# Mixed Hybrid Finite Element Eddington Acceleration of Discrete Ordinates Source Iteration

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Mathematics and Computation

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## **Overview**

- 1. Motivation
- 2. Source Iteration Background
- 3. Eddington Acceleration
- 4. Results
- 5. Conclusions

# Motivation

#### **Motivation**

#### Radiation Hydrodynamics

- Describes the effects of emission, absorption, scattering on fluid momentum and energy
- Required in high energy density laboratory experiments (NIF, Z Machine) and astrophysics

Mixed Hybrid Finite Element Method (MHFEM) hydrodynamics

#### Problems

- MHFEM and first-order form of transport are incompatible ⇒ can't use linear acceleration scheme
- Radiation transport is expensive

#### Goal

Develop a transport algorithm that

- 1. Accelerates Discrete Ordinates Source Iteration
- Bridges Linear Discontinuous Galerkin (LDG) transport and MHFEM multiphysics

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# Source Iteration Background

## **Boltzmann Equation**

Steady-state, mono-energetic, istropically-scattering, fixed-source Linear Boltzmann Equation in 1D slab geometry:

$$\mu \frac{\partial \psi}{\partial x}(x,\mu) + \Sigma_t(x)\psi(x,\mu) = \frac{\Sigma_s(x)}{2} \int_{-1}^1 \psi(x,\mu')d\mu' + \frac{Q(x)}{2}$$

 $\mu=\cos\theta$  the cosine of the angle of flight  $\theta$  relative to the x-axis  $\Sigma_t(x), \ \Sigma_s(x) \ \text{total and scattering macroscopic cross sections}$   $Q(x) \ \text{the isotropic fixed-source}$   $\psi(x,\mu) \ \text{the angular flux}$ 

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#### Integro-differential equation

# Discrete Ordinates $(S_N)$ Angular Discretization

Compute angular flux on N discrete angles

$$\psi(x,\mu) \xrightarrow{\mathsf{S}_N} \begin{cases} \psi_1(x), & \mu = \mu_1 \\ \psi_2(x), & \mu = \mu_2 \\ \vdots \\ \psi_N, & \mu = \mu_N \end{cases}$$

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 $\mu_1$ ,  $\mu_2$ , ...,  $\mu_N$  defined by N-point Gauss Quadrature Rule

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 $\mu_1,\ \mu_2,\ \dots,\ \mu_N$  defined by N-point Gauss Quadrature Rule Integrate order 2N-1 polynomials exactly with

$$\phi(x) = \int_{-1}^{1} \psi(x, \mu) d\mu \xrightarrow{S_N} \sum_{n=1}^{N} w_n \psi_n(x)$$

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$$\mu_n \frac{\mathrm{d}\psi_n}{\mathrm{d}x}(x) + \Sigma_t(x)\psi_n(x) = \frac{\Sigma_s(x)}{2}\phi(x) + \frac{Q(x)}{2}, \ 1 \le n \le N$$
$$\phi(x) = \sum_{n=1}^N w_n \psi_n(x)$$

#### N coupled, ordinary differential equations

All coupling in scattering term

#### Source Iteration

Decouple by lagging scattering term

$$\mu_n \frac{\mathrm{d}\psi_n^{\ell+1}}{\mathrm{d}x}(x) + \Sigma_t(x)\psi_n^{\ell+1}(x) = \frac{\Sigma_s(x)}{2}\phi^{\ell}(x) + \frac{Q(x)}{2}, 1 \le n \le N$$

#### N independent, first-order, ordinary differential equations

Solve each equation with well-known sweeping process

#### **Source Iteration**

- 1. Given previous estimate for  $\phi^\ell(x),$  solve for  $\psi_n^{\ell+1}$
- 2. Compute  $\phi^{\ell+1}(x) = \sum_{n=1}^{N} w_n \psi_n^{\ell+1}(x)$
- 3. Update scattering term with  $\phi^{\ell+1}(x)$  and repeat until:

$$\frac{\|\phi^{\ell+1}(x)-\phi^{\ell}(x)\|}{\|\phi^{\ell+1}(x)\|}<\epsilon$$

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Slow to converge in optically thick systems with minimal losses to absorption and leakage

Radiation Hydrodynamics problems often contain highly diffusive regions

 $S_N$  is too expensive in these regions

Need an acceleration scheme that rapidly increases the rate of convergence of source iteration

