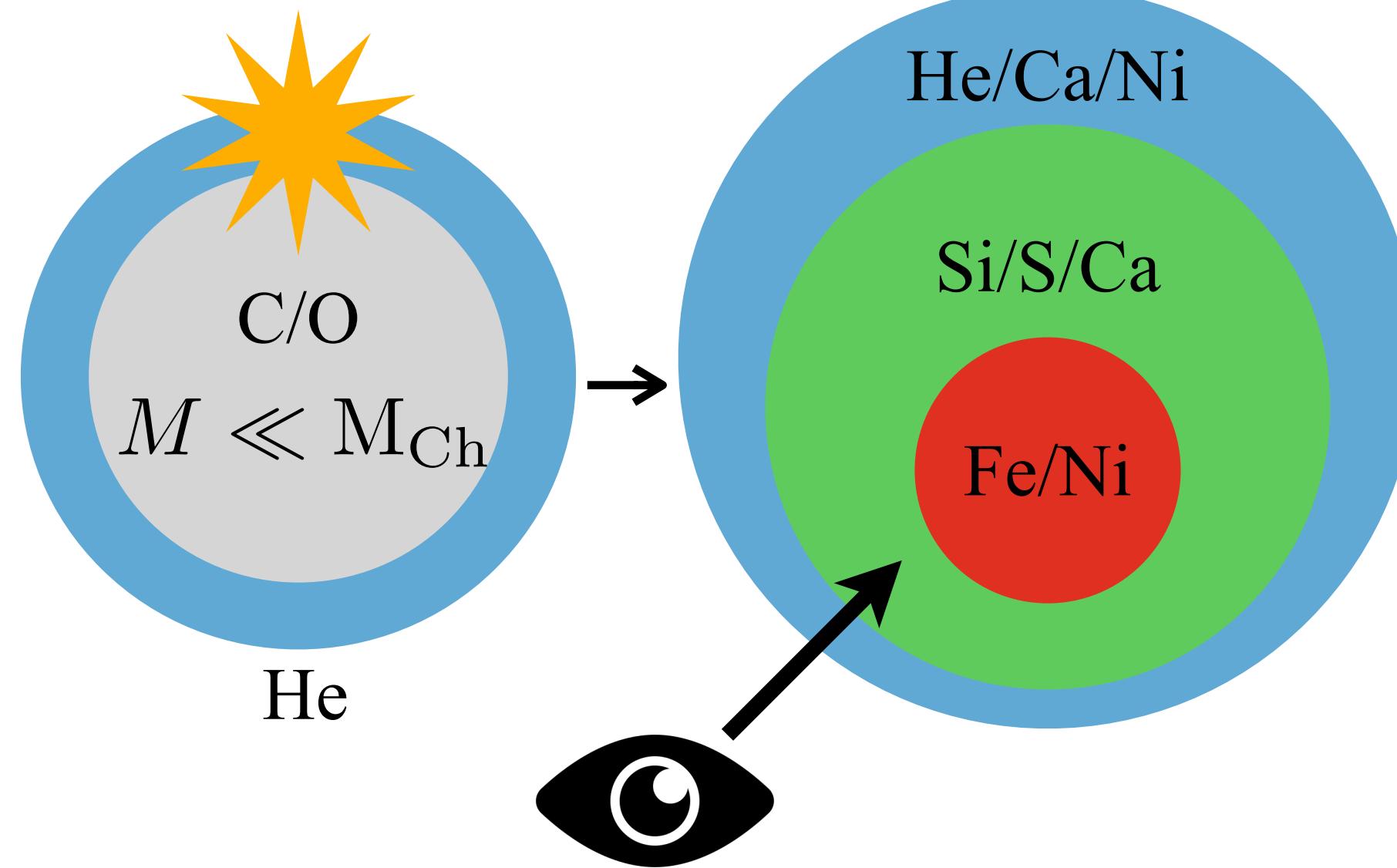
Suppression of flux $< 5000 \text{ \AA}$



SN2022joj: Asymmetric Double Detonation?

