

Ray Tracing: shading

CS 4620 Lecture 6

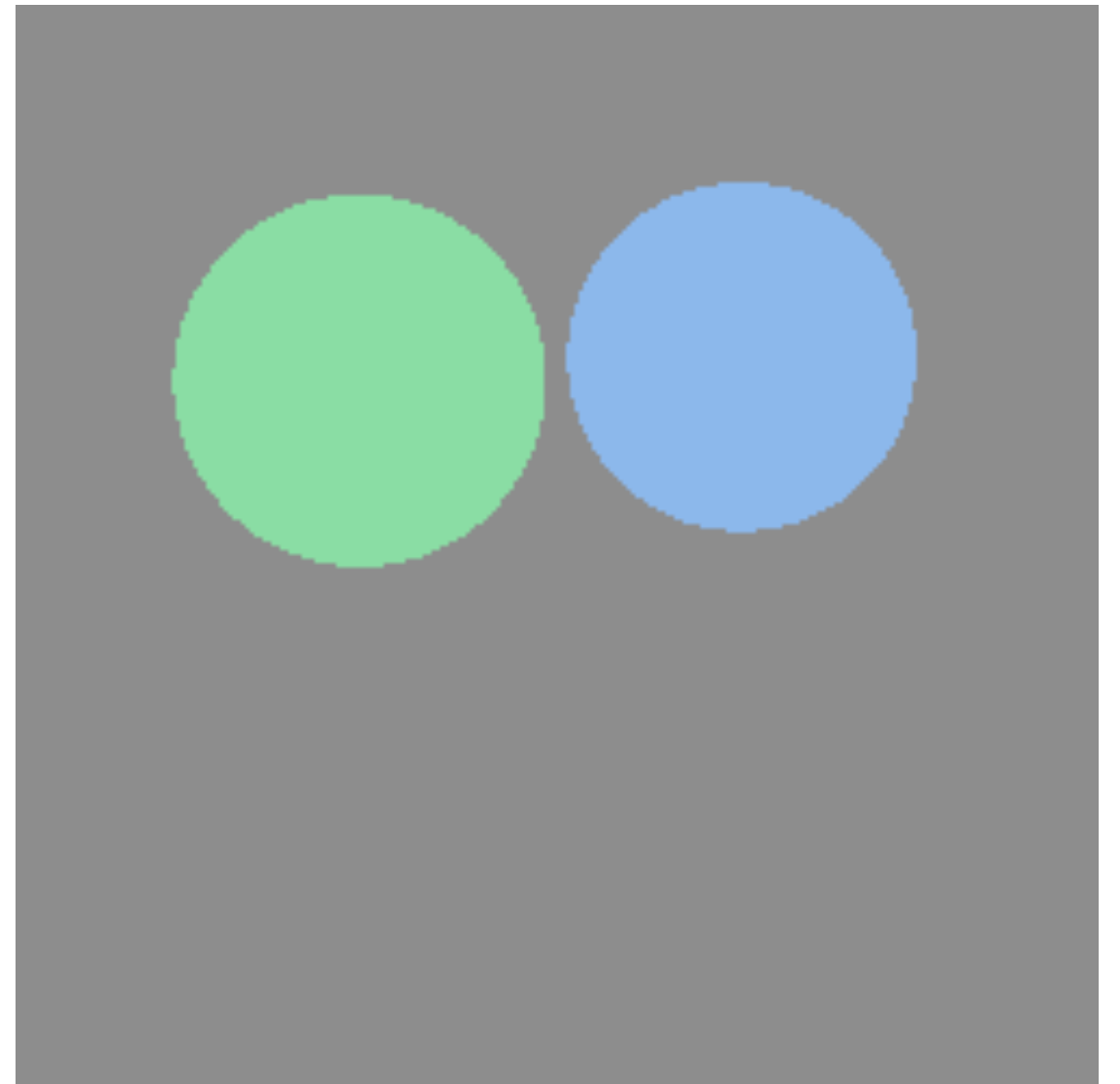
Image so far

- **With eye ray generation and scene intersection**

```
for 0 <= iy < ny
  for 0 <= ix < nx {
    ray = camera.getRay(ix, iy);
    c = scene.trace(ray, 0, +inf);
    image.set(ix, iy, c);
  }

...

Scene.trace(ray, tMin, tMax) {
  surface, t = surfs.intersect(ray, tMin, tMax);
  if (surface != null) return surface.color();
  else return black;
}
```

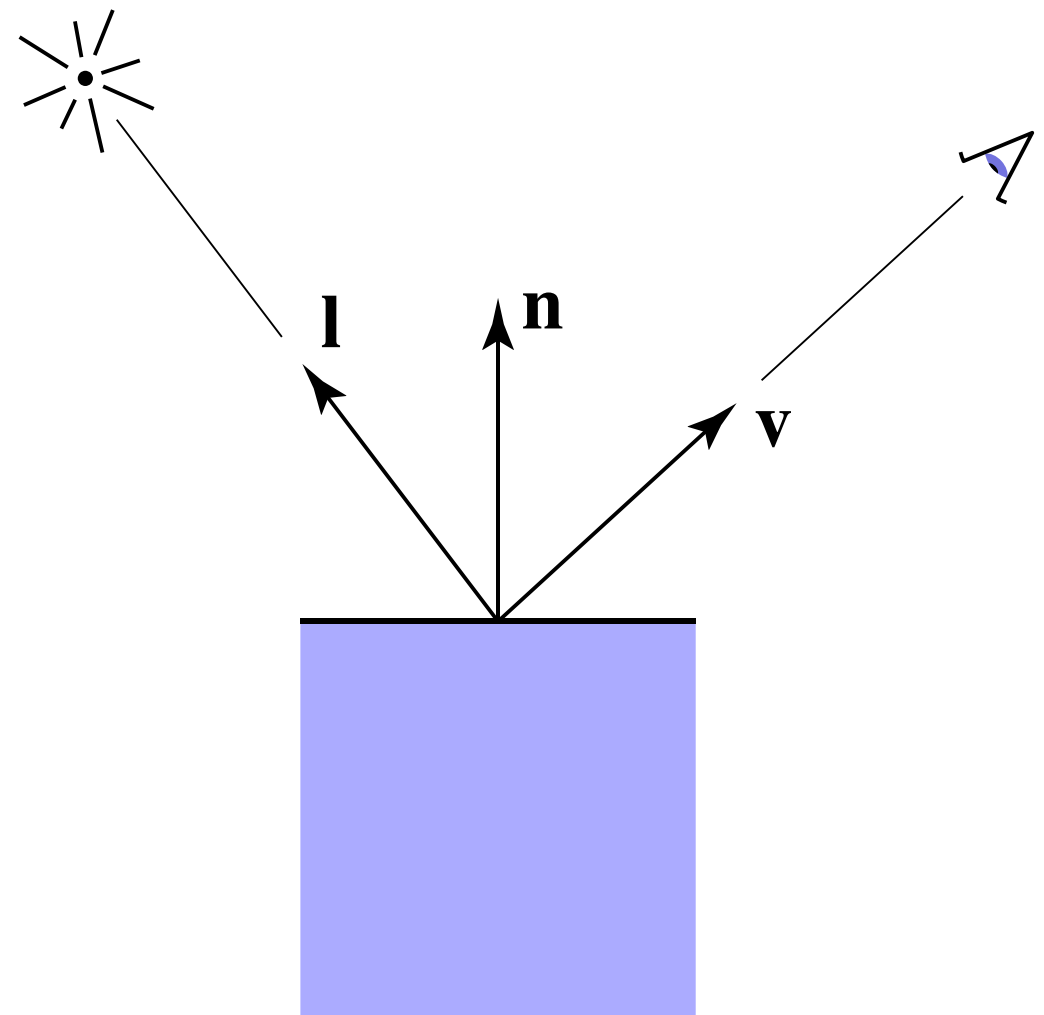


Shading

- **Compute light reflected toward camera**

- **Inputs:**

- eye direction
- light direction
(for each of many lights)
- surface normal
- surface parameters
(color, roughness, ...)



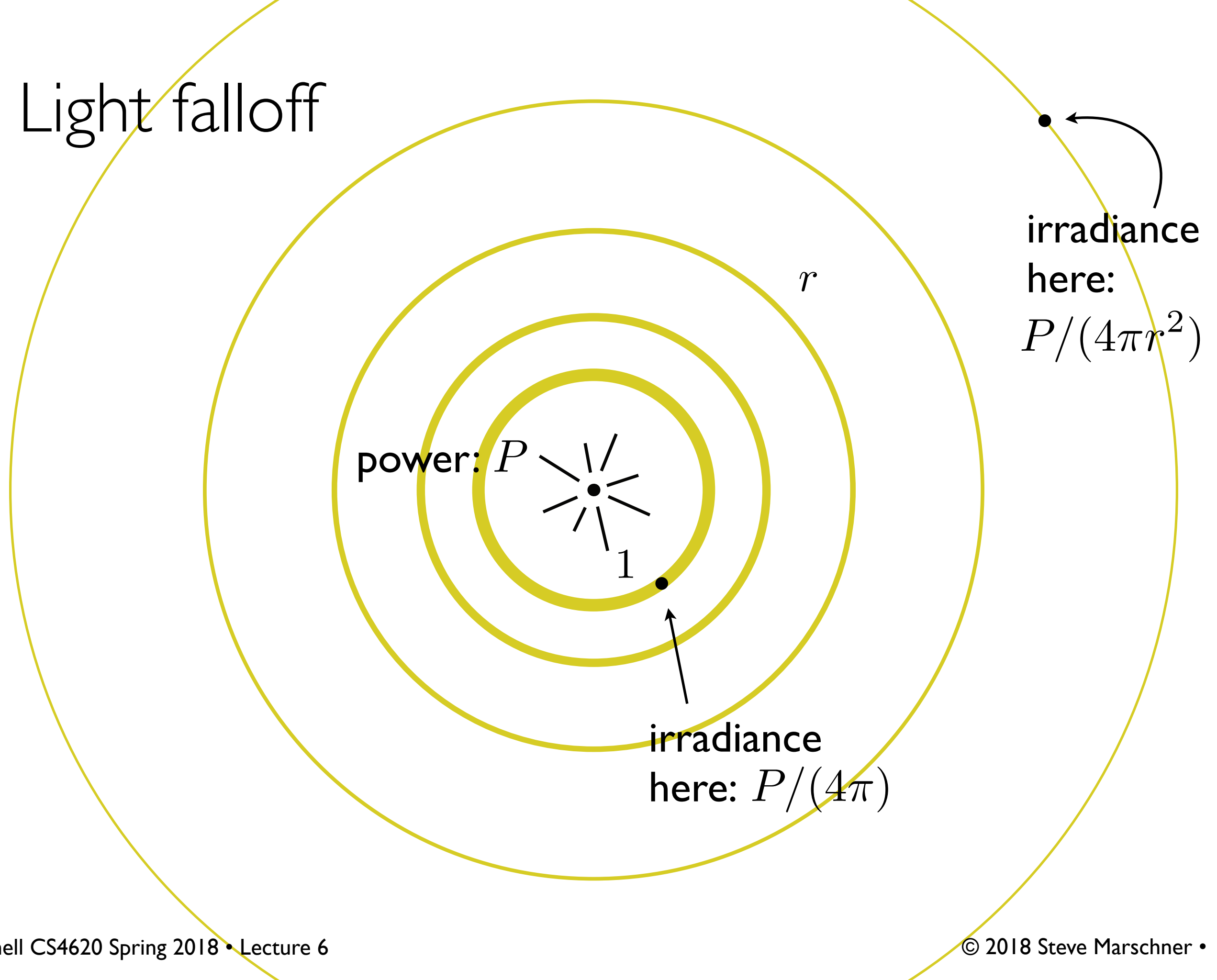
Shading philosophy

- **Goals of shading depend on purpose of image**
 - visualization, CAD: maximize visual clarity
 - visual effects, advertising: maximize resemblance to reality
 - animation, games: somewhere in between
- **Basic starting point: physics of light reflection**
 - a set of useful approximations to real surfaces
 - can remove things for simplicity/clarity
 - can add things for increased accuracy/realism

