

Isospin splitting of nucleon effective mass in light of the decoded symmetry energy

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Outline

I. Motivation

II. ImQMD model (-23 version)

III. Effective mass splitting from HICs.

IV.Summary and Outlook

I. Motivation

Definition

The nonrelativistic nucleon effective mass is defined as:

$$\frac{m_q^*}{m} = \left[1 + \frac{m}{p} \frac{\partial V_q}{\partial p}\right]^{-1}, \qquad q = n, p.$$

The effective mass splitting:

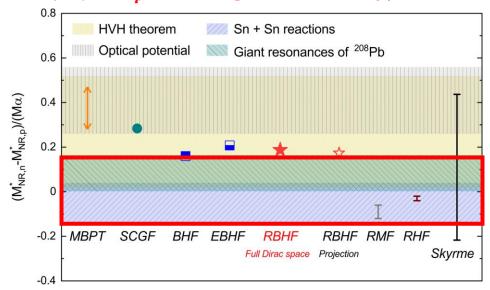
$$\Delta m_{np}^* = \frac{m_n^* - m_p^*}{m}.$$

- ➤ The study of effective mass splitting has important physical significance for us to understand various physical phenomena in nuclear physics and astrophysics.
- symmetry energy constraints
- properties of neutron stars
- properties of rare isotopes
- •

Current status of constraints on effective mass splitting

1) Microscopic model: MBPT, SCGF, BHF, EBHF, HVH, RBHF, $(m_n^* > m_p^*)$ RMF, RHF $(m_n^* < m_p^*)$...

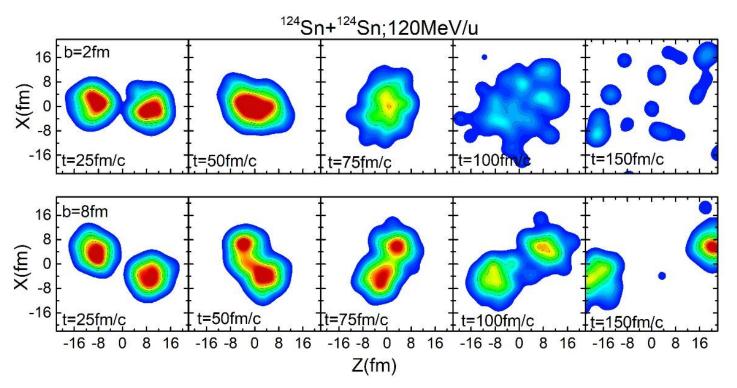
- 2) Optical potential $(m_n^* > m_p^*)$
- 3) HICs $(m_n^* < m_p^*)$, with larger uncertianty)



- The sign and magnitude of the effective mass splitting are still unclear.
- What is behind the door of the constraints of Δm_{np}^* from HICs?

Sibo Wang, Hui Tong, Qiang Zhao, Chencan Wang, Peter Ring, and Jie Meng. Phys. Rev. C. 108, L031303 (2023).

I. Motivation



The effective mass splitting constrained by HICs is not at a certain density or momentum.

ImQMD-23(Improved quantum molecular dynamics model-23):

- 1. The standard Skyrme potential energy density can be used: $u_{sky} = u_{loc} + u_{md}$, connect the nuclear reaction and structure based on the same footing.
- 2. Extended Skyrme MDI $(g(\mathbf{p} \mathbf{p}') = \delta(\mathbf{r} \mathbf{r}') \sum_{l=0}^{N} b_l (\mathbf{p} \mathbf{p}')^{2l})$.
- 3. Accurate calculations on the nonlinear density-dependent interaction
- 4. New Pauli blocking with antisymmetry wave function effects.