Maximum luminosity

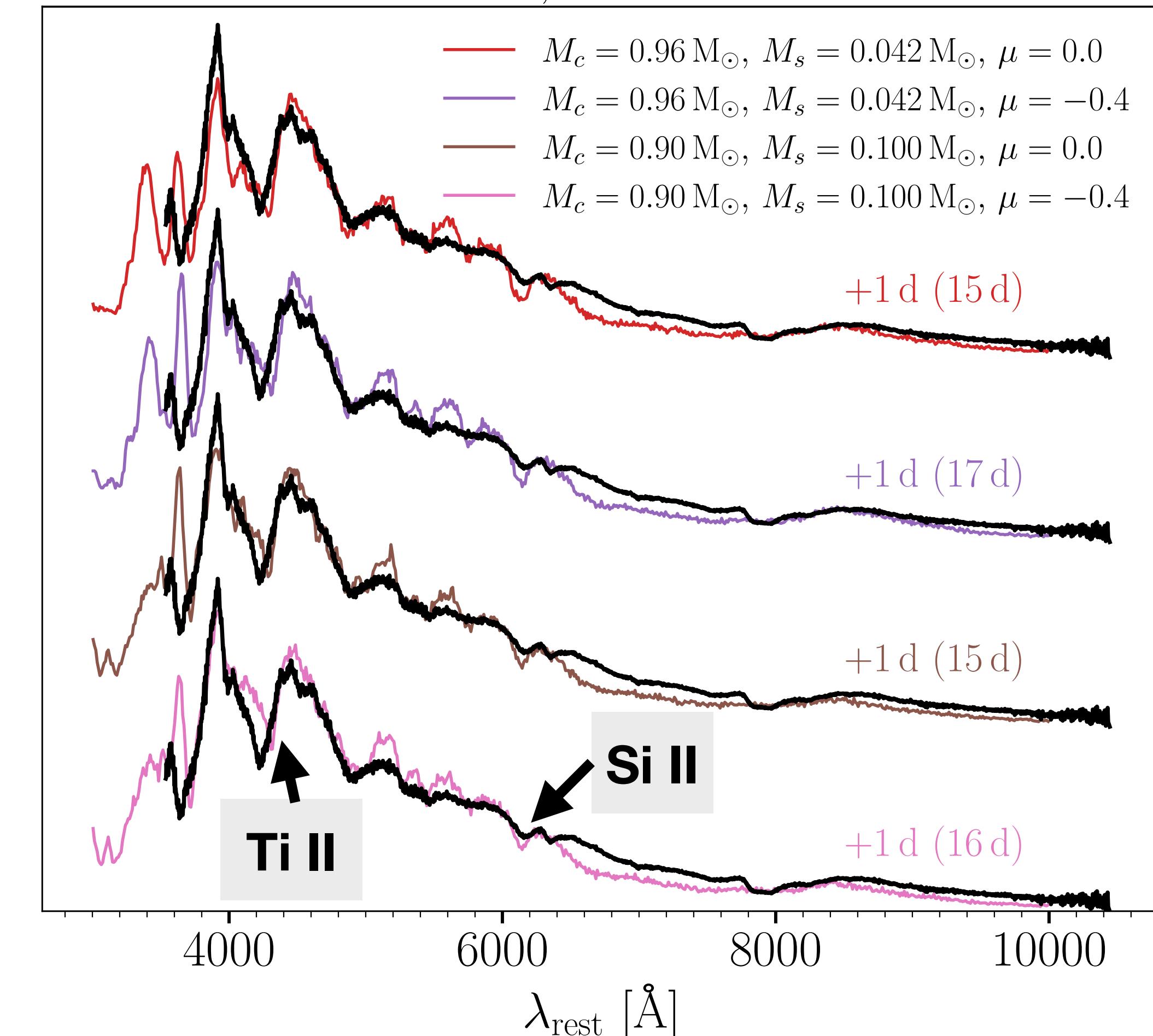
Strong Ti II: Helium shell ashes



Weak, slow Si II: sightline opposite
to the He detonation

Liu et al., 2023b

$$t - t_{B,\max} = +0 \text{ d}$$

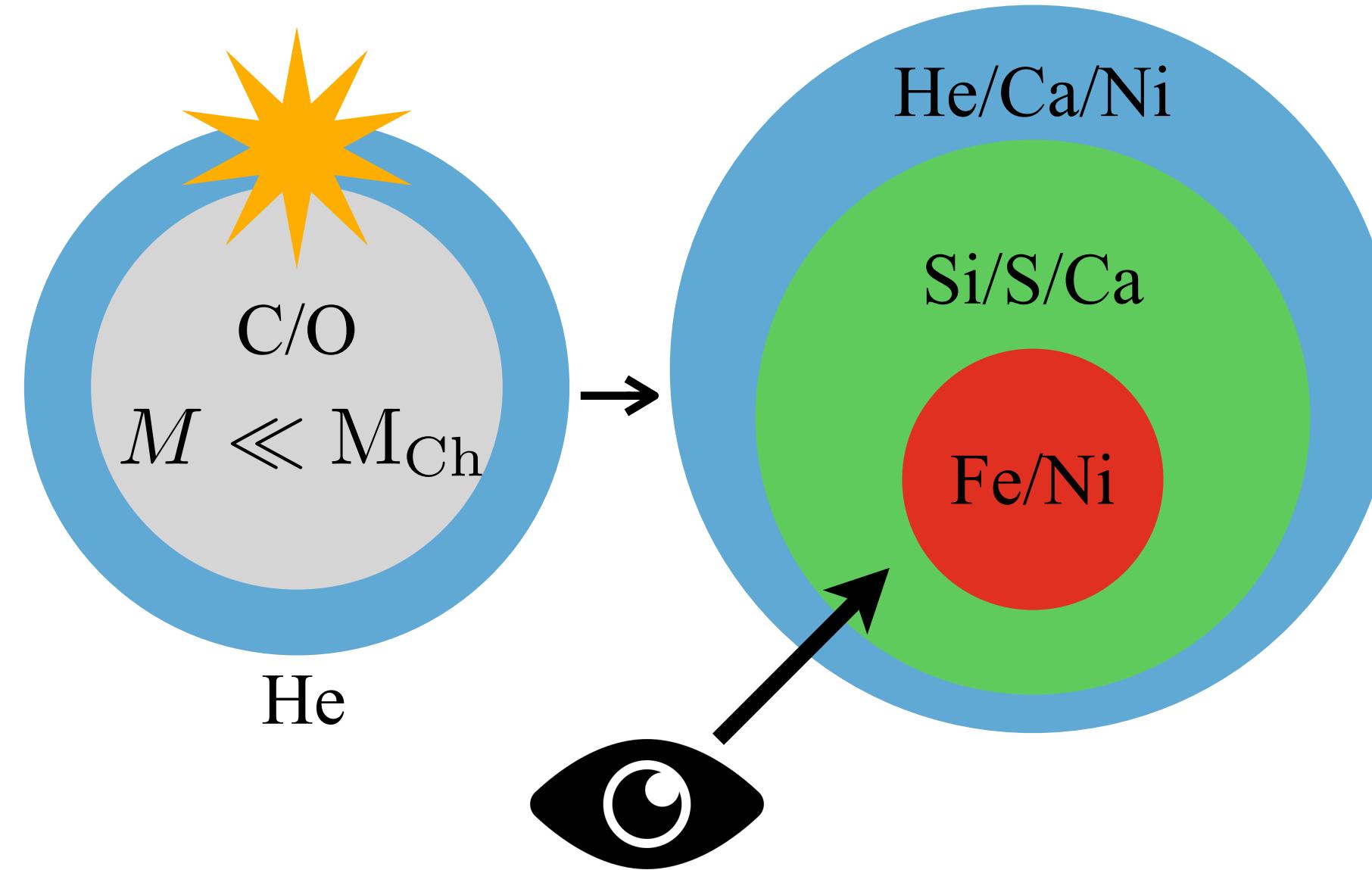


SN2022joj: Asymmetric Double Detonation?

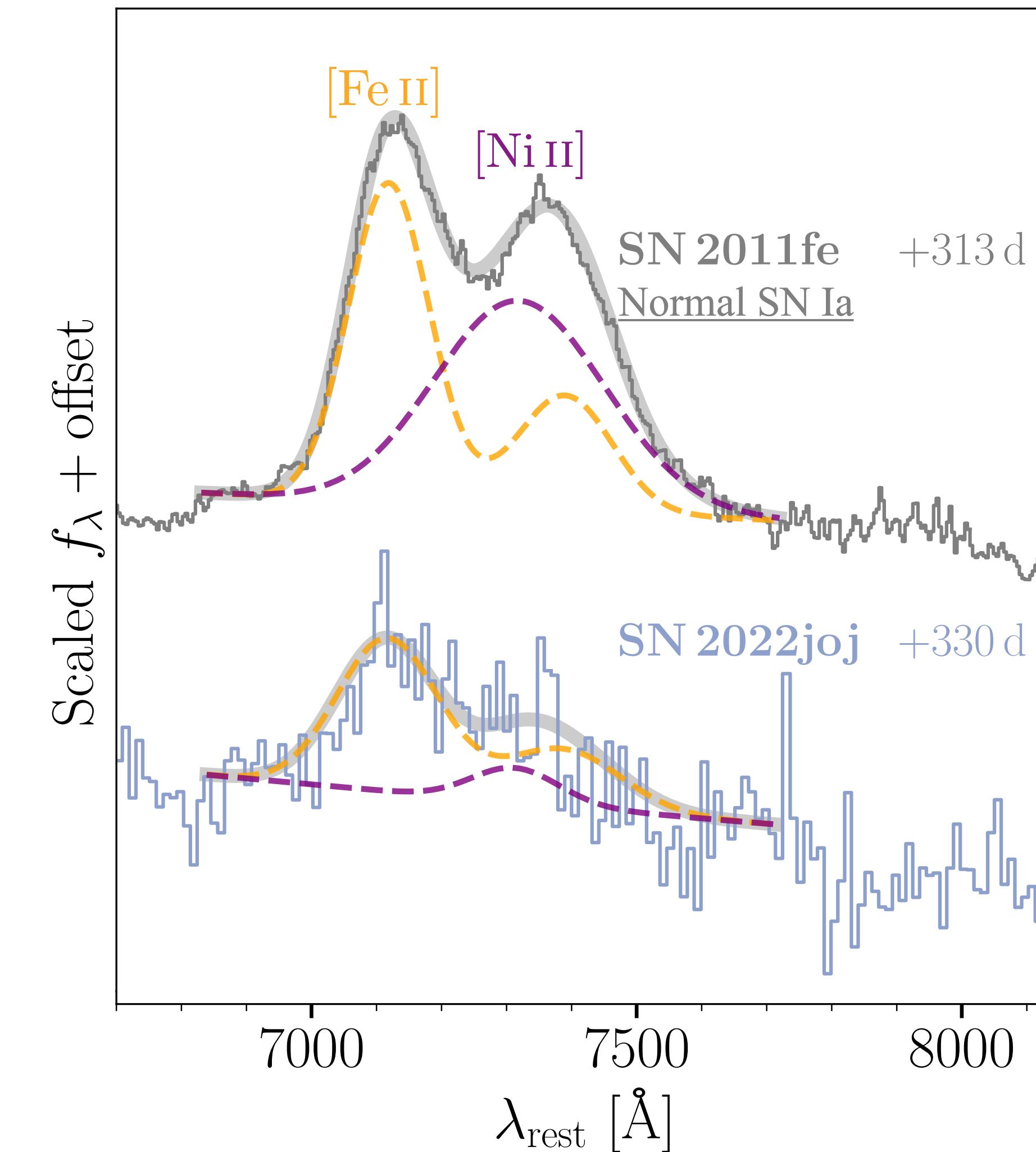
>200 days after explosion

Liu et al., 2023b

Absence of [Ni II]: low central density in the progenitor



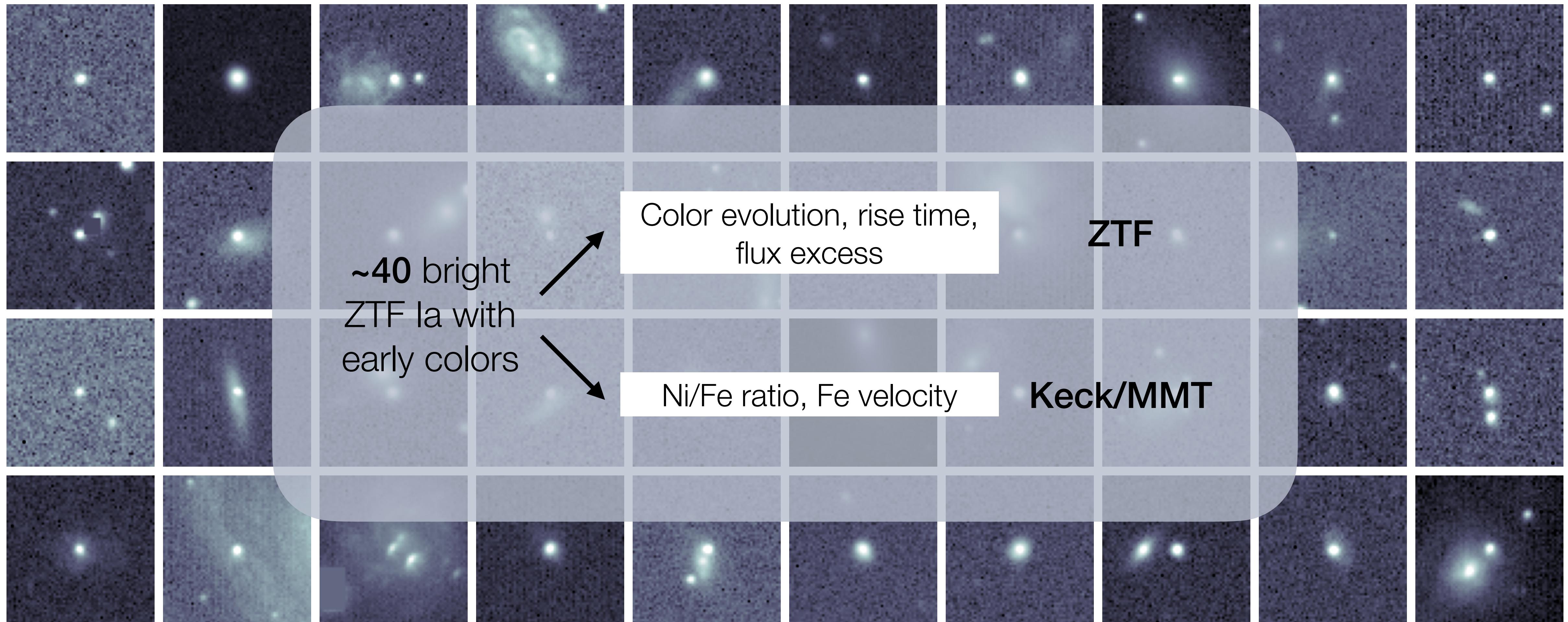
Blueshifted Fe core: sightline opposite to the He detonation



But what about the normal la?