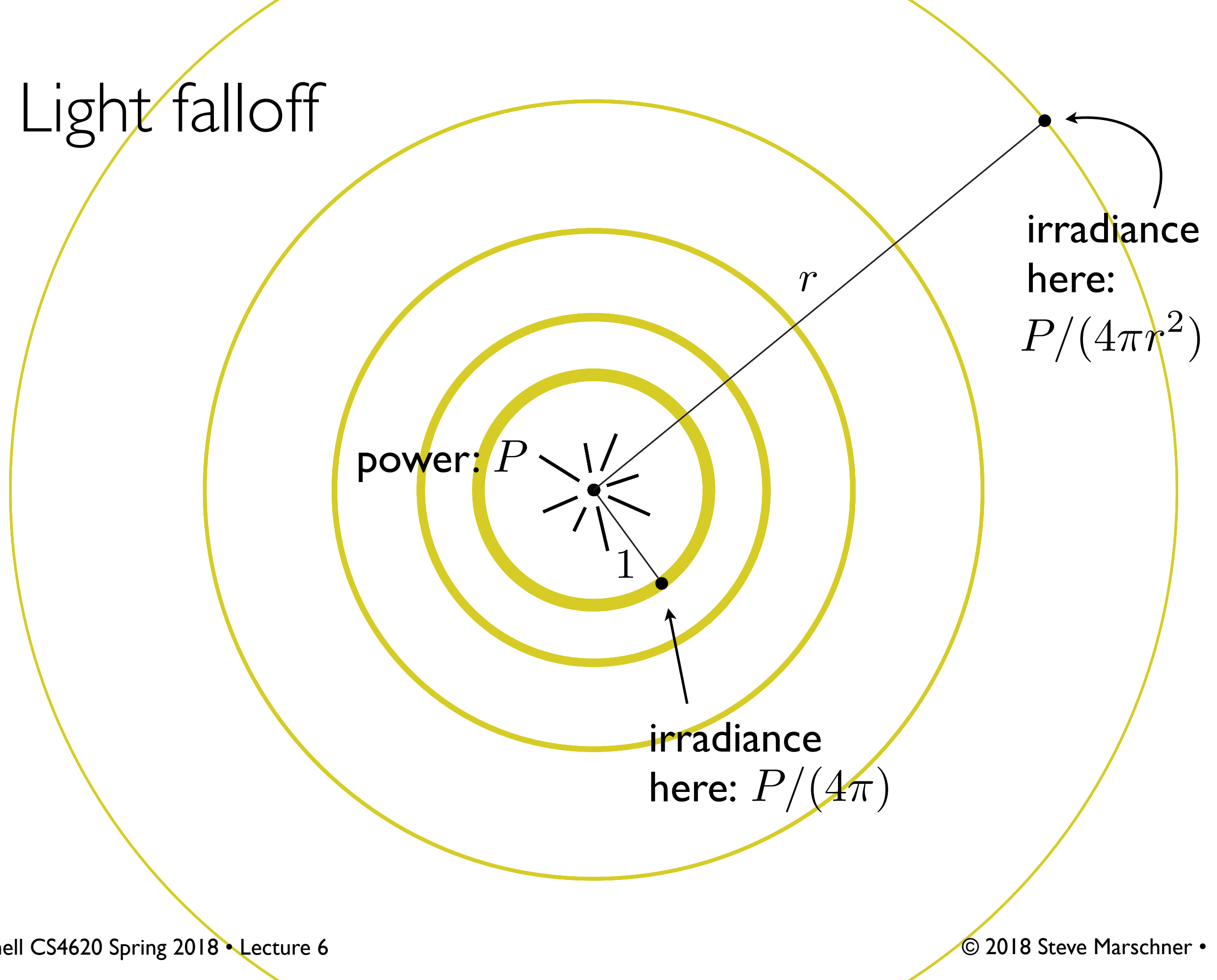
Light

- **Think of light as a flow of particles through space**
 - disregarding wave nature: polarization, interference, diffraction
 - for now disregarding color: only how much light
- **Sources of light**
 - point sources (a flashlight) ← we will stick to this for now.
 - directional sources (the sun)
 - area sources (a fluorescent tube)
 - environment sources (the sky)

Light falloff



Light falloff



Irradiance from isotropic point source

- **A sphere surrounding the source receives all the power**
- **A small, flat surface of area A facing the source receives a fraction (area of surface) / (area of sphere) of that power:**

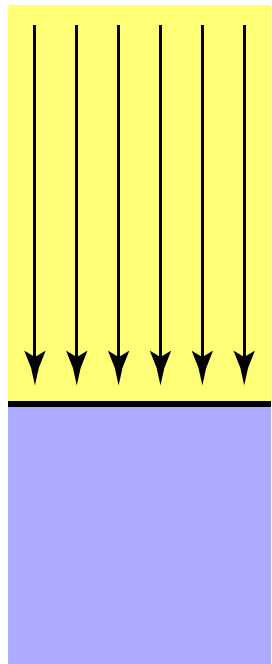
$$P_A = P \frac{A}{4\pi r^2}$$

- **Irradiance is power per unit area:**

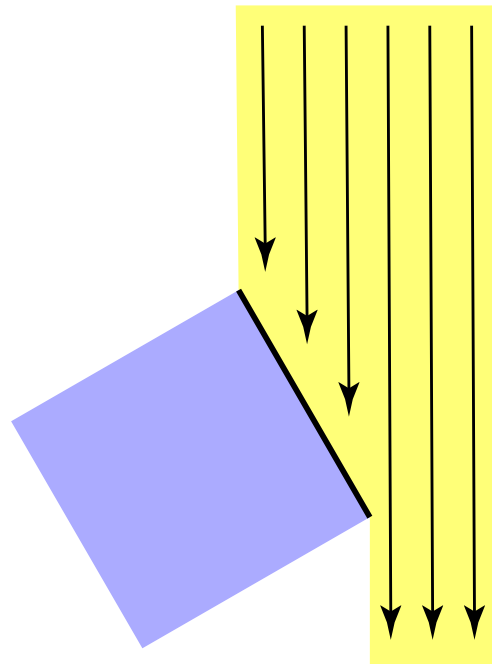
$$E = P_A / A = \frac{P}{4\pi r^2} = \frac{P}{4\pi} \frac{1}{r^2}$$

↑ ↑
intensity geometry factor

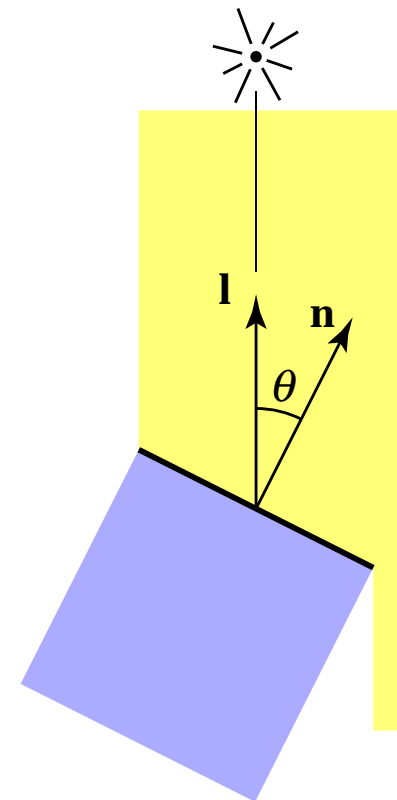
Lambert's cosine law



Top face of cube
receives a certain
amount of light



Top face of
 60° rotated cube
intercepts half the light



In general, light per unit
area is proportional to
 $\cos \theta = \mathbf{l} \cdot \mathbf{n}$

Irradiance from isotropic point source

- **A surface of area A facing at an angle to the source receives a factor of $\cos \theta$ less light:**

$$P_A = P \frac{A \cos \theta}{4\pi r^2}$$

- **Irradiance is power per unit area:**

$$E = P_A / A = \frac{P}{4\pi} \frac{\cos \theta}{r^2}$$

↑
intensity

↑
geometry factor

Diffuse reflection

- **Simplest reflection model**
- **Reflected light is independent of view direction**
- **Reflected light is proportional to irradiance**
 - constant of proportionality is the diffuse reflection coefficient

$$L_d = k_d E$$

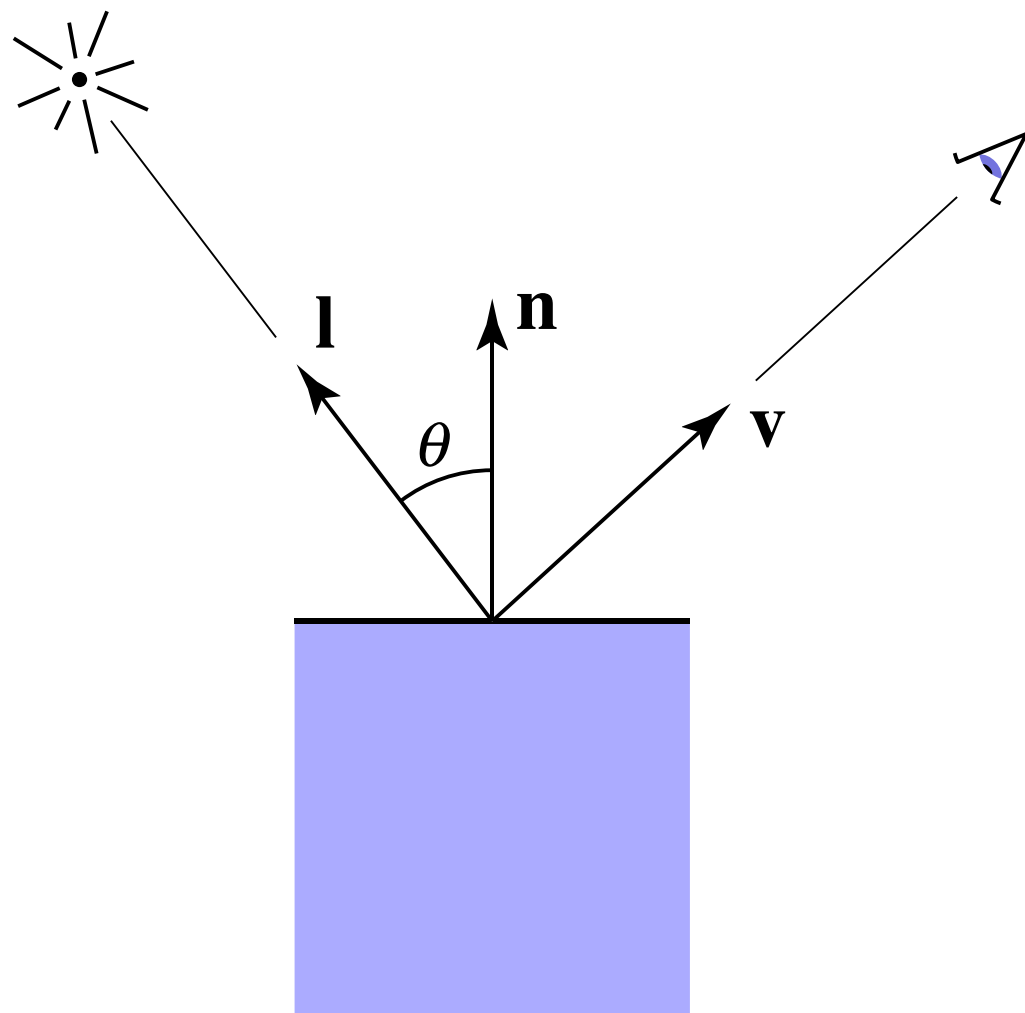
- **More useful to think in terms of reflectance**
 - reflectance is the fraction reflected (between 0 and 1)

$$L_d = \frac{R_d}{\pi} E$$

- will have to explain the factor of pi later

Lambertian shading

- **Shading independent of view direction**



irradiance from source

diffuse reflectance

$$L_d = \frac{R}{\pi} \frac{\max(0, \mathbf{n} \cdot \mathbf{l})}{r^2} I$$

diffusely reflected radiance

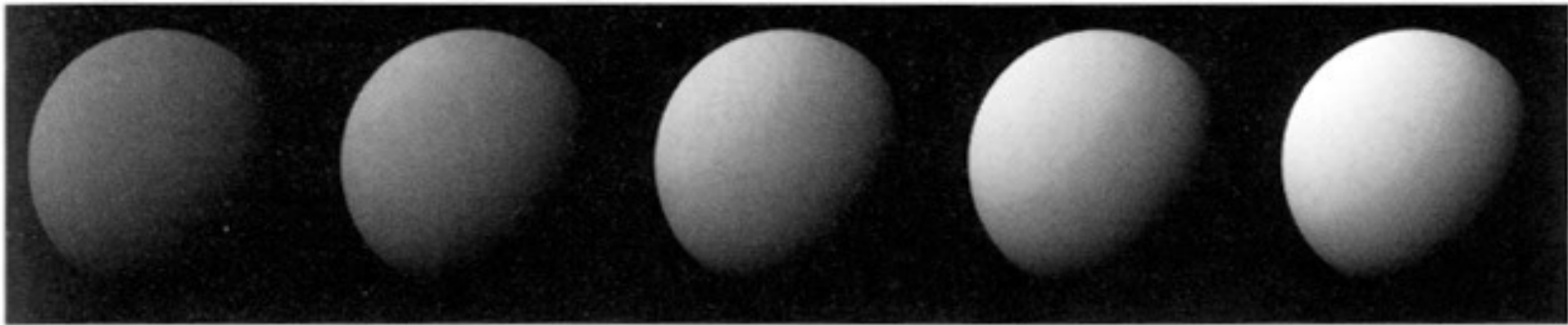
diffuse coefficient

distance to source

intensity of source

Lambertian shading

- **Produces matte appearance**



$k_d \longrightarrow$

[Foley et al.]

