ASTR HW 7

PROBLEM 1

```
clear all
close all
clc
afile = readmatrix("ssp_datav3.csv");
% Assigning the variables
r_a = afile(:,1); % Aphelion [km]
r_p = afile(:,2); % Perihelion [km]
M_planets = afile(:,4); % Mass of planets [kg]
v_a = afile(:,7); % Velocity at aphelion [km/s]
v_p = afile(:, 8); % Velocity at perhelion [km/s]
v_esc = afile(:, 10); % Escape velocity [km/s]
M_sun = [1.989*10^30 1.989*10^30 1.989*10^30 1.989*10^30 M_planets(1)]; % Mass of Sun [kg]
m_i = [10 100 1000].^3*4*pi/3*3.3/1000*10^6; % Mass of impactor [kg]
G = 6.67*10^(-20); % Graviational constant [km^3/kg/s^2]
```

Condition 1 : At Aphelion

```
% Earth
[v_E1_1, prob_E1_1] = cal\_object\_vel(M_planets(1), m_i(1), v_a(1), v_esc(1));
                                                                                 % 10m impactor
[v_E1_2, prob_E1_2] = cal_object_vel(M_planets(1), m_i(2), v_a(1), v_esc(1));
                                                                                 % 100m impactor
[v E1 3, prob E1 3] = cal_object_vel(M_planets(1), m_i(3), v_a(1), v_esc(1));
                                                                                 % 1000m impactor
% Mars
[v M1 1, prob M1 1] = cal object vel(M planets(2), m i(1), v a(2), v esc(2));
                                                                                 % 10m impactor
[v M1 2, prob_M1_2] = cal_object_vel(M_planets(2), m_i(2),v_a(2),v_esc(2));
                                                                                 % 100m impactor
[v M1 3, prob M1 3] = cal_object_vel(M_planets(2), m_i(3), v_a(2), v_esc(2));
                                                                                 % 1000m impactor
% Jupiter
[v_{J1}, prob_{J1}] = cal_object_vel(M_planets(3), m_i(1), v_a(3), v_esc(3));
                                                                                % 10m impactor
[v_{J1}_{2}, prob_{J1}_{2}] = cal_object_vel(M_planets(3), m_i(2), v_a(3), v_esc(3));
                                                                                 % 100m impactor
[v_J1_3, prob_J1_3] = cal_object_vel(M_planets(3), m_i(3), v_a(3), v_esc(3));
                                                                                 % 1000m impactor
% Saturn
[v S1 1, prob S1 1] = cal object vel(M planets(4), m i(1), v a(4), v esc(4));
                                                                                 % 10m impactor
[v_S1_2, prob_S1_2] = cal_object_vel(M_planets(4), m_i(2), v_a(4), v_esc(4));
                                                                                 % 100m impactor
[v_{S1_3}, prob_{S1_3}] = cal_object_vel(M_planets(4), m_i(3), v_a(4), v_esc(4));
                                                                                % 1000m impactor
% Moon
[v_Mo1_1, prob_Mo1_1] = cal_object_vel(M_planets(5), m_i(1), v_a(5), v_esc(5));
                                                                                   % 10m impactor
[v\_Mo1\_2, prob\_Mo1\_2] = cal\_object\_vel(M\_planets(5), m\_i(2), v\_a(5), v\_esc(5));
                                                                                  % 100m impactor
[v Mo1_3, prob_Mo1_3] = cal_object_vel(M_planets(5), m_i(3),v_a(5),v_esc(5));
                                                                                   % 1000m impactor
```

Condition 2: At Perihelion

```
% Earth
[v_E2_1, prob_E2_1] = cal_object_vel(M_planets(1), m_i(1),v_p(1),v_esc(1));  % 10m impactor
[v_E2_2, prob_E2_2] = cal_object_vel(M_planets(1), m_i(2),v_p(1),v_esc(1));  % 100m impactor
[v_E2_3, prob_E2_3] = cal_object_vel(M_planets(1), m_i(3),v_p(1),v_esc(1));  % 1000m impactor
% Mars
[v_M2_1, prob_M2_1] = cal_object_vel(M_planets(2), m_i(1),v_p(2),v_esc(2));  % 100m impactor
[v_M2_2, prob_M2_2] = cal_object_vel(M_planets(2), m_i(2),v_p(2),v_esc(2));  % 100m impactor
```

