



# Image Formation

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**Image:** Projection of 3D scene onto 2D plane. We need to understand the geometric and photometric relation between the scene and its image.

## Topics:

- (1) Pinhole and Perspective Projection
- (2) Image formation using Lenses
- (3) Lens Related Issues
- (4) Wide Angle Cameras



# Image Formation



Screen

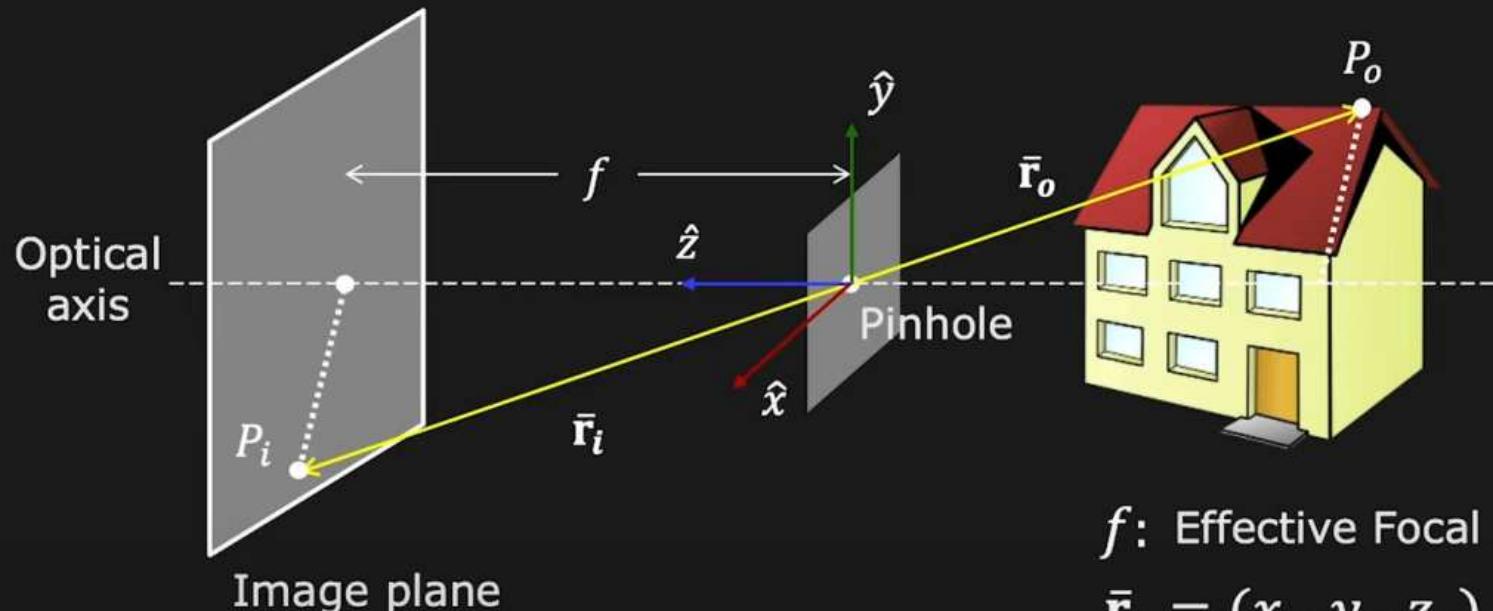


Scene

Is an image being formed on the screen?



# Perspective Imaging with Pinhole



$f$ : Effective Focal Length

$$\bar{r}_o = (x_o, y_o, z_o)$$

$$\bar{r}_i = (x_i, y_i, f)$$

Using similar triangles:

$$\frac{\bar{r}_i}{f} = \frac{\bar{r}_o}{z_o}$$

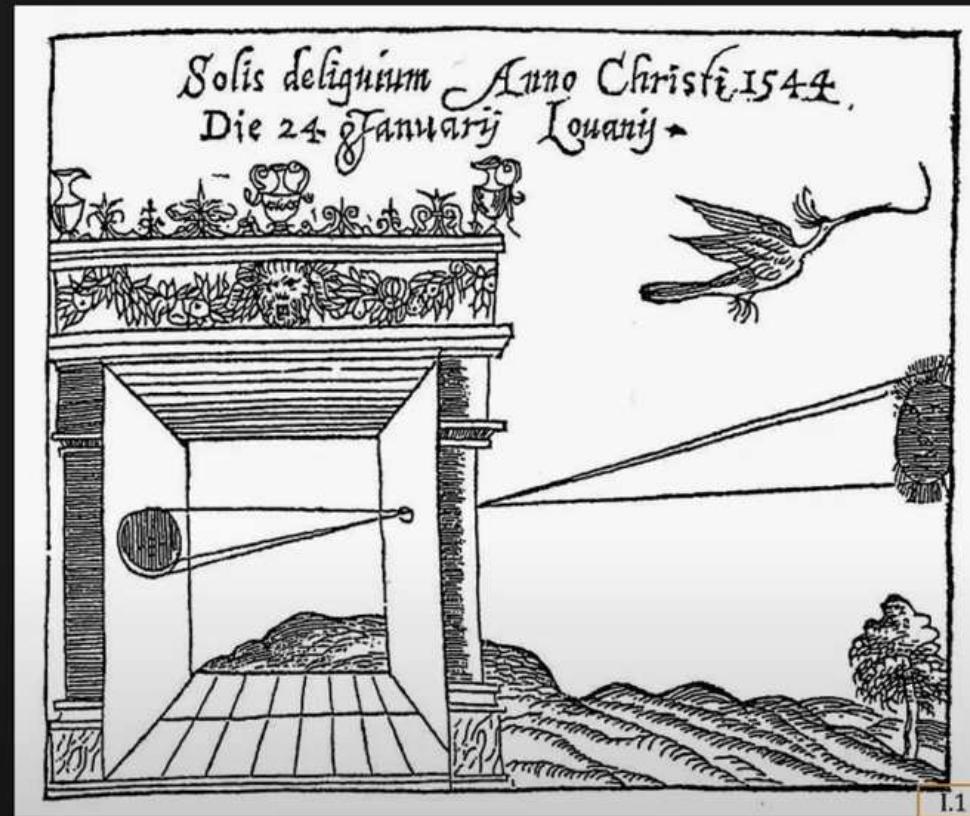
→

$$\frac{x_i}{f} = \frac{x_o}{z_o}, \quad \frac{y_i}{f} = \frac{y_o}{z_o}$$





# Camera Obscura



"Dark Chamber"

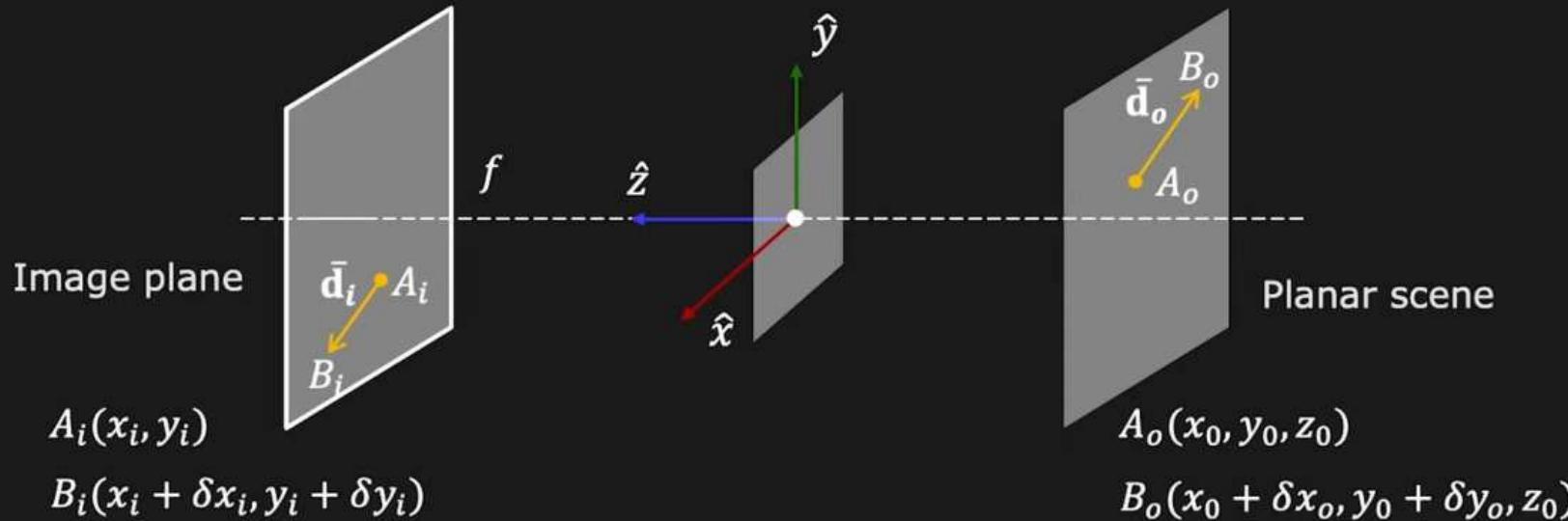




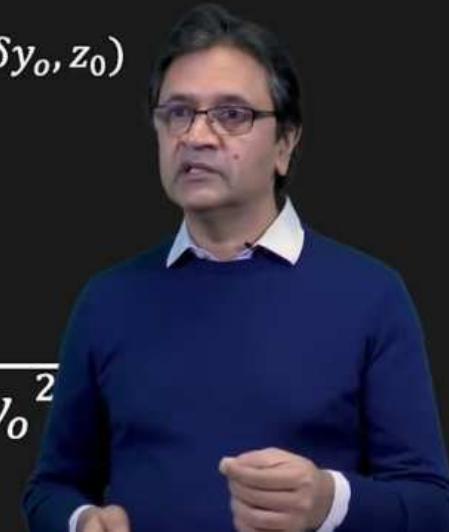
# Pinhole Eye of *Nautilus pompilius*



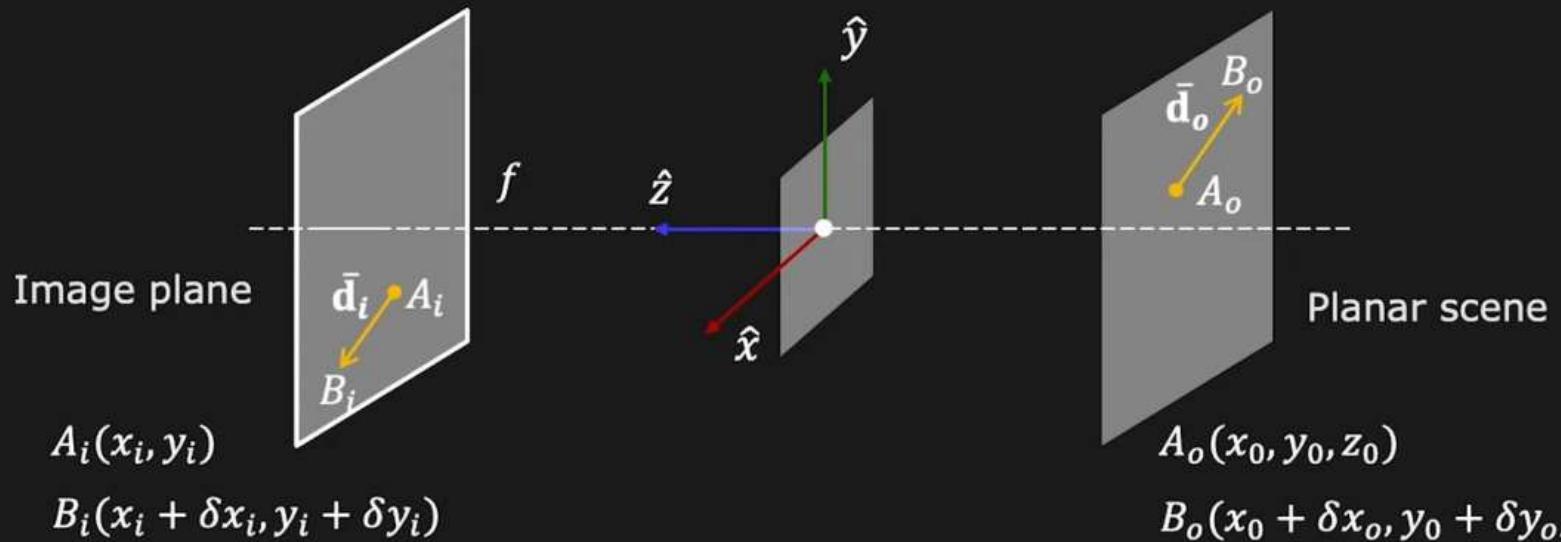
# Image Magnification



$$\text{Magnification: } |m| = \frac{\|\bar{d}_i\|}{\|\bar{d}_o\|} = \sqrt{\delta x_i^2 + \delta y_i^2} / \sqrt{\delta x_o^2 + \delta y_o^2}$$



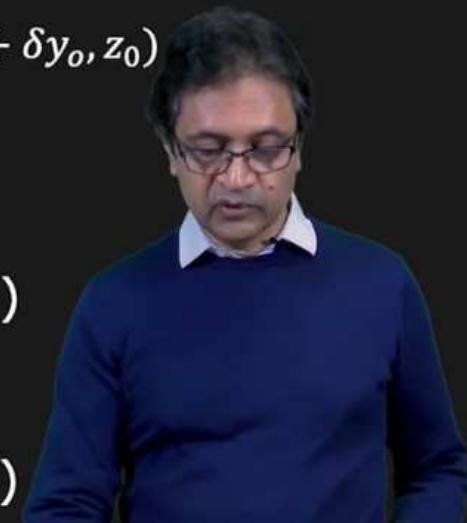
# Image Magnification



From Perspective Projection:

$$\frac{x_i}{f} = \frac{x_o}{z_o} \quad \text{and} \quad \frac{y_i}{f} = \frac{y_o}{z_o} \quad \dots \quad (\text{A})$$

$$\frac{x_i + \delta x_i}{f} = \frac{x_o + \delta x_o}{z} \quad \text{and} \quad \frac{y_i + \delta y_i}{f} = \frac{y_o + \delta y_o}{z} \quad \dots \quad (\text{B})$$



# Image Magnification

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From (A) and (B) we get:

$$\frac{\delta x_i}{f} = \frac{\delta x_o}{z_o} \quad \text{and} \quad \frac{\delta y_i}{f} = \frac{\delta y_o}{z_o}$$

Magnification:

$$|m| = \frac{\|\bar{\mathbf{d}}_i\|}{\|\bar{\mathbf{d}}_o\|} = \sqrt{{\delta x_i}^2 + {\delta y_i}^2} / \sqrt{{\delta x_o}^2 + {\delta y_o}^2} = \left| \frac{f}{z_o} \right|$$



# Image Magnification

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1.4

$$m = \frac{f}{z_o}$$

Image size **inversely proportional** to depth



# Image Magnification

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$$m = \frac{f}{z_o}$$

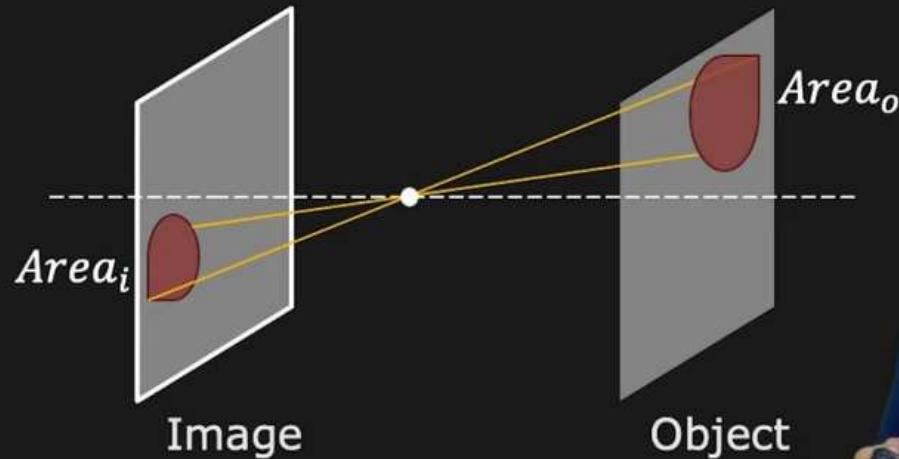
Image size **inversely proportional** to depth



# Image Magnification

Remarks:

- $m$  can be assumed to be constant if the range of scene depth  $\Delta z$  is much smaller than the average scene depth  $\tilde{z}$
- $$\frac{Area_i}{Area_o} = m^2$$



# Vanishing Point

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Parallel straight lines converge at a single image point



# Vanishing Point

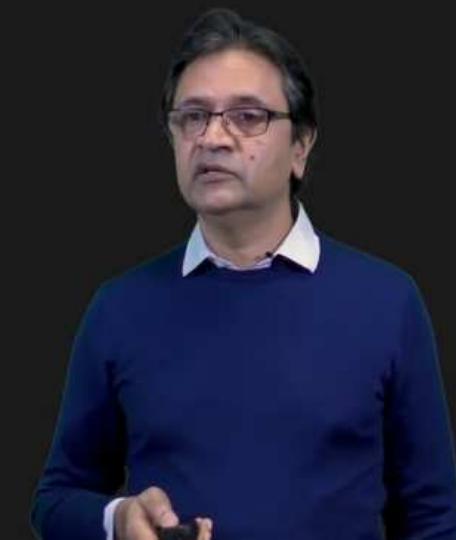
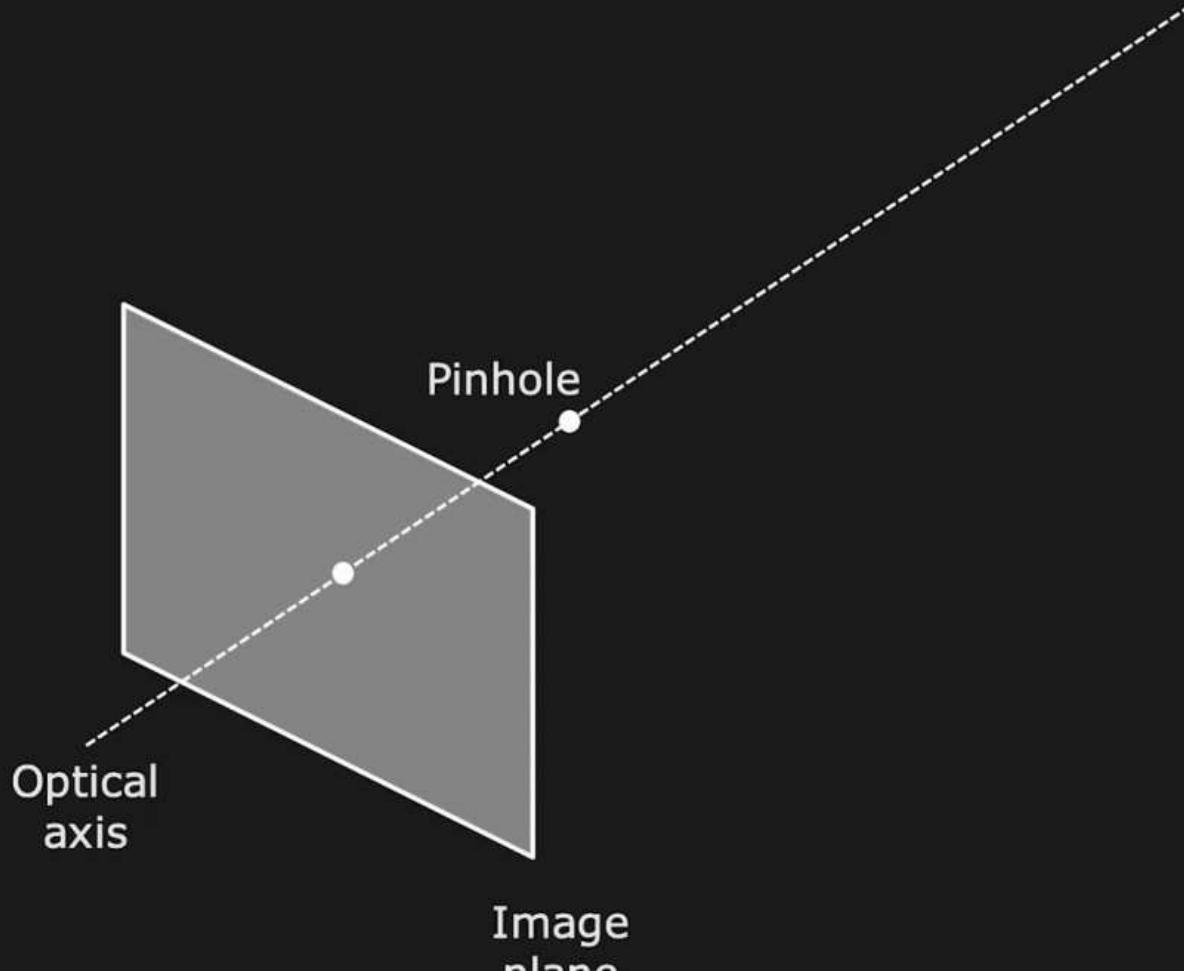


Location of Vanishing Point depends on the orientation of parallel straight lines.

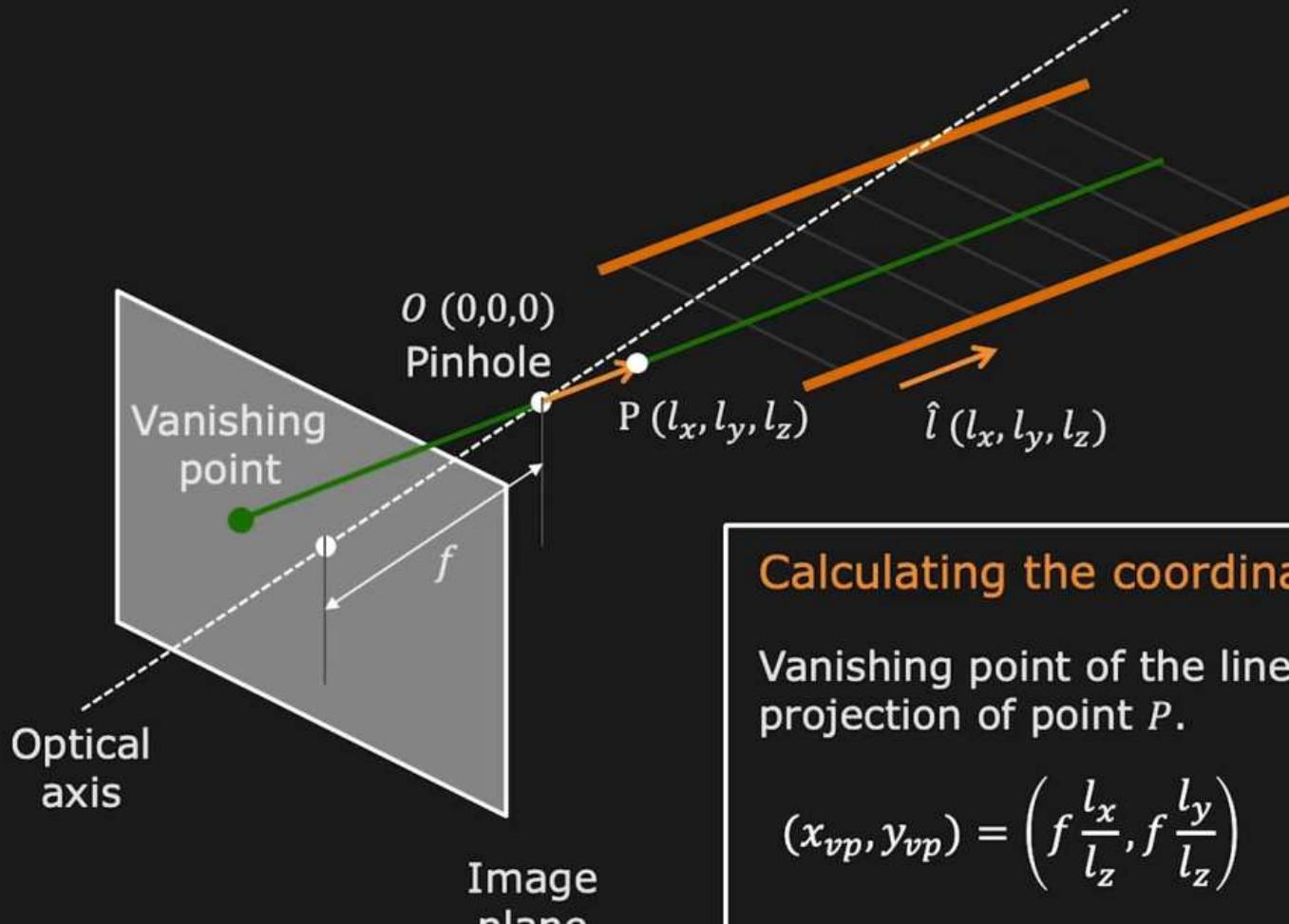


# Finding the Vanishing Point

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# Finding Vanishing Point



Calculating the coordinates:

Vanishing point of the line is the projection of point  $P$ .

$$(x_{vp}, y_{vp}) = \left( f \frac{l_x}{l_z}, f \frac{l_y}{l_z} \right)$$



# Use of Vanishing Point in Art

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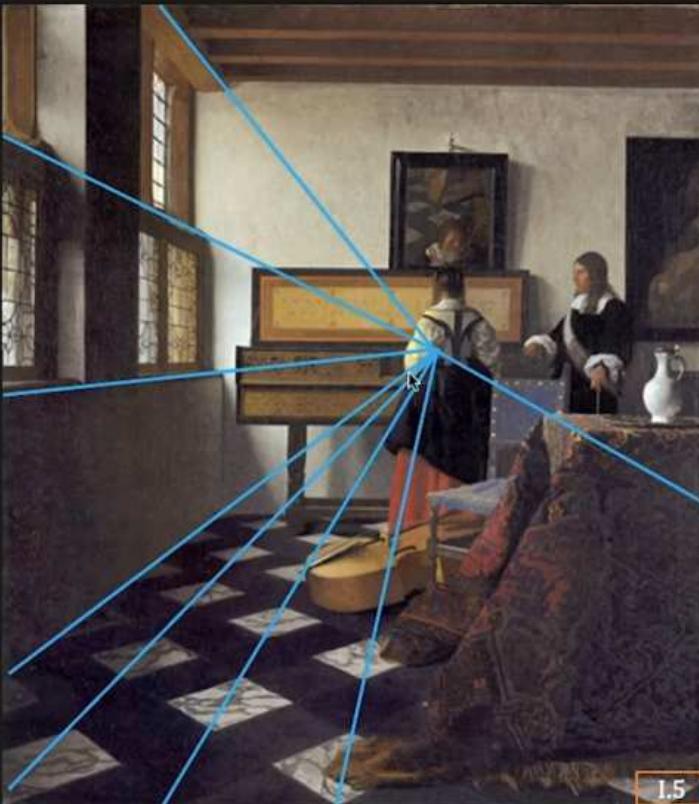


*The Music Lesson*, Johannes Vermeer, c. 1662-1664



# Use of Vanishing Point in Art

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*The Music Lesson*, Johannes Vermeer, c. 1662-1664



# False Perspective



Depth appears to be ~155 feet



Depth is actually ~30 feet



Galleria Spada, Francesco Borromini, 1652

# What is the Ideal Pinhole Size?



2 mm

1 mm

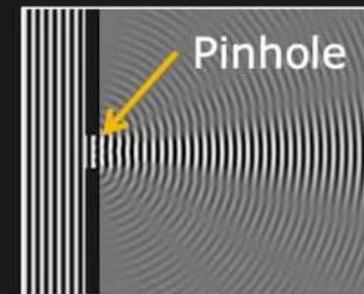
0.6 mm

0.35 mm

0.15 mm

0.07 mm  
I.8

The pinhole must be **tiny**,  
but if it's too tiny it will cause **diffraction**.

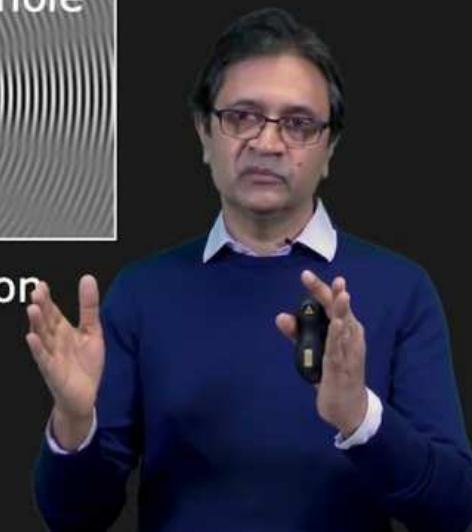


Diffraction

Ideal pinhole diameter:

$$d \approx 2\sqrt{f\lambda}$$

$f$ : effective focal length  
 $\lambda$ : wavelength



# What about Exposure Time?

Pinholes pass less light and hence require **long exposures** to capture bright images.



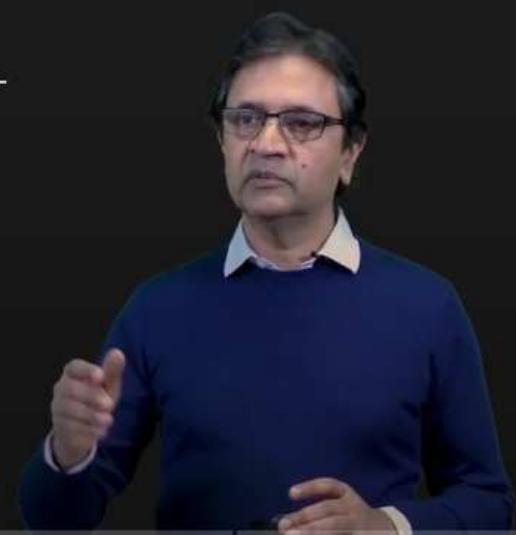
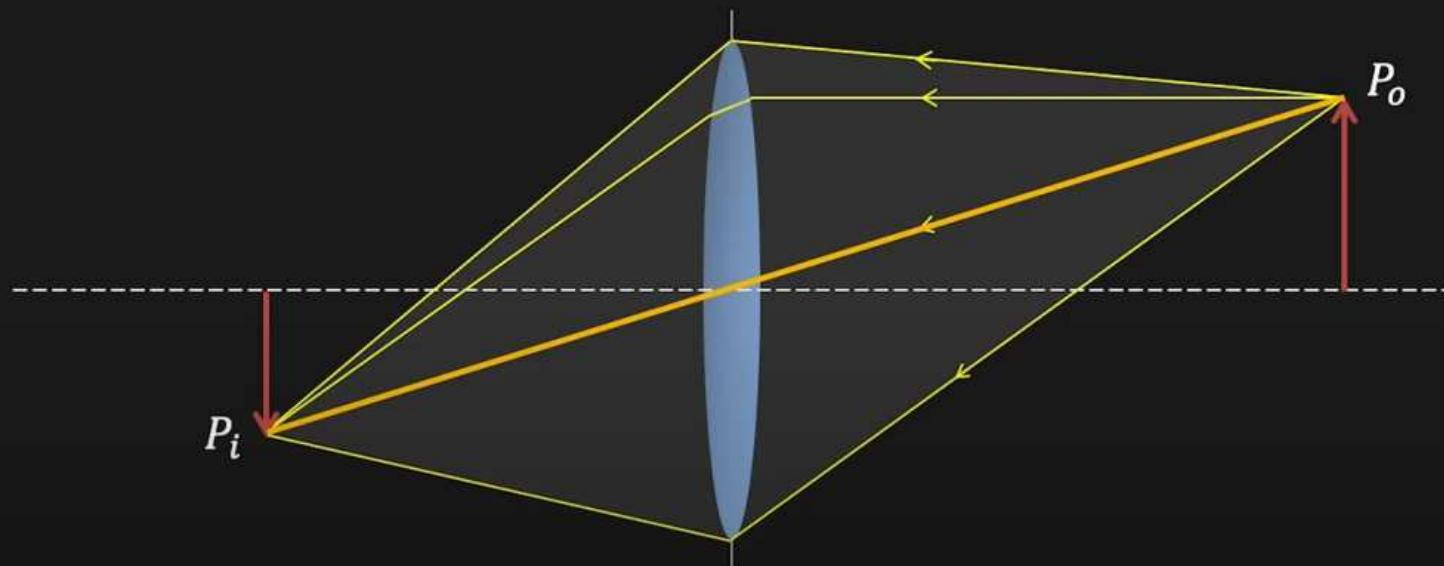
$f = 73 \text{ mm}$ ,  $d = 0.2 \text{ mm}$ ,  
Exposure,  $T = 12 \text{ s}$





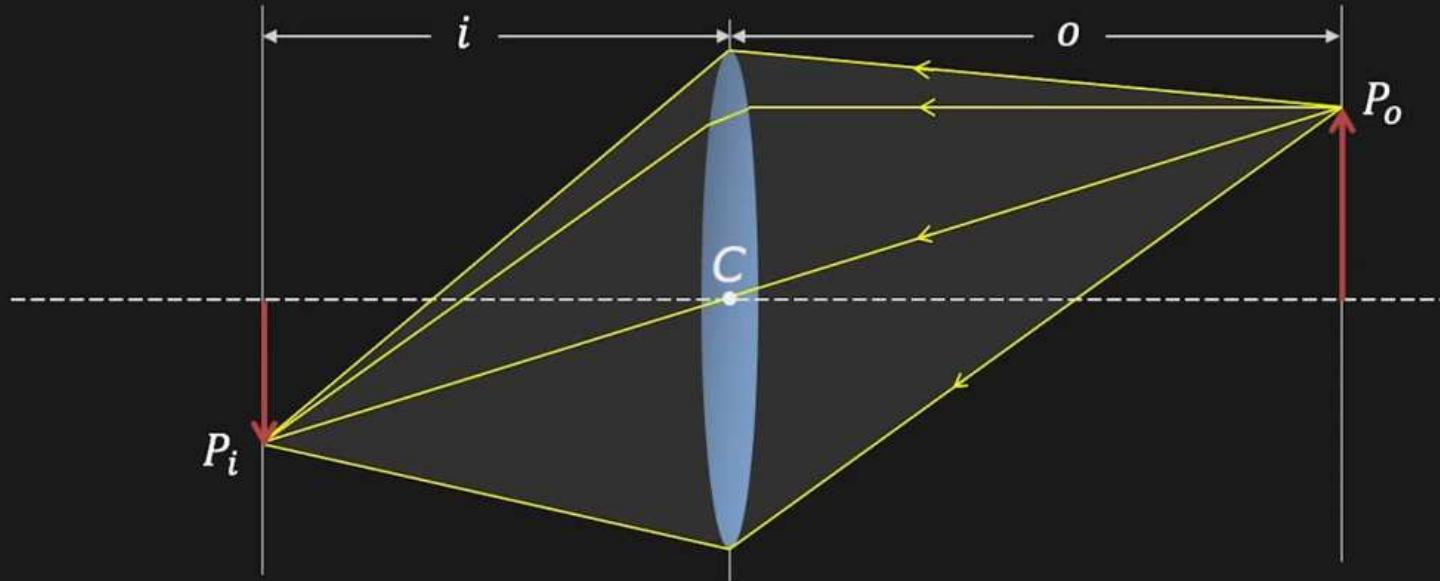
# Lenses

Same projection as pinhole, but gather more light!



Focal length ( $f$ ) determines the lens' bending power

# Gaussian Lens (Thin Lens) Law



$f$ : focal length

$i$ : image distance

$o$ : object distance

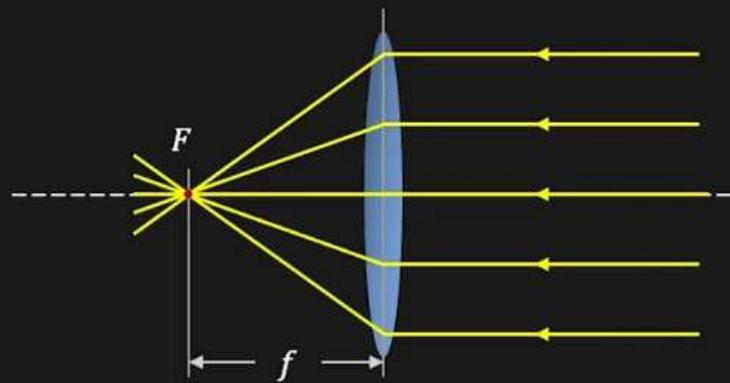
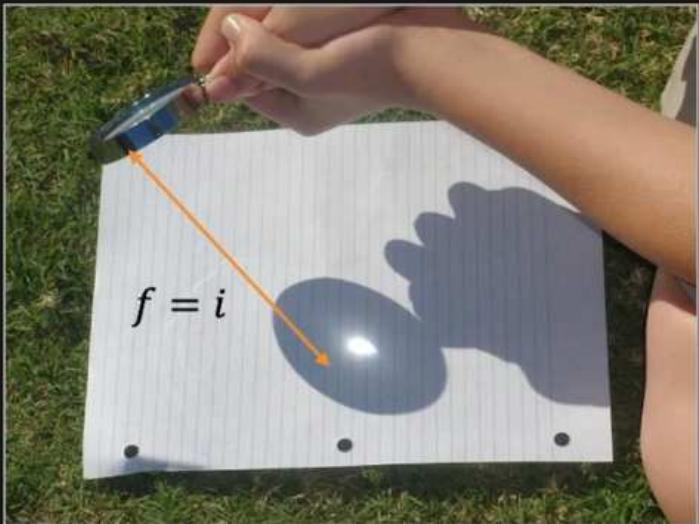
$$\frac{1}{i} + \frac{1}{o} = \frac{1}{f}$$

Example: If  $f = 50\text{mm}$  &  $o = 300\text{mm}$ , then image distance  $i = 60\text{mm}$



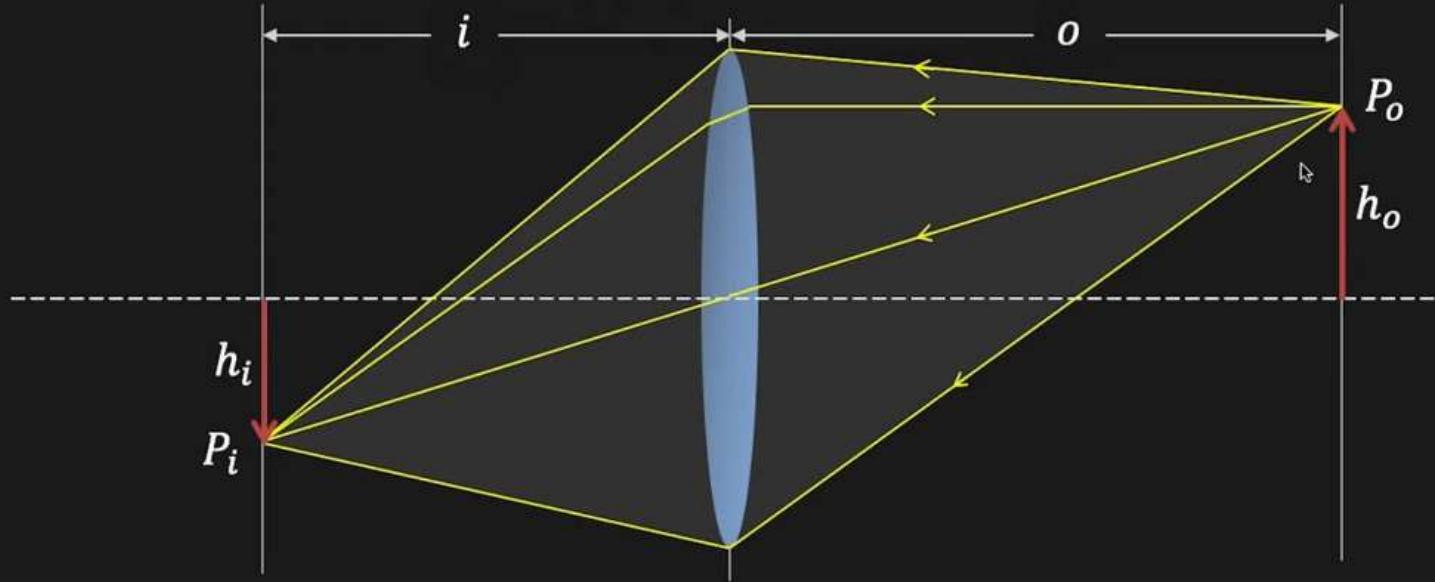
# How to Find the Focal Length?

$$\frac{1}{i} + \frac{1}{o} = \frac{1}{f} \quad \Rightarrow \quad \text{If } o = \infty, \text{ then } f = i$$



Focal length: Distance at which incoming rays that are parallel to the optical axis converge.

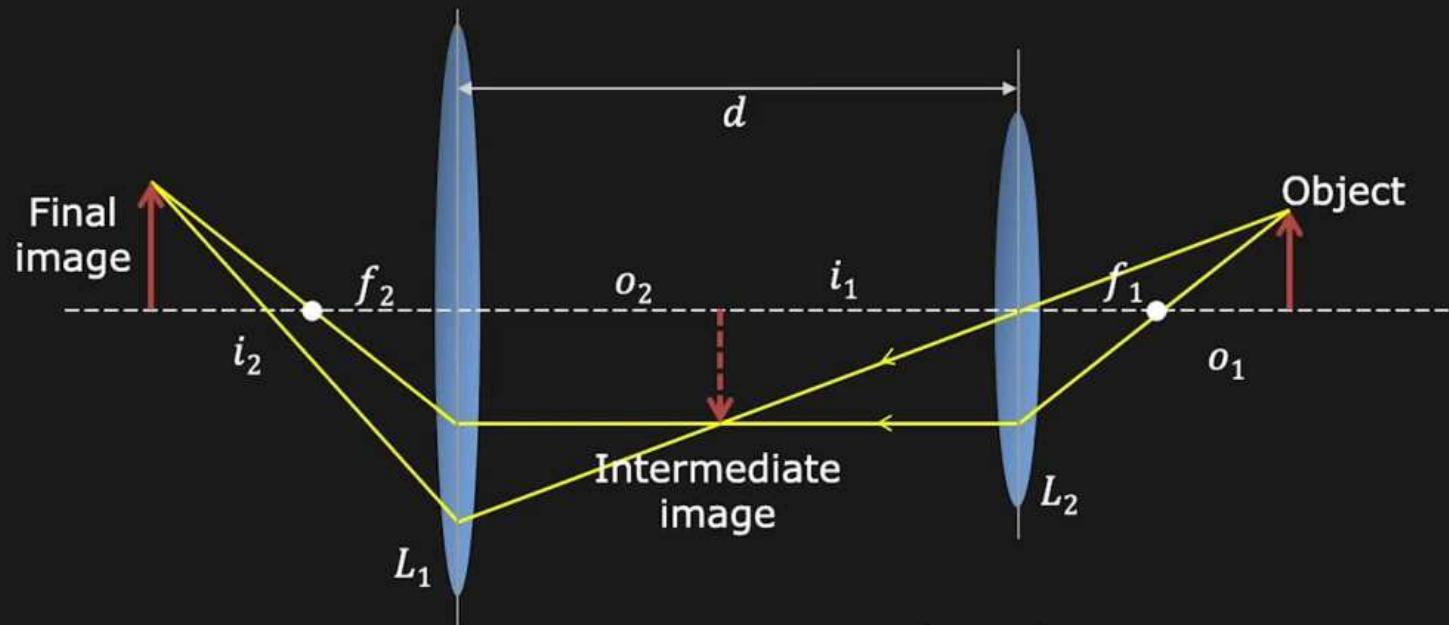
# Image Magnification



$$\text{Magnification: } m = \frac{h_i}{h_o}$$

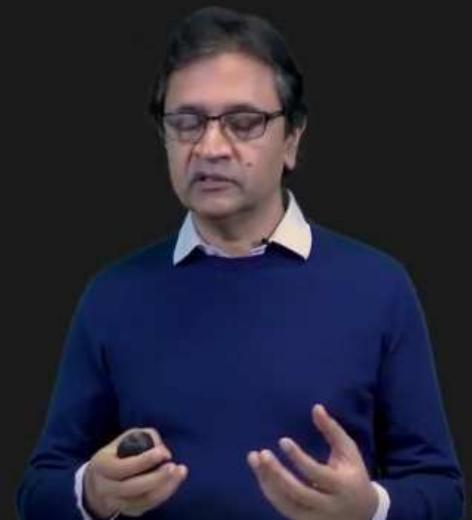


# Two Lens System



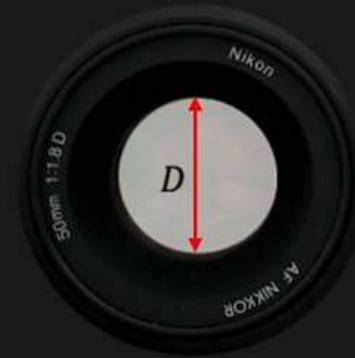
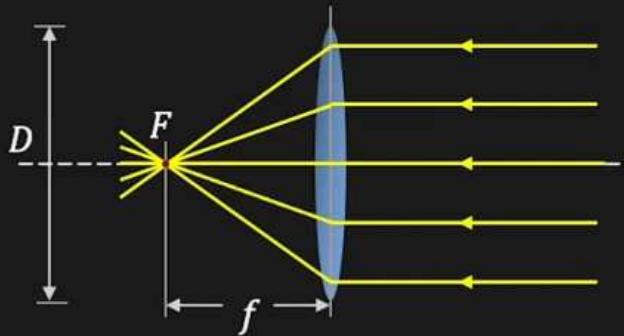
$$\text{Magnification: } m = \frac{i_2}{o_2} \cdot \frac{i_1}{o_1}$$

**Zooming:** Move lenses to change magnification



# Aperture of Lens

Light receiving area of lens, indicated by lens diameter.



Aperture can be reduced/increased to control image brightness



# f-number (f-stop, f-ratio) of Lens

Convenient to represent aperture as a fraction of focal length

$$\text{Aperture: } D = f/N$$

$$\text{f-Number: } N = f/D$$

where **N** is called the **f-Number** of lens.

Ex: A 50mm focal length, f/1.8 lens implies:

$N = 1.8$  ( $D = 27.8\text{mm}$ ) when aperture is fully open



$N = 1.8$



$N = 4$



$N = 8$



$N = 11$



# Blur Circle (Defocus)

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Focused Point

$$\frac{1}{i} + \frac{1}{o} = \frac{1}{f}$$

$$i = \frac{of}{o - f}$$

Defocused Point

$$\frac{1}{i'} + \frac{1}{o'} = \frac{1}{f}$$

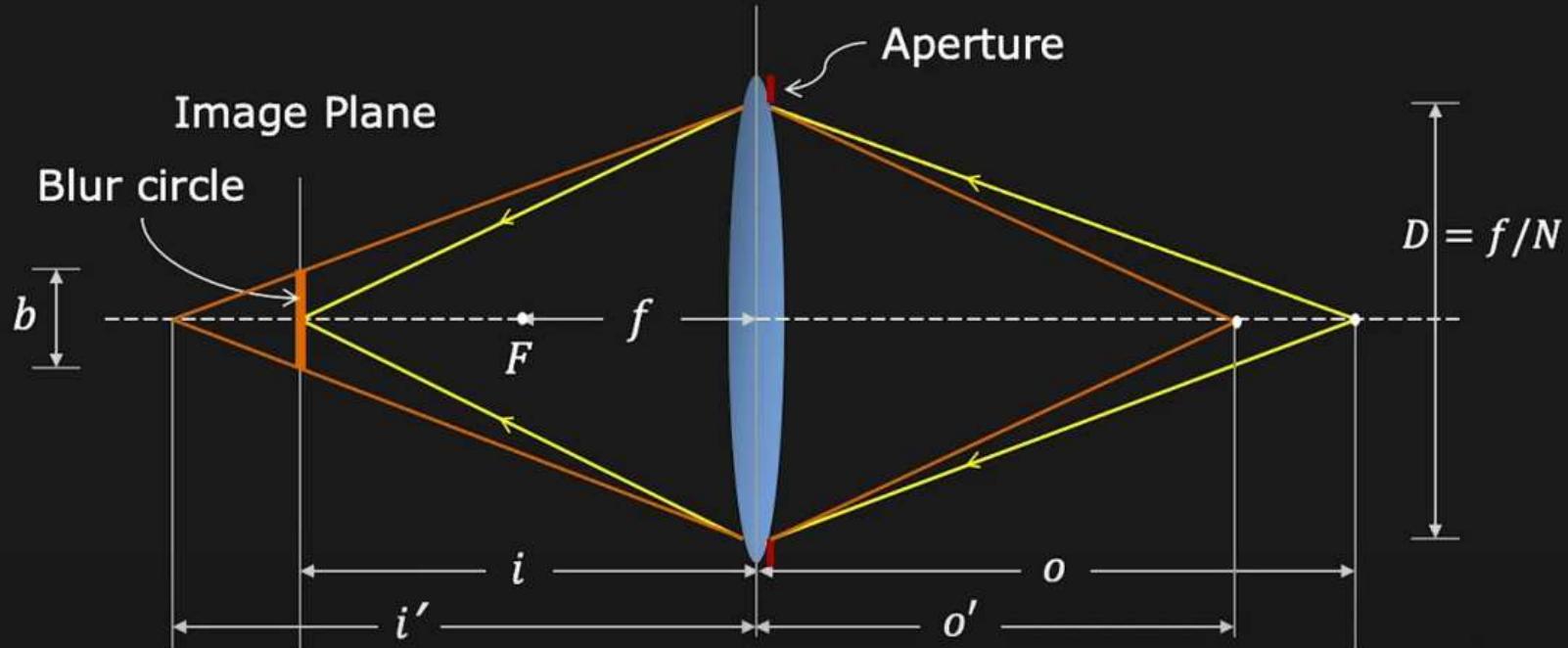
(Gaussian Lens Law)

$$i' = \frac{o'f}{o' - f}$$





# Lens Defocus



From similar triangles:

$$\frac{b}{D} = \frac{|i' - i|}{i'}$$

Blur circle diameter:

$$b = \frac{D}{i'} |i' - i|$$

$$b \propto D \propto \frac{1}{N}$$



# Blur Circle (Defocus)

---

Focused Point

$$\frac{1}{i} + \frac{1}{o} = \frac{1}{f}$$

$$i = \frac{of}{o - f}$$

Defocused Point

$$\frac{1}{i'} + \frac{1}{o'} = \frac{1}{f}$$

(Gaussian Lens Law)

$$i' = \frac{o'f}{o' - f}$$





# Blur Circle (Defocus)

Focused Point

$$\frac{1}{i} + \frac{1}{o} = \frac{1}{f}$$

$$i = \frac{of}{o-f}$$

Defocused Point

$$\frac{1}{i'} + \frac{1}{o'} = \frac{1}{f}$$

(Gaussian Lens Law)

$$i' = \frac{o'f}{o'-f}$$

$$i' - i = \frac{f}{(o'-f)} \cdot \frac{f}{(o-f)} \cdot (o - o')$$

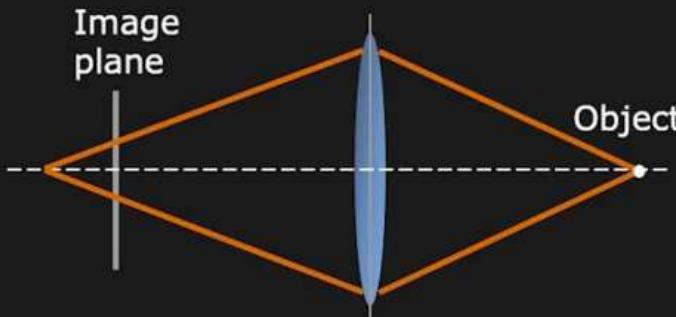
$$b = Df \left| \frac{(o - o')}{o'(o - f)} \right|$$

$$b = \frac{f^2}{N} \left| \frac{(o - o')}{o'(o - f)} \right|$$

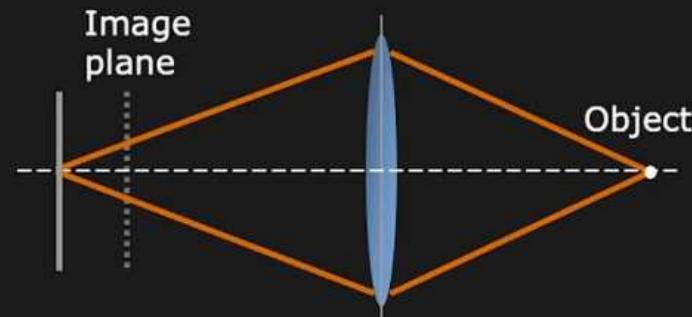




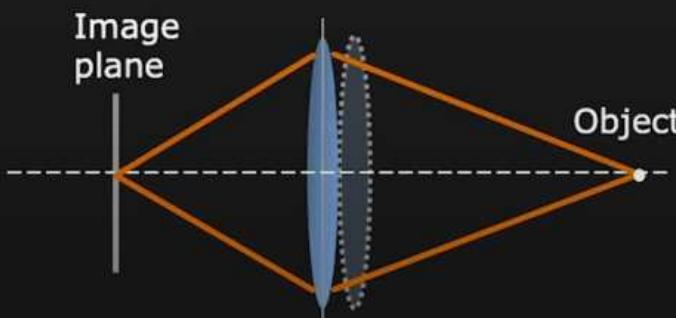
# Focusing



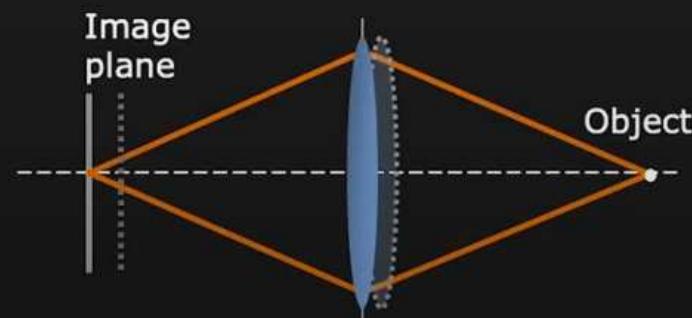
Defocused System



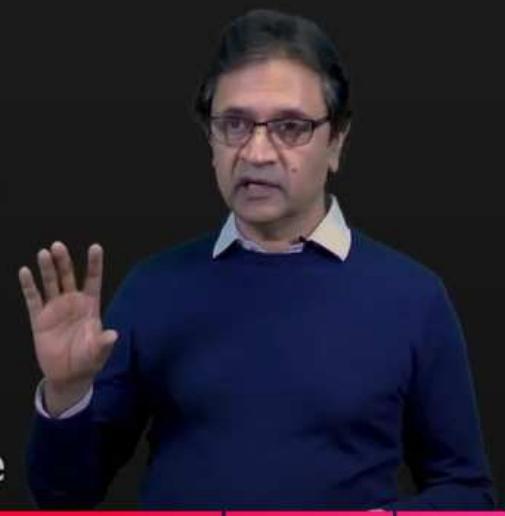
Move the image plane



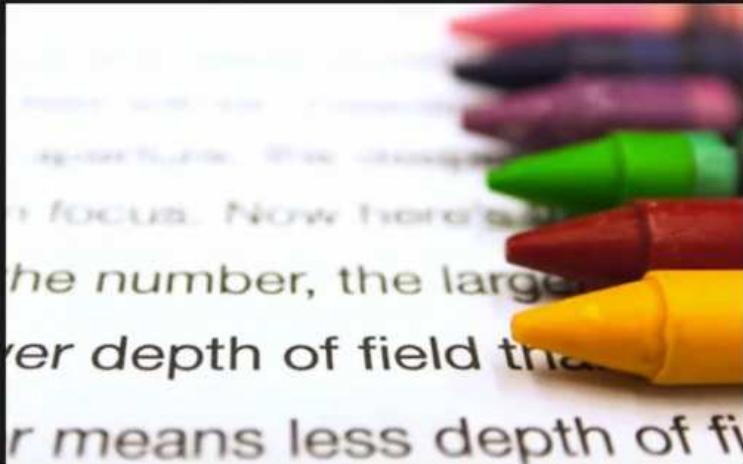
Move the lens



Move both lens and image plane



# Depth of Field (DoF)

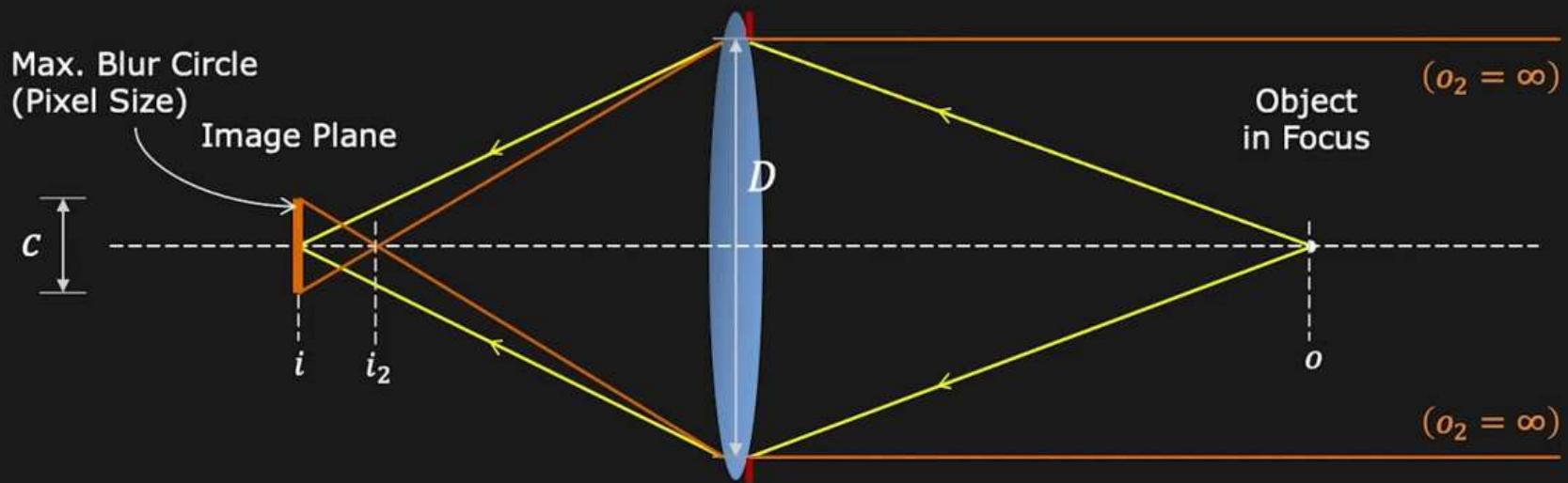


I.10

Range of object distances over which the image is “sufficiently well” focused, i.e., range over which blur  $b$  is less than pixel size.



# Hyperfocal Distance



The closest distance  $o = h$  the lens must be focused to keep objects at infinity ( $o_2 = \infty$ ) acceptably sharp (blur circle  $\leq c$ ).

Hyperfocal Distance:

$$h = \frac{f^2}{Nc} + f$$



# Aperture Size: DOF vs. Brightness



# Aperture Size: DOF vs. Brightness

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Focal Length 50 mm, Focus = 1m, Aperture D = 3.125mm, f-Number N = 16



# Aperture Size: DOF vs. Brightness

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## Large Aperture (Small f-Number)

- Bright Image or Short Exposure Time
- Shallow Depth of Field

## Small Aperture (Large f-Number)

- Dark Image or Long Exposure Time
- Large Depth of Field



# Tissue Box Camera

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Camera



Image





# Blocking the Lens



Camera

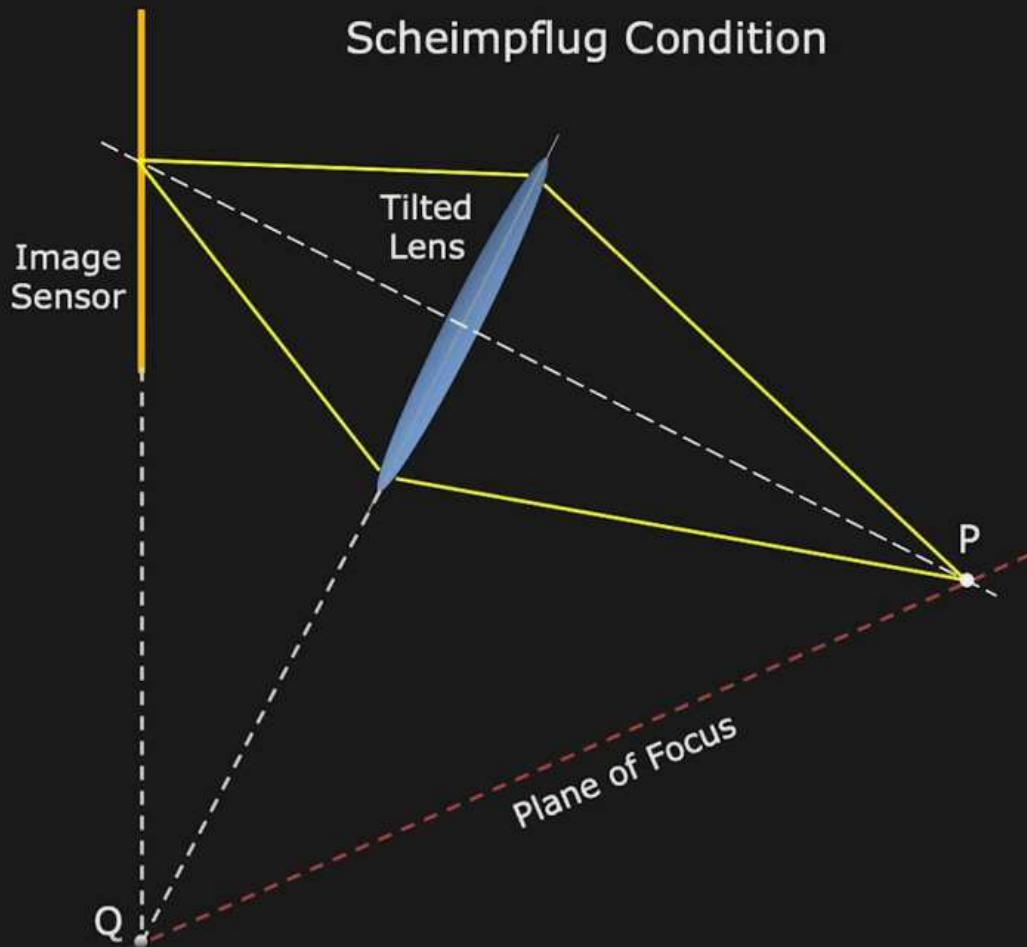


Image

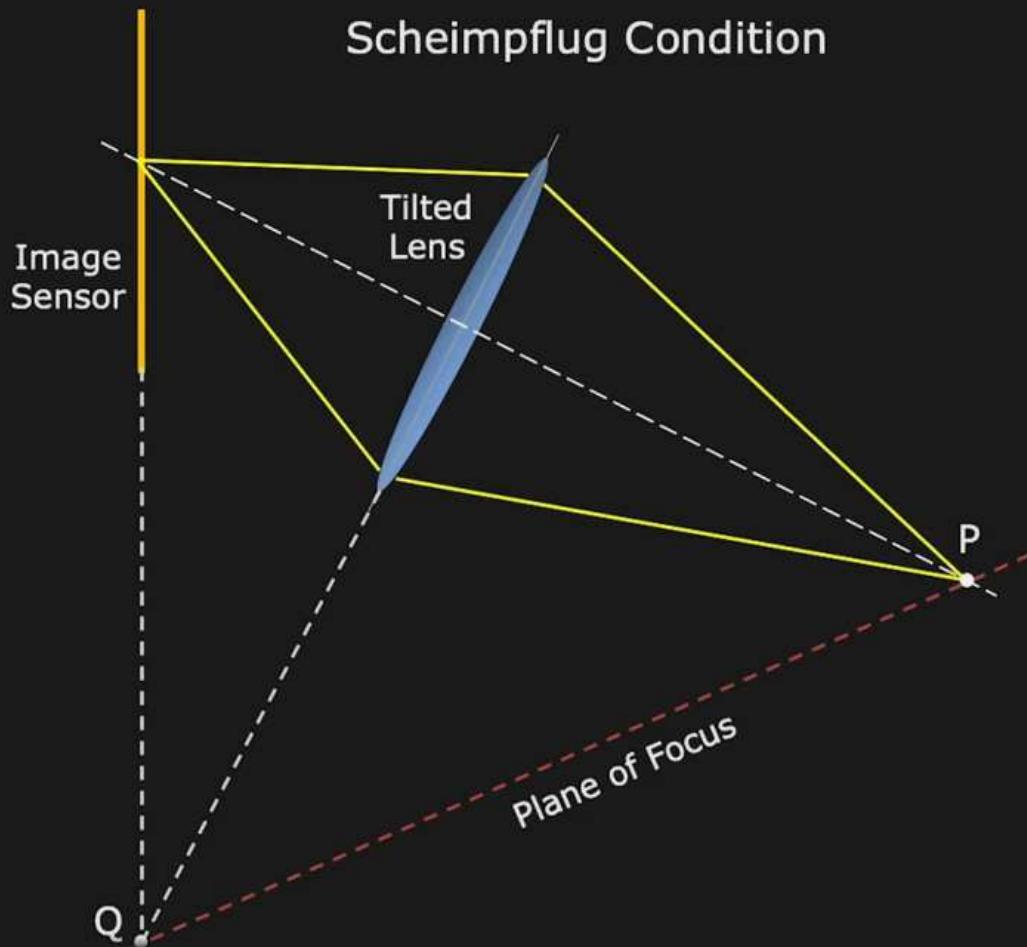


# Tilting the Lens

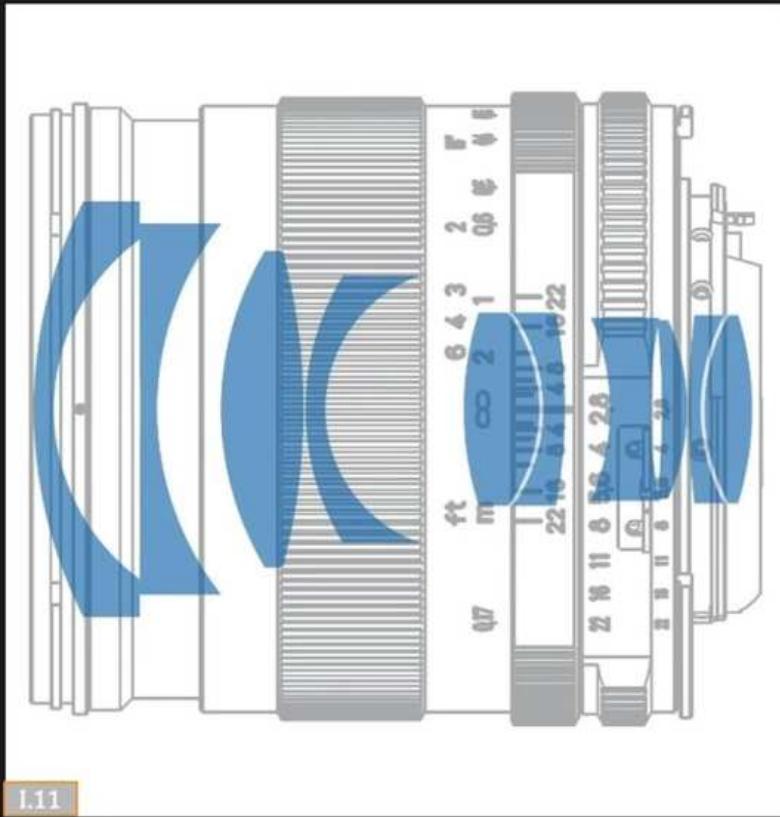
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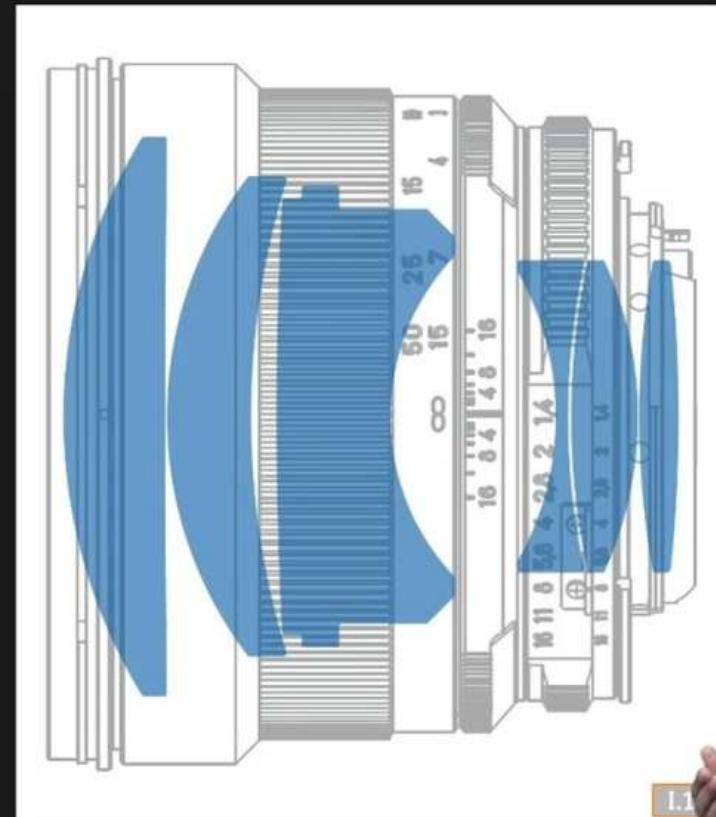
# Tilting the Lens



# Compound Lenses



Zeiss 25mm F2.8

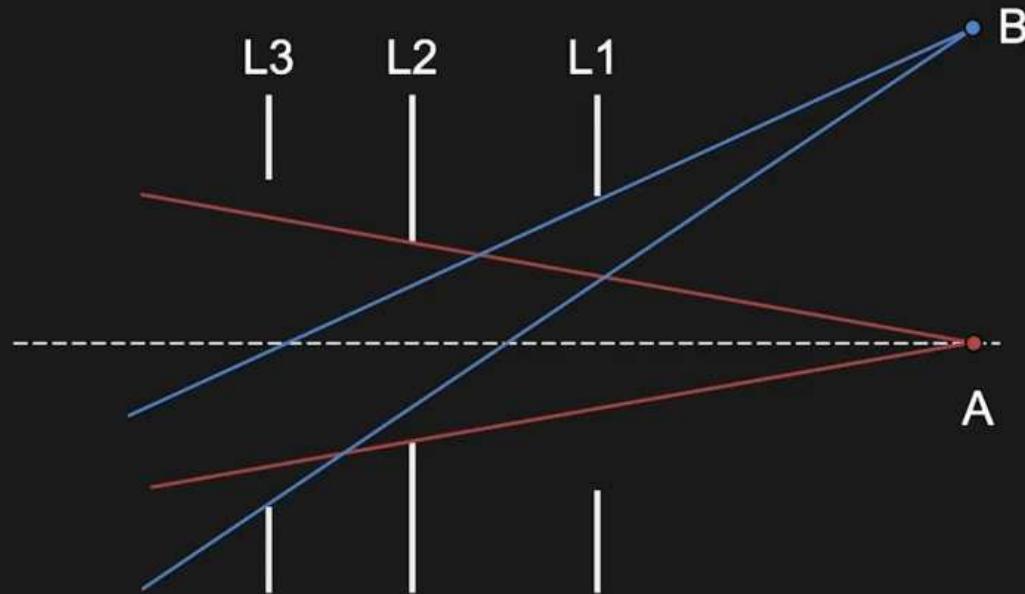


Zeiss 85mm F1.4



# Vignetting

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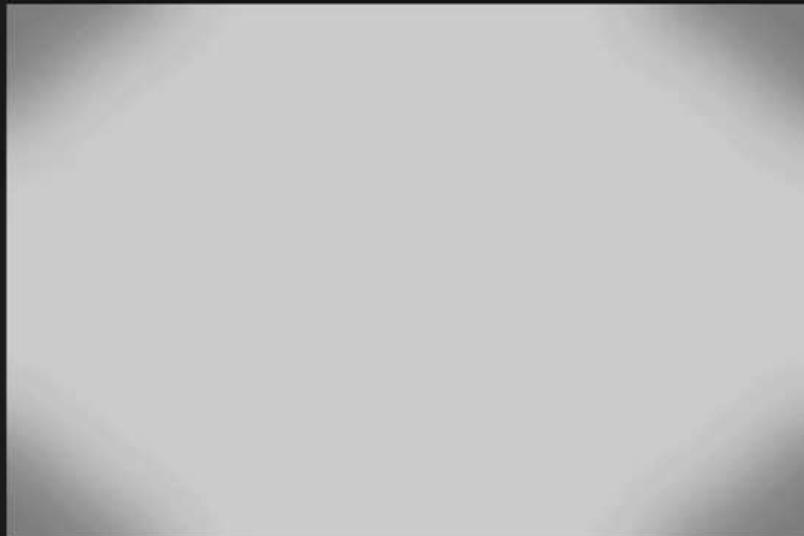


More light passes through L3 from point A than point B.  
Results in a smooth fall-off in brightness from A to B.