Diffuse shading

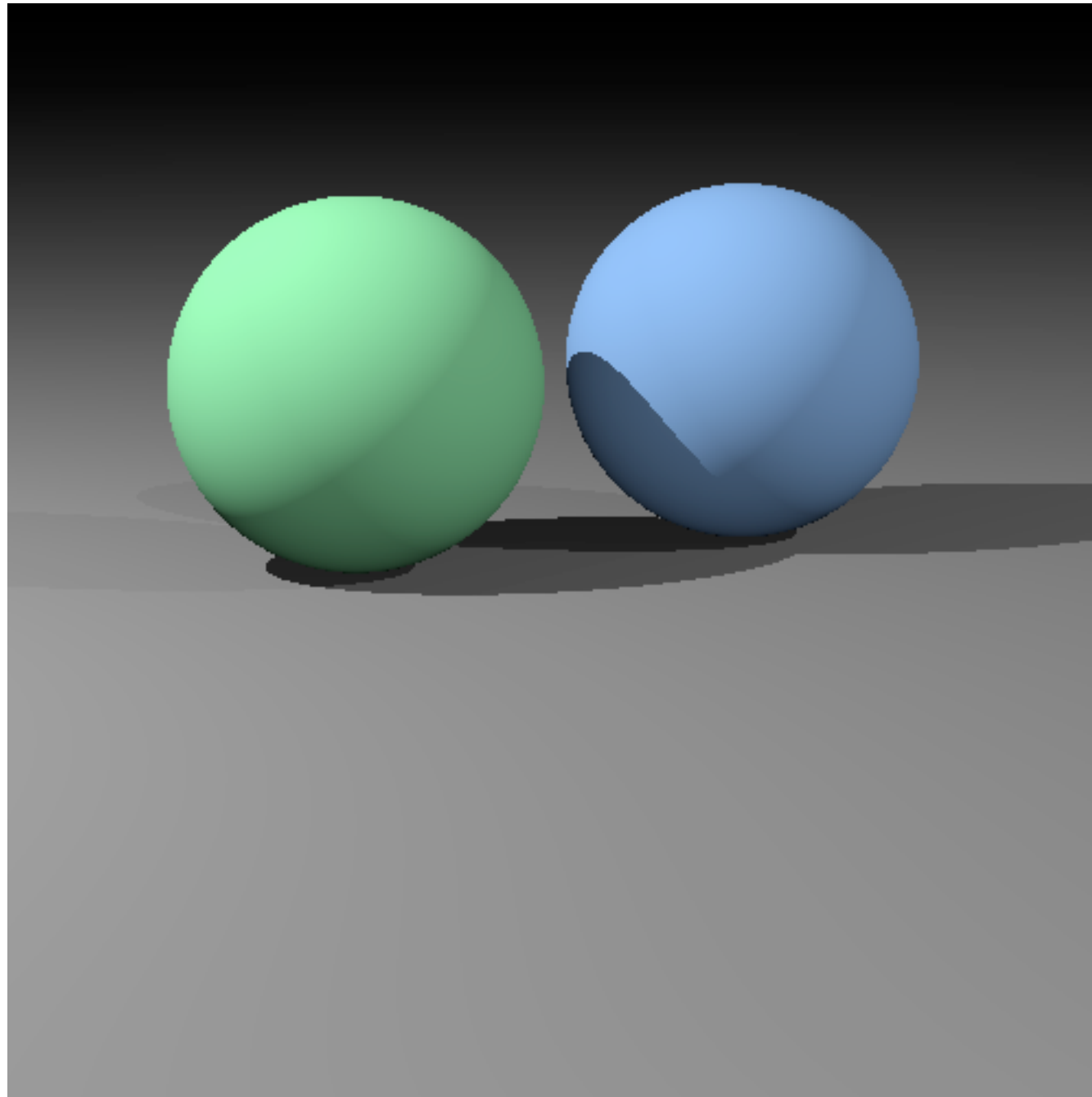
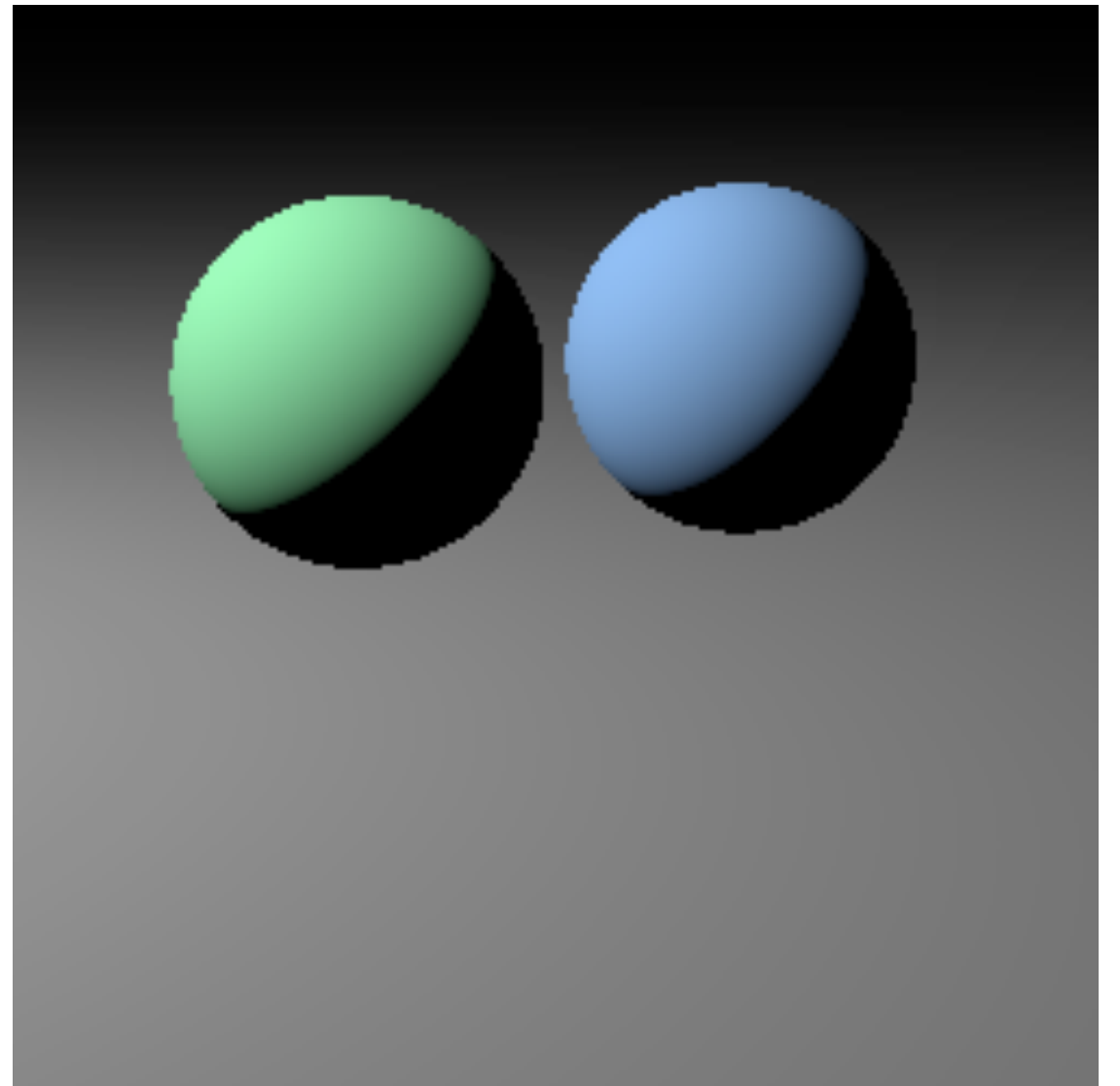


Image so far

```
Scene.trace(Ray ray, tMin, tMax) {  
    surface, t = hit(ray, tMin, tMax);  
    if surface is not null {  
        point = ray.evaluate(t);  
        normal = surface.getNormal(point);  
        return surface.shade(ray, point,  
                               normal, light);  
    }  
    else return backgroundColor;  
}
```

...

```
Surface.shade(ray, point, normal, light) {  
    v = -normalize(ray.direction);  
    l = normalize(light.pos - point);  
    // compute shading  
}
```

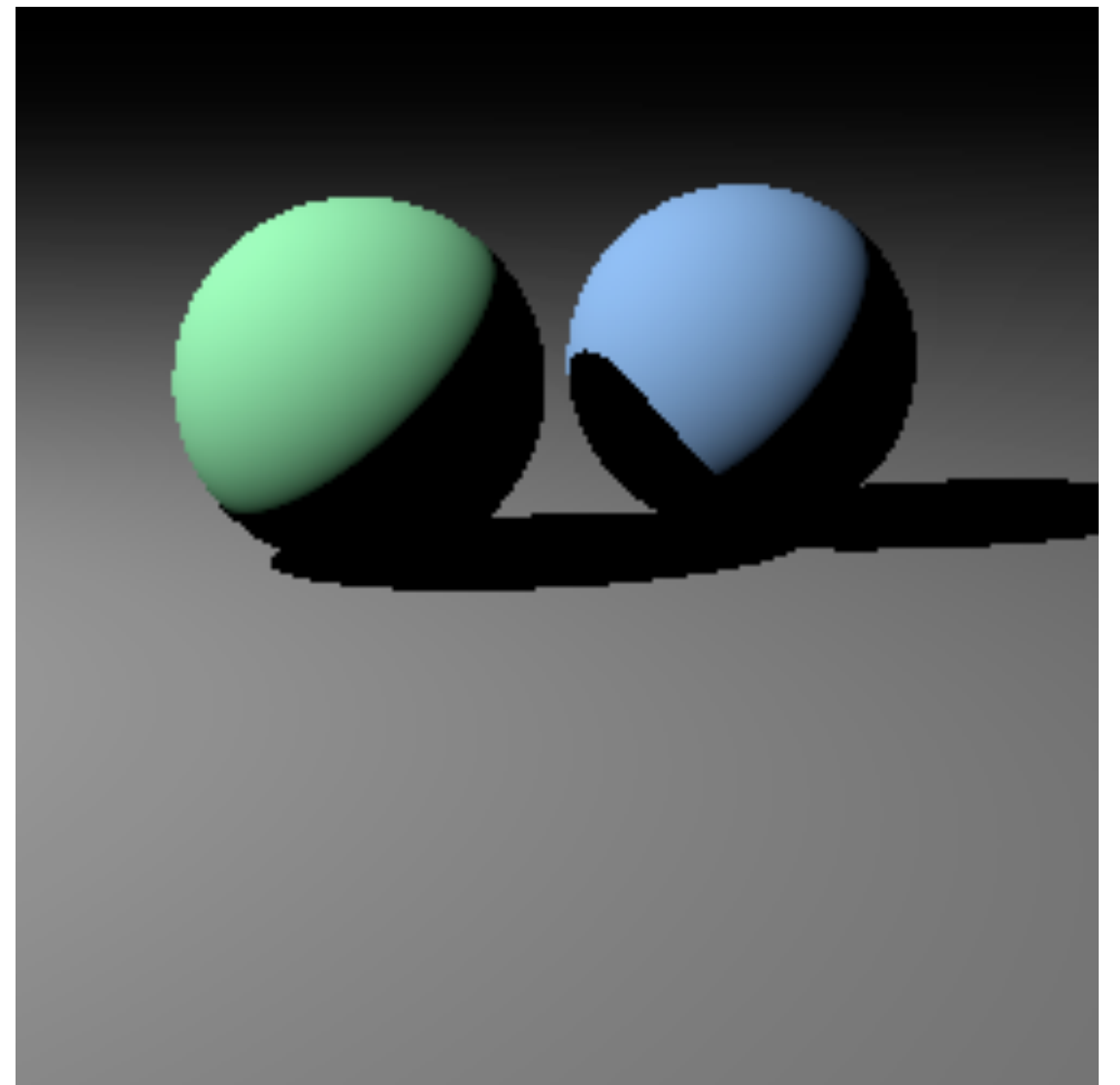


Shadows

- **Surface is only illuminated if nothing blocks the light**
 - i.e. if the surface can “see” the light
- **With ray tracing it’s easy to check**
 - just intersect a ray with the scene!

