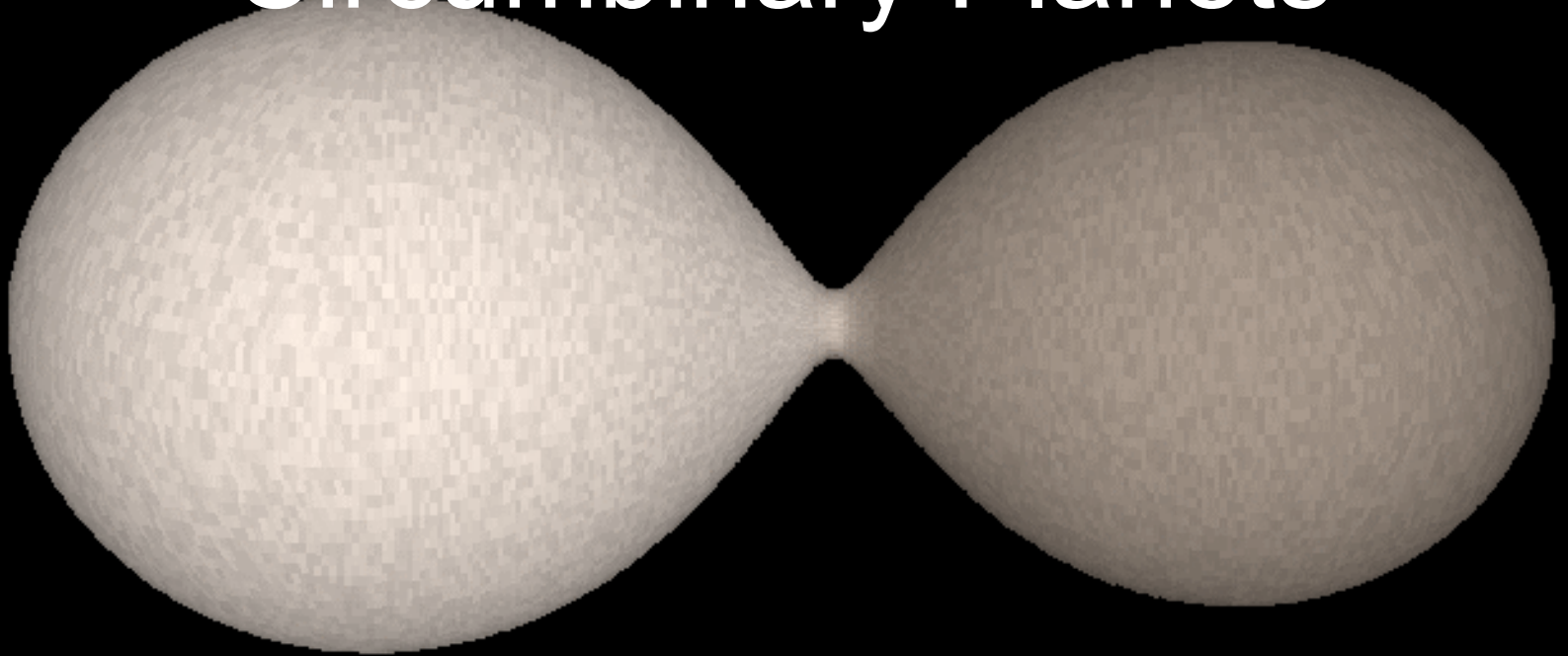


CB-BLS: Identifying Transiting Circumbinary Planets