# Vignetting

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Brightness fall-off (Vignetting)  
in image of a White Wall

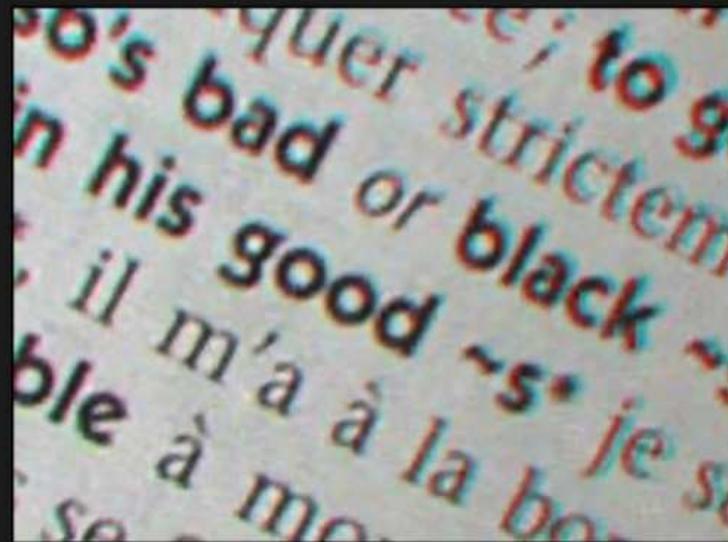
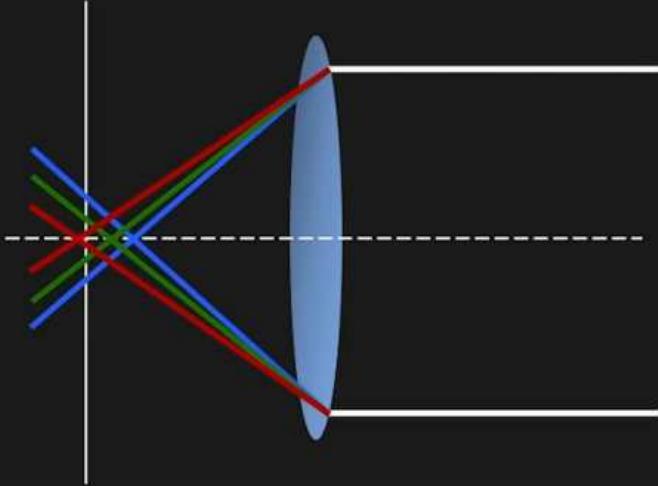


Brightness fall-off (Vignetting)  
in image of a Natural Scene



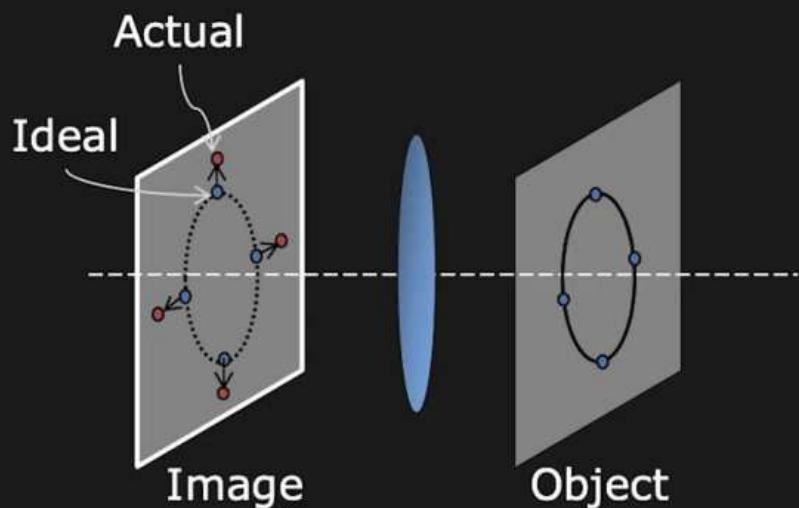
# Chromatic Aberration

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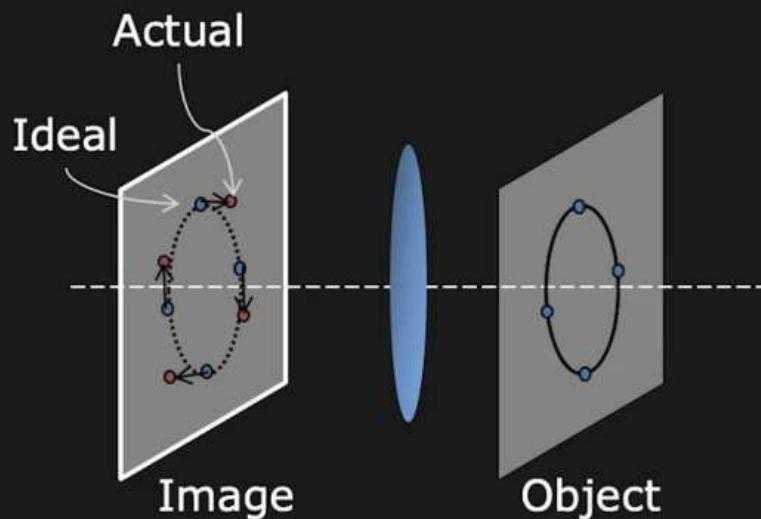


Refractive index (and hence focal length) of lens is different for different wavelengths.

# Geometric Distortion



Radial distortion



Tangential distortion

Due to lens imperfections



# Geometric Distortion Correction



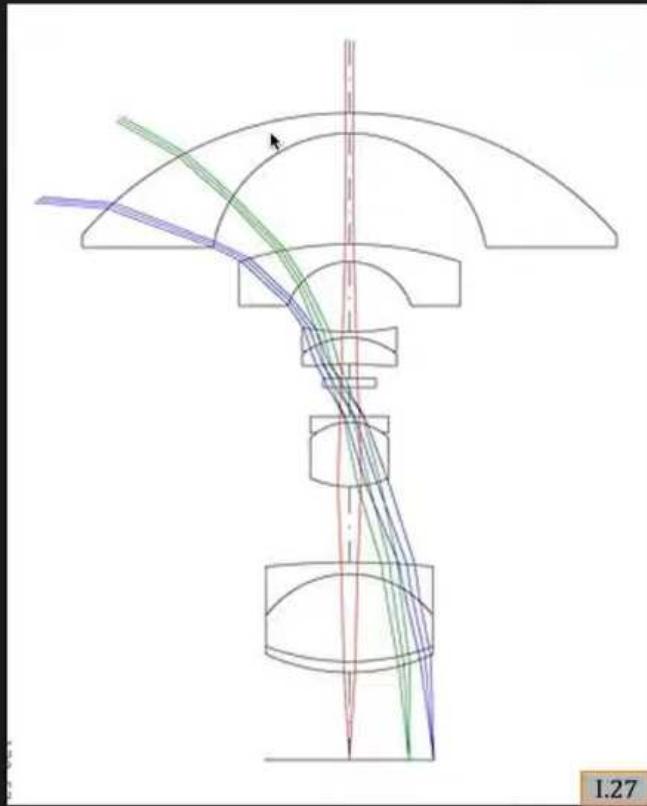
Radial (Barrel) distortion

Undistorted image



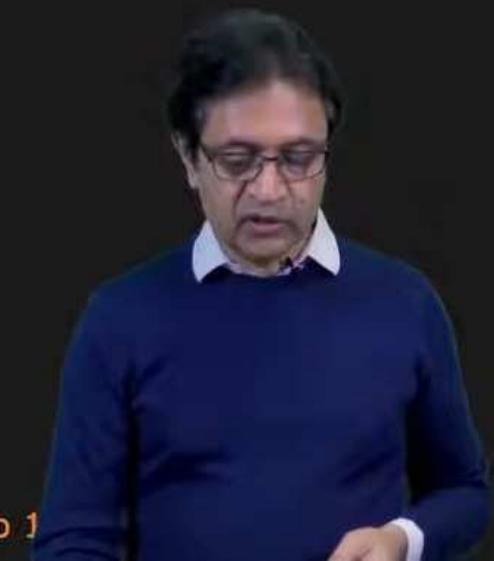
# Fisheye Lens Camera

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170° Fisheye Lens

[Miyamoto 1]



# Fisheye Image

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Fisheye Lens



Hemispherical Field of View



# Capturing the Complete Sphere

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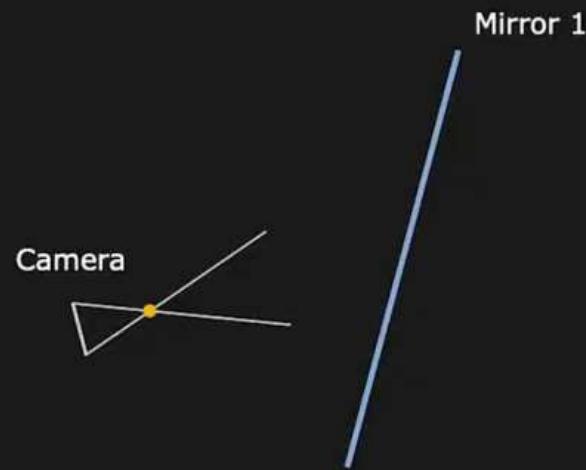


Ricoh Theta



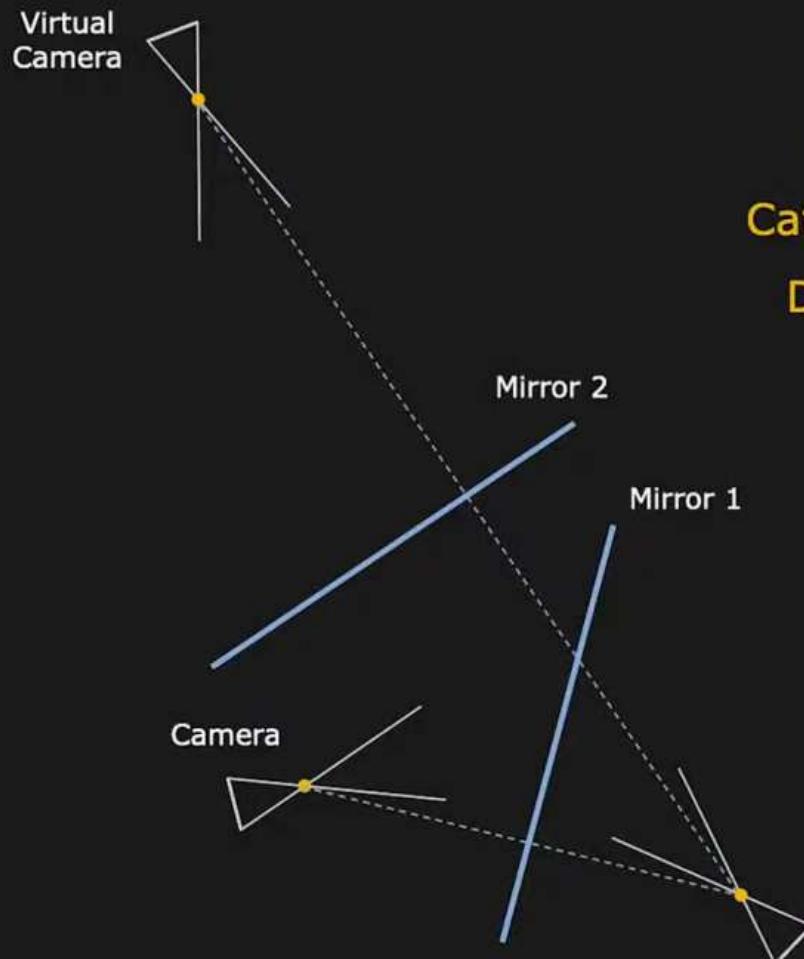
# Planar Mirrors and Reflected Cameras

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# Planar Mirrors and Reflected Cameras

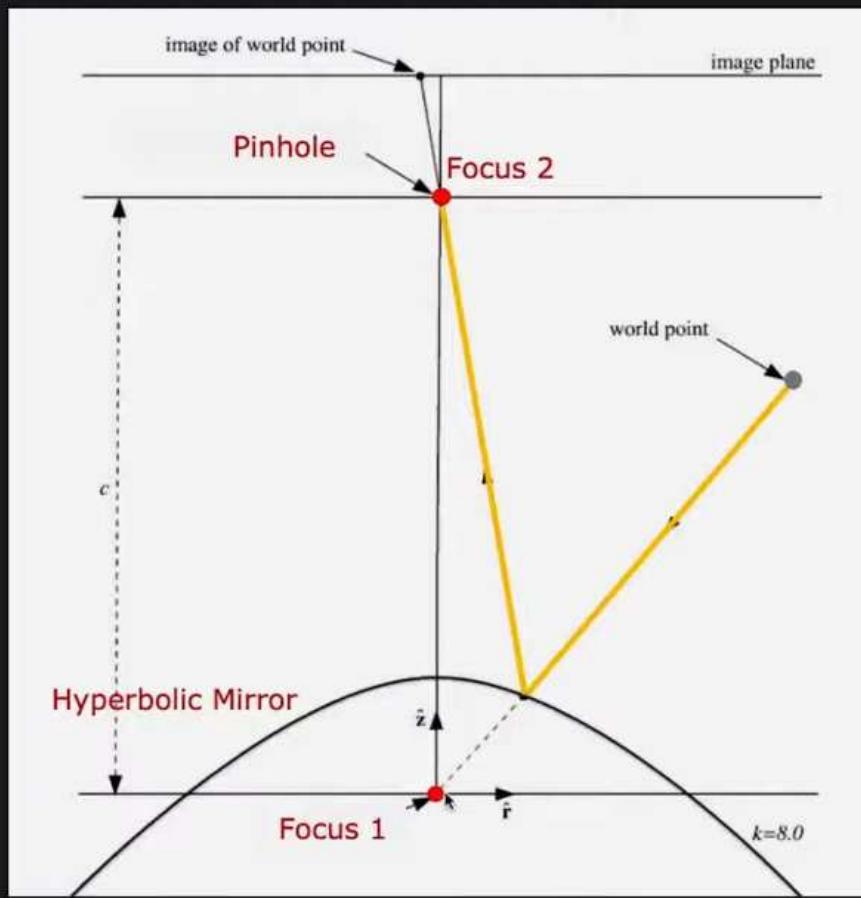
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Catadioptrics =  
Catoptrics (Mirrors) +  
Dioptrics (Lenses)



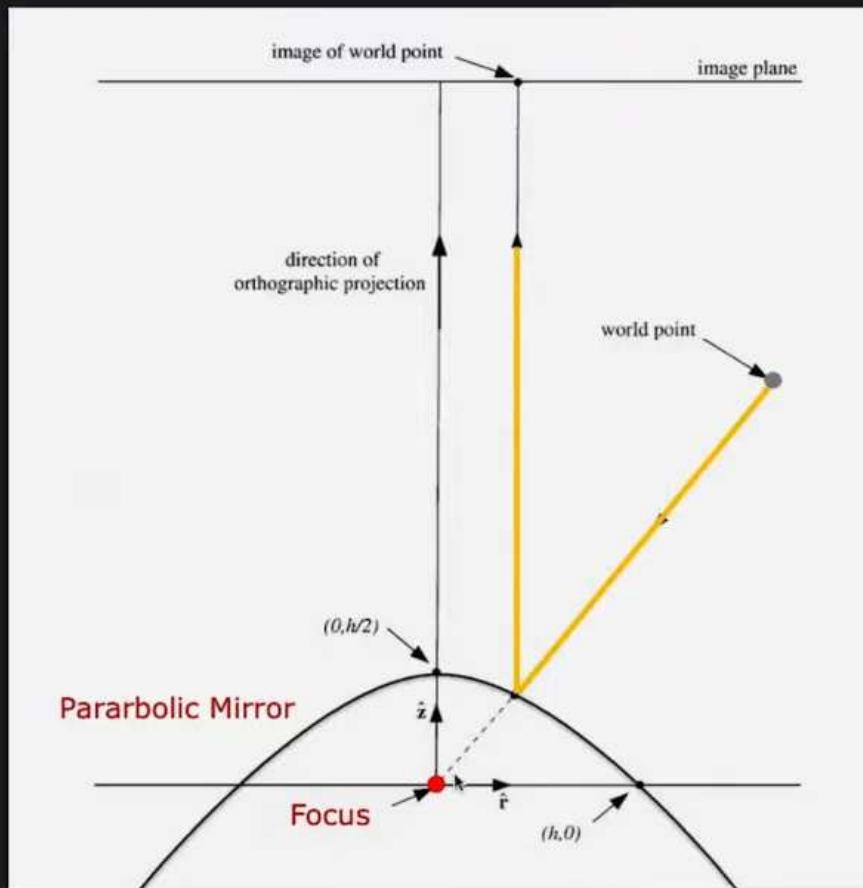
# Hyperbolic Mirror Camera



[Yagi 1999, Baker 1]



# Parabolic Mirror Camera



# Parabolic Mirror Image

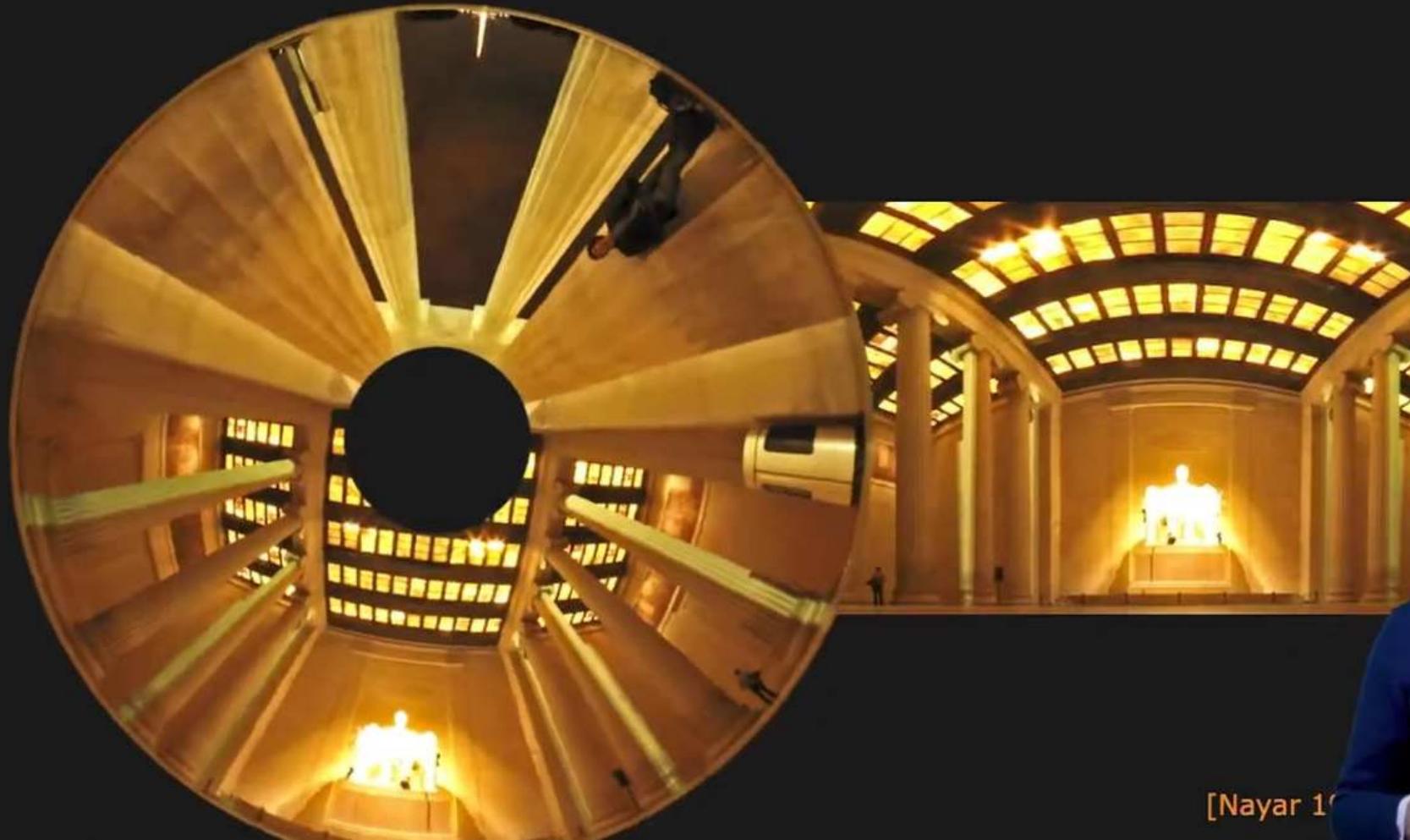
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[Nayar 1

# Parabolic Mirror Image

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[Nayar 10]



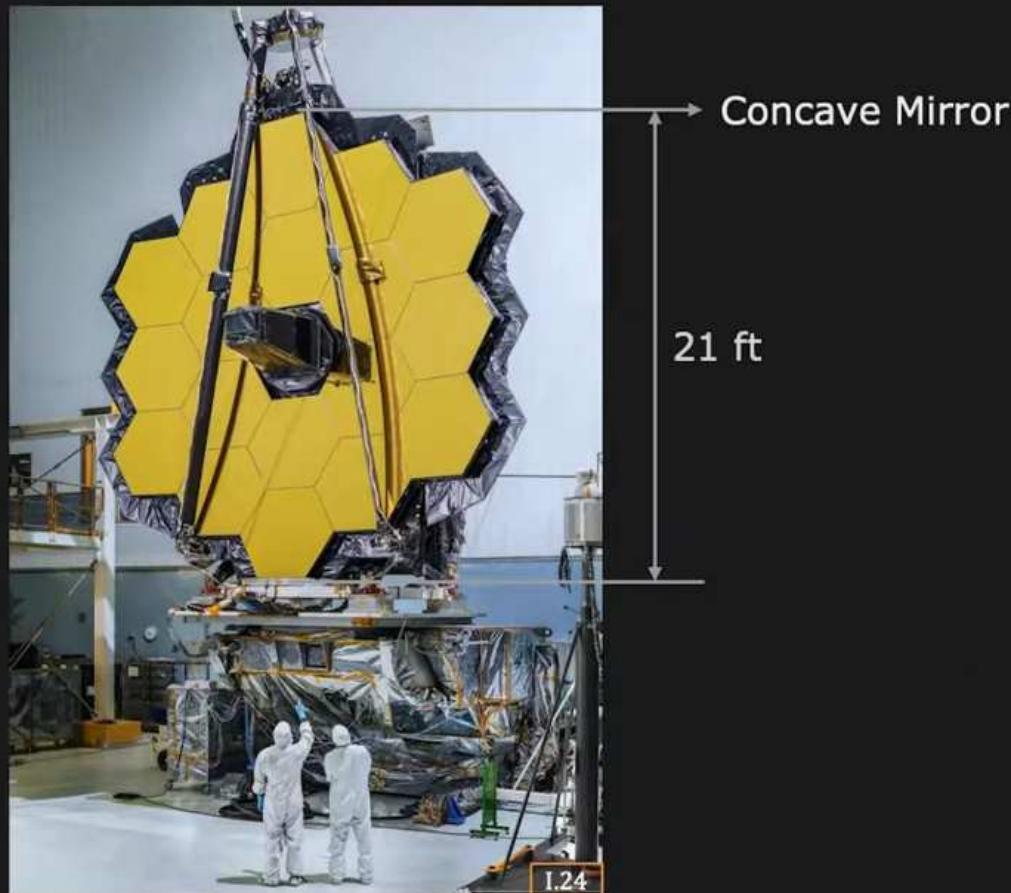
Sony "Bloggie"



Kogeto "Dot" for iPhone



# Concave Mirrors and Telescopes



James Webb Space Telescope



# Like Scallops?

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I.29



# Scallop Eyes

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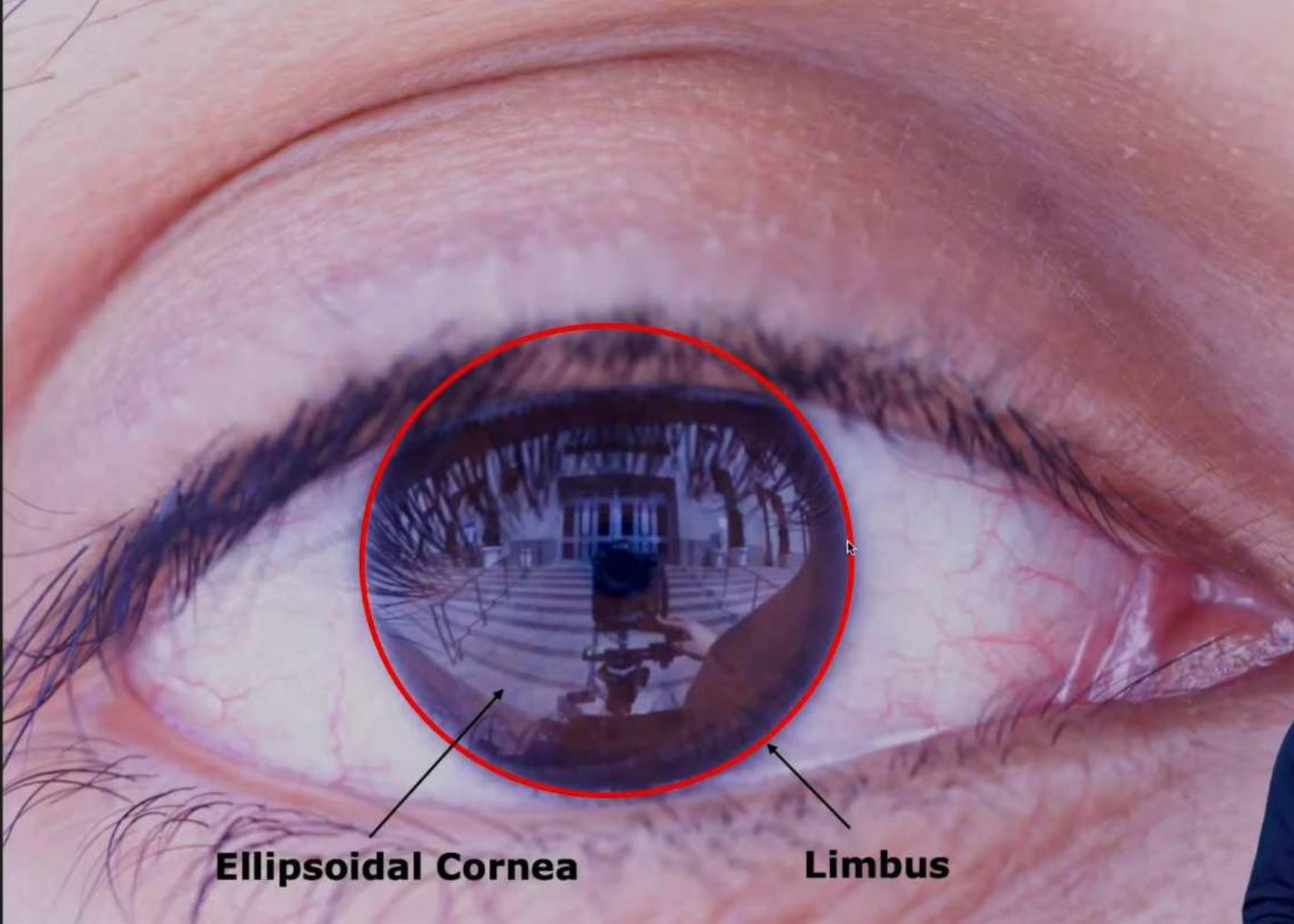
I.25

Telescopic Eyes with Parabolic Mirrors



# The World in an Eye



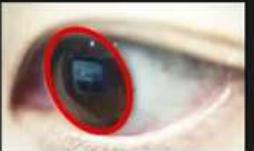


**Ellipsoidal Cornea**

**Limbus**



Eye Images



Environment Images

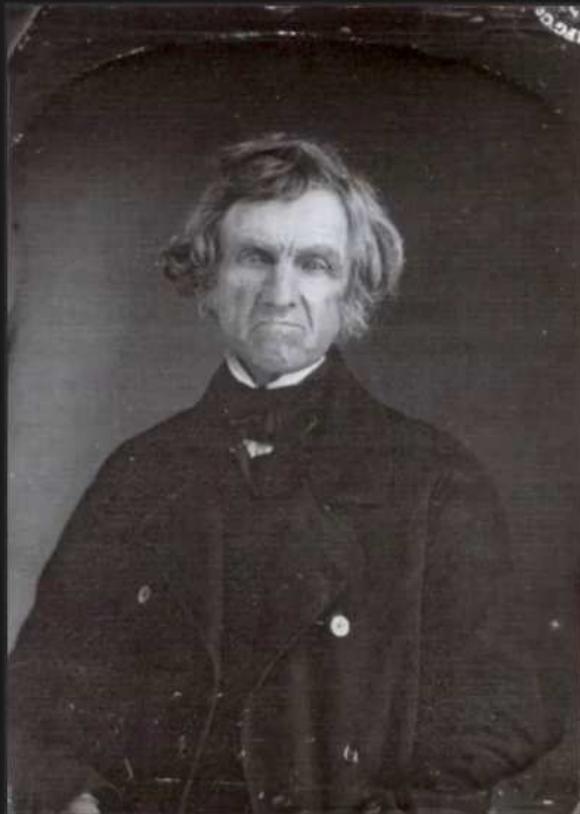


Retinal Images

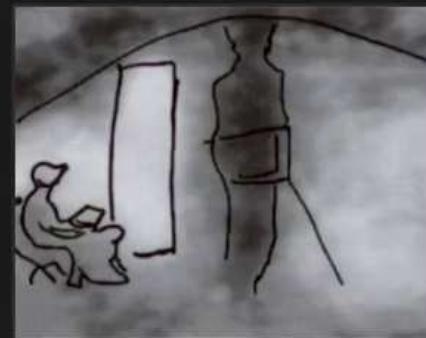


# Going Back in Time

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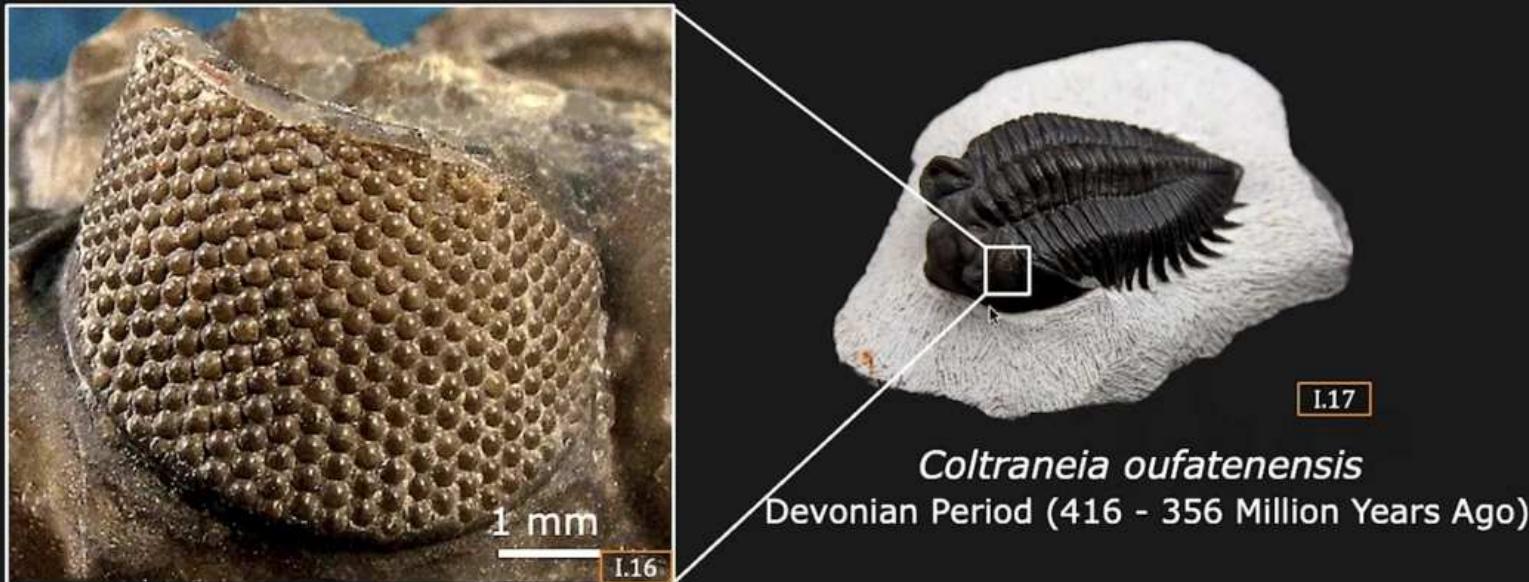
"Grumpy Grandpa", c. 1845



[Courtesy: Joe Bauman, Daguerreian Society]



# Fossilized Eye of Trilobite

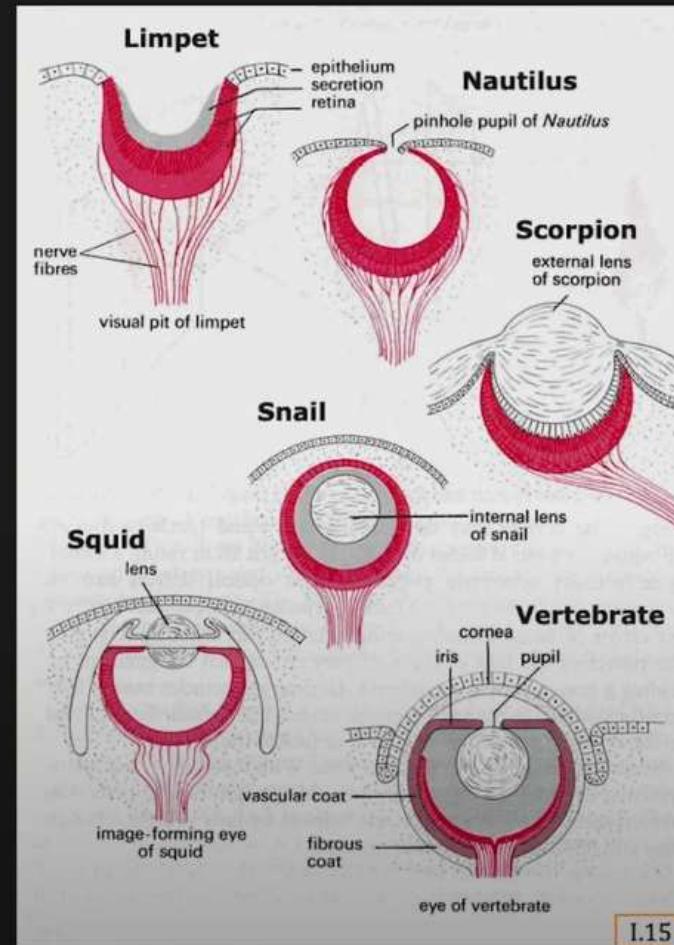


Earliest kind of eye preserved as a fossil. The facets are the corneal lenses made from transparent Calcite ( $\text{CaCO}_3$ ).

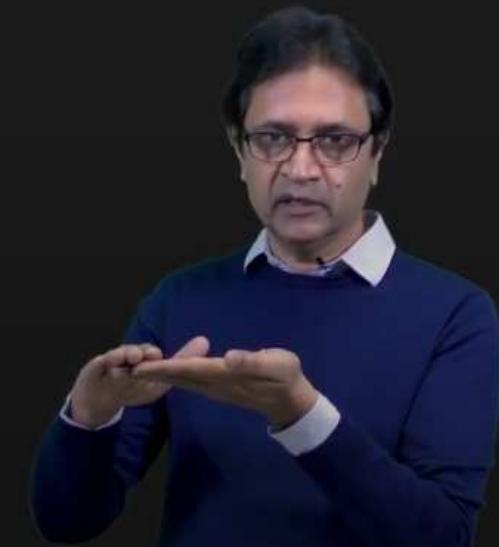




# Primitive Eyes

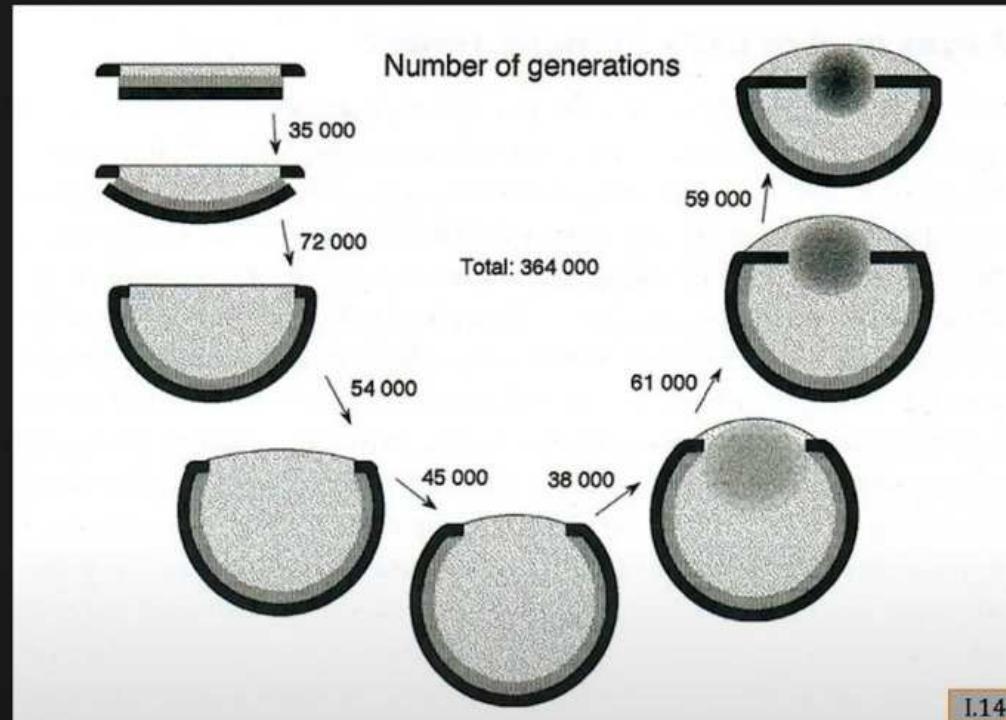


I.15





# Evolution of Eye: A Simulation

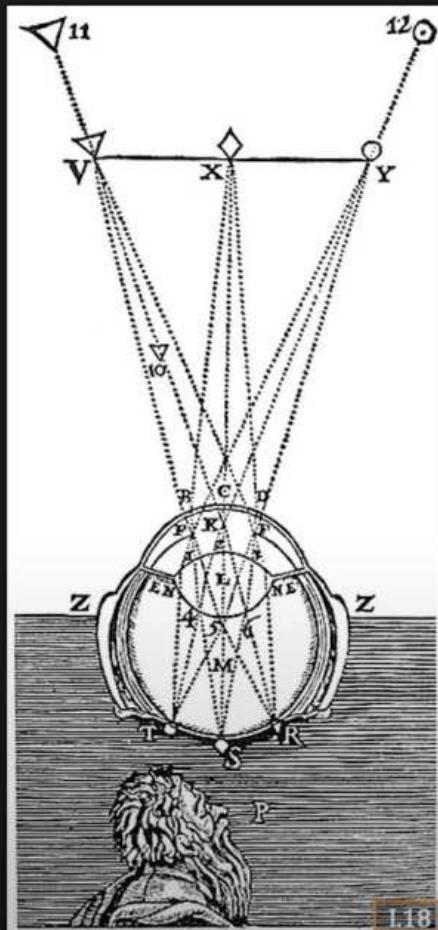


A patch of light sensitive epithelium can be gradually turned into a perfectly focused camera-type eye, if there is a continuous natural selection for improved spatial vision. This simulation reveals that the complete evolution can be accomplished in about 400,000 generations. First a pigment cup eye evolves, and then a lens.

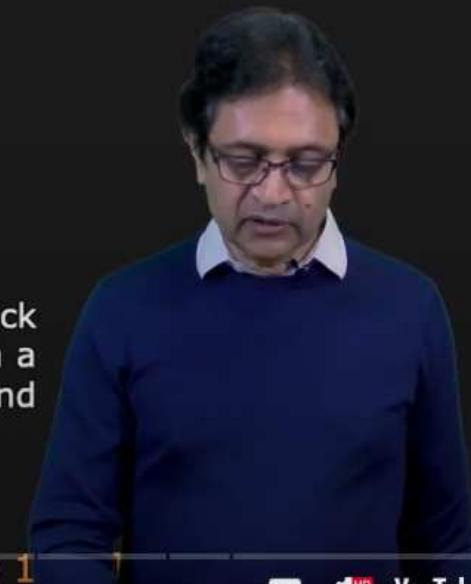




# Image Formation in the Eye

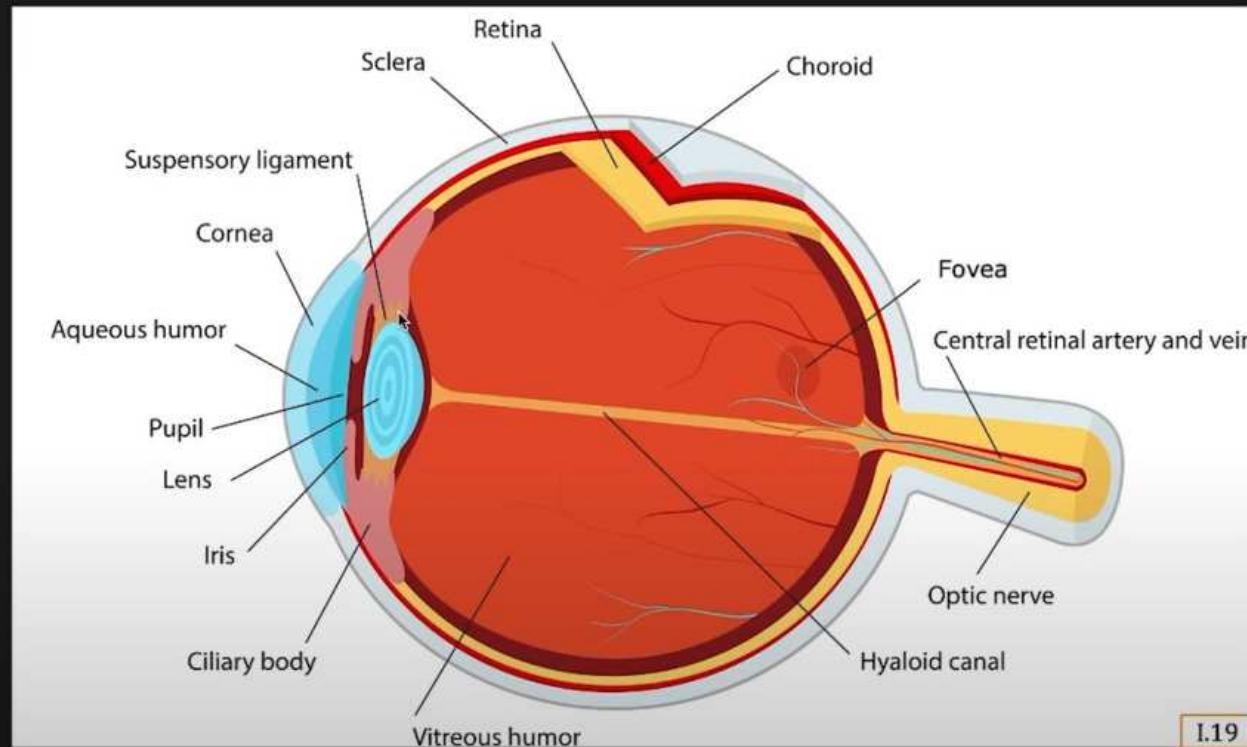


Descartes removed the eye of an ox, scraped its back to make it transparent, and then observed on it from a darkened room "not perhaps without wonder and pleasure" the inverted image of the scene.

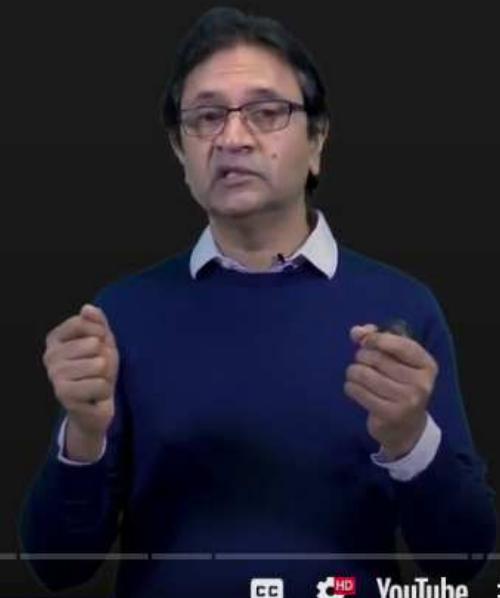




# Optics in Human Eye

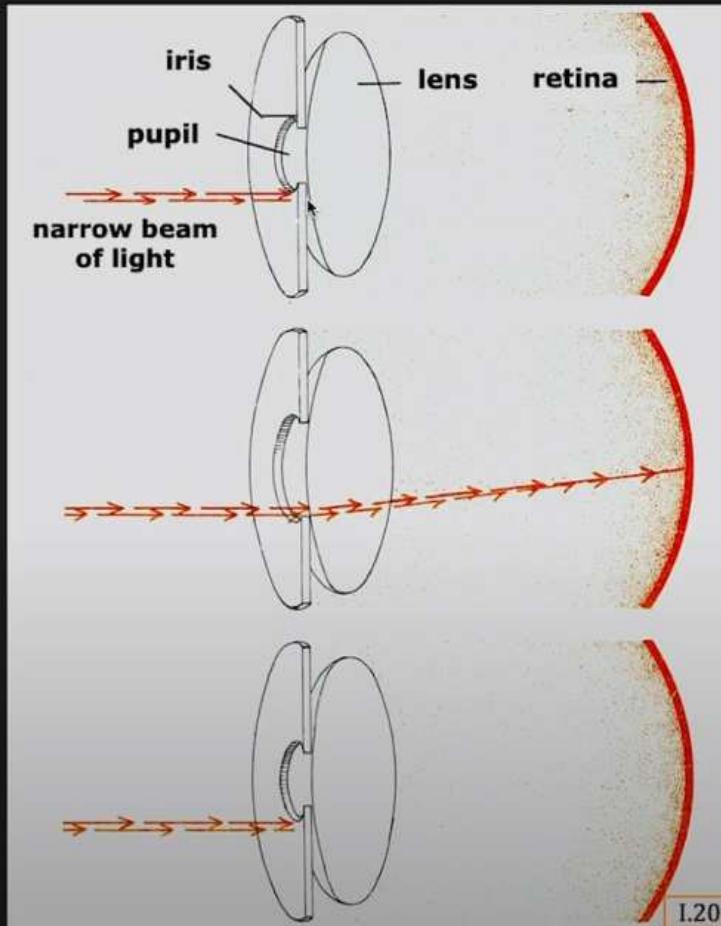


I.19





# Human Eye: Iris Control System

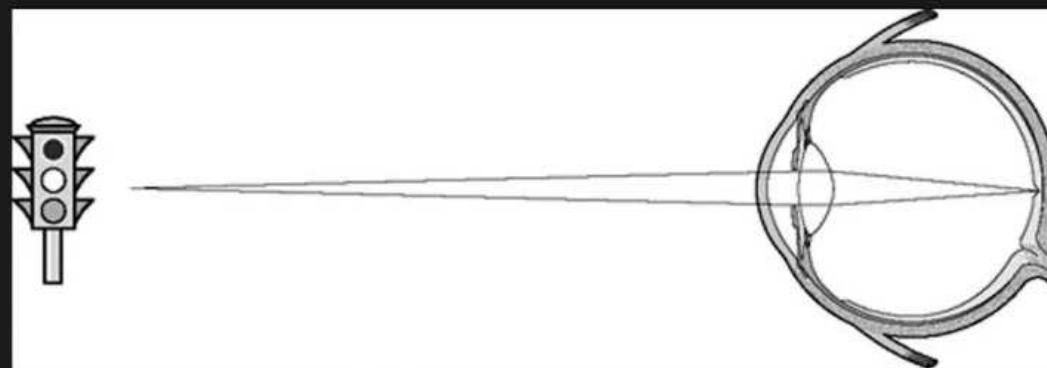


Making the iris oscillate with a narrow beam of light. When the iris opens up, strong light reaches the retina, which causes the retina to close. When it closes, no light is received by the retina and the iris opens again. The frequency and amplitude of this oscillation of the iris reveals the response of its control system.





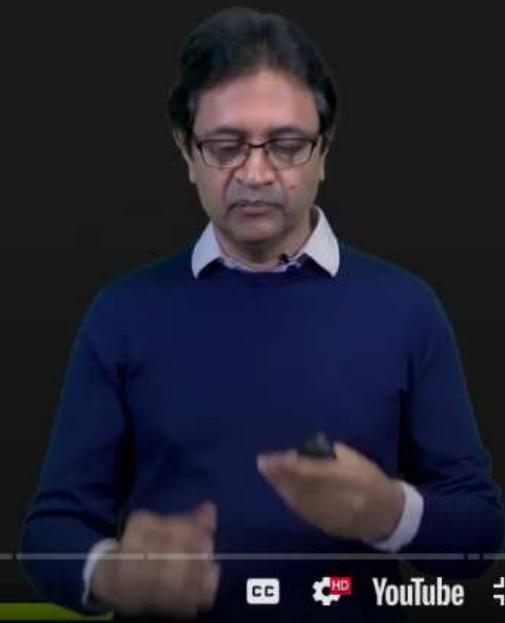
# Accommodation (Focusing) in the Eye



Focusing on distant objects: Lens is relaxed

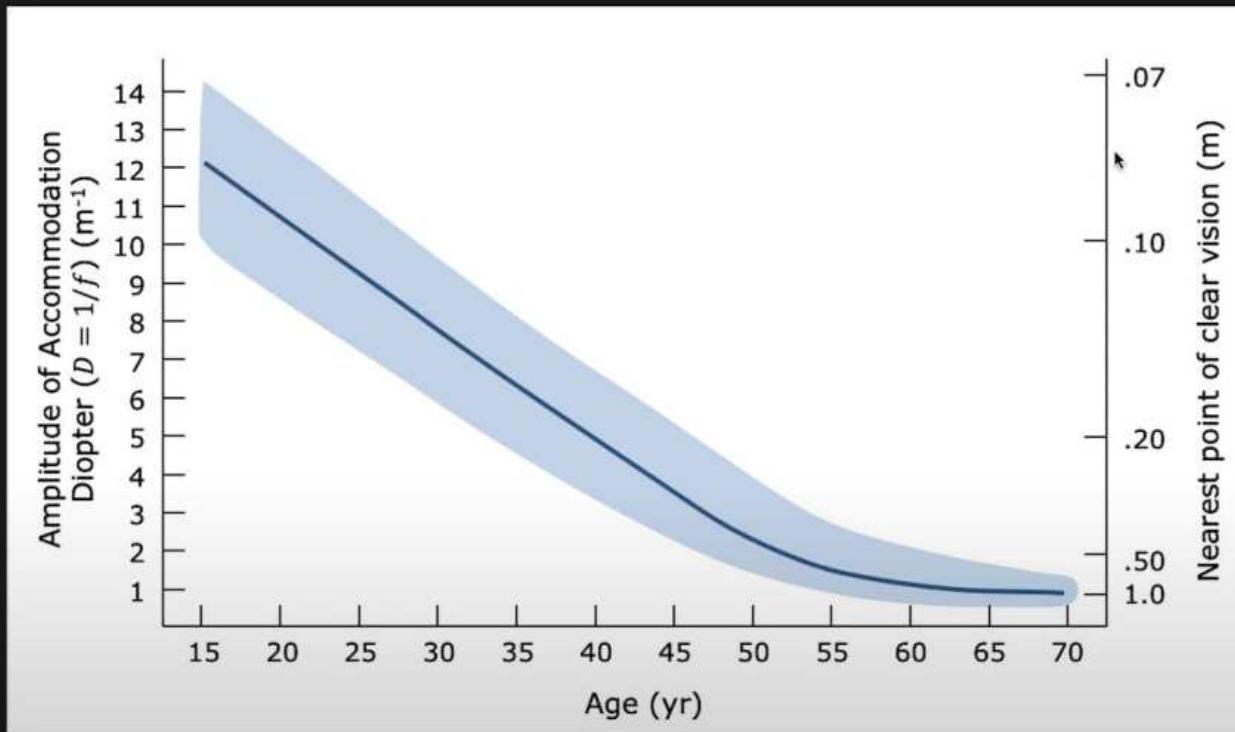


Focusing on nearby objects: Lens is squished





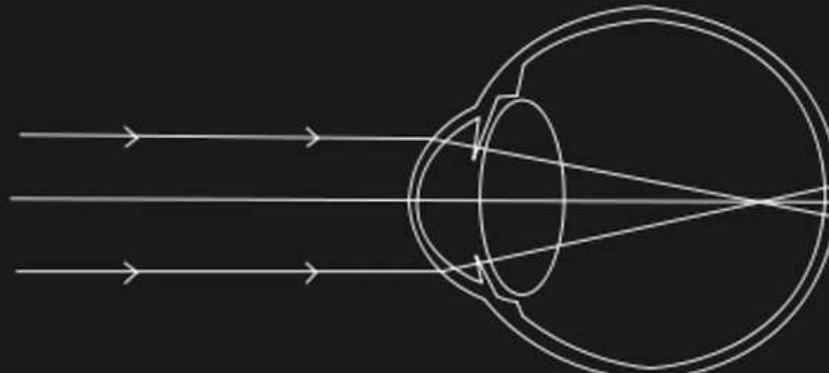
# Change in Accommodation with Age



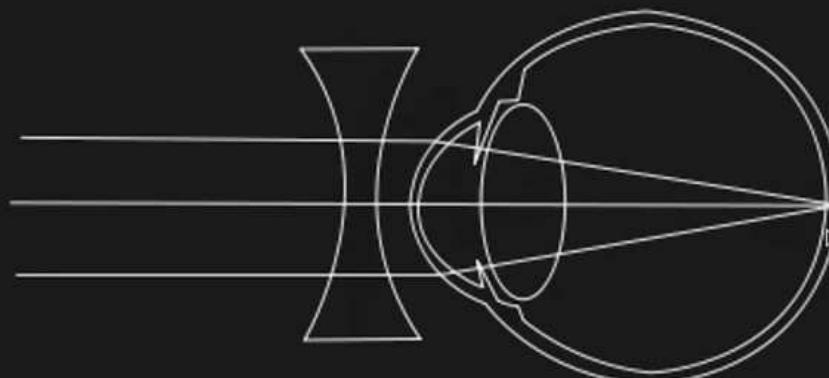
Accommodation decreases with age



# Myopia (Near-Sightedness)



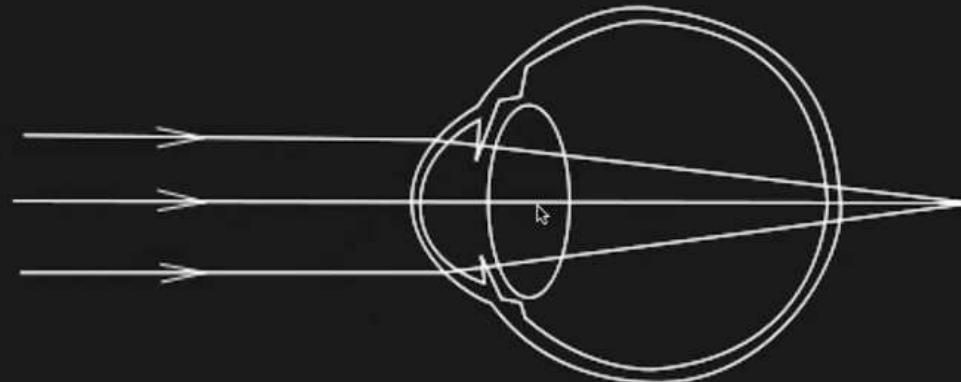
Inability to focus on objects far away



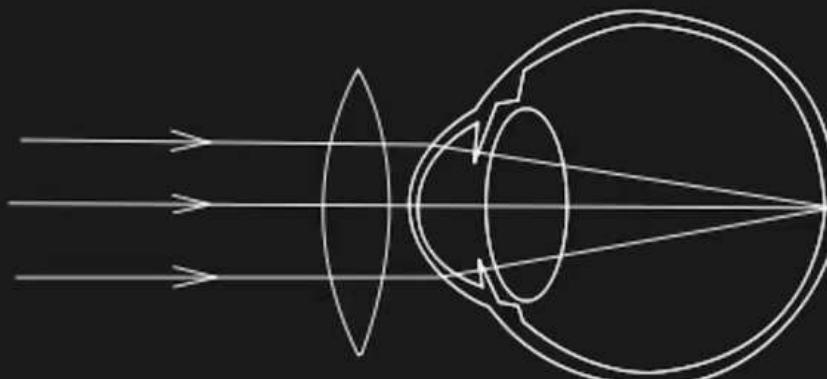
Can be fixed by using a diverging (concave) lens



# Hyperopia (Far-Sightedness)



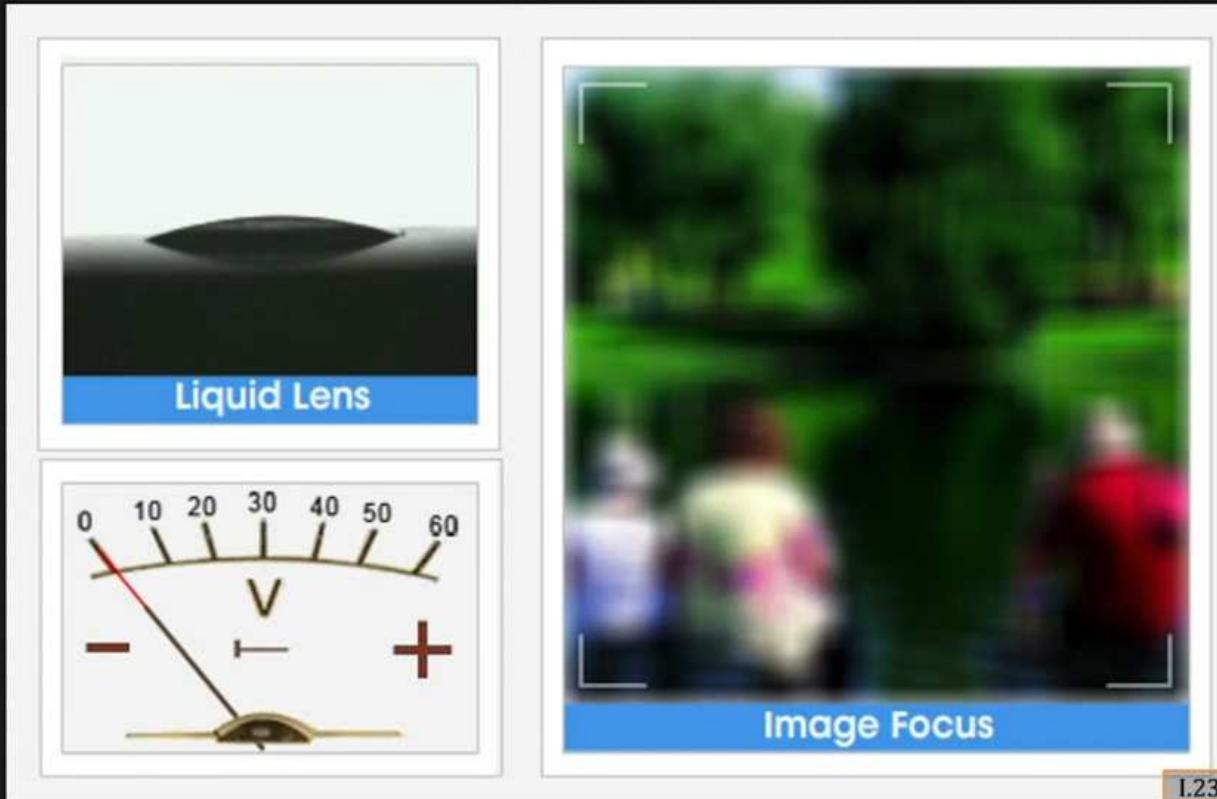
Inability to focus on nearby objects



Can be fixed by using a converging (convex) lens



# Liquid Lens

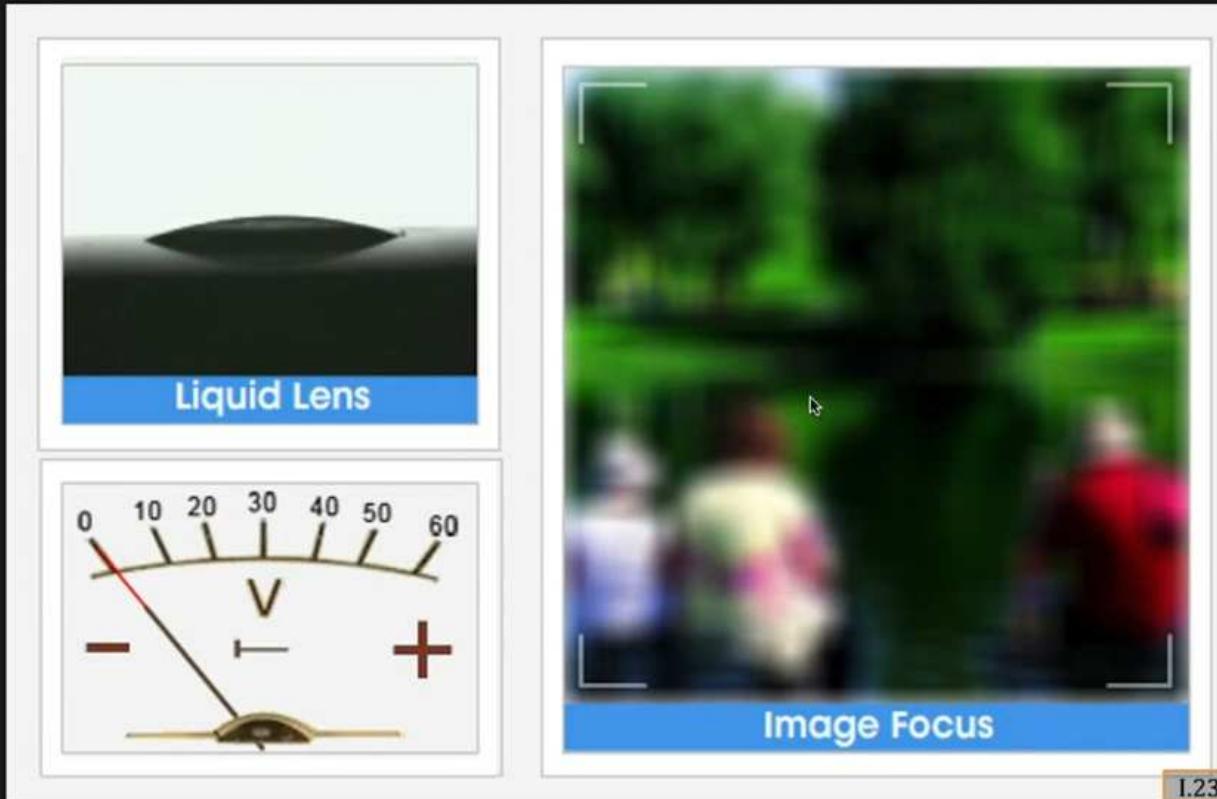


I.23

The shape and hence the focal length of the liquid lens can be precisely controlled by applying a voltage.