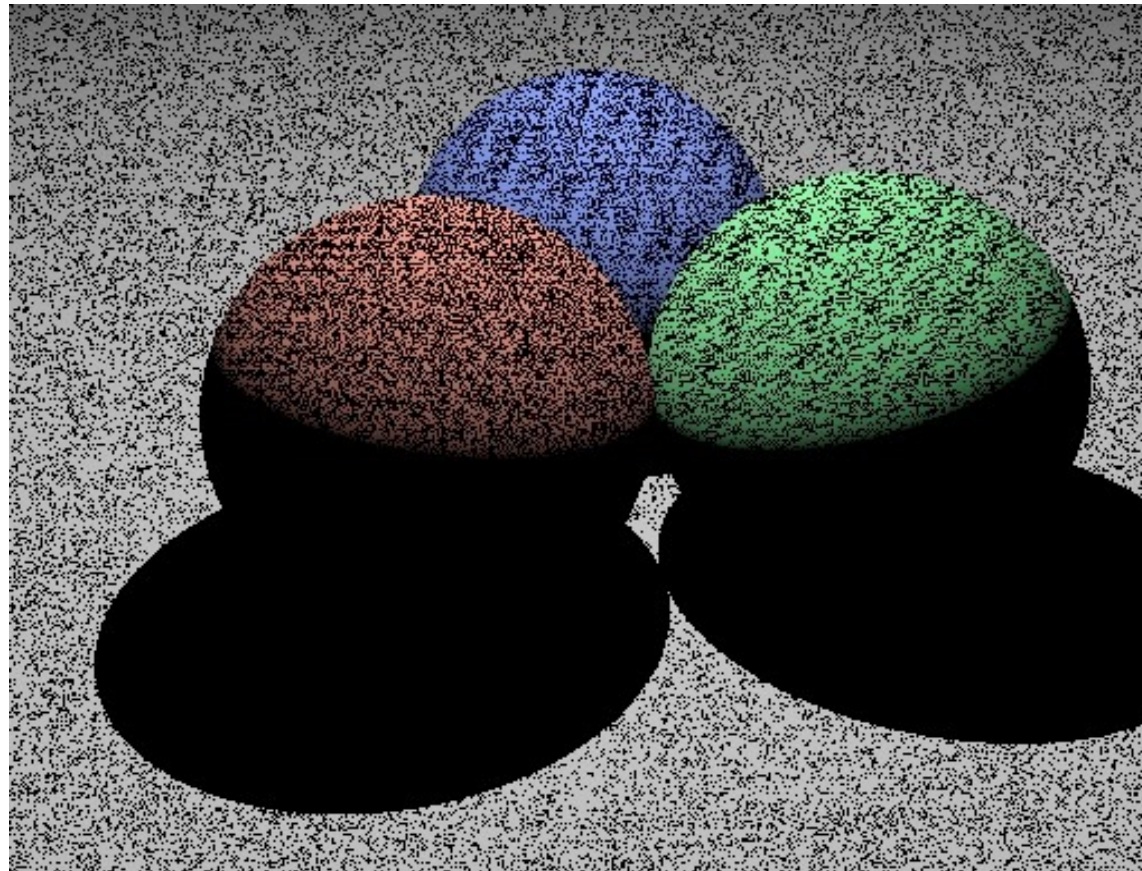
Image so far

```
Surface.shade(ray, point, normal, light) {  
  shadRay = (point, light.pos - point);  
  if (shadRay not blocked) {  
    v = -normalize(ray.direction);  
    l = normalize(light.pos - point);  
    // compute shading  
  }  
  return black;  
}
```



Shadow rounding errors

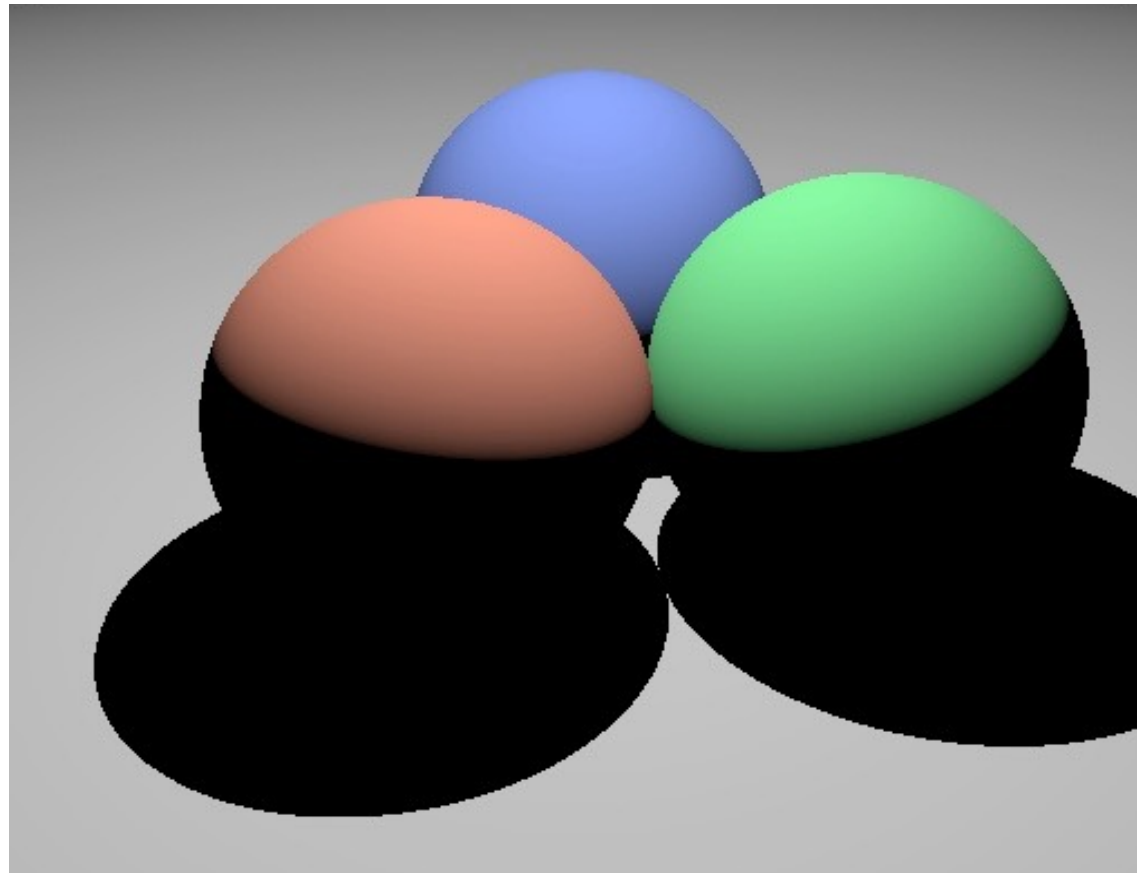
- **Don't fall victim to one of the classic blunders:**



- **What's going on?**
 - hint: at what t does the shadow ray intersect the surface you're shading?

Shadow rounding errors

- **Solution: shadow rays start a tiny distance from the surface**



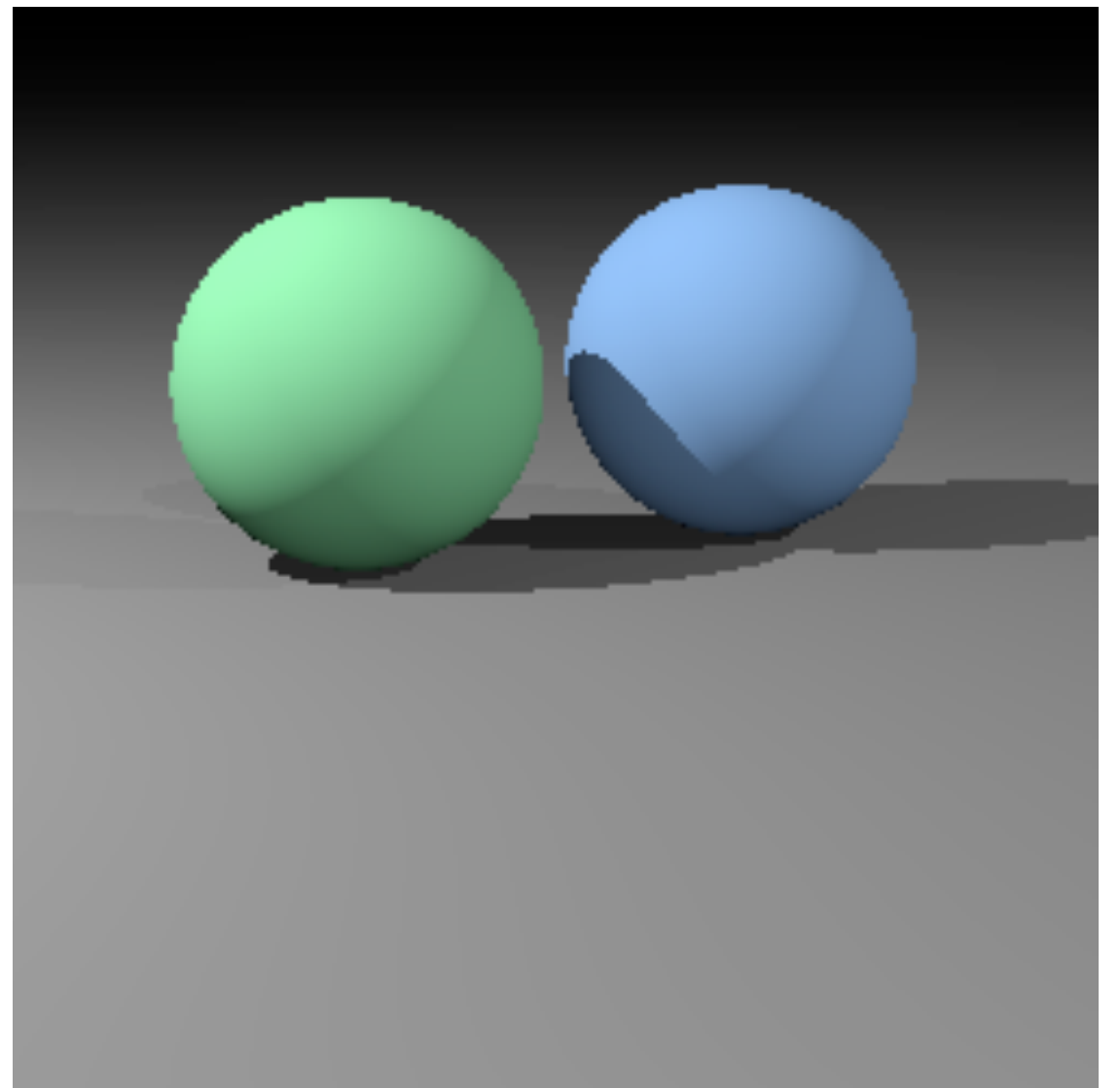
- **Do this by moving the start point, or by limiting the t range**

Multiple lights

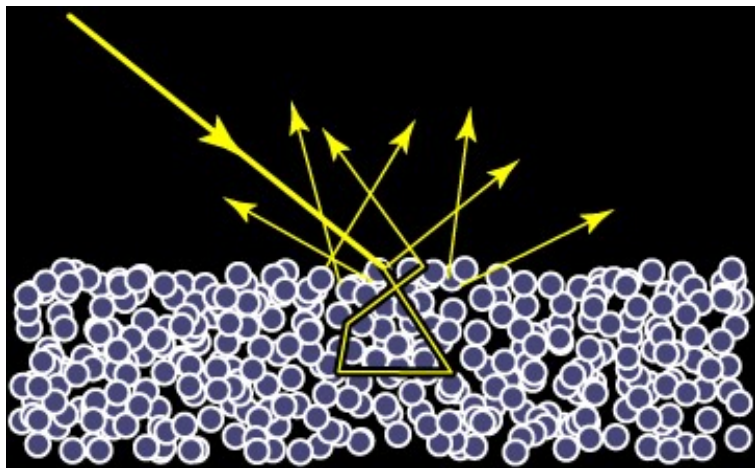
- **Important to fill in black shadows**
- **Just loop over lights, add contributions**
- **Ambient shading**
 - black shadows are not really right
 - one solution: dim light at camera
 - alternative: add a constant “ambient” color to the shading...

