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## Colorimetry - Part 3: CIE tristimulus values (ISO/CIE 11664-3:2019)

Colorimétrie - Partie 3: Composantes trichromatiques CIE (ISO/CIE11664-3:2019)

Farbmetrik - Teil 3: CIE-Farbwerte (ISO/CIE 11664-3:2019)

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### European foreword

"International Commission on Illumination" in collaboration with Technical Committee CEN/TC 139 Committee Technical ģ (ENISO/CIE 11664-3:2019) has been prepared "Paints and varnishes" the secretariat of which is held by DIN. document

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# INTERNATIONAL STANDARD

First edition 2019-06

### Colorimetry —

Part 3: CIE tristimulus values

Colorimétrie —

Partie 3: Composantes trichromatiques CIE





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ISO copyright office CP 401 • Ch. de Blandonnet 8

CH-1214 Vernier, Geneva Phone: +41 22 749 01 11 Fax: +41 22 749 09 47 Email: copyright@iso.org Website: www.iso.org

Published in Switzerland

CIE Central Bureau

Babenberger straße 9/9A A-1010 Vienna, Austria Phone: +43 1 714 3187 Fax: +41 22 749 09 47 Email: ded@de.co.at Website: www.cie.co.at

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#### Foreword

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This document was prepared by the International Commission on Illumination (CIE) in cooperation with Technical Committee ISO/TC 274, Light and lighting. This first edition of ISO/CIE 11664-3 cancels and replaces ISO 11664-3:2012 | CIE S 014-3:2011, of which it constitutes a minor revision, incorporating minor editorial updates.

A list of all parts in the ISO 11664 and ISO/CIE 11664 series can be found on the ISO website.

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### Introduction

Colourstimuli with different spectral distributions can look alike. An important function of colorimetry is to determine which stimuli look alike to a given observer with a given set of colour-matching functions. This is done by calculating a set of three tristimulus values for each stimulus. Equality of tristimulus values indicates equality of colour appearance under equal irradiation and viewing conditions. This document is based on long-standing CIE recommendations (see CIE 15[1]) for the calculation of tristimulus values.

### Colorimetry —

Part 3:

## **CIE tristimulus values**

#### 1 Scope

This document specifies methods of calculating the tristimulus values of colour stimuli for which the spectral distributions are provided. These colour stimuli can be produced by self-luminous light sources or by reflecting or transmitting objects. This document requires that the colour stimulus function be tabulated at measurement intervals of 5 nm or less in a wavelength range of at least 380 nm to 780 nm. Extrapolation methods are suggested for cases where the measured wavelength range is less than 380 nm to 780 nm. The standard method is defined as summation at 1 nm intervals over the wavelength range from 360 nm to 830 nm. Alternative abridged methods are defined for larger intervals (up to 5 nm) and shorter ranges (down to 380 nm to 780 nm). The alternative methods are to be used only when appropriate and when the user has reviewed the impact on the final results. This document can be used in conjunction with the CIE 1931 standard colorimetric observer or the CIE 1964 standard colorimetric observer.

### 2 Normative references

The following documents are referred to in the text in such a way that some or all of their content constitutes requirements of this document. For dated references, only the edition cited applies. For undated references, the latest edition of the referenced document (including any amendments) applies.

ISO/CIE 11664-1, Colorimetry — Part 1: CIE standard colorimetric observers

ISO 23539, Photometry — The CIE system of physical photometry

CIES 017, ILV: International Lighting Vocabulary

### 3 Terms and definitions

For the purposes of this document, the terms and definitions given in CIE S 017 apply.