# Liquid Lens



I.23

The shape and hence the focal length of the liquid lens can be precisely controlled by applying a voltage.



# Image Sensing

---

Need to convert Optical Images to Digital Images  
*(numbers)* for computer representation and use.

## Topics:

- (1) A Brief History of Imaging
- (2) Types of Image Sensors
- (3) Resolution, Noise, Dynamic Range
- (4) Sensing Color
- (5) Camera Response and HDR Imaging



# Image Sensing

---

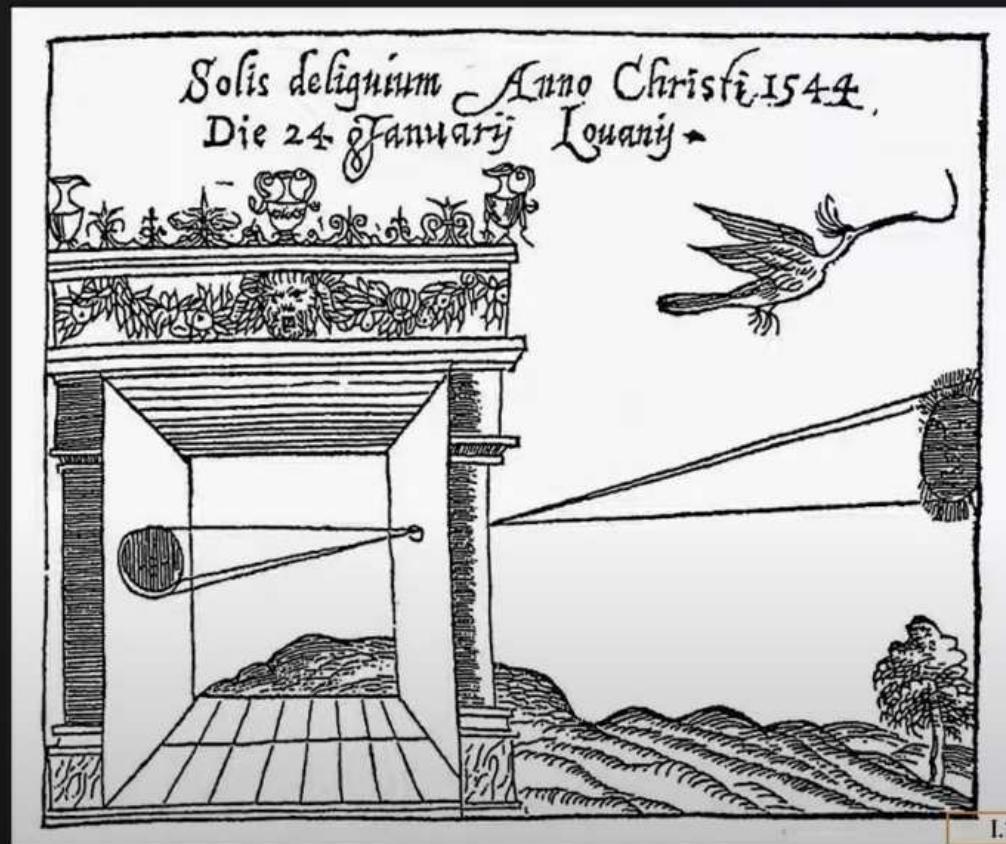
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- (5) Camera Response and HDR Imaging
- (6) Nature's Image Sensors



# Pinhole Camera

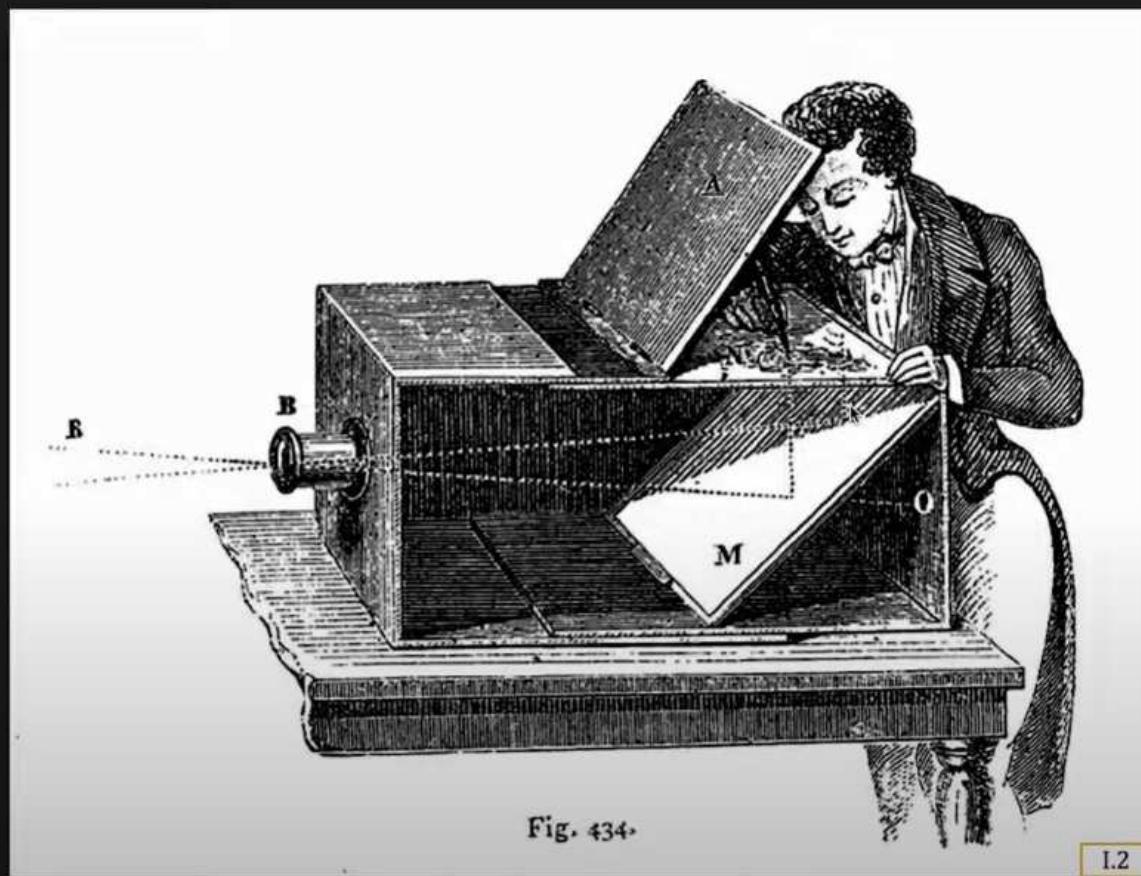


*Camera Obscura*

1558



# Lens Based Camera Obscura



1558  
1568



# Invention of Film



Still Life, Louis Jaques Mande Daguerre, 1837



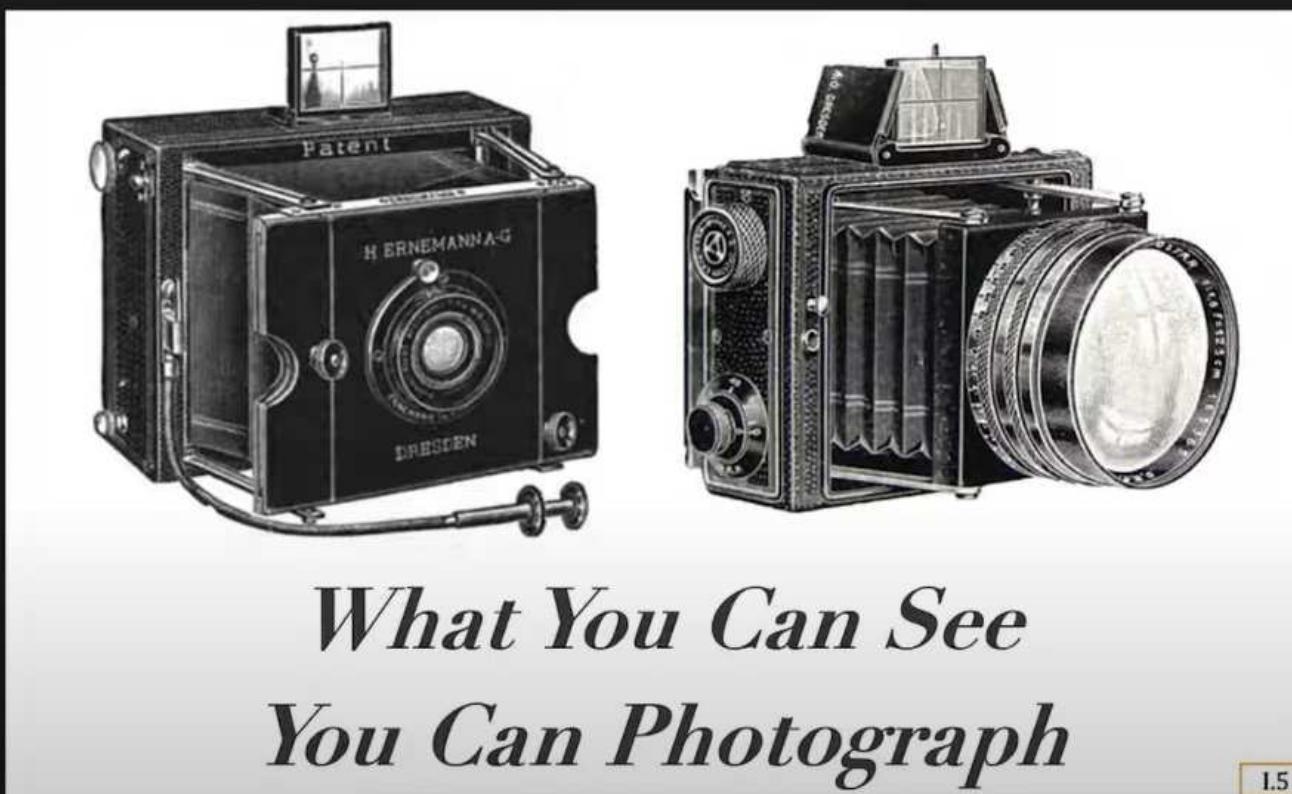
# Color Film



Louis Ducos du Hauron, 1887

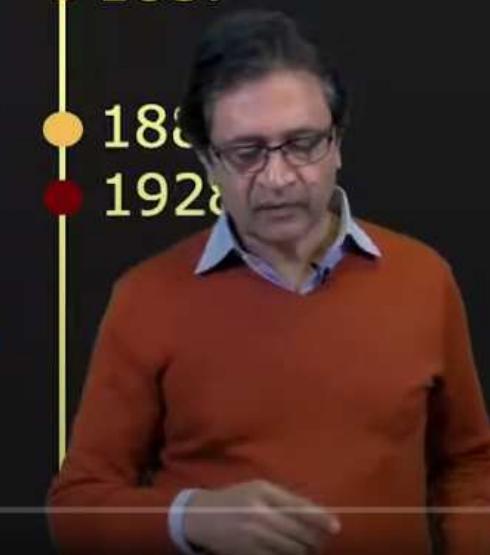


# Ernemann Camera



*What You Can See  
You Can Photograph*

- 1558
- 1568
- 1837
- 1851
- 1928

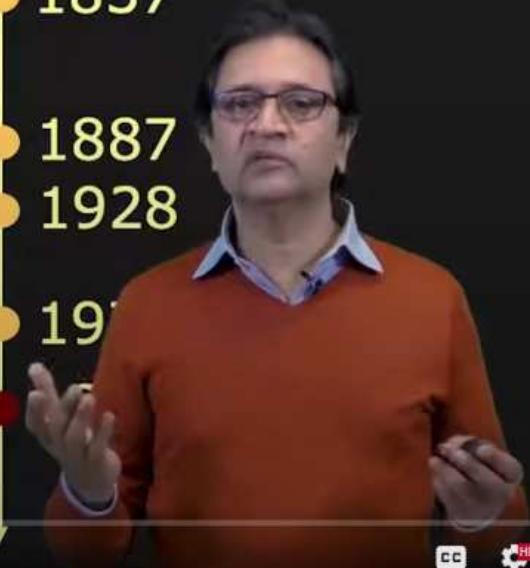


1.5

# Silicon Image Detector



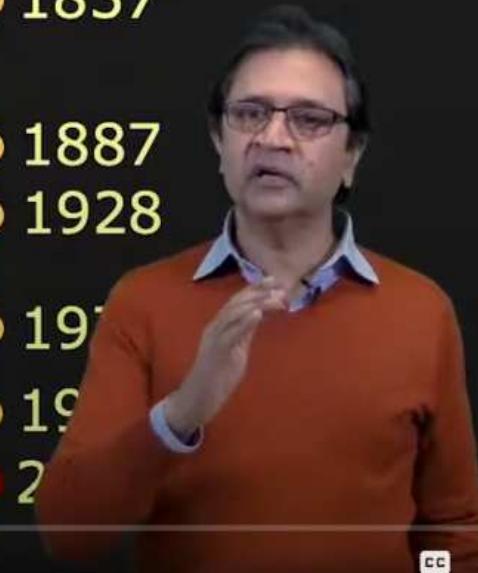
# Digital Cameras



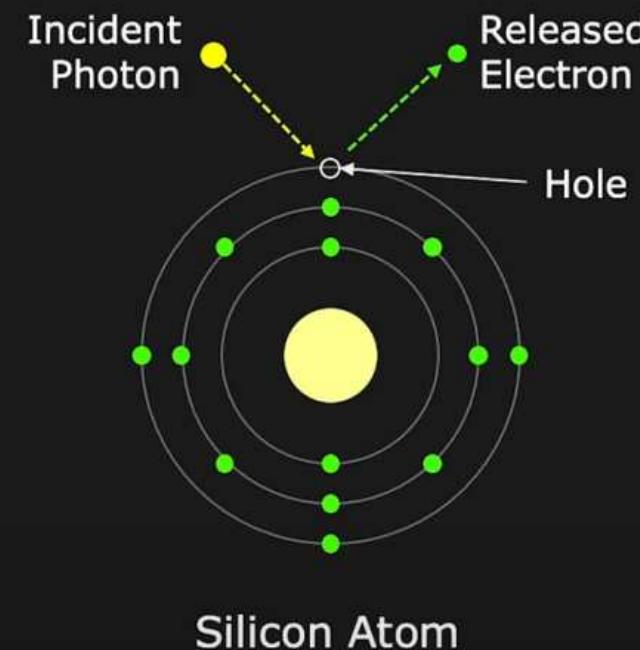
# Phones with Cameras



iPhone 1



# Converting Light into Electric Charge

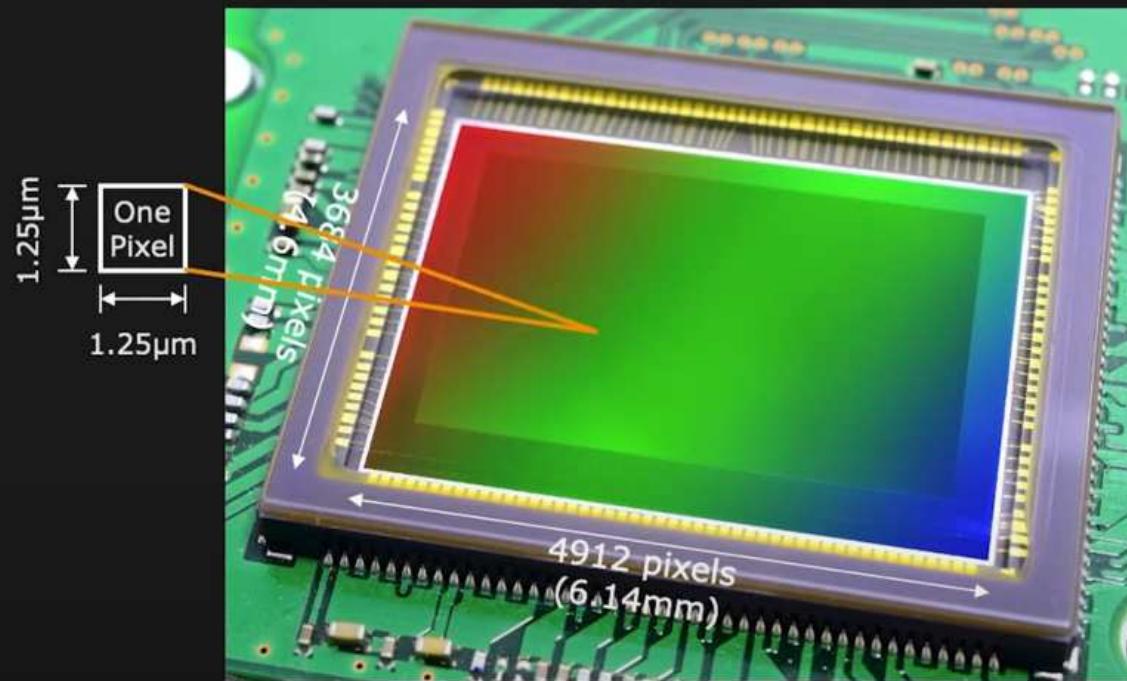


Photon with sufficient energy incident on a Si atom creates an **electron-hole pair**.



▶ Jump ahead

# Image Sensor: A Closer Look

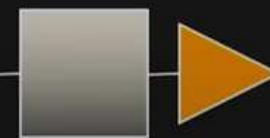
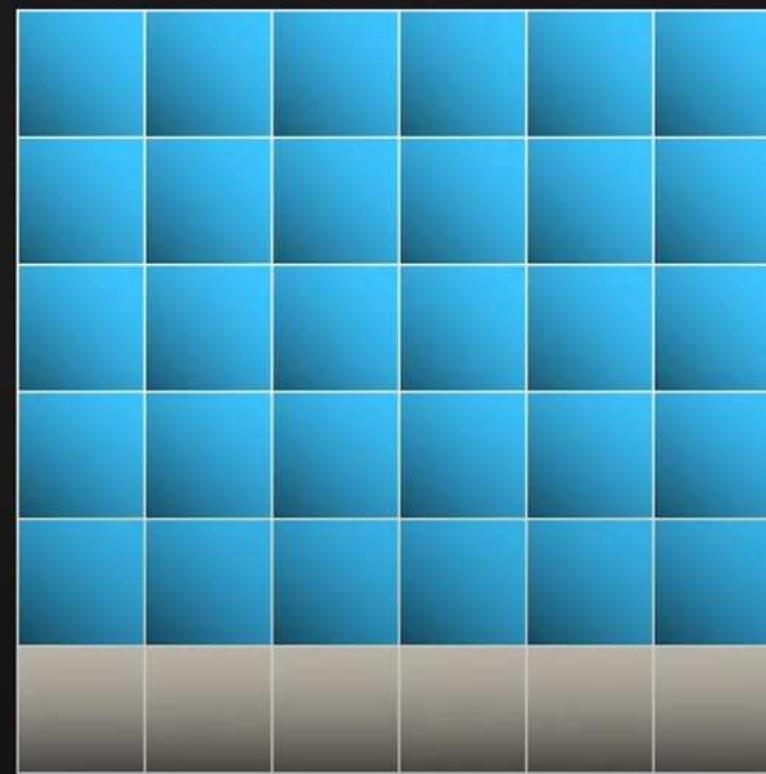


18 Megapixels



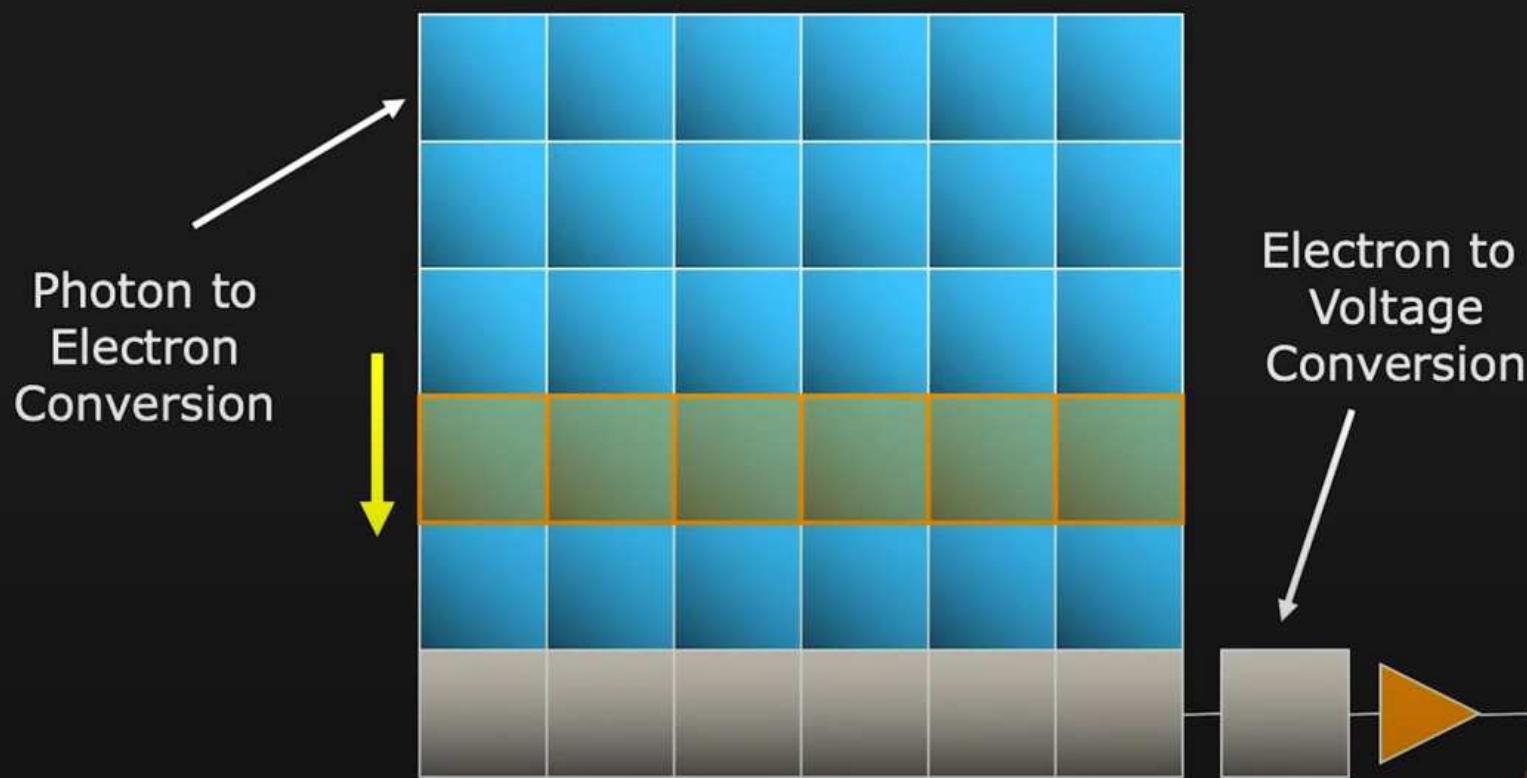
# Types of Image Sensors: CCD

CCD: Charge Coupled Device



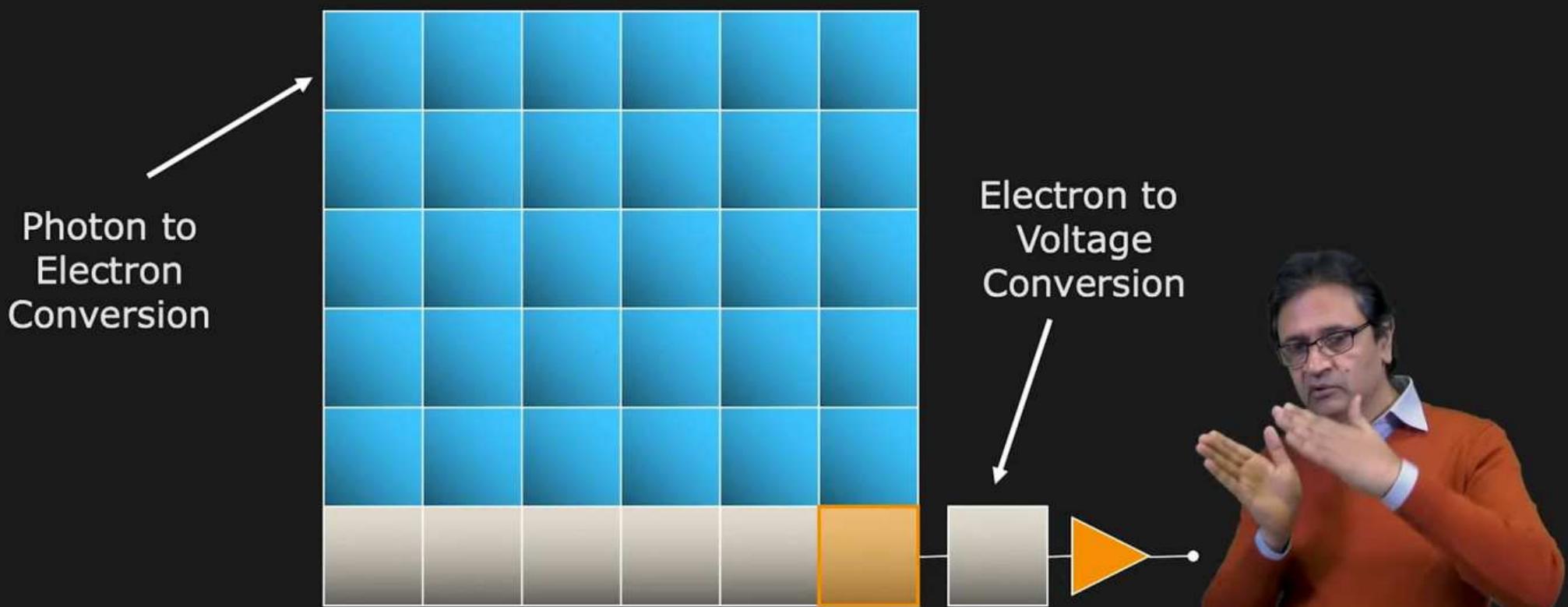
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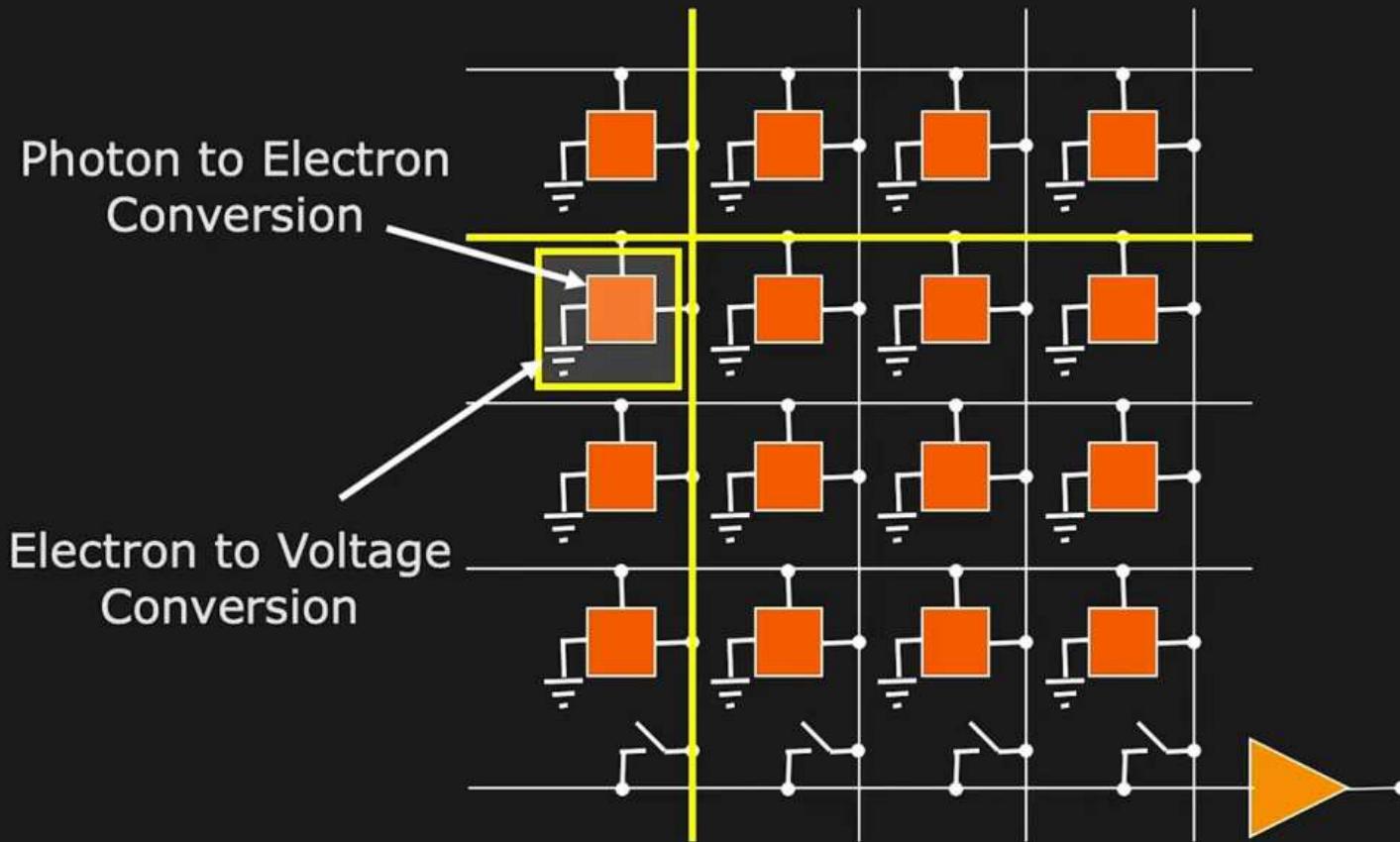
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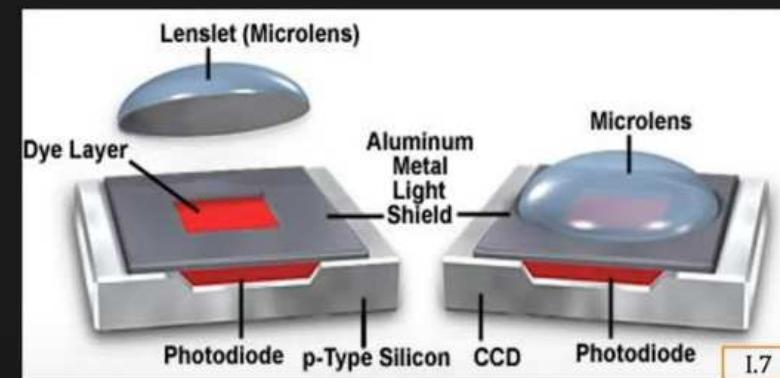
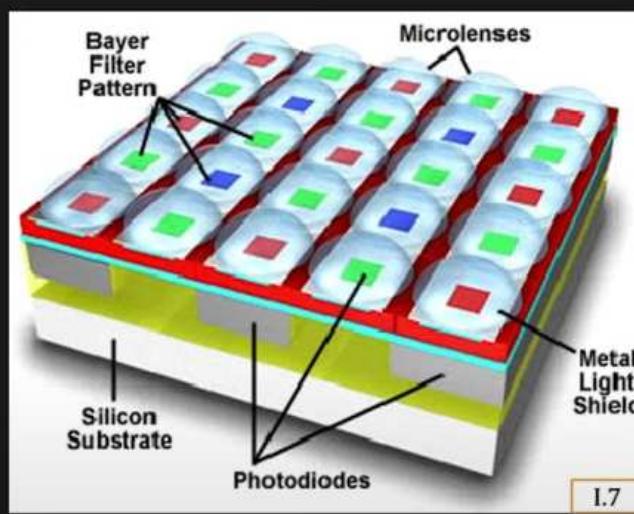


# Types of Image Sensors: CMOS

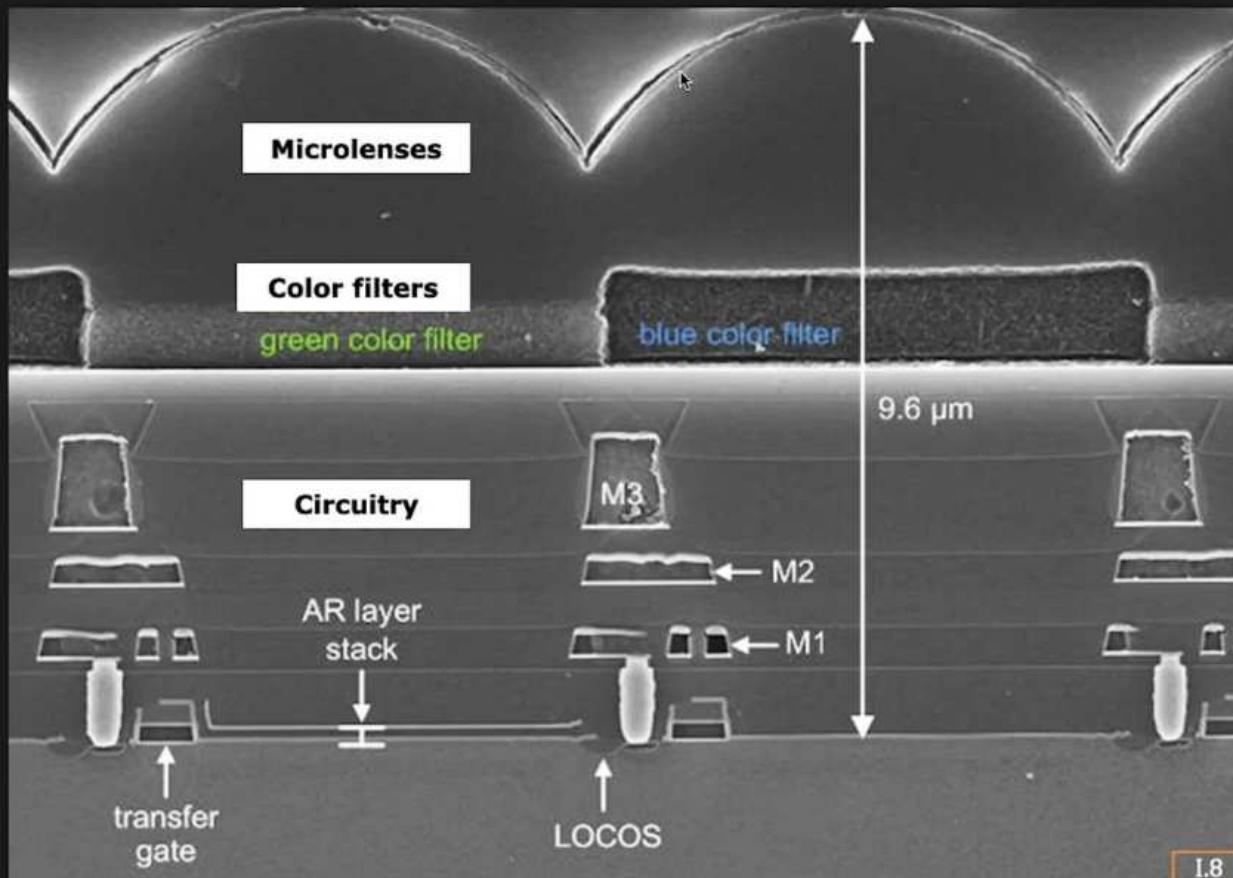
CMOS: Complimentary Metal-Oxide Semiconductor



# Image Sensor: A Closer Look



# Image Sensor: A Closer Look

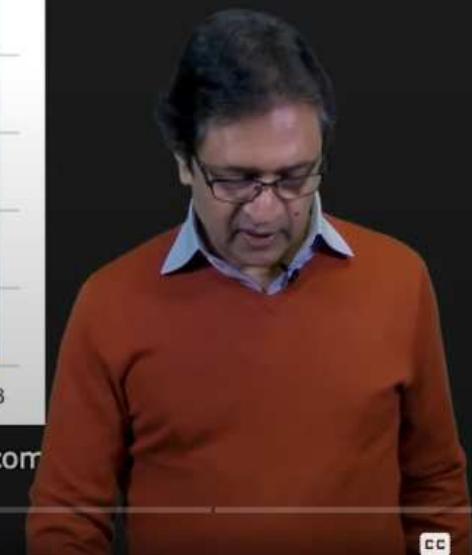
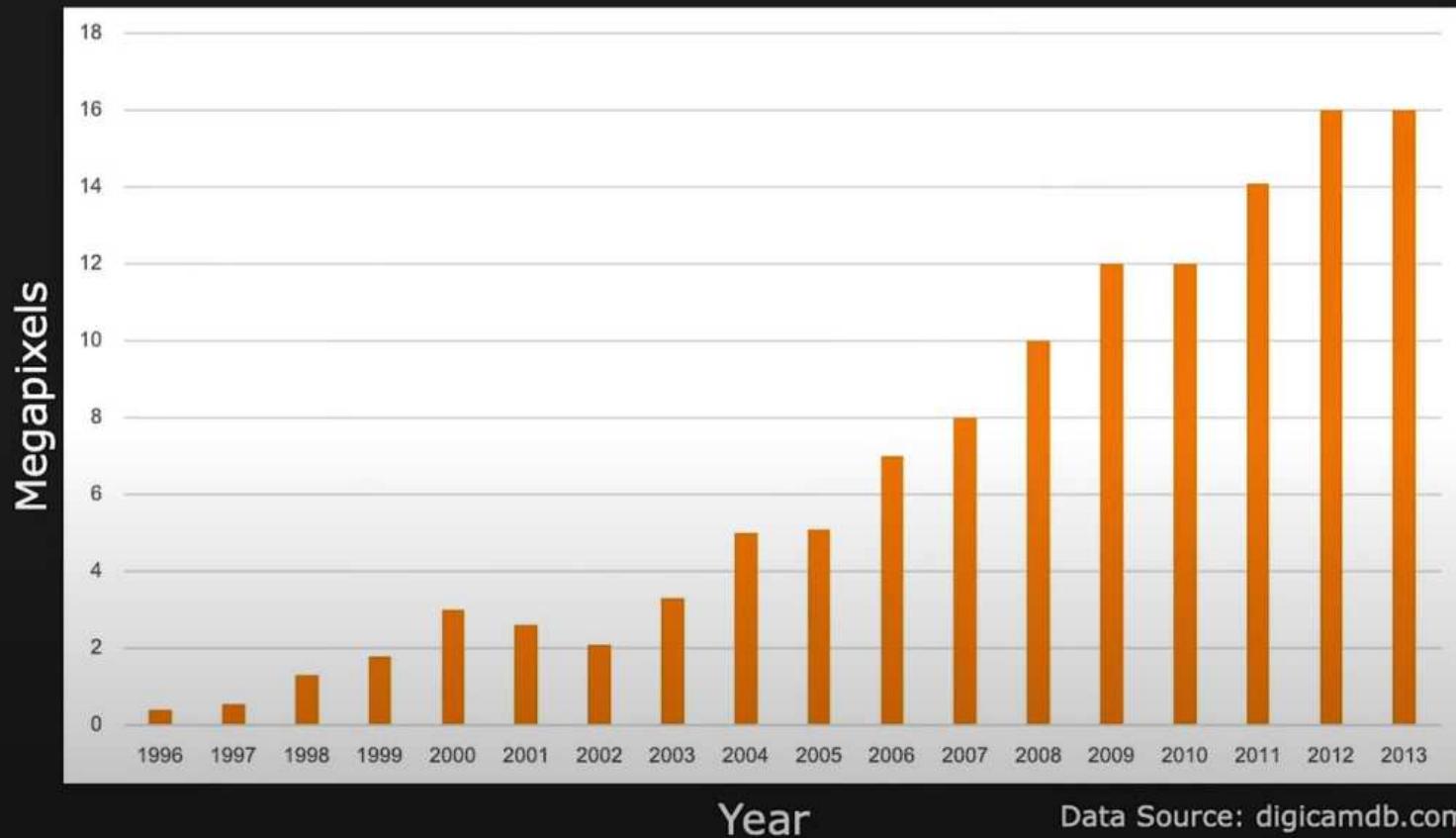


Cross Section of an Image Sensor  
(Scanning Electron Microscope Image)



# Image Sensor Resolution

Median Sensor Resolution in Consumer Cameras



# Noise in Image Sensors

---

**Noise:** Unwanted modification of signal during capture, conversion, transmission, processing.

- **Photon Shot Noise (Scene Dependent)**
  - Quantum nature of light
  - Random arrival of photons
- **Readout Noise (Scene Independent)**
  - Electronic Noise: Pre analog-to-digital conversion
  - Quantization Noise: Post analog-to-digital conversion
- **Other Sources (Scene Independent)**
  - Dark Current Noise: Thermally generated electrons
  - Fixed Pattern Noise: Defective pixels



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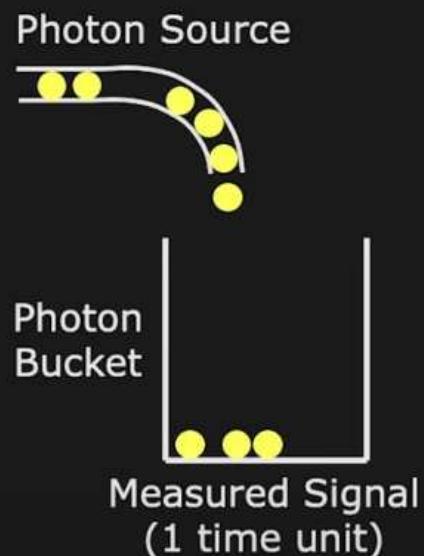
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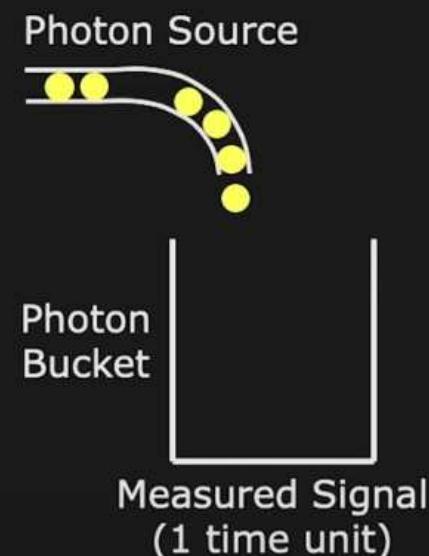
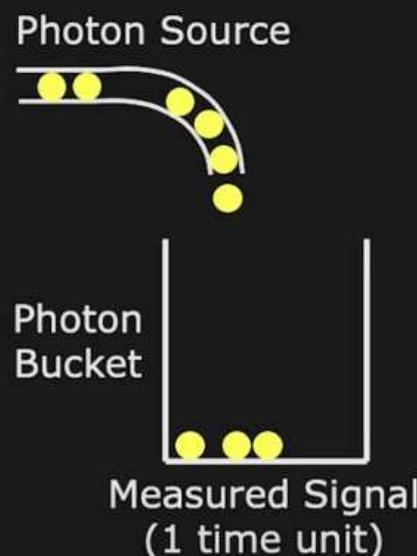
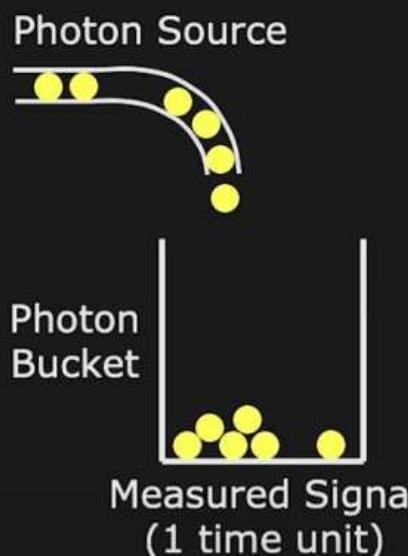
# Photon Shot Noise



Average Photon Flux (Per Unit Time) = 3 Photons



# Photon Shot Noise

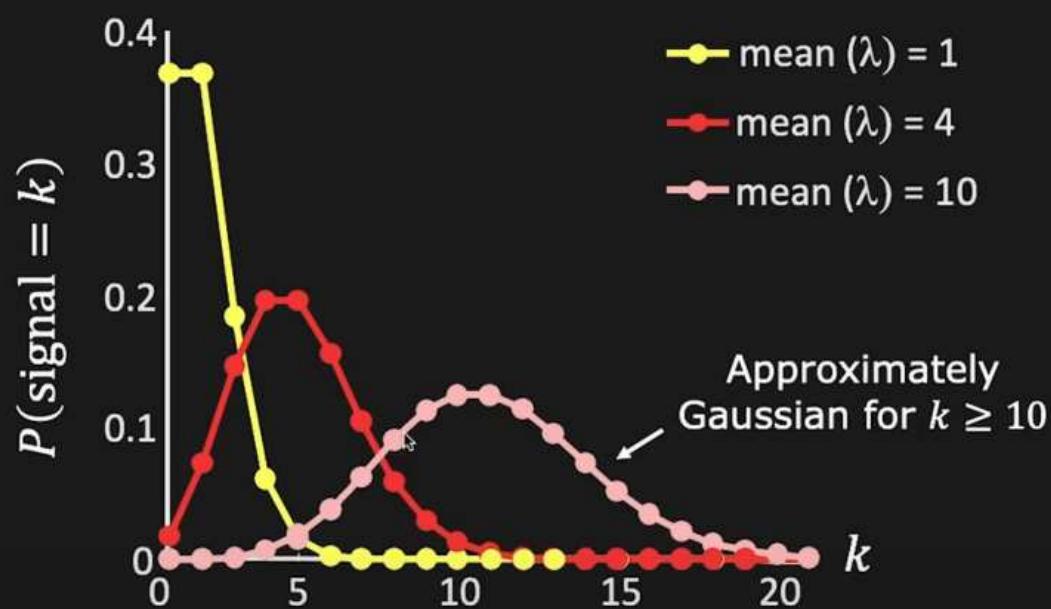


Average Photon Flux (Per Unit Time) = 3 Photons

Variation Due to Random Generation of Photons



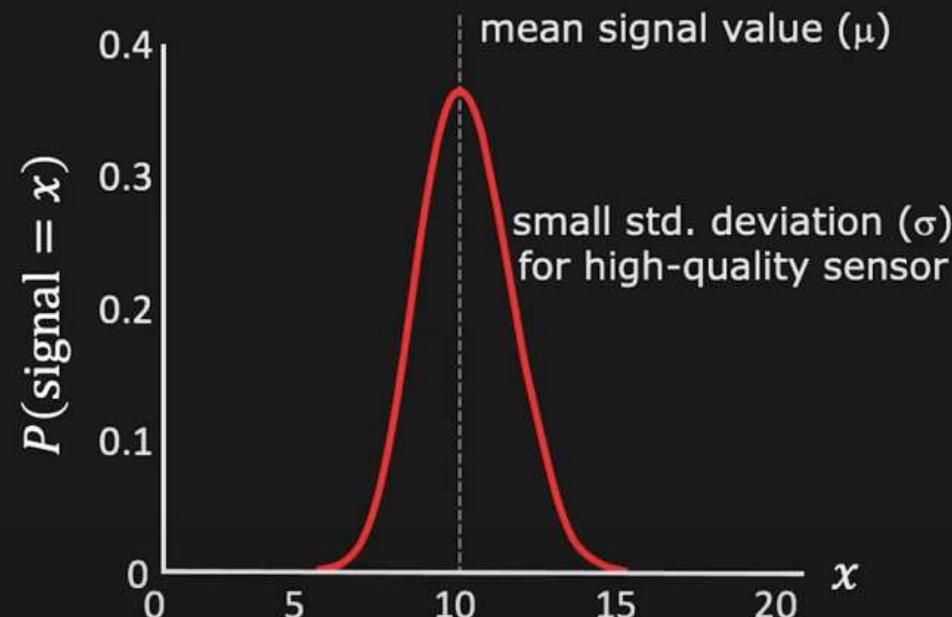
# Photon Noise: Poisson Distribution



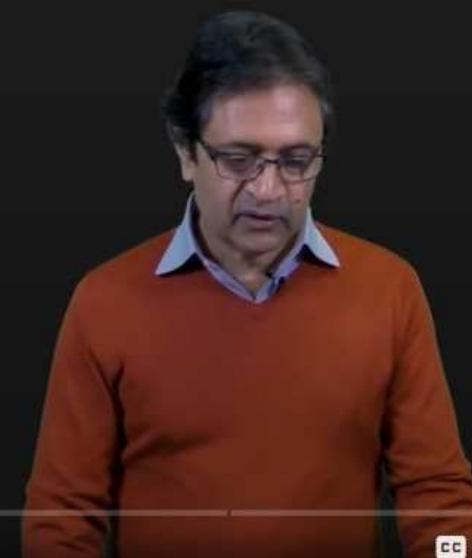
$$P(\text{signal} = k) = \frac{\lambda^k e^{-\lambda}}{k!}$$

$\text{Var} [\text{signal}] = \text{Mean} [\text{signal}] \Rightarrow \text{Scene Dependent Noise}$

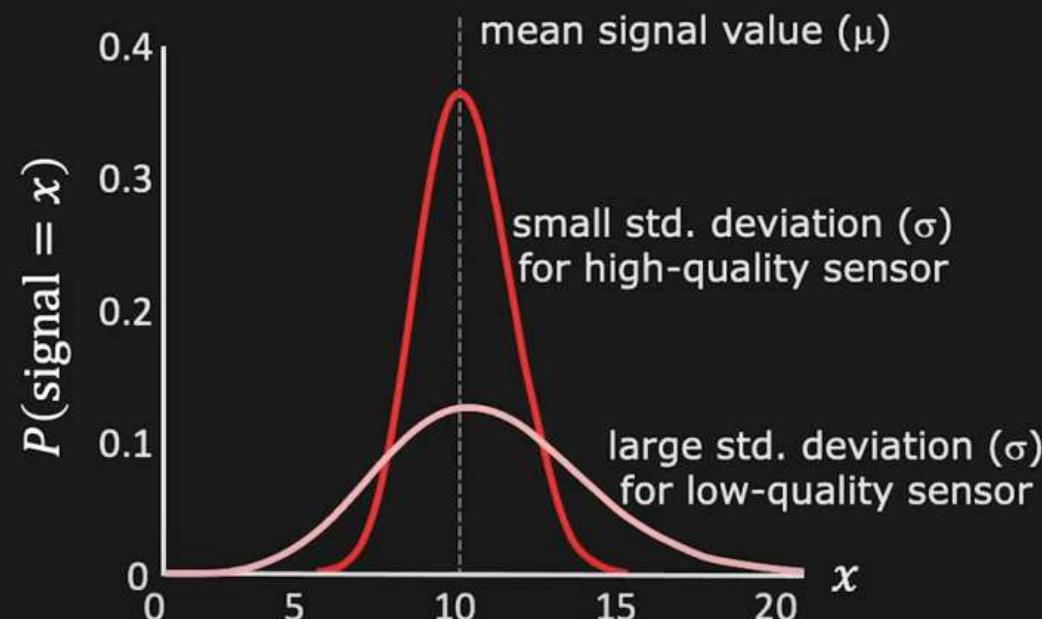
# Read Noise: Gaussian Distribution



$$P(\text{signal} = x) = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^2}$$



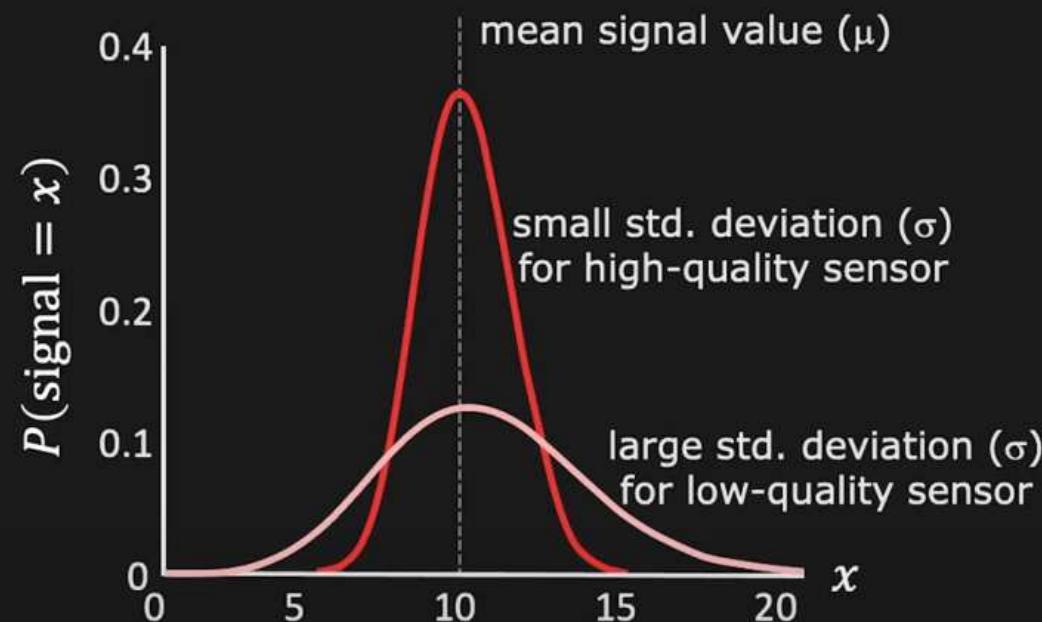
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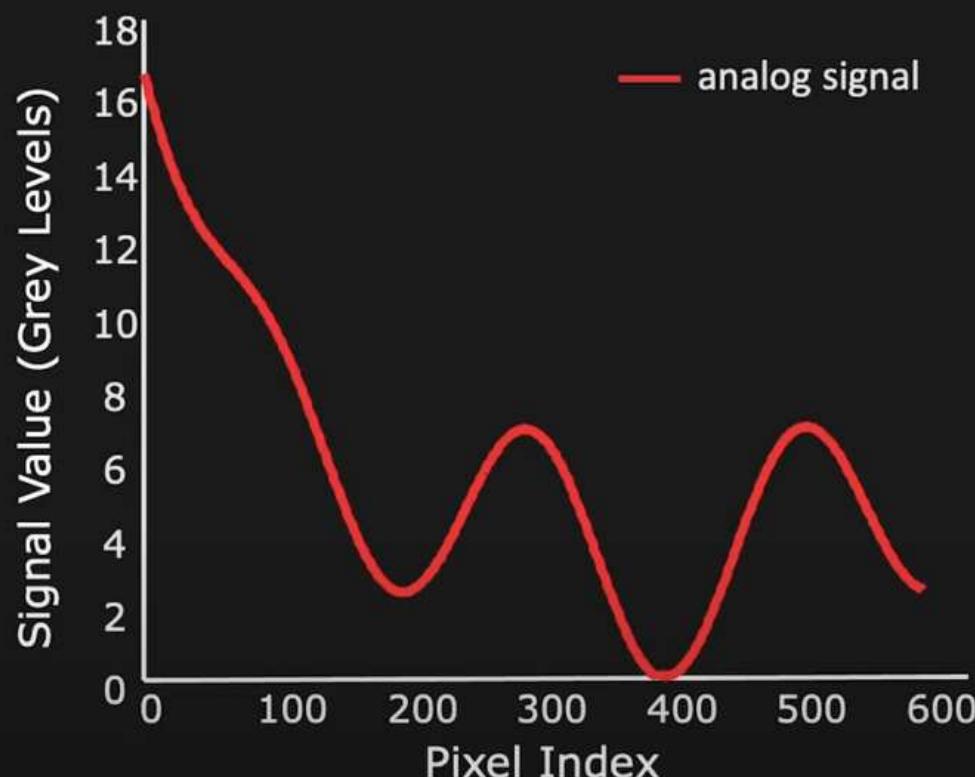


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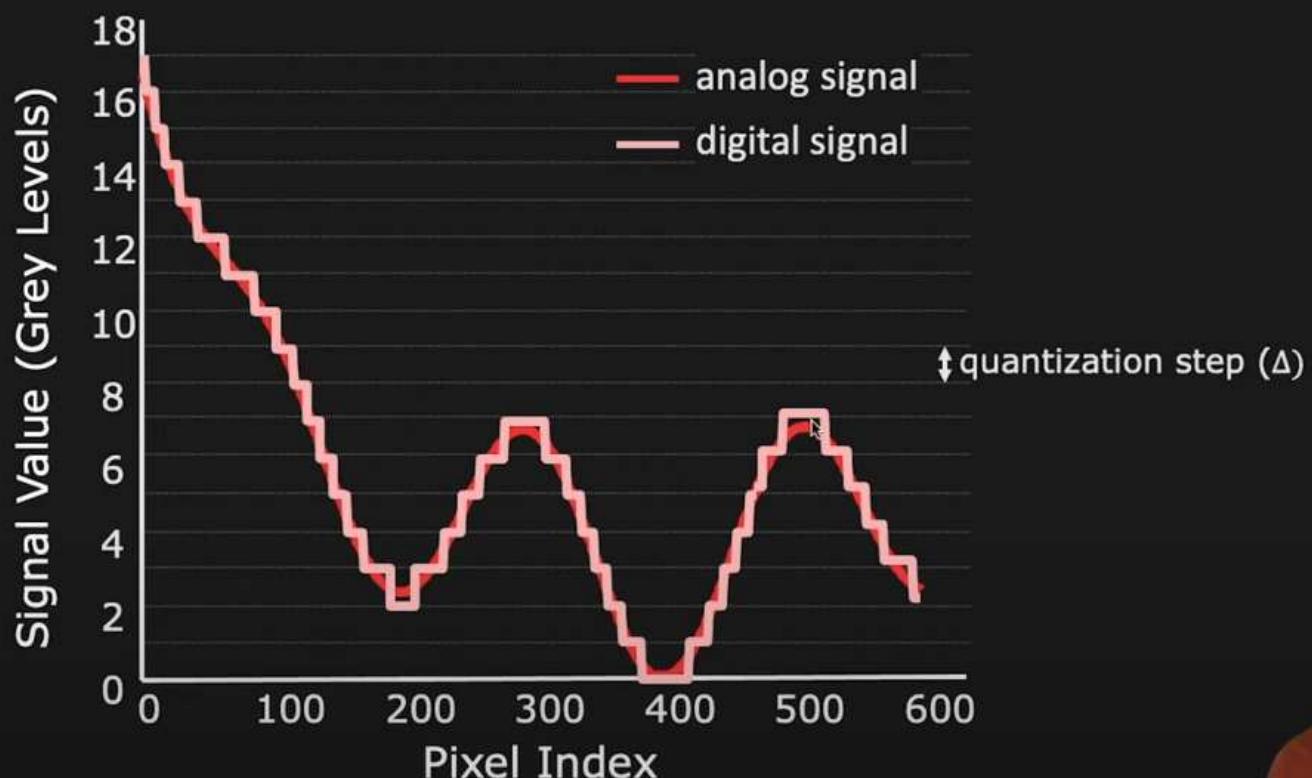


Depends on Sensor Quality (Scene Independent)

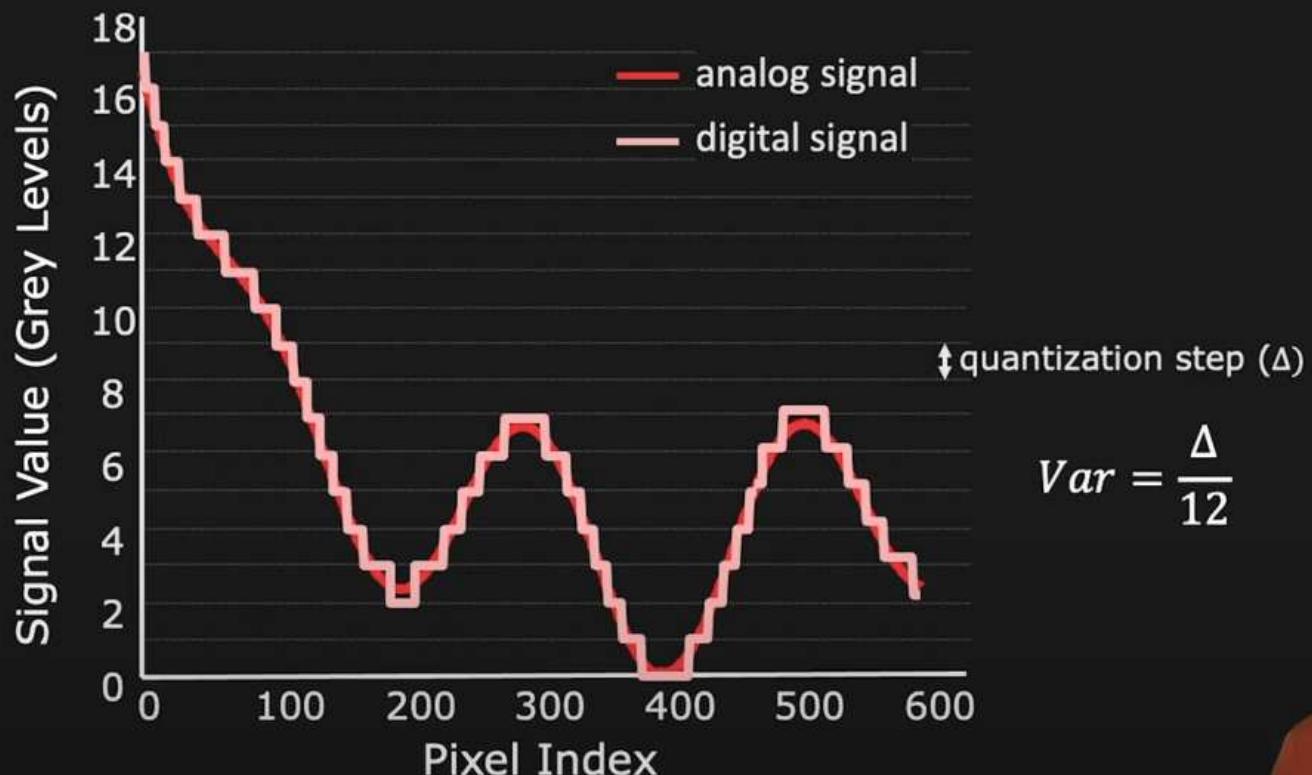
# Quantization Noise



# Quantization Noise



# Quantization Noise



Negligible in Modern Sensors  
Due to High Intensity Resolution (12-14 bits)



# Other Noise Sources

## Dark Current Noise (Thermal)



Follows Poisson Distribution

Significant Only for Long (>2 min) Exposures (Astronomy)

## Fixed Pattern Noise (Defective Pixels) Random Noise



Fixed Pattern Noise

Can be Reduced by  
Dark Frame Subtraction



# Sensor Dynamic Range

$$\text{Dynamic Range} = 20 \log \left( \frac{B_{max}}{B_{min}} \right) \text{ decibels (dB)}$$

$B_{max}$ : The maximum possible photon energy  
(full potential well)

$B_{min}$ : The minimum detectable photon energy  
(in the presence of noise)

Sensor	$B_{max}:B_{min}$	dB
Human Eye	1,000,000:1	120
HDR Display	200,000:1	106
Digital Camera	4096:1	72.2
Film Camera	2948:1	66.2
Digital Video	45:1	33.1



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# Quantum Efficiency

Incoming light can vary in **Wavelength ( $\lambda$ )**

Pixel Intensity  
(Electron Flux)



Incoming Light  
(Photon Flux)

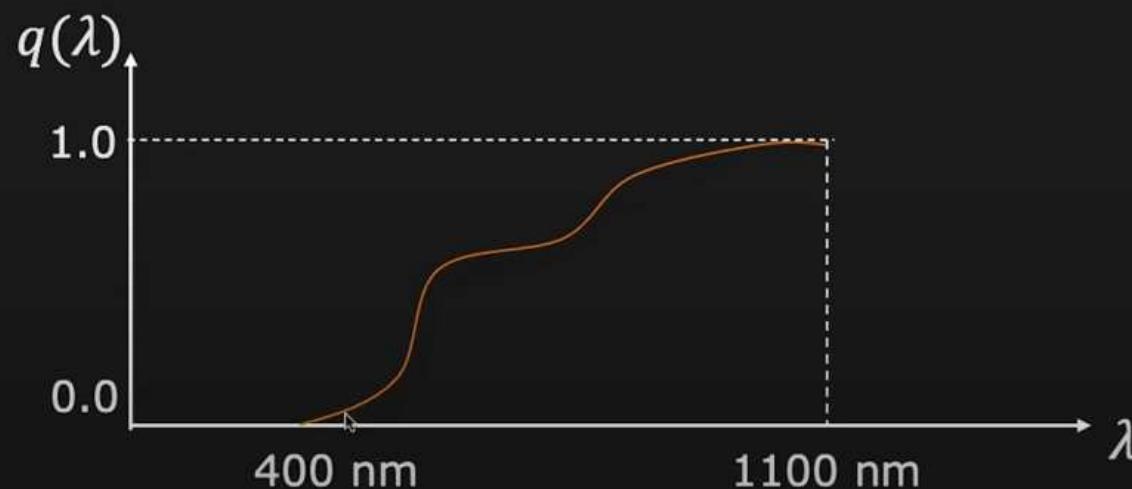
**Quantum Efficiency:**

$$q(\lambda) = \frac{\text{Electron Flux Generated}}{\text{Photon Flux of wavelength } \lambda}$$



# Quantum Efficiency of Silicon

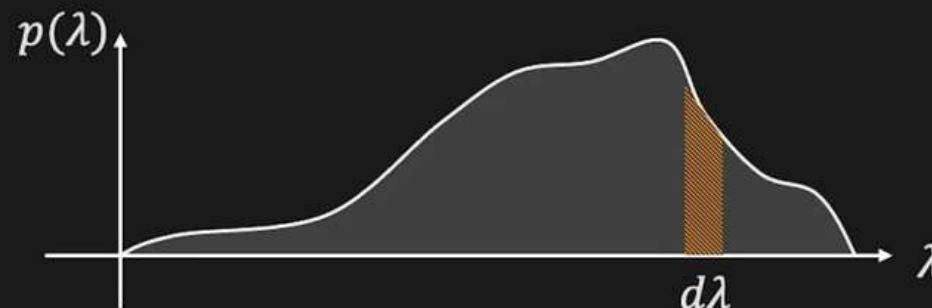
Silicon is: 
$$\begin{cases} \text{Transparent} & \text{for } \lambda > 1100 \text{ nm} \\ \text{Opaque} & \text{for } \lambda < 400 \text{ nm} \end{cases}$$



Assume Monochromatic Light  $\lambda = \lambda_i$  with flux  $p(\lambda_i)$ :

$$I = q(\lambda_i)p(\lambda_i)$$

What if incoming light has Spectral Distribution  $p(\lambda)$ ?