ZTF Early-Late Ia Sample (2023-2024)

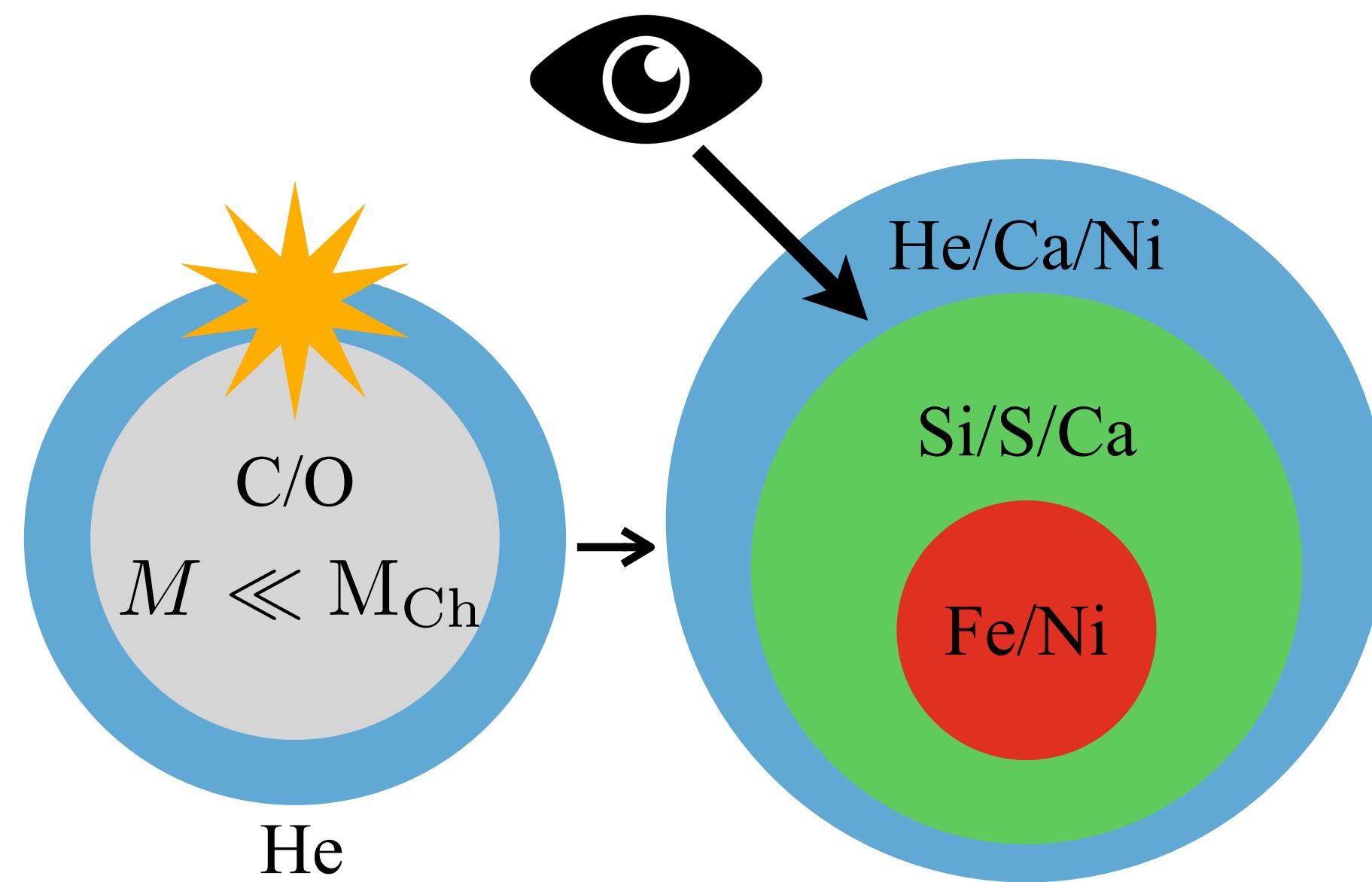
Combining early light curves and nebular spectroscopy



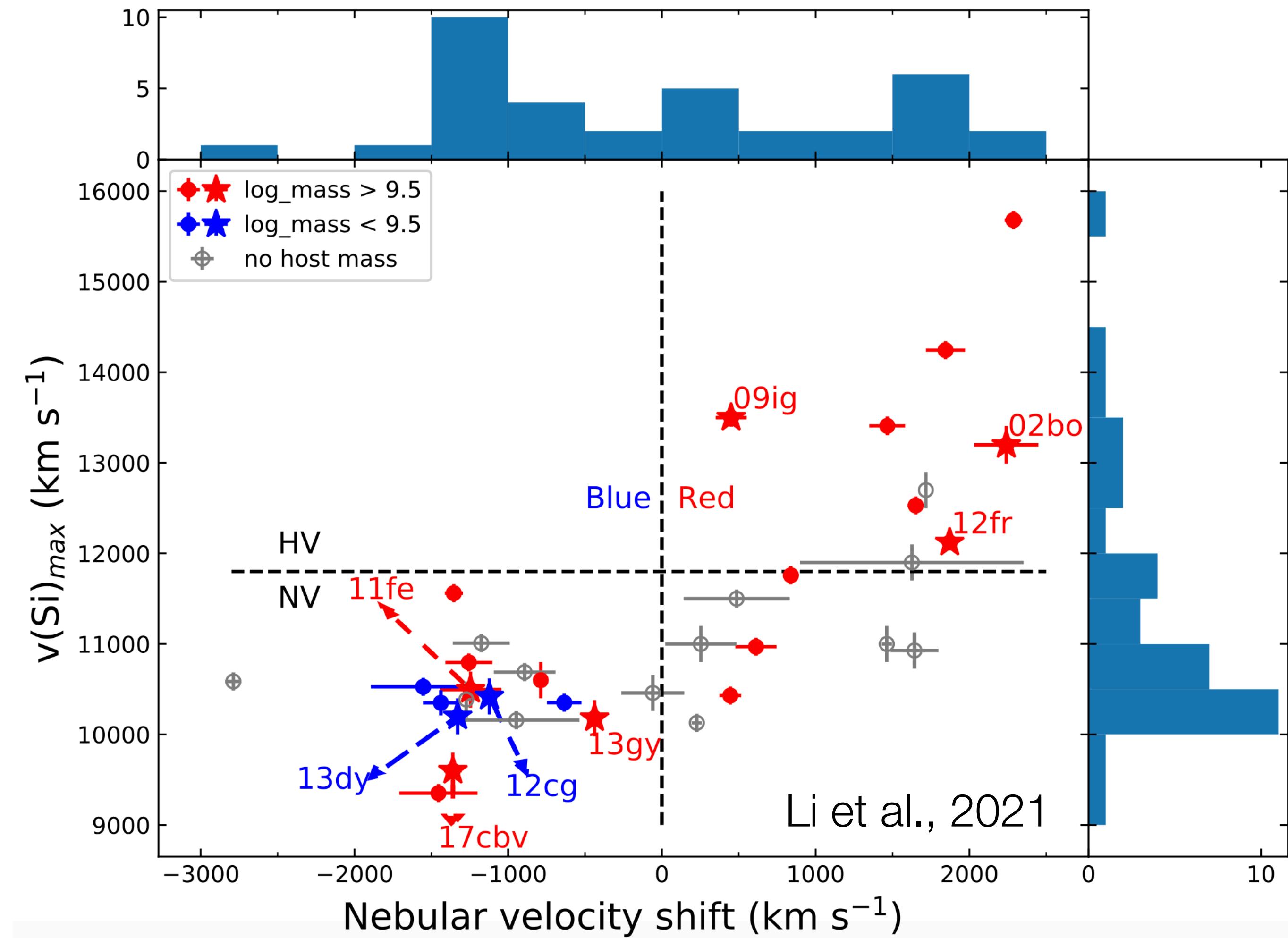
(Not quite) Early + Late Independent probes of the population

Consistent with double detonations at certain viewing angles (Boos et al., 2021)

