

# Strong lensing with upcoming wide-field radio surveys

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# Why do strong lensing science in the radio?

Extremely high angular resolution with VLBI

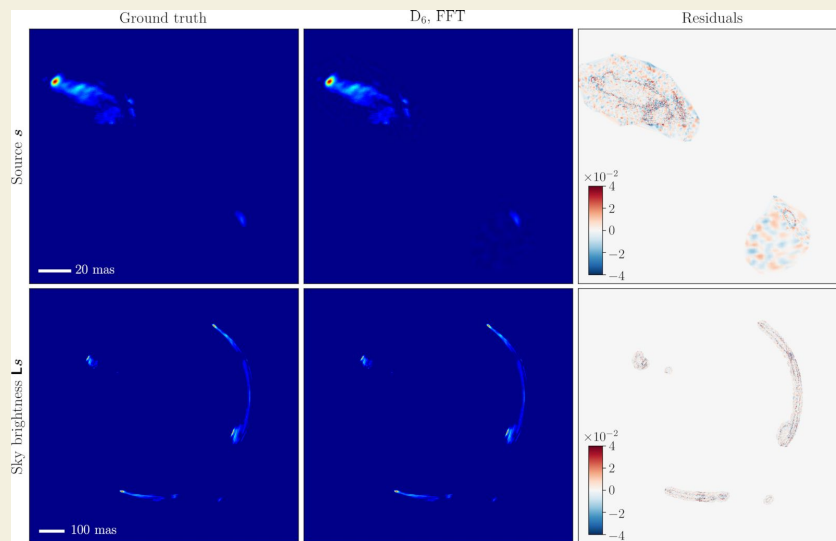
Radio sources observable to high redshift

Polarization (Stokes IQUV is preserved)

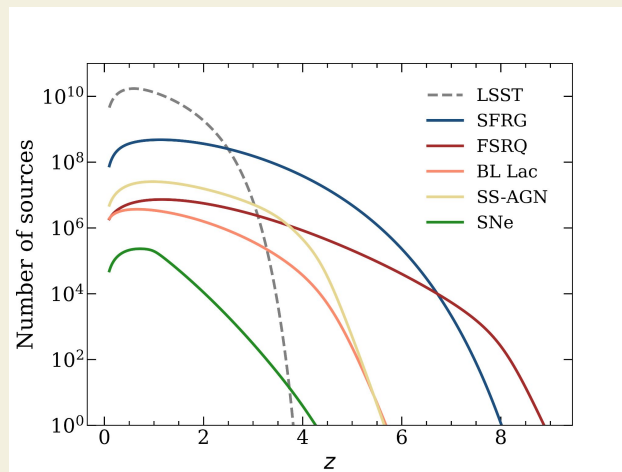
Clean – deflector galaxy is often radio quiet

No dust, no atmosphere (seeing)

Complementary information to other wavelengths!



0.25% error in recovered lens parameters with 3 mas resolution from VLBI!



$\langle z_{\text{source}} \rangle \approx 2$   
For DSA-2000

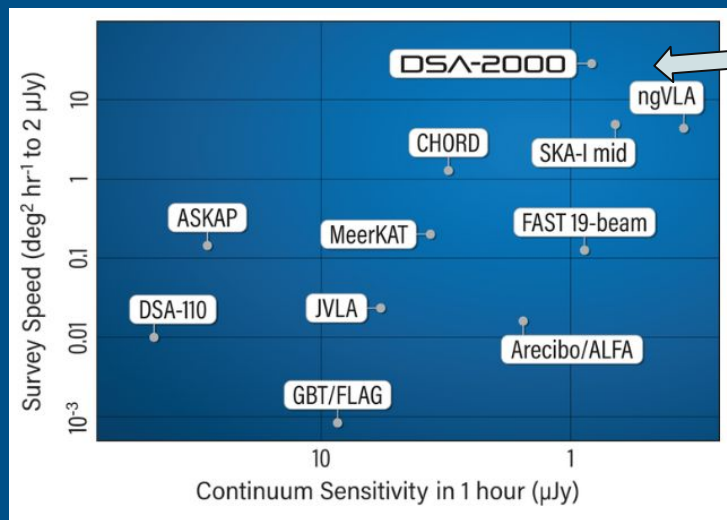
**Strong lens studies in the radio have been limited in past decades by the less than ~100 known radio lenses. How will this change with next generation radio telescopes?**