Flux (Energy) of Light with wavelength between  $\lambda$  and  $\lambda + d\lambda$  is  $p(\lambda)d\lambda$ . Therefore:

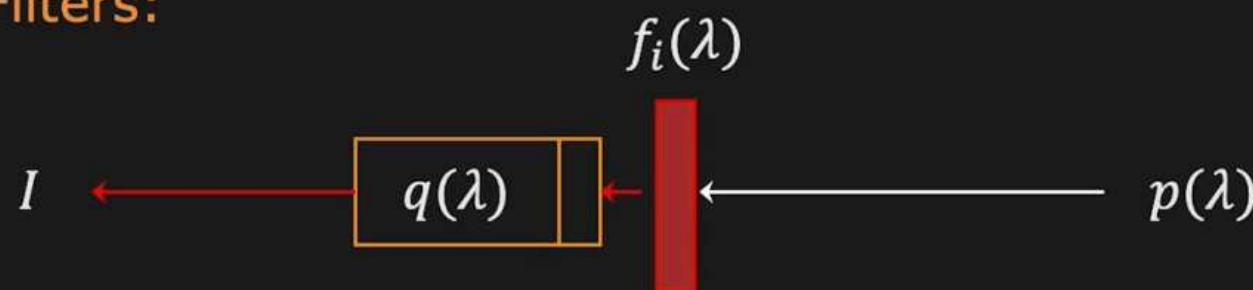
$$I = \int_0^{\infty} q(\lambda)p(\lambda)d\lambda$$



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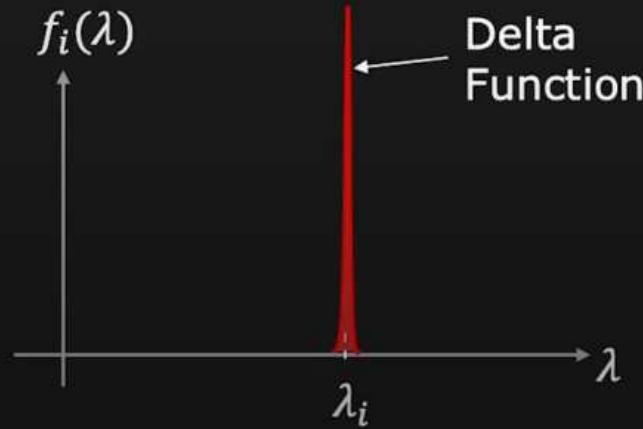
If we know  $I$  and  $q(\lambda)$ ,  
can we find  $p(\lambda)$ ?

Use Filters:



Let  $f_i(\lambda) = \delta(\lambda - \lambda_i)$

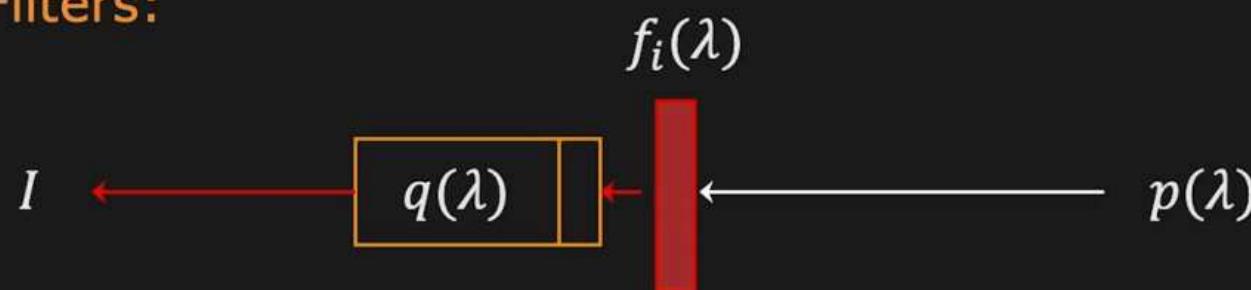
Note:  $\int_0^{\infty} \delta(\lambda - \lambda_i) d\lambda = 1$



$$I = \int_0^{\infty} q(\lambda)p(\lambda)d\lambda$$

If we know  $I$  and  $q(\lambda)$ ,  
can we find  $p(\lambda)$ ?

Use Filters:



$$I = \int_0^{\infty} q(\lambda)p(\lambda)f_i(\lambda)d\lambda = \int_0^{\infty} q(\lambda)p(\lambda)\delta(\lambda - \lambda_i)d\lambda$$

$$I = q(\lambda_i)p(\lambda_i)$$



How many filters do we need to recover  $p(\lambda)$ ? Infinite?

# What is “Color”?

Human Response to different wavelengths

Visible light:

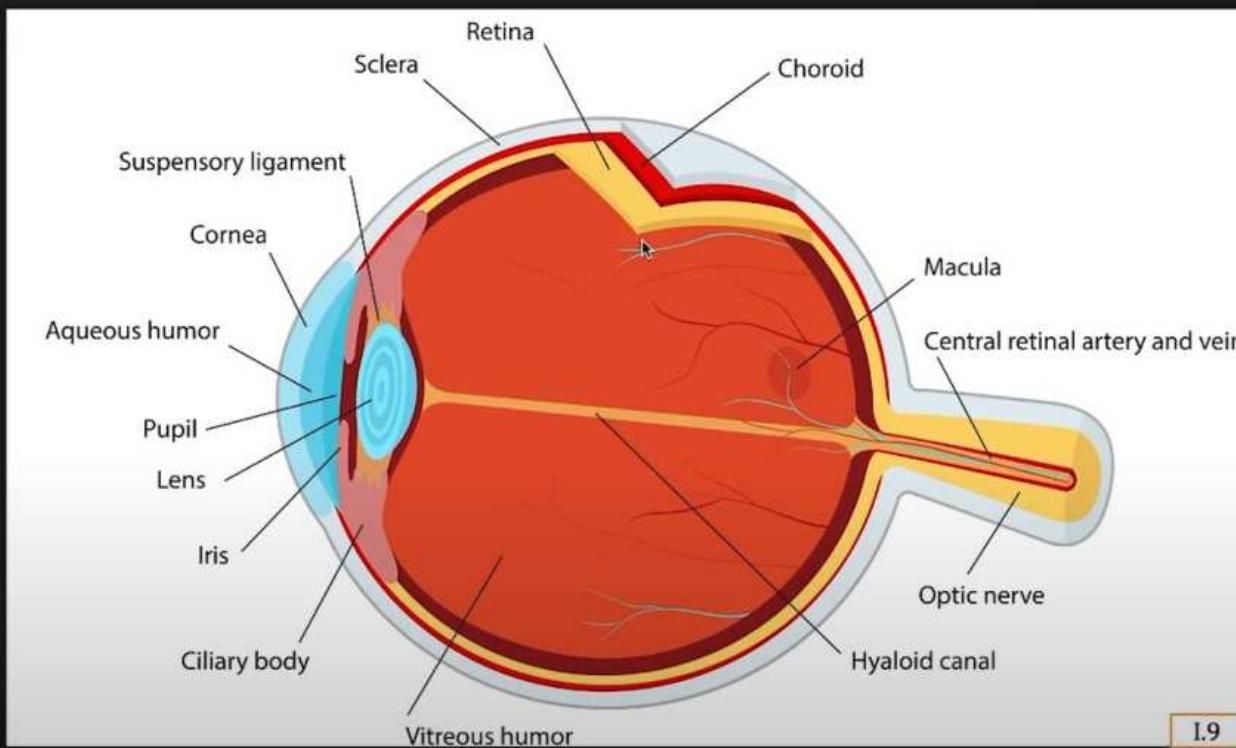


Do We recover spectral distribution  $p(\lambda)$ ?

Sensors in the human eye: Rods & Cones  
Neurochemical Sensors (3 types)



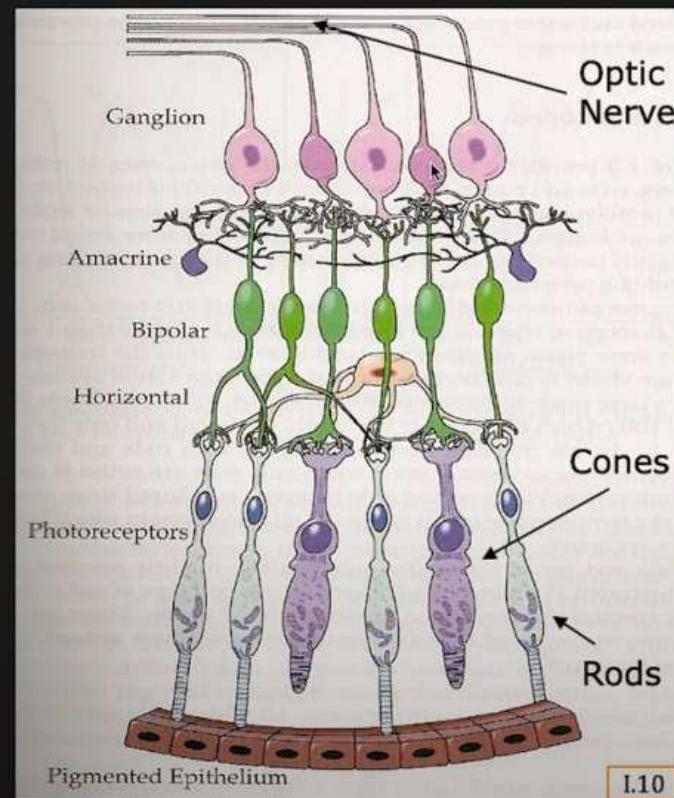
# The Human Eye



I.9



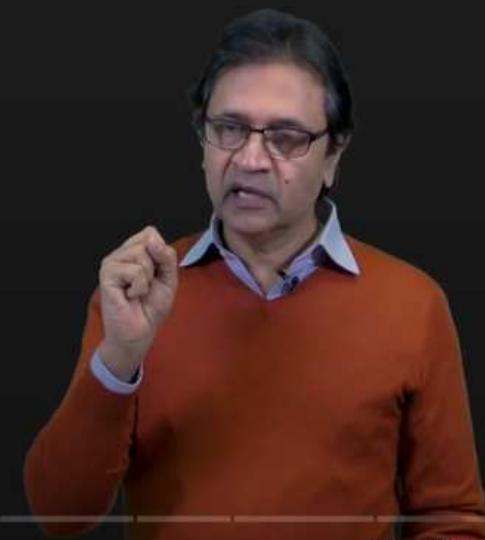
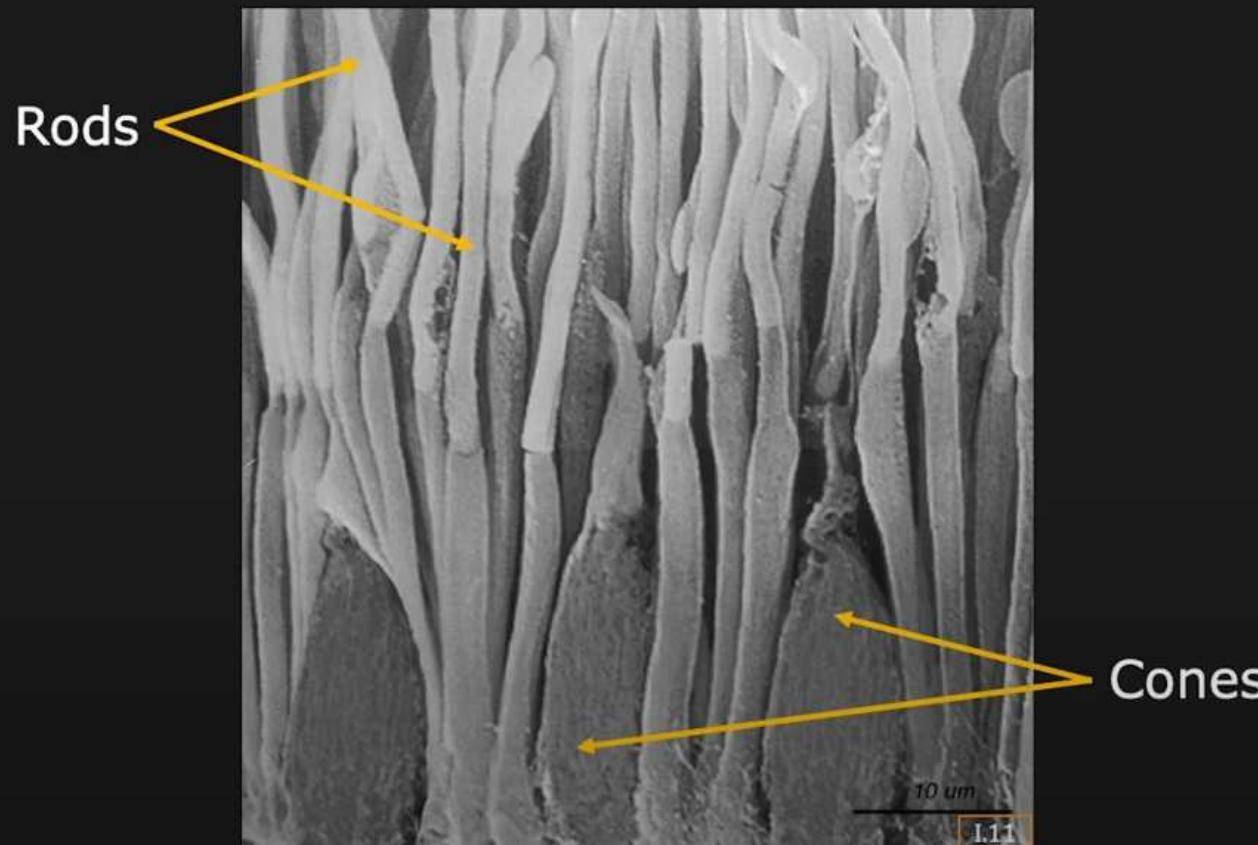
# A Cross-Section of the Retina



[Fairchild 200

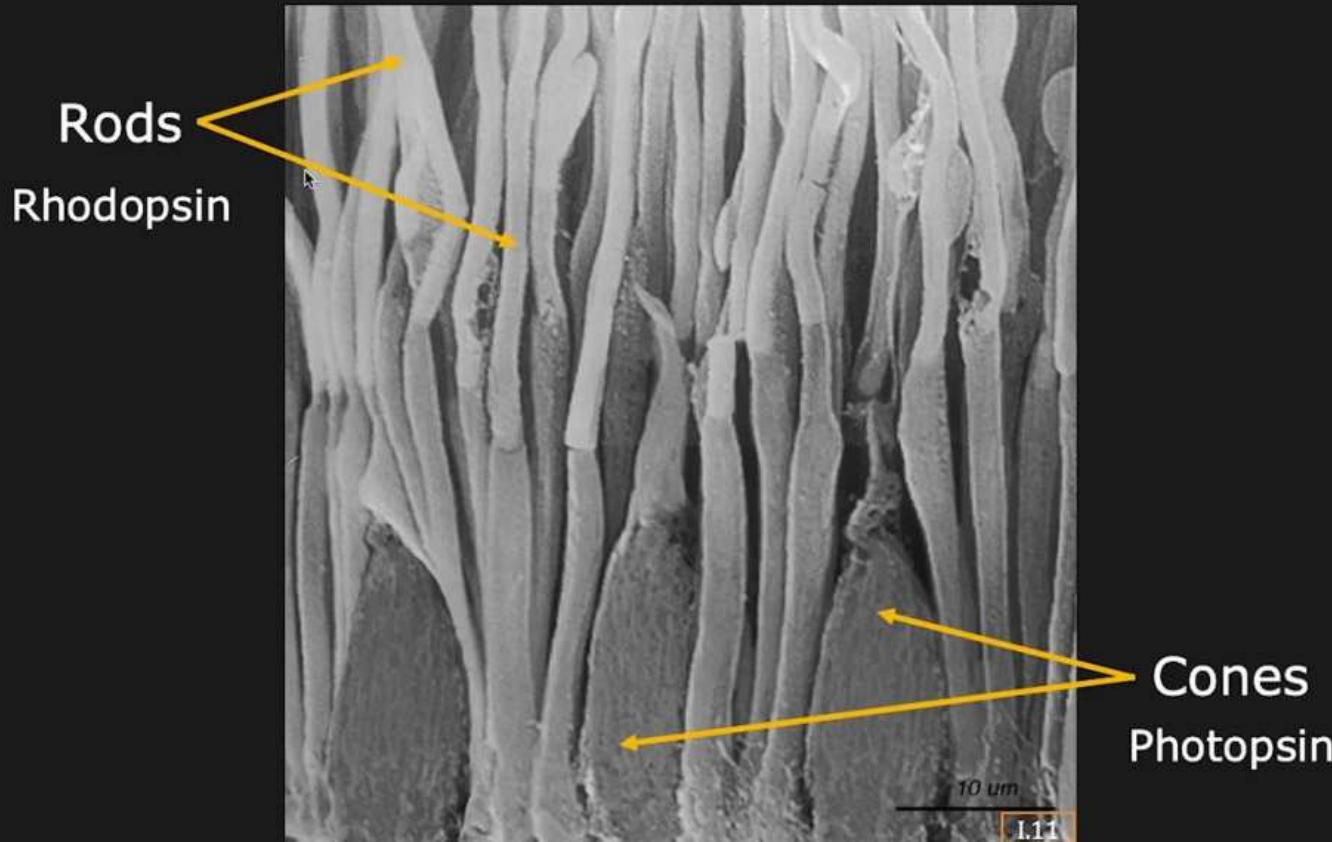


# The Eye's Pixels



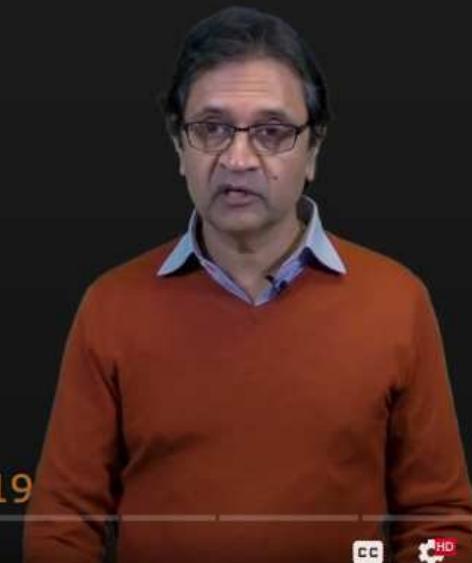
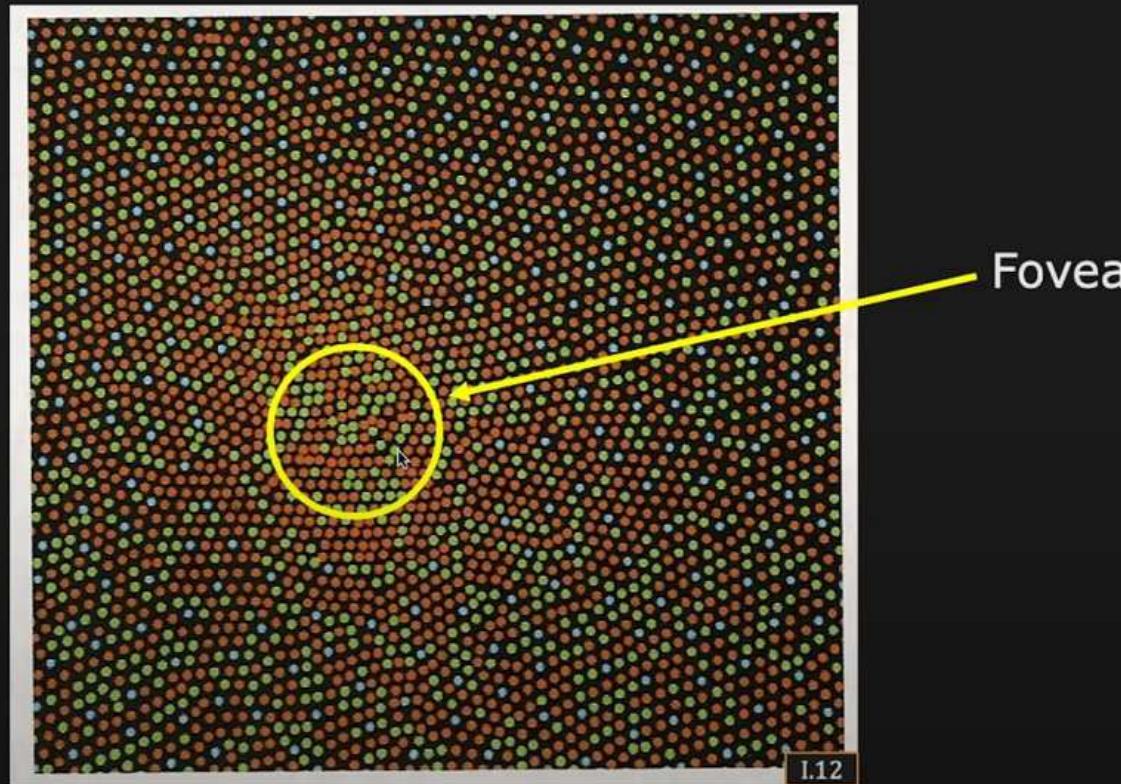
# The Eye's Pixels

---



# Distribution of Cones in Human Retina

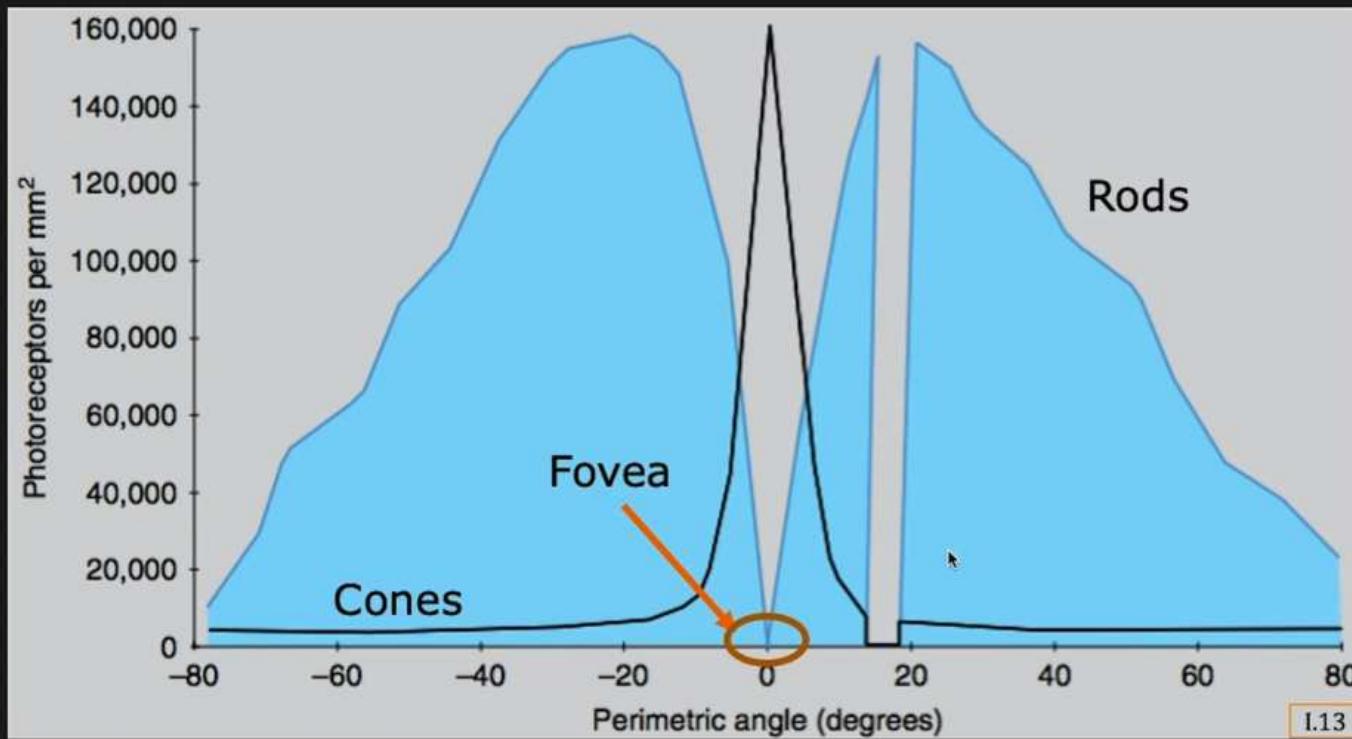
Three types of cones for sensing red, green, blue



[William 19]

# Resolution of Rods and Cones

On average, ~120M rods and ~7M cones per retina.



[Fairchild 2006]



# Tristimulus Values

Three Intensities ( $R$ ,  $G$ ,  $B$ ):

$$R = \int_{-\infty}^{\infty} h_r(\lambda) p(\lambda) d\lambda$$

$$G = \int_{-\infty}^{\infty} h_g(\lambda) p(\lambda) d\lambda$$

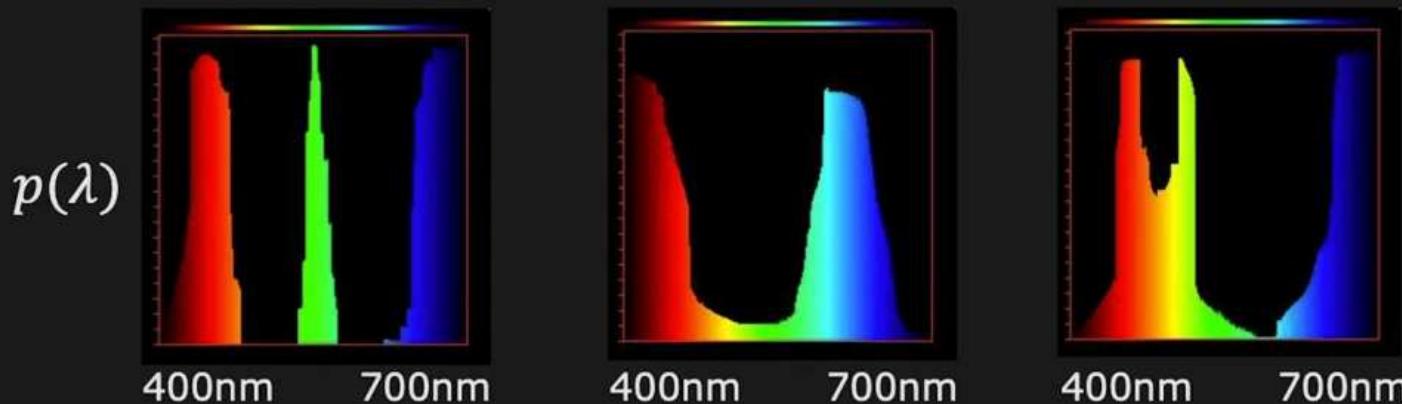
$$B = \int_{-\infty}^{\infty} h_b(\lambda) p(\lambda) d\lambda$$



# Metamers

Metamers: Different  $p(\lambda)$  that produce same ( $R, G, B$ )

For example, different spectra:

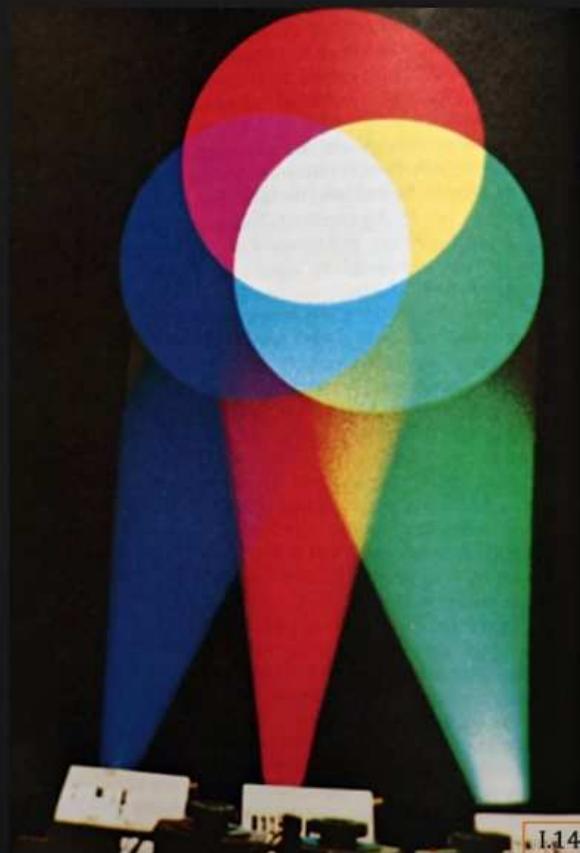


Same perceived color:

$$(R, G, B) = (115, 60, 108) =$$



# The Mixing of Colors



Human Sensation of nearly all colors can be produced using 3 wavelengths!

$$(\lambda_r, \lambda_g, \lambda_b) = (650, 530, 410) \text{ nm}$$

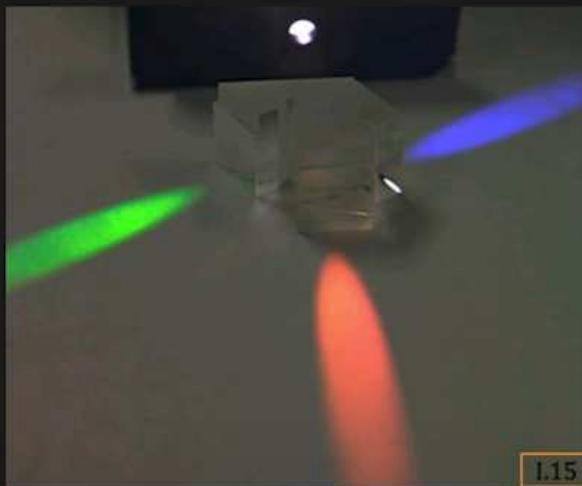
Hence, cameras and displays often use 3 filters:

(red, green, blue)

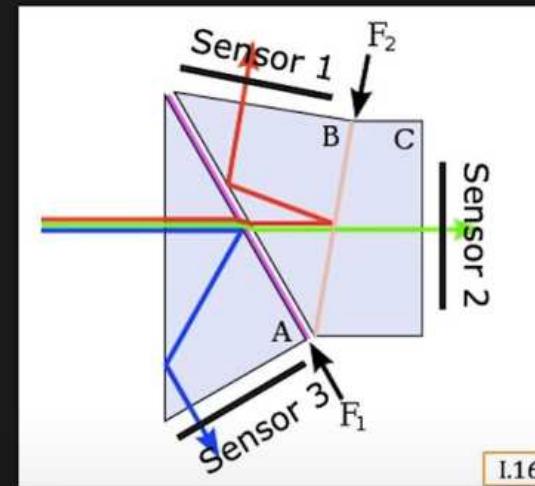


Young's Experiment on Color Mixture

# Sensing Color using Dichroic Prism



Dichroic Prism

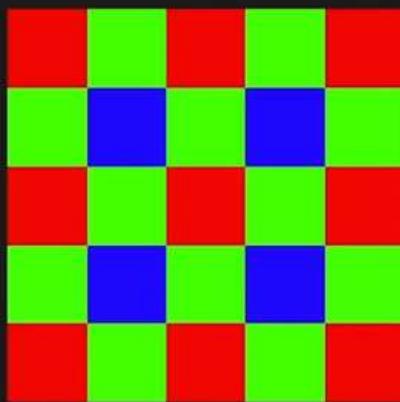


3-CCD Camera using  
Dichroic Prism

Each Sensor Detects One Color



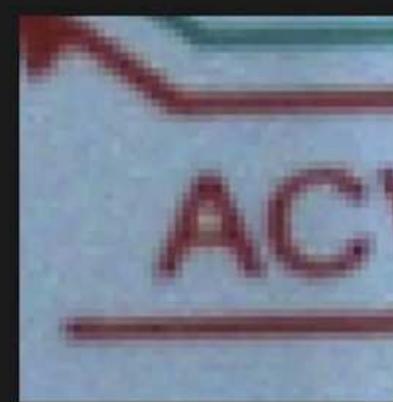
# Sensing Color Using Color Mosaic



Bayer Pattern  
(Color Filter Mosaic)

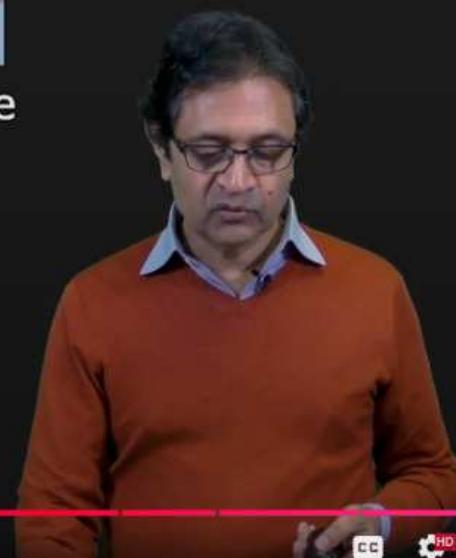


## Raw Image



## Interpolated Image

## Color Filled in by Interpolation (Demosaicing)



# Camera Response Function $f(\cdot)$

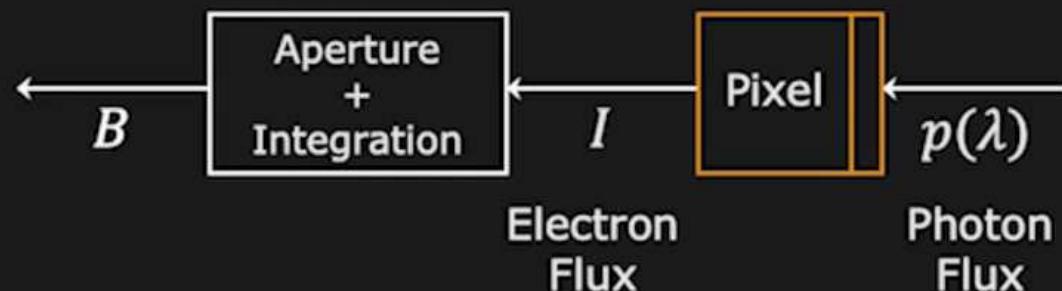


Image Brightness:

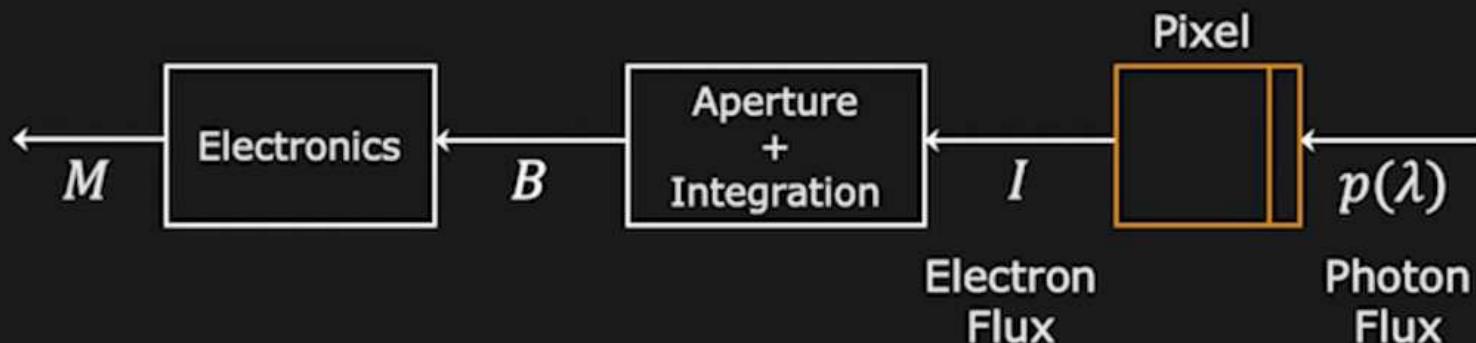
$$B = I \cdot \left( \frac{\pi d^2}{4} \right) \cdot t$$

Aperture Area (diameter  $d$ )      Integration Time



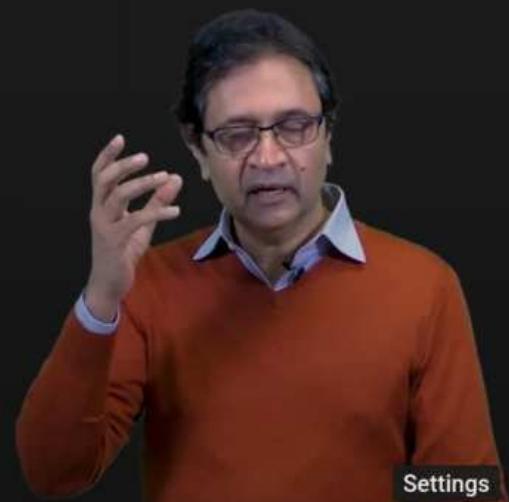
Settings

# Camera Response Function $f(\cdot)$

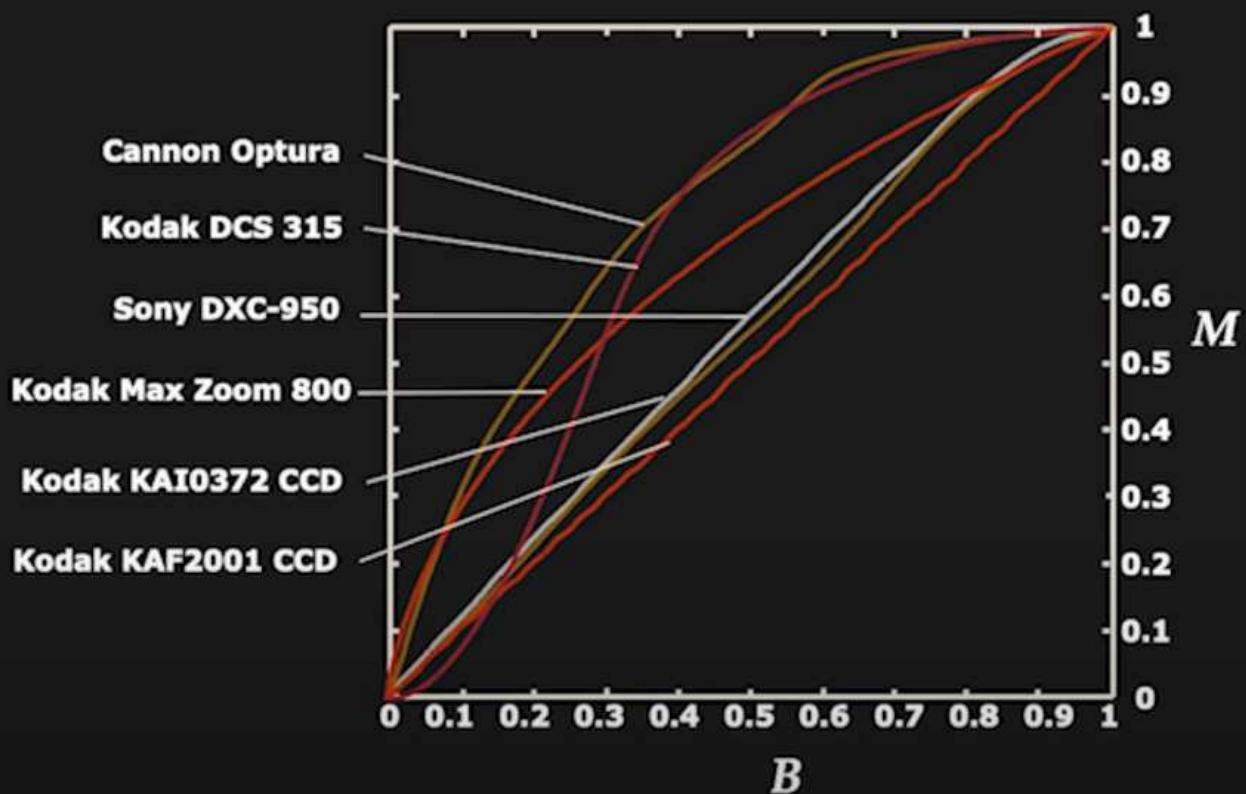


Measured Brightness:  $M = f(B) = f(I \cdot e)$

where Exposure,  $e = \left(\frac{\pi d^2}{4}\right) \cdot t$



# Camera Response Function $f(\cdot)$



"Gamma Curves"

[Grossberg 2]

Settings

# Radiometric Calibration: Finding $f(\cdot)$

Calibration using a chart:

1. Patches with known reflectance (when uniformly lit)
2. Fit linear segments or curve

Macbeth Chart

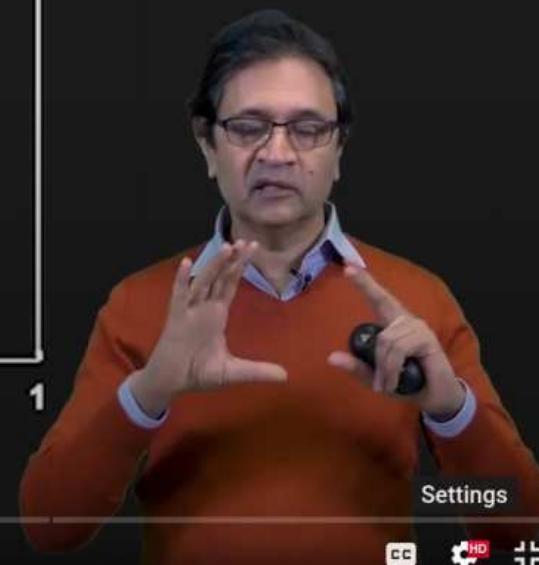
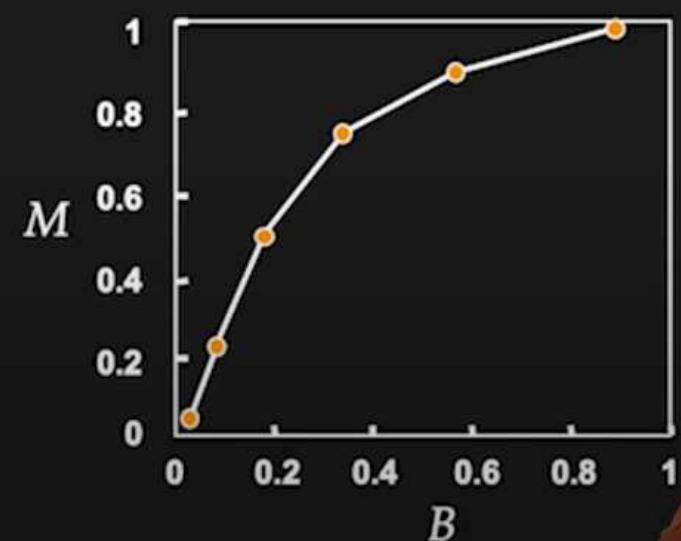
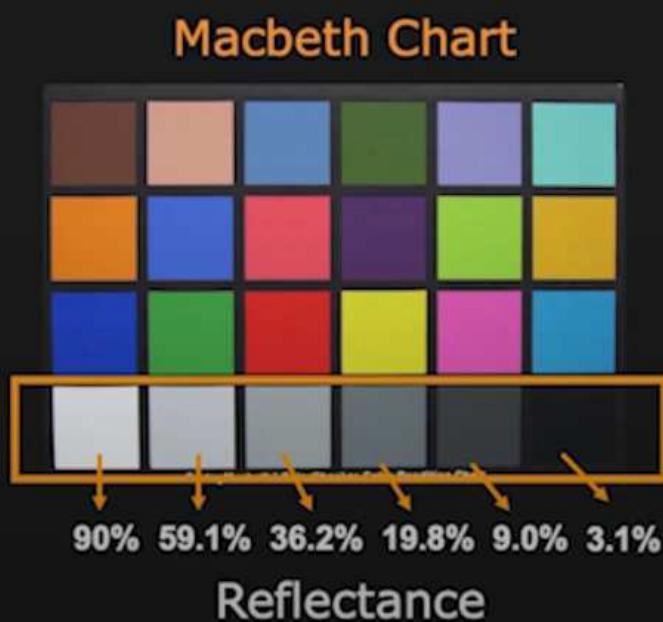


Settings

# Radiometric Calibration: Finding $f(\cdot)$

Calibration using a chart:

1. Patches with known reflectance (when uniformly lit)
2. Fit linear segments or curve



Settings

# High Dynamic Range: Multiple Exposures

Assume Camera Response  $f(\cdot)$  is Linear



with  $e_0$



$e_1$



$e_2$



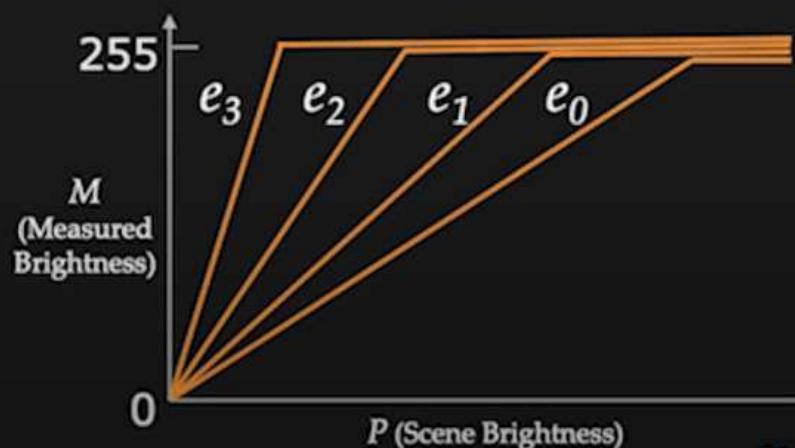
$e_3$

$$M_0 = \min(e_0 \cdot P, 255)$$

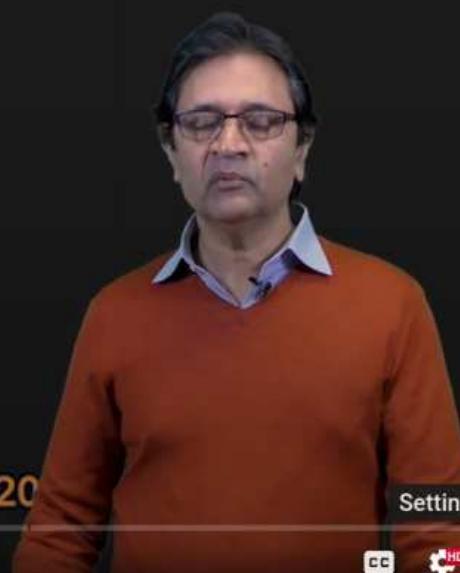
$$M_1 = \min(e_1 \cdot P, 255)$$

$$M_2 = \min(e_2 \cdot P, 255)$$

$$M_3 = \min(e_3 \cdot P, 255)$$



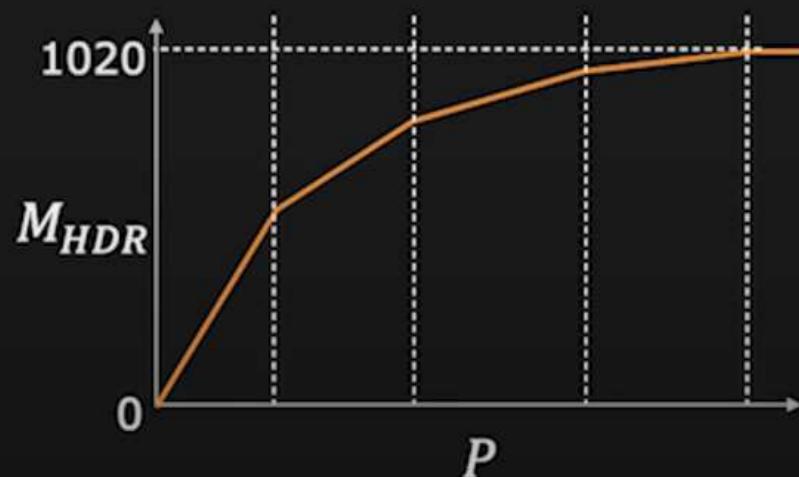
[Nayar 20]



# High Dynamic Range: Multiple Exposures

Aggregate Image:  $M_{HDR} = M_0 + M_1 + M_2 + M_3$

Camera Response  $f(\cdot)$  for Aggregate Image:



[Nayar 2000]

Settings

# The Motion Problem

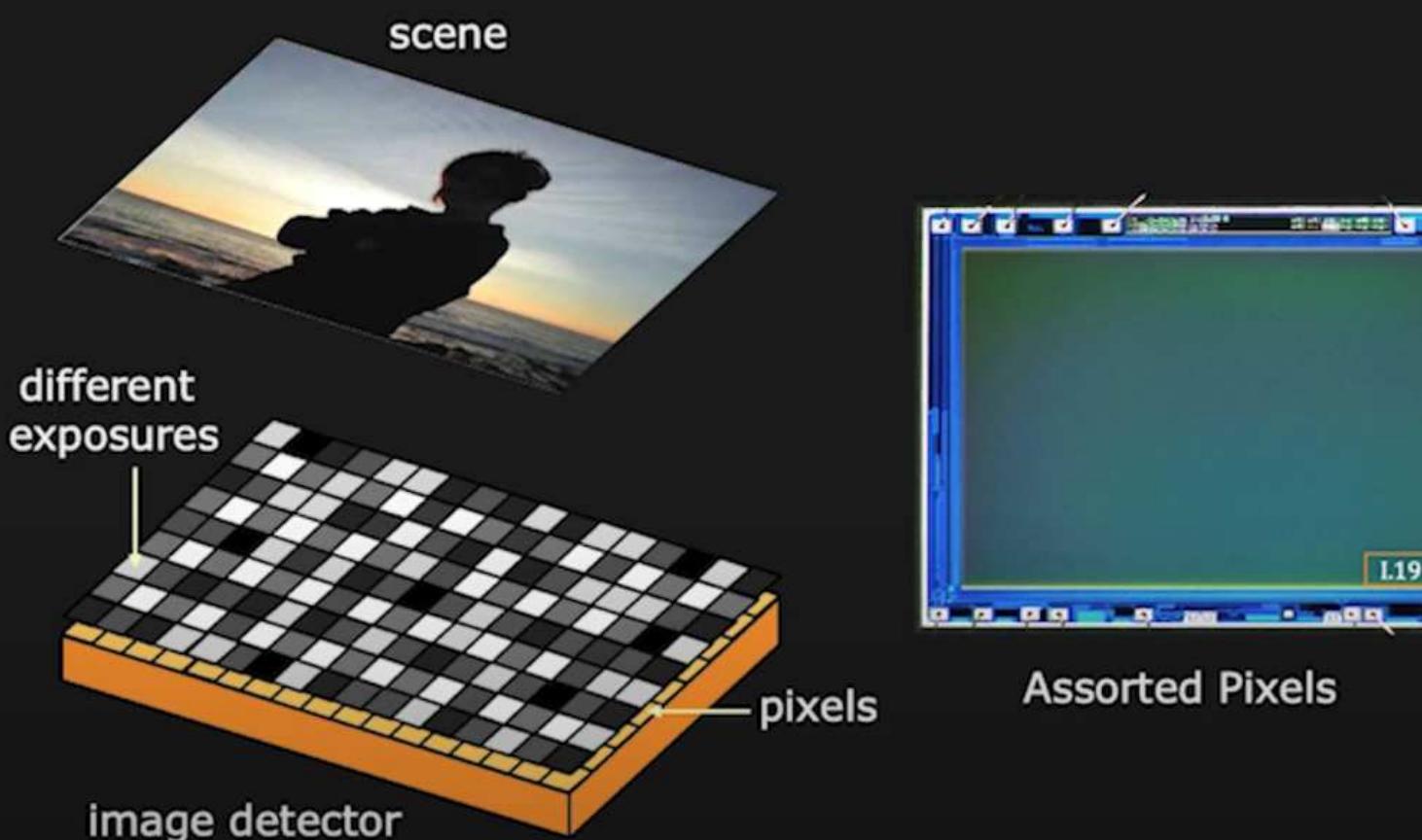


iPhone 4 HDR Mode



Settings

# Single Shot HDR Imaging



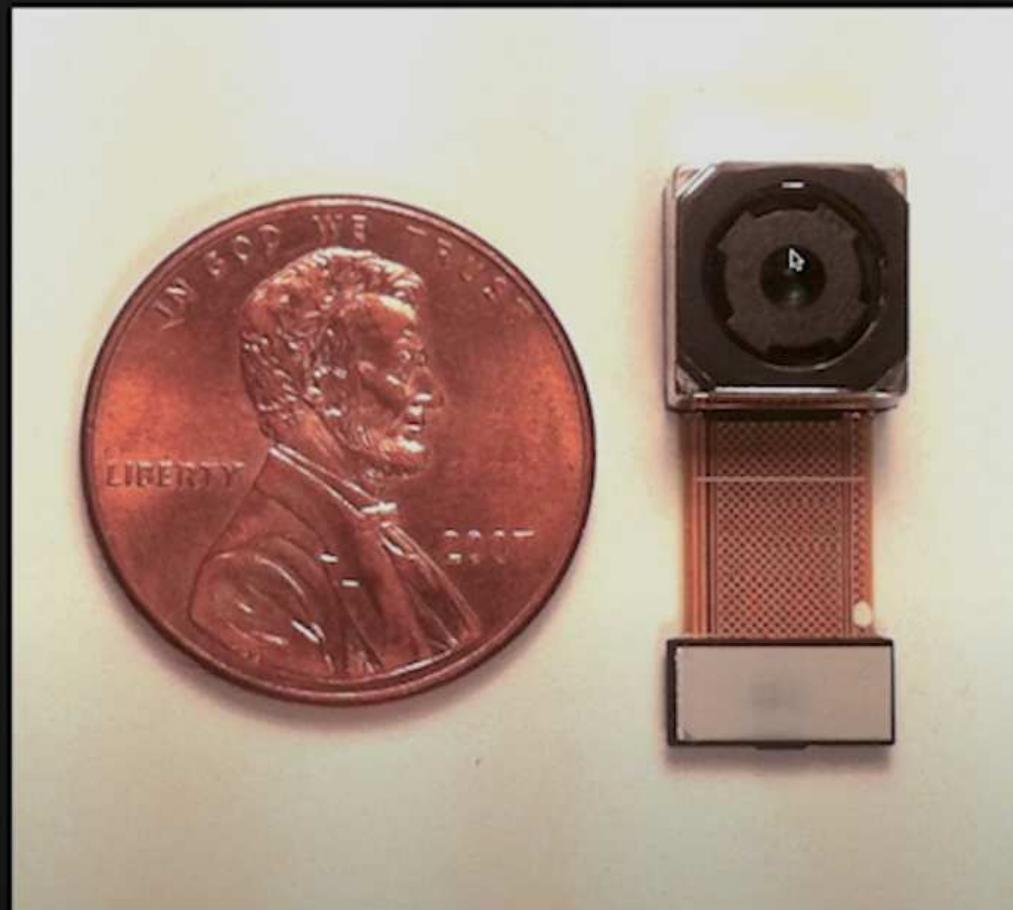
[Nayar]

Settings

## Good Old Camera

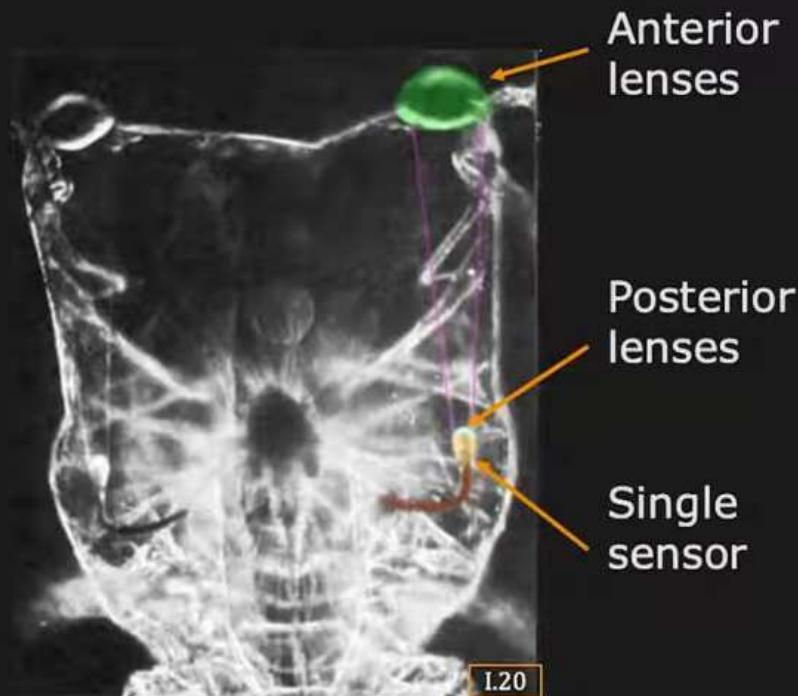


Settings



Settings

# The Curious Eye of Copilia



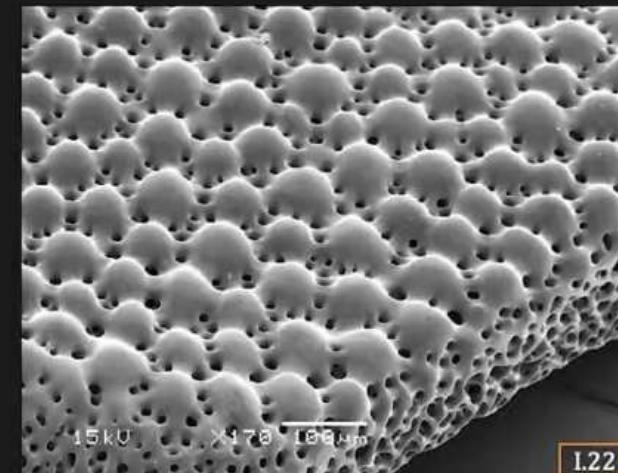
[Gregory D.

# It's All Eyes



I.21

Brittle Star  
(*Ophiocoma wendtii*)



I.22

SEM image of the lenses

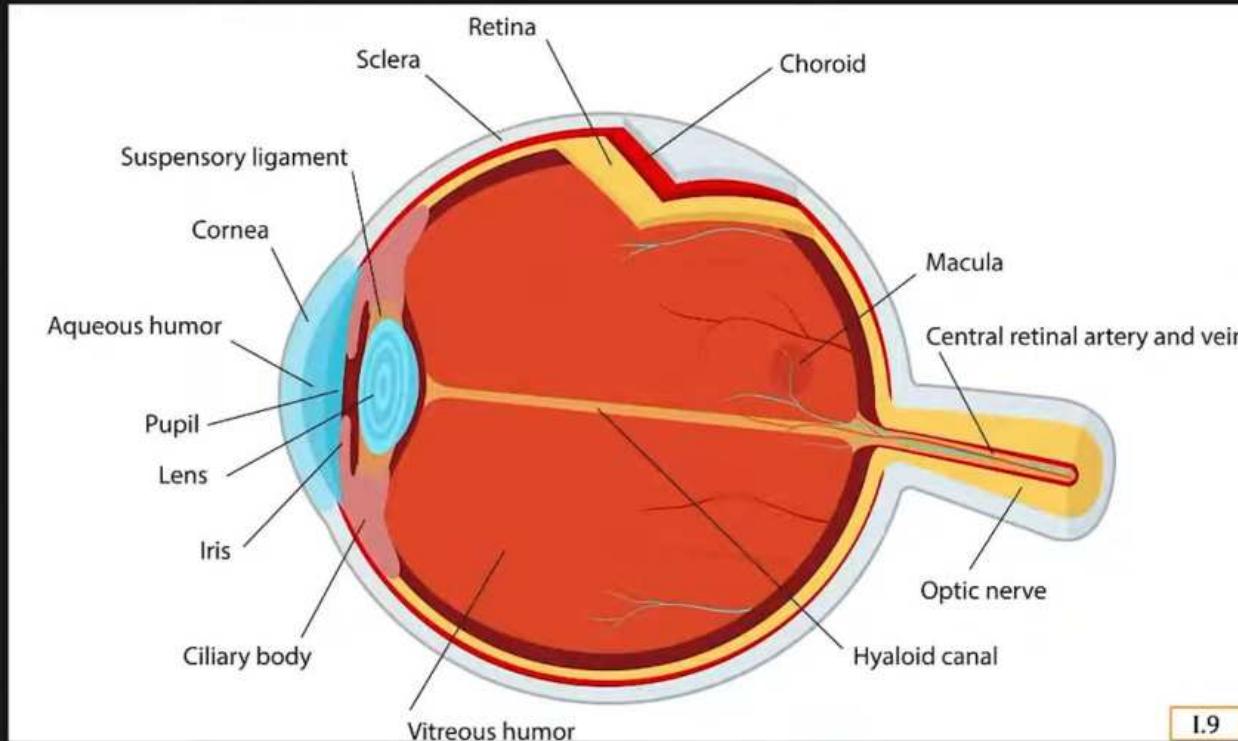
The entire body is covered with lenses

[Aizenberg]

# Camouflage in Octopus



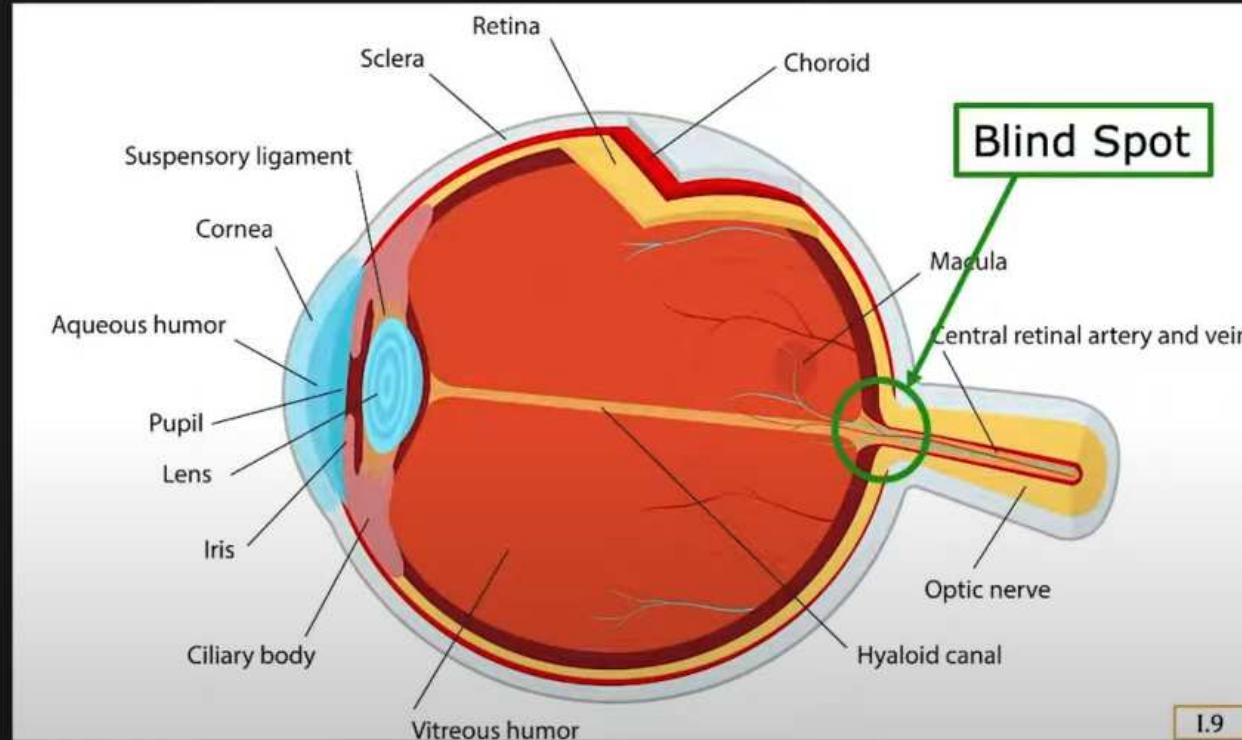
# Blind Spot



1.9



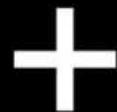
# Blind Spot



1.9



# Blind Spot



Close your left eye and look at the cross with your right eye.  
Move your head slowly towards or away from the image,  
until at some point the white disc disappears.



# Binary Images

Binary Image: Can have only two values (0 or 1).  
Simple to process and analyze.



# Making Binary Images

---

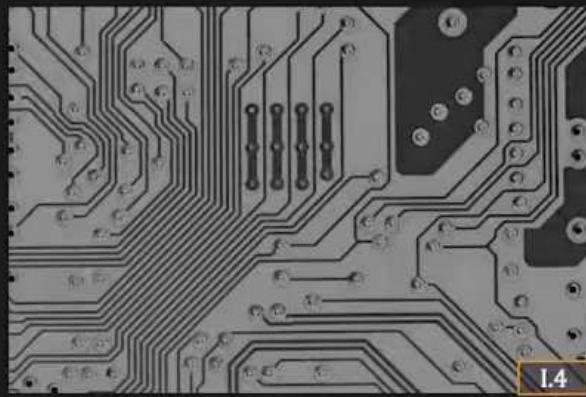
**Binary Image  $b(x,y)$ :** Usually obtained from Gray-level image  $g(x,y)$  by **Thresholding**.