High Si velocity at peak:
sightline near the He detonation



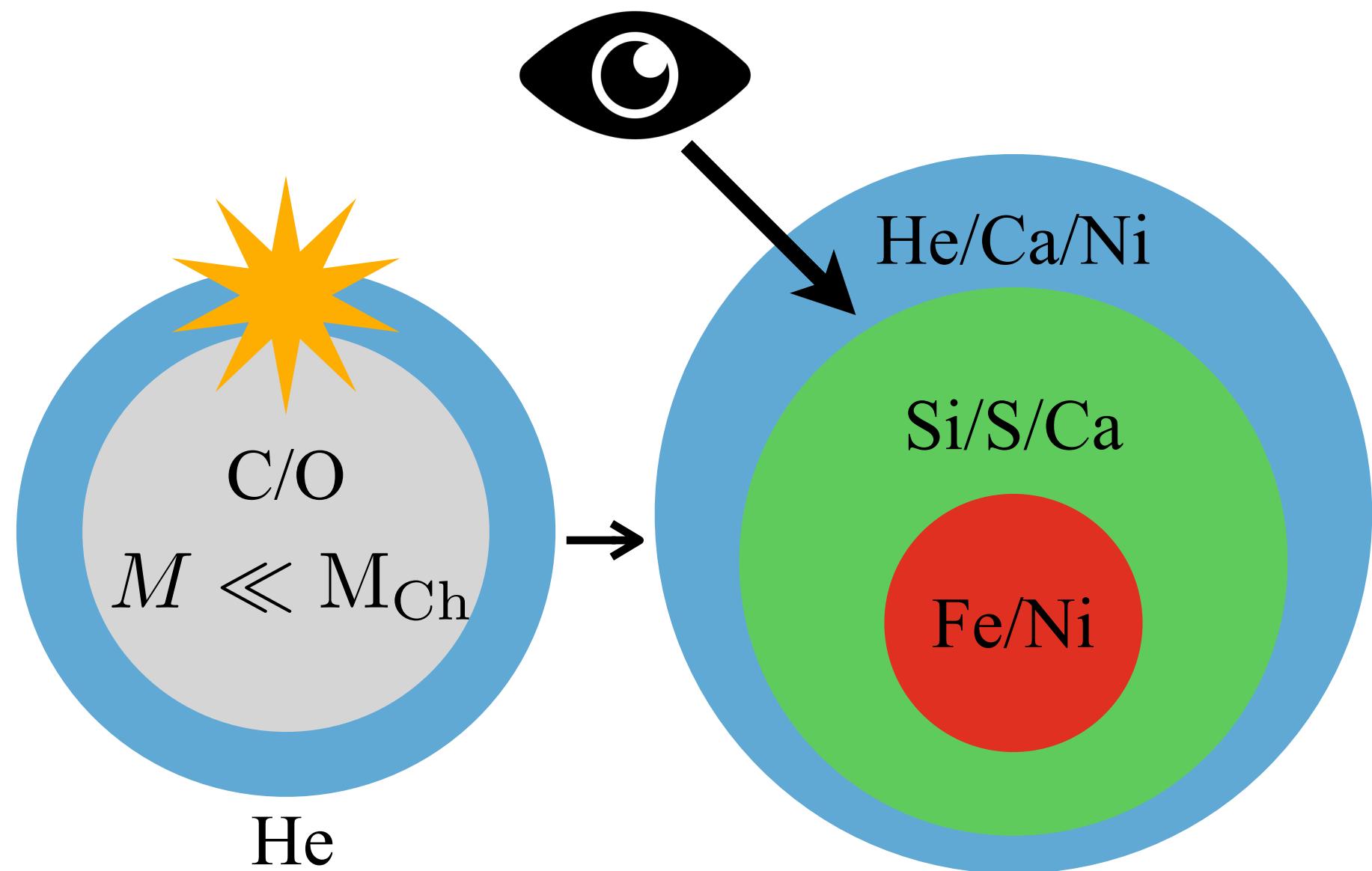
Redshift Fe core at late time:
sightline near the He detonation



Early + Late

Independent probes of the population

Light curves **at early time**



He shell ashes?

Ni-56 asymmetry?

IME asymmetry?

Redshift Fe core at late time:
sightline near the He detonation

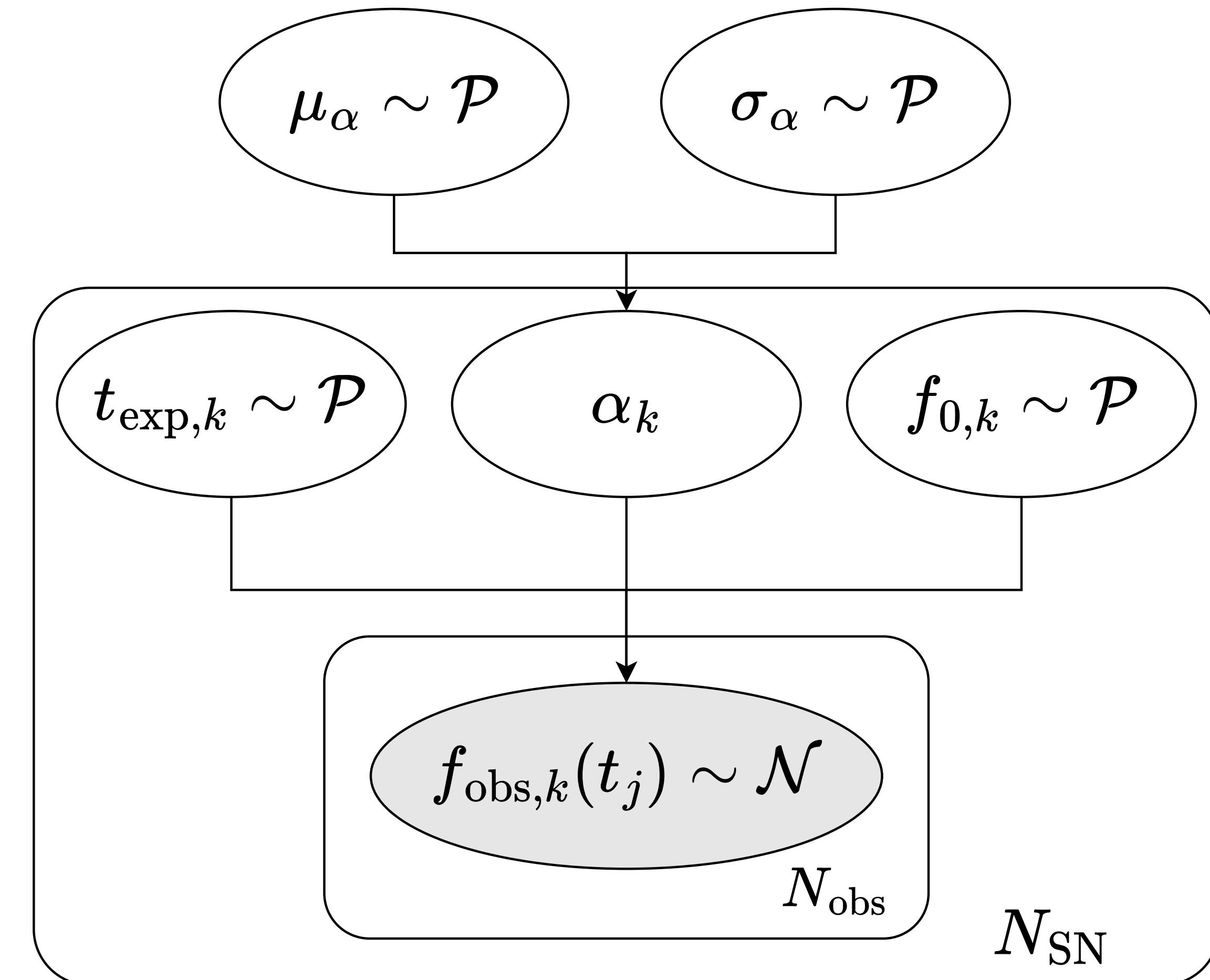
A Systematic Study on the Rise

$\sim 10^3$ SNe Ia from ZTF DR2 (Rigault et al. 2025)