ImQMD-23 team group:

Yingxun Zhang, Junping Yang, Xiang Chen, Cui Ying, Zhuxia Li, Kai Zhao

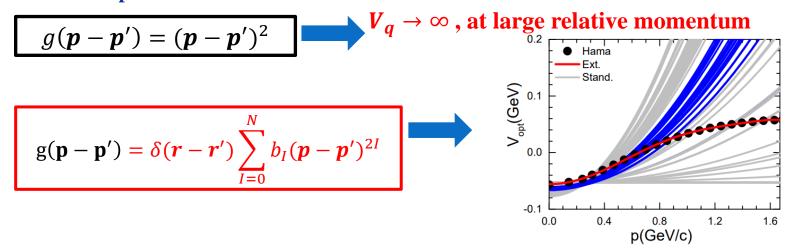
The Skyrme potential energy density is used: $u_{sky} = u_{loc} + u_{md}$

$$u_{loc} = \frac{\alpha \rho^{2}}{2 \rho_{0}} + \frac{\beta}{\eta + 1} \frac{\rho^{\eta + 1}}{\rho_{0}^{\eta}} + \frac{g_{sur}}{2\rho_{0}} (\nabla \rho)^{2} \qquad u_{md} = C_{0} \sum_{ij} \int d^{3}p d^{3}p' f_{i}(\mathbf{r}, \mathbf{p}) f_{j}(\mathbf{r}, \mathbf{p}') g(\mathbf{p} - \mathbf{p}')$$

$$+ \frac{g_{sur,iso}}{\rho_{0}} [\nabla (\rho_{n} - \rho_{p})]^{2} \qquad + D_{0} \sum_{ij \in n} \int d^{3}p d^{3}p' f_{i}(\mathbf{r}, \mathbf{p}) f_{j}(\mathbf{r}, \mathbf{p}') g(\mathbf{p} - \mathbf{p}')$$

$$+ A_{sym} \frac{\rho^{2}}{\rho_{0}} \delta^{2} + B_{sym} \frac{\rho^{\eta + 1}}{\rho_{0}^{\eta}} \delta^{2} \qquad + D_{0} \sum_{ij \in p} \int d^{3}p d^{3}p' f_{i}(\mathbf{r}, \mathbf{p}) f_{j}(\mathbf{r}, \mathbf{p}') g(\mathbf{p} - \mathbf{p}')$$

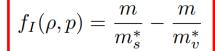
Momentum-dependent interaction:



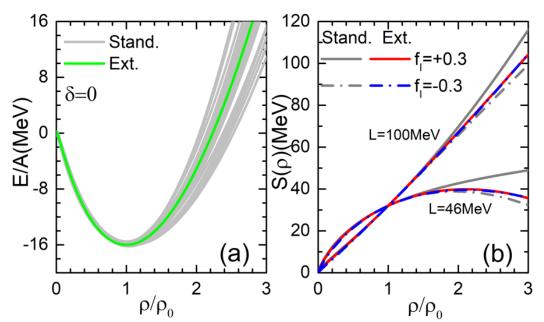
Equation of state and the density dependence of the symmetry energy:

TABLE I: Four sets of nuclear matter parameters used in this work.

$\rho_0(\mathrm{fm}^{-3})$	$E_0(MeV)$	$K_0({ m MeV})$	$S_0({ m MeV})$	L(MeV)	m_s^*/m	f_I
0.16	-16	230	32	$46,\!100$	0.77	0.3
0.16	-16	230	32	$46,\!100$	0.77	-0.3

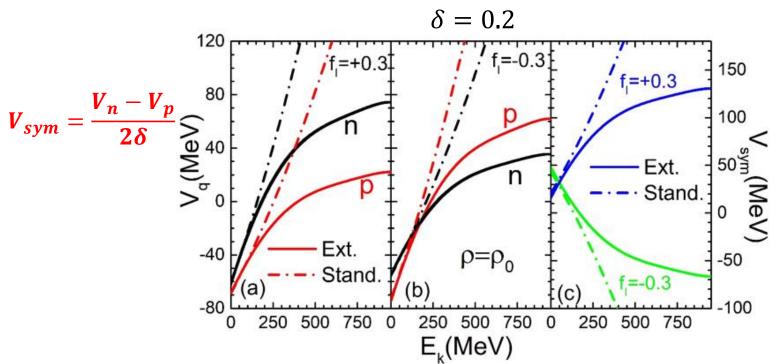


$$(m_n^* < m_p^*)$$
$$(m_n^* > m_p^*)$$



Junping Yang, Xiang Chen, Ying Cui, Yangyang Liu, Zhuxia Li, and