**Characteristic Function:**

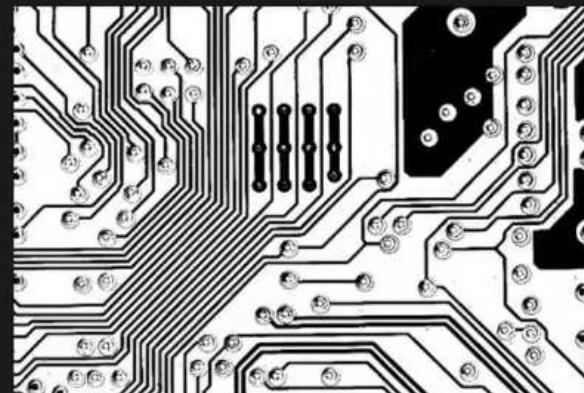
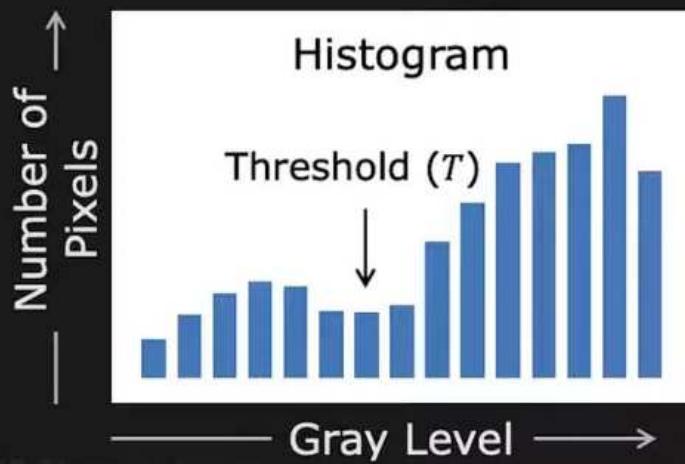
$$b(x,y) = \begin{cases} 0, & g(x,y) < T \\ 1, & g(x,y) \geq T \end{cases}$$



# Selecting a Threshold ( $T$ )



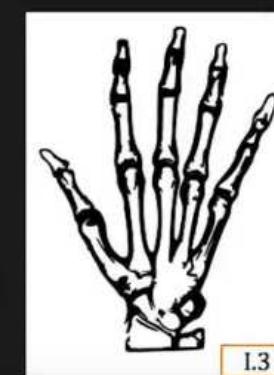
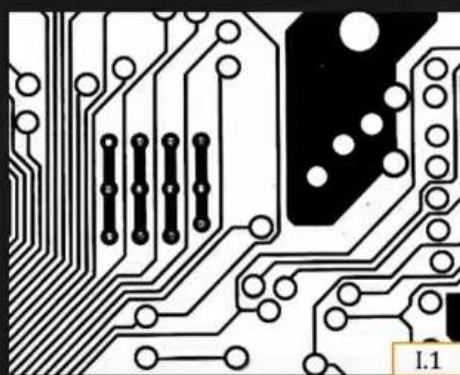
Gray Image  $g(x,y)$



Binary Image  $b(x,y)$



# Examples of Binary Images



# Capturing a Binary Image

---



# Capturing a Binary Image

---



Backlighting



# Binary Images

Binary Image: Can have only two values (0 or 1).  
Simple to process and analyze.

## Topics:

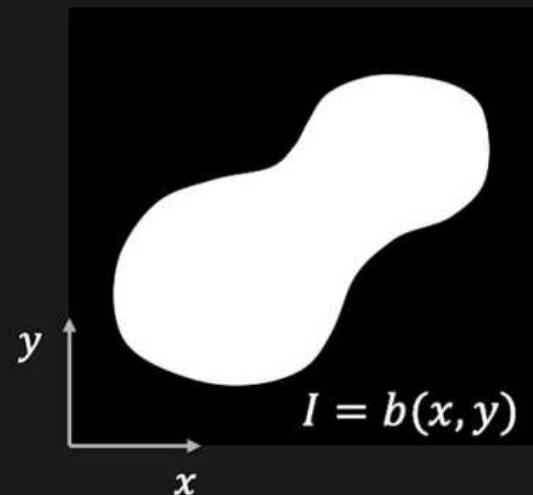
- (1) Geometric Properties
- (2) Segmenting Binary Images
- (3) Iterative Modification



# Geometric Properties of Binary Images

Assume:

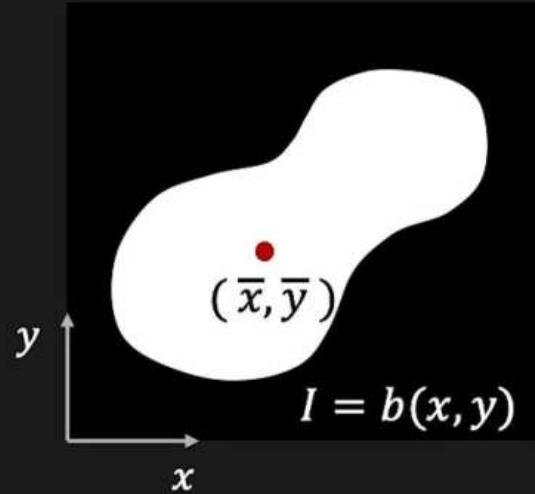
- $b(x, y)$  is continuous
- Only one object



# Area and Position

Area: (Zeroth Moment)

$$A = \iint_I b(x, y) dx dy$$



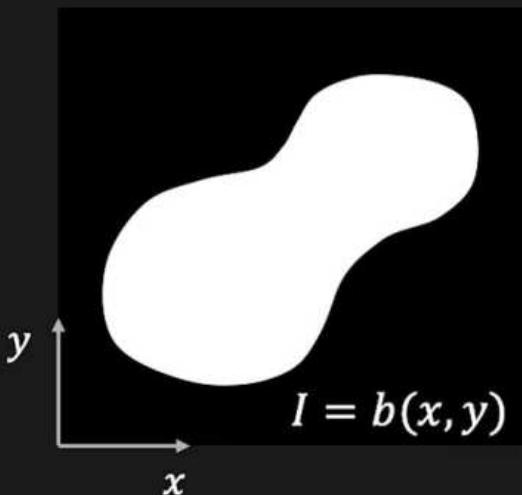
Position: Center of Area (First Moment)

$$\bar{x} = \frac{1}{A} \iint_I x b(x, y) dx dy , \quad \bar{y} = \frac{1}{A} \iint_I y b(x, y) dx dy$$



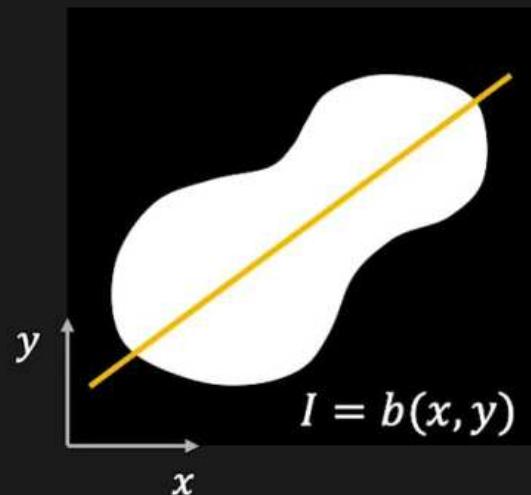
# Orientation

Difficult to define!



# Orientation

Difficult to define!



Use: Axis of Least Second Moment

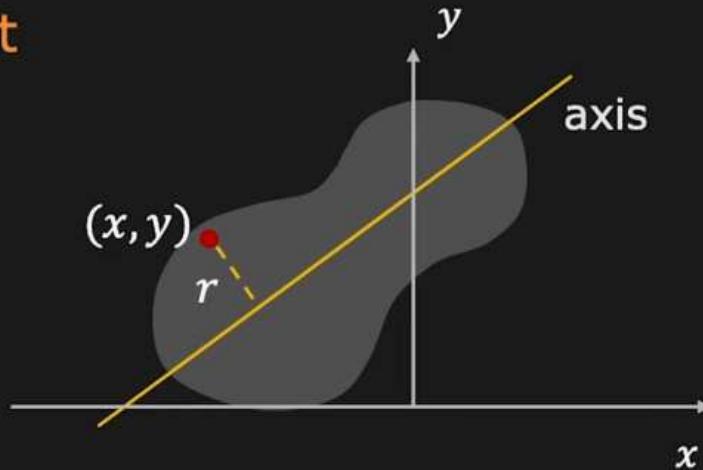


# Orientation

---

Axis of Least Second Moment  
minimizes:

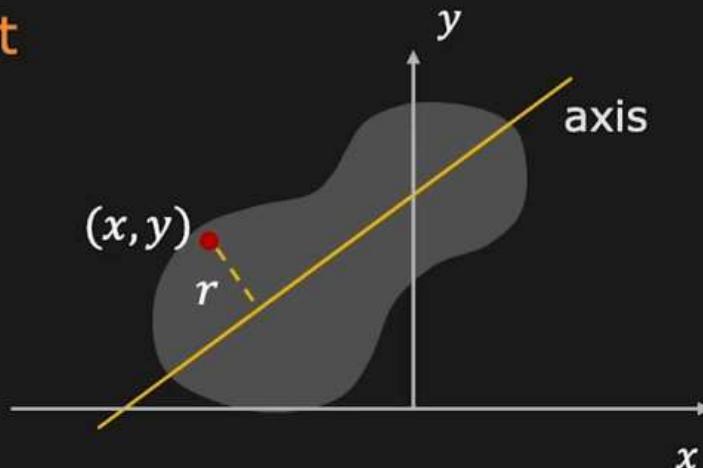
$$E = \iint_I r^2 b(x, y) dx dy$$



# Orientation

Axis of Least Second Moment  
minimizes:

$$E = \iint_I r^2 b(x, y) dx dy$$



Which equation to use for axis?

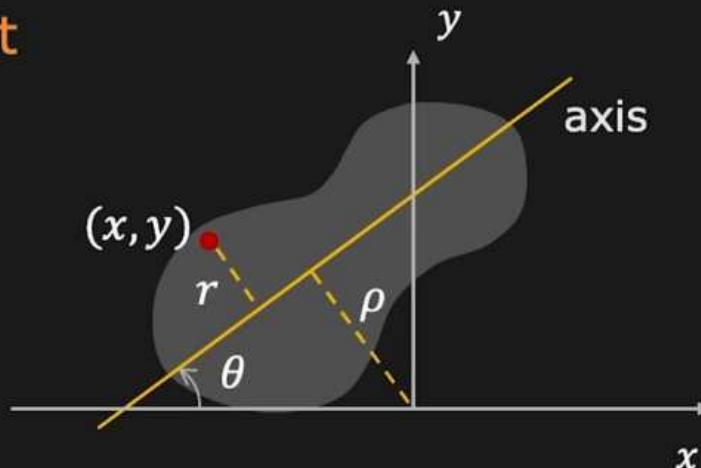
$$y = mx + b ? \quad -\infty \leq m \leq \infty$$



# Orientation

Axis of Least Second Moment  
minimizes:

$$E = \iint_I r^2 b(x, y) dx dy$$



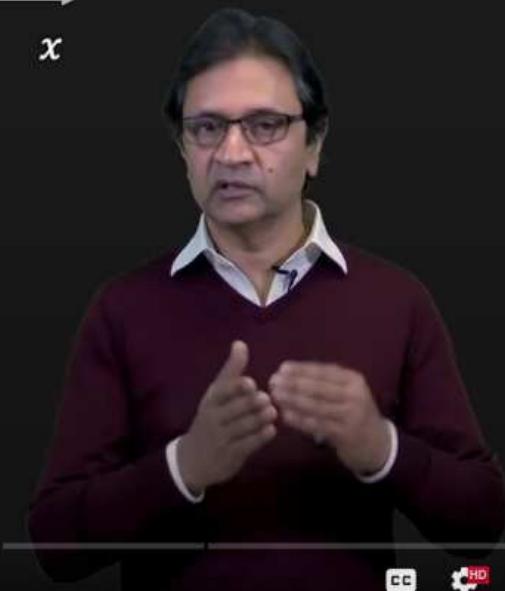
Which equation to use for axis?

$$y = mx + b ? \quad -\infty \leq m \leq \infty$$

Use:

$$x \sin \theta - y \cos \theta + \rho = 0$$

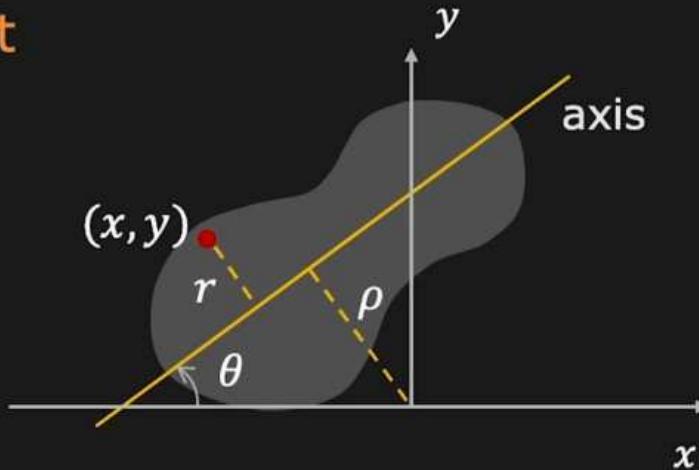
$\rho, \theta$  are finite



# Orientation

Axis of Least Second Moment  
minimizes:

$$E = \iint_I r^2 b(x, y) dx dy$$



Which equation to use for axis?

$$y = mx + b ? \quad -\infty \leq m \leq \infty$$

Use:

$$x \sin \theta - y \cos \theta + \rho = 0$$

$\rho, \theta$  are finite

Find  $\rho$  and  $\theta$  that minimize  $E$  for given  $b(x, y)$

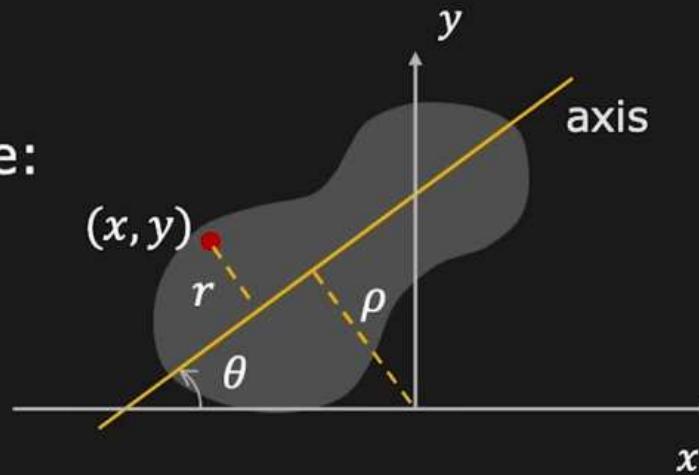


# Distance Between Point and Line

Given a line  $ax + by + c = 0$

Distance of point  $(x, y)$  from line:

$$r = \left| \frac{ax + by + c}{\sqrt{a^2 + b^2}} \right|$$



Similarly, given axis  $x \sin \theta - y \cos \theta + \rho = 0$

Distance of point  $(x, y)$  from axis:

$$r = \left| \frac{x \sin \theta - y \cos \theta + \rho}{\sqrt{\sin^2 \theta + \cos^2 \theta}} \right|$$

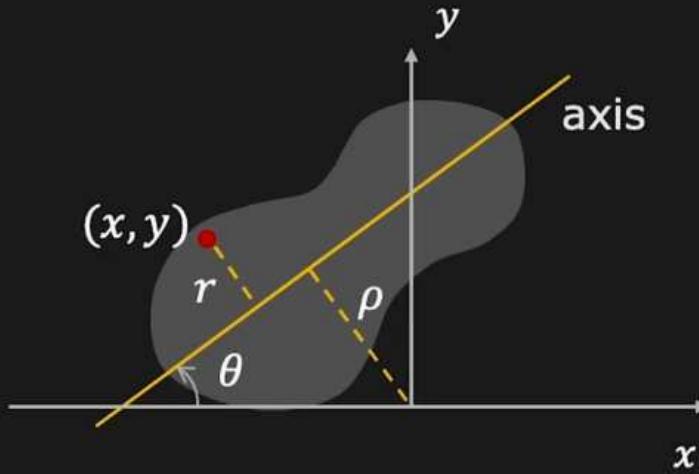
$$r = |x \sin \theta - y \cos \theta + \rho|$$



# Minimizing Second Moment

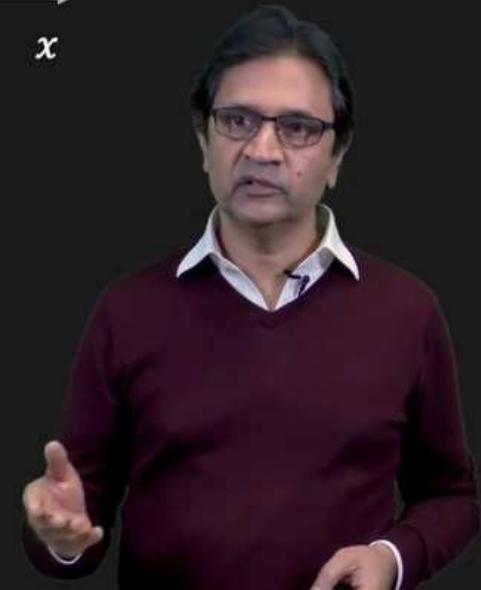
Axis of Least Second Moment  
minimizes:

$$E = \iint_I r^2 b(x, y) dx dy$$



So, minimize:

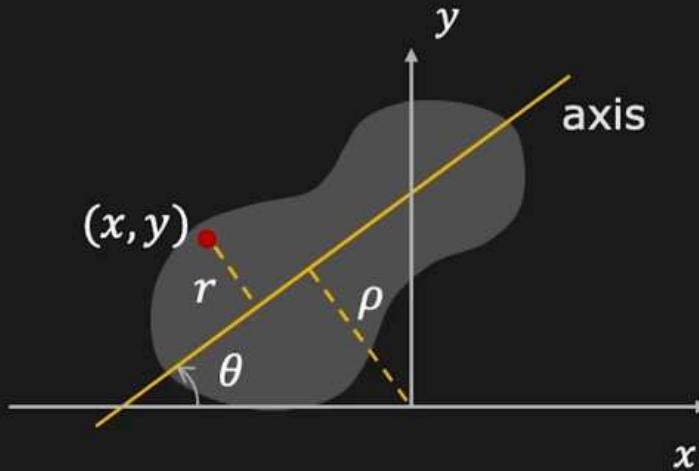
$$E = \iint_I (x \sin \theta - y \cos \theta + \rho)^2 b(x, y) dx dy$$



# Minimizing Second Moment

Axis of Least Second Moment  
minimizes:

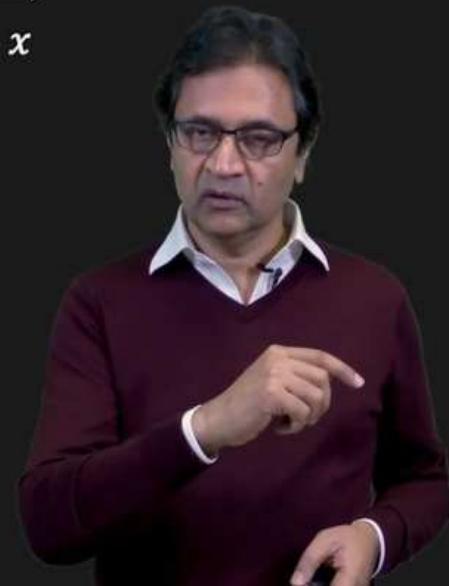
$$E = \iint_I r^2 b(x, y) dx dy$$



So, minimize:

$$E = \iint_I (x \sin \theta - y \cos \theta + \rho)^2 b(x, y) dx dy$$

Using  $\frac{\partial E}{\partial \rho} = 0$  we get:  $A(\bar{x} \sin \theta - \bar{y} \cos \theta + \rho) = 0$

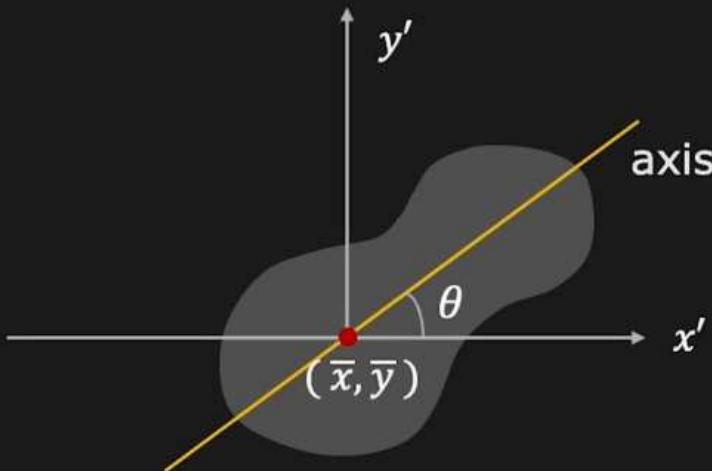


# Shift the Coordinate System

Change coordinates:

$$x' = x - \bar{x}, y' = y - \bar{y}$$

$$\begin{aligned}x \sin \theta - y \cos \theta + \rho \\= x' \sin \theta - y' \cos \theta\end{aligned}$$



Therefore, we can rewrite  $E$  as:

$$E = a \sin^2 \theta - b \sin \theta \cos \theta + c \cos^2 \theta$$

where: 
$$\begin{cases} a = \iint_{I'} (x')^2 b(x, y) dx' dy' \\ b = 2 \iint_{I'} (x'y') b(x, y) dx' dy' \\ c = \iint_{I'} (y')^2 b(x, y) dx' dy' \end{cases} \quad (a, b, c \text{ are easy to compute})$$



## Finally, Minimize $E$

Using  $\frac{dE}{d\theta} = (a - c) \sin 2\theta - b \cos 2\theta = 0$  we get:

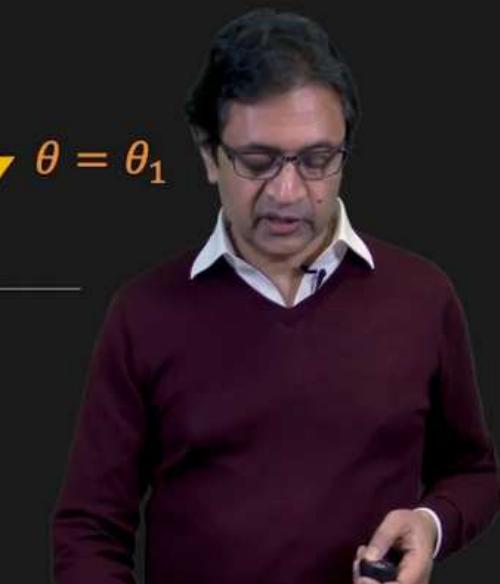
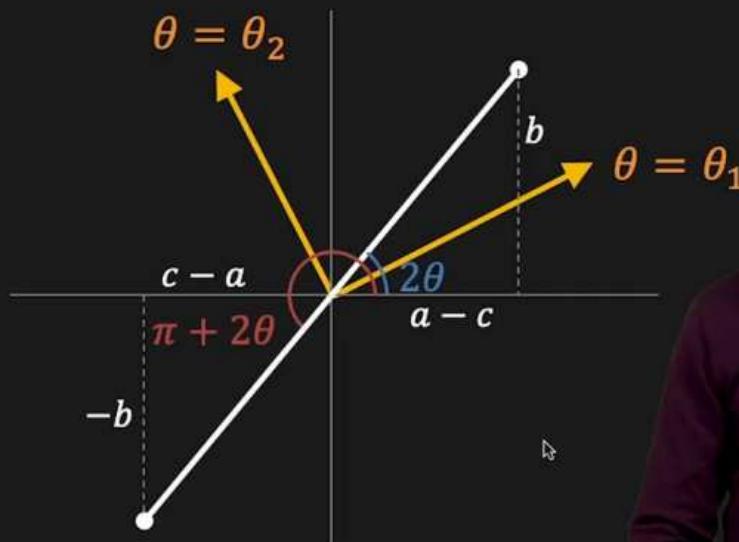
$$\tan 2\theta = \frac{b}{a - c}$$

We know that:  $\tan 2\theta = \tan(2\theta + \pi) = \frac{-b}{c - a}$

$\theta$  has two solutions.

1.  $\theta = \theta_1$
2.  $\theta = \theta_2 = \theta_1 + \frac{\pi}{2}$

One gives Minimum of  $E$   
and the other Maximum of  $E$



## Finally, Minimize $E$

Using  $\frac{dE}{d\theta} = (a - c) \sin 2\theta - b \cos 2\theta = 0$  we get:

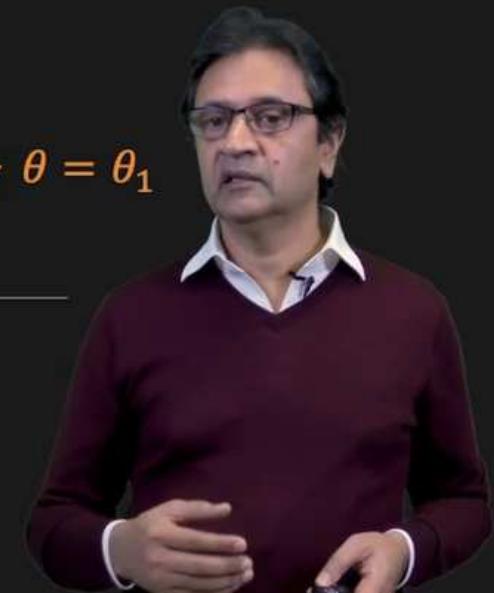
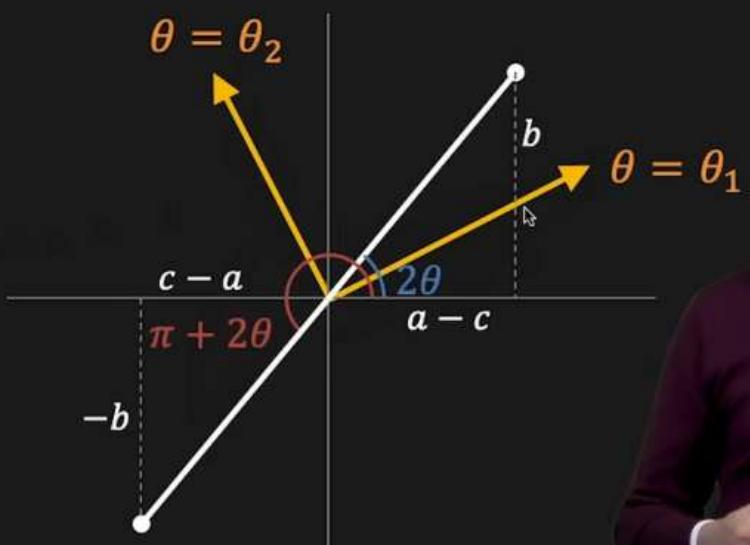
$$\tan 2\theta = \frac{b}{a - c}$$

We know that:  $\tan 2\theta = \tan(2\theta + \pi) = \frac{-b}{c - a}$

$\theta$  has two solutions.

1.  $\theta = \theta_1$
2.  $\theta = \theta_2 = \theta_1 + \frac{\pi}{2}$

One gives Minimum of  $E$   
and the other Maximum of  $E$



# Which One To Use?

---

Using second derivative test:

If  $\frac{d^2E}{d\theta^2} = (a - c) \cos 2\theta + b \sin 2\theta$

$> 0$	then <b>Minimum</b>
$< 0$	then <b>Maximum</b>

Substituting  $\cos 2\theta_1$ ,  $\sin 2\theta_1$ ,  $\cos 2\theta_2$  and  $\sin 2\theta_2$ :

$$\frac{d^2E}{d\theta^2}(\theta_1) > 0 \quad \text{and} \quad \frac{d^2E}{d\theta^2}(\theta_2) < 0$$

Therefore,

Orientation:

$$\theta = \theta_1 = \frac{\operatorname{atan2}(b, a - c)}{2}$$



# Which One To Use?

Using second derivative test:

If  $\frac{d^2E}{d\theta^2} = (a - c) \cos 2\theta + b \sin 2\theta$

$> 0$	then <b>Minimum</b>
$< 0$	then <b>Maximum</b>

Substituting  $\cos 2\theta_1, \sin 2\theta_1, \cos 2\theta_2$  and  $\sin 2\theta_2$ :

$$\frac{d^2E}{d\theta^2}(\theta_1) > 0 \quad \text{and} \quad \frac{d^2E}{d\theta^2}(\theta_2) < 0$$



Therefore,

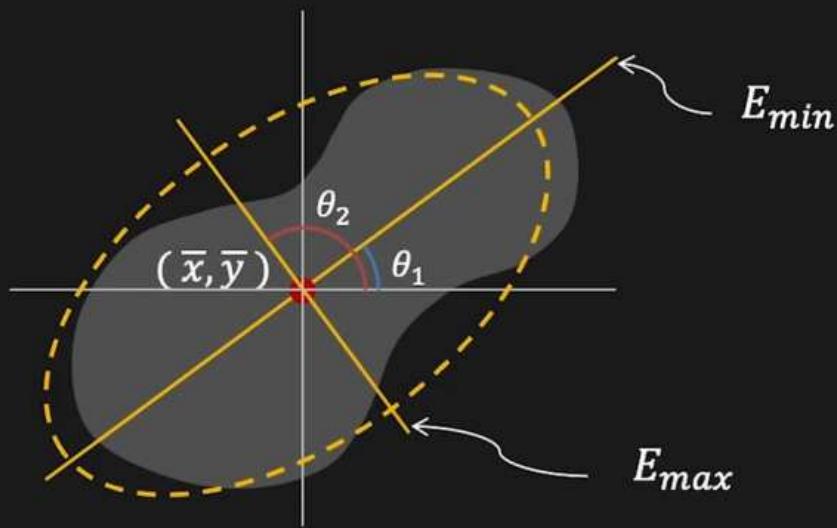
Orientation:

$$\theta = \theta_1 = \frac{\operatorname{atan2}(b, a - c)}{2}$$



# Roundedness

---

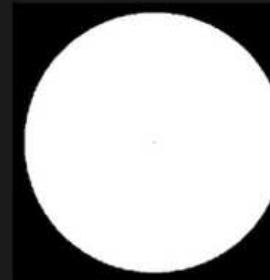
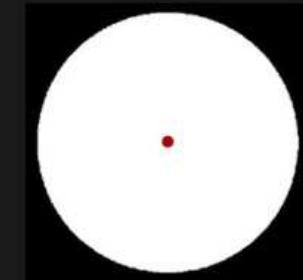


$$\text{Roundedness} = \frac{E_{min}}{E_{max}}$$

where:  $E_{min} = E(\theta_1)$  and  $E_{max} = E(\theta_2)$



# Examples

Gray Image	Binary Image	Orientation	Roundedness
			0.19
			0.49
			1.0

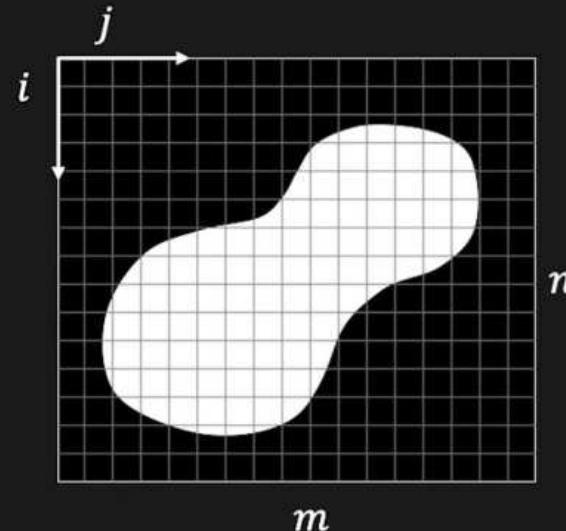


# Discrete Binary Images

$b_{ij}$ : Value at cell (pixel) in row  $i$  and column  $j$ .

Assume pixel area = 1.

**Area:** 
$$A = \sum_{i=1}^n \sum_{j=1}^m b_{ij}$$



**Position:** Center of Area (First Moment)

$$\bar{x} = \frac{1}{A} \sum_{i=1}^n \sum_{j=1}^m i b_{ij}$$

$$\bar{y} = \frac{1}{A} \sum_{i=1}^n \sum_{j=1}^m j b_{ij}$$



# Discrete Binary Images

---

Second Moments:

$$a' = \sum_{i=1}^n \sum_{j=1}^m i^2 b_{ij} \quad b' = 2 \sum_{i=1}^n \sum_{j=1}^m ij b_{ij} \quad c' = \sum_{i=1}^n \sum_{j=1}^m j^2 b_{ij}$$

Note:  $a'$ ,  $b'$ ,  $c'$  are second moments w.r.t origin.

$a$ ,  $b$ ,  $c$  (w.r.t. center) can be found from  $a'$ ,  $b'$ ,  $c'$ ,  $\bar{x}$ ,  $\bar{y}$ ,  $A$

Hint: Expand  $a = \sum_{i=1}^n \sum_{j=1}^m (i - \bar{x})^2 b_{ij}$  and represent in terms of  $a'$ ,  $\bar{x}$ ,  $A$ .



# Discrete Binary Images

## Second Moments:

$$a' = \sum_{i=1}^n \sum_{j=1}^m i^2 b_{ij} \quad b' = 2 \sum_{i=1}^n \sum_{j=1}^m ij b_{ij} \quad c' = \sum_{i=1}^n \sum_{j=1}^m j^2 b_{ij}$$

Note:  $a'$ ,  $b'$ ,  $c'$  are second moments w.r.t origin.

$a$ ,  $b$ ,  $c$  (w.r.t. center) can be found from  $a'$ ,  $b'$ ,  $c'$ ,  $\bar{x}$ ,  $\bar{y}$ ,  $A$

Hint: Expand  $a = \sum_{i=1}^n \sum_{j=1}^m (i - \bar{x})^2 b_{ij}$  and represent in terms of  $a'$ ,  $\bar{x}$ ,  $A$ .



# Multiple Objects

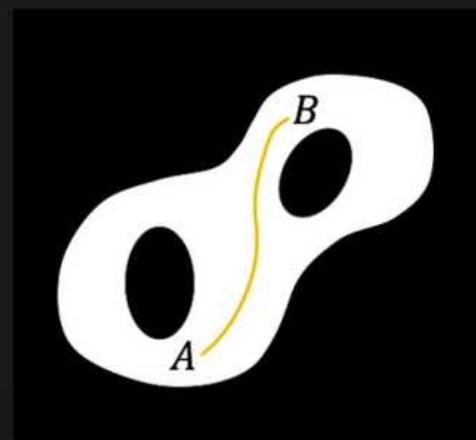


Need to **Segment** image into separate **Components**



# Connected Component

Maximal Set of Connected Points



A and B are connected if path exists between A and B along which  $b(x, y)$  is constant.



# Connected Component Labeling

---

## Region Growing Algorithm

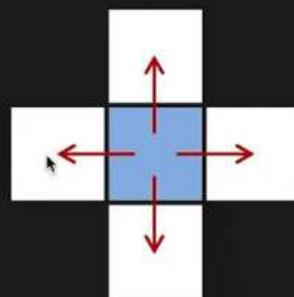
- (a) Find **Unlabeled "Seed"** point with  $b = 1$ .  
If not found, Terminate.
- (b) Assign **New Label** to seed point
- (c) Assign **Same Label** to its Neighbors with  $b = 1$
- (d) Assign **Same Label** to Neighbors of Neighbors  
with  $b = 1$ . Repeat until no more Unlabeled  
Neighbors with  $b=1$ .
- (e) Go to (a)



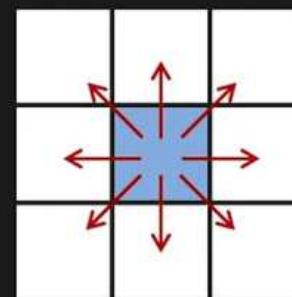
# What do we mean by Neighbors?

---

Connectedness



4-Connectedness  
4-C



8-Connectedness  
8-C

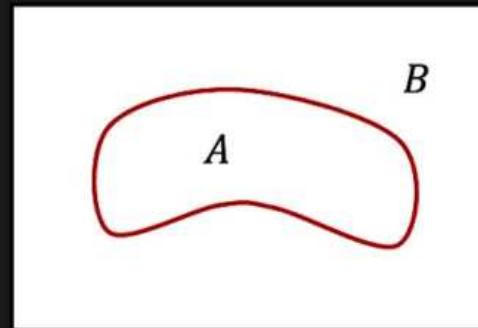
Neither is Perfect!



# Connectedness

Jordan's Curve Theorem

Closed curve  
→ 2 Connected Regions



Consider

0	1	0
1	0	1
0	1	0



B1	O1	B1
O4	B2	O2
B1	O3	B1

4-C  
Hole without a  
closed loop!

B	O	B
O	B	O
B	O	B

8-C  
Connected backgrounds  
with a closed loop!



# Hexagonal Tessellation

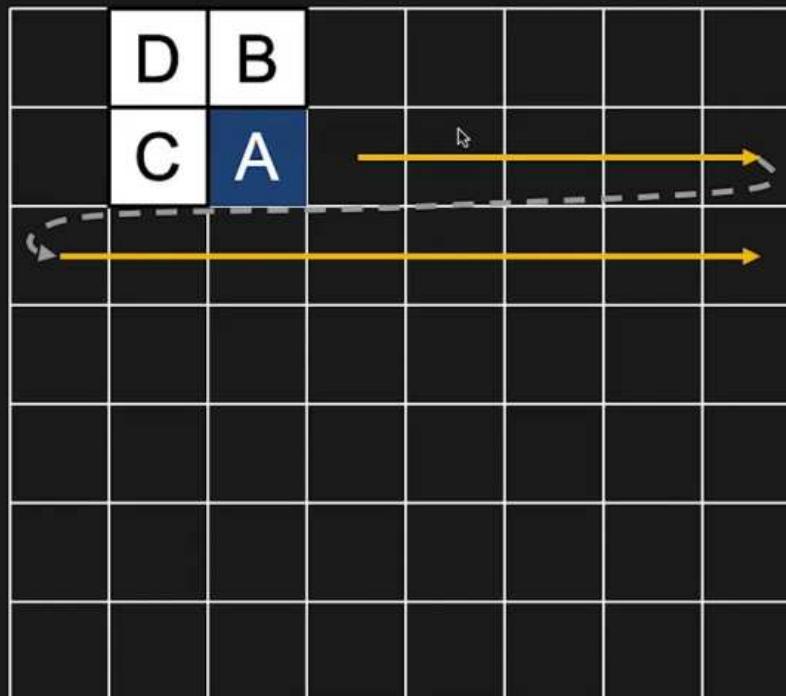
---



Above asymmetry makes a Square Grid behave  
like a Hexagonal Grid



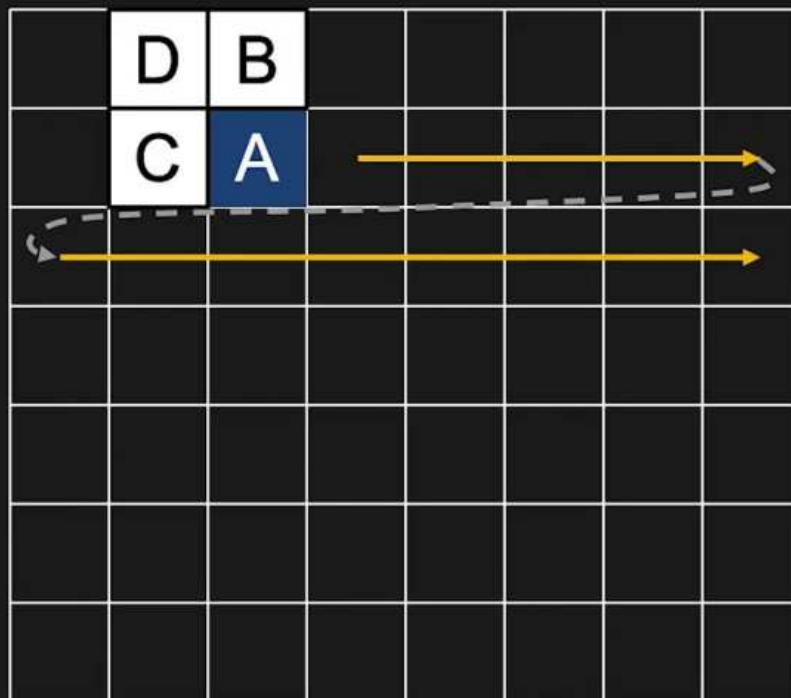
# Sequential Labeling Algorithm



We want to label A.  
B, C, D are already labeled.



# Sequential Labeling Algorithm



Raster  
Scanning

We want to label A.  
B, C, D are already labeled.



# Sequential Labeling Algorithm

X	X
X	0

→  $\text{label}(A) = \text{"background"}$

0	0
0	1

→  $\text{label}(A) = \text{new label}$

D	X
X	1

→  $\text{label}(A) = \text{label}(D)$

0	0
C	1

→  $\text{label}(A) = \text{label}(C)$

0	B
0	1

→  $\text{label}(A) = \text{label}(B)$

0	B
C	1

→ If  
     $\text{label}(B) = \text{label}(C)$   
then,  
     $\text{label}(A) = \text{label}(B)$

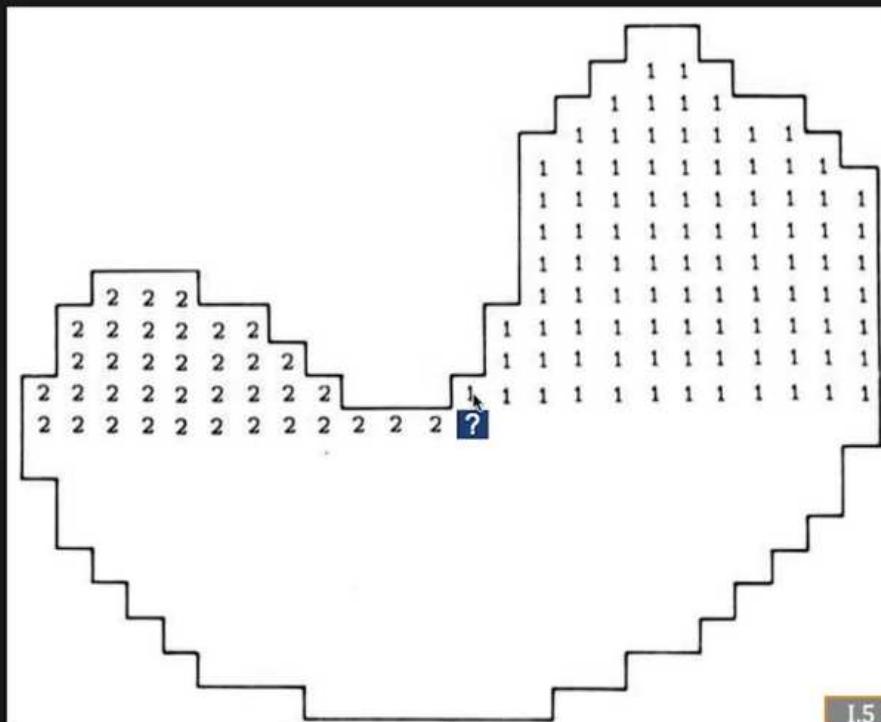


X: Value does not matter (Can be 0 or 1)

# Sequential Labeling Algorithm

0	B
C	1

→ What if  $\text{label}(B) \neq \text{label}(C)$ ?



# Sequential Labeling Algorithm

---

0	B
C	1

→ What if  $\text{label}(B)$  not equal to  $\text{label}(C)$ ?

Solution: Create **Equivalence Table**

- Note down that  $\text{label}(B) \equiv \text{label}(C)$
- Assign  $\text{label}(A) = \text{label}(B)$

2 ≡ 1
7 ≡ 3, 6, 4
8 ≡ 5
⋮



# Euler Number (E)

---

No. of Bodies ( $B$ ) – No. of Holes ( $H$ )



Letter B:  $E_b = -1$

Letter i:  $E_i = 2$

Letter n:  $E_n = 1$

Image :  $E = 4$

$$E_{image} = \sum E_{non-overlapping\ regions}$$



# Euler Differential ( $E^*$ )

---

Change in the Euler number of the image.



$$E = 0$$



$$E = 1$$

Euler Differential:  $E^* = 1$

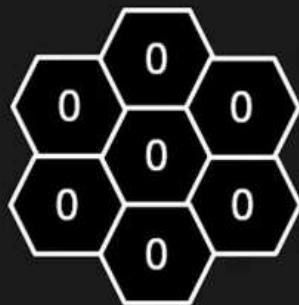


# Neighborhood Sets Based on $E^*$

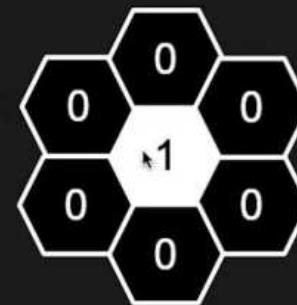
---

Each pixel has  $2^6 = 64$  possible neighborhoods

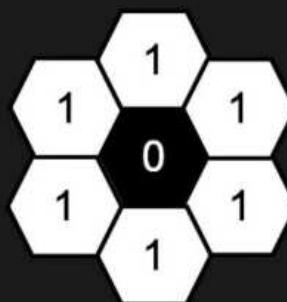
Neighborhood patterns are classified based on the Euler Differential they generate, assuming the center pixel goes from 0 to 1.



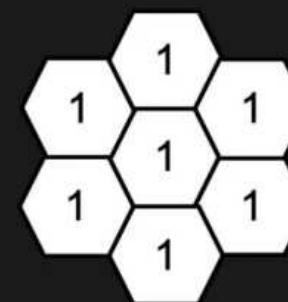
Neighborhood  $\in N_{+1}$   
 $0 \rightarrow 1, E^* = 1$



# Neighborhood Sets Based on $E^*$



Neighborhood  $\in N_{+1}$   
 $0 \rightarrow 1, E^* = 1$



Neighborhood  $\in N_0$   
 $0 \rightarrow 1, E^* = 0$

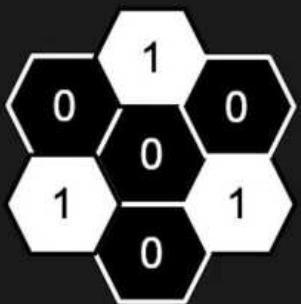
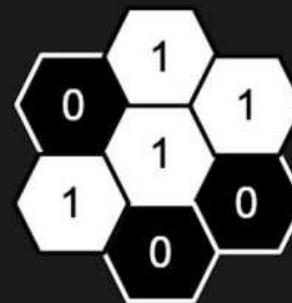


# Neighborhood Sets Based on $E^*$

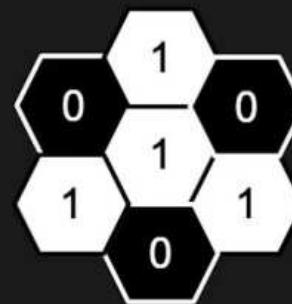
---



Neighborhood  $\in N_{-1}$   
 $0 \rightarrow 1, E^* = -1$



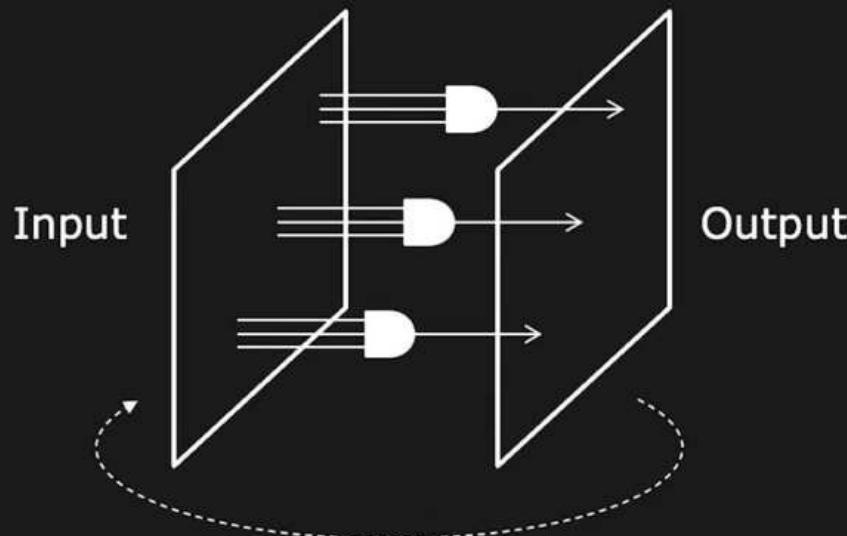
Neighborhood  $\in N_{-2}$   
 $0 \rightarrow 1, E^* = -2$



# Iterative Neighborhood Operations

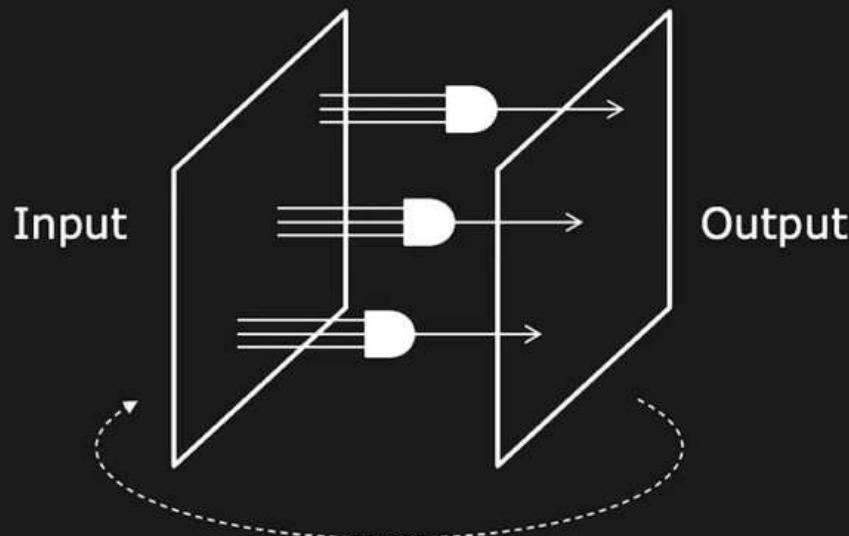
---

Incrementally apply neighborhood operations on images



# Iterative Neighborhood Operations

Incrementally apply neighborhood operations on images



**Conservative Operations** do **not** change  
the Euler number of the image.



# Notation for Iterative Modification

Specify Neighborhood Set,  $S$ . Ex:  $S$  can be  $N_{+1}$ ,  $N_0$ ,  $N_{-1}$  or  $N_{-2}$  or a combination of these.

Consider pixel  $(i, j)$ . Let:

- $a_{ij} = 1$  if Neighborhood of  $(i, j) \in S$  else 0
- $b_{ij}$  = current value of pixel  $(i, j)$
- $c_{ij}$  = new value of pixel  $(i, j)$

$a_{ij}$	$b_{ij}$		$c_{ij}$
0	0		?
0	1		?
1	0		?
1	1		?



# Iterative Modification Algorithms

Specify Neighborhood Set  $S$  and apply one of the 16 algorithms to each pixel.

$\longleftrightarrow$  16 algorithms  $\longrightarrow$

$a_{ij}$	$b_{ij}$	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
0	0	0	0	0	0	0	0	0	1	1	1	1	1	1	1	1	
0	1	0	0	0	0	1	1	1	1	0	0	0	0	1	1	1	
1	0	0	0	1	1	0	0	1	1	0	0	1	1	0	0	1	
1	1	0	1	0	1	0	1	0	1	0	1	0	1	0	1	0	



# Iterative Modification Algorithms

Specify Neighborhood Set  $S$  and apply one of the 16 algorithms to each pixel.

← 16 algorithms →

$a_{ij}$	$b_{ij}$	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
0	0	0	0	0	0	0	0	0	0	1	1	1	1	1	1	1	
0	1	0	0	0	0	1	1	1	1	0	0	0	0	1	1	1	
1	0	0	0	1	1	0	0	1	1	0	0	1	1	0	0	1	
1	1	0	1	0	1	0	1	0	1	0	1	0	1	0	1	0	

Growing Objects:  $S \in N_0$  and Algorithm 7

Thinning Objects:  $S \in N_0$  and Algorithm 4



# Finding Skeletons

---

Thinning without changing the Euler number

$S \in N_0$  and *Algorithm 4*



# Finding Skeletons

---

Thinning without changing the Euler number

$S \in N_0$  and *Algorithm 4*



# Image Processing I

Transform image to new one that is clearer or easier to analyze.

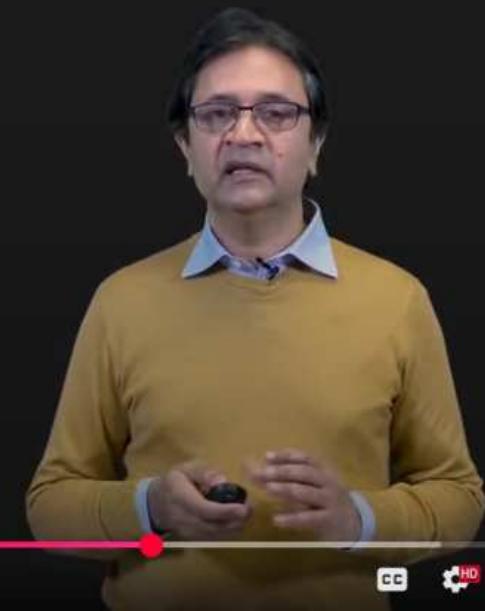


# Image Processing I

Transform image to new one that is clearer or easier to analyze.

## Topics:

- (1) Pixel Processing
- (2) LSIS and Convolution
- (3) Linear Image Filters



# Image Processing I

---

Transform image to new one that is clearer or easier to analyze.

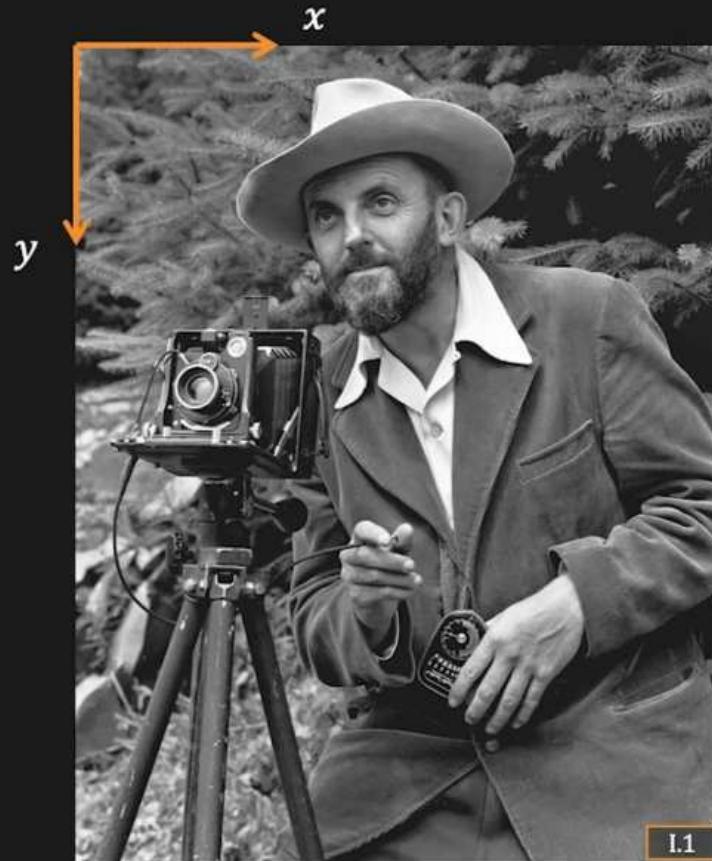
## Topics:

- (1) Pixel Processing
- (2) LSIS and Convolution
- (3) Linear Image Filters
- (4) Non-Linear Image Filters



# Image as a Function

---



$f(x, y)$  is the image intensity at position  $(x, y)$



# Pixel (Point) Processing

Transformation  $T$  of intensity  $f$  at each pixel to intensity  $g$ :

$$g(x, y) = T(f(x, y))$$



# Point Processing



Original ( $f$ )



Lighten ( $f + 128$ )



Darken ( $f - 128$ )

Invert ( $255 - f$ )



# Pixel Processing



Original ( $f$ )



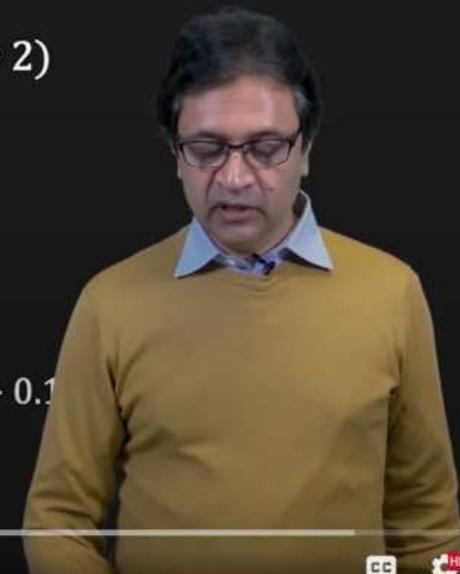
Low Contrast ( $f/2$ )



High Contrast ( $f * 2$ )



Gray ( $0.3f_R + 0.6f_G + 0.1f_B$ )

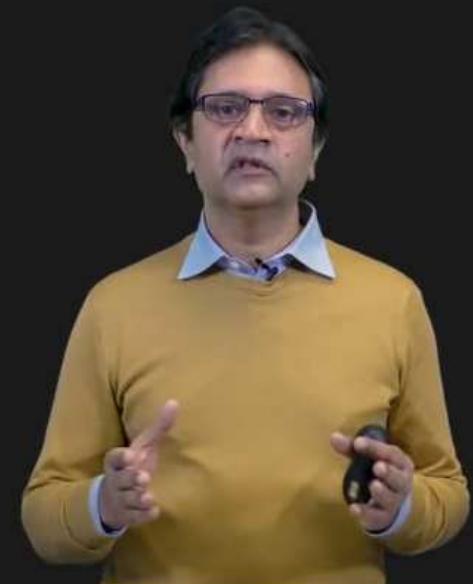


# Linear Shift Invariant System (LSIS)

---



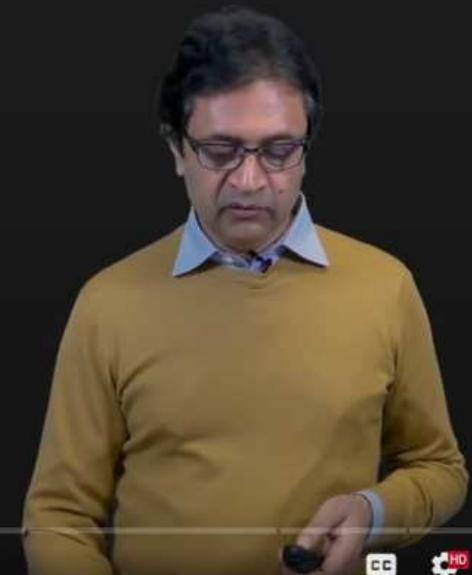
Study of Linear Shift Invariant Systems (**LSIS**)  
leads to useful image processing algorithms.