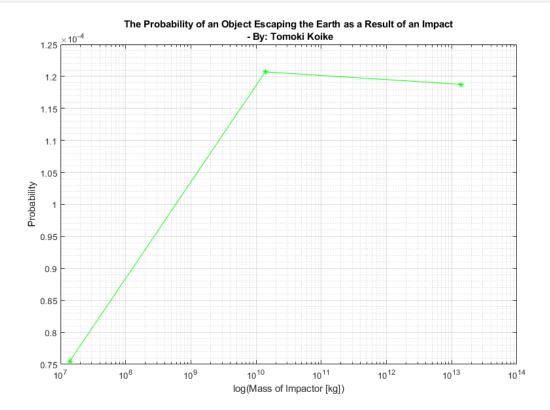
```
[v M2 3, prob_M2_3] = cal_object_vel(M_planets(2), m_i(3), v_p(2), v_esc(2)); % 1000m impactor
% Jupiter
[v_J2_1, prob_J2_1] = cal_object_vel(M_planets(3), m_i(1), v_p(3), v_esc(3));
                                                                               % 10m impactor
[v J2 2, prob_J2_2] = cal_object_vel(M_planets(3), m_i(2),v_p(3),v_esc(3));
                                                                               % 100m impactor
[v_{J2_3}, prob_{J2_3}] = cal\_object\_vel(M_planets(3), m_i(3), v_p(3), v_esc(3));
                                                                               % 1000m impactor
% Saturn
[v S2 1, prob S2 1] = cal object vel(M planets(4), m i(1), v p(4), v esc(4));
                                                                               % 10m impactor
[v_S2_2, prob_S2_2] = cal\_object\_vel(M_planets(4), m_i(2), v_p(4), v_esc(4));
                                                                              % 100m impactor
[v_S2_3, prob_S2_3] = cal_object_vel(M_planets(4), m_i(3),v_p(4),v_esc(4));
                                                                              % 1000m impactor
% Moon
[v_Mo2_1, prob_Mo2_1] = cal_object_vel(M_planets(5), m_i(1), v_p(5), v_esc(5));
                                                                                 % 10m impactor
[v_Mo2_2, prob_Mo2_2] = cal_object_vel(M_planets(5), m_i(2), v_p(5), v_esc(5)); % 100m impactor
[v_Mo2_3, prob_Mo2_3] = cal_object_vel(M_planets(5), m_i(3), v_p(5), v_esc(5)); % 1000m impactor
```

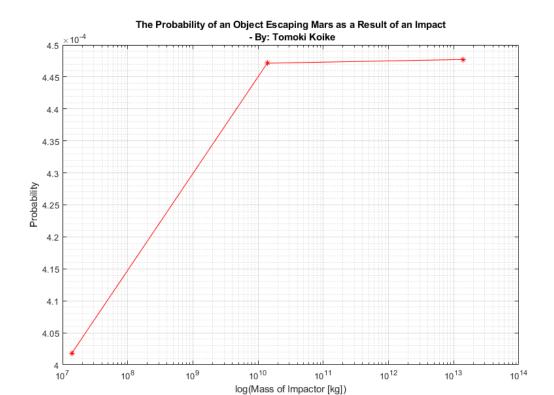
Total Probability: (both aphelion and perhelion together)

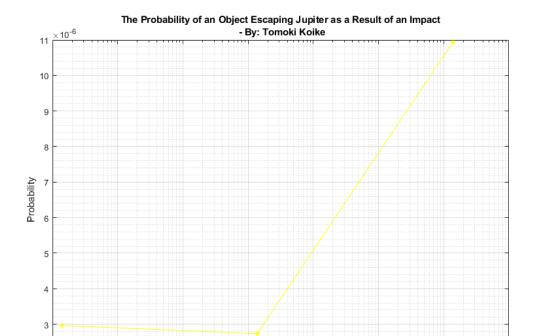
```
% Earth
prob_E_tot1 = prob_E1_1*prob_E2_1; % Impactor 10m
prob_E_tot2 = prob_E1_2*prob_E2_2;
                                   % Impactor 100m
prob E tot3 = prob E1 3*prob E2 3;
                                   % Impactor 1000m
prob_E = [prob_E_tot1, prob_E_tot2, prob_E_tot3];
% Mars
prob M tot1 = prob M1 1*prob M2 1;
                                   % Impactor 10m
prob_M_tot2 = prob_M1_2*prob_M2_2;
                                   % Impactor 100m
prob M tot3 = prob M1 3*prob M2 3;
                                   % Impactor 1000m
prob_M = [prob_M_tot1, prob_M_tot2, prob_M_tot3];
% Jupiter
prob J tot1 = prob J1 1*prob J2 1;
                                   % Impactor 10m
prob_J_tot2 = prob_J1_2*prob_J2_2;
                                   % Impactor 100m
prob_J_tot3 = prob_J1_3*prob_J2_3;
                                   % Impactor 1000m
prob_J = [prob_J_tot1, prob_J_tot2, prob_J_tot3];
% Saturn
                                   % Impactor 10m
prob_S_tot1 = prob_S1_1*prob_S2_1;
prob S tot2 = prob S1 2*prob S2 2;
                                   % Impactor 100m
prob S tot3 = prob S1 3*prob S2 3; % Impactor 1000m
prob_s = [prob_S_tot1, prob_S_tot2, prob_S_tot3];
% Moon
prob_Mo_tot1 = prob_Mo1_1*prob_Mo2_1;
                                      % Impactor 10m
prob Mo tot2 = prob Mo1 2*prob Mo2 2; % Impactor 100m
prob Mo tot3 = prob Mo1 3*prob Mo2 3; % Impactor 1000m
prob_Mo = [prob_Mo_tot1, prob_Mo_tot2, prob_Mo_tot3];
```