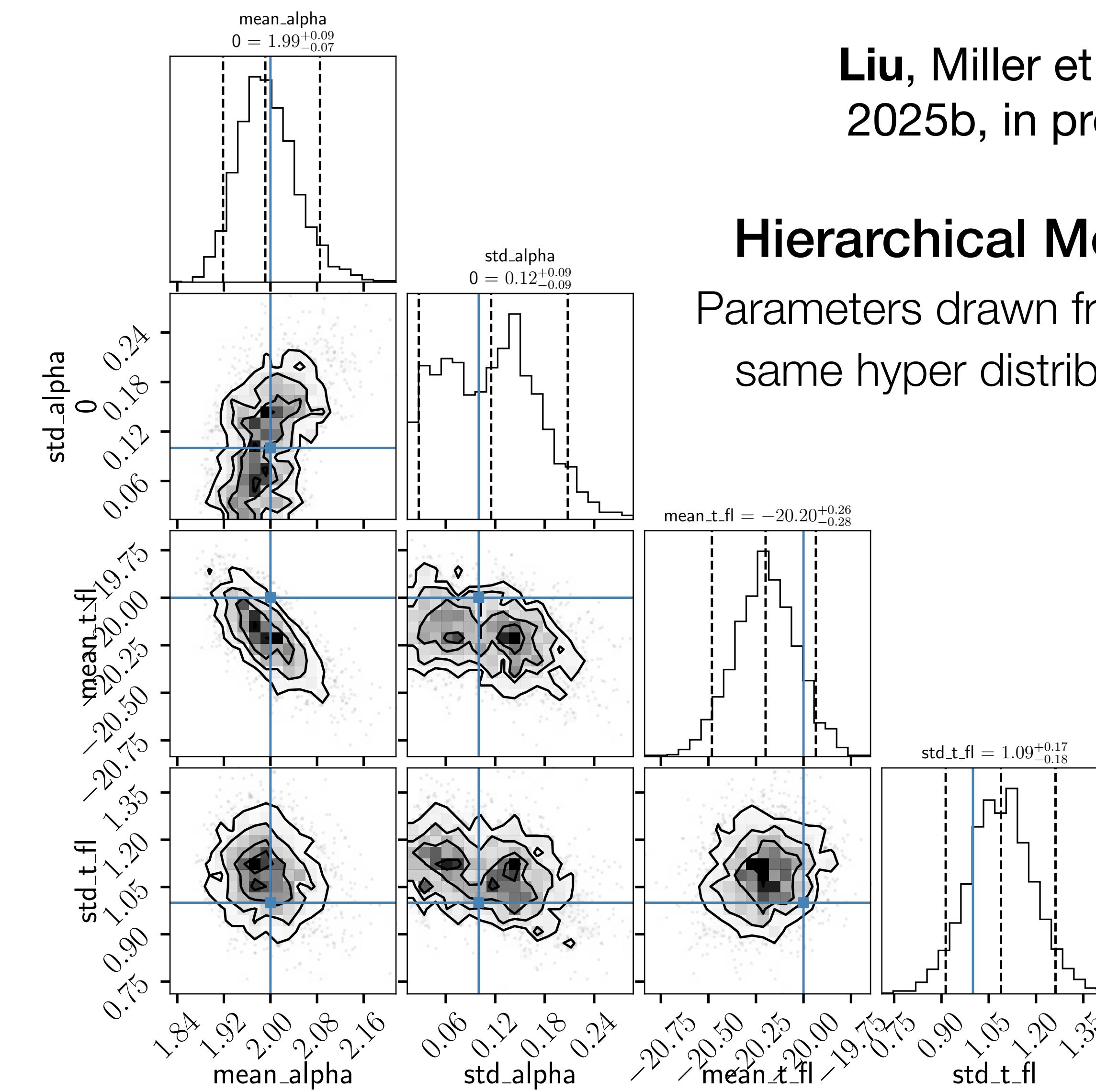
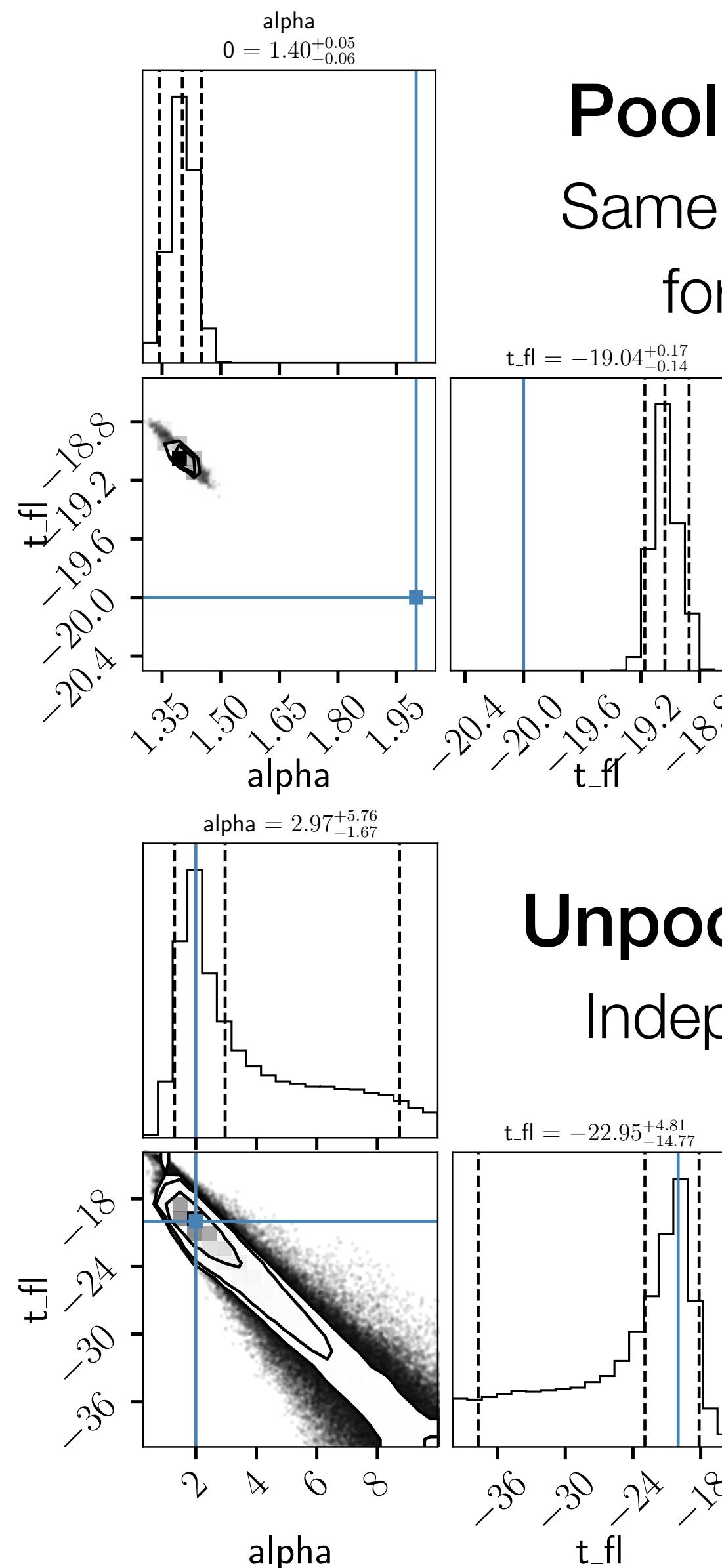
$$f_{\text{obs},k}(t_j) = f_{0,k}(t - t_{\text{exp}})^{\alpha_k}$$

- **Previous model:** fit each SN independently to infer α_k
- **Hierarchical model:** draw α_k from the same distribution controlled by hyper-parameters

Fit **ALL** SNe simultaneously!

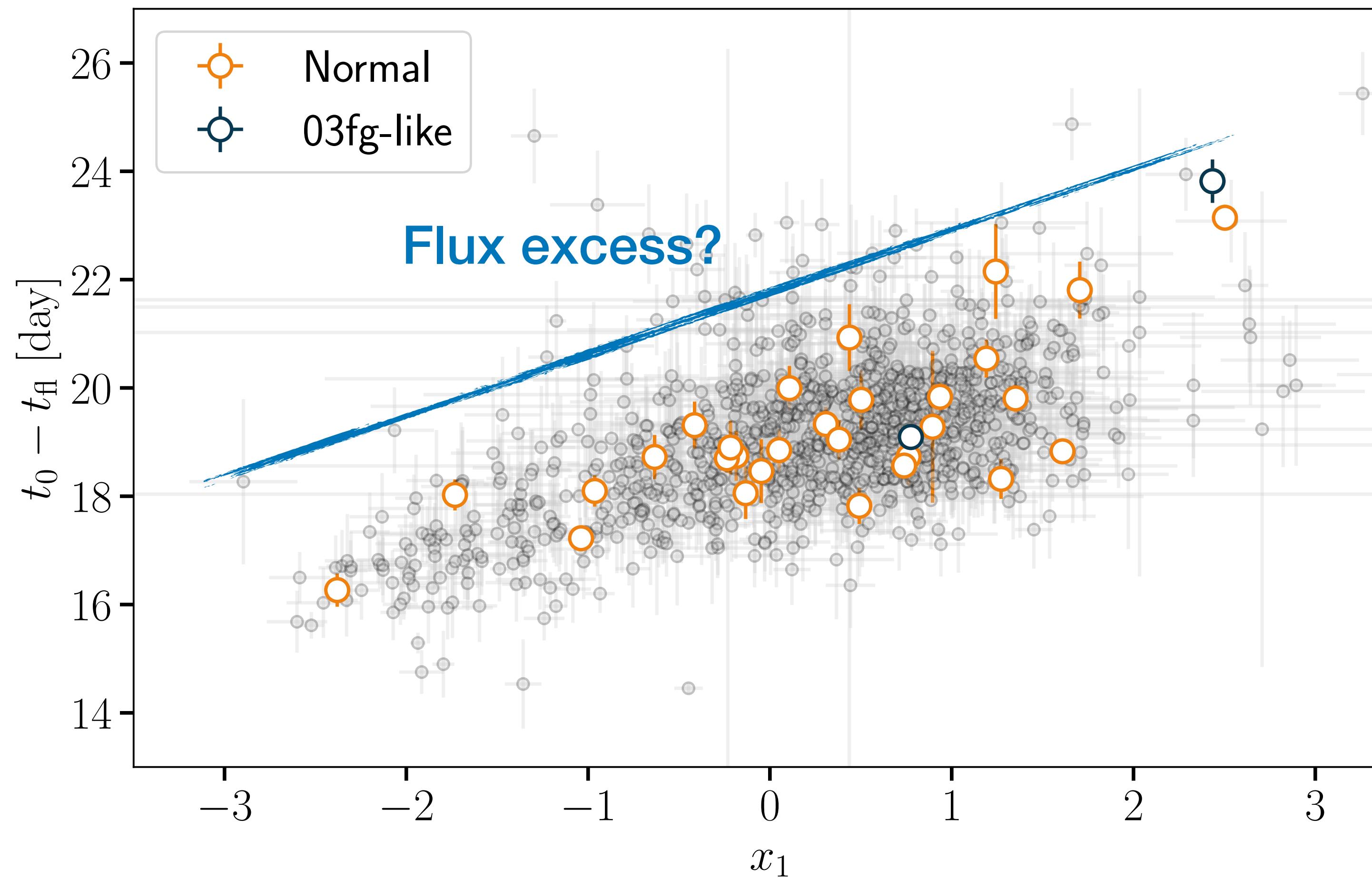


100 Synthetic Light Curves with: $\alpha_i \sim N(2, 0.1)$ $t_{\text{fl}} \sim N(-20, 1)$



A Systematic Study on the Rise

$\sim 10^3$ SNe Ia from ZTF DR2 (Rigault et al. 2025)