# DSA-2000

The next generation radio survey

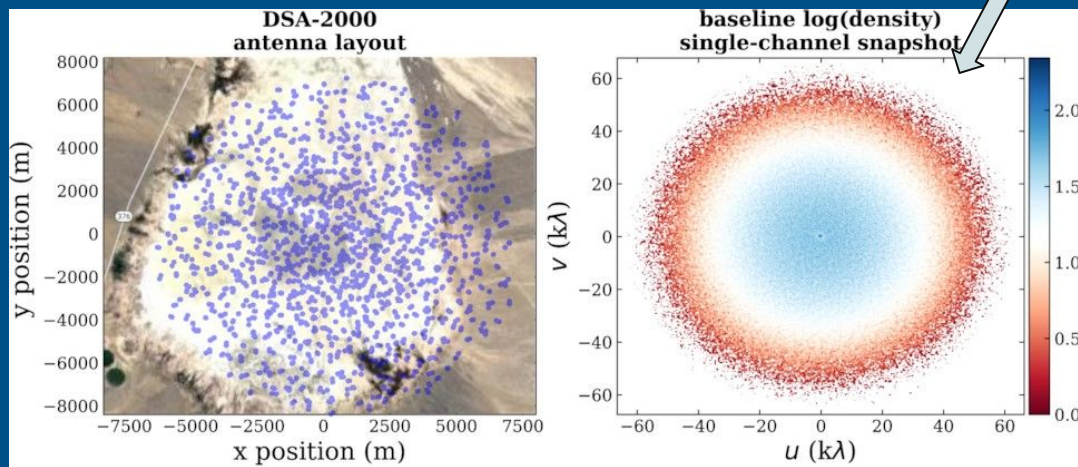
Unprecedented survey speed and sensitivity

2'' resolution at top of 0.7–2 GHz radio band



Order of magnitude better survey speed!

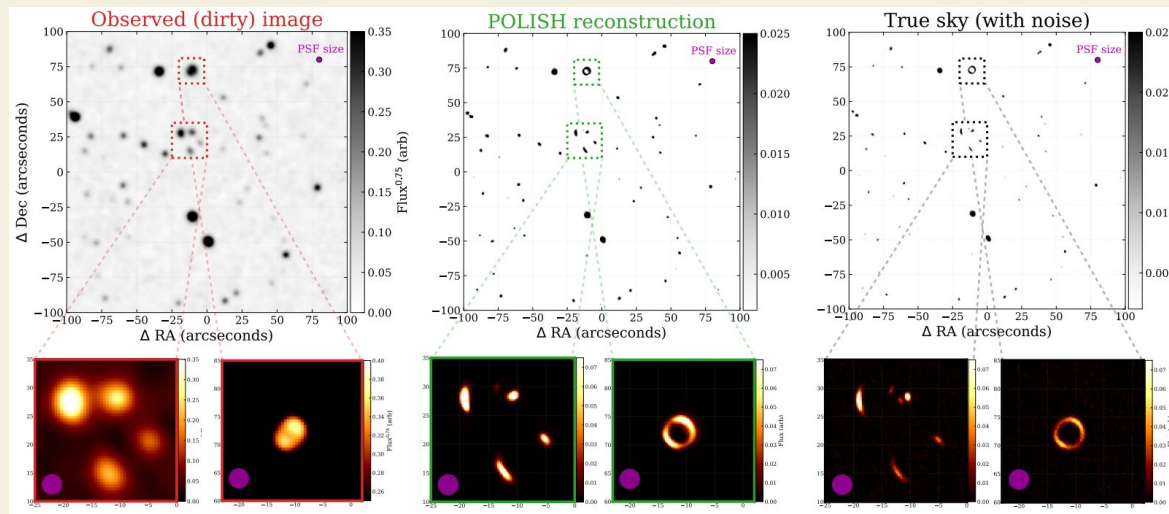
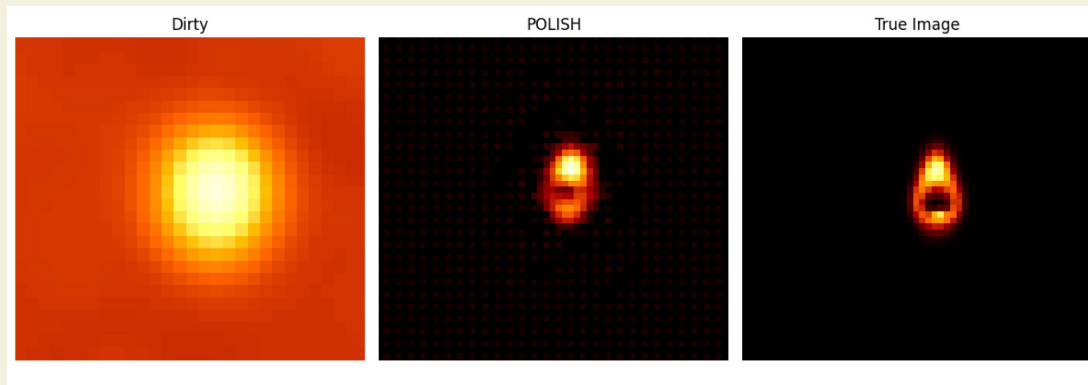
Filled aperture radio "camera"!