Plotting

```
% Earth
figure('Renderer', 'painters', 'Position', [10 10 900 600]);
semilogx(m_i, prob_E, '*-g')
title({['The Probability of an Object Escaping the Earth as a Result of an ' ...
    'Impact'],' - By: Tomoki Koike'})
xlabel('log(Mass of Impactor [kg])')
ylabel('Probability')
grid on
grid minor
```







10¹⁰

log(Mass of Impactor [kg])

10⁷

10⁸

10⁹

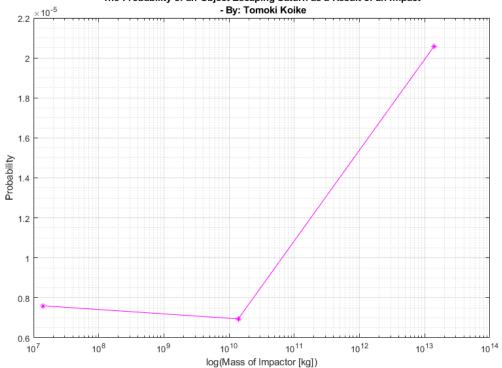
10¹¹

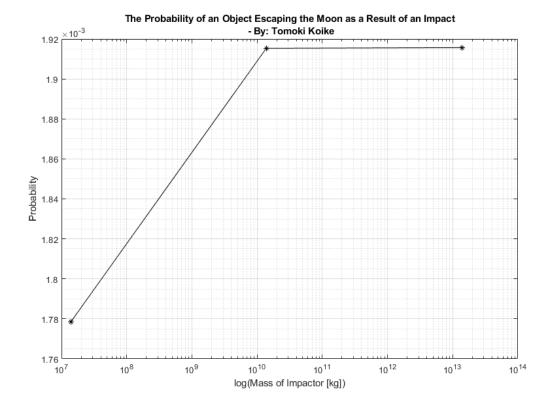
10¹³

10¹²

10¹⁴



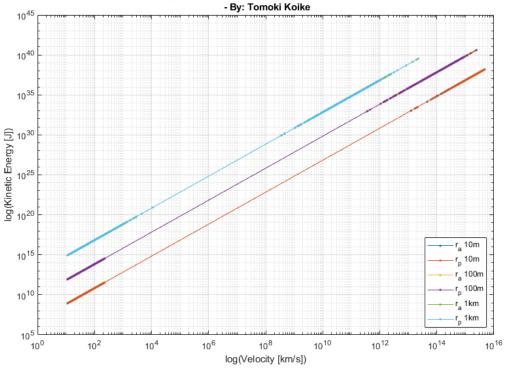




Kinetic Energy: Plotting

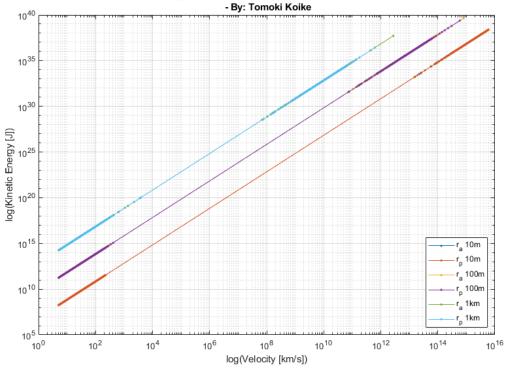
```
% Earth
figure('Renderer', 'painters', 'Position', [10 10 900 600]);
loglog(sort(v_E1_1), cal_KE(sort(v_E1_1), m_i(1)), '.-')
title({['Kinetic Energy of Object Escaping the Earth as a Result of an ' ...
    'Impact'],' - By: Tomoki Koike'})
xlabel('log(Velocity [km/s])')
ylabel('log(Kinetic Energy [J])')
hold on
loglog(sort(v_E2_1), cal_KE(sort(v_E2_1), m_i(1)), '.-')
loglog(sort(v_E1_2), cal_KE(sort(v_E1_2), m_i(2)), '.-')
loglog(sort(v_E2_2), cal_KE(sort(v_E2_2), m_i(2)),
loglog(sort(v_E1_3), cal_KE(sort(v_E1_3), m_i(3)), '.-')
loglog(sort(v_E2_3), cal_KE(sort(v_E2_3), m_i(3)),
hold off
legend('r_a 10m', 'r_p 10m', 'r_a 100m', 'r_p 100m', 'r_a 1km', 'r_p 1km', ...
    'Location', 'southeast')
grid on
grid minor
box on
```