**Eddington Acceleration** 

Zeroth Moment: integrate over all angles

$$\int_{-1}^{1} \mu \frac{d\psi}{dx}(x,\mu) d\mu + \int_{-1}^{1} \Sigma_{t}(x)\psi(x,\mu) d\mu = \int_{-1}^{1} \frac{\Sigma_{s}(x)}{2} \phi(x) d\mu + \int_{-1}^{1} \frac{Q(x)}{2} d\mu$$

#### Zeroth Moment: integrate over all angles

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Use 
$$J(x) = \int_{-1}^{1} \mu \psi(x, \mu) \, \mathrm{d}\mu$$
,  $\phi(x) = \int_{-1}^{1} \psi(x, \mu) \, \mathrm{d}\mu$ 

#### **Zeroth Angular Moment**

$$\frac{\mathrm{d}}{\mathrm{d}x}J(x) + \Sigma_a(x)\phi(x) = Q(x)$$

$$\int_{-1}^1 \mu^2 \frac{\mathrm{d} \psi}{\mathrm{d} x}(x,\mu) \, \mathrm{d} \mu + \int_{-1}^1 \mu \Sigma_t(x) \psi(x,\mu) \, \mathrm{d} \mu \ = \ \int_{-1}^1 \mu \frac{\Sigma_s(x)}{2} \phi(x) \, \mathrm{d} \mu + \int_{-1}^1 \mu \frac{Q(x)}{2} \, \mathrm{d} \mu$$

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Rearrange derivative

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Multiply and divide by  $\int_{-1}^{1} \psi(x,\mu) d\mu$ 

$$\frac{\mathrm{d}}{\mathrm{d}x} \int_{-1}^{1} \psi(x,\mu) \,\mathrm{d}\mu \frac{\int_{-1}^{1} \mu^{2} \psi(x,\mu) \,\mathrm{d}\mu}{\int_{-1}^{1} \psi(x,\mu) \,\mathrm{d}\mu}$$

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**Eddington Factor** 

$$\langle \mu^2 \rangle(x) = \frac{\int_{-1}^1 \mu^2 \psi(x, \mu) \,\mathrm{d}\mu}{\int_{-1}^1 \psi(x, \mu) \,\mathrm{d}\mu}$$

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Angular flux weighted average of  $\mu^2$ 

# **Moment Equations**

$$\frac{\mathrm{d}}{\mathrm{d}x}J(x)+\Sigma_a(x)\phi(x)=Q(x) \tag{Zeroth Moment}$$
 
$$\frac{\mathrm{d}}{\mathrm{d}x}\langle\mu^2\rangle(x)\phi(x)+\Sigma_t(x)J(x)=0 \tag{First Moment}$$

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Solving the Moment Equations requires knowledge of the angular flux (the solution)

#### **Eddington Acceleration**

Use  $\mathsf{S}_N$  to compute  $\langle \mu^2 \rangle(x)$  and Moment Equations to find  $\phi(x)$ 

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- 3. Solve the Moment Equations for  $\phi^{\ell+1}(x)$  using  $\langle \mu^2 \rangle^{\ell+1/2}(x)$
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#### Acceleration occurs because

- 1. Angular shape of the angular flux converges quickly  $\Rightarrow$  Eddington factor quickly converges
- Moment Equations model all scattering at once ⇒ dependence on source iterations to introduce scattering information is reduced

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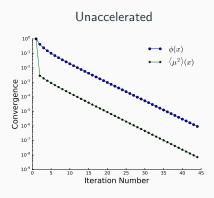
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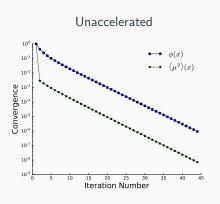
Difference between  $\mathsf{S}_N$  and Moment solutions can be used as a measure of mesh convergence

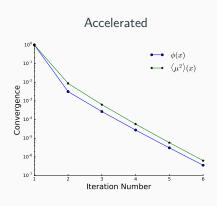
# Results

# **Convergence Rate Comparison**

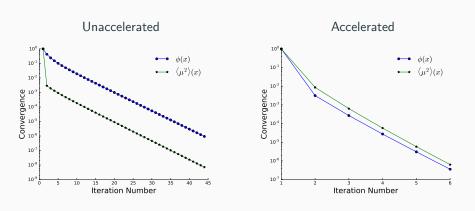


# **Convergence Rate Comparison**





## **Convergence Rate Comparison**



Fast rate of convergence of  $\langle \mu^2 \rangle(x)$  is transfered to  $\phi(x)$ 