Image so far

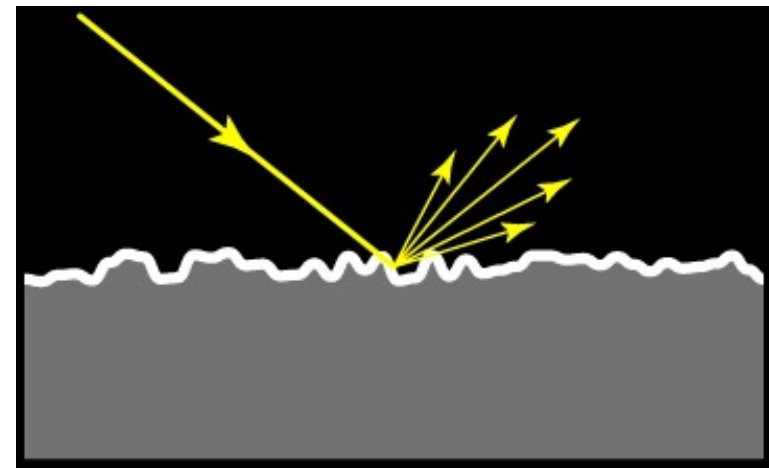
```
shade(ray, point, normal, lights) {  
    result = ambient;  
    for light in lights {  
        if (shadow ray not blocked) {  
            result += shading contribution;  
        }  
    }  
    return result;  
}
```



Specular shading



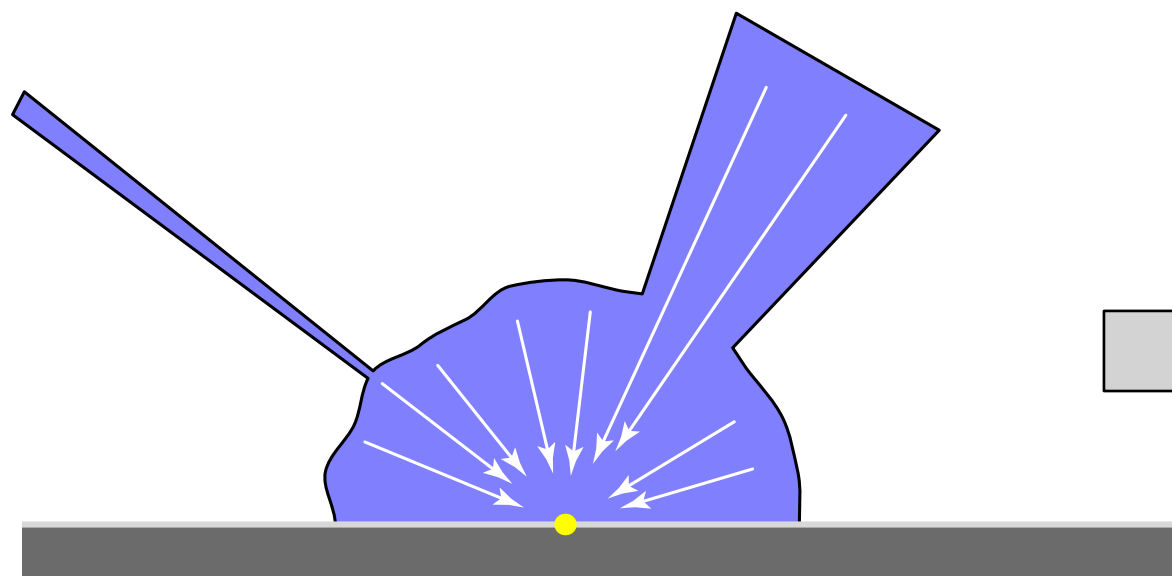
diffuse



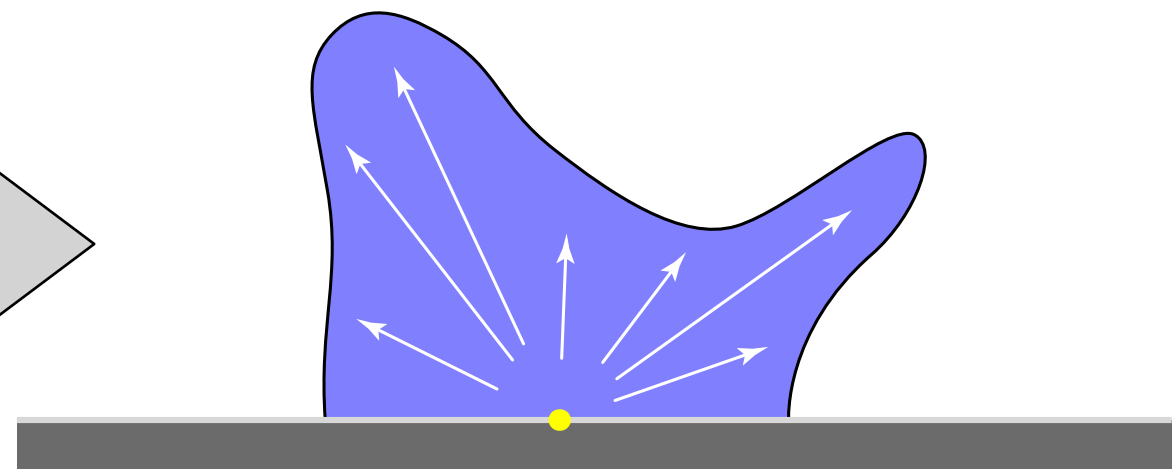
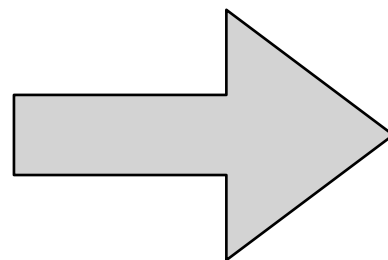
specular

Light reflection: full picture

- **when writing a shader, think like a bug standing on the surface**
 - bug sees an incident distribution of light arriving at the surface
 - physics question: what is the outgoing distribution of light?

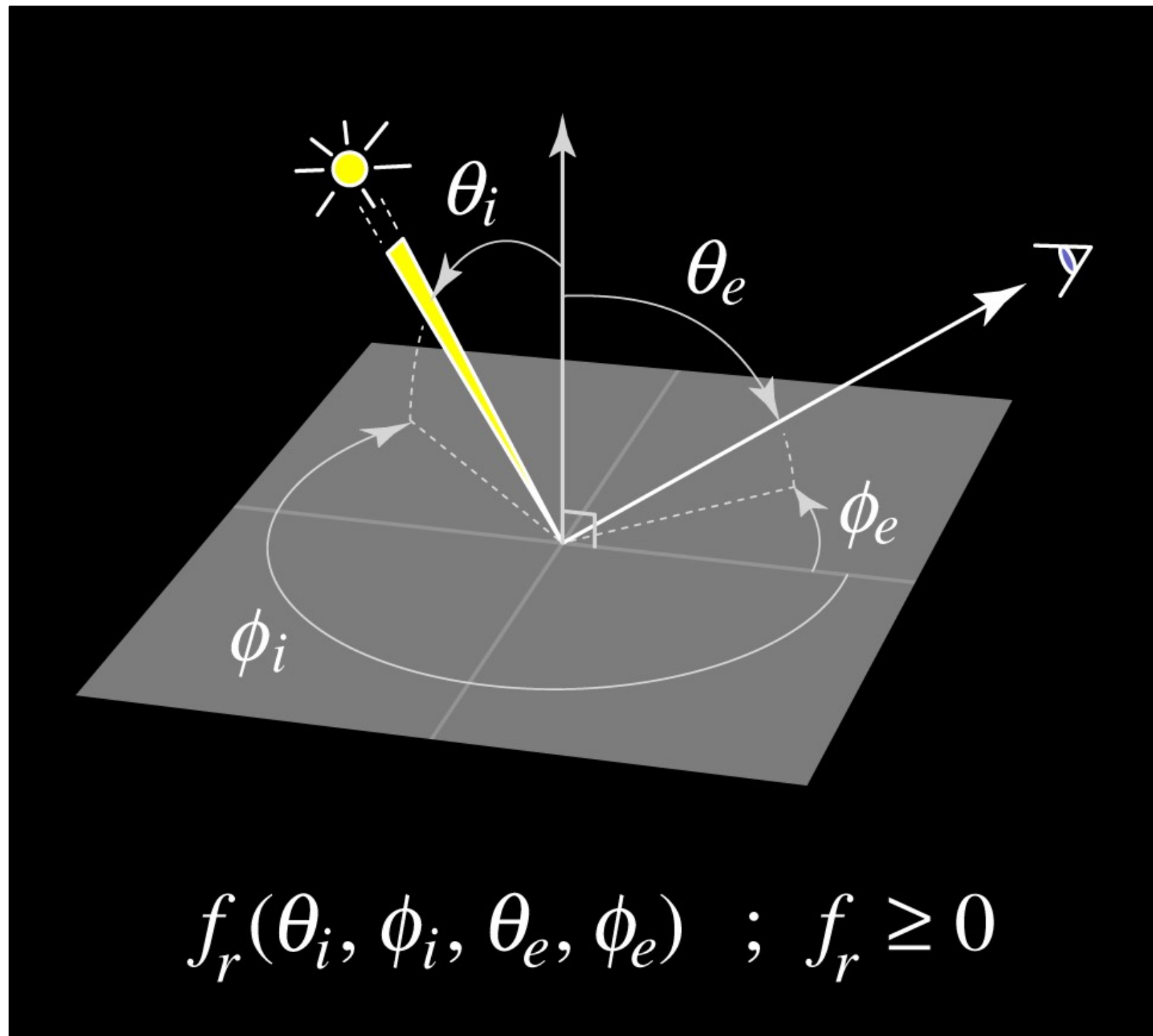


incident distribution
(function of direction)



reflected distribution
(function of direction)