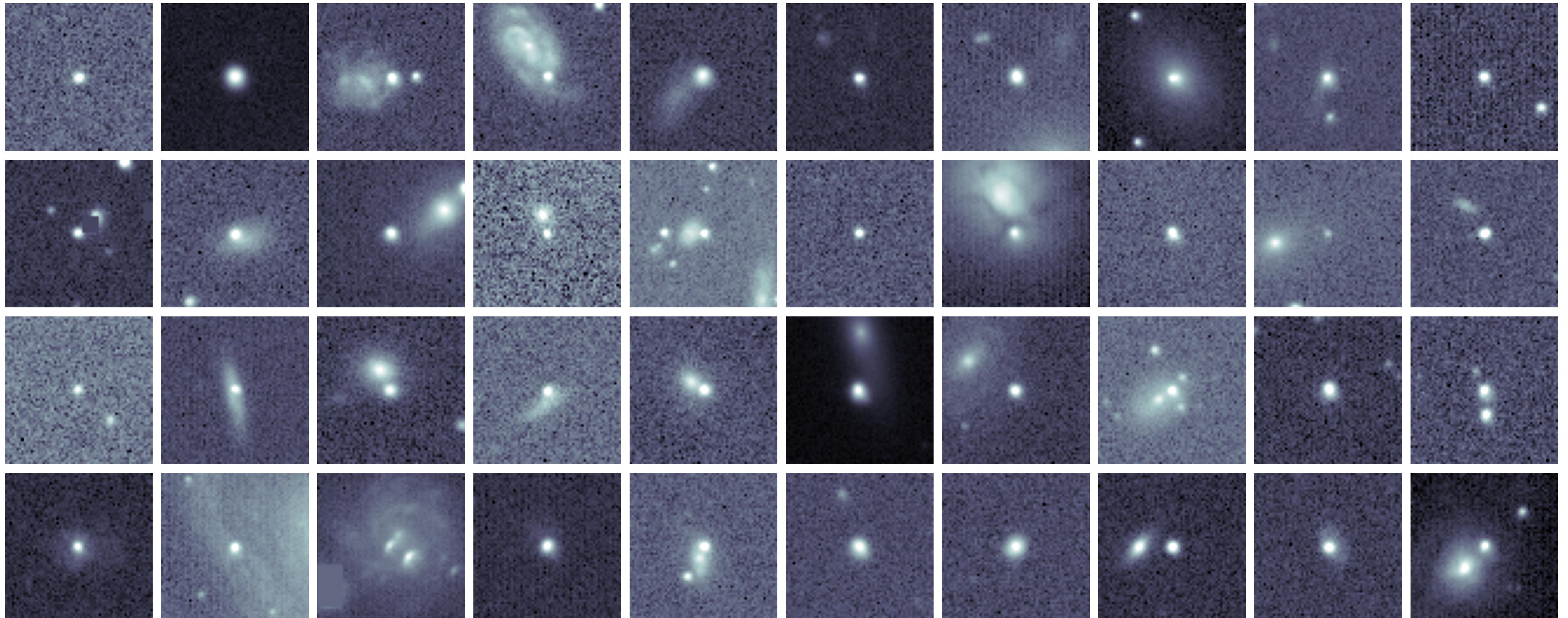
$$t_0 - t_{\text{fl}} \text{ [day]} \sim \mathcal{N}(19.1 \pm 1.6)$$

$$\alpha_g \sim \mathcal{N}(2.28 \pm 0.36)$$

$$\alpha_r \sim \mathcal{N}(2.08 \pm 0.28)$$

ZTF Early-Late Ia Sample (2023-2024)

Significant host contamination in >50% of the targets



HostSub_GP

Liu & Miller 2025 - on arXiv today!

Modeling the 2d spectrum of host galaxies with Gaussian process (GP) for better background subtraction in supernova spectroscopy.

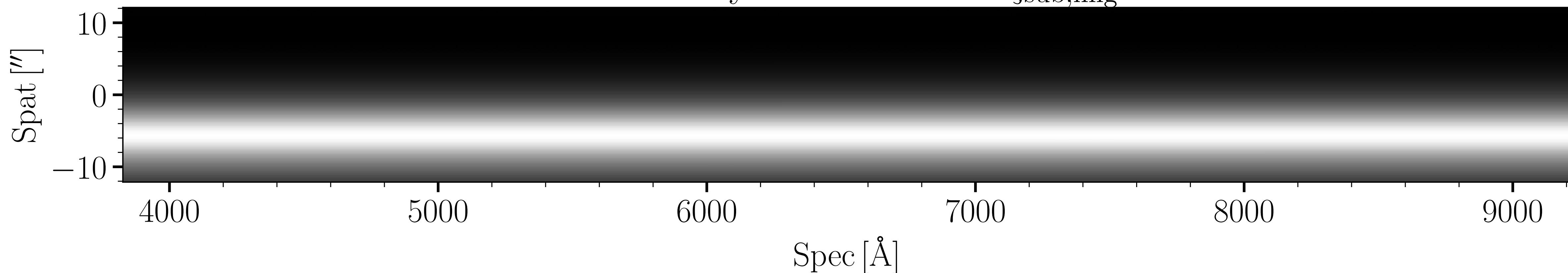
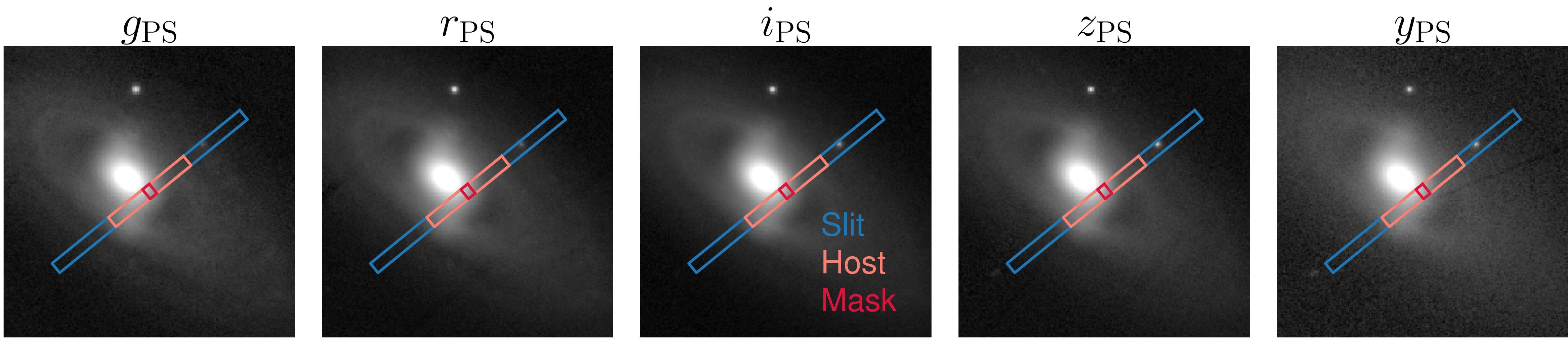
- Fully **Python** + available at GitHub
- Leveraging **archival imaging** as priors
- Accelerated by JAX: **Just-in-time** compilation + **automatic differentiation**

Robust galaxy background removal in a few
minutes on your laptop!

