DERIVING THE OPEN AND CLOSED BOX EQUATIONS FOR OUR HODEL

WE KNOW FROM THE FOLLOWING ASSUMPTIONS:

. SINGLE BURST . SFR = E Mgos wid = 1, E = const. that

SFR = E Mgas $M_* = \int_{-\infty}^{\infty} dt SFR(t)$ $Z_{neb} = \frac{M(0)}{M(H)}$

. ASSUMING ALL METALS ARE MOSTLY IN STARS . AND GAS IS MOSTLY HYDROGEN

 $Z = \frac{M_2}{\Pi_{gos}}$ $M_2 = metals$ in Galaxy

CLOSED BOX

MZ = 4 SNR - Z SFR ASSUMING DULY TYPE II

YIELD DUE STAR FORMATION SAR = SFR

TO SN EXPL.

M_Z = [y-Z] SFR ⇒ M_Z + M_Z SFR = ySFR

⇒ MZ + MZ E = 4 SFR which is OF THE FORM

 $\dot{x} + Ax = B \implies x(t) = e^{-A(t)} \int_{t_0}^{t} B(t) e^{-A(t)} dt$

WITH $A(t) = \int_{t}^{t} \tilde{A}(t) dt = \varepsilon (t-t_0)$ IF $\varepsilon = const.$

THE FOLLOWING SOLUTION WILL THEN BE

 $M_{\underline{z}}(t) = e = \varepsilon(t-t_0) \int_{t_0}^{t} y \operatorname{SFR} e = \varepsilon(t-t_0) dt$

 $= e \qquad \begin{cases} \xi(t-t_0) & \xi(t-t_0) \\ \xi(t-t_0) & \xi(t-t_0) \end{cases} dt$

BUT FROM d Mg = -SER WE FOUND Mg(t) = Mg(to) e

$$M(t) = e \cdot (t - t_0) \int_{t_0}^{t} y \in M_g(t_0) e^{-\epsilon(t - t_0)} e^{-\epsilon(t - t_0)} dt$$

$$= e \cdot (t - t_0) M_0 y \in (t - t_0)$$

so
$$\Pi_{\xi}(t) = \Pi_{\xi}^{closed} = \Psi M_0 \quad \mathcal{E}(t-t_0) = \mathcal{E}(t-t_0)$$

WHICH TEANS

OPEN BOX

Gas presteen
$$Z_{16H} = 0$$
 (no H_2^{in})

Him = H^{out} \Rightarrow this way $H_g = ESFR$

we know solution
$$\chi(t) = e^{-\int a(t) dt}$$

$$now A(t) = \int_{t_0}^{t} a(t) dt = ? = (1+\eta) \varepsilon (t-t_0)$$

WHICH HAS SOLUTION (1+17) E = a OPEN OPEN = $d(t-t_0)$ $\begin{cases} t \\ t \end{cases}$ $d(t-t_0)$ $d(t-t_0)$ $d(t-t_0)$ $d(t-t_0)$ FROM OUR ASSUMPT. EMg = SFR $M_{\chi}^{OPEN} = e^{-\lambda(t-t_0)} \int_{t_0}^{t} \frac{1}{4} \frac{1}{\epsilon} \frac{1$ IF Y, E & t $M_{t}^{OPEN} = e^{-\lambda(t-t_{0})} \int_{t}^{t} e^{\epsilon \eta (t-t_{0})} dt \epsilon y M_{0}$ $= \frac{1}{\epsilon} \epsilon \eta \quad \text{Mo} \quad = \frac{\epsilon}{\epsilon} (1+\eta)(t-t_0) \quad = \frac{\epsilon}{\epsilon} \eta \quad (t-t_0) \quad + \quad (-...)$ $= \frac{1}{2} \int_{\eta} \frac{\varepsilon}{\varepsilon} \frac{1}{1} \int_{0}^{\eta} e^{-\varepsilon(t-t_{0})} = \frac{\varepsilon}{\eta} \left(\frac{1}{1-t_{0}} \right) \int_{0}^{\eta} \frac{1}{1-t_{0}} \int_{0}^{\eta} \frac{$ M2 /4(f-f0) $M_{\xi}^{\text{OPEN}} = \frac{\mathcal{E}}{\eta} \frac{1}{(t-t_{0})} M_{\xi}^{2}$ BIG RESULT HALING SAID THIS WE NOW CAN USE RETALLICITY, SINCE MOSS = MOSS $\frac{\varepsilon}{\eta} = \frac{1}{(t-t_0)} = \frac{1}{(t-t_0)} = \frac{1}{\varepsilon}$ DOESN'T DEPEND ON (t-to) NOT OF THE BALAXY to = time obs. = NOW SO Mo > Ma

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HAVING SAID THIS WE NOW HAVE THAT WE NEED TO EXPRESS THIS WTO
STELLAR MASS.
MA CON LIVETIMA EQ. WE TUST NEED THE CLOSED BOX ONE
SO WE CAN SAY ALREADY
   ZOPEN =
                   y n-1 x n-1 (M*)
   WE KNOW M_{*} = \int_{t_{0}}^{t} SFR dt = E M_{0} \int_{t_{0}}^{t} e^{-E(t-t_{0})} dt
                     = & Mo (- 1) (e- E(t-to) - 1)
                     = _ Mo (1-e - E(t-to))
So M* = Mo(1-e E(t-to)) WHICH MEANS, SINCE Z = 4 E(t-to)
= \frac{\mathcal{E}(t-t_0)}{\mathcal{H}_0} = 1 - e = 1 - \frac{\mathcal{E}(t-t_0)}{\mathcal{H}_0} = 1 - \frac{\mathcal{H}_{\star}}{\mathcal{H}_0}
    Q: HOW DO WE DEFINE NO?
                                            IF Mo >> MEAS
                                            · ξ CLDSED & 4 M*

Mo
  Q: ASSUMPTIONS CORRECT ?
  \chi of \epsilon N = 4 \eta^{-1} (M*) \propto \eta^{-1} (M*)
                                             M & Mx FROM DISPEASE
                                                 MALL MASSES DAVE ANAUTIC
               941 2
                                            4 ~ 0.01 (TREPLONT ET AL.)
                                -4-
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