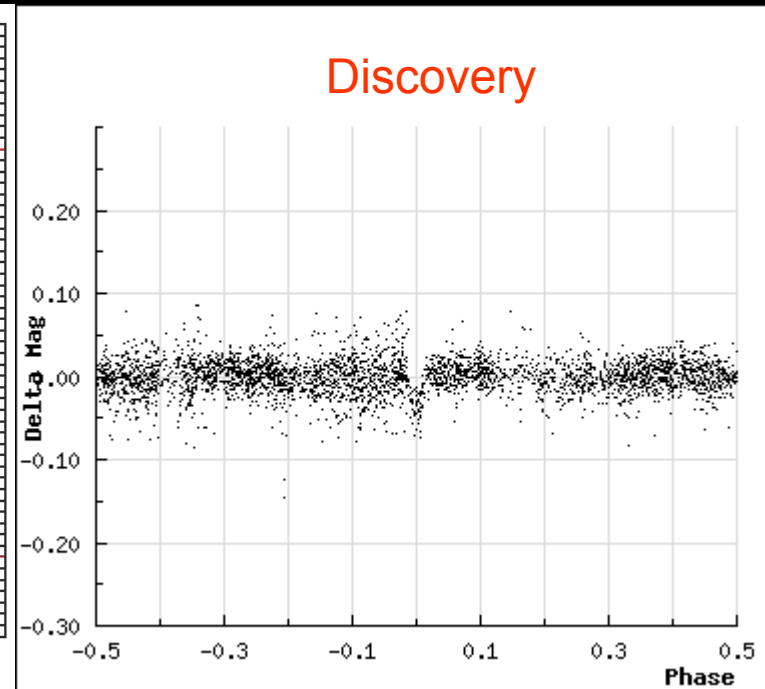
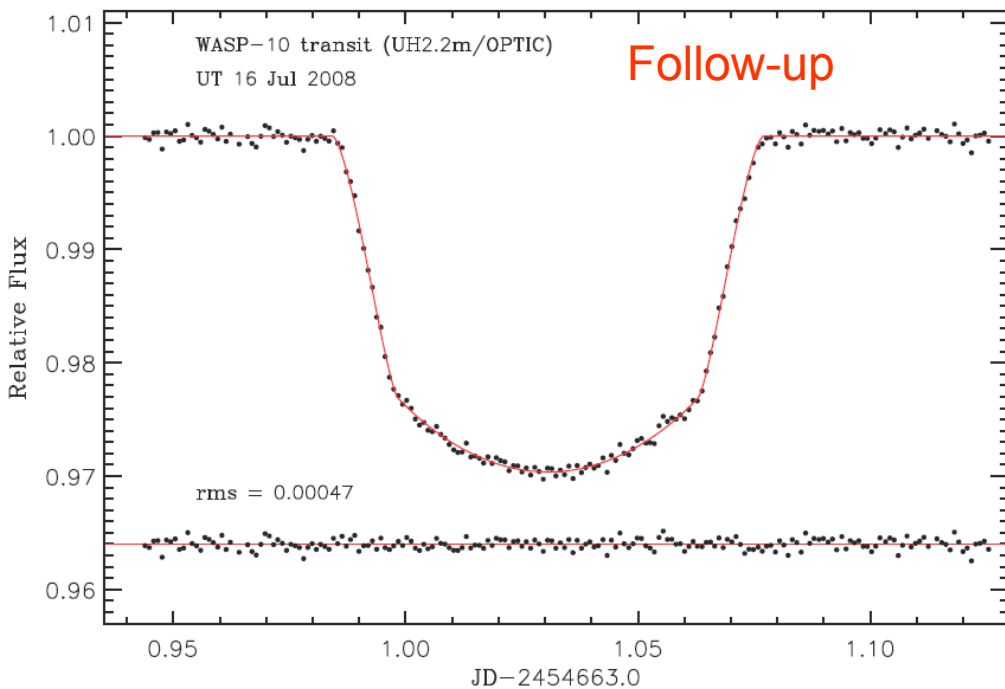
Aviv Ofir