HOSTSUB_GP Pipeline

Building the prior - normalized 2D profiles

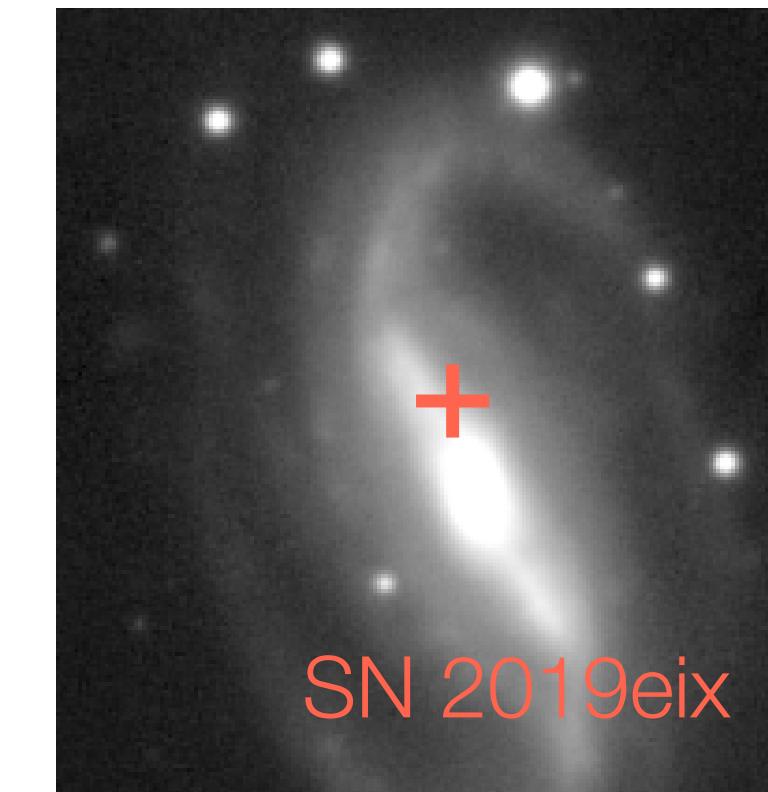
Liu & Miller 2025



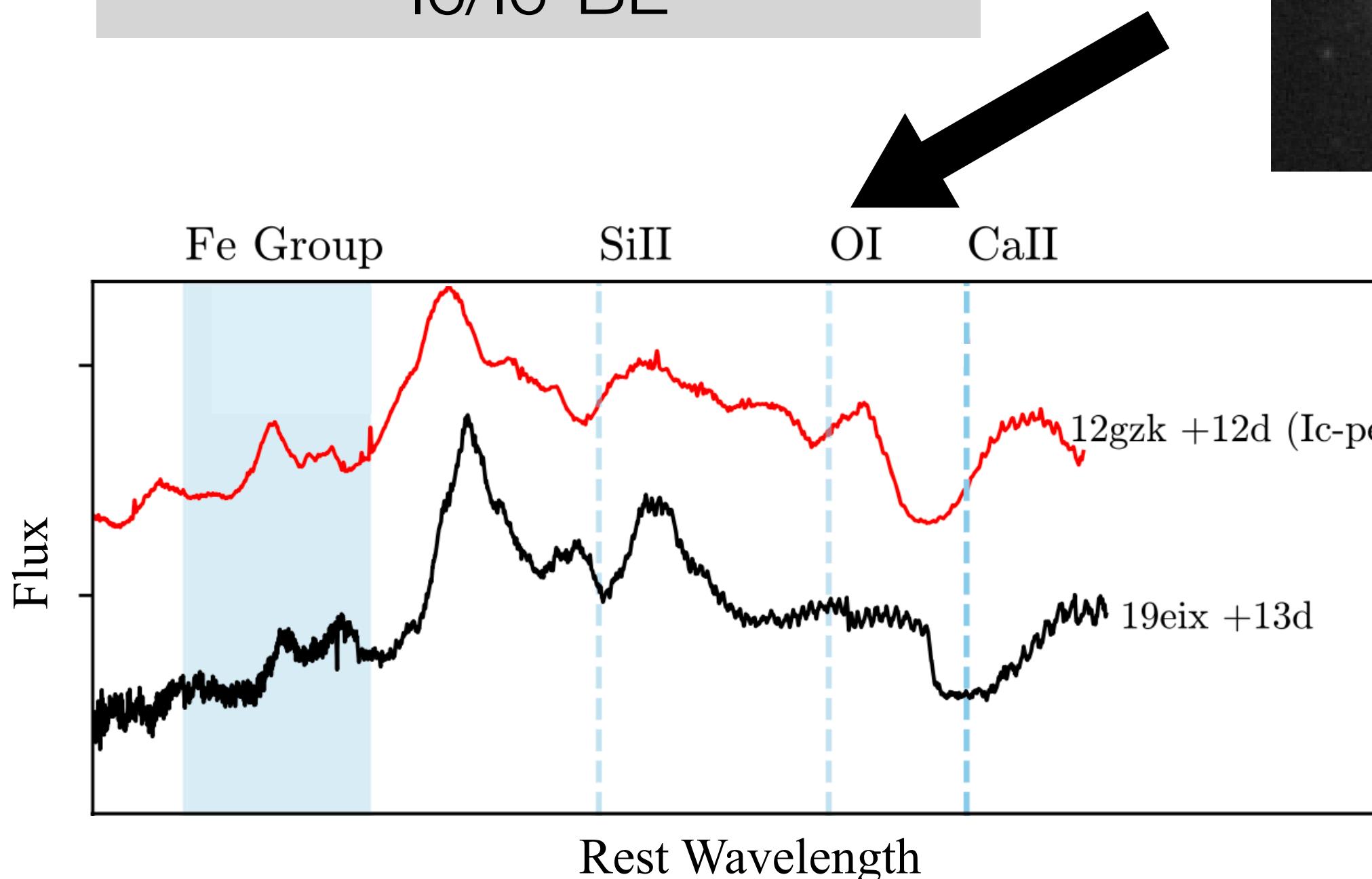
HOSTSUB_GP Pipeline

Revisiting SN 2019eix – peculiar Ia or Ic?

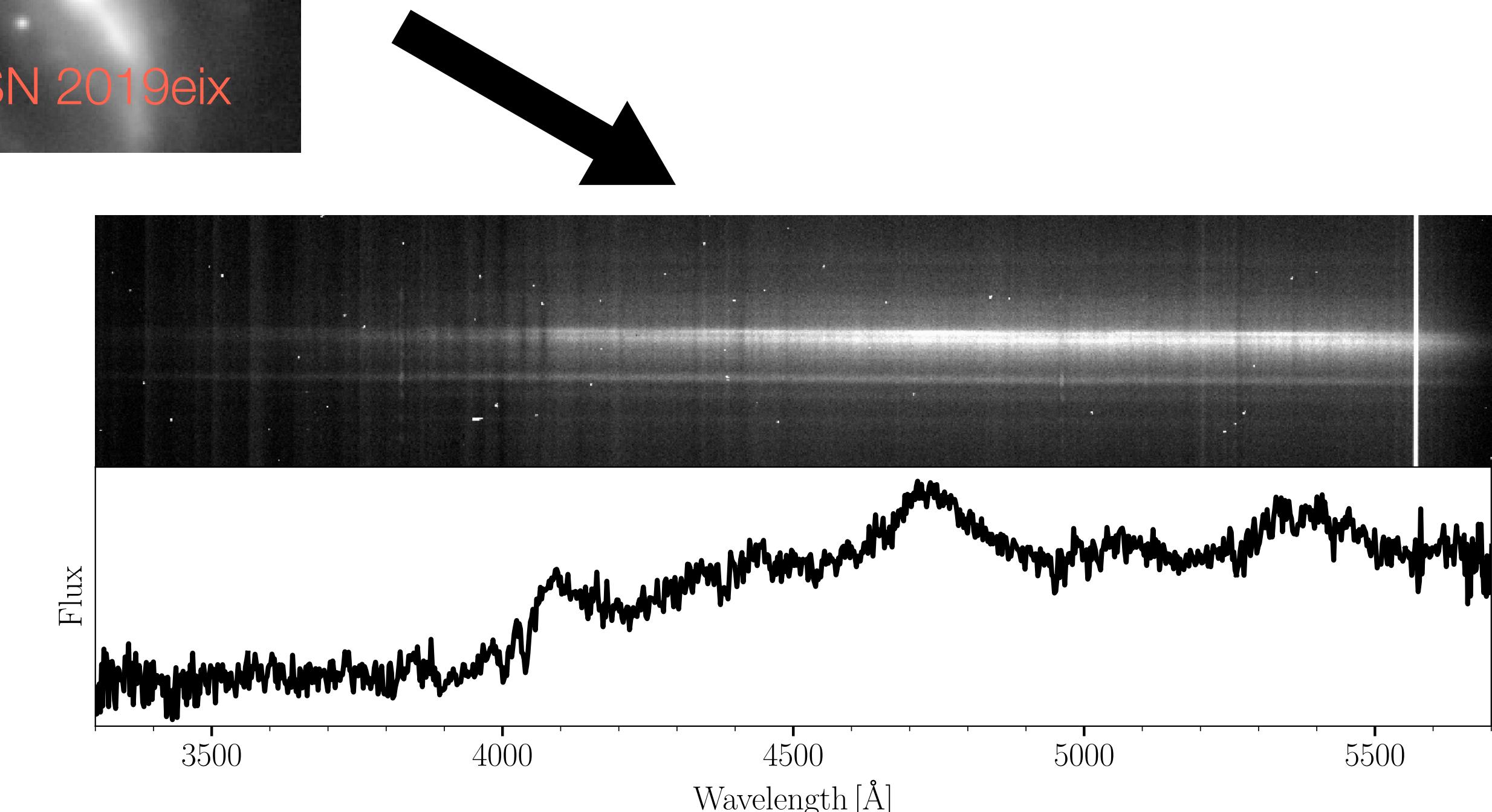
Photospheric phase
Resembling both double-detonation Ia and some Ic/Ic-BL