Kinetic Energy of Object Escaping the Earth as a Result of an Impact



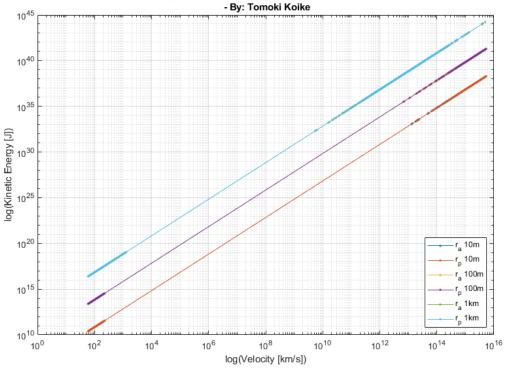
```
% Mars
figure('Renderer', 'painters', 'Position', [10 10 900 600]);
loglog(sort(v_M1_1), cal_KE(sort(v_M1_1), m_i(1)), '.-')
title({['Kinetic Energy of Object Escaping Mars as a Result of an ' ...
    'Impact'],' - By: Tomoki Koike'})
xlabel('log(Velocity [km/s])')
ylabel('log(Kinetic Energy [J])')
hold on
loglog(sort(v_M2_1), cal_KE(sort(v_M2_1), m_i(1)), '.-')
loglog(sort(v_M1_2), cal_KE(sort(v_M1_2), m_i(2)),
loglog(sort(v_M2_2), cal_KE(sort(v_M2_2), m_i(2)),
loglog(sort(v_M1_3), cal_KE(sort(v_M1_3), m_i(3)),
loglog(sort(v_M2_3), cal_KE(sort(v_M2_3), m_i(3)),
hold off
legend('r_a 10m', 'r_p 10m', 'r_a 100m', 'r_p 100m', 'r_a 1km', 'r_p 1km', ...
    'Location', 'southeast')
grid on
grid minor
box on
```

Kinetic Energy of Object Escaping Mars as a Result of an Impact



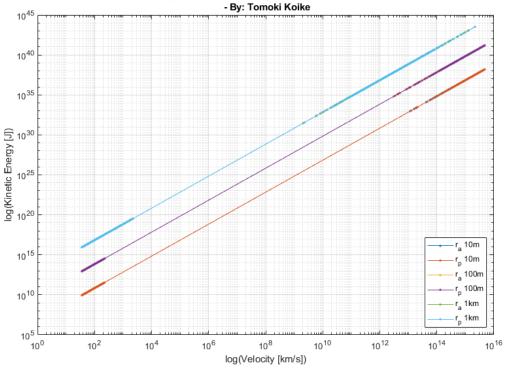
```
% Jupiter
figure('Renderer', 'painters', 'Position', [10 10 900 600]);
loglog(sort(v_J1_1), cal_KE(sort(v_J1_1), m_i(1)), '.-')
title({['Kinetic Energy of Object Escaping Jupiter as a Result of an ' ...
    'Impact'],' - By: Tomoki Koike'})
xlabel('log(Velocity [km/s])')
ylabel('log(Kinetic Energy [J])')
hold on
loglog(sort(v_J2_1), cal_KE(sort(v_J2_1), m_i(1)), '.-')
loglog(sort(v_J1_2), cal_KE(sort(v_J1_2), m_i(2)),
loglog(sort(v_J2_2), cal_KE(sort(v_J2_2), m_i(2)),
loglog(sort(v_J1_3), cal_KE(sort(v_J1_3), m_i(3)),
loglog(sort(v_J2_3), cal_KE(sort(v_J2_3), m_i(3)),
hold off
legend('r_a 10m', 'r_p 10m', 'r_a 100m', 'r_p 100m', 'r_a 1km', 'r_p 1km', ...
    'Location', 'southeast')
grid on
grid minor
box on
```