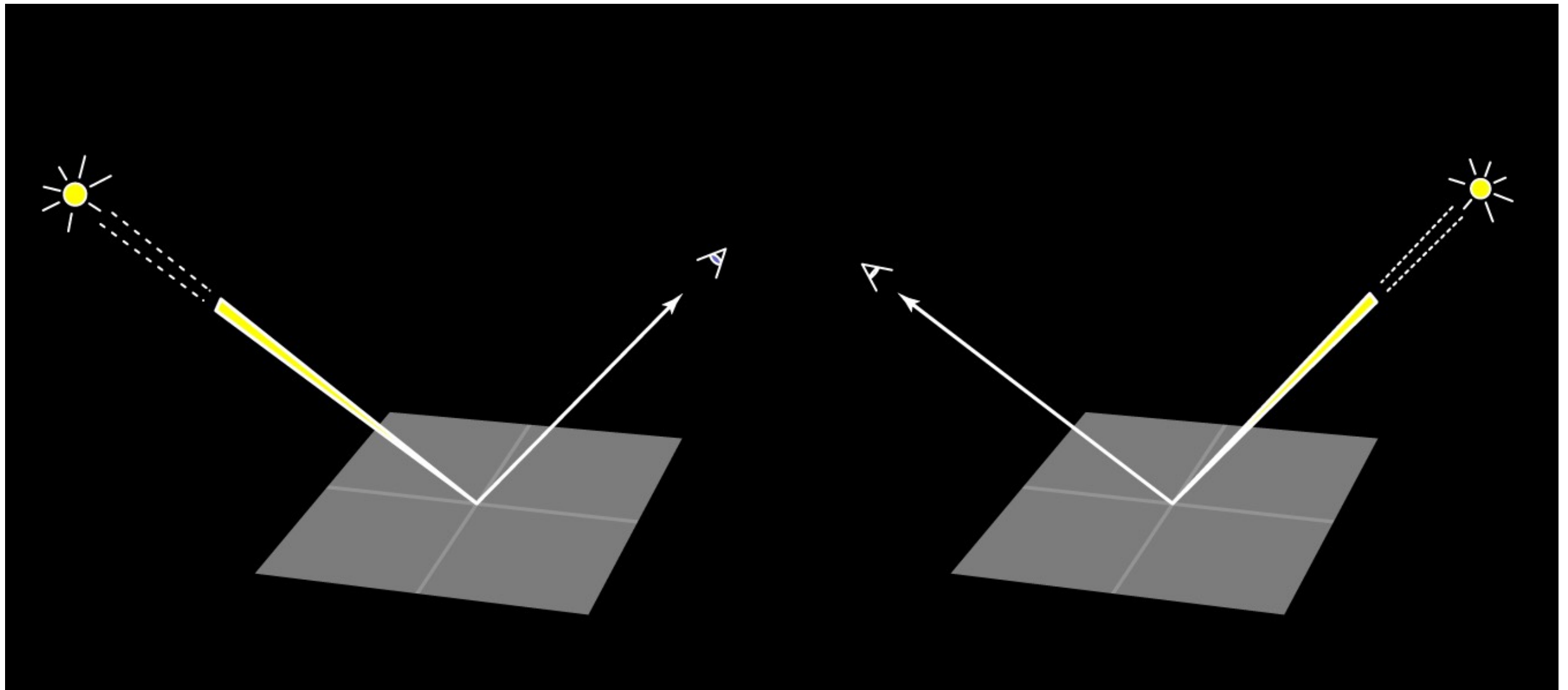
BRDF



Bidirectional Reflectance Distribution Function

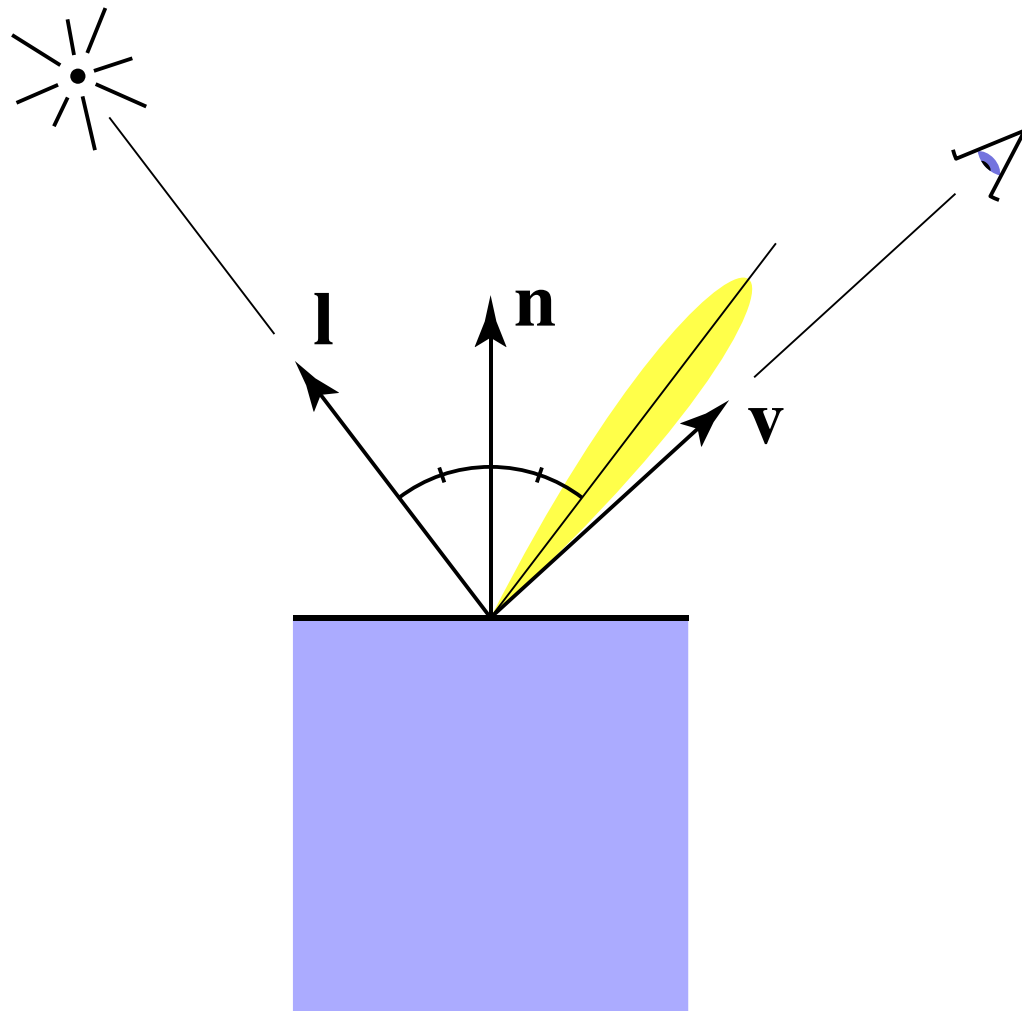
Reciprocity

- Interchanging arguments
- Physical requirement



Specular reflection

- **Intensity depends on view direction**
 - bright near mirror configuration

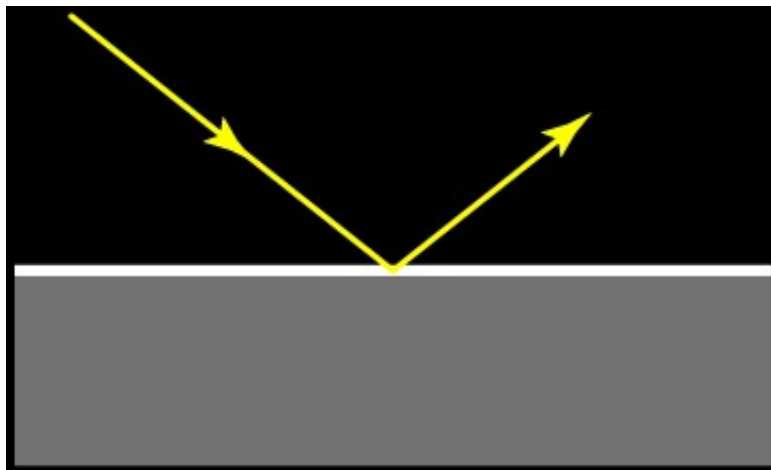


Caution: in notes and assignment, \mathbf{v} is called ω_r and \mathbf{l} is called ω_i . No meaningful difference, just notational.

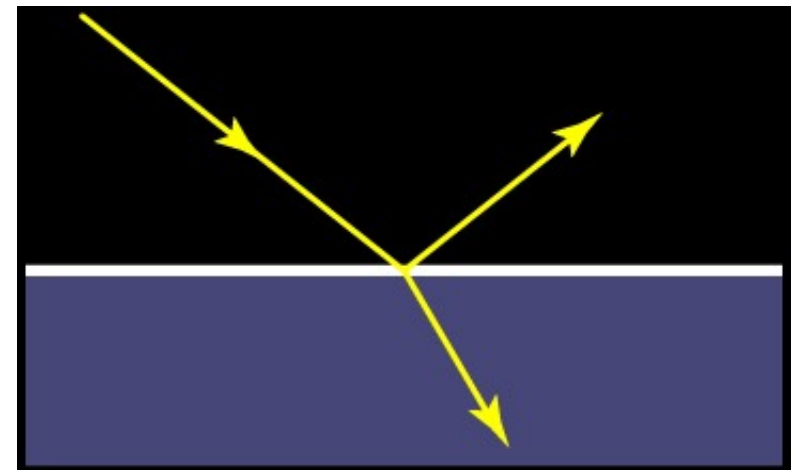
Smooth surfaces



metal

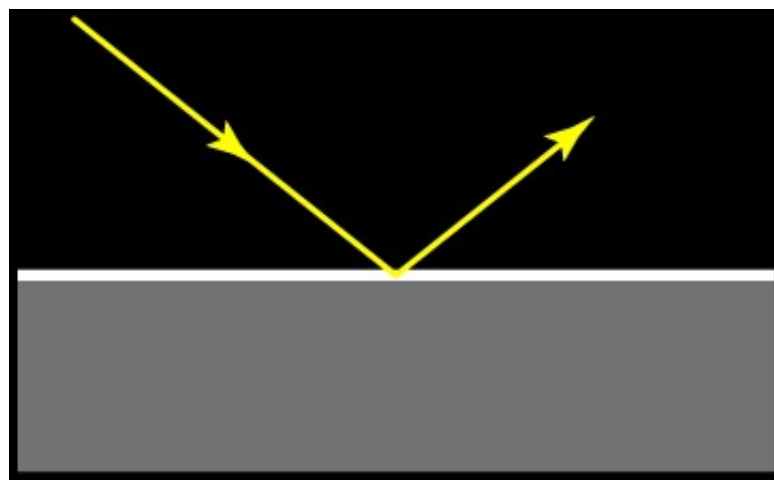


dielectric

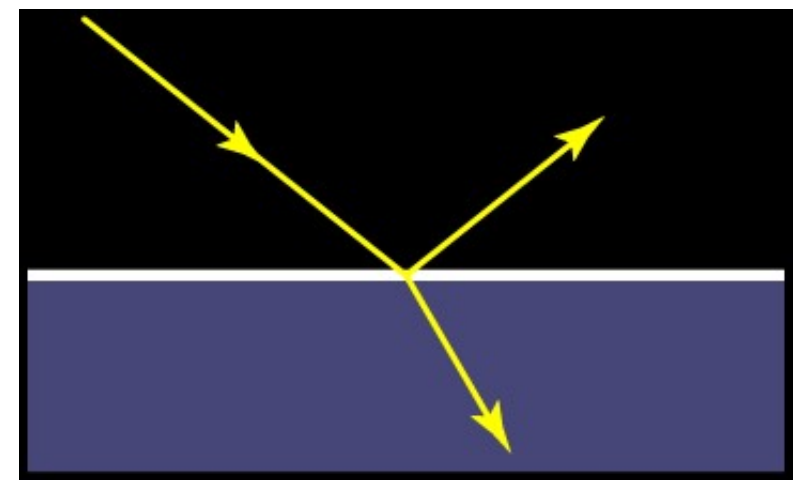


Ideal specular reflection

- **Smooth surfaces of pure materials have ideal specular reflection**
 - Metals (conductors) and dielectrics (insulators) behave differently
- **Reflectance (fraction of light reflected) depends on angle**