ISO and IEC maintain terminological databases for use in standardization at the following addresses:

- ISO Online browsing platform: available at https://www.iso.org/obp
- IEC Electropedia: available at http://www.electropedia.org/

# 4 Symbols and abbreviations

$k, k_{10}$	normalizing constants
$K_{\mathrm{m}}$	maximum spectral luminous efficacy of radiation in the CIE standard system of physical photometry
$K_{\mathrm{m,10}}$	maximum spectral luminous efficacy of radiation when the $V_{10}(\lambda)$ function is used for photometry
R(λ)	spectral reflectance factor
5(7)	relative spectral distribution of an illuminant
V(X)	spectral luminous efficiency function in the CIE standard system of physical photometry
V <sub>10</sub> (λ)	spectral luminous efficiency function when the $\overline{y}_{10}(\lambda)$ function is used for photometry
$W_{\mathbf{x}}(\lambda), W_{\mathbf{y}}(\lambda), W_{\mathbf{z}}(\lambda)$	pre-calculated weighting functions for tristimulus integration using the CIE 1931 standard colorimetric observer
$W_{x,10}(\lambda), W_{y,10}(\lambda), W_{z,10}(\lambda)$	pre-calculated weighting functions for tristimulus integration using the CIE 1964 standard colorimetric observer
$X, \mathcal{Y}, Z$	chromaticity coordinates calculated using the CIE 1931 standard colorimetric observer
$^{X}_{10}$ , $^{y}_{10}$ , $^{z}_{10}$	chromaticity coordinates calculated using the CIE 1964 standard colorimetric observer
$\overline{x}(\lambda), \overline{y}(\lambda), \overline{z}(\lambda)$	colour-matching functions of the CIE 1931 standard colorimetric observer (also known as the CIE 2° standard colorimetric observer)
$\overline{x}_{10}(\lambda), \overline{y}_{10}(\lambda), \overline{z}_{10}(\lambda)$	colour-matching functions of the CIE 1964 standard colorimetric observer (also known as the CIE 10° standard colorimetric observer)
X,Y,Z	tristimulus values calculated using the CIE 1931 standard colorimetric observer
$X_{10'} Y_{10'} Z_{10}$	tristimulus values calculated using the CIE 1964 standard colorimetric observer
$\beta(\lambda)$	spectral radiance factor
Δλ	wavelength interval
$\varphi_{\lambda}(\lambda)$	colour stimulus function (description of a colour stimulus by the spectral concentration of a radiometric quantity, such as radiance or radiant power, as a function of wavelength)
$\varphi(\lambda)$	relative colour stimulus function (relative spectral distribution of the colour stimulus function)
У	wavelength
p(3)	spectral reflectance

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spectral transmittance

### 5 Standard method

#### 5.1 General

CIE 1964 standard colorimetric observer. If the angle subtended at the eye by the colour stimulus (or This document may be used in conjunction with the CIE 1931 standard colorimetric observer or the fields to be matched in colour) is between about 1° and 4° the CIE 1931 standard colorimetric observer shall be used. If this angular subtense is greater than 4° the CIE 1964 standard colorimetric observer shall be used. The same colorimetric observer shall be used for all stimuli to be compared with each other.

# 5.2 Calculation of tristimulus values

In the CIE 1931 standard colorimetric system, tristimulus values X, Y and Z are defined as integrals over the spectral range 360 nm to 830 nm according to Formulae (1) to (3):

$$X = k \int_{\mathcal{Q}} \varphi_{\lambda}(\lambda) \, \overline{x}(\lambda) \, d\lambda \tag{1}$$

$$Y = k \int_{\mathbf{Q}} \varphi_{\lambda}(\lambda) \, \overline{y}(\lambda) \, d\lambda \tag{2}$$

$$Z = k \int_{\lambda} \varphi_{\lambda}(\lambda) \bar{z}(\lambda) d\lambda \tag{3}$$

where

is the colour stimulus function to be evaluated;

are the colour-matching functions of the CIE 1931 standard colorimetric observer  $\overline{x}(\lambda), \overline{y}(\lambda), \overline{z}(\lambda)$ 

is a normalizing constant defined in 5.3 and 5.4.