# LSIS: Linearity

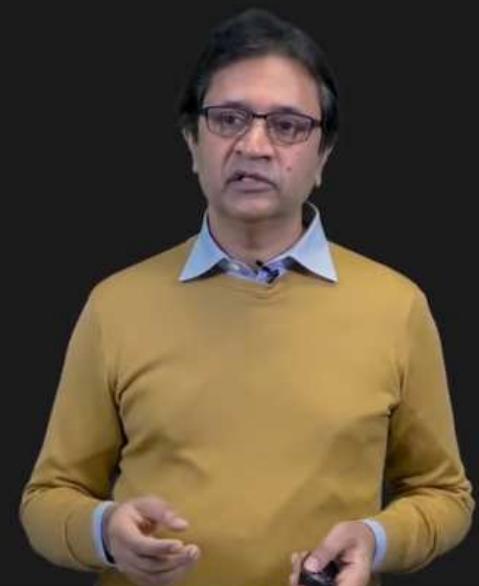
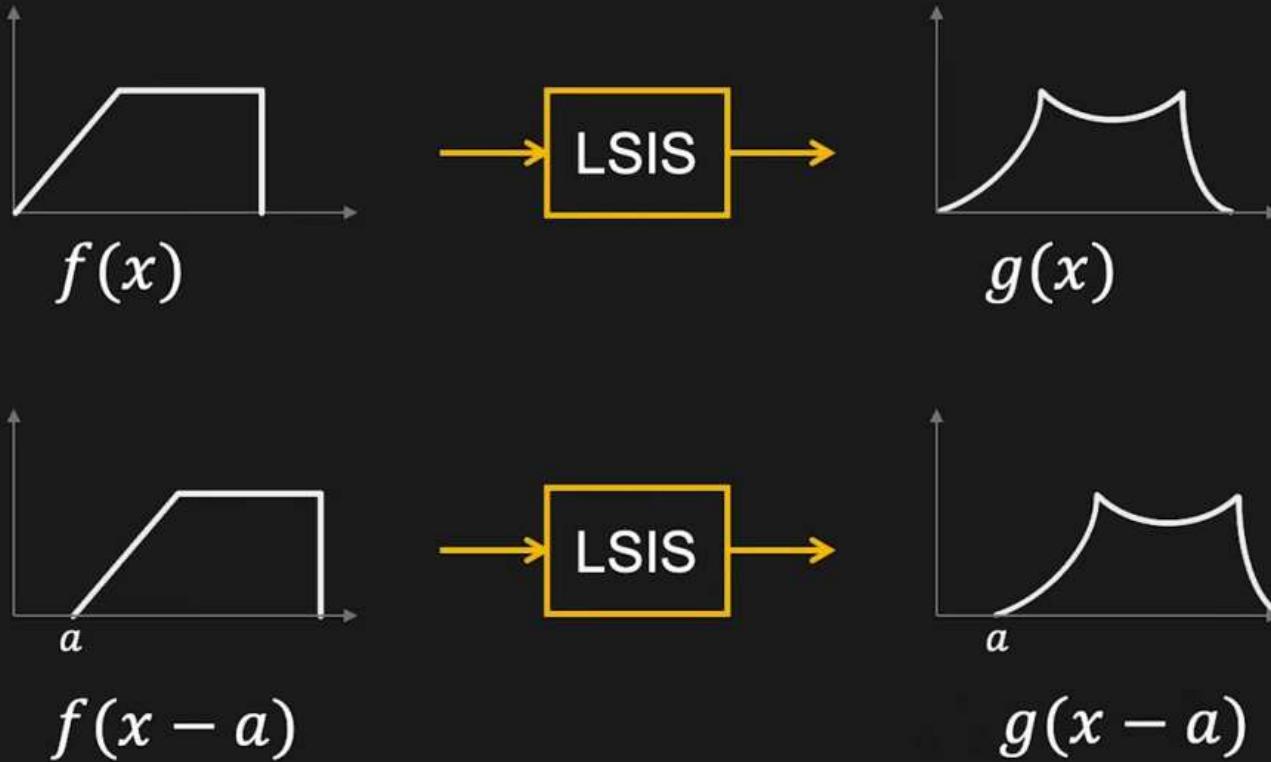
$$f_1 \rightarrow \boxed{\text{LSIS}} \rightarrow g_1 \quad f_2 \rightarrow \boxed{\text{LSIS}} \rightarrow g_2$$

$$\alpha f_1 + \beta f_2 \rightarrow \boxed{\text{LSIS}} \rightarrow \alpha g_1 + \beta g_2$$



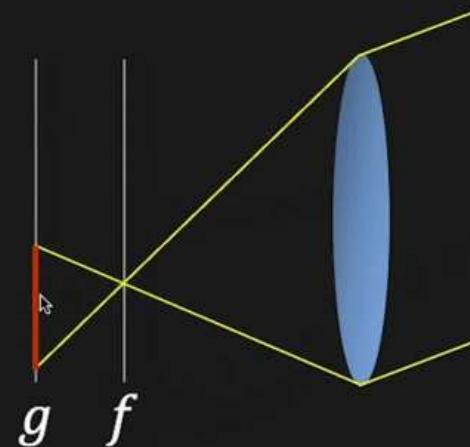
# LSIS: Shift Invariance

---

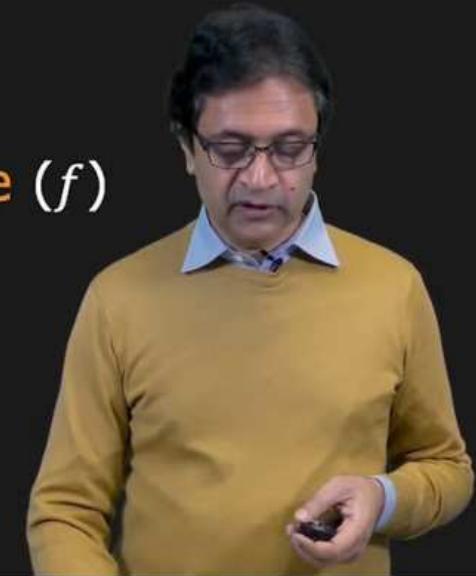


# Ideal Lens is an LSIS

---

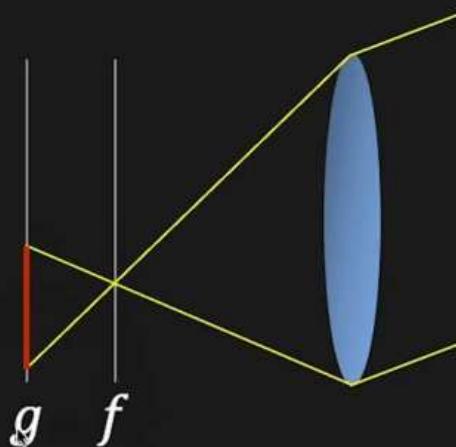


Defocused Image ( $g$ ): Processed version of Focused Image ( $f$ )



# Ideal Lens is an LSIS

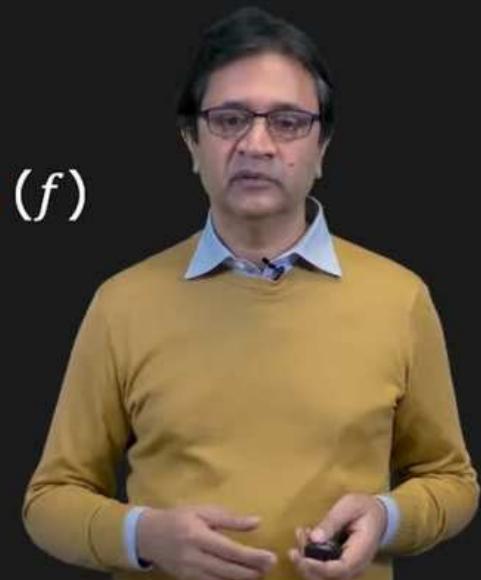
---



Defocused Image ( $g$ ): Processed version of Focused Image ( $f$ )

Linearity: Brightness variation

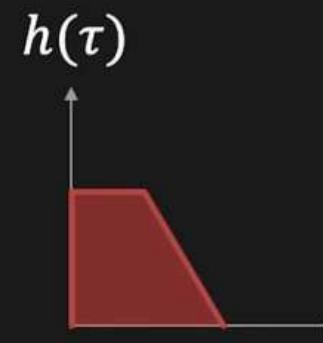
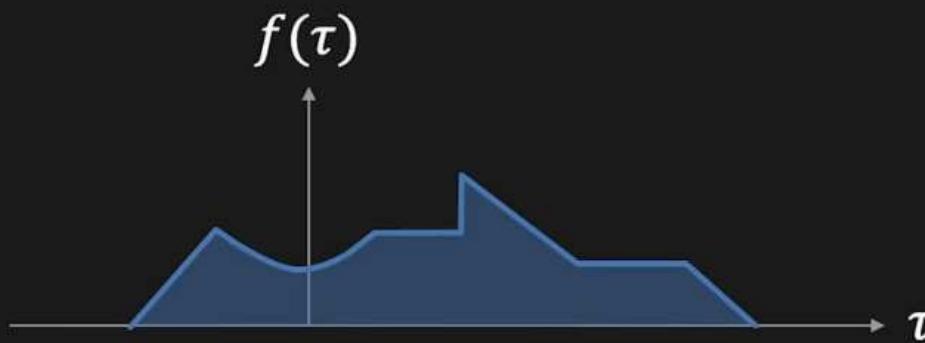
Shift invariance: Scene movement



# Convolution

Convolution of two functions  $f(x)$  and  $h(x)$

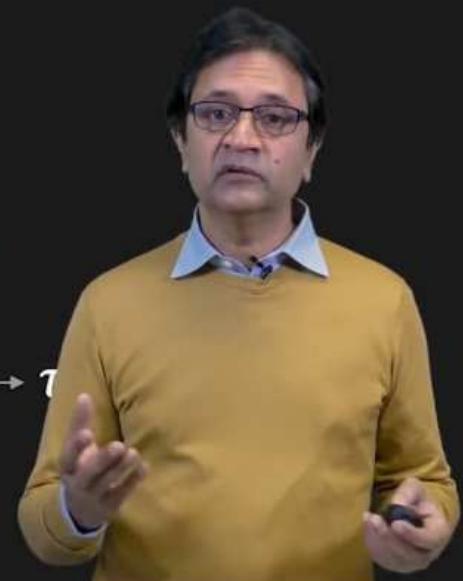
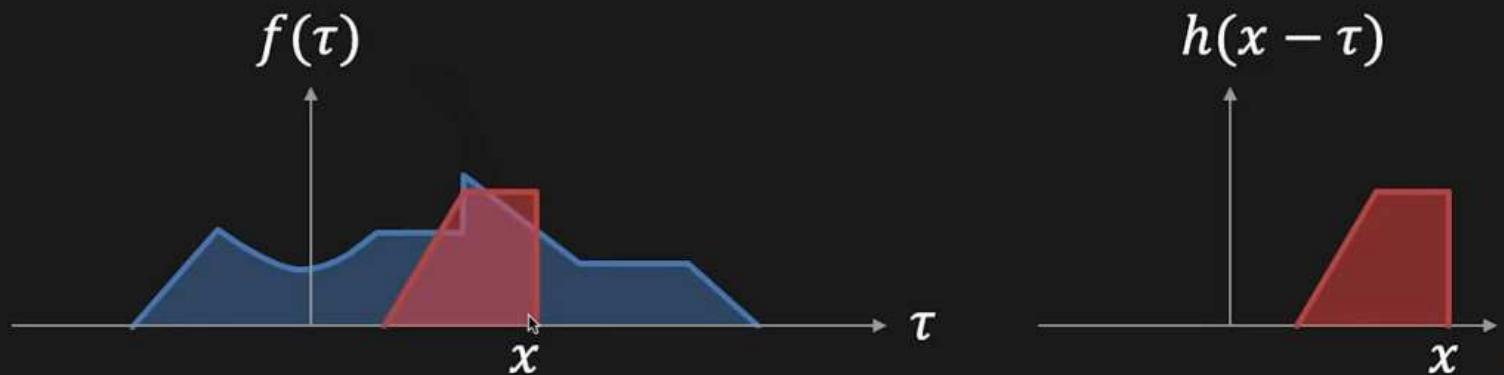
$$g(x) = f(x) * h(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$$



# Convolution

Convolution of two functions  $f(x)$  and  $h(x)$

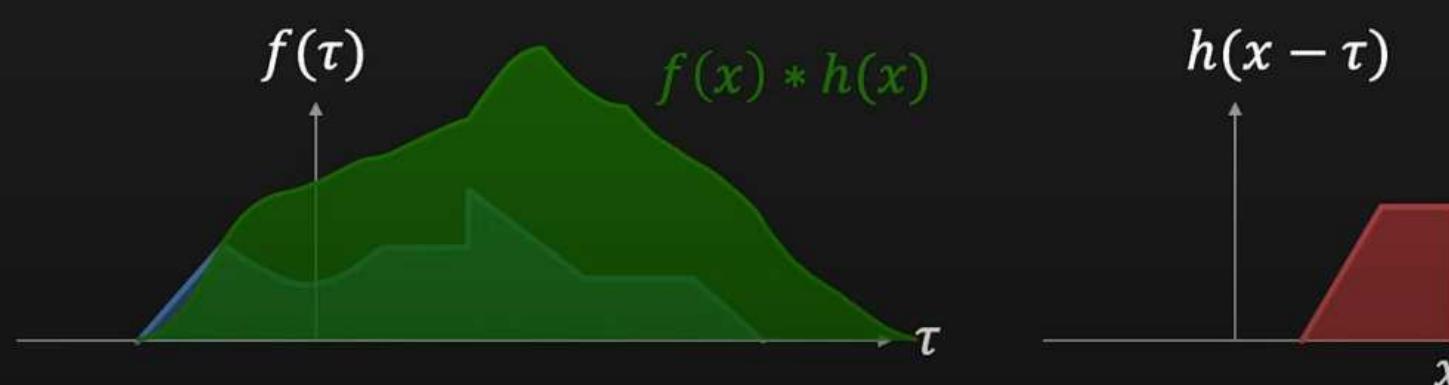
$$g(x) = f(x) * h(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$$



# Convolution

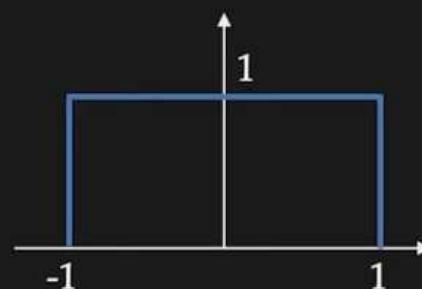
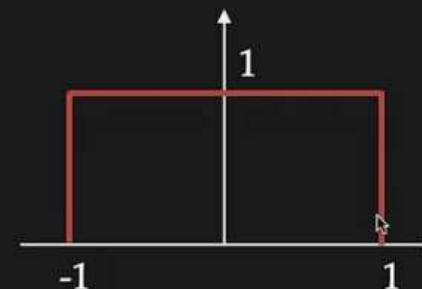
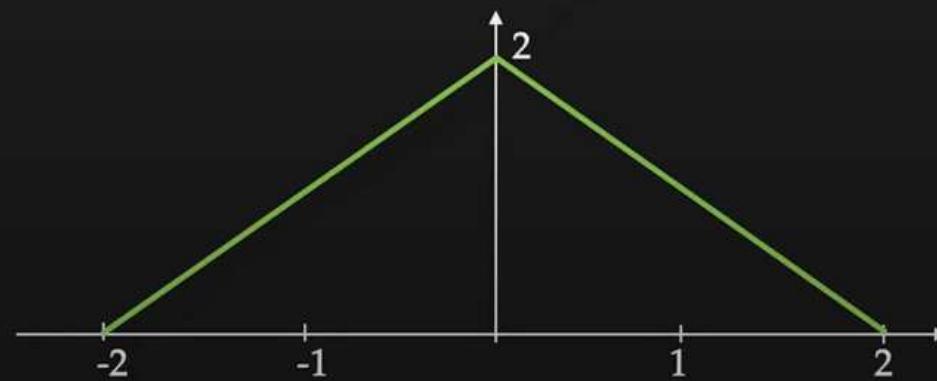
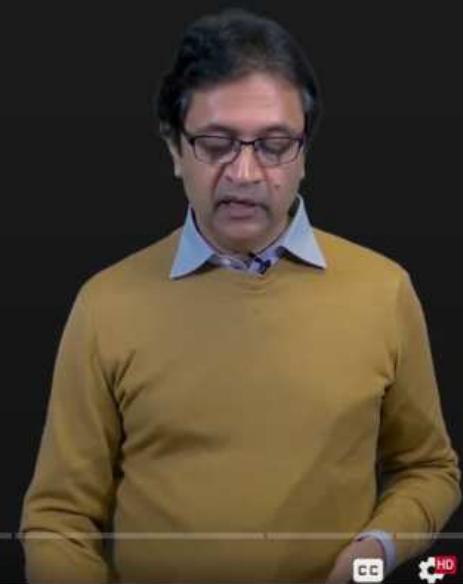
Convolution of two functions  $f(x)$  and  $h(x)$

$$g(x) = f(x) * h(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$$



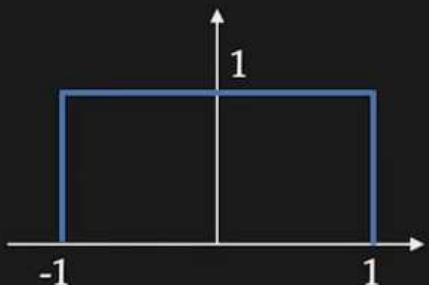
LSIS implies Convolution and Convolution implies LS

# Convolution: Example

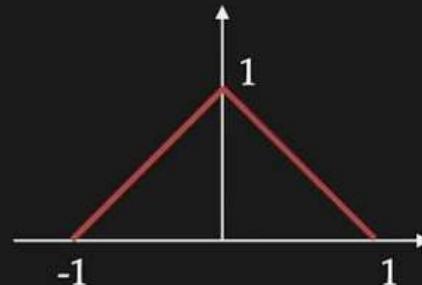

$$f(x)$$

$$h(x)$$

$$f(x) * h(x)$$


# Convolution: Example

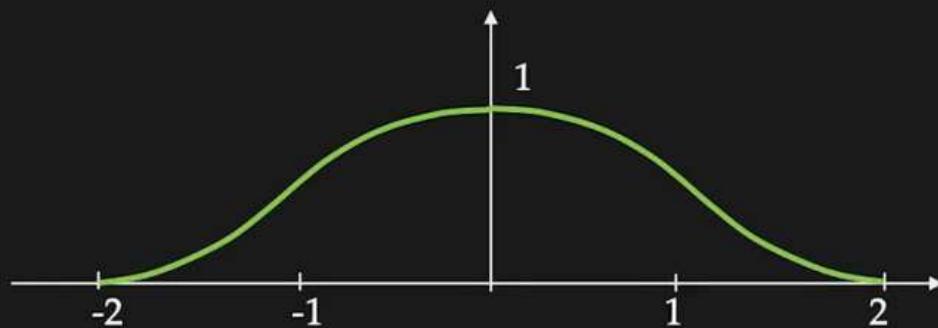
---



$f(x)$



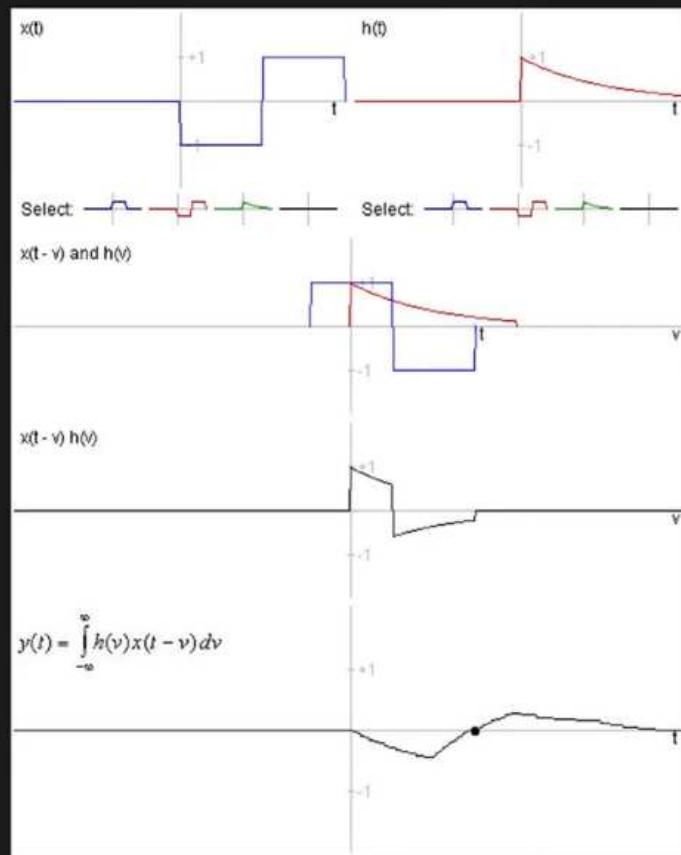
$h(x)$



$f(x) * h(x)$



# Convolution: Online Demo



<http://www.jhu.edu/signals/convolve/>

# Convolution is LSIS

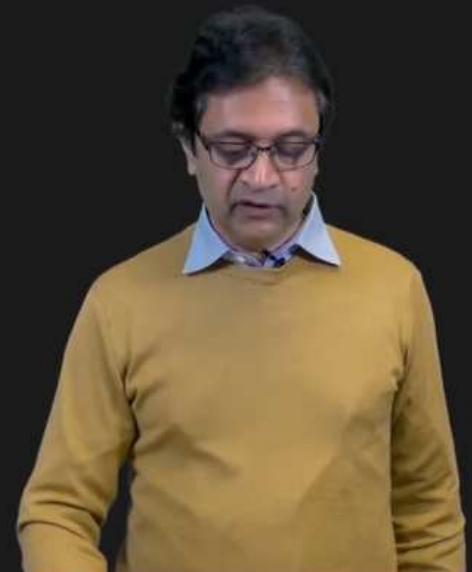
---

Linearity:

Let:  $g_1(x) = \int_{-\infty}^{\infty} f_1(\tau)h(x - \tau) d\tau$  and  $g_2(x) = \int_{-\infty}^{\infty} f_2(\tau)h(x - \tau) d\tau$

Then:

$$\begin{aligned} & \int_{-\infty}^{\infty} (\alpha f_1(\tau) + \beta f_2(\tau))h(x - \tau) d\tau \\ &= \alpha \int_{-\infty}^{\infty} f_1(\tau)h(x - \tau) d\tau + \beta \int_{-\infty}^{\infty} f_2(\tau)h(x - \tau) d\tau \\ &= \alpha g_1(x) + \beta g_2(x) \end{aligned}$$



# Convolution is LSIS

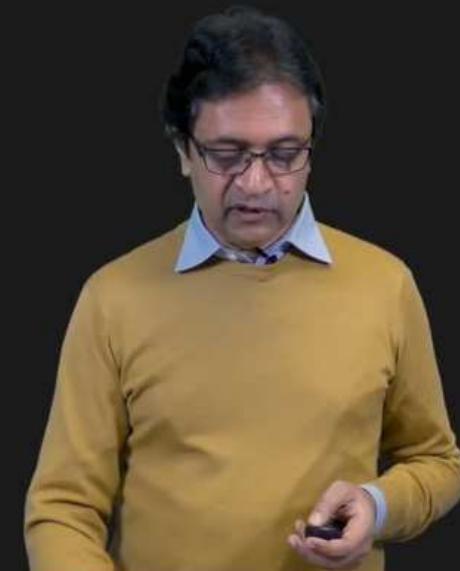
---

Shift Invariance:

Let: 
$$g(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$$

Then:

$$\begin{aligned} & \int_{-\infty}^{\infty} f(\tau - a)h(x - \tau) d\tau \\ &= \int_{-\infty}^{\infty} f(\mu)h(x - a - \mu) d\mu && \text{(Substituting } \mu = \tau - a\text{)} \\ &= g(x - a) \end{aligned}$$



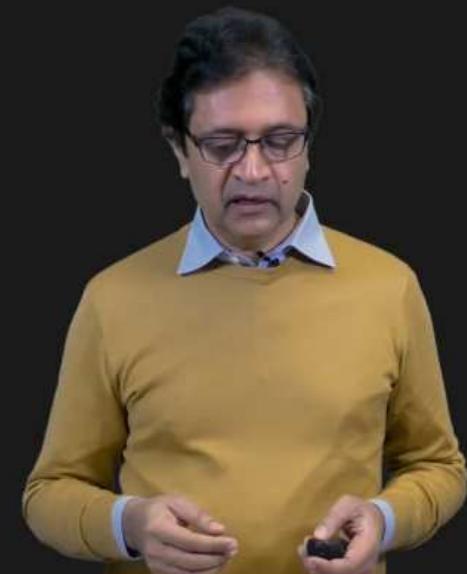
## Can we find $h$ ?

---

$$f \rightarrow \boxed{h} \rightarrow g \quad g(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$$

What input  $f$  will produce output  $g = h$  ?

$$h(x) = \int_{-\infty}^{\infty} ?(\tau)h(x - \tau) d\tau$$

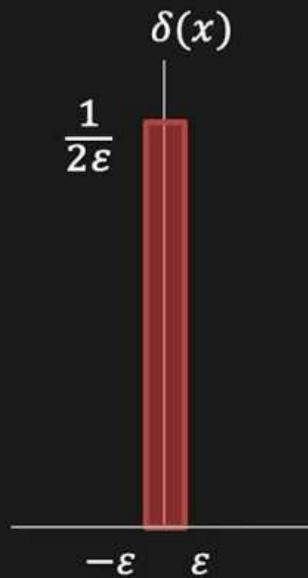


# Unit Impulse Function

$$\delta(x) = \begin{cases} 1/2\varepsilon, & |x| \leq \varepsilon \\ 0, & |x| > \varepsilon \end{cases}$$

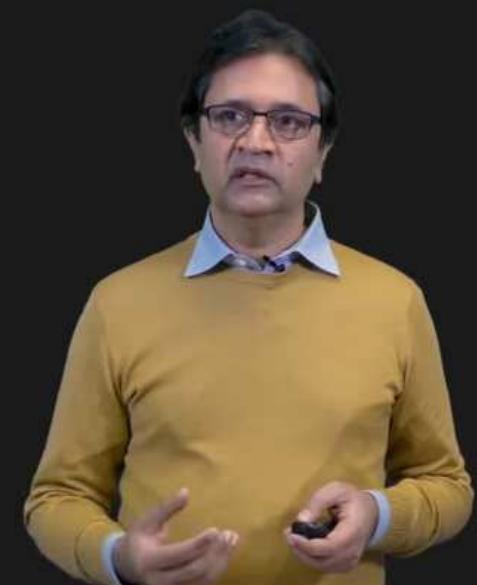
$\varepsilon \rightarrow 0$

$$\int_{-\infty}^{\infty} \delta(\tau) d\tau = \frac{1}{2\varepsilon} \cdot 2\varepsilon = 1$$

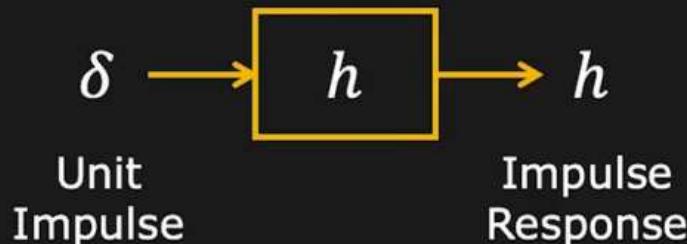


$$\int_{-\infty}^{\infty} \delta(\tau) b(x - \tau) d\tau = b(x)$$

Sifting Property



# Impulse Response

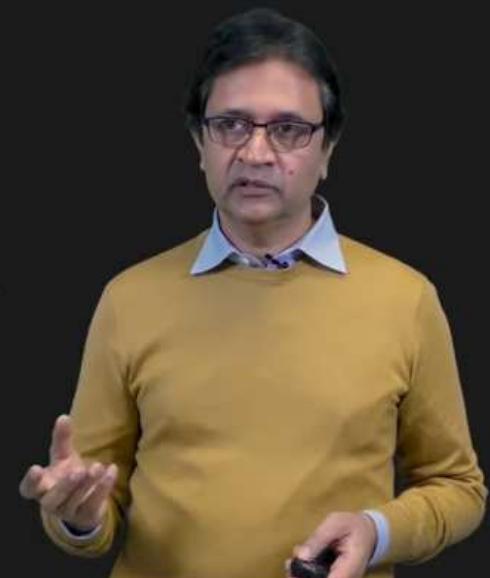


$$g(x) = f(x) * h(x)$$

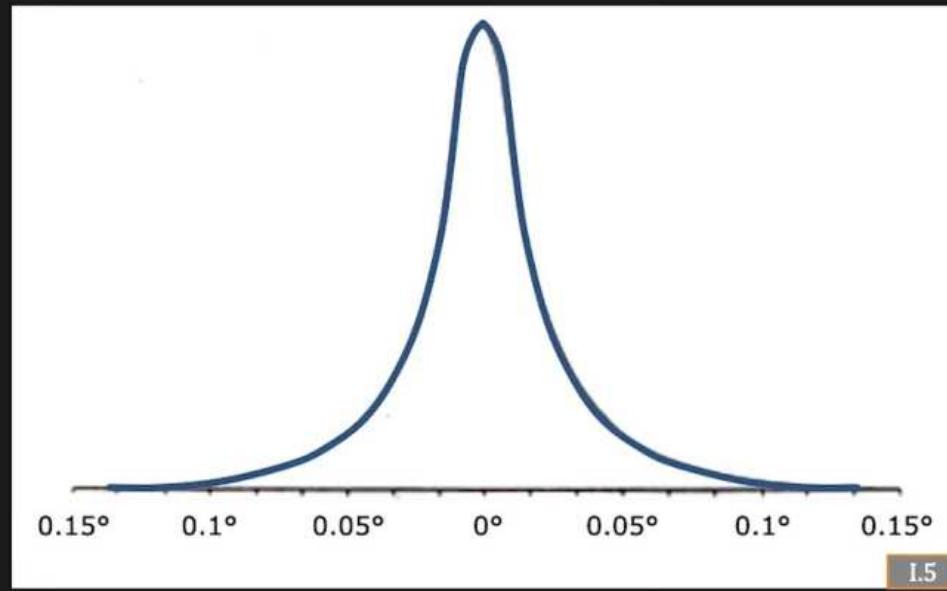
$$h(x) = \delta(x) * h(x)$$

$$g(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$$

$$h(x) = \int_{-\infty}^{\infty} \delta(\tau)h(x - \tau) d\tau$$



# Impulse Response of Human Eye



Human Eye PSF



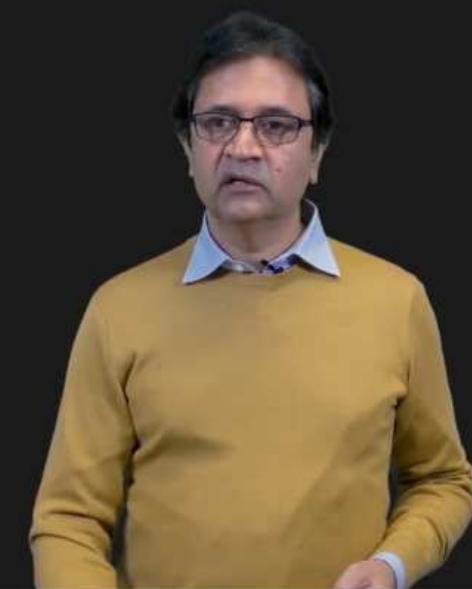
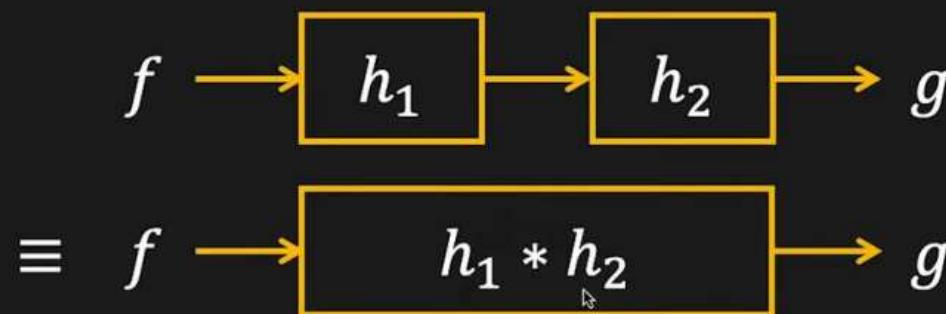
# Properties of Convolution

---

Commutative       $a * b = b * a$

Associative       $(a * b) * c = a * (b * c)$

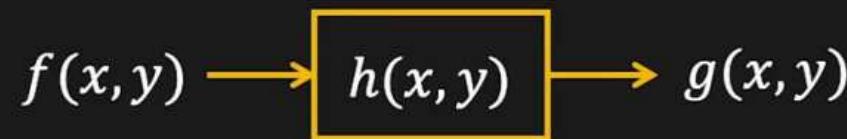
Cascaded System



# 2D Convolution

---

LSIS:



Convolution:

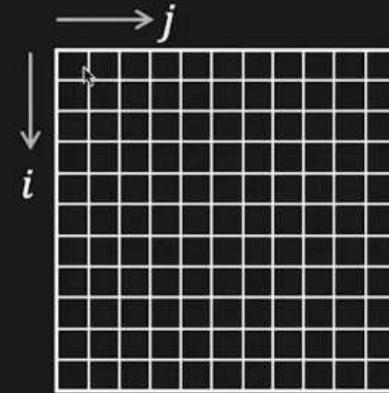
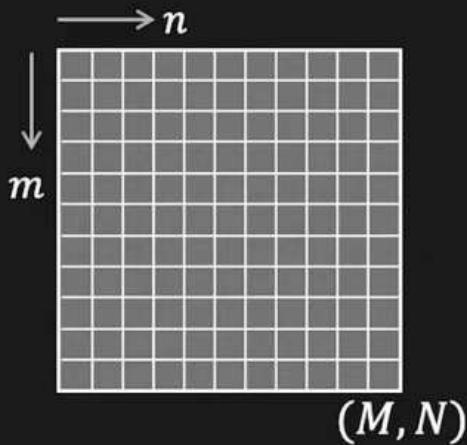
$$g(x, y) = \iint_{-\infty}^{\infty} f(\tau, \mu) h(x - \tau, y - \mu) d\tau d\mu$$



# Convolution with Discrete Images



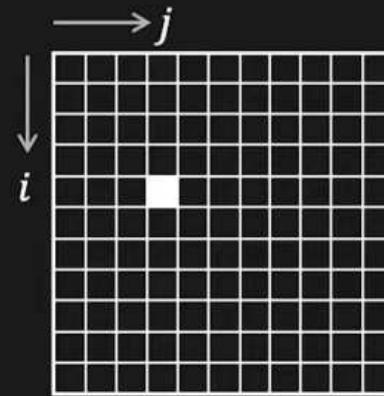
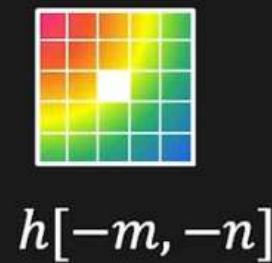
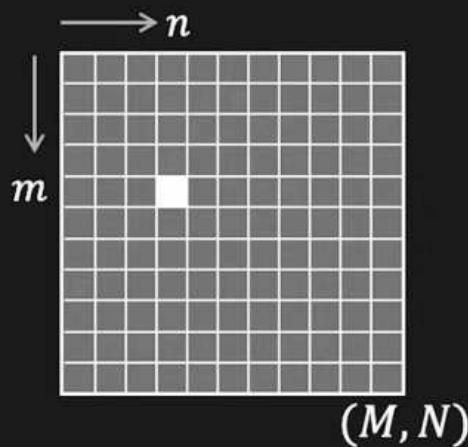
$$g[i,j] = \sum_{m=1}^M \sum_{n=1}^N f[m,n]h[i-m, j-n]$$



# Convolution with Discrete Images



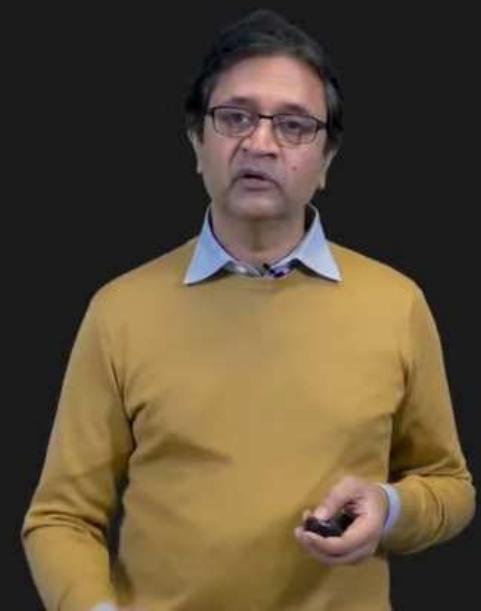
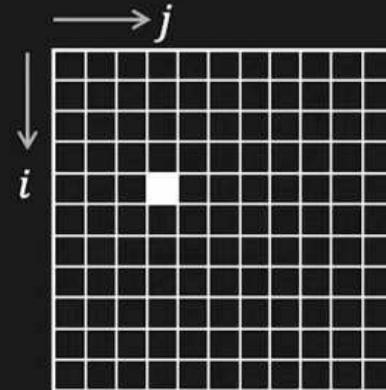
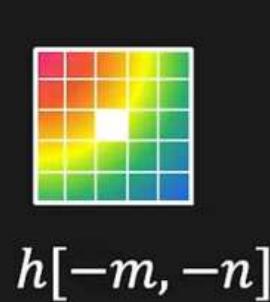
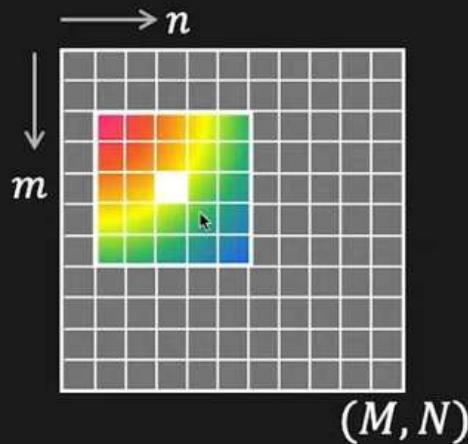
$$g[i,j] = \sum_{m=1}^M \sum_{n=1}^N f[m,n] \underbrace{h[i-m, j-n]}_{\text{"Mask," "Kernel," "Filter"}}$$



# Convolution with Discrete Images

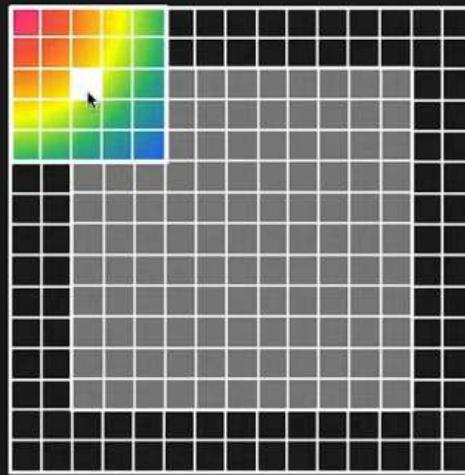


$$g[i,j] = \sum_{m=1}^M \sum_{n=1}^N f[m,n] \underbrace{h[i-m, j-n]}_{\text{"Mask," "Kernel," "Filter"}}$$



# Border Problem

---



Solution:

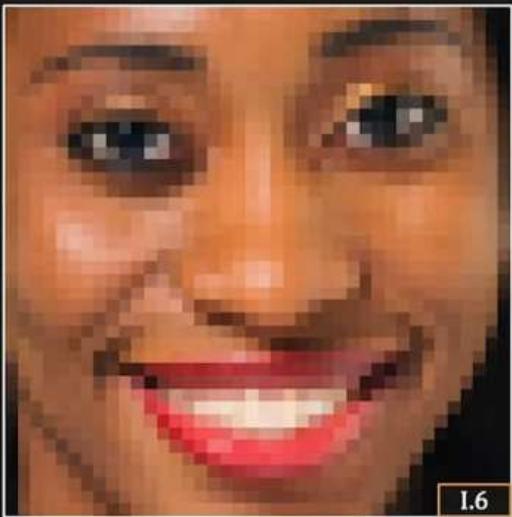
- Ignore border
- Pad with constant value
- Pad with reflection



# Example: Impulse Filter

---

Input



$f(x, y)$

$$* \begin{array}{|c|c|c|} \hline & & \\ \hline & \blacksquare & \\ \hline & & \\ \hline \end{array} =$$

Output



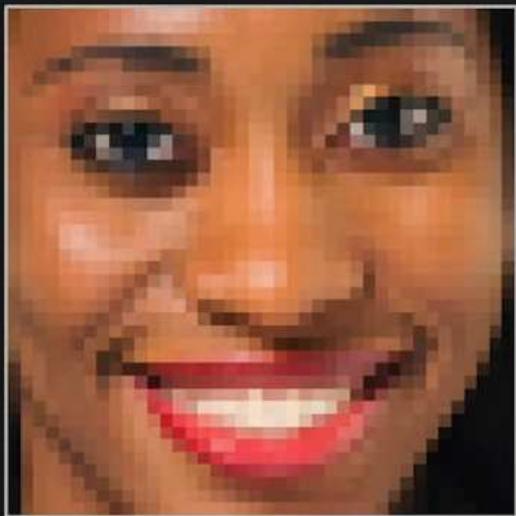
$f(x, y)$



# Example: Image Shift

---

Input



$$f(x, y)$$

$$\ast \quad \begin{array}{|c|c|c|c|} \hline & & & \\ \hline \end{array} =$$

Output



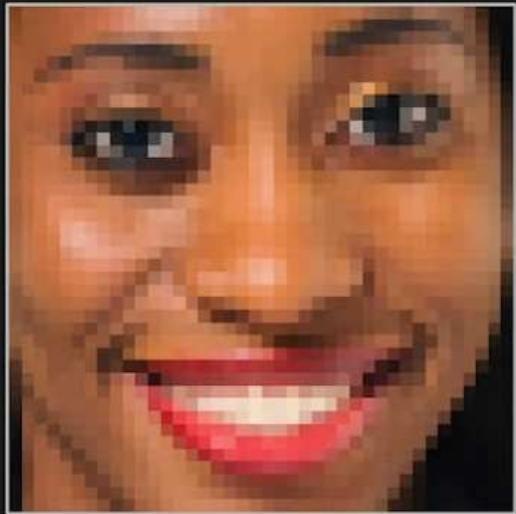
$$f(x - u, y - v)$$



# Example: Averaging

---

Input



$f(x, y)$

$$* \begin{array}{|c|c|c|c|c|} \hline & & & & \\ \hline \end{array} =$$

“Box Filter”  
 $5 \times 5$

Output



$g(x, y)$

Result Image is saturated. Why?



# Example: Averaging

---

Input



$f(x, y)$

$$* \begin{array}{|c|c|c|c|c|} \hline & & & & \\ \hline \end{array} =$$

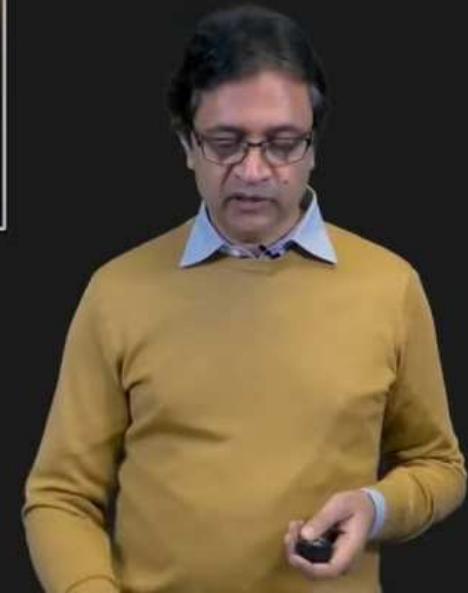
“Box Filter”  
 $5 \times 5$

Output



$g(x, y)$

Sum of all the filter (kernel) weights should be 1.



# Smoothing With Box Filter

Input



$$f(x, y)$$

$$\ast \begin{array}{|c|}\hline \text{Box Filter} \\ \hline 21 \times 21 \end{array} =$$

Output



$$g(x, y)$$

Image smoothed with a box filter does not look “natural.” Has blocky artifacts.



# Smoothing With “Fuzzy” Filter

---

Input



$f(x, y)$

$$* \begin{array}{|c|} \hline \text{ } \\ \hline \end{array} =$$

“Fuzzy Filter”  
 $21 \times 21$

Output



$g(x, y)$

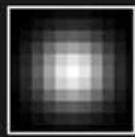
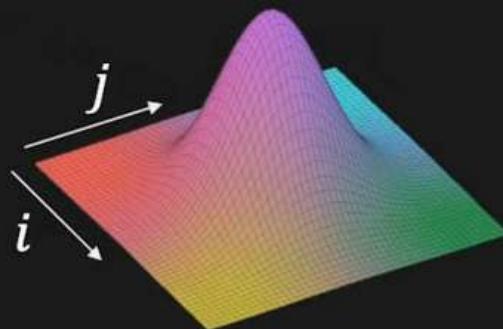


# Gaussian Kernel: A Fuzzy Filter

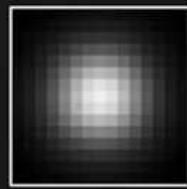
---

$$n_{\sigma}[i,j] = \frac{1}{2\pi\sigma^2} e^{-\frac{1}{2}\left(\frac{i^2+j^2}{\sigma^2}\right)}$$

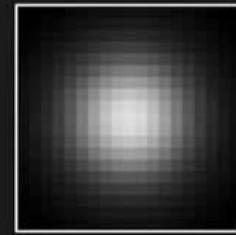
$\sigma^2$ : Variance



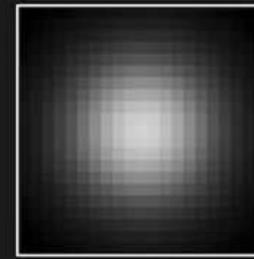
$\sigma = 2$



$\sigma = 3$



$\sigma = 4$



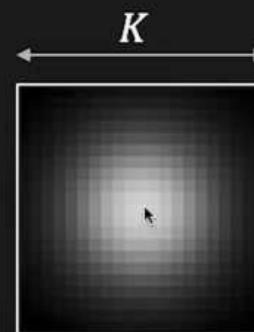
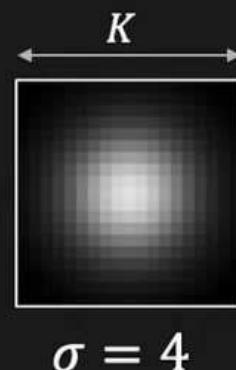
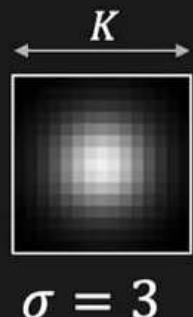
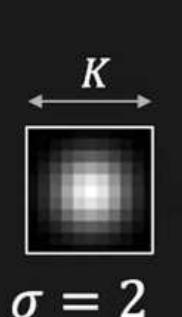
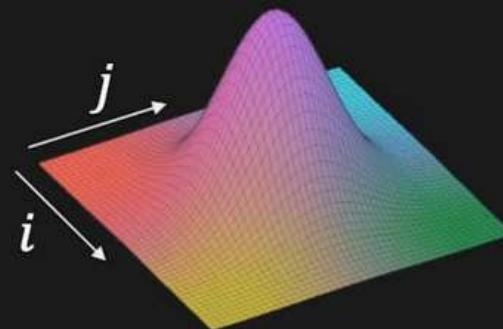
$\sigma = 5$



# Gaussian Kernel: A Fuzzy Filter

$$n_{\sigma}[i,j] = \frac{1}{2\pi\sigma^2} e^{-\frac{1}{2}\left(\frac{i^2+j^2}{\sigma^2}\right)}$$

$\sigma^2$ : Variance



Rule of thumb: Set kernel size  $K \approx 2\pi\sigma$



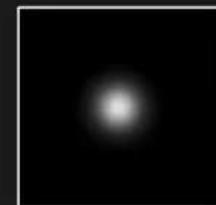
# Gaussian Smoothing

---

Input



\*



=

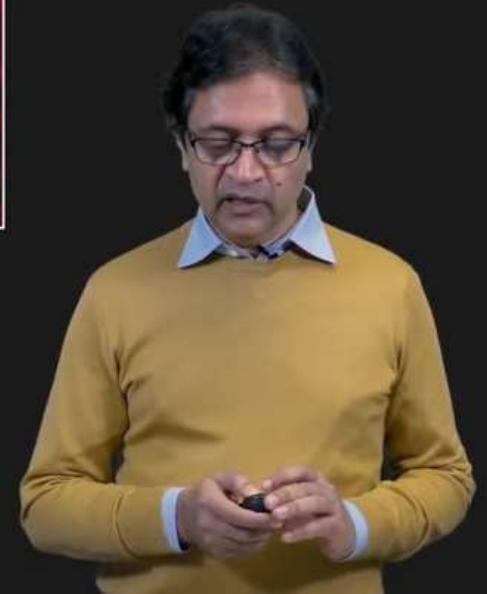


$f(x, y)$

$n_8(x, y)$

$g(x, y)$

Larger the kernel (or  $\sigma$ ), more the blurring



# Gaussian Smoothing is Separable

$$g[i, j] = \frac{1}{2\pi\sigma^2} \sum_{m=1}^K \sum_{n=1}^K e^{-\frac{1}{2}\left(\frac{m^2+n^2}{\sigma^2}\right)} f[i - m, j - n]$$

$$g[i, j] = \frac{1}{2\pi\sigma^2} \sum_{m=1}^K e^{-\frac{1}{2}\left(\frac{m^2}{\sigma^2}\right)} \cdot \sum_{n=1}^K e^{-\frac{1}{2}\left(\frac{n^2}{\sigma^2}\right)} f[i - m, j - n]$$



# Gaussian Smoothing is Separable

$$g[i,j] = \frac{1}{2\pi\sigma^2} \sum_{m=1}^K \sum_{n=1}^K e^{-\frac{1}{2}\left(\frac{m^2+n^2}{\sigma^2}\right)} f[i-m, j-n]$$

$$g[i,j] = \frac{1}{2\pi\sigma^2} \sum_{m=1}^K e^{-\frac{1}{2}\left(\frac{m^2}{\sigma^2}\right)} \cdot \sum_{n=1}^K e^{-\frac{1}{2}\left(\frac{n^2}{\sigma^2}\right)} f[i-m, j-n]$$

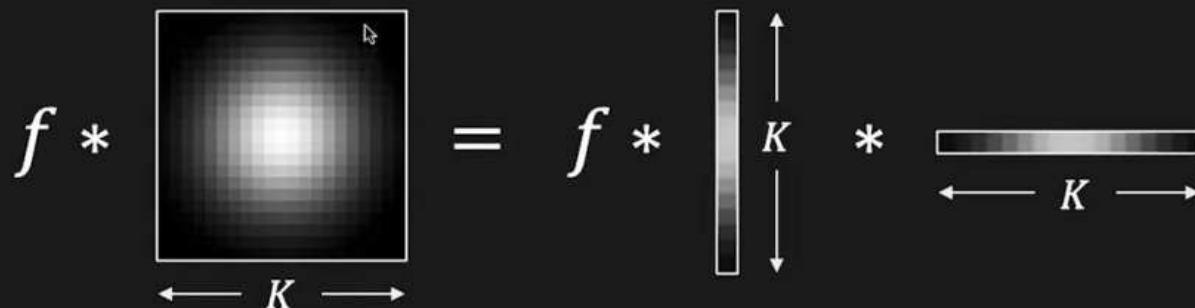
Using One 2D Gaussian Filter  $\equiv$  Using Two 1D Gaussian Filters

$$f * \begin{bmatrix} & & \\ & \text{Blur} & \\ & & \end{bmatrix} = f * \begin{bmatrix} & & \\ K & & \\ & & \end{bmatrix} * \begin{bmatrix} & & \\ & & \\ & & K \end{bmatrix}$$



# Gaussian Smoothing is Separable

Using One 2D Gaussian Filter  $\equiv$  Using Two 1D Gaussian Filters



Which one is faster? Why?

$K^2$  Multiplications

$K^2 - 1$  Additions

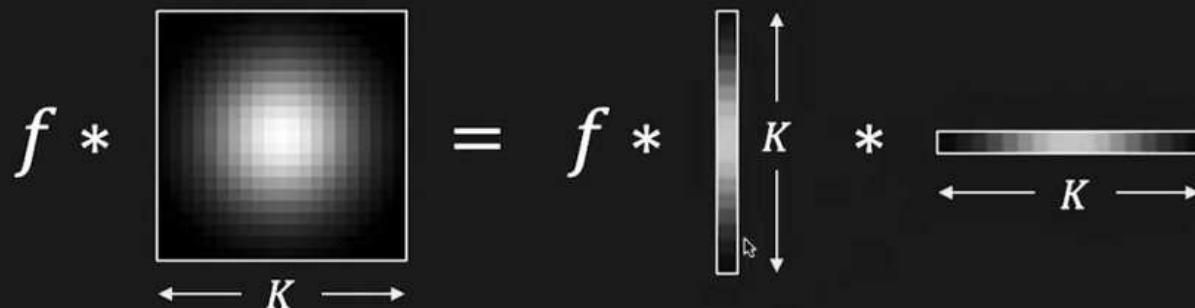
$2K$  Multiplications

$2(K - 1)$  Additions



# Gaussian Smoothing is Separable

Using One 2D Gaussian Filter  $\equiv$  Using Two 1D Gaussian Filters



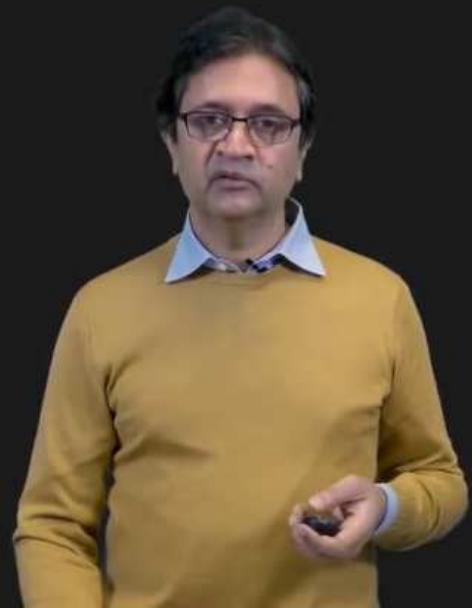
Which one is faster? Why?

$K^2$  Multiplications

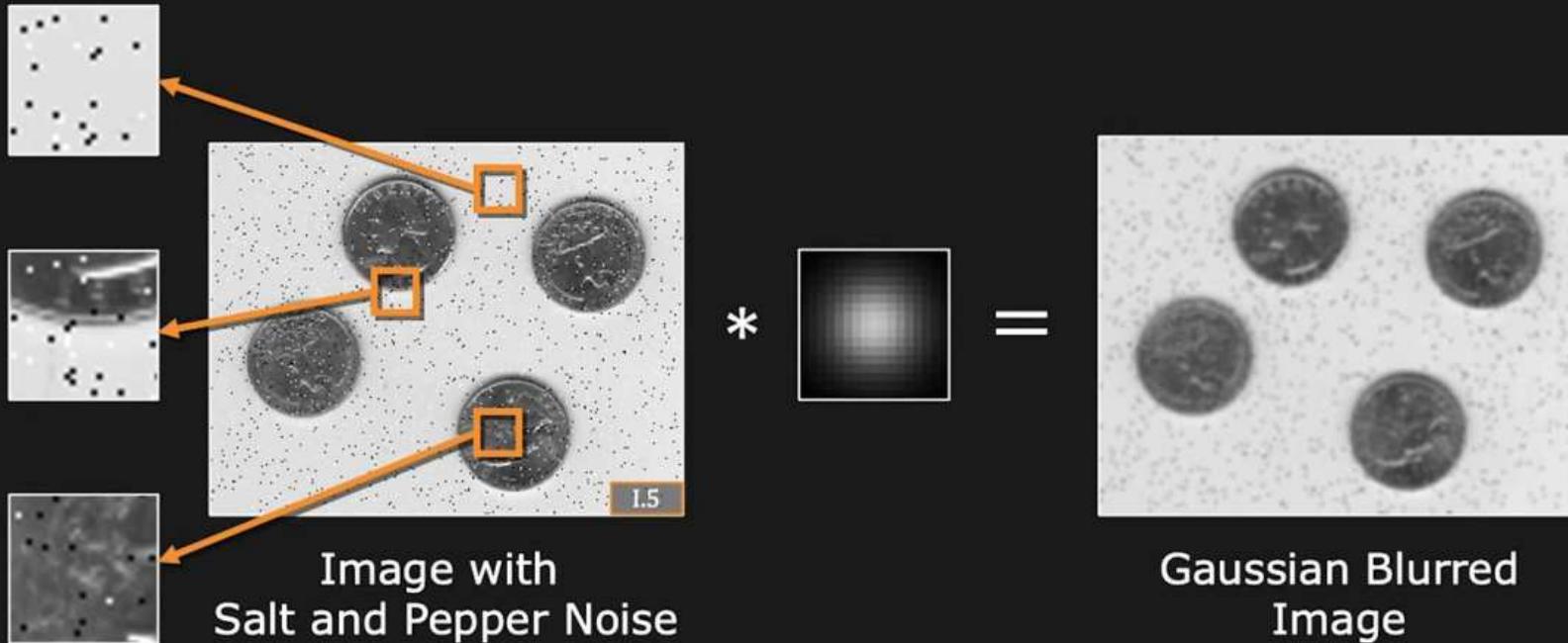
$K^2 - 1$  Additions

$2K$  Multiplications

$2(K - 1)$  Additions



# Smoothing to Remove Image Noise



Problem with Smoothing:

- Does not remove outliers (**Noise**)
- Smooths edges (**Blur**)



# Median Filtering

1. Sort the  $K^2$  values in window centered at the pixel
2. Assign the Middle Value (**Median**) to pixel

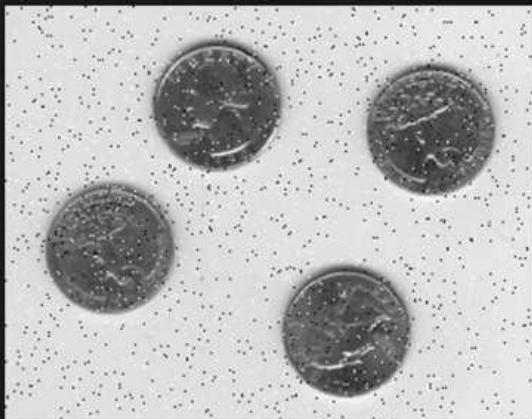


Image with  
Salt and Pepper Noise



Median Filtered  
Image ( $K = 3$ )

**Non-linear Operation**  
(Cannot be implemented using convolution)



# Median Filtering

---

Not Effective when Image Noise is not a Simple Salt and Pepper Noise.

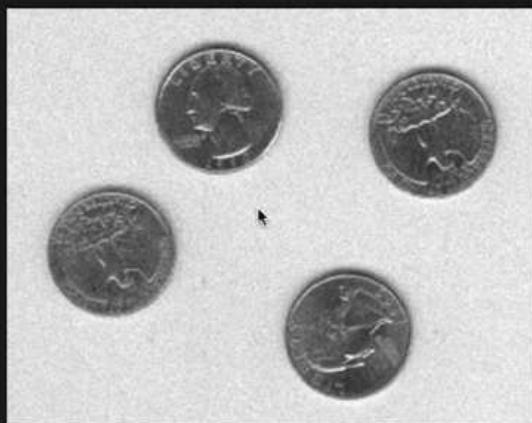
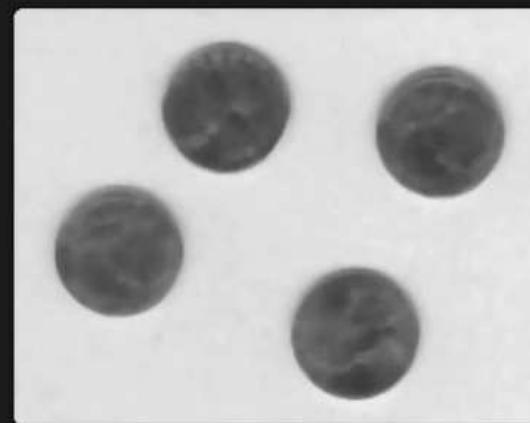
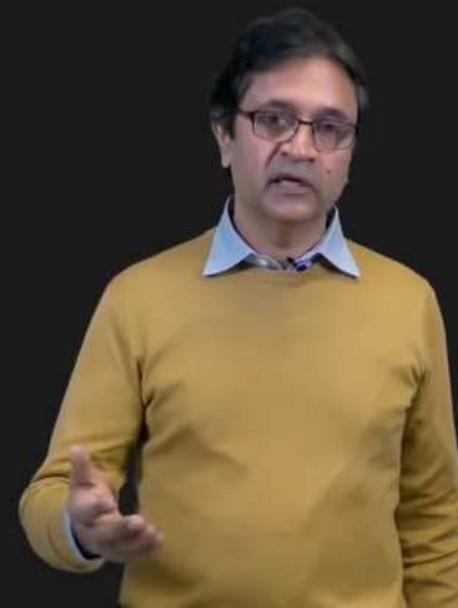


Image with Noise

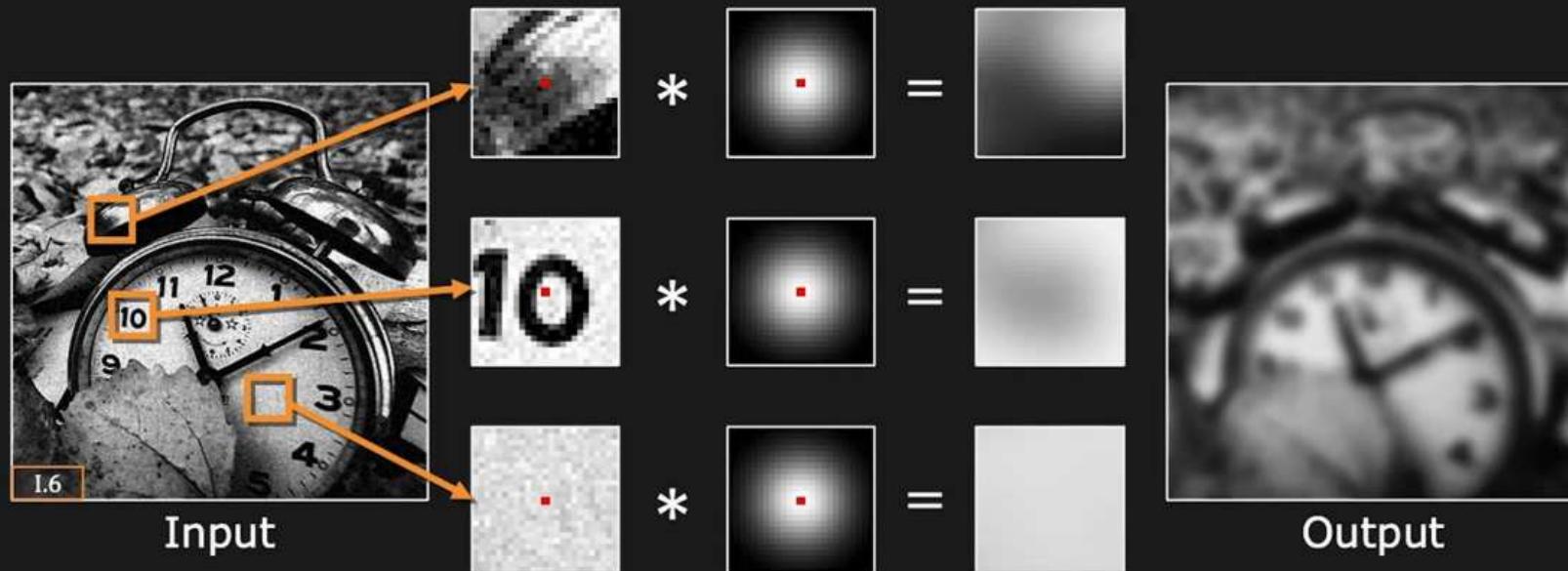


Median Filtered  
Image ( $K = 7$ )

Larger  $K$  causes blurring of image detail



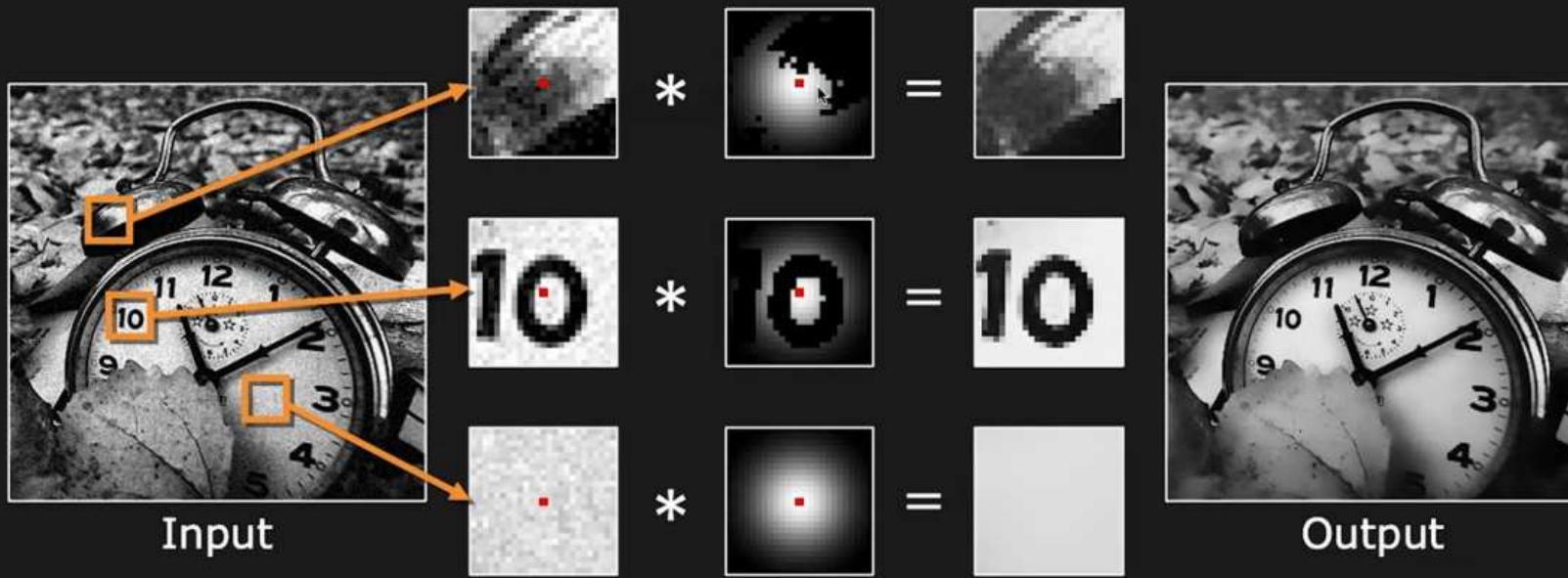
# Revisiting Gaussian Smoothing



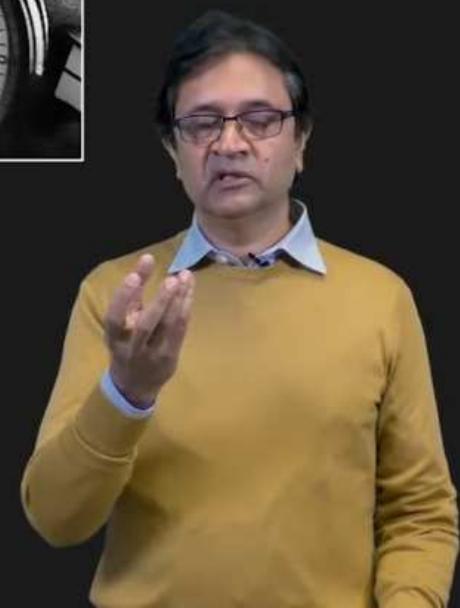
Same Gaussian kernel is used everywhere.  
Blurs across edges.



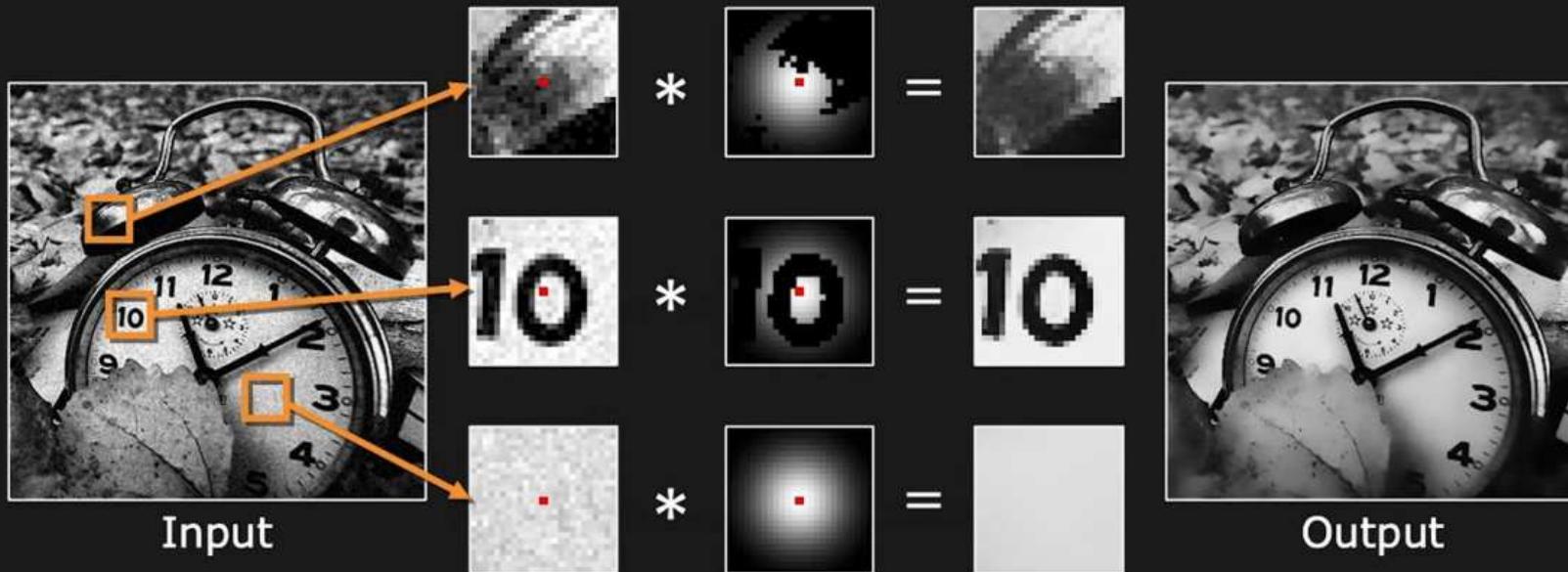
# Blur Similar Pixels Only



**"Bias"** Gaussian Kernel such that pixels not similar in intensity to the center pixel receive a lower weight.



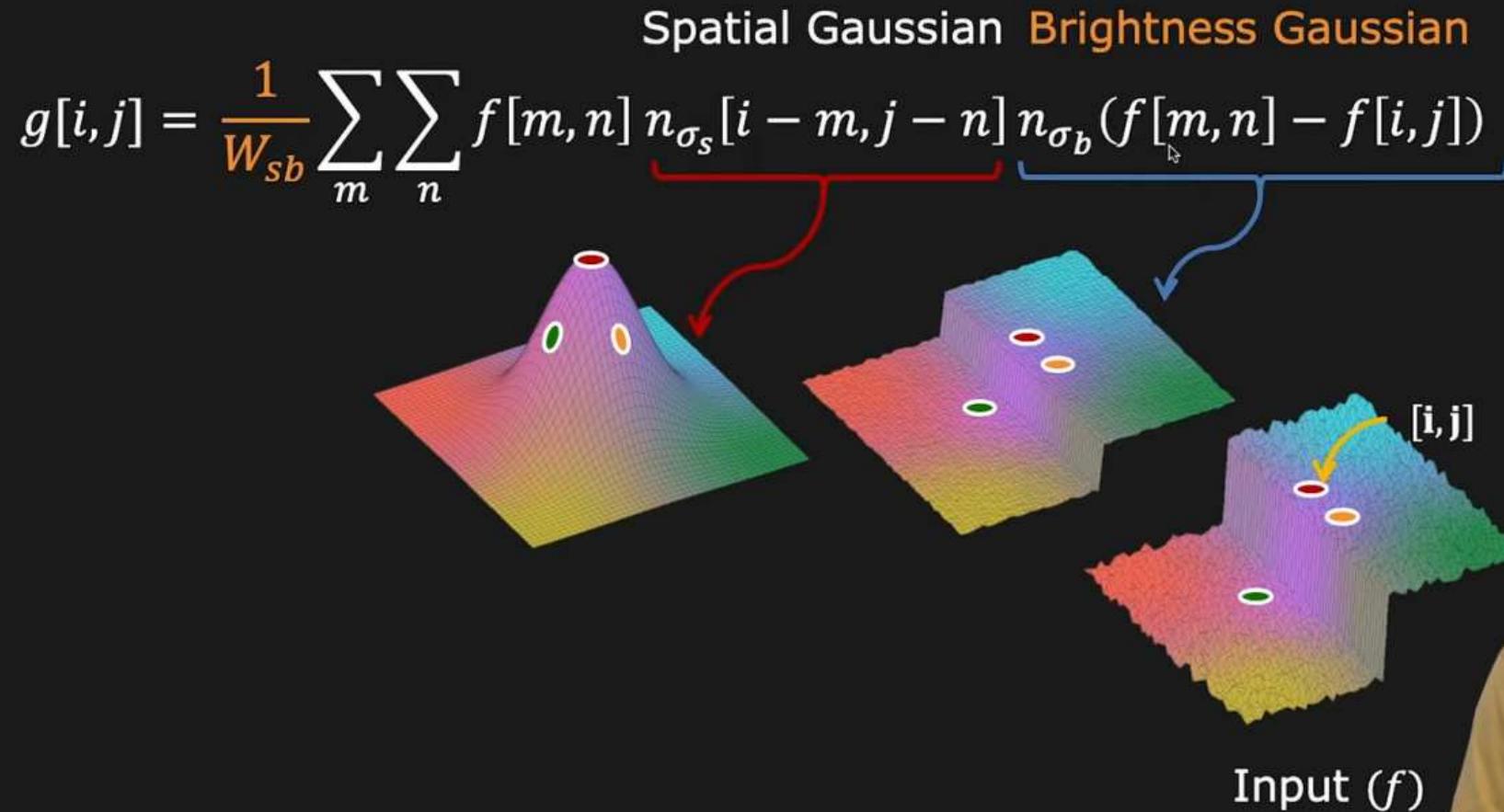
# Blur Similar Pixels Only



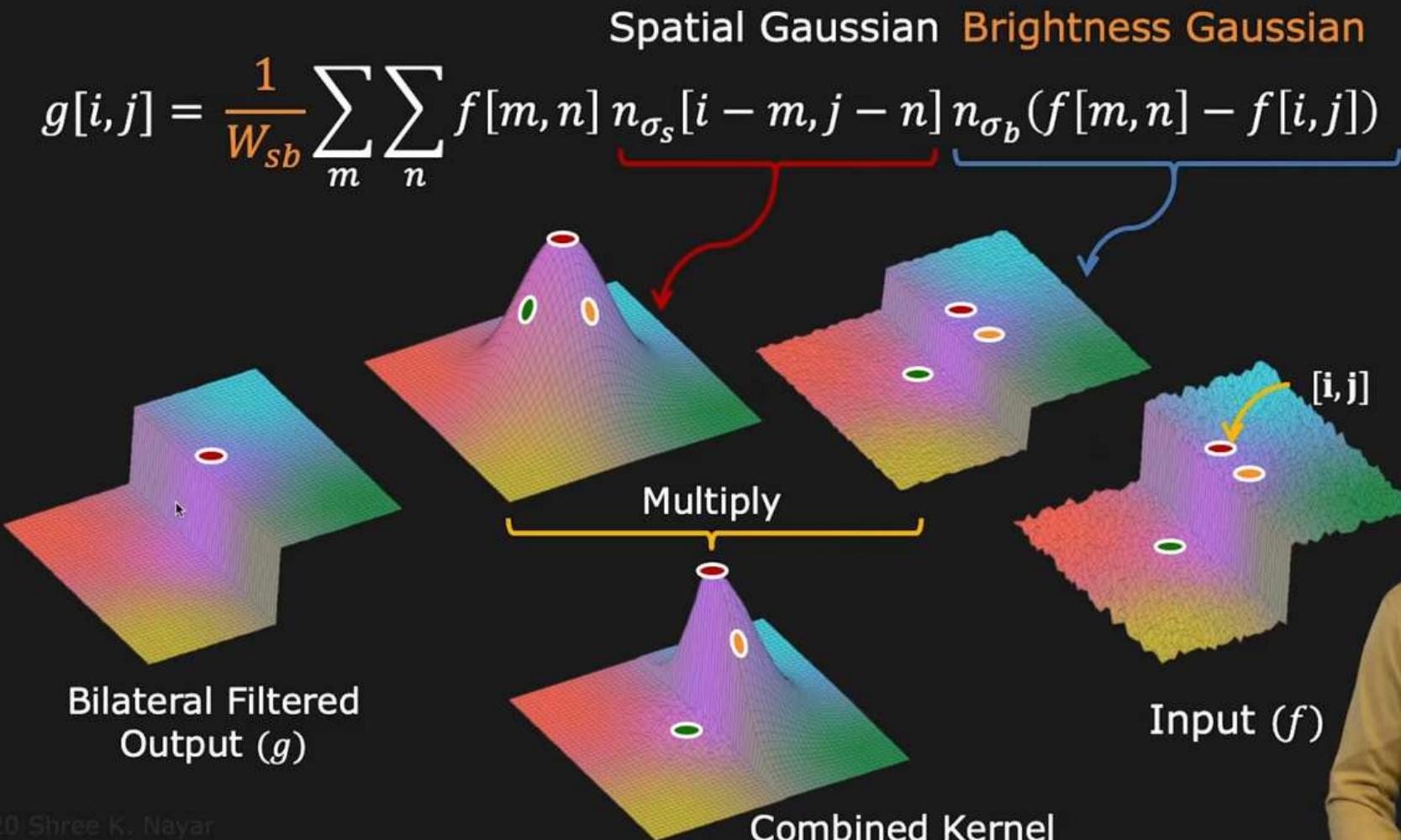
**"Bias"** Gaussian Kernel such that pixels not similar in intensity to the center pixel receive a lower weight.