Kinetic Energy of Object Escaping Jupiter as a Result of an Impact



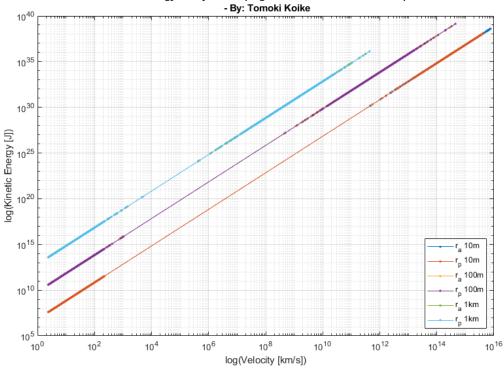
```
% Saturn
figure('Renderer', 'painters', 'Position', [10 10 900 600]);
loglog(sort(v_S1_1), cal_KE(sort(v_S1_1), m_i(1)), '.-')
title({['Kinetic Energy of Object Escaping Saturn as a Result of an ' ...
    'Impact'],' - By: Tomoki Koike'})
xlabel('log(Velocity [km/s])')
ylabel('log(Kinetic Energy [J])')
hold on
loglog(sort(v_S2_1), cal_KE(sort(v_S2_1), m_i(1)), '.-')
loglog(sort(v_S1_2), cal_KE(sort(v_S1_2), m_i(2)),
loglog(sort(v_S2_2), cal_KE(sort(v_S2_2), m_i(2)),
loglog(sort(v_S1_3), cal_KE(sort(v_S1_3), m_i(3)),
loglog(sort(v_S2_3), cal_KE(sort(v_S2_3), m_i(3)),
hold off
legend('r_a 10m', 'r_p 10m', 'r_a 100m', 'r_p 100m', 'r_a 1km', 'r_p 1km', ...
    'Location', 'southeast')
grid on
grid minor
box on
```

Kinetic Energy of Object Escaping Saturn as a Result of an Impact



```
% Moon
figure('Renderer', 'painters', 'Position', [10 10 900 600]);
loglog(sort(v_Mo1_1), cal_KE(sort(v_Mo1_1), m_i(1)), '.-')
title({['Kinetic Energy of Object Escaping the Moon as a Result of an ' ...
    'Impact'],' - By: Tomoki Koike'})
xlabel('log(Velocity [km/s])')
ylabel('log(Kinetic Energy [J])')
hold on
loglog(sort(v Mo2 1), cal KE(sort(v Mo2 1), m i(1)), '.-')
loglog(sort(v_Mo1_2), cal_KE(sort(v_Mo1_2), m_i(2)),
loglog(sort(v_Mo2_2), cal_KE(sort(v_Mo2_2), m_i(2)),
loglog(sort(v_Mo1_3), cal_KE(sort(v_Mo1_3), m_i(3)),
loglog(sort(v_Mo2_3), cal_KE(sort(v_Mo2_3), m_i(3)),
hold off
legend('r_a 10m', 'r_p 10m', 'r_a 100m', 'r_p 100m', 'r_a 1km', 'r_p 1km', ...
    'Location', 'southeast')
grid on
grid minor
box on
```

Kinetic Energy of Object Escaping the Moon as a Result of an Impact



```
function m_o = cal_m_o(M, m_i, V, v_esc, beta, gamma, delta)
    m o = ((M+m i)*V*(cos(delta)+sin(delta)*tan(gamma)) - ...
        m_i*v_esc*(cos(beta)+sin(beta)*tan(gamma)) - M*V) / ...
        (V*(cos(delta)+sin(delta)*tan(gamma)));
end
function [v_suf, prob] = cal_object_vel(M, m_i, V, v_e)
    v = zeros([181, 181, 21]);
    for beta = 0:180
        for gamma = 0:180
            for delta = 0:20
                m_o = cal_m_o(M, m_i, V, v_e, beta, gamma, delta);
                temp = ((M-m o+m i)*sin(delta) - m i*v e*sin(beta))/m o*cos(beta);
                if temp ~= inf
                    v(beta+1, gamma+1, delta+1) = temp;
                end
            end
        end
    end
    v_{temp} = v(\sim(isnan(v)|isinf(v)));
    v_num = length(v_temp);
    v_suf = v_temp(v_temp>=v_e);
    v_suf = sort(v_suf);
    v_suf_num = length(v_suf);
    prob = v_suf_num / v_num;
end
function KE = cal_KE(v, m)
```

 $KE = 0.5*m*v.^2;$ end