The standard method for evaluating these integrals is numerical summation from 360 nm to 830 nm at wavelength intervals,  $\Delta \lambda$ , equal to 1 nm according to Formulae (4) to (6):

$$X = k \sum_{j} \varphi_{\lambda}(\lambda) \, \overline{x}(\lambda) \, \Delta \lambda \tag{4}$$

$$Y = k \sum_{j} \varphi_{\lambda}(\lambda) \, \overline{y}(\lambda) \, \Delta \lambda \tag{5}$$

$$Z = k \sum_{\lambda} \varphi_{\lambda}(\lambda) \, \overline{z}(\lambda) \, \Delta \lambda \tag{6}$$

ISO/CIE 11664-1 and a colour stimulus function,  $arphi_\lambda(\lambda)$ , measured using a symmetrical triangular or Ξ using colour-matching functions  $\bar{x}(\lambda), \bar{y}(\lambda), \bar{z}(\lambda)$  defined with seven significant figures trapezoidal bandpass with a half width equal to 1 nm.

function,  $\varphi(\lambda)$ , may be used instead of the colour stimulus function,  $\varphi_{\lambda}(\lambda)$ . It is essential that, for stimuli that will be considered together, all the spectral distributions involved be assessed on the same relative basis. The tristimulus values obtained are then relative in the sense that all the values involved may Tristimulus values are often evaluated on a relative basis. In such cases the relative colour stimulus

be multiplied by the same single arbitrary constant, k. In certain cases, however, k shall be chosen according to agreed conventions; these conventions are explained in 5.3 and 5.4. NOTE The wavelength range of 360 nm to 830 nm is in accordance with established CIE practice (ISO/CIE 11664-1 and Reference [1]). Clause 6 of this document specifies abridged methods that can be used when data are not available over the full range of 360 nm to 830 nm at 1 nm intervals.

## Normalizing constant for self-luminous light sources 53

of the radiometric quantity corresponding to the photometric quantity required, and the constant, k, shall be set equal to 683 lm·W<sup>-1</sup> which is the numerical value of  $K_{\rm m}$ , the maximum spectral luminous efficacy in the CIE system of physical photometry as specified in ISO 23539 | CIE S 010. For self-luminous objects, the normalizing constant, k, is usually chosen on the grounds of convenience. If, however, in the CIE 1931 standard colorimetric system, the Y value is required to be numerically equal to the absolute value of a photopic photometric quantity,  $\varphi_{\lambda}(\lambda)$  shall be the spectral concentration

## Normalizing constant for reflecting or transmitting objects 4.5

For reflecting or transmitting object colours, the colour stimulus function,  $\varphi_{\lambda}(\lambda)$ , shall be replaced by the relative colour stimulus function,  $\varphi(\lambda)$ , evaluated as shown in Formulae (7) to (10).

$$\varphi(\lambda) = R(\lambda)S(\lambda) \tag{7}$$

or

$$\varphi(\lambda) = \beta(\lambda)S(\lambda) \tag{8}$$

o

$$\phi(\lambda) = \rho(\lambda)S(\lambda) \tag{9}$$

or

$$\phi(\lambda) = \tau(\lambda)S(\lambda) \tag{10}$$

where

- R(λ) is the spectral reflectance factor;
- $\beta(\lambda)$  is the spectral radiance factor;
- $\rho(\lambda)$  is the spectral reflectance;
- τ(λ) is the spectral transmittance;
- S(以 is the relative spectral distribution of the illuminant.

In all these cases, the constant, k, shall be chosen so that Y = 100 for objects for which  $R(\lambda)$  [or  $\beta(\lambda)$ ,  $\rho(\lambda)$ , τ(λ)] equals 1 for all wavelengths. Hence, as shown in Formula (11):

$$k = 100 / \sum_{\lambda} S(\lambda) \bar{y}(\lambda) \Delta \lambda \tag{11}$$

where the summation range and interval, and the values of  $\overline{y}(\lambda)$ , are the same as in Formulae (4) to (6).