# Bilateral Filter: Add Bias to Gaussian



# Bilateral Filter: Add Bias to Gaussian



# Bilateral Filter: Summary

$$g[i, j] = \frac{1}{W_{sb}} \sum_m \sum_n f[m, n] n_{\sigma_s}[i - m, j - n] n_{\sigma_b}(f[m, n] - f[i, j])$$

Where:

$$n_{\sigma_s}[m, n] = \frac{1}{2\pi\sigma_s^2} e^{-\frac{1}{2}\left(\frac{m^2+n^2}{\sigma_s^2}\right)} \quad n_{\sigma_b}(k) = \frac{1}{\sqrt{2\pi}\sigma_b} e^{-\frac{1}{2}\left(\frac{k^2}{\sigma_b^2}\right)}$$

$$W_{sb} = \sum_m \sum_n n_{\sigma_s}[i - m, j - n] n_{\sigma_b}(f[m, n] - f[i, j])$$



# Gaussian vs. Bilateral Filtering: Example



Original



Gaussian

$$\sigma_s = 2$$



Bilateral

$$\sigma_s = 2, \sigma_b = 10$$



# Bilateral Filtering: Changing $\sigma_b$

---



Bilateral  
 $\sigma_s = 6, \sigma_b = 10$



Bilateral  
 $\sigma_s = 6, \sigma_b = 20$



Bilateral  
 $\sigma_s = 6, \sigma_b = \infty$   
(Gaussian Smoothing)



# Template Matching

---



Template

How do we locate the template in the image?



# Template Matching



Template

How do we locate the template in the image?

Minimize:

$$E[i,j] = \sum_m \sum_n (f[m,n] - t[m-i,n-j])^2$$

$$E[i,j] = \sum_m \sum_n (f^2[m,n] + t^2[m-i,n-j] - \underbrace{2f[m,n]t[m-i,n-j]}_{\text{Maximize}})$$



# Template Matching



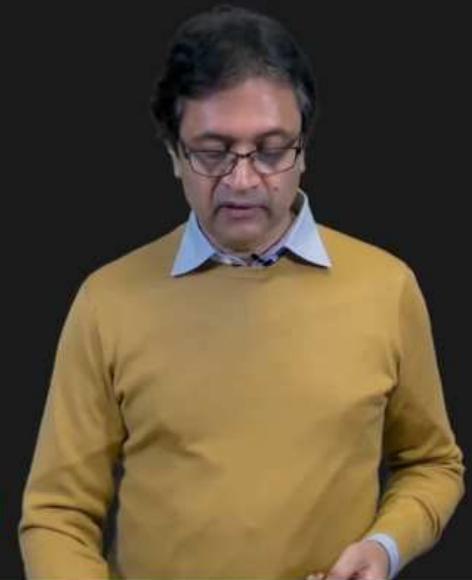
Template

How do we locate the template in the image?

Maximize:

$$R_{tf}[i,j] = \sum_m \sum_n f[m,n]t[m-i,n-j] = t \otimes f$$

(Cross-Correlation)



# Convolution vs. Correlation

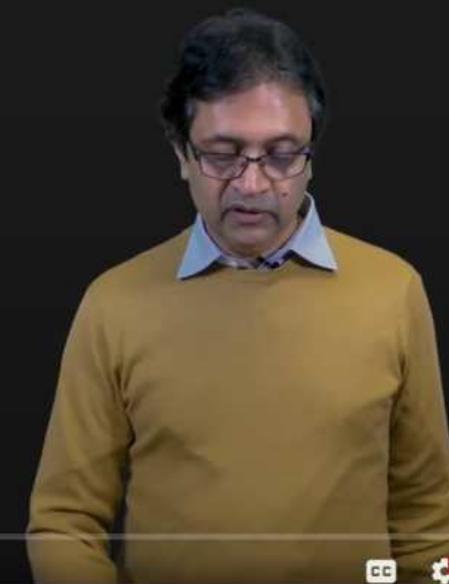
Convolution:

$$g[i, j] = \sum_m \sum_n f[m, n] t[i - m, j - n] = t * f$$

Correlation:

$$R_{tf}[i, j] = \sum_m \sum_n f[m, n] t[m - i, n - j] = t \otimes f$$

No Flipping in Correlation



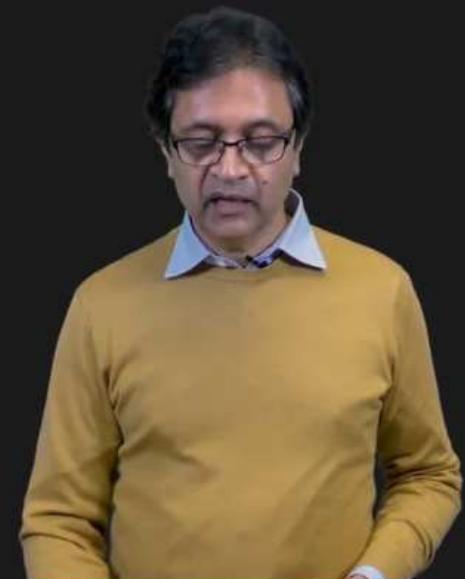
# Problem with Cross-Correlation

$$R_{tf}[i,j] = \sum_m \sum_n f[m,n]t[m-i,n-j] = t \otimes f$$



$$R_{tf}(C) > R_{tf}(B) > R_{tf}(A)$$

We need  $R_{tf}(A)$  to be the maximum!



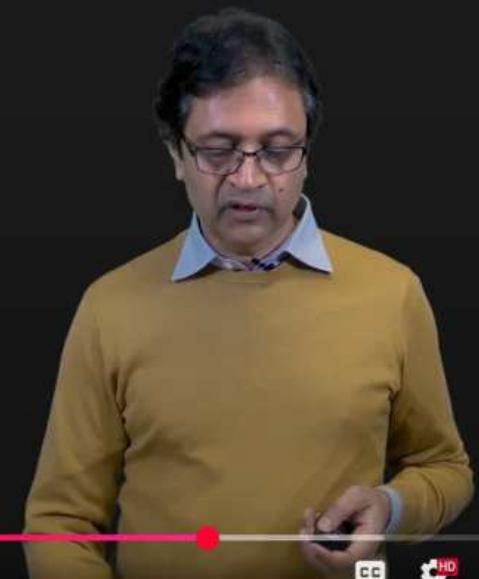
# Normalized Cross-Correlation

Account for energy differences

$$N_{tf}[i, j] = \frac{\sum_m \sum_n f[m, n] t[m - i, n - j]}{\sqrt{\sum_m \sum_n f^2[m, n]} \sqrt{\sum_m \sum_n t^2[m - i, n - j]}}$$



$$N_{tf}(A) > N_{tf}(B) > N_{tf}(C)$$



# Normalized Cross-Correlation

Account for energy differences

$$N_{tf}[i, j] = \frac{\sum_m \sum_n f[m, n] t[m - i, n - j]}{\sqrt{\sum_m \sum_n f^2[m, n]} \sqrt{\sum_m \sum_n t^2[m - i, n - j]}}$$



$$\otimes \quad \text{[Template Image]} \quad =$$



# Recap: Image Processing I

---

Transform image to new one that is clearer or easier to analyze.

## Topics:

- (1) Pixel Processing
- (2) LSIS and Convolution
- (3) Linear Image Filters
- (4) Non-Linear Image Filters
- (5) Template Matching by Correlation



# Image Processing II

---

Transform image to new one that is clearer or easier to analyze.

## Topics:

- (1) Fourier Transform
- (2) Convolution Theorem



# Image Processing II

---

Transform image to new one that is clearer or easier to analyze.

## Topics:

- (1) Fourier Transform
- (2) Convolution Theorem
- (3) Deconvolution in Frequency Domain
- (4) Sampling Theory and Aliasing



# Jean Baptiste Joseph Fourier

---



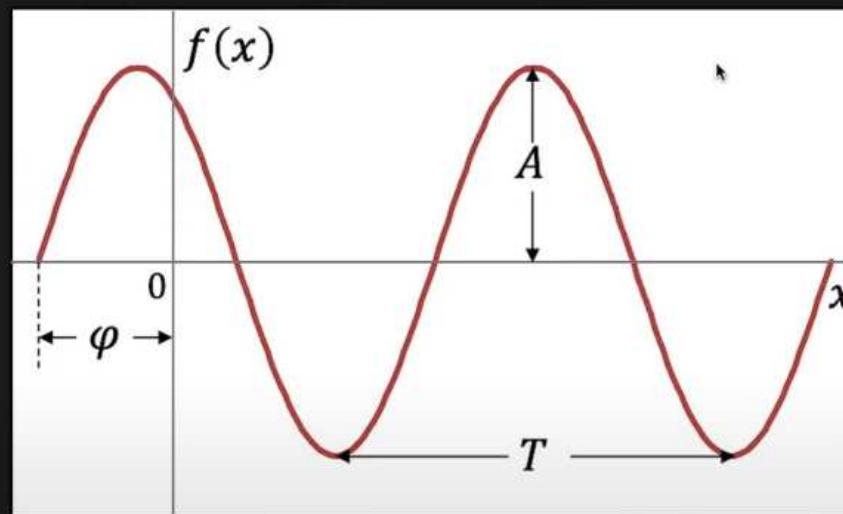
(1768-1830)

Any **Periodic Function** can be rewritten as a **Weighted Sum**  
of **Infinite Sinusoids** of **Different Frequencies**.



# Sinusoid

$$f(x) = A \sin(2\pi ux + \varphi)$$



$A$ : Amplitude

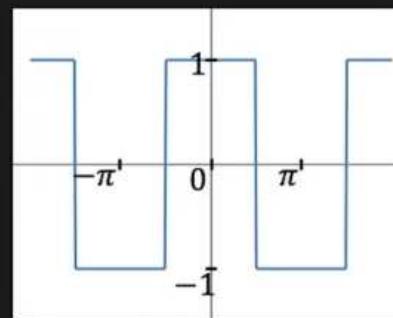
$T$ : Period

$\varphi$ : Phase

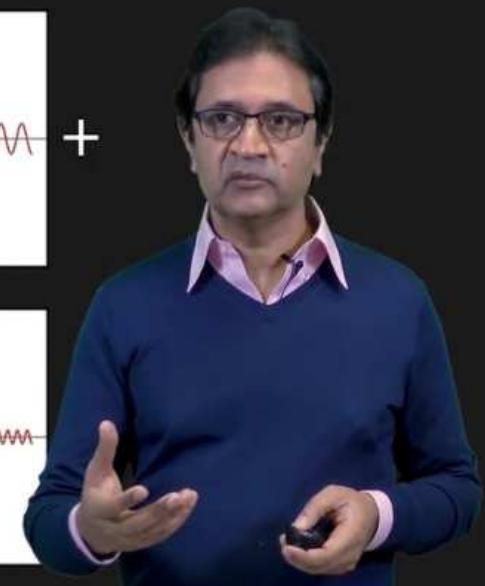
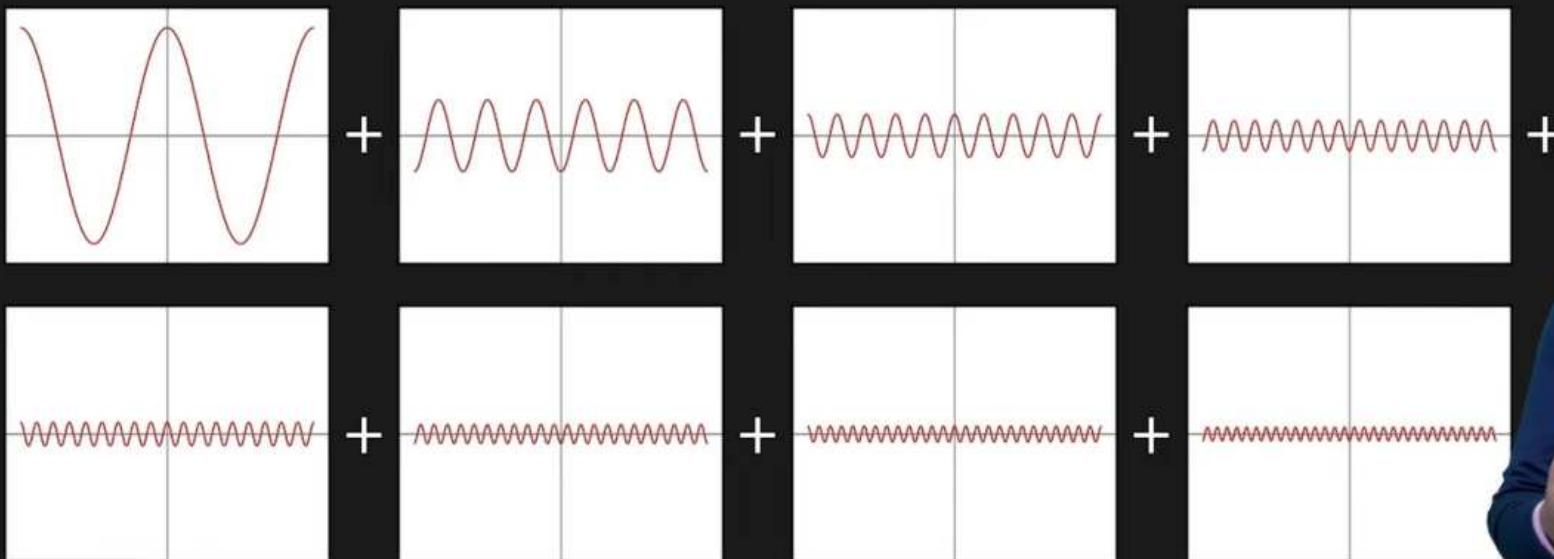
$u$ : Frequency ( $1/T$ )



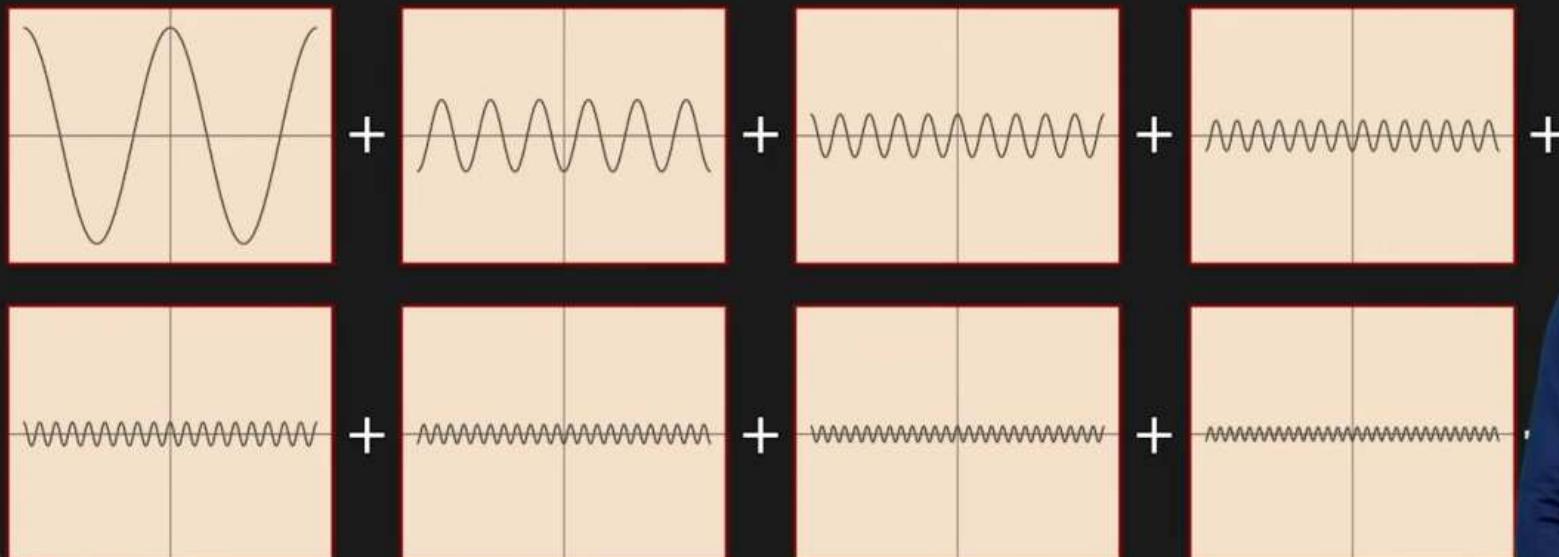
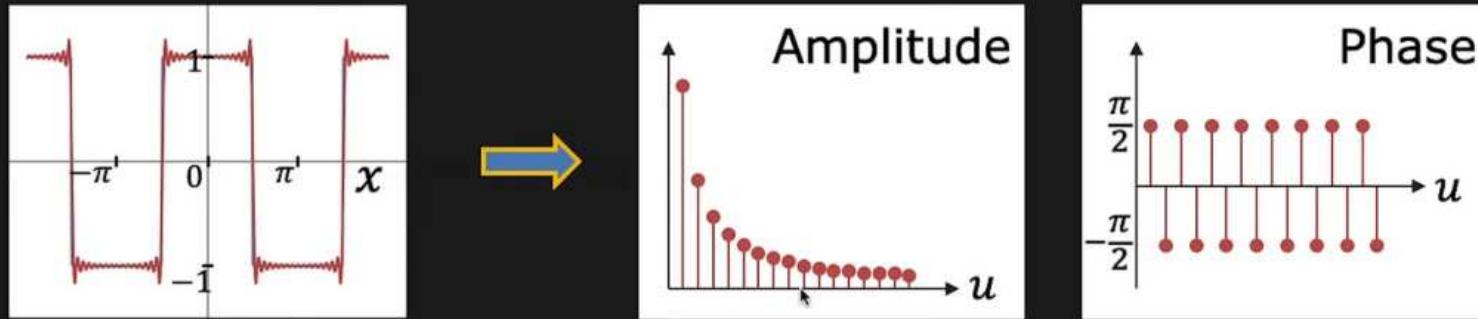
# Fourier Series



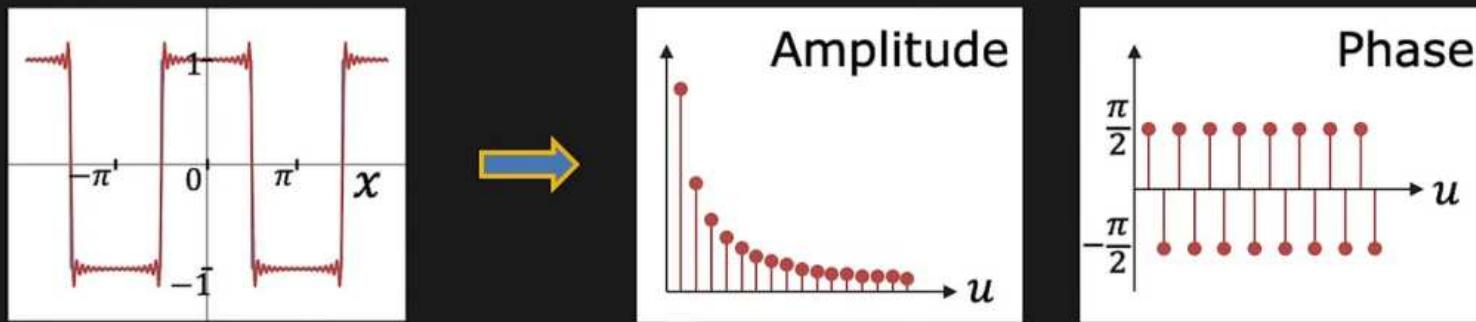
Square Wave  
(Period  $2\pi$ )



# Frequency Representation of Signal



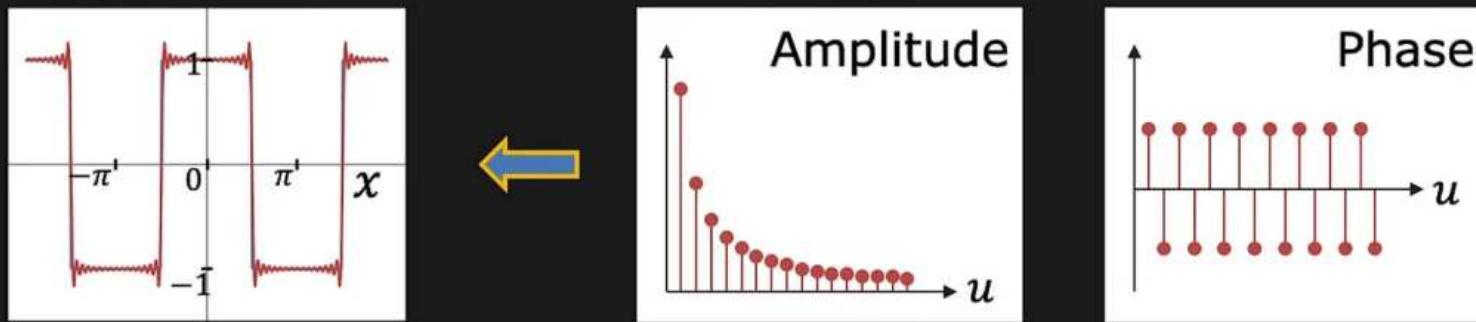
# Fourier Transform (FT)



Represents a signal  $f(x)$  in terms of Amplitudes and Phases of its Constituent Sinusoids.



# Inverse Fourier Transform (IFT)



Computes the signal  $f(x)$  from the Amplitudes and Phases of its Constituent Sinusoids.



# Finding FT and IFT

---

Fourier Transform:

$$F(u) = \int_{-\infty}^{\infty} f(x)e^{-i2\pi ux}dx$$

$x$ : space

$u$ : frequency

$$e^{i\theta} = \cos \theta + i \sin \theta$$

$$i = \sqrt{-1}$$

Inverse Fourier Transform:

$$f(x) = \int_{-\infty}^{\infty} F(u)e^{i2\pi ux}du$$



# Complex Exponential (Euler Formula)

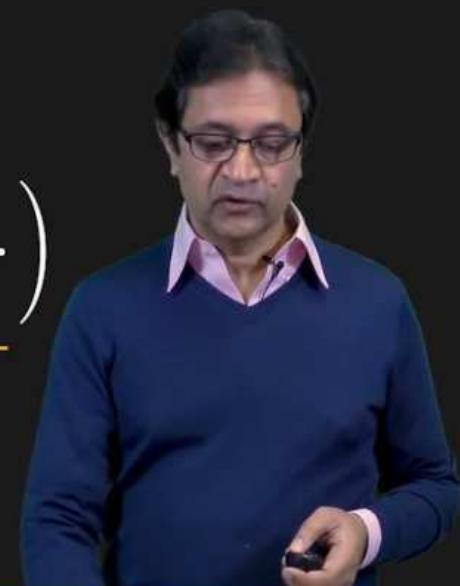
$$e^{i\theta} = \cos \theta + i \sin \theta$$

$$i = \sqrt{-1}$$

Expand  $e^{i\theta}$  using Taylor Series:

$$e^{i\theta} = 1 + i\theta + \frac{(i\theta)^2}{2!} + \frac{(i\theta)^3}{3!} + \frac{(i\theta)^4}{4!} + \frac{(i\theta)^5}{5!} + \frac{(i\theta)^6}{6!} + \dots$$

$$e^{i\theta} = \underbrace{\left(1 - \frac{\theta^2}{2!} + \frac{\theta^4}{4!} - \frac{\theta^6}{6!} + \dots\right)}_{\cos \theta} + \underbrace{i\left(\theta - \frac{\theta^3}{3!} + \frac{\theta^5}{5!} - \frac{\theta^7}{7!} + \dots\right)}_{\sin \theta}$$



# Fourier Transform is Complex!

$F(u)$  holds the **Amplitude** and **Phase** of the sinusoid of frequency  $u$ .

$$F(u) = \int_{-\infty}^{\infty} f(x)e^{-i2\pi ux}dx$$

$$F(u) = \Re{F(u)} + i \Im{F(u)}$$

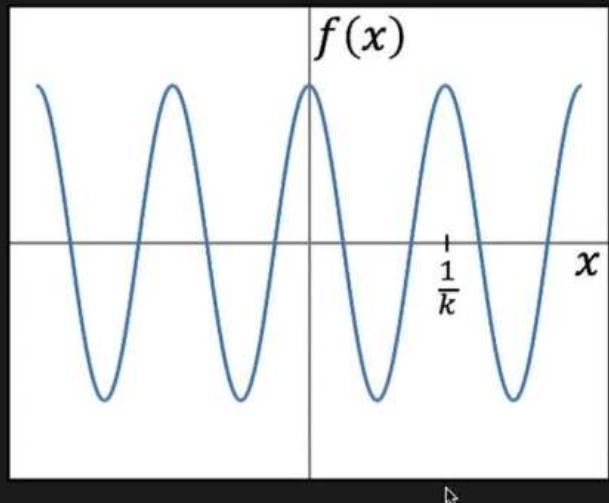
**Amplitude:**  $A(u) = \sqrt{\Re{F(u)}^2 + \Im{F(u)}^2}$

**Phase:**  $\varphi(u) = \text{atan2}(\Im{F(u)}, \Re{F(u)})$



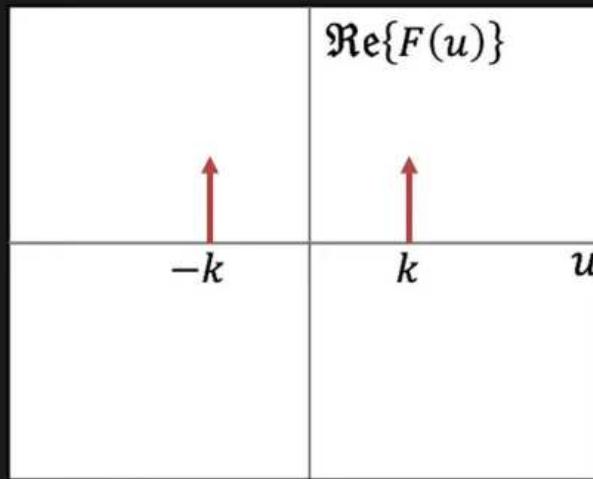
# Fourier Transform Examples

Signal  $f(x)$



$$f(x) = \cos 2\pi kx$$

Fourier Transform  $F(u)$

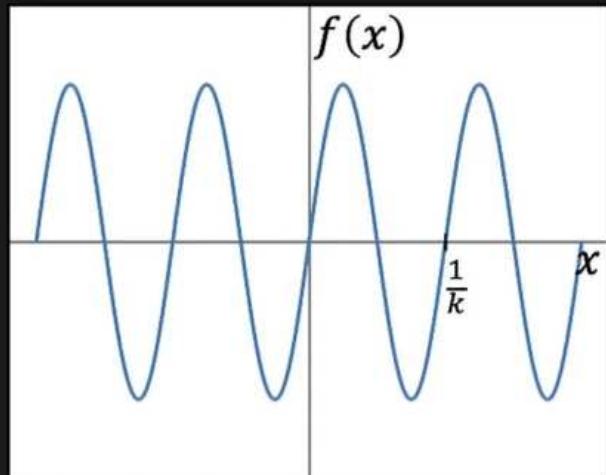


$$F(u) = \frac{1}{2}[\delta(u + k) + \delta(u - k)]$$



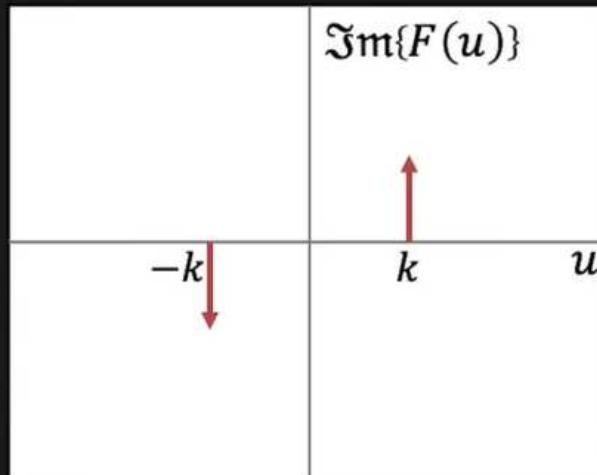
# Fourier Transform Examples

Signal  $f(x)$

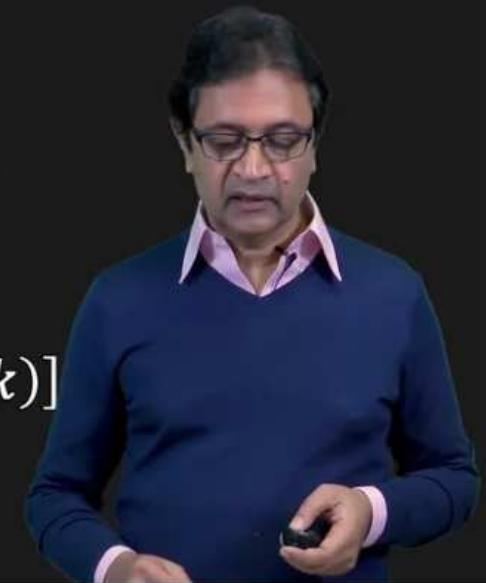


$$f(x) = \sin 2\pi kx$$

Fourier Transform  $F(u)$

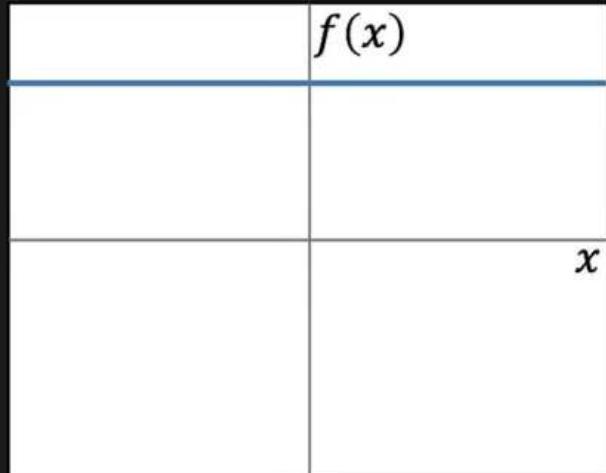


$$F(u) = \frac{1}{2}i[\delta(u + k) - \delta(u - k)]$$



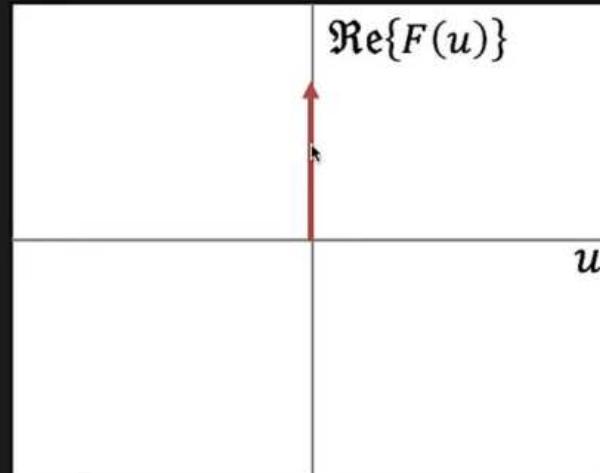
# Fourier Transform Examples

Signal  $f(x)$



$$f(x) = 1$$

Fourier Transform  $F(u)$

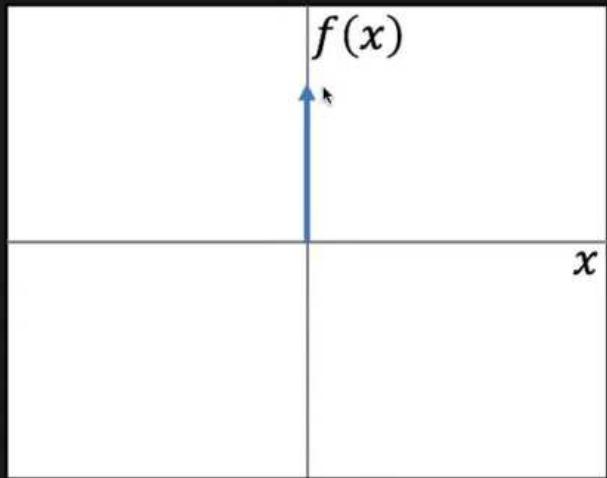


$$F(u) = \delta(u)$$



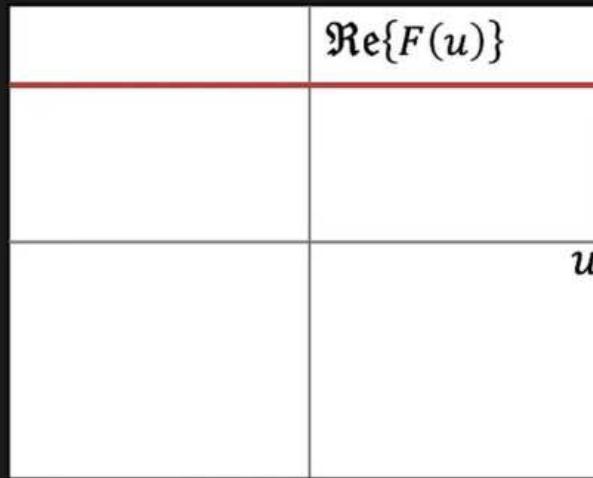
# Fourier Transform Examples

Signal  $f(x)$



$$f(x) = \delta(x)$$

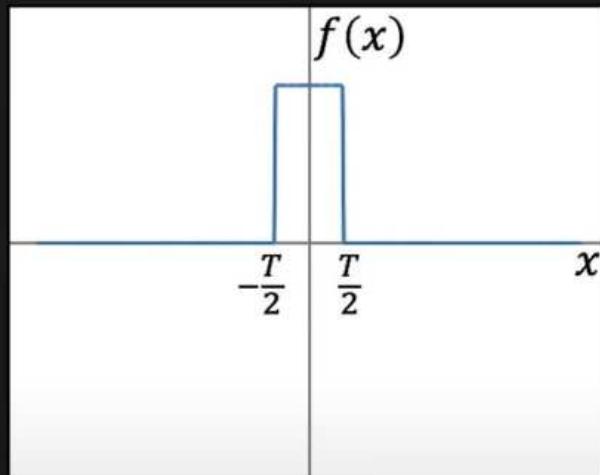
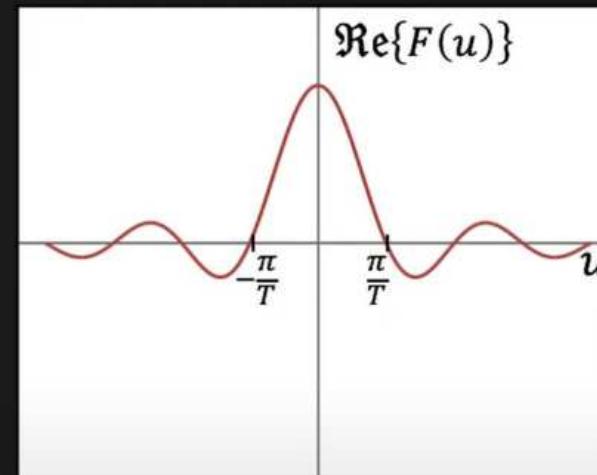
Fourier Transform  $F(u)$



$$F(u) = 1$$



# Fourier Transform Examples

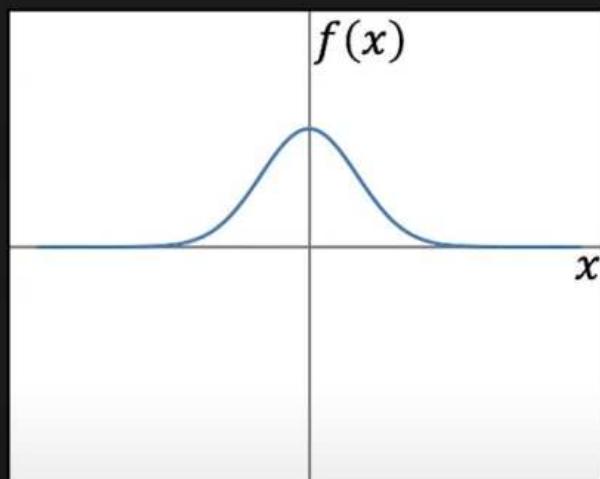
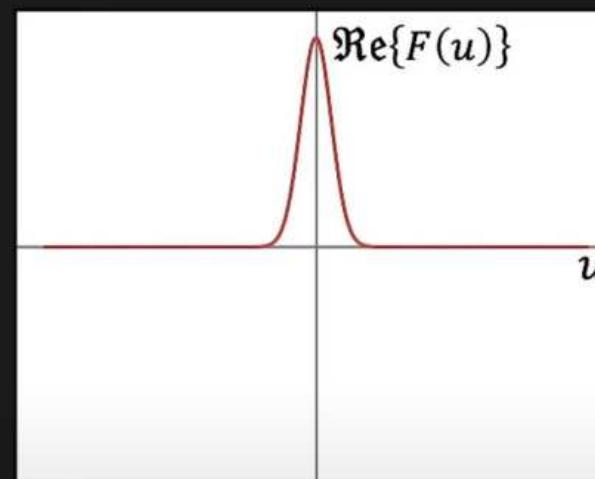
Signal  $f(x)$ Fourier Transform  $F(u)$ 

$$f(x) = \text{Rect}\left(\frac{x}{T}\right)$$

$$F(u) = T \operatorname{sinc} Tu$$



# Fourier Transform Examples

Signal  $f(x)$ Fourier Transform  $F(u)$ 

$$f(x) = e^{-ax^2}$$

$$F(u) = \sqrt{\pi/a} e^{-\pi^2 u^2 / a}$$



# Properties of Fourier Transform

Property	Spatial Domain	Frequency Domain
Linearity	$\alpha f_1(x) + \beta f_2(x)$	$\alpha F_1(u) + \beta F_2(u)$
Scaling	$f(ax)$	$\frac{1}{ a } F\left(\frac{u}{a}\right)$
Shifting	$f(x - a)$	$e^{-i2\pi u a} F(u)$
Differentiation	$\frac{d^n}{dx^n}(f(x))$	$(i2\pi u)^n F(u)$



# Convolution

Convolution of two functions  $f(x)$  and  $h(x)$

$$g(x) = f(x) * h(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$$

$f(\tau)$



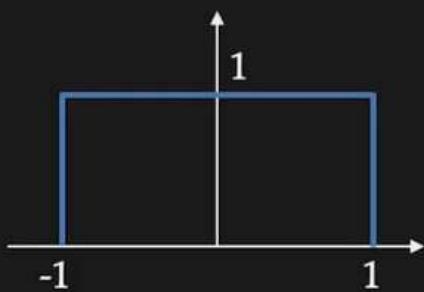
$h(\tau)$



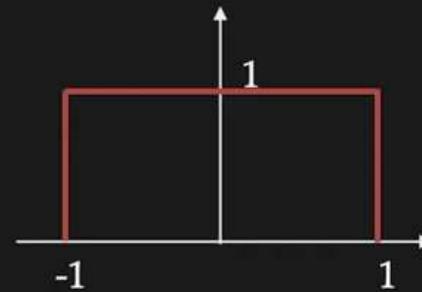
▶▶ Jump ahead

# Convolution: Example

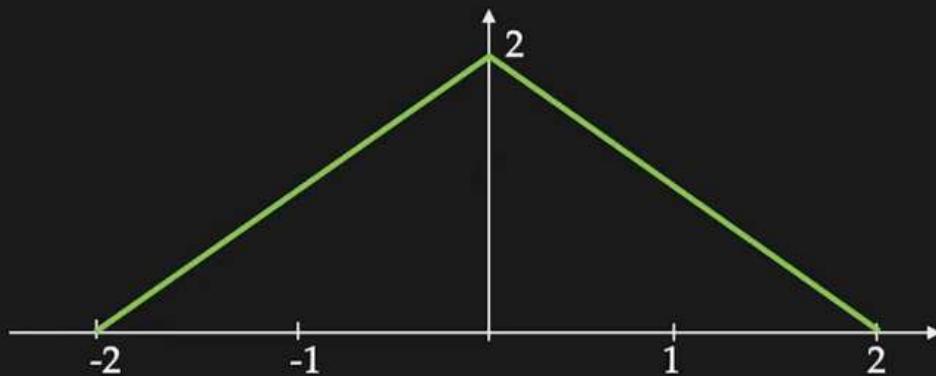
---



$$f(x)$$



$$h(x)$$



$$f(x) * h(x)$$



# Convolution and Fourier Transform

Convolution:  $g(x) = f(x) * h(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$

Fourier Transform of  $g(x)$ :

$$G(u) = \int_{-\infty}^{\infty} g(x)e^{-i2\pi ux} dx$$

$$G(u) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(\tau)h(x - \tau)e^{-i2\pi ux} d\tau dx$$

$$G(u) = \int_{-\infty}^{\infty} f(\tau)e^{-i2\pi u\tau} d\tau \int_{-\infty}^{\infty} h(x - \tau)e^{-i2\pi u(x - \tau)} dx$$



# Convolution and Fourier Transform

---

Convolution:  $g(x) = f(x) * h(x) = \int_{-\infty}^{\infty} f(\tau)h(x - \tau) d\tau$

Fourier Transform of  $g(x)$ :

$$G(u) = \int_{-\infty}^{\infty} g(x)e^{-i2\pi ux} dx$$

$$G(u) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(\tau)h(x - \tau)e^{-i2\pi ux} d\tau dx$$

$$G(u) = \int_{-\infty}^{\infty} f(\tau)e^{-i2\pi u\tau} d\tau \quad \int_{-\infty}^{\infty} h(x - \tau)e^{-i2\pi u(x - \tau)} dx$$

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$\triangleleft F(u)$

$H(u)$

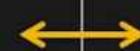


# Convolution and Fourier Transform

Spatial Domain

Frequency Domain

$$g(x) = f(x) * h(x)$$



$$G(u) = F(u) H(u)$$

Convolution

Multiplication

$$g(x) = f(x) \cdot h(x)$$



$$G(u) = F(u) * H(u)$$

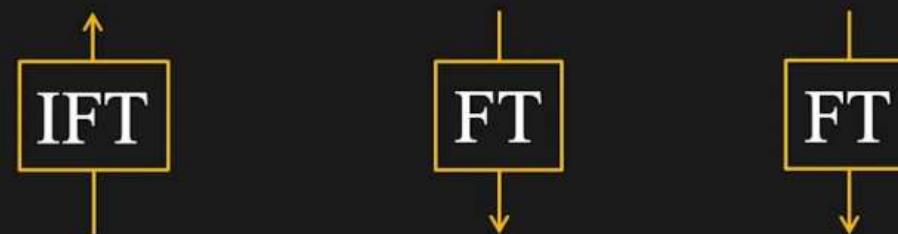
Multiplication

Convolution



# Convolution Using Fourier Transform

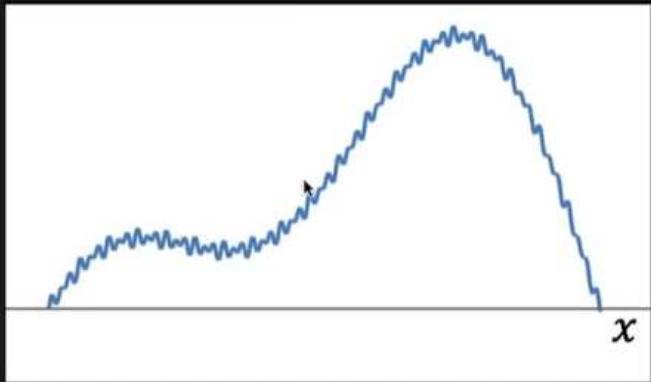
$$g(x) = f(x) * h(x)$$



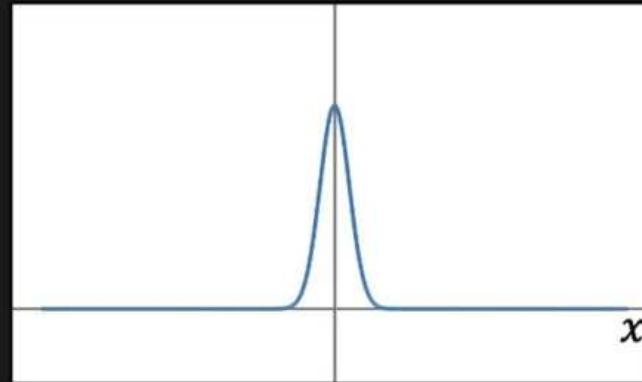
$$G(u) = F(u) \times H(u)$$



# Gaussian Smoothing in Fourier Domain



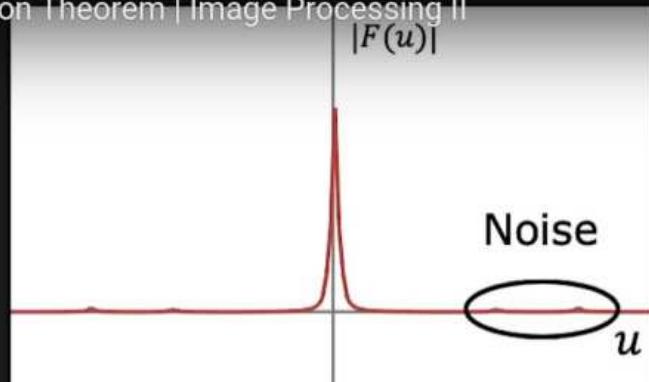
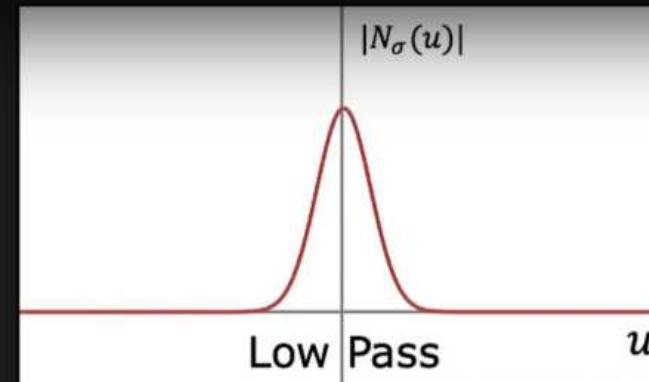
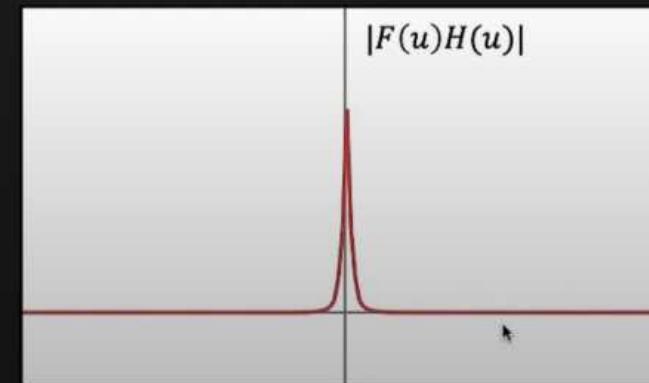
Noisy Signal  $f(x)$

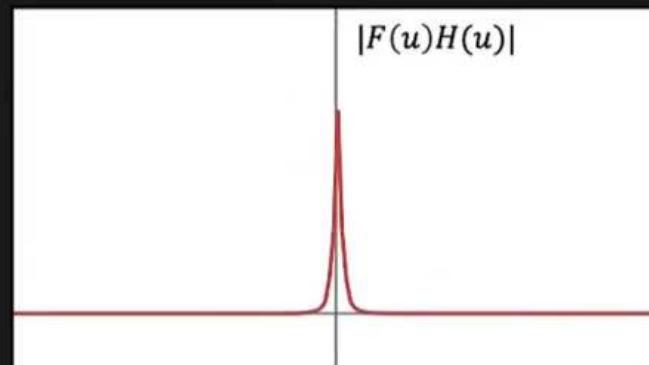


Gaussian Kernel  $n_\sigma(x)$

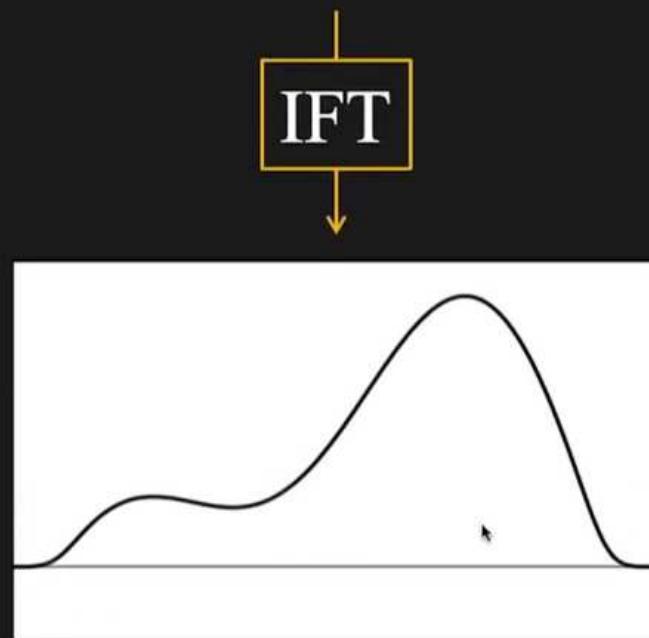
Convolve the Noisy Signal with a Gaussian Kernel



 $F(u)$  $N_\sigma(u)$  $F(u)H(u)$ 

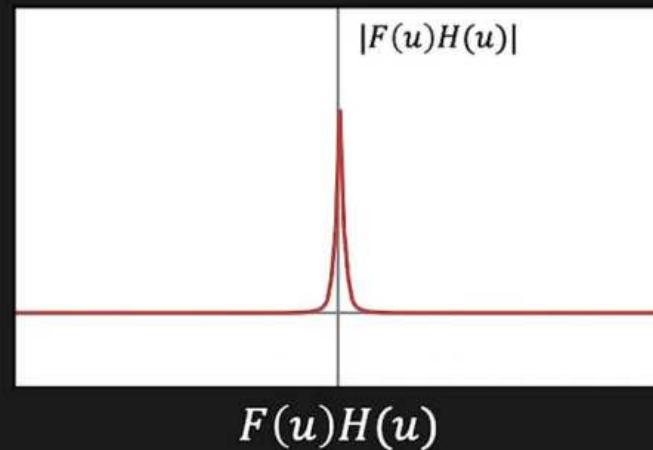


$F(u)H(u)$

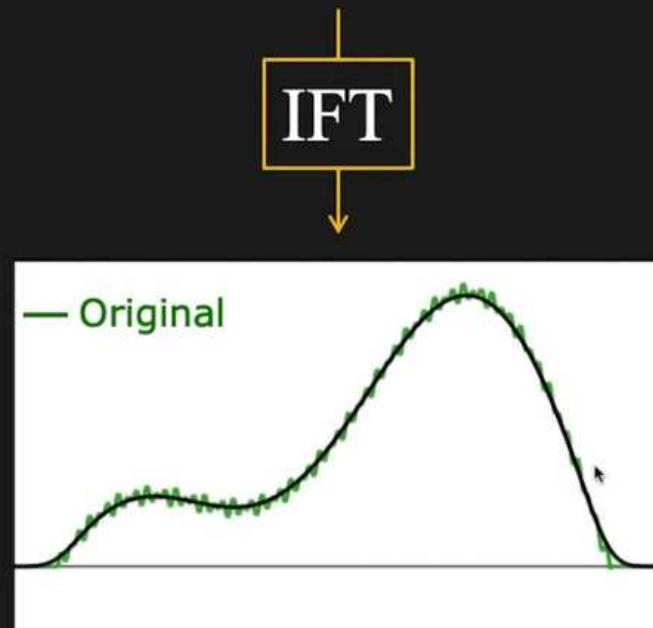


Gaussian Blurred Signal  $g(x)$





$F(u)H(u)$



Gaussian Blurred Signal  $g(x)$

IFT



# 2D Fourier Transform

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Fourier Transform:

$$F(u, v) = \iint_{-\infty}^{\infty} f(x, y) e^{-i2\pi(ux+vy)} dx dy$$

$u$  and  $v$  are frequencies along  $x$  and  $y$ , respectively

Inverse Fourier Transform:

$$f(x, y) = \iint_{-\infty}^{\infty} F(u, v) e^{i2\pi(xu+yv)} du dv$$



# 2D Fourier Transform: Discrete Images

Discrete Fourier Transform (DFT):

$$F[p, q] = \sum_{m=0}^{M-1} \sum_{n=0}^{N-1} f[m, n] e^{-i2\pi pm/M} e^{-i2\pi qn/N}$$

$$p = 0 \dots M - 1$$

$$q = 0 \dots N - 1$$

$p$  and  $q$  are frequencies along  $m$  and  $n$ , respectively

Inverse Discrete Fourier Transform (IDFT):

$$f[m, n] = \frac{1}{MN} \sum_{p=0}^{M-1} \sum_{q=0}^{N-1} F[p, q] e^{i2\pi pm/M} e^{i2\pi qn/N}$$

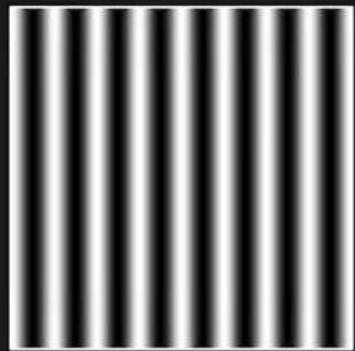
$$m = 0 \dots M -$$

$$n = 0 \dots N -$$

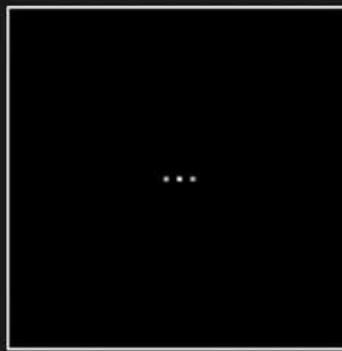


# 2D Fourier Transform: Example 1

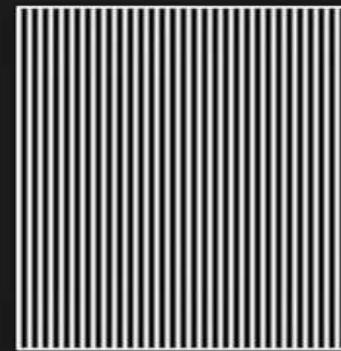
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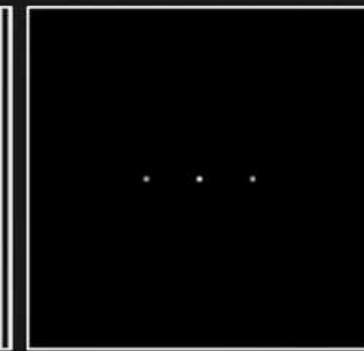
$f(m, n)$



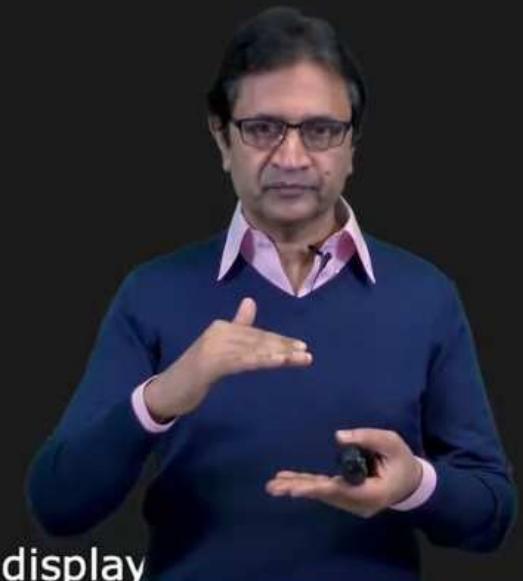
$\log(|F(p, q)|)$



$g(m, n)$

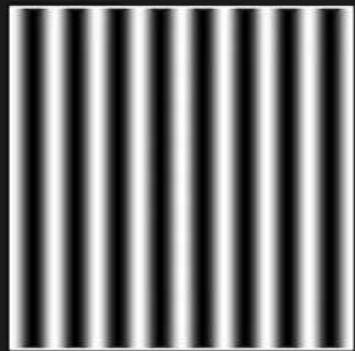


$\log(|G(p, q)|)$



**Note:**  $\log(|F|)$  is used just for display

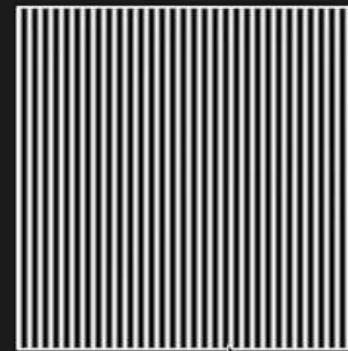
# 2D Fourier Transform: Example 1



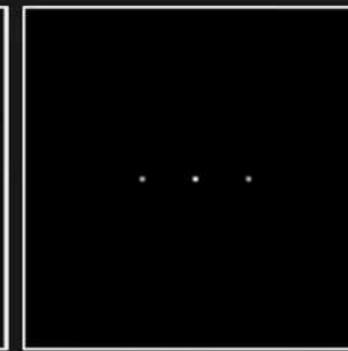
$f(m, n)$



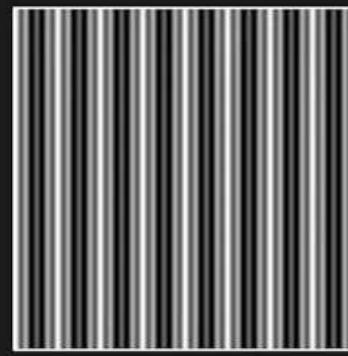
$\log(|F(p, q)|)$



$g(m, n)$



$\log(|G(p, q)|)$



$f(m, n) + g(m, n)$

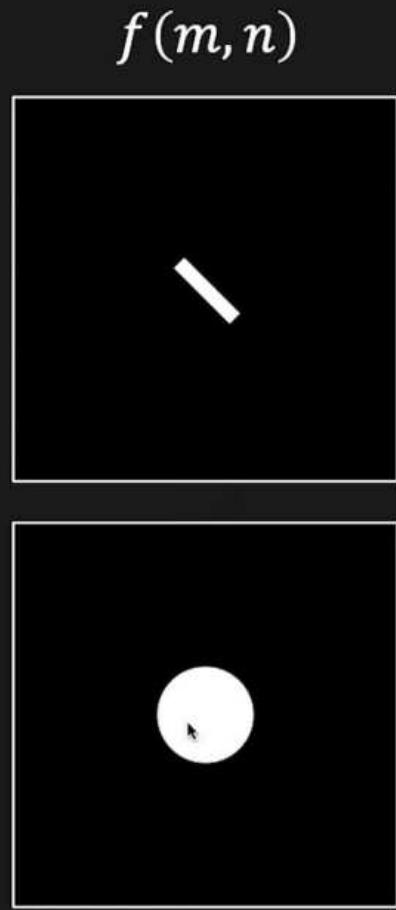


$\log(|F(p, q) + G(p, q)|)$

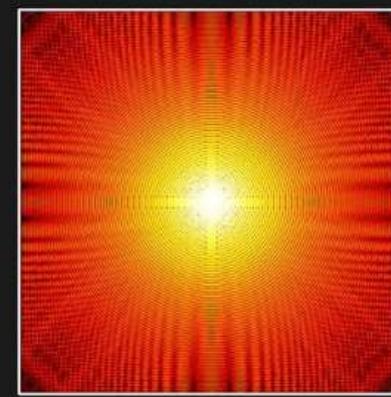
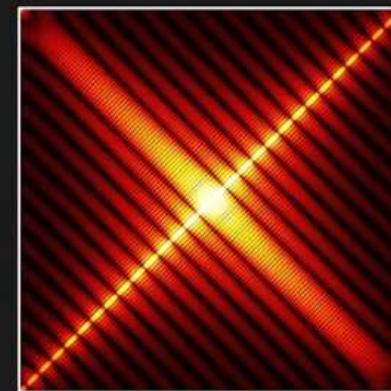


**Note:**  $\log(|F|)$  is used just for dis

# 2D Fourier Transform: Example 2



$\log(|F(p, q)|)$



Min



Max

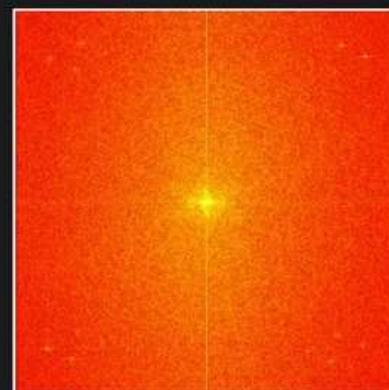
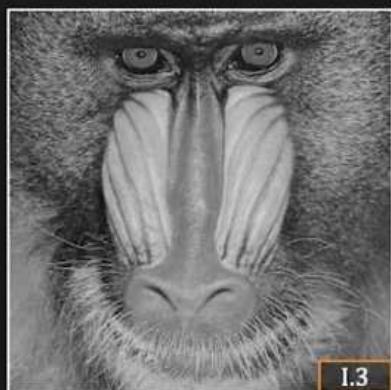
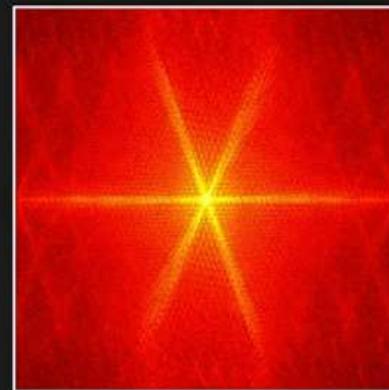