Tel Aviv University

MNRAS 387, 1597 + updates

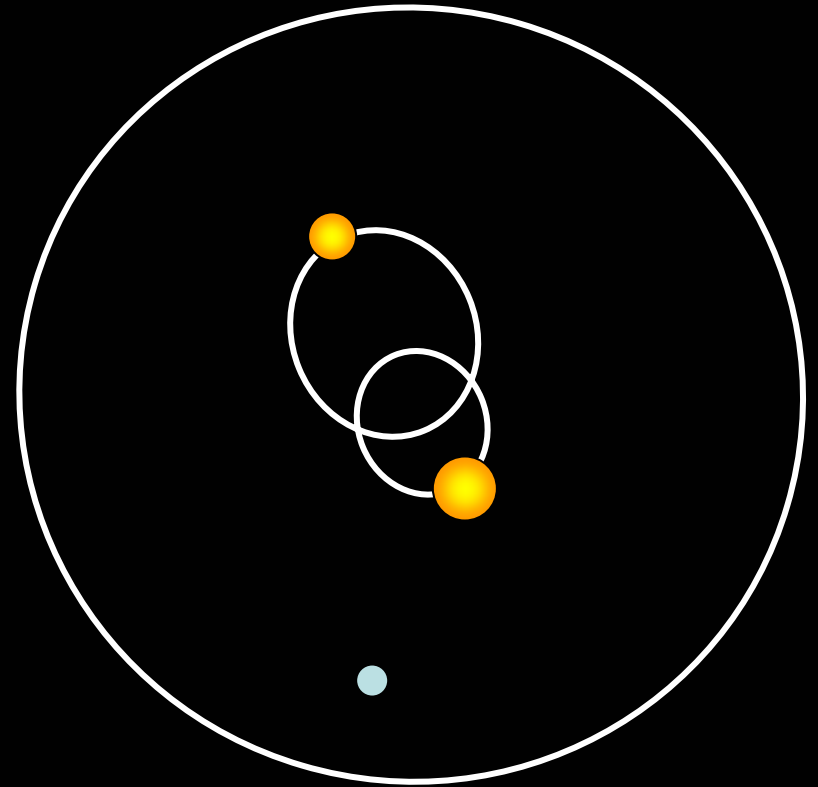
Background

- ~330 extra-solar planets are known.
- Detected in variety of techniques:
 - RV, transits, microlensing, imaging, timing, ...
- The basic transit light curve is simple:



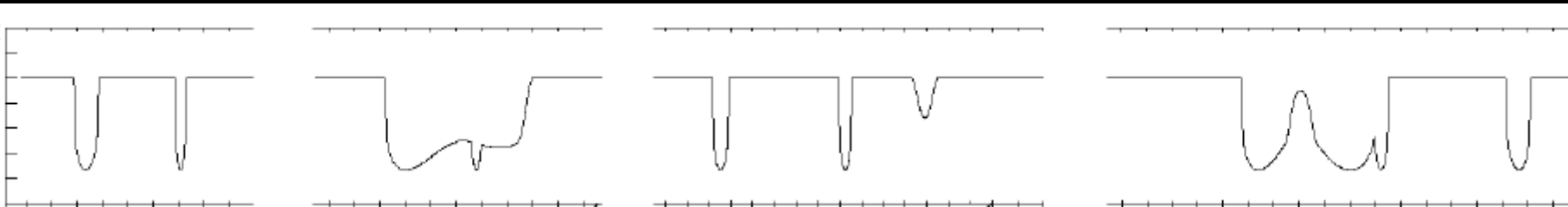
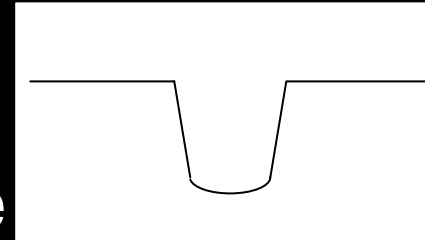
Identifying Transiting Circumbinary Planets

- The Challenges
- Solution: “CB-BLS”
- Updates
- Conclusions



The Challenges (1)

- **Goal: detecting transiting exoplanets around eclipsing binaries (EBs).**
- Transiting planets around single stars have precise depth, duration, time
- Transiting circumbinary (CB) planets are expected to have special photometric and temporal characteristics





The Challenges (2)

Nothing is constant:

- **Photometric:** Signal not constant in depth
 - Signal = blocked / total instantaneous flux, but:

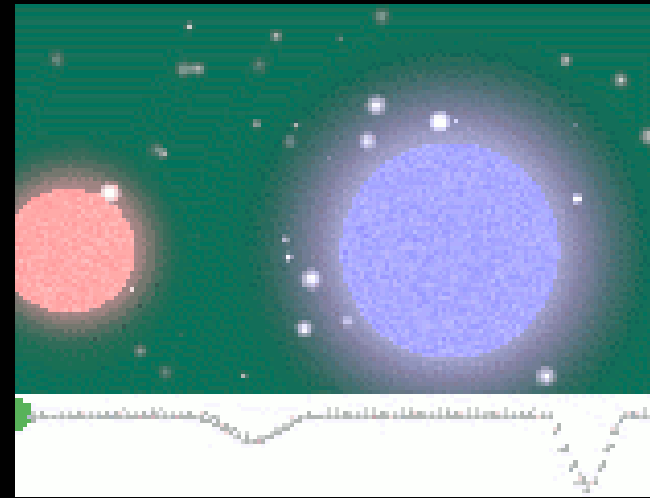
Different surface
brightness of
components

Depends on the binary orbital
phase (ellipsoidal variation,
eclipses).

- **Temporal:**
 - The transits are not periodic
 - Transit durations are highly non-uniform