The values of Y for all objects are then equal to the percentage values of luminous reflectance factor [in the case of  $R(\lambda)$ ], luminous reflectance [in the case of  $\rho(\lambda)$ ] or

luminous transmittance [in the case of  $r(\lambda)$ ]. This is because the  $\overline{y}(\lambda)$  function is identical to the CIE spectral luminous efficiency function  $V(\lambda)$  NOTE All four quantities,  $R(\lambda)$ ,  $\beta(\lambda)$ ,  $\rho(\lambda)$  and  $\tau(\lambda)$ , are ratios. If, for convenience, any of these quantities are reported as percentages, the numerical values must be divided by 100 for the above derivation of k to be correct.

# 5.5 CIE 1964 standard colorimetric system

The colour-matching functions  $\overline{x}_{10}(\lambda)$ ,  $\overline{y}_{10}(\lambda)$ ,  $\overline{z}_{10}(\lambda)$  of the CIE 1964 standard colorimetric observer (ISO/CIE 11664-1) may be used in place of  $\overline{x}(\lambda)$ ,  $\overline{y}(\lambda)$ ,  $\overline{z}(\lambda)$ . In this case, the symbols X, Y, Z and k shall be replaced by  $X_{10}$ ,  $Y_{10}$ ,  $Z_{10}$  and  $k_{10}$  in all formulae in this document.

(683,6 lm·W<sup>-1</sup>) has not been standardized or approved by the General Conference on Weights and Measures (Conférence Générale des Poids et Mesures, CGPM). However, CIE 165:2005[2] has recommended that the function Use of the  $\bar{y}_{10}(\lambda)$  function for photometry (as in 5.3), with the appropriate value of  $K_{m,10}$ , can be so used, especially if luminance has to be determined parafove ally.

### 6 Abridged methods

#### 6.1 General

In some cases, the standard method defined in Clause 5 cannot be used because the colour stimulus function or relative colour stimulus function is not available over the full range of 360 nm to 830 nm in 1 nm intervals.

## Abridged method for data at 5 nm intervals or less 6.2

If it is demonstrated that the resulting errors are insignificant for the purpose of the user, tristimulus values X, Y, Z shall be calculated by numerical summation from 380 nm to 780 nm at wavelength intervals,  $\Delta\lambda$ , equal to 5 nm according to Formulae (4) to (6) using colour-matching functions  $\overline{x}(\lambda), \overline{y}(\lambda), \overline{z}(\lambda)$ , as defined with seven significant figures in ISO/CIE 11664-1 (see  $\overline{24}$  for bandwidth requirements If the colour stimulus function or relative colour stimulus function data are provided at wavelength intervals of 2 nm, 3 nm or 4 nm, the same method shall be used, subject to the same conditions. If the wavelength interval is less than 5 nm but not an integer multiple of 1 nm, either the colour-matching functions and the illuminant or the colour stimulus data shall be interpolated so that they match. See 7.3 for guidance on this, NOTE 1 Some publications give the colour-matching functions from ISO/CIE 11664-1 with values rounded to four significant figures. These rounded values can be used provided that it is demonstrated that the resulting errors are insignificant for the purpose of the user.

NOTE2 Some CCD array spectrometers record data at unequal wavelength intervals. In this case,  $\Delta\lambda$  Formulae (4), (5), (6) and (11) will vary.

# Abridged method for 10 nm or 20 nm data for reflecting or transmitting objects 6.3

nm This document does not coverabridged methods for 10 nm or 20 nm data. It applies only to data at 5 intervals or less.