# Superresolution with the DSA-2000 for lens finding

Superresolution down to  $1''$  for DSA-2000 enables discovery of more galaxy scale lenses

Not necessary for lens finding, but possible with modern computer vision and deterministic PSF in the radio

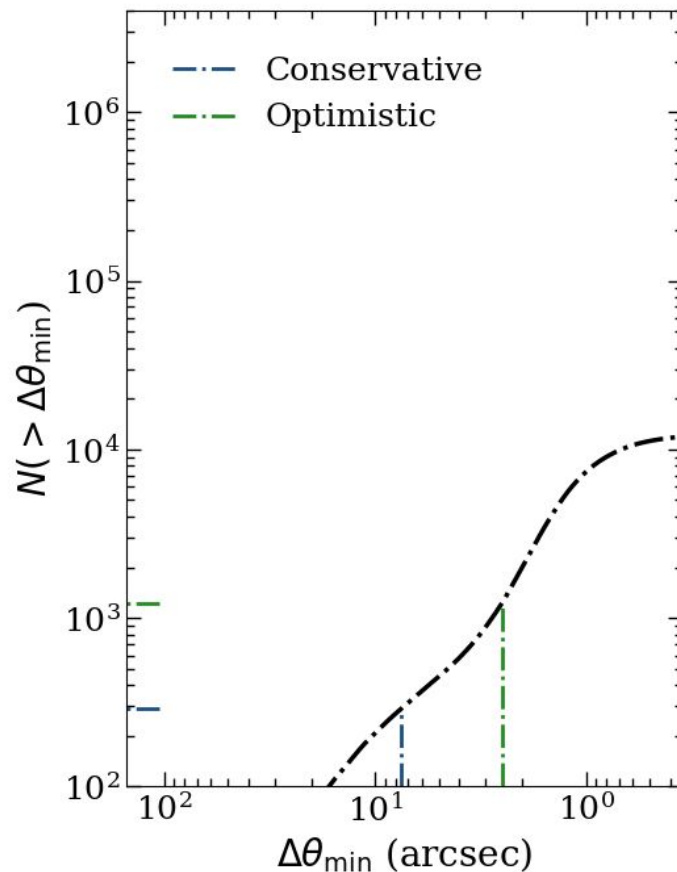


# Our work

Forecasting expected lensing yields in DSA-2000, SKA-mid, and VLA all sky surveys, using up to date expected performances

Modeling galaxy, galaxy group, and galaxy cluster lenses

$\Delta\theta > 3 \times \text{PSF FWHM}$  (conservative) or  $\Delta\theta > \text{PSF FWHM}$  (optimistic)