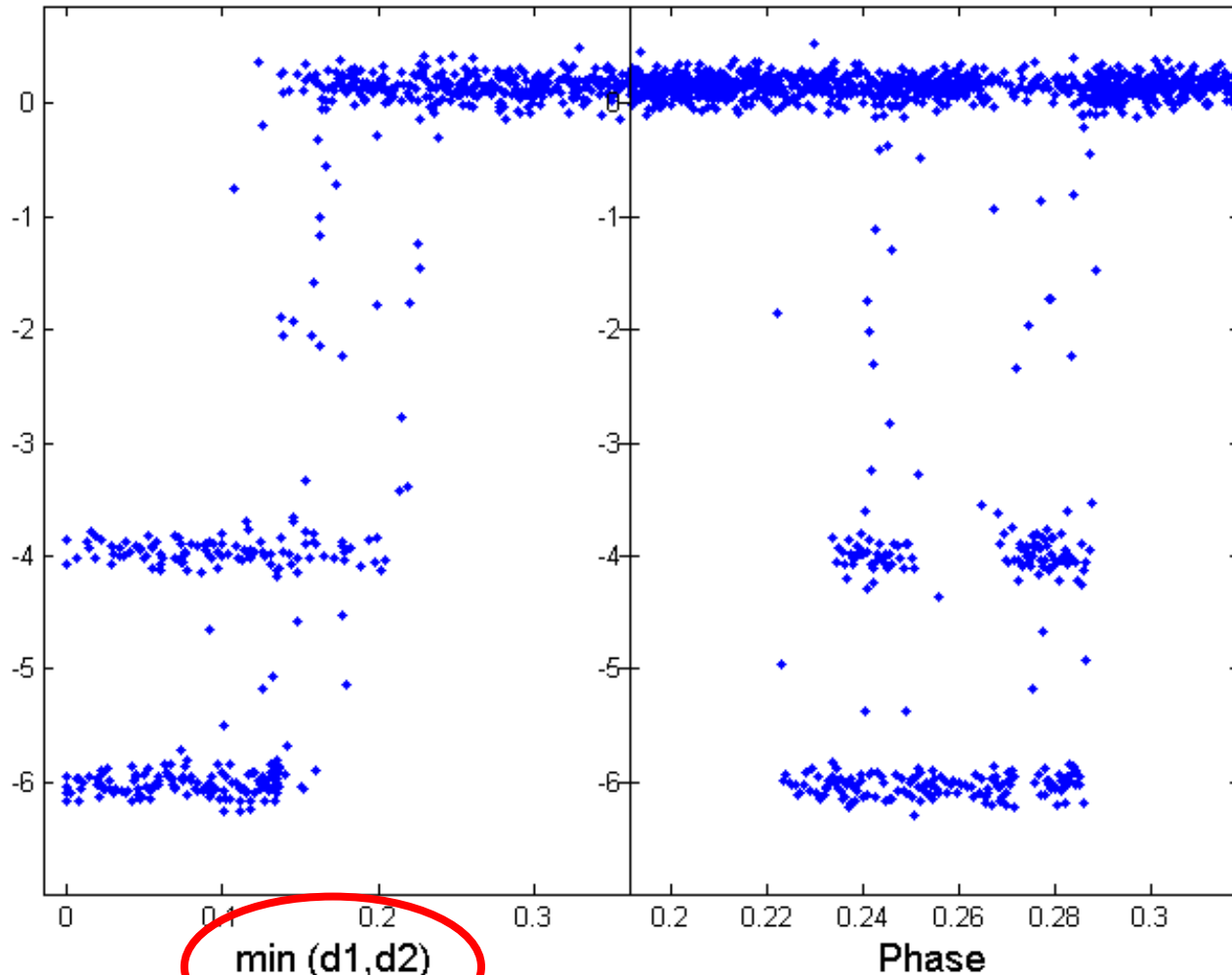
The **CB-BLS** Solution - Overview

- Solve EB (P_B , T_0 , e , ω , i_B , $R_{1,2}$, J)
- **Photometric** characteristics
 - Regularize the depths:
 - normalize LC,EB model to $\max(\text{model flux})=1$
 - multiply LC residuals (and errors) by the EB model
 - Define a CB-BLS statistic that allows for different surface brightness / use J
- **Temporal** characteristics → **Geometric**
 - Transits are NOT a function of time, but of geometry: the alignment of celestial bodies