A common method for calculating tristimulus values X, Y, Z of reflecting or transmitting object colours at wavelength intervals, Δλ, equal to 10 nm or 20 nm is to use Formulae (12) to (14):

$$X = \sum_{\lambda} R(\lambda) W_{x}(\lambda) \tag{12}$$

$$Y = \sum_{\lambda} R(\lambda) W_{y}(\lambda) \tag{13}$$

$$Z = \sum_{\lambda} R(\lambda) W_{z}(\lambda) \tag{14}$$

where

is the spectral reflectance factor measured using a symmetrical triangular or trapezoidal bandpass with a half width equal to the wavelength interval (10 nm or 20 nm);

to give the best fit to the standard method (see Clause 5) based on the assumption that  $R(\lambda)$  varies smoothly between the measured (10 nm or the wavelength interval, the bandwidth and the normalizing constant, are pre-calculated weighting functions that take into account the colour-matching functions, the relative spectral distribution of the illuminant, 20 nm interval) values.  $W_{\mathbf{x}}(\lambda)$ ,  $W_{\mathbf{y}}(\lambda)$  and  $W_{\mathbf{z}}(\lambda)$ 

The spectral reflectance factor R(λ) may be replaced in Formulae (12) to (14) by the spectral radiance factor  $\beta(\lambda)$ , the spectral reflectance  $\rho(\lambda)$  or the spectral transmittance  $\tau(\lambda)$ . Examples of pre-calculated weighting functions prepared for this purpose are given in AST M E308-15<sup>[3]</sup>,

Details of the calculation of weighting functions have been published (see References [4], [5], [6] and [Z]). NOTE3

# Abridged method for 10 nm or 20 nm data for self-luminous light sources 6.4

as fluorescent lamps, gas discharge lamps and light-emitting diodes, calculation of tristimulus values from stimulus functions measured at wavelength intervals,  $\Delta\lambda$ , greater than 5 nm will not give accurate This document does not cover abridged methods for 10 nm or 20 nm data. It applies only to data at 5 nm intervals or less. Data intervals of larger than 5 nm should not be used for light sources, except when it is demonstrated that the error for a larger interval is negligible for the particular light source being evaluated. For many self-luminous light sources, particularly those with narrow-band features such

# 7 Supplementary treatment of input data

#### 7.1 Genera

This clause outlines supplementary treatment of data necessary to apply the methods of Clauses 5 and 6 or to correct measured data for improved accuracy.

wavelength range, interval and bandwidth throughout for any set of calculations in which data for may not be available because the measurement was made at intervals greater than specified, and/or unequal wavelength intervals were used, and/or data at the spectral extremes were omitted, and/or the The use of the methods described in Clauses 5 and 6 of this document requires that the colour stimulus range, at a specified interval, and with a specified bandwidth. It is important to use the same different colours are to be compared precisely. In practical applications, however, all the required data bandwidth was not equal to the sampling interval, and /or the bandpass shape was not a symmetrical calculation from predicted data may be inexact. Thus, prediction methods should only be used if the function,  $\varphi_{\lambda}(\lambda)$ , or the relative colour stimulus function,  $\varphi(\lambda)$ , be known over a specified wavelength triangle or trapezium. Sometimes it is possible to predict the needed but unmeasured data, although

user has reviewed the impact on the final results and has demonstrated that the resulting errors insignificant for the purpose of the user. Some guidance and warnings are given in Z2, Z3 and Z4.

### 7.2 Extrapolation

Extrapolation of measured data if the measurement range is less than 380 nm to 780 nm may cause errors and shall be used only if it can be demonstrated that the resulting errors are insignificant for the purpose of the user: NOTE 1 When predicting needed values of  $\varphi_{\lambda}(\lambda)$ ,  $\varphi(\lambda)$ ,  $\beta(\lambda)$ ,  $\rho(\lambda)$  or  $\tau(\lambda)$  beyond the measured range, unmeasured values can, as a rough approximation, be set equal to the nearest measured value of the appropriate quantity(1) or simple linear extrapolation can be used(8). Missing values can be given other values, such as zero or 100 %, if there is a reason to do so based on the data or other experience(8).

NOTE2 The sum of the weights at the unmeasured wavelengths divided by the sum of the weights at all wavelengths is ameasure of the maximum error introduced by using predicted values instead of measured ones.