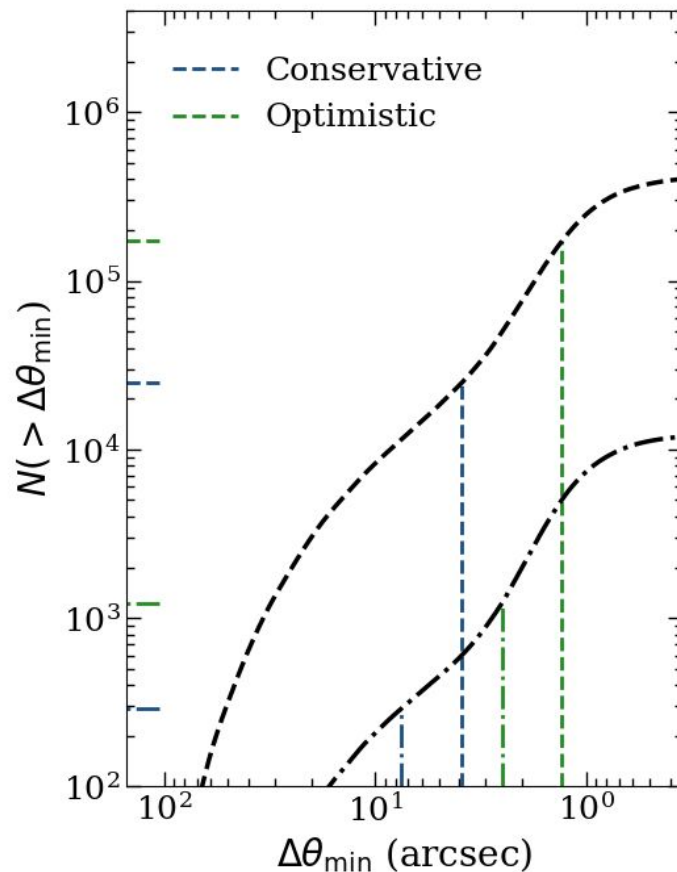
VLA

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SKA-mid AA\*

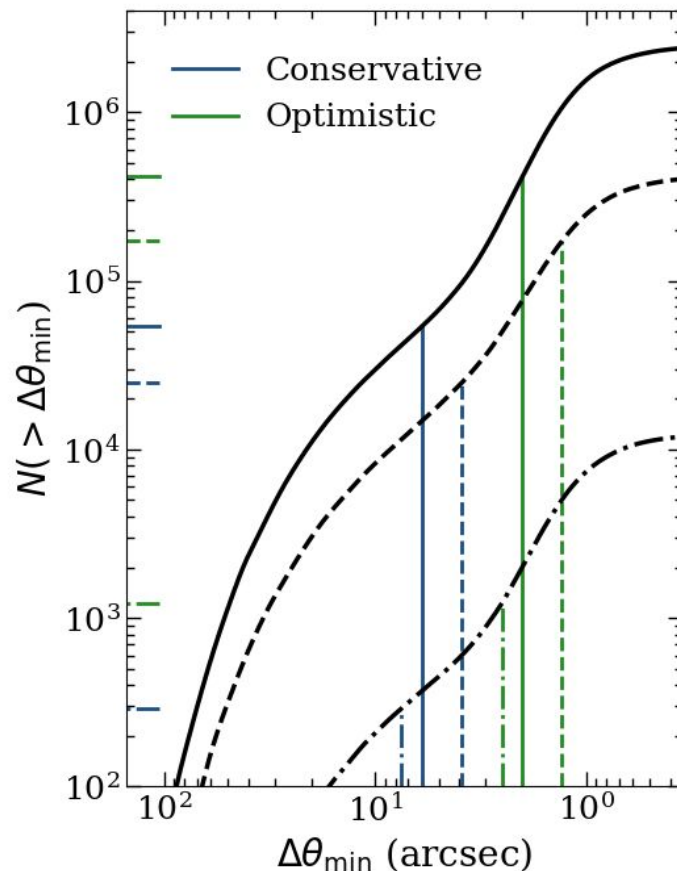
VLASS

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DSA-2000

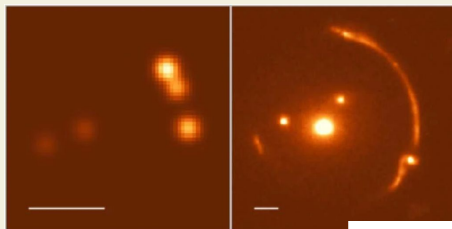
SKA-mid AA\*

VLASS



# Applications

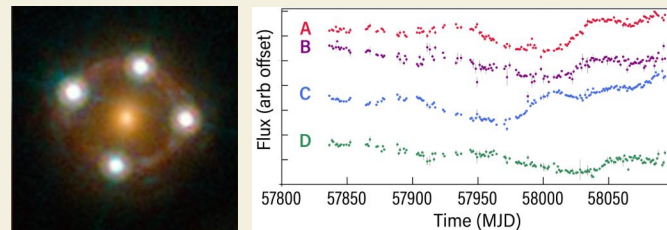
**Dark Matter (Sub)structure**  
Increase sample size by OOMs



High res ngVLA  
follow-up



**Time-delay Cosmography**  
100s of lensed time-variable sources

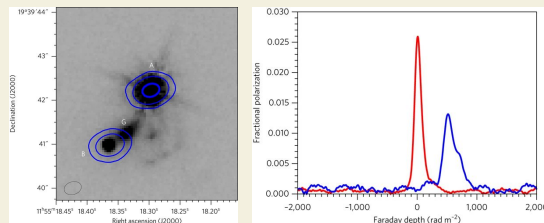


High cadence  
multiwavelength  
follow-up



DSA-2000 is a  
discovery engine

**High-redshift Universe**  
Thousands of lenses at  $z_{\text{source}} > 5$



Magnetic fields at cosmological  
distances and more

**Significant Survey Overlap**  
30,000 deg<sup>2</sup> footprint