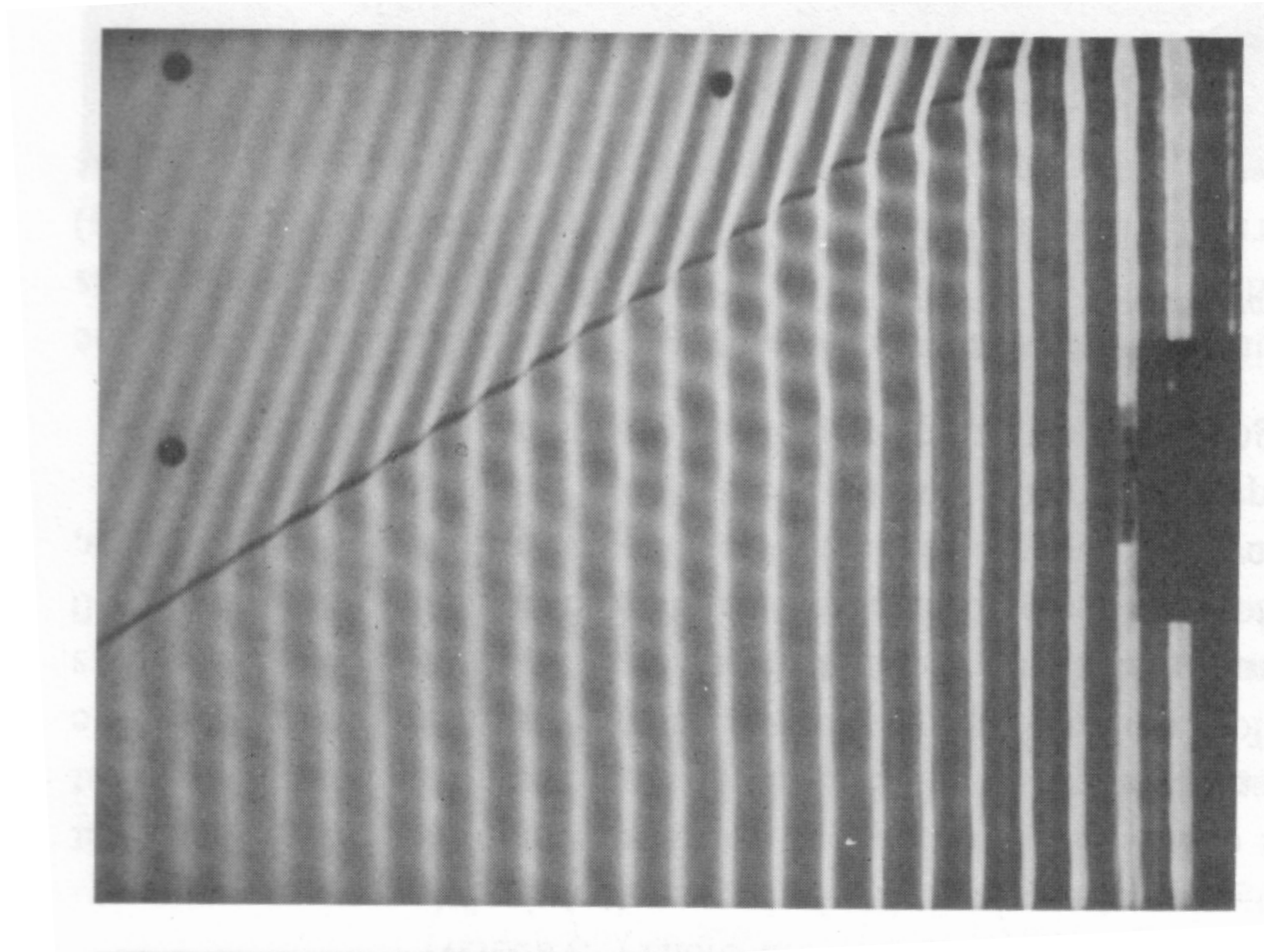


metal



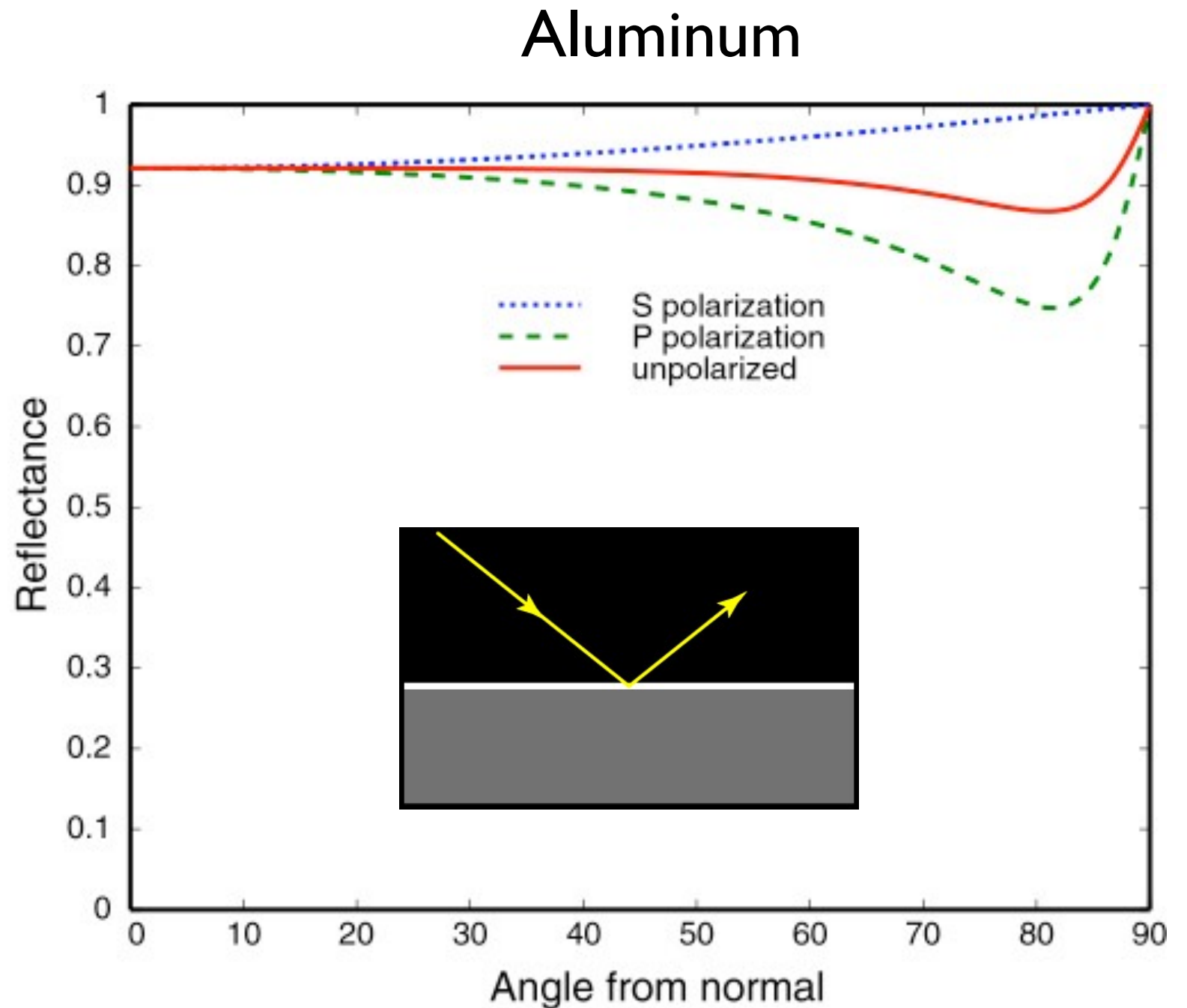
dielectric

Refraction at boundary of media



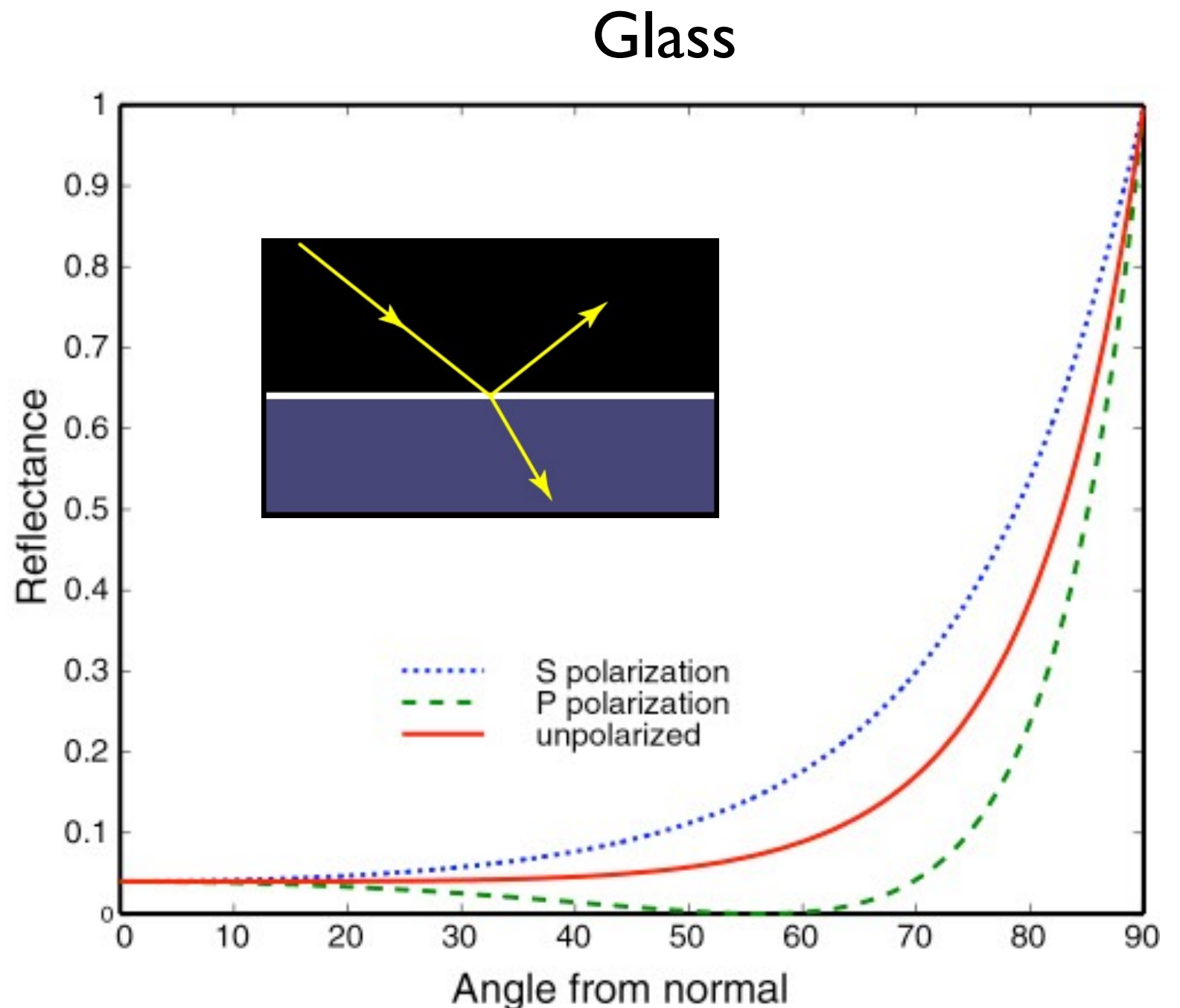
Specular reflection from metal

- **Reflectance does depend on angle**
 - but not much
 - safely ignored in basic rendering



Specular reflection from glass/water

- **Dependence on angle is dramatic!**
 - about 4% at normal incidence
 - always 100% at grazing
 - remaining light is transmitted
- **This is important for proper appearance**



Fresnel's formulas

- **They predict how much light reflects from a smooth interface between two materials**

- usually one material is empty space

$$F_p = \frac{\eta_2 \cos \theta_1 - \eta_1 \cos \theta_2}{\eta_2 \cos \theta_1 + \eta_1 \cos \theta_2}$$

$$F_s = \frac{\eta_1 \cos \theta_1 - \eta_2 \cos \theta_2}{\eta_1 \cos \theta_1 + \eta_2 \cos \theta_2}$$

$$R = \frac{1}{2} (F_p^2 + F_s^2)$$

where

$$\eta_1 \sin \theta_1 = \eta_2 \sin \theta_2$$

note: the formula in the notes and assignment is different but equivalent.

- R is the fraction that is reflected
- $(1 - R)$ is the fraction that is transmitted



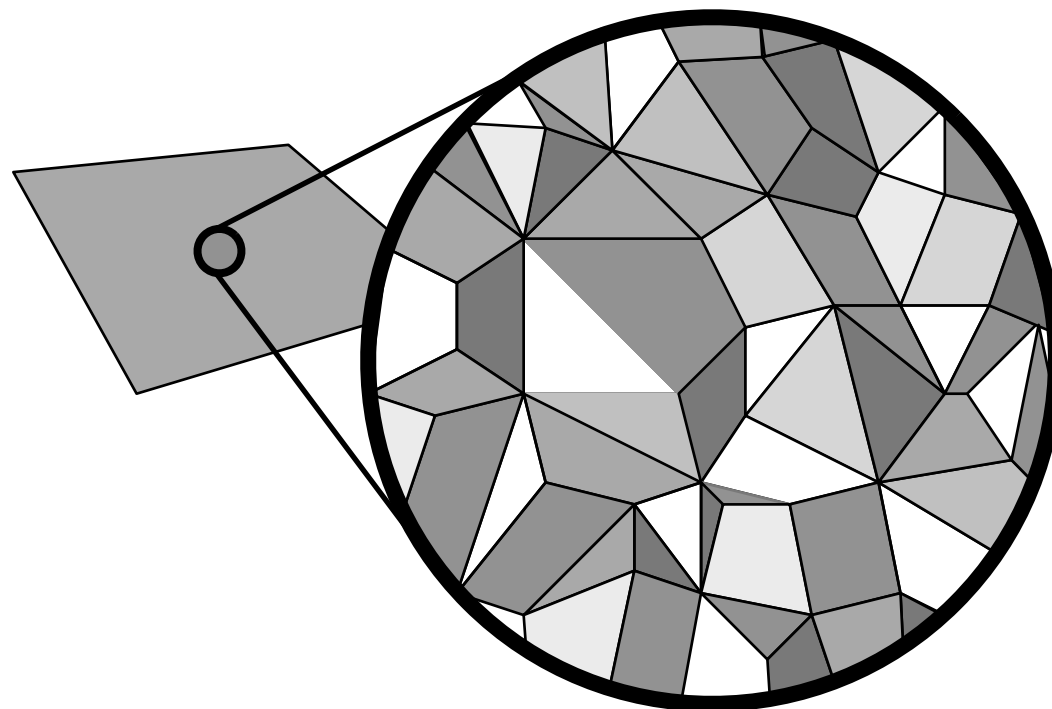
Fresnel reflection



[Mike Hill & Gaain Kwan | Stanford cs348 competition 2001]

