

$\hat{\theta}$: local horizon
 b semiminor axis

shape size $\epsilon < 0$



$\gamma > 0$ ascending
 $\gamma < 0$ descending

$$\hat{h} = \frac{\hat{h}}{\hat{h}}$$

$$r_p = \frac{p}{1+e} = a(1-e)$$

$$r_a = \frac{p}{1-e} = a(1+e)$$

$$\hat{h} = \hat{e} \times \hat{p}$$
$$p = \frac{h^2}{\mu}$$
$$p = a(1 - e^2)$$

$$a = r = p$$

$$a = r = p \quad \varepsilon = -\frac{\mu}{2} = \frac{v^2}{2} - \frac{\mu}{2}$$

$$N_c = \sqrt{\frac{\mu}{r}}$$

top half orbit

γ is always above
L.H. (local horizon)

positive γ

bottom half orbit

γ is always below L.H.

negative δ

- distance between 2 particles are growing

descending

true anomaly $180^\circ \sim 360^\circ$ (or $-180^\circ \sim 0^\circ$)

distance is decreasing

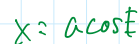
$$\vec{v} = v_r \hat{r} + v_\theta \hat{\theta}$$

→ v_r is positive

descending

true anomaly $180^\circ \sim 360^\circ$ (or $-180^\circ \sim 0^\circ$)

distance is decreasing



$$y^2 = b \sin E$$

$$\frac{dy}{dx} = \frac{\frac{dy}{dE}}{\frac{dx}{dE}} = \frac{b \cos E}{-a \sin E}$$



$$\varphi = \arctan\left(-\frac{b \cos k}{a \sin k}\right)$$

General Ellipse:

General Ellipse:

$$E = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a}$$

$v_c = \sqrt{\frac{\mu}{r}} \Rightarrow v_c^2 = \frac{\mu}{r}$ $E = -\frac{\mu}{2a}$
negative

$$v^2 = \frac{2\mu}{r} - \frac{\mu}{a} = 2v_c^2 - \underbrace{\frac{\mu}{a}}_{\text{positive by definition}}$$

positive by definition $\frac{\mu}{a} > 0$

$$v^2 = 2v_c^2 - \text{stuff}$$

$v < \sqrt{2} v_c$ → always!!

$$\frac{dA}{dt} = \frac{h}{2} \rightarrow dt = \frac{2}{h} dA$$

period: $IP = \frac{2}{h} (\pi ab)$ area of ellipse $b = a\sqrt{1-e^2}$

$$= \frac{2}{h} \pi a (a \sqrt{1-e^2})$$

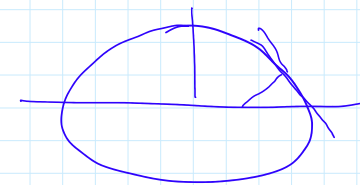
$$= \frac{2\pi a^2 \sqrt{1-e^2}}{\sqrt{\mu p}} \quad \because p = a(1-e^2)$$

$$P = 2\pi \sqrt{\frac{a^3}{\mu}}$$

$$\hat{r} = (\cos \theta^*, \sinh \theta^*)$$

$$\hat{\theta} = \hat{r} \begin{pmatrix} \cos 90^\circ & -\sin 90^\circ \\ \sin 90^\circ & \cos 90^\circ \end{pmatrix}$$

$$0^{\frac{1}{2}} + 90^0$$



F3

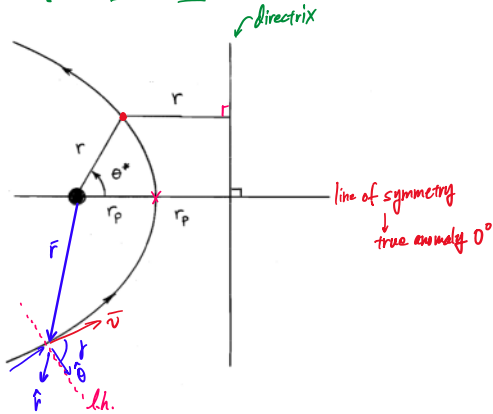
Parabola

$e=1$

$a=\infty$

$\varepsilon=0$

$\varepsilon = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a} \xrightarrow{a \rightarrow \infty} 0$



Orbit NOT closed; particle leaves vicinity of attracting body

$$\frac{v^2}{2} - \frac{\mu}{r} = 0 \rightarrow v^2 = \frac{2\mu}{r} = 2v_c^2 \Rightarrow v = \sqrt{2} v_c$$

escape speed at distance r

$$\frac{v_\infty^2}{2} - \frac{\mu}{r_\infty} = 0 \rightarrow v_\infty = 0$$

can just barely escape grav influence of attracting particle

relative velocity $\rightarrow \emptyset$

F4

Hyperbola

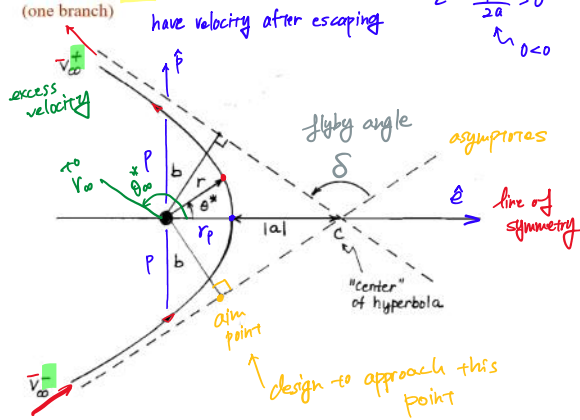
$e > 1$

$a < 0$ (by convention)

$\varepsilon > 0$

$$\varepsilon = \frac{\mu}{2a} > 0$$

$\hookrightarrow a < 0$

rel \rightarrow

$$\varepsilon = \frac{v^2}{2} - \frac{\mu}{r} \text{ const.}$$

$$\varepsilon = -\frac{\mu}{2a} = +\frac{\mu}{2|a|} = \frac{v_\infty^2}{2} - \frac{\mu}{r_\infty} = \frac{v_\infty^2}{2} \quad \varepsilon > 0 \quad v_\infty \neq 0$$

$$r_p = a(1-e) = |a|(e-1) \quad e > 1 \quad a < 0 \Rightarrow \oplus$$

$$p = a(1-e^2) = |a|(e^2-1) \quad e > 1 \quad a < 0 \rightarrow \oplus$$

$$r = \frac{p}{1+e \cos \theta^*} = \frac{p}{1+e \cos \theta^*}$$

F5

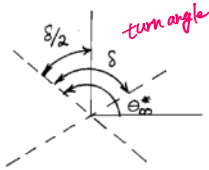
parallel to asymptote

$$r_{\infty} = \frac{|a|(e^2 - 1)}{1 + e \cos \theta_{\infty}^*}$$

OR

$$1 + e \cos \theta_{\infty}^* = \frac{|a|(e^2 - 1)}{r_{\infty}}$$

$$\rightarrow \cos \theta_{\infty}^* = -\frac{1}{e}$$



$$\theta_{\infty}^* = \frac{\delta}{2} + 90^\circ$$

$$\cos \theta_{\infty}^* = -\sin \frac{\delta}{2}$$

$$\sin \frac{\delta}{2} = \frac{1}{e}$$

$$\varepsilon = \frac{+\mu}{2|a|} = \frac{v^2}{2} - \frac{\mu}{r}$$

$$v^2 = \frac{2\mu}{r} + \underbrace{\frac{\mu}{|a|}}_{\text{positive}} v_{\infty}^2$$



$$v > \sqrt{2} v_{\infty}$$

always