Scale cross sections, source

$$\Sigma_t \to \Sigma_t/\epsilon$$

$$\Sigma_a \to \epsilon \Sigma_a$$

$$Q \to \epsilon Q$$

System becomes diffusive as  $\epsilon \to 0$ 

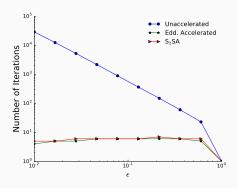
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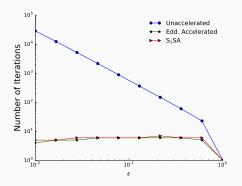
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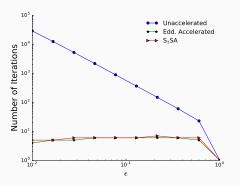
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Survives diffusion limit

Performs similarly to consistently differenced, linear acceleration ( $S_2SA$ )

# **Solution Convergence**

Compare

$$\frac{\|\phi_{\mathsf{S}_N}(x) - \phi_{\mathsf{Moment}}(x)\|}{\|\phi_{\mathsf{Moment}}(x)\|}$$

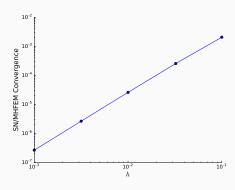
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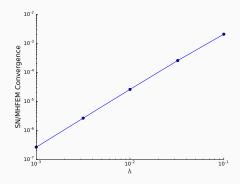


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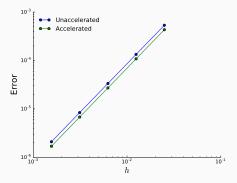
 $\mathsf{S}_N$  and Moment solutions converge as mesh is refined

Set Q(x) to force solution to

$$\phi(x) = \sin\left(\frac{\pi x}{x_b}\right)$$

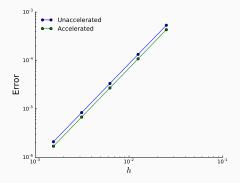
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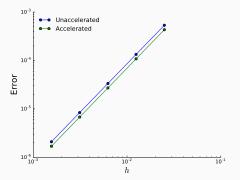
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Both second order accurate

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Both second order accurate

Eddington Acceleration did not effect the order of accuracy of lumped LDG

# Conclusions

#### Summary

#### Conclusions

- Scheme successfully accelerated source iteration in 1D slab geometry
- Eddington Acceleration is uniquely suited for radiation hydrodynamics
  - Transport and acceleration steps can be differenced with different methods
  - Reduces expense of source iteration
  - Provides inexpensive, conservative solution
- Showed MHFEM and lumped LDG can be paired

#### Summary

#### Conclusions

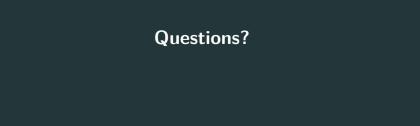
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#### Future Work

- Add temperature for radiative transfer
- Show still works in higher dimensions
- Develop an efficient rad hydro algorithm that makes use of the inexpensive Moment solution in multiphysics iterations

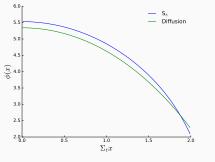
#### References

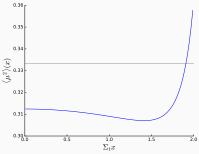
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# S<sub>8</sub> v. Diffusion

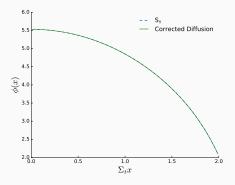
Small system  $\Rightarrow$  diffusion not expected to be accurate





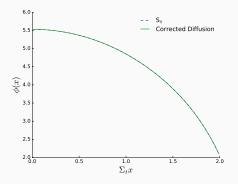
#### S<sub>8</sub> v. Drift Diffusion

Use  $\langle \mu^2 \rangle(x)$  from S<sub>8</sub> in Moment Equations



## S<sub>8</sub> v. Drift Diffusion

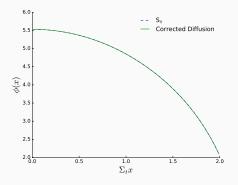
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Moment Equations and  $S_N$  match!

#### **S**<sub>8</sub> v. Drift Diffusion

Use  $\langle \mu^2 \rangle(x)$  from S<sub>8</sub> in Moment Equations



Moment Equations and  $S_N$  match!

Requires knowledge of angular flux