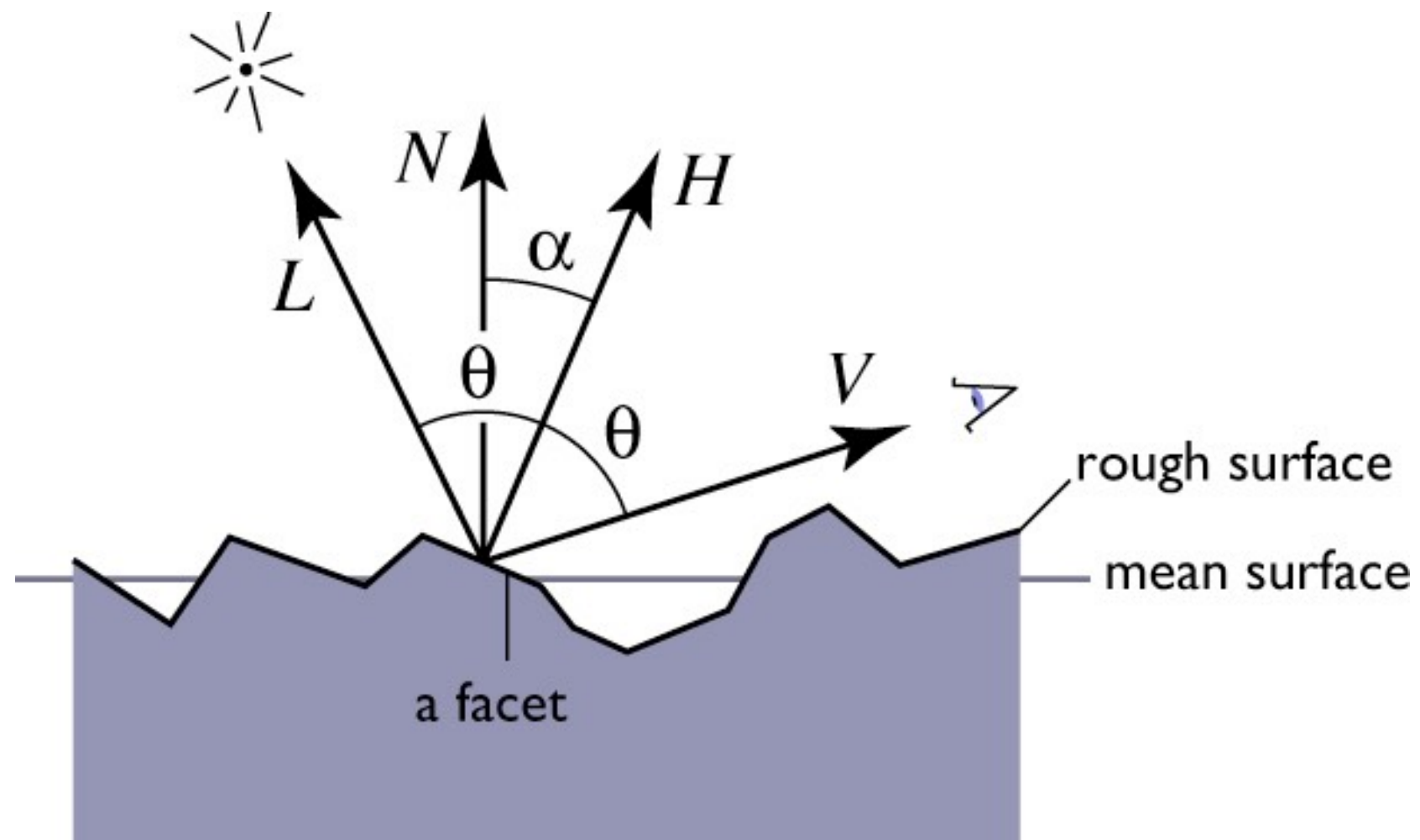
Microfacet BRDF Model

- **The *microfacet* idea**
 - surface modeled as random collection of planar facets
 - an incoming ray hits exactly one facet, at random
- **Key input: probability distribution of facet angle**



Facet Reflection

- **H vector used to define facets that contribute**
 - L and V determine H ; only facets with that normal matter
 - reflected light is proportional to number of facets



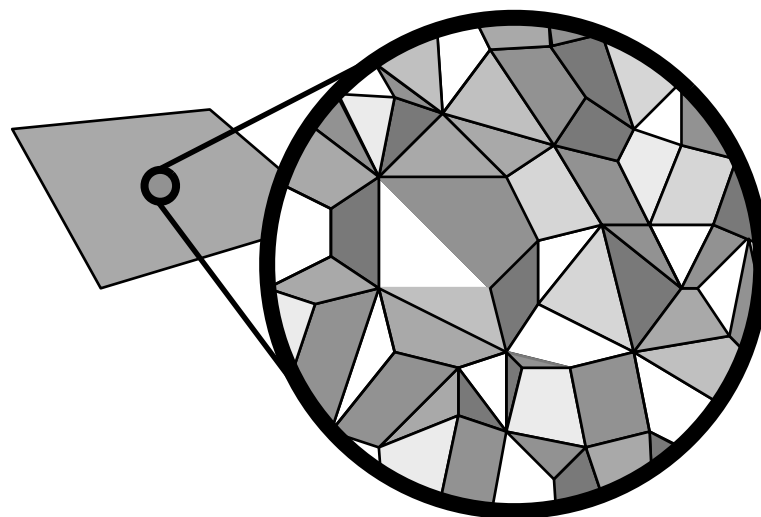
Microfacet BRDF Model

$$f_r(\mathbf{n}, \mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h})D(\mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})}{4|\mathbf{n} \cdot \mathbf{l}||\mathbf{n} \cdot \mathbf{v}|}$$

Microfacet BRDF Model

Facet distribution

$$f_r(\mathbf{n}, \mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h}) D(\mathbf{h}) G(\mathbf{l}, \mathbf{v}, \mathbf{h})}{4|\mathbf{n} \cdot \mathbf{l}||\mathbf{n} \cdot \mathbf{v}|}$$



[Stephen Westin]

Facet Distribution

- **D function describes distribution of \mathbf{h}**
- **Many choices, depending on surface characteristics**
- **A classic choice is due to Beckmann**
 - derivation based on Gaussian random processes

$$D(\mathbf{h}) = \frac{e^{-\frac{\tan^2(\mathbf{h}, \mathbf{n})}{m^2}}}{\pi m^2 \cos^4(\mathbf{h}, \mathbf{n})}$$

Cook-Torrance BRDF Model

Fresnel Reflectance

$$f_r(\mathbf{n}, \mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h}) D(\mathbf{h}) G(\mathbf{l}, \mathbf{v}, \mathbf{h})}{4 |\mathbf{n} \cdot \mathbf{l}| |\mathbf{n} \cdot \mathbf{v}|}$$

- **Fresnel reflectance for smooth facet**
 - more light reflected at grazing angles

Cook-Torrance BRDF Model

Masking/shadowing

$$f_r(\mathbf{n}, \mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h}) D(\mathbf{h}) G(\mathbf{l}, \mathbf{v}, \mathbf{h})}{4 |\mathbf{n} \cdot \mathbf{l}| |\mathbf{n} \cdot \mathbf{v}|}$$

Masking and Shadowing

- **Many options; Smith shadowing-masking follows from D**
 - long story, it is an integral related to D that often doesn't have a closed form solution
 - generally it is assumed that masking and shadowing are statistically independent
 - for Beckmann, recommend using this rational approximation due to Bruce Walter:

$$G(\omega_i, \omega_r, \mathbf{h}) = G_1(\omega_i, \mathbf{h})G_1(\omega_r, \mathbf{h})$$

$$G_1(\mathbf{v}, \mathbf{h}) = \chi^+ \left(\frac{\mathbf{v} \cdot \mathbf{h}}{\mathbf{v} \cdot \mathbf{n}} \right) \begin{cases} \frac{3.535a + 2.181a^2}{1 + 2.276a + 2.577a^2}, & \text{if } a < 1.6 \\ 1 & \text{otherwise} \end{cases}$$

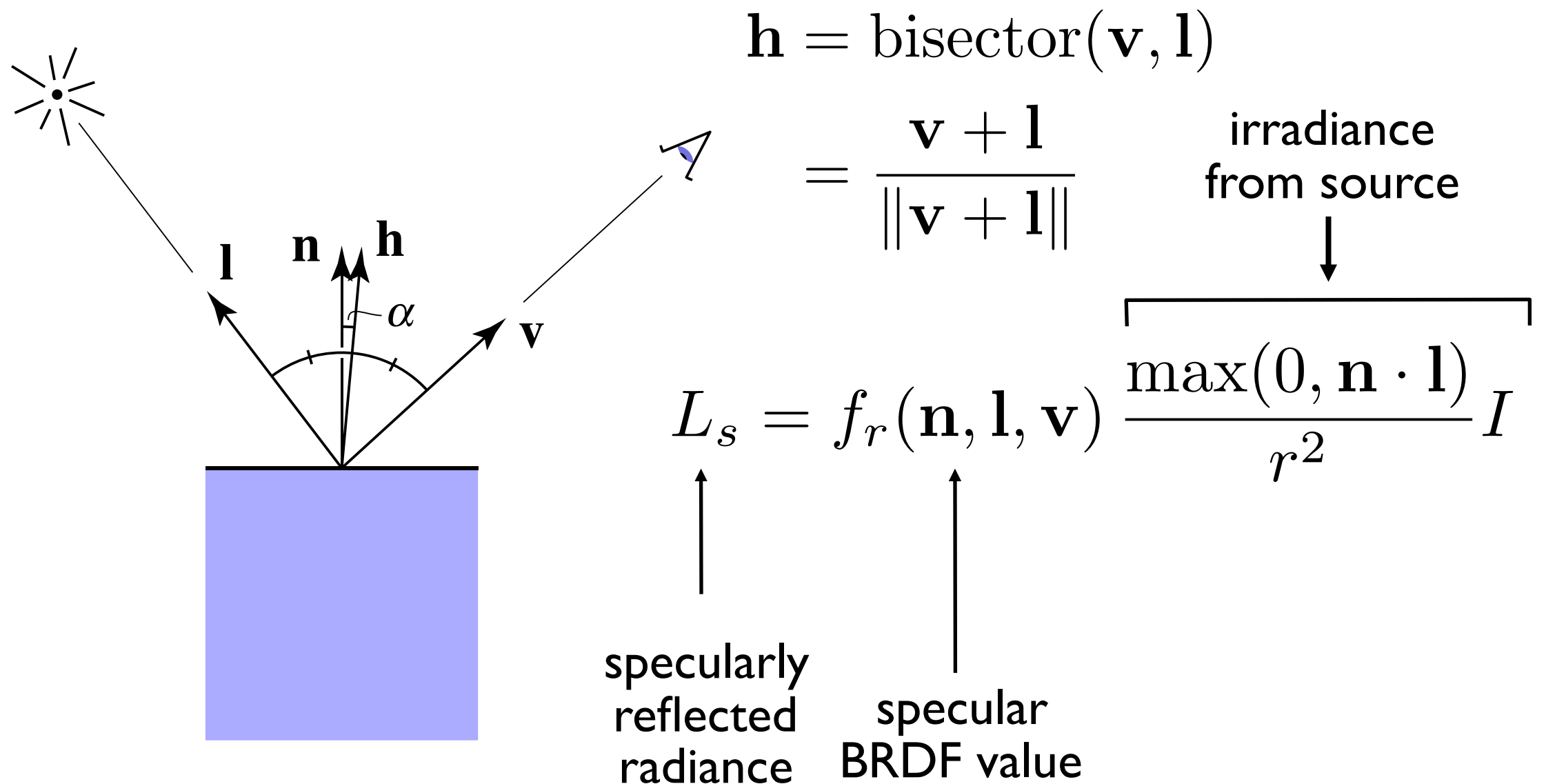
Microfacet BRDF Model

$$f_r(\mathbf{n}, \mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h}) D(\mathbf{h}) G(\mathbf{l}, \mathbf{v}, \mathbf{h})}{4|\mathbf{n} \cdot \mathbf{l}||\mathbf{n} \cdot \mathbf{v}|}$$

- reasons for cosine terms in denominator
- if one is there they clearly both have to be there (by reciprocity)

Specular shading (Microfacet)

- **Close to mirror \Leftrightarrow half vector near normal**
 - reflectance is high when $\mathbf{n} \cdot \mathbf{h}$ is near 1.0 (α is near 0)



Ray tracer architecture 101

- **You want a class called Ray**
 - point and direction; evaluate(t)
 - possible: t_{Min} , t_{Max}
- **Some things can be intersected with rays**
 - individual surfaces
 - groups of surfaces (acceleration goes here)
 - the whole scene
 - make these all subclasses of Surface
 - limit the range of valid t values (e.g. shadow rays)
- **Once you have the visible intersection, compute the color**
 - may want to separate shading code from geometry
 - separate class: Material (each Surface holds a reference to one)
 - its job is to compute the color

Architectural practicalities

- **Return values**

- surface intersection tends to want to return multiple values
 - t , surface or shader, normal vector, maybe surface point
- in many programming languages (e.g. Java) this is a pain
- typical solution: an *intersection record*
 - a class with fields for all these things
 - keep track of the intersection record for the closest intersection
 - be careful of accidental aliasing (which is very easy if you're new to Java)

- **Efficiency**

- in Java the (or, a) key to being fast is to minimize creation of objects
- what objects are created for every ray? try to find a place for them where you can reuse them.
- Shadow rays can be cheaper (any intersection will do, don't need closest)
- but: “First Get it Right, Then Make it Fast”