Solution - Geometric

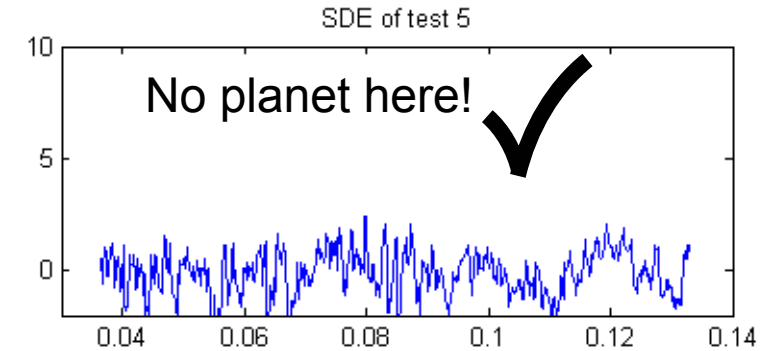
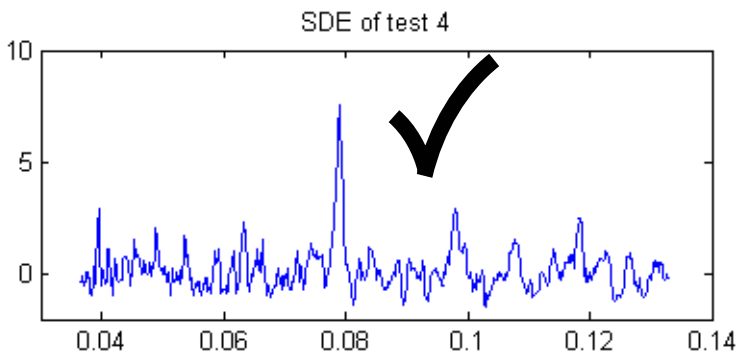
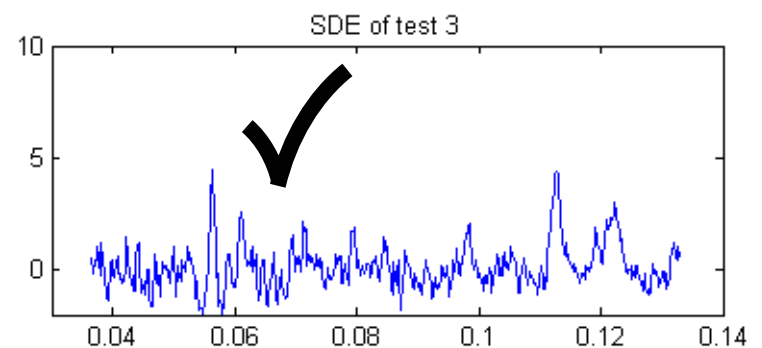
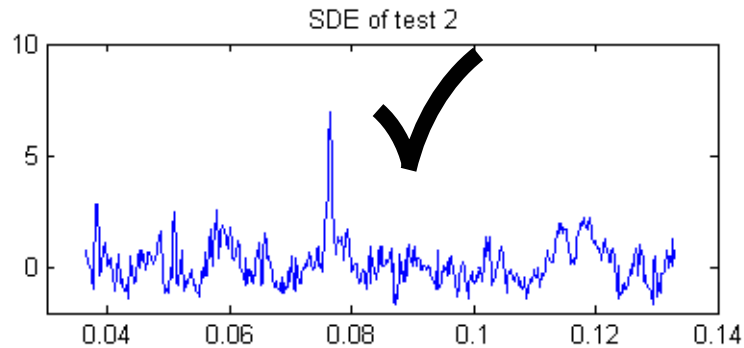
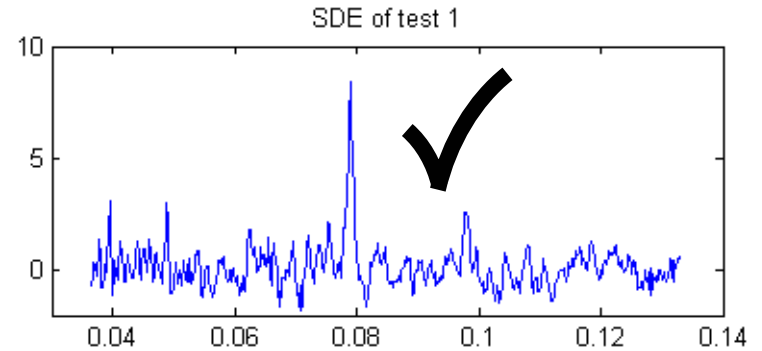
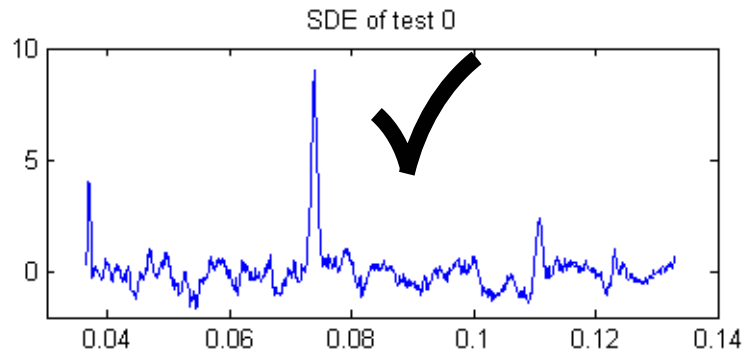
Folding a CB planet light curve in phase and in projected distance (0.01% noise)



- Assume \rightarrow sk
- Assume \rightarrow sk
- At ea
exac
- Visual

$+m_2)$
 $\varphi_0)$
 ed
 t is
)

Updates Beyond Ofir (2008)



Conclusions – **CB-BLS**: A New Tool

- CB-BLS is general, sensitive, quick
 - CB-BLS allows to harness **existing datasets** to the detection of transiting CB planets
 - CB-BLS efficient even when **noise > signal** (~100% recovery for 1_{Jup} around 2 x ~Sun @ 1% noise)
-
- Remaining difficulty: EB modeling errors
 - Will add red noise to the residuals
 - But: Different time scale, periodic, and false-alarms can be identified



Thank You.

Questions?