# 2D Fourier Transform: Example 3

$f(m, n)$



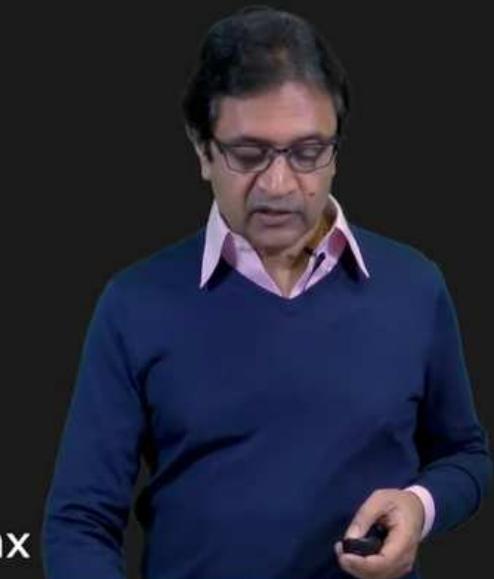
$\log(|F(p, q)|)$



Min



Max

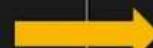
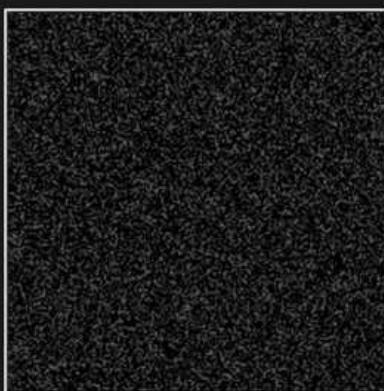
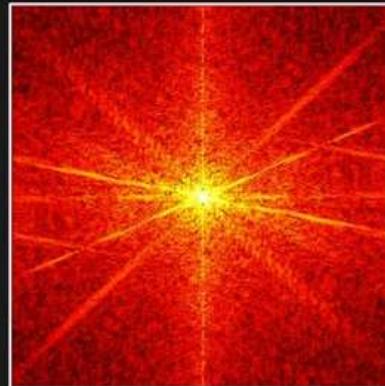


# 2D Fourier Transform: Example 4

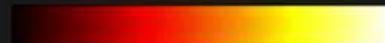
$f(m, n)$



$\log(|F(p, q)|)$



Min



Max

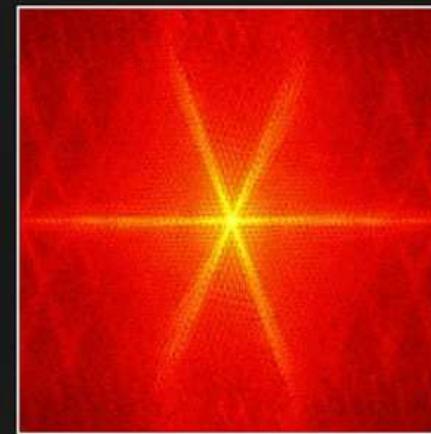


# Low Pass Filtering

$f(m, n)$



$\log(|F(p, q)|)$

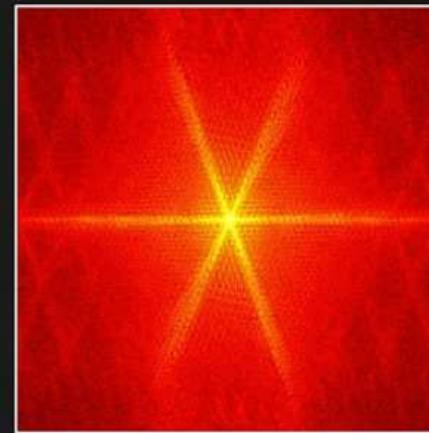


# Low Pass Filtering

$f(m, n)$

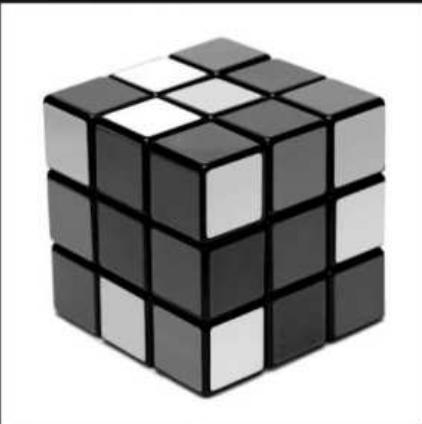


$\log(|F(p, q)|)$

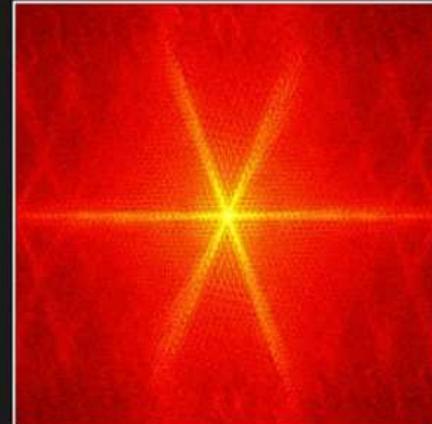


# High Pass Filtering

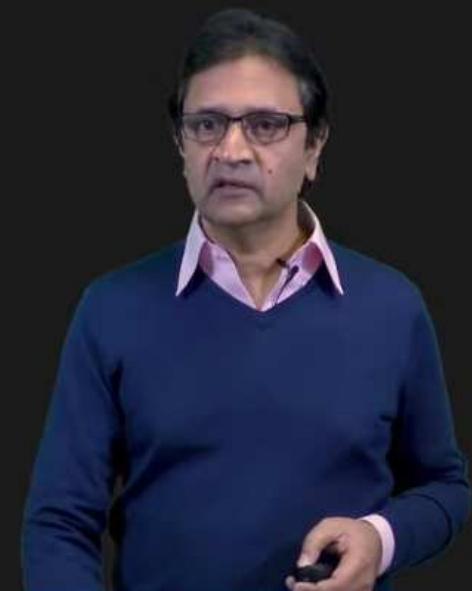
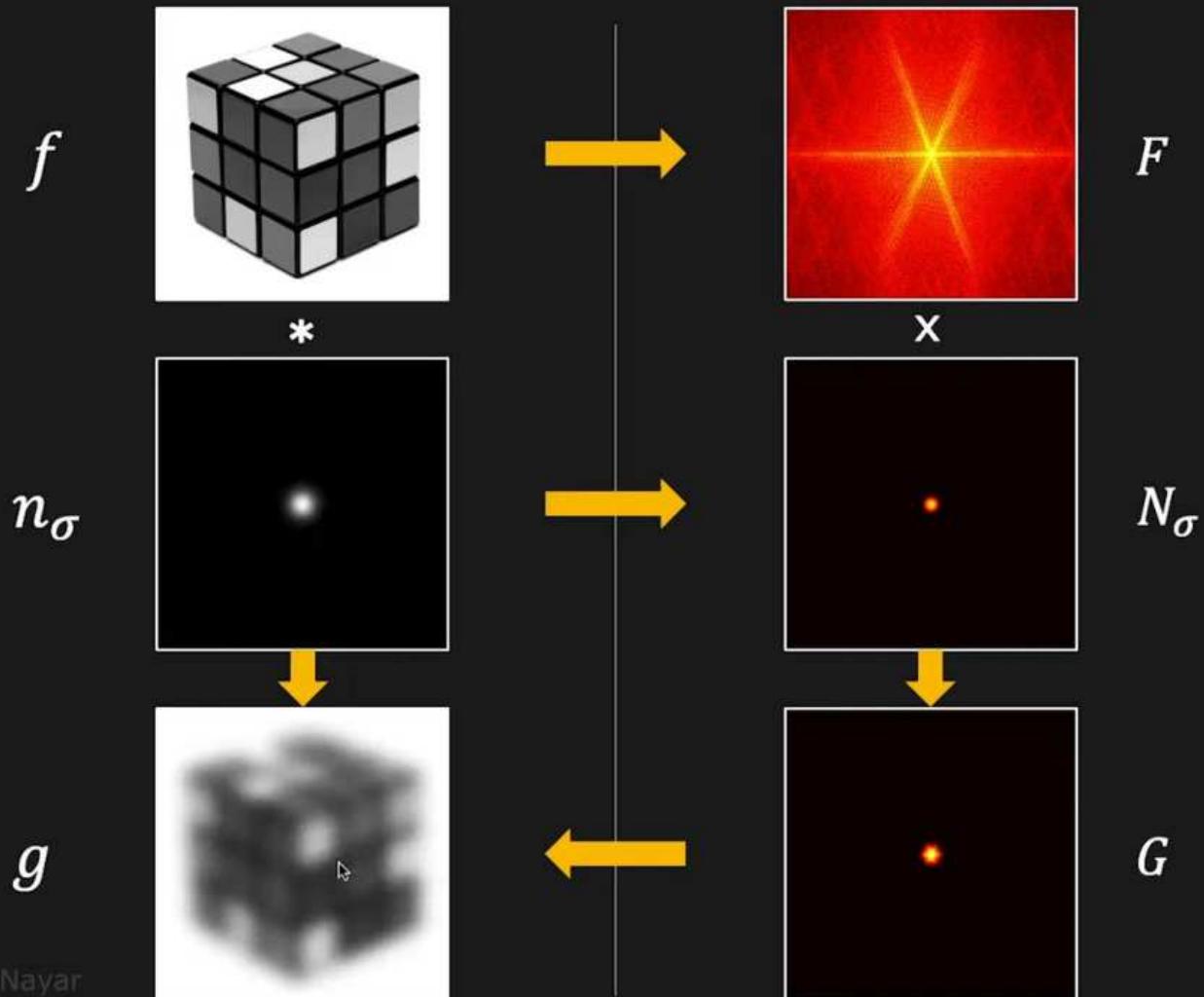
$f(m, n)$



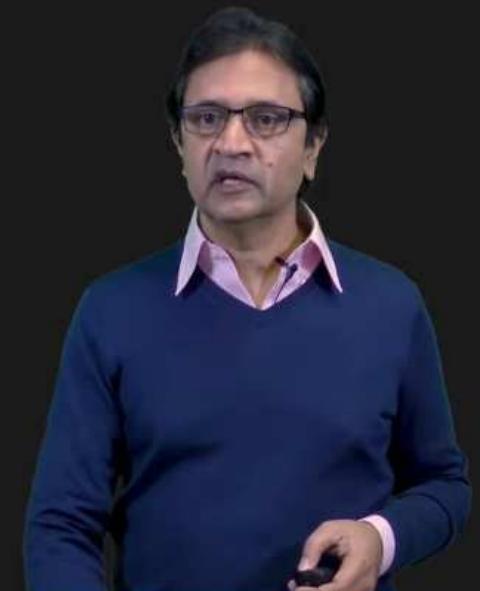
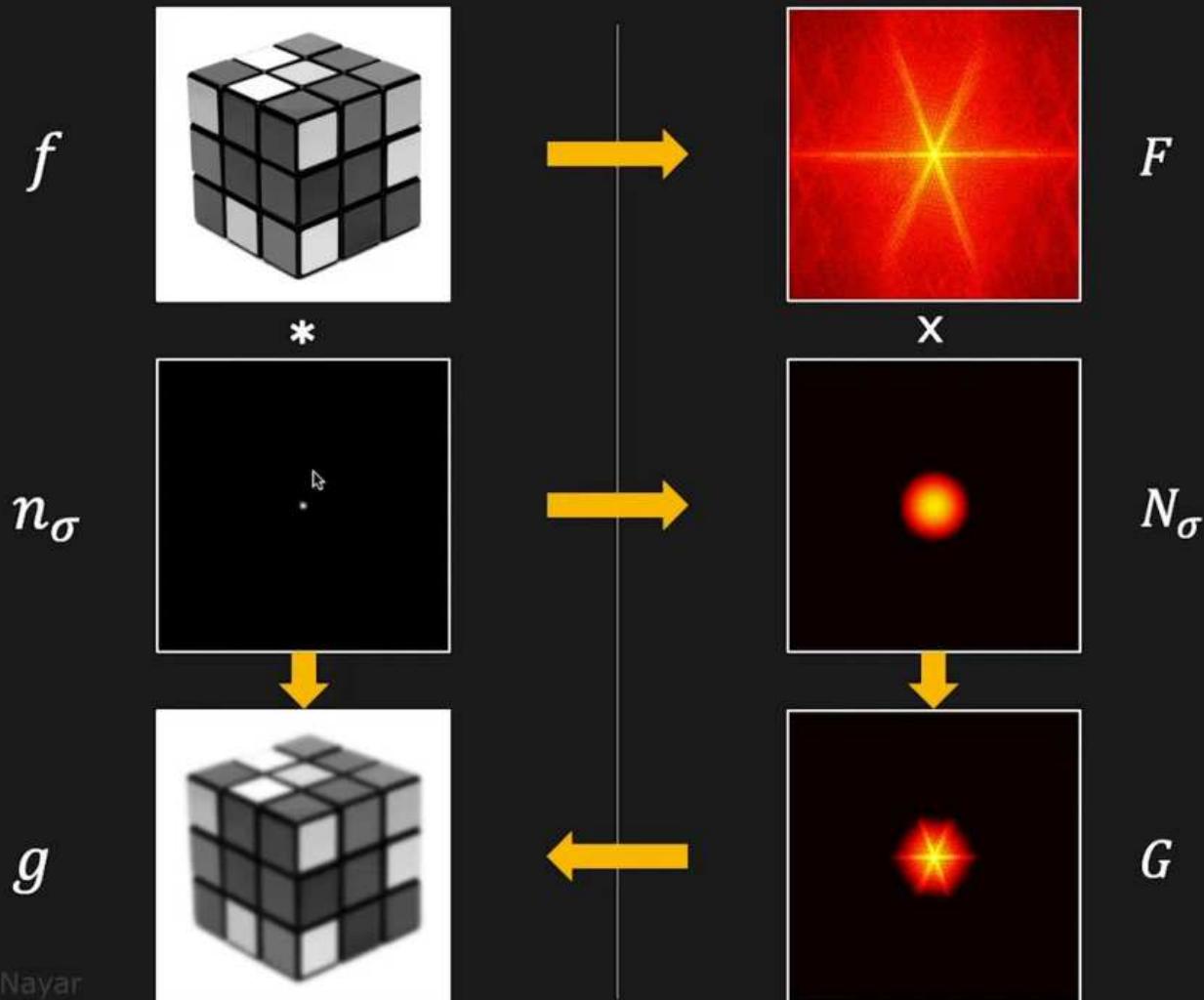
$\log(|F(p, q)|)$



# Gaussian Smoothing



# Gaussian Smoothing

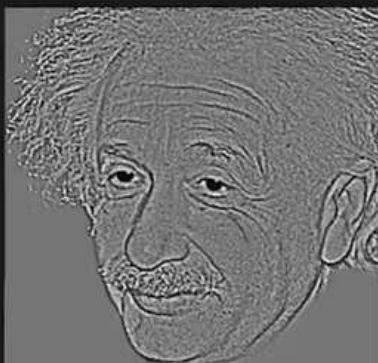


# Hybrid Images

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Low Freq Only



High Freq Only



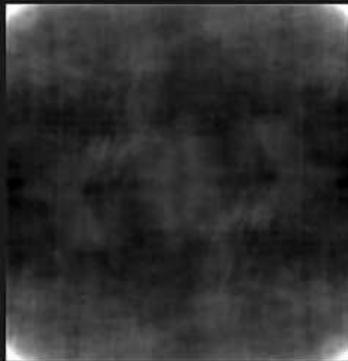
[Oliva 2]

# Importance of Phase

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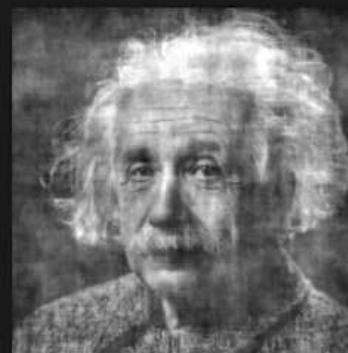
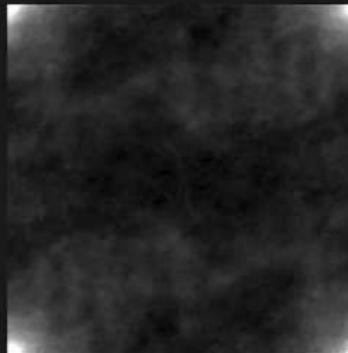
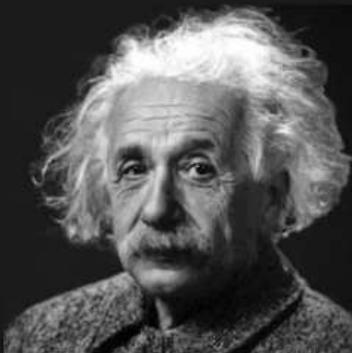
Original Image



Magnitude Preserved,  
Phase Set to Zero

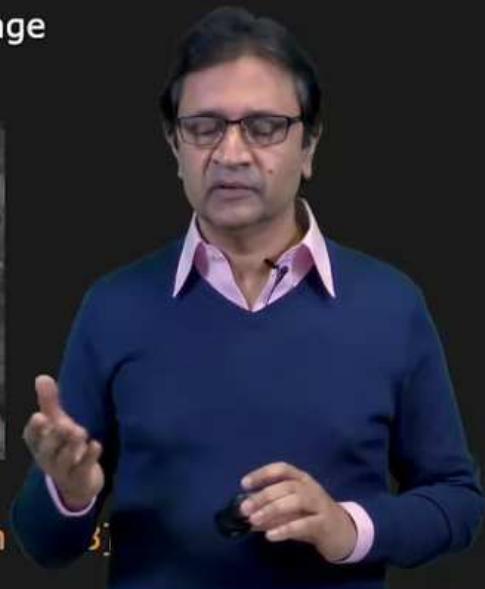


Phase Preserved,  
Magnitude Set to Average  
of Natural Images



[Oppenheim

5]

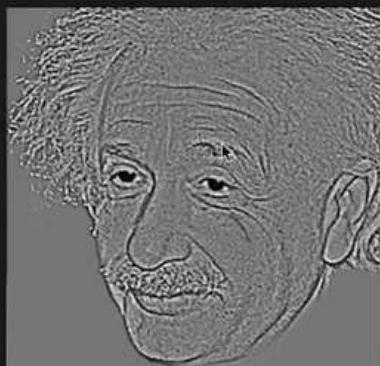


# Hybrid Images

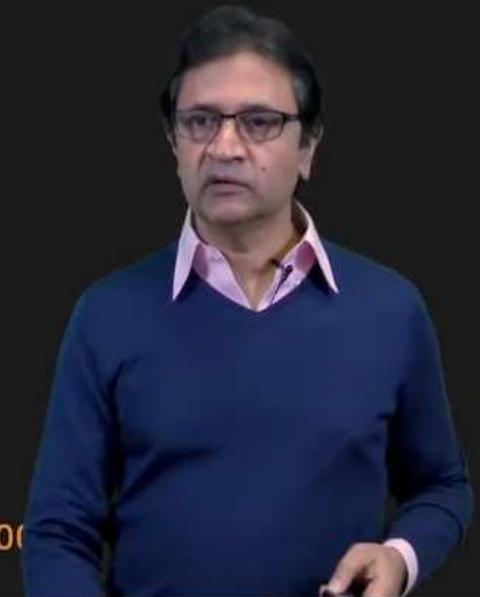
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Low Freq Only



High Freq Only



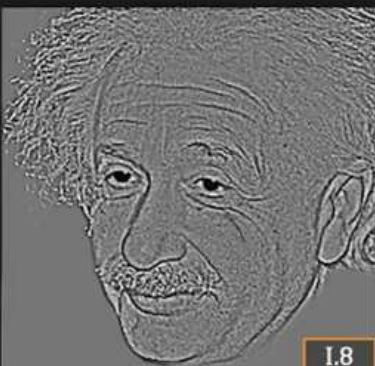
[Oliva 2006]

# Hybrid Images

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Low Freq Only



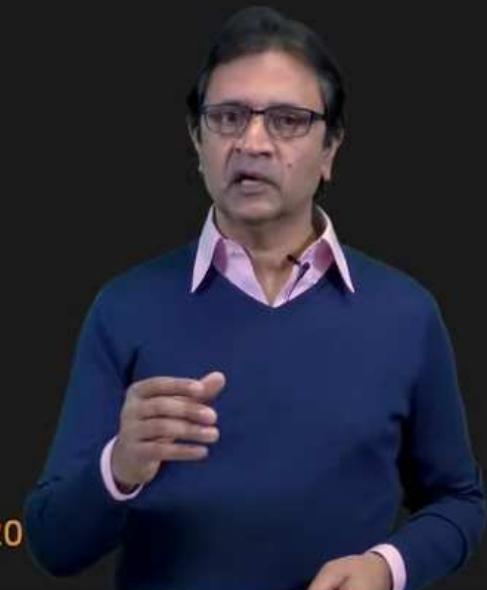
I.8

High Freq Only



Hybrid (Sum) Image

[Oliva 20]

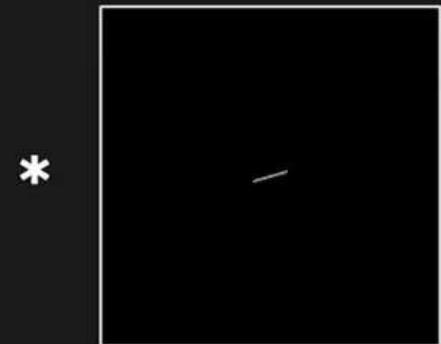


# Motion Blur

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Scene  $f(x, y)$



PSF  $h(x, y)$   
(Camera Shake)



Image  $g(x, y)$

$$f(x, y) * h(x, y) = g(x, y)$$

Given captured image  $g(x, y)$  and PSF  $h(x, y)$ ,  
can we estimate actual scene  $f(x, y)$ ?

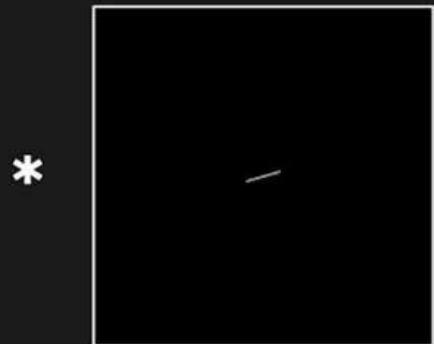


# Motion Deblur: Deconvolution

---



Scene  $f(x, y)$



PSF  $h(x, y)$   
(Camera Shake)



Image  $g(x, y)$

Let  $f'$  be the recovered scene.

$$f'(x, y) * h(x, y) = g(x, y)$$



# Motion Deblur: Deconvolution



Let  $f'$  be the recovered scene.

$$f'(x, y) * h(x, y) = g(x, y)$$

$$F'(u, v)H(u, v) = G(u, v)$$

$$F'(u, v) = \frac{G(u, v)}{H(u, v)} \rightarrow \boxed{\text{IFT}} \rightarrow f'(x, y)$$

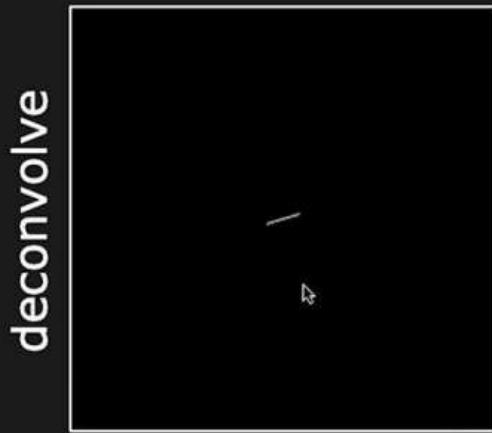


# Motion Deblur: Deconvolution

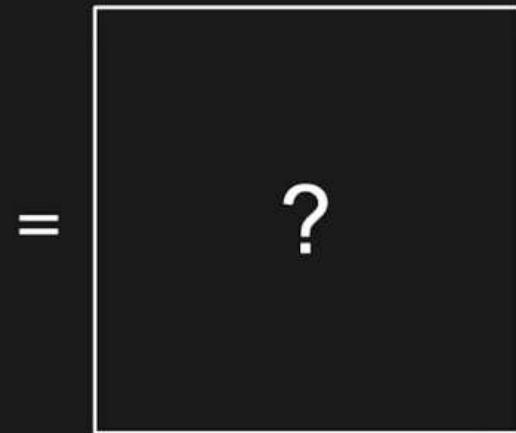
$$F'(u, v) = \frac{G(u, v)}{H(u, v)} \xrightarrow{\text{IFT}} f'(x, y)$$



Image  $g(x, y)$



PSF  $h(x, y)$



Recovered  $f'(x, y)$



# Motion Deblur: Deconvolution

$$F'(u, v) = \frac{G(u, v)}{H(u, v)} \xrightarrow{\text{IFT}} f'(x, y)$$

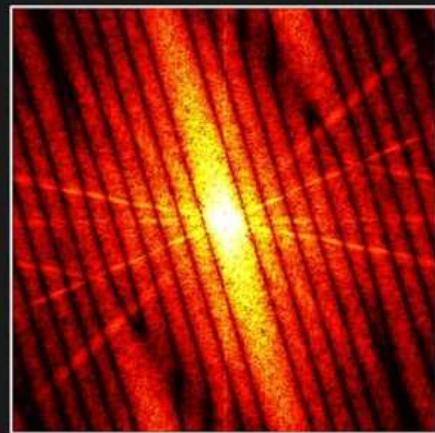
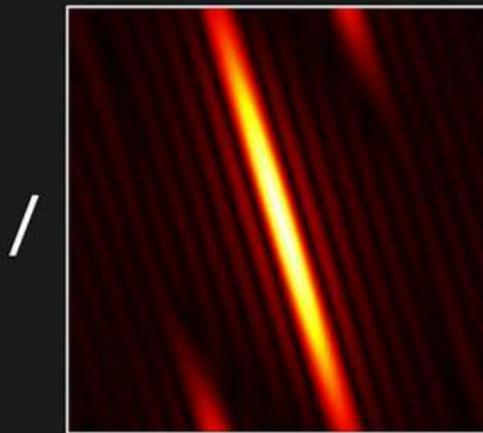
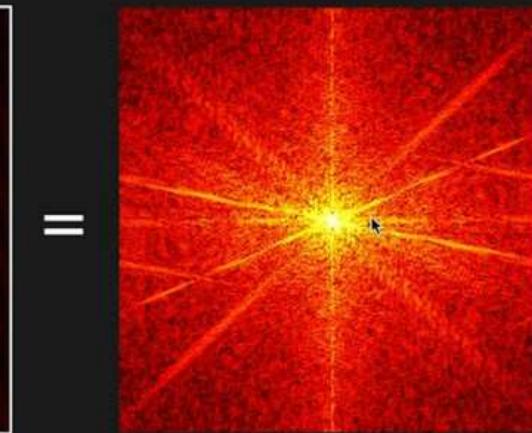


Image  $G(u, v)$

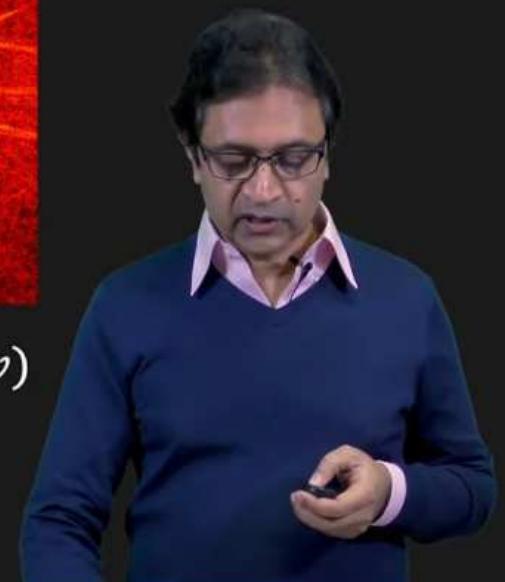


PSF  $H(u, v)$



Recovered  $F'(u, v)$

Step 1: Recover  $F'(u, v)$  in Fourier Domain

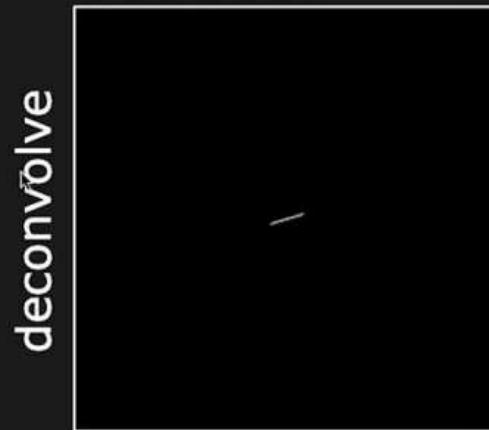


# Motion Deblur: Deconvolution

$$F'(u, v) = \frac{G(u, v)}{H(u, v)} \xrightarrow{\text{IFT}} f'(x, y)$$



Image  $g(x, y)$



PSF  $h(x, y)$



Recovered  $f'(x, y)$

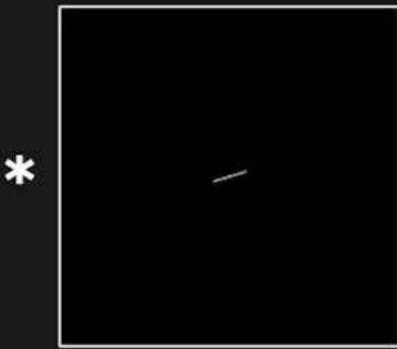
Step 2: Compute IFT of  $F'(u, v)$  to recover scene



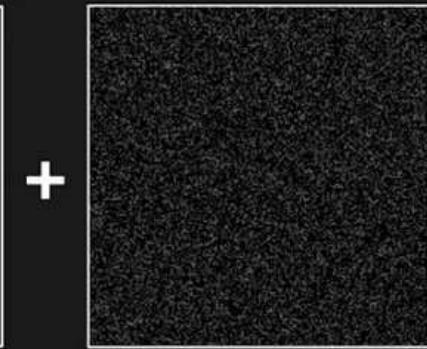
# Adding Noise to the Problem



Scene  $f(x, y)$



PSF  $h(x, y)$   
(Camera Shake)



Noise  $\eta(x, y)$



Image  $g(x, y)$

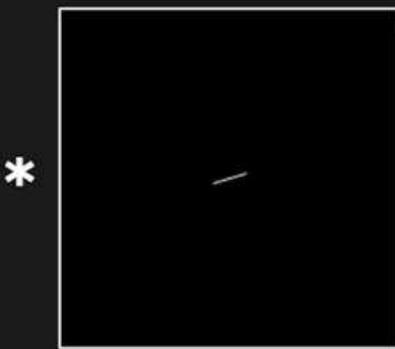
$$f(x, y) * h(x, y) + \eta(x, y) = g(x, y)$$



# Adding Noise to the Problem



Scene  $f(x, y)$



PSF  $h(x, y)$   
(Camera Shake)



Noise  $\eta(x, y)$



Image  $g(x, y)$

$$f(x, y) * h(x, y) + \eta(x, y) = g(x, y)$$

Can we afford to ignore noise?



# Motion Deblur: Deconvolution

If we ignore the noise ( $\eta(x, y)$ ):

$$\frac{G(u, v)}{H(u, v)} = F'(u, v) \xrightarrow{\text{IFT}} f'(x, y)$$

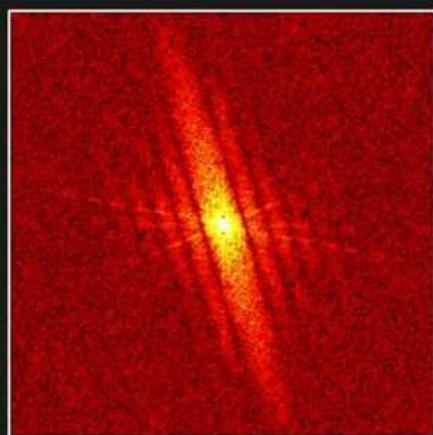
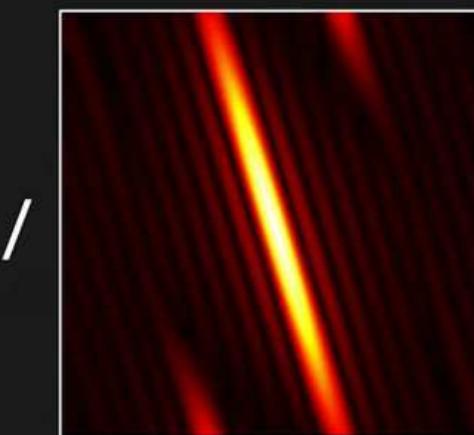
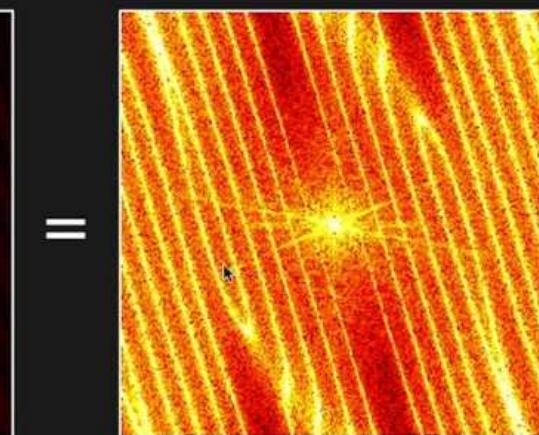


Image  $G(u, v)$



PSF  $H(u, v)$



Recovered  $F'(u, v)$

Higher frequencies in  $F'(u, v)$  are amplified



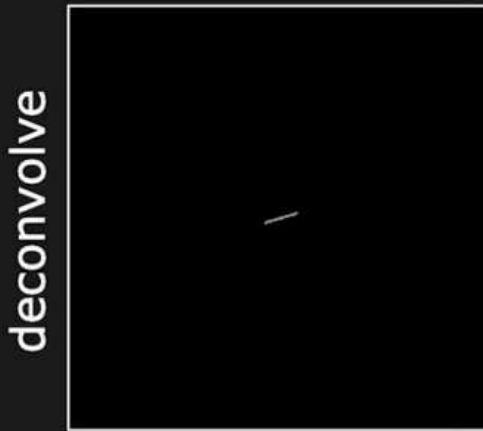
# Motion Deblur: Deconvolution

If we ignore the noise ( $\eta(x, y)$ ):

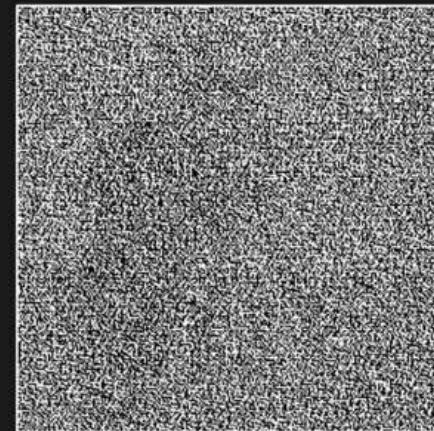
$$\frac{G(u, v)}{H(u, v)} = F'(u, v) \xrightarrow{\text{IFT}} f'(x, y)$$



Image  $g(x, y)$   
**(with noise)**

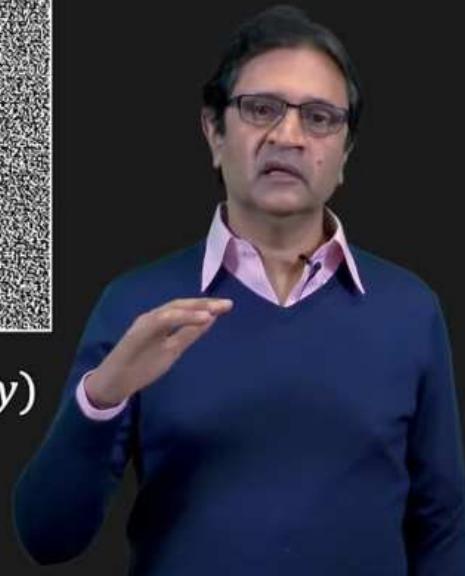


PSF  $h(x, y)$



Recovered  $f'(x, y)$

Noise is significantly amplified



# Deconvolution: Issues

---

$$\frac{G(u, v)}{H(u, v)} = F'(u, v) \longrightarrow \boxed{\text{IFT}} \longrightarrow f'(x, y)$$

1. Where  $H(u, v) = 0, F'(u, v) = \infty \rightarrow$  Not recoverable
2. Motion blur filter  $H(u, v)$  is a low pass filter.

For high frequencies  $(u, v)$ :

- Noise  $N(u, v)$  in  $G(u, v)$  is high
- Filter  $H(u, v) \approx 0$



# Deconvolution: Issues

---

$$\frac{G(u, v)}{H(u, v)} = F'(u, v) \longrightarrow \boxed{\text{IFT}} \longrightarrow f'(x, y)$$

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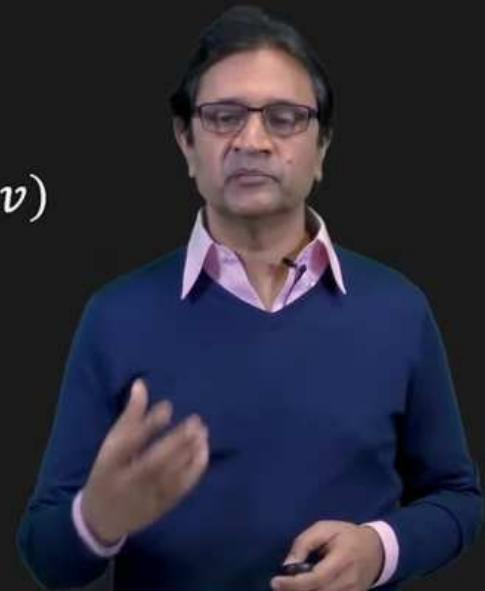
For high frequencies  $(u, v)$ :

- Noise  $N(u, v)$  in  $G(u, v)$  is high
- Filter  $H(u, v) \approx 0$



Noise in  $G(u, v)$   
is amplified

We need some kind of Noise Suppression.



# Noise Suppression: Weiner Deconvolution

---

$$F'(u, v) = \frac{G(u, v)}{H(u, v)} \left[ \frac{1}{1 + \frac{NSR(u, v)}{|H(u, v)|^2}} \right]$$

Noise-to-Signal Ratio,  $NSR(u, v)$ :

$$NSR(u, v) = \frac{\text{Power of Noise at } (u, v)}{\text{Power of Signal (Scene) at } (u, v)} = \frac{|N(u, v)|^2}{|F(u, v)|^2}$$



# Noise Suppression: Weiner Deconvolution

---

$$F'(u, v) = \frac{G(u, v)}{H(u, v)} \left[ \frac{1}{1 + \frac{NSR(u, v)}{|H(u, v)|^2}} \right]$$

Where:

Weiner Filter  $\stackrel{\text{def}}{=}$

$$W(u, v) = \frac{1}{H(u, v)} \left[ \frac{1}{1 + \frac{NSR(u, v)}{|H(u, v)|^2}} \right]$$

Noise-to-Signal Ratio,  $NSR(u, v)$ :

$$NSR(u, v) = \frac{\text{Power of Noise at } (u, v)}{\text{Power of Signal (Scene) at } (u, v)} = \frac{|N(u, v)|^2}{|F(u, v)|^2}$$



# Noise Suppression: Weiner Deconvolution

---

$$F'(u, v) = \frac{G(u, v)}{H(u, v)} \left[ \frac{1}{1 + \frac{NSR(u, v)}{|H(u, v)|^2}} \right]$$

- Determining  $NSR$  requires us to have prior knowledge of the noise “pattern” and the scene (or of a similar scene).

$$NSR(u, v) = \frac{|N(u, v)|^2}{|F(u, v)|^2}$$

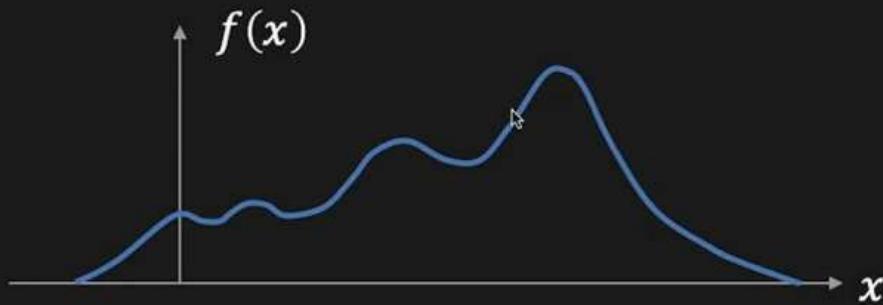
- Often  $NSR$  is set to a single suitable constant  $\lambda$ .

$$NSR(u, v) = \lambda$$

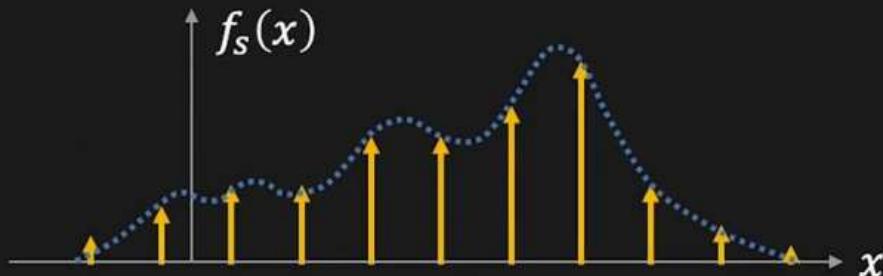


# From Continuous to Digital Image

Continuous Signal:

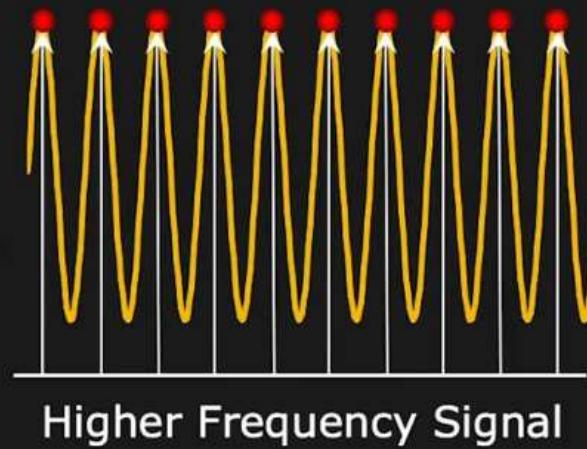
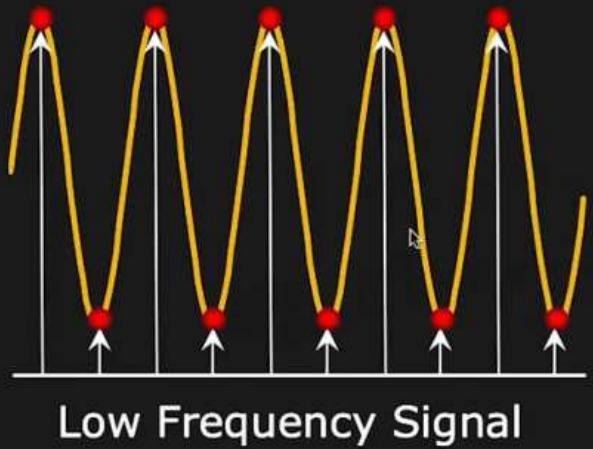


Digital Signal:

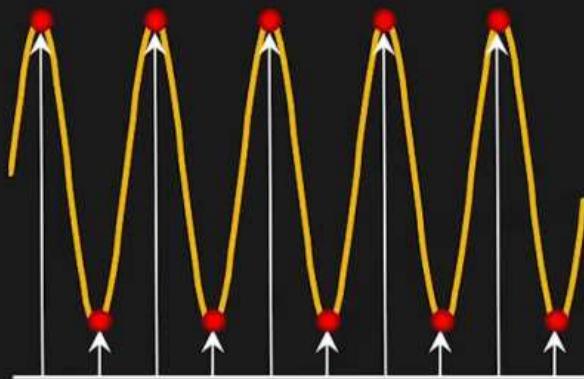


How “dense” should the samples be?

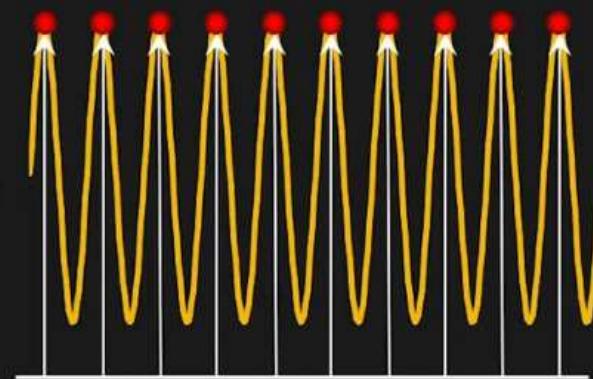
# Sampling Problem



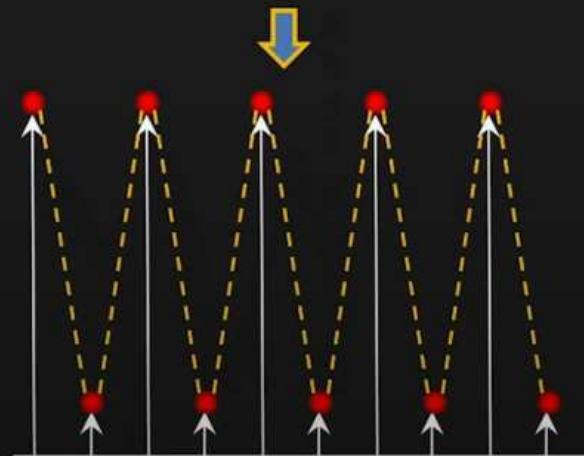
# Sampling Problem



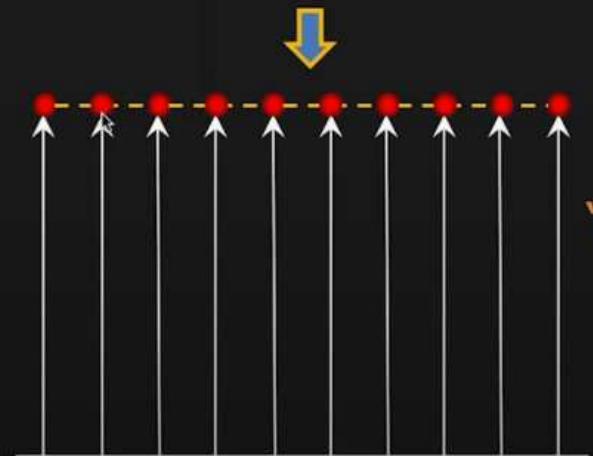
Low Frequency Signal



Higher Frequency Signal



Reconstructed Signal



Reconstructed Signal

"Aliasing"



# Sampling Problem



"Well sampled" image

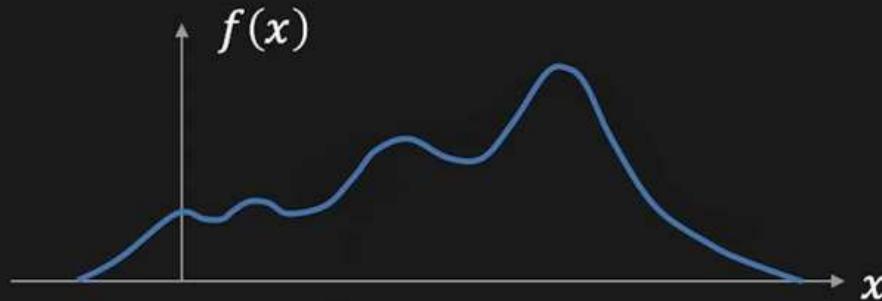


"Under sampled" image  
(visible **aliasing** artifacts)



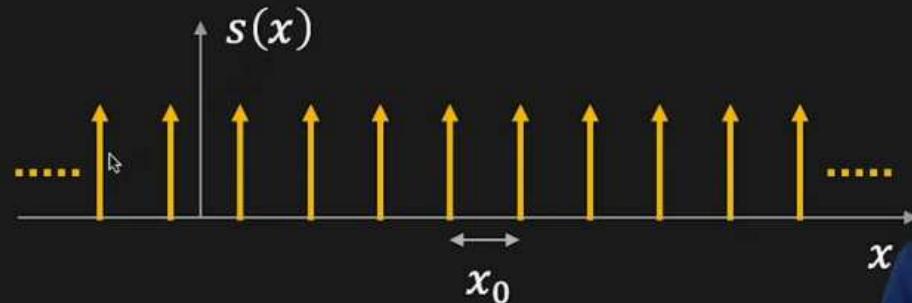
# Sampling Theory

Continuous Signal:



Shah Function (Impulse Train):

$$s(x) = \sum_{n=-\infty}^{\infty} \delta(x - nx_0)$$



Sampled Function:

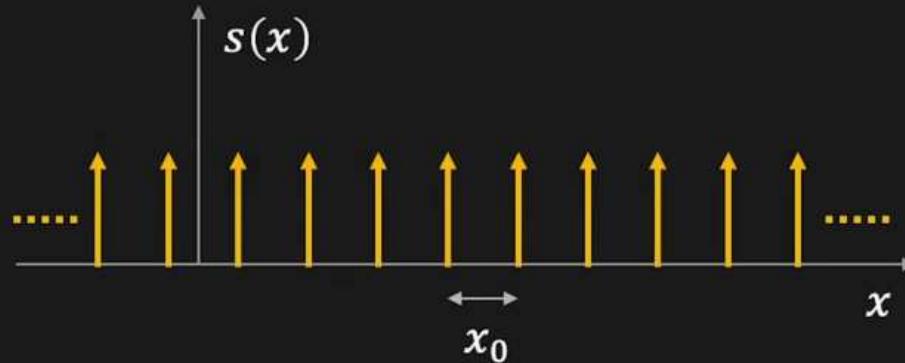
$$f_s(x) = f(x)s(x)$$



# Shah Function (Impulse Train)

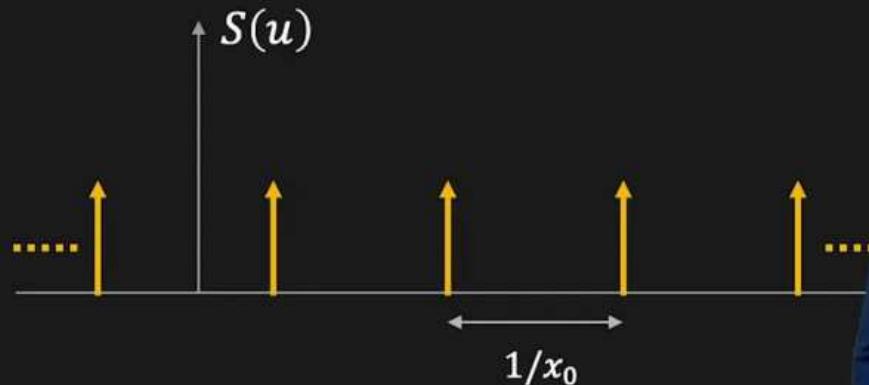
Shah Function (Spatial Domain):

$$s(x) = \sum_{n=-\infty}^{\infty} \delta(x - nx_0)$$



Shah Function (Fourier Domain):

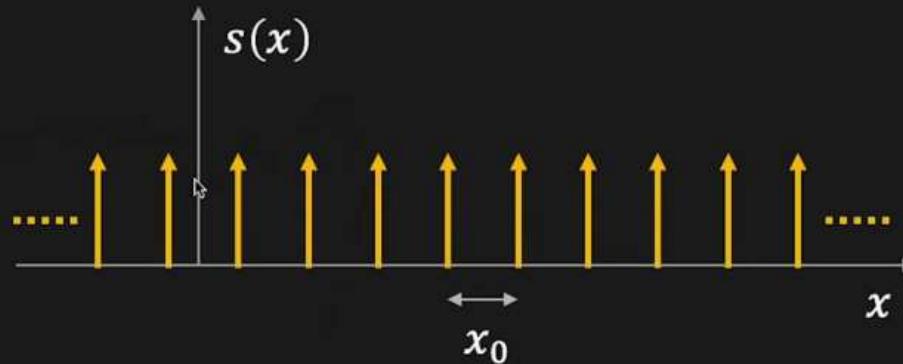
$$S(u) = \frac{1}{x_0} \sum_{n=-\infty}^{\infty} \delta\left(u - \frac{n}{x_0}\right)$$



# Shah Function (Impulse Train)

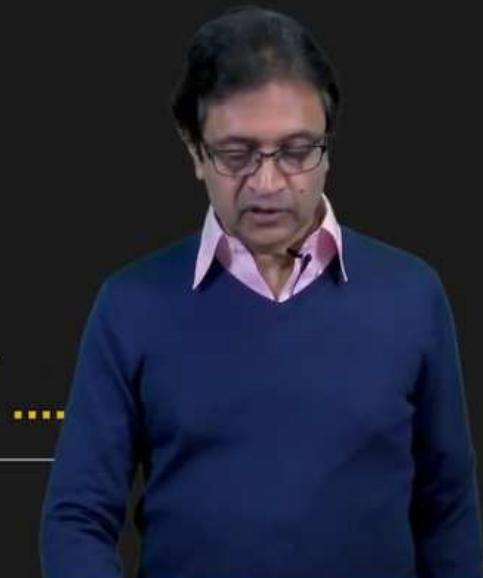
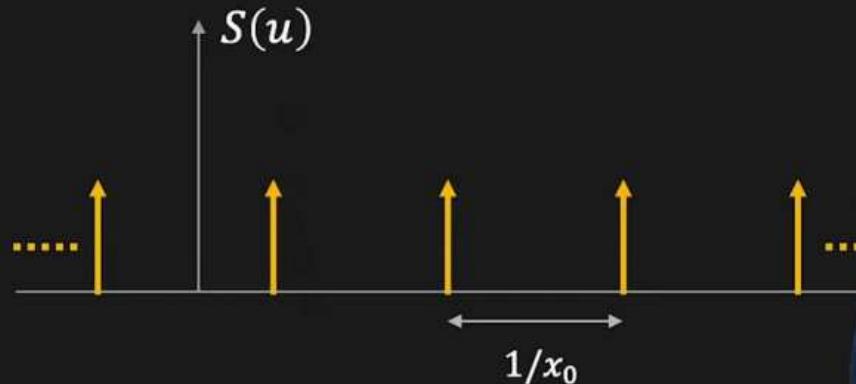
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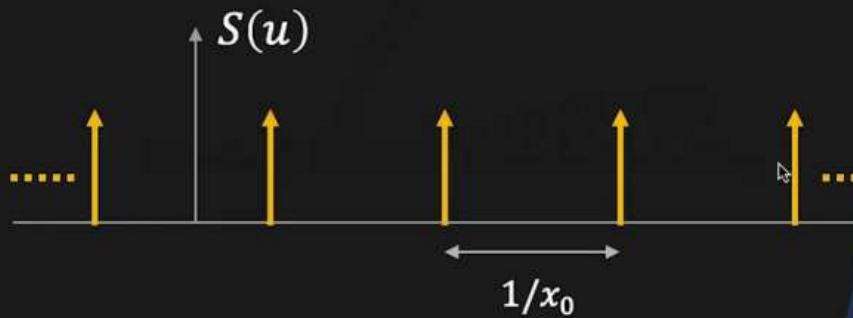
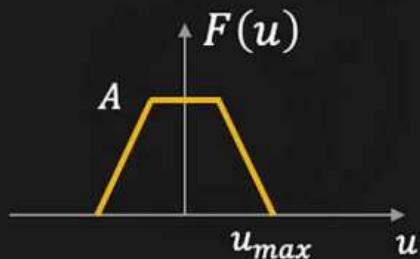
# Fourier Analysis of Sampled Signal

Sampled Signal:

$$f_s(x) = f(x)s(x) = f(x) \sum \delta(x - nx_0)$$

$$F_s(u) = F(u) * S(u) = F(u) * \frac{1}{x_0} \sum \delta(u - n/x_0)$$

For example:



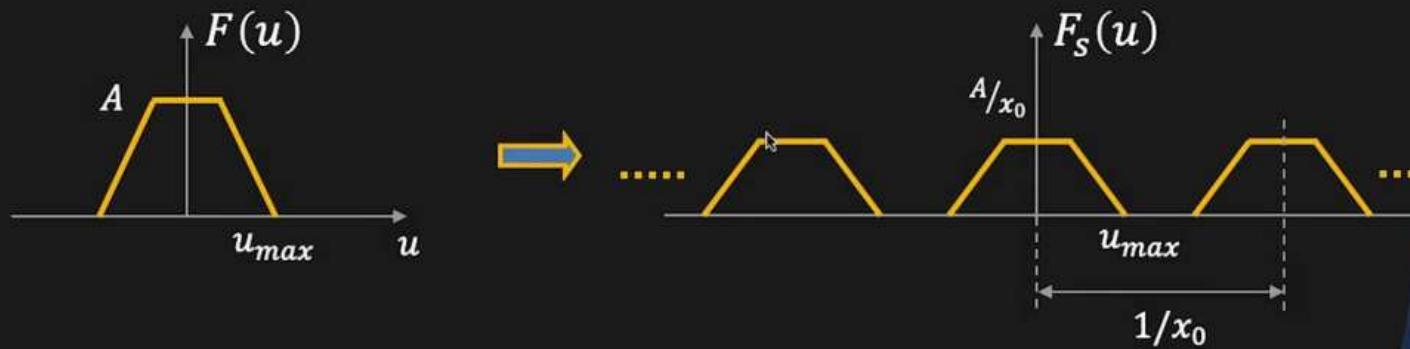
# Fourier Analysis of Sampled Signal

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If  $u_{max} \leq \frac{1}{2x_0}$



# Aliasing

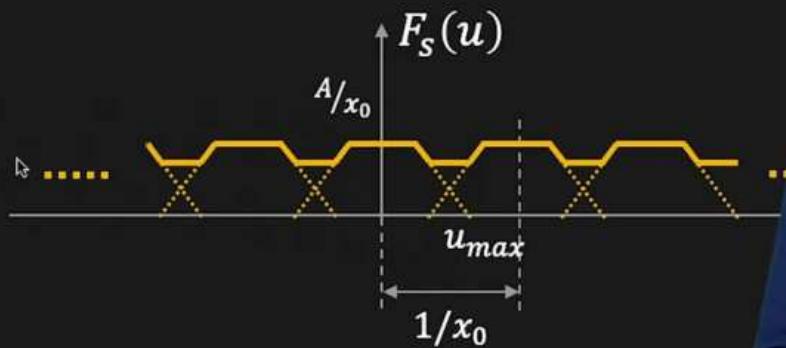
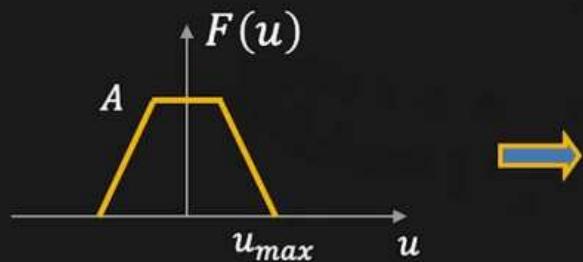
Sampled Signal:

$$f_s(x) = f(x)s(x) = f(x) \sum \delta(x - nx_0)$$

$$F_s(u) = F(u) * S(u) = F(u) * \frac{1}{x_0} \sum \delta(u - n/x_0)$$

If  $u_{max} > \frac{1}{2x_0}$

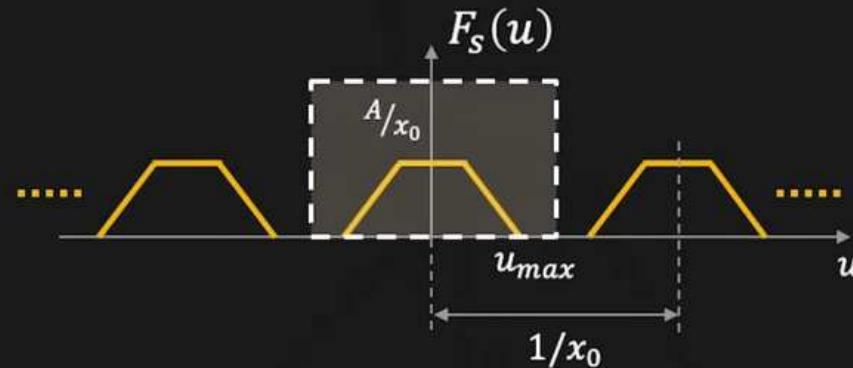
Aliasing



# Nyquist Theorem

Can we recover  $f(x)$  from  $f_s(x)$ ? In other words,  
can we recover  $F(u)$  from  $F_s(u)$ ?

Only if  $u_{max} \leq \frac{1}{2x_0}$  (**Nyquist Frequency**)



$$C(u) = \begin{cases} x_0, & |u| < 1/2x_0 \\ 0, & \text{Otherwise} \end{cases}$$



# Nyquist Theorem

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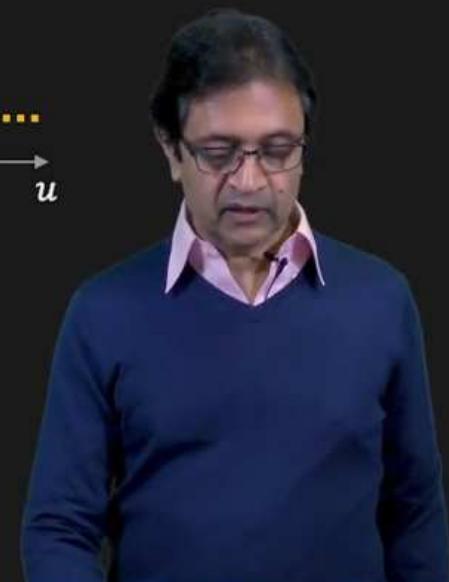
Only if  $u_{max} \leq \frac{1}{2x_0}$  (**Nyquist Frequency**)



$$F(u) = F_s(u)C(u)$$

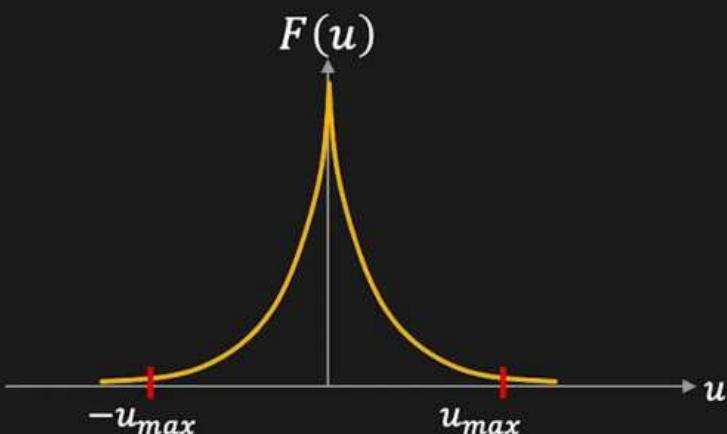
$$f(x) = IFT(F(u))$$

$$C(u) = \begin{cases} x_0, & |u| < 1/2x_0 \\ 0, & \text{Otherwise} \end{cases}$$



# Aliasing in Digital Imaging

Aliasing occurs when imaging a scene (signal) that has frequencies above the image sensor's Nyquist Frequency



Typical Power Spectrum  
of Natural Scenes

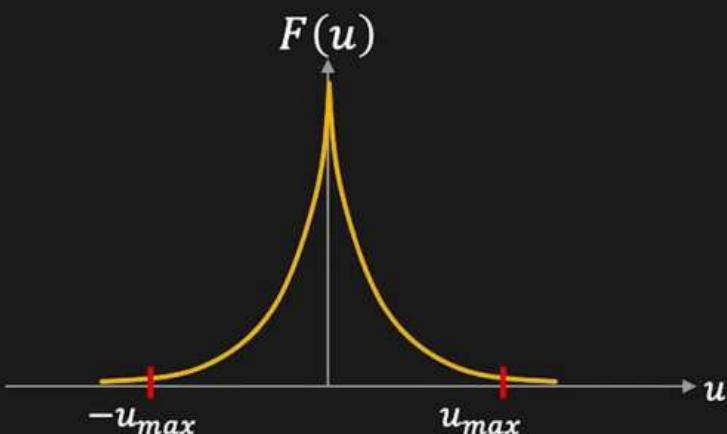


Aliasing artifacts usually occur in the form of Moiré patterns



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Typical Power Spectrum  
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Aliasing artifacts usually occur in  
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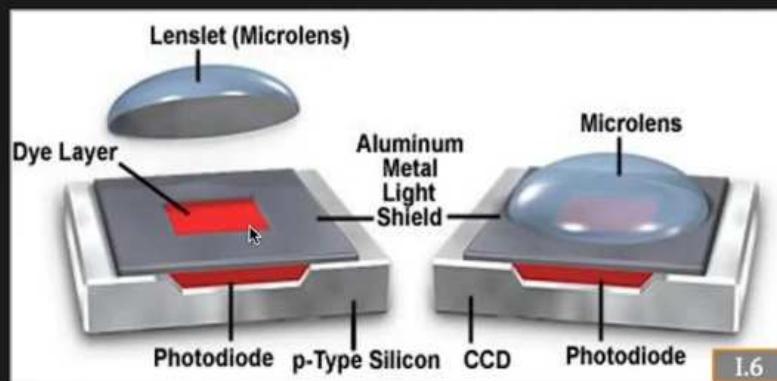
How do sensors combat aliasing?

# Minimizing the Effects of Aliasing

**Band Limit:** Clip the signal above the Nyquist frequency.

Effectively, “blur” the scene before sampling.

Sensors use two strategies.



Pixels are area-samplers  
(box-averaging filter)

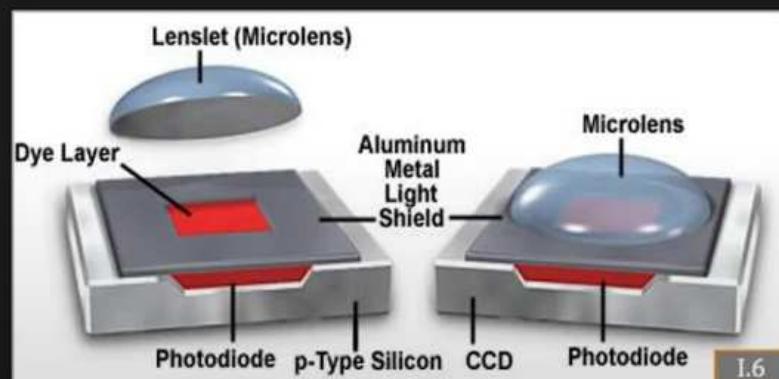


# Minimizing the Effects of Aliasing

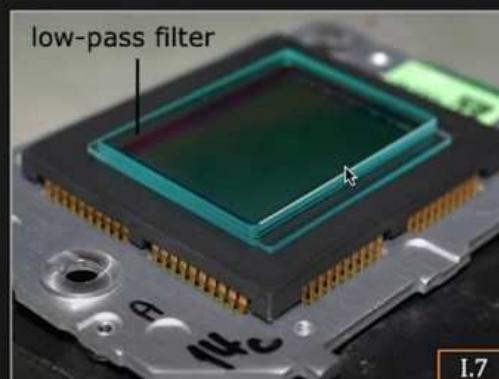
**Band Limit:** Clip the signal above the Nyquist frequency.

Effectively, “blur” the scene before sampling.

Sensors use two strategies.



Pixels are area-samplers  
(box-averaging filter)



Use optical low-pass filter  
(anti-aliasing filter)

