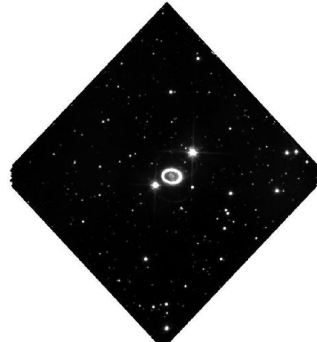


# Image-to-Text Retrieval

Image	Top classes (fine-tuned)	Top classes (base)	Abstract
	<ol style="list-style-type: none"><li>dwarf galaxies</li><li>RR Lyrae variables</li><li>red giants</li><li>trans-Neptunian objects</li></ol>	<ol style="list-style-type: none"><li>high-redshift quasars</li><li>gravitational lensing</li><li>low surface brightness galaxies</li><li>white dwarfs</li></ol>	isolated dwarf galaxies, WLM, Pegasus Dwarf Irregular Galaxy, stellar mass, main sequence stars; understand growth of stellar mass in WLM and PegDIG, constrain effects of reionization, analyze radial population gradients, study variable star populations in WLM and PegDIG
	<ol style="list-style-type: none"><li>supernova remnants</li><li>protostars</li><li>galactic structure</li><li>core-collapse supernova</li></ol>	<ol style="list-style-type: none"><li>gravitational lensing</li><li>supernovae</li><li>binary star systems</li><li>circumstellar disks</li></ol>	supernova SN 1987A, circumstellar ring, supernova remnant, shocked ring, radioactive isotopes; determine mass loss of Sanduleak -69 202 progenitor, test ideas for SN 1987A progenitor, analyze shape of SN 1987A explosion, resolve inner debris of SN 1987A, reveal true extent of gas outside the ring
	<ol style="list-style-type: none"><li>galaxy clusters</li><li>lyman alpha</li><li>intracluster medium</li><li>dark energy</li></ol>	<ol style="list-style-type: none"><li>gravitational lensing</li><li>high-redshift quasars</li><li>ultra diffuse galaxies</li><li>galaxy clusters</li></ol>	X-ray luminous galaxy clusters, eMACS clusters, Balmer Break Galaxies, Lyman-break galaxies, gravitational telescopes; identify most powerful gravitational telescopes for future studies, constrain mass distribution within extreme systems, improve understanding of galaxy-galaxy and galaxy-gas interactions, unveil Balmer Break Galaxies at $z \sim 2$ , detect Lyman-break galaxies at $z > 6$
	<ol style="list-style-type: none"><li>globular clusters</li><li>star clusters</li><li>galactic structure</li><li>crowded stellar field</li></ol>	<ol style="list-style-type: none"><li>star clusters</li><li>globular clusters</li><li>open clusters</li><li>stellar populations</li></ol>	pre-main sequence stars, Large Magellanic Cloud, young clusters, color-magnitude diagrams, main-sequence turn offs; disentangle between age and rotation in pre-MS stars, assess multiple bursts of star formation in young clusters, understand the eMSTO phenomenon, investigate the multiple stellar populations phenomenon, analyze the luminosity function of young clusters's Turn-On region