### 7.3 Interpolation

an equation exists, or by mathematical curve fitting. It is recommended that Sprague interpolation is used if the wavelength intervals are constant, and third order spline functions for non-even wavelength increments[8]. If the colour-matching functions are interpolated, linear interpolation shall be used to If the wavelengths of the measured data do not match exactly those of the colour-matching functions, one or the other need to be interpolated so that they match. For reflectance and transmittance data, and matching functions may be interpolated. For interpolation of measured data, an estimate of values between the measured values may be found by a theoretical equation representing the data if such interpolate points between the 1 nm intervals (see ISO/CIE 11664-1). For light source data that can include narrowband peaks and emission lines, the measured data shall not be interpolated, and the for light sources having smooth spectral distribution curves, either the measured data or the colourcolour-matching functions shall be interpolated to match the wavelengths of the measured data. NOTE1 Interpolation of measured data to smaller intervals usually does not improve calculated colour accuracy. The data intervals described in  $\overline{\text{Clauses 5}}$  and  $\underline{6}$  apply to the original measured data. If original data interval is much larger than 5 nm, the data do not meet the requirements of this document even if they are converted to 5 nm or smaller intervals by interpolation. NOTE 2 The wavelength interval of instruments is often matched with the instrument bandwidth. In such cases, if the data are interpolated to a different interval, the integrity of the matching will be lost, and the converted data will not be applicable for bandpass corrections such as those referenced in 24.

#### 7.4 Bandwidth

All spectral data obtained from instruments have finite bandwidth, which will propagate to errors in calculated tristimulus values. Errors arising from bandwidth are generally much larger (by an order of magnitude) than the calculation errors associated with data intervals. The bandwidth of the instruments used to produce the input data for this document shall be 5 nm or less, unless the bandwidth is corrected by one of the available methods such as given in References [9], [10], [11], [12], [13], [14], [15], [5] and [16].

sampled relative to other wavelength points. For this, the bandpass should be a symmetrical triangular shape, and its half-bandwidth should be matched with the wavelength interval or an integer multiple thereof. This matching condition is not critical lifthe interval is much smaller than the bandwidth. This For the measurement of light sources having emission lines, such as discharge lamps, the data interval and the bandwidth shall be matched so that any wavelength points will not be under-sampled or overmatching condition is not required for reflectance/transmittance measurements of object colours, because these spectra are relatively smooth. An exception is, however, when a bandpass correction is to be applied, in which case the matching condition is required.

## 8 Chromaticity coordinates

If required, chromaticity coordinates x, y and z shall be calculated from the tristimulus values X, Y and Z according to Formulae (15), (16) and (17):

$$\frac{X}{X+Y+X} = 1$$

$$y = \frac{\gamma}{X + Y + Z} \tag{16}$$

$$(17) Z = \frac{Z}{X + Y + Z}$$

Because of the relation x + y + z = 1, it suffices to quote x and y only.

The diagram produced by plotting x as abscissa and y as ordinate shall be referred to as the CIE 1931 chromaticity diagram or the CIE (x, y) diagram. The chromaticity coordinates  $x_{10}$ ,  $y_{10}$  and  $z_{10}$  shall be computed similarly from  $X_{10}$ ,  $Y_{10}$  and  $Z_{10}$ , and the diagram produced by plotting  $x_{10}$  as abscissa and  $y_{10}$  as ordinate shall be referred to as the CIE 1964 chromaticity diagram or the CIE  $(x_{10}, y_{10})$  diagram.

## 9 Numerical procedures

All numerical calculations shall be carried out using the full number of significant digits provided by the input data. Final results shall be rounded to the number of significant digits indicated by the uncertainty of the measurements.

## 10 Presentation of results

Any reporting of tristimulus values and derived parameters shall be accompanied by a statement concerning the measurement geometry, observer, illuminant (for object colours) and sample backing (for reflecting object colours) used to generate them. The report shall also specify the wavelength range and the interval of the summation.

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