Nebular phase
Host galaxy subtraction is extremely challenging



Padilla Gonzalez et al., 2023

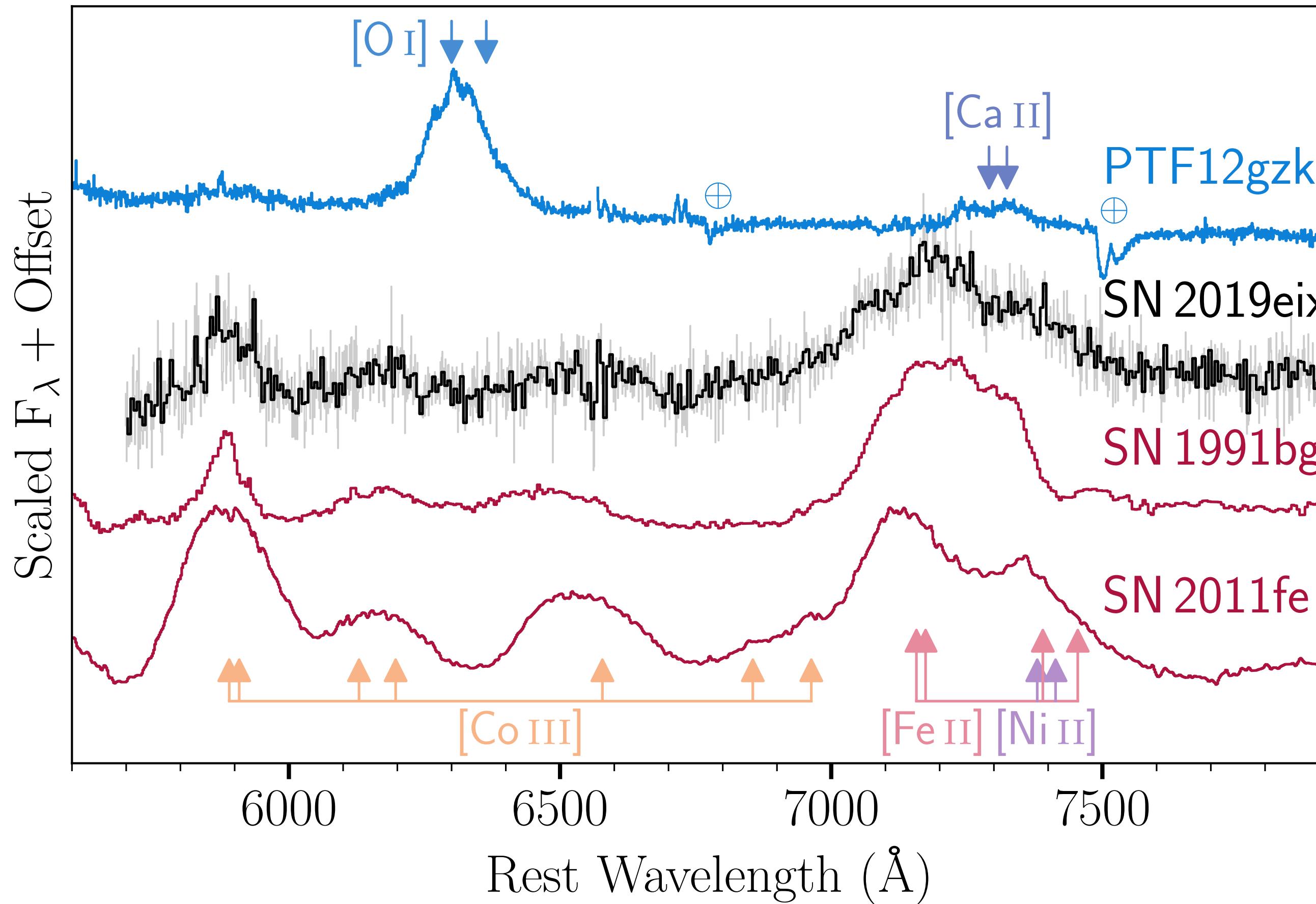


PI: Fremling

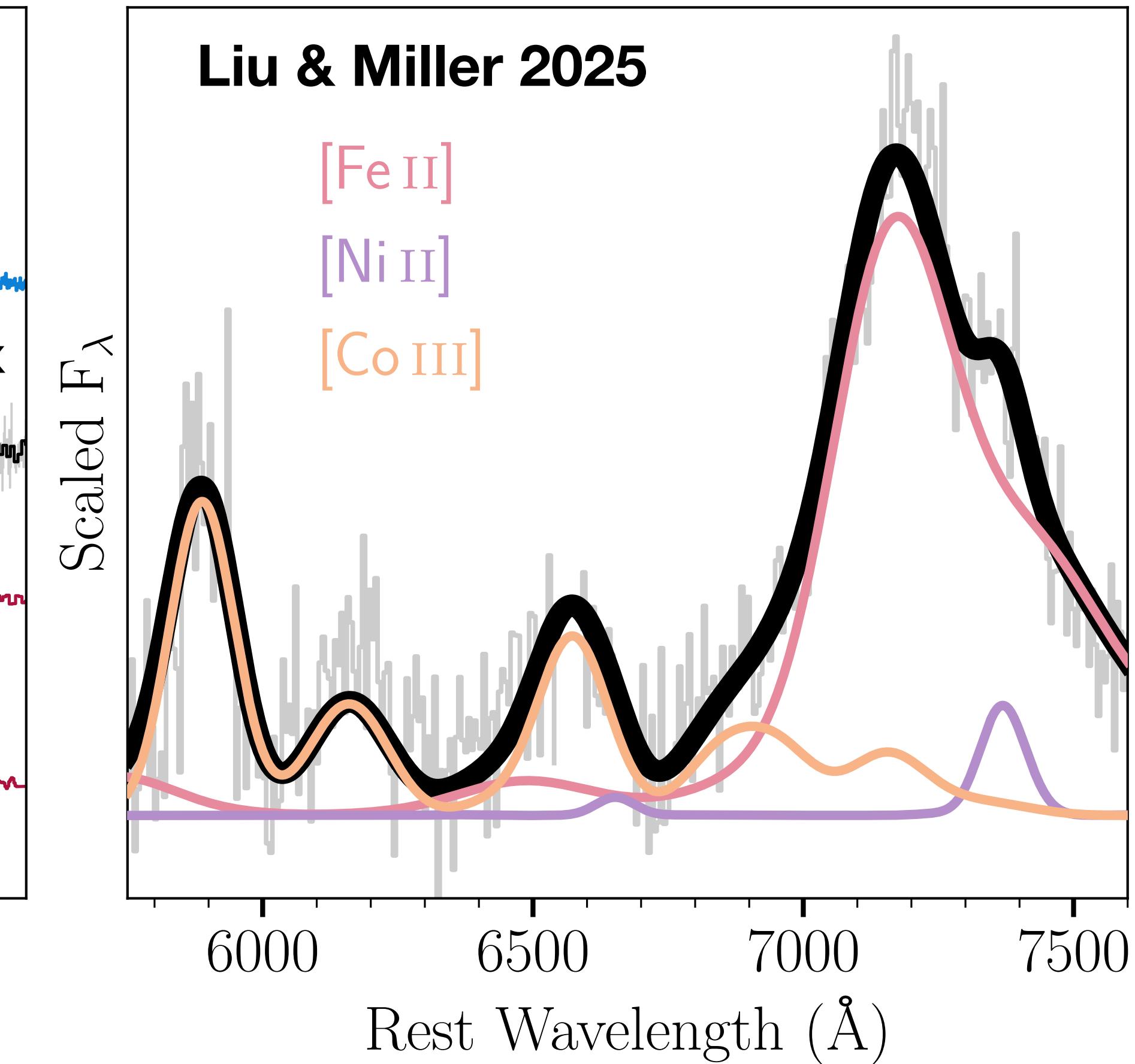
HOSTSUB_GP Pipeline

Revisiting SN 2019eix – double-detonation Ia!

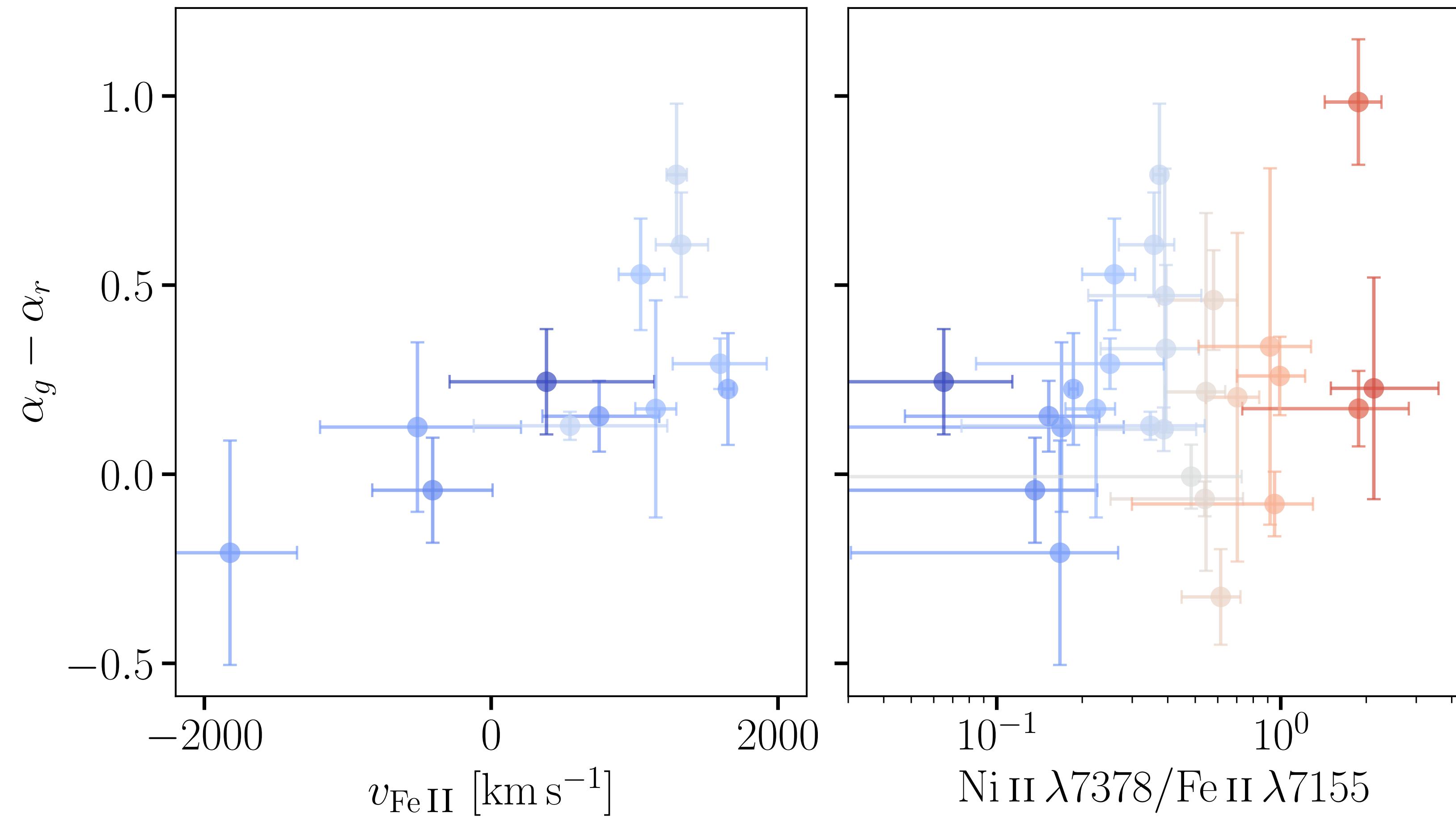
Strong Fe, Ni, Co, no [O I]



$$n_{\text{Ni II}}/n_{\text{Fe II}} \simeq 0.5\%$$



Early v.s. Late Color evolution v.s. Fe core velocity



Key Takeaways

Still Far from Fully Understanding Ia Progenitors, but...

- **DO NOT** ignore rare weirdos!
- **DO NOT** fit early rises individually!
- **DO NOT** exclude host contaminated SNe from your sample!
- **Do** analyze early and late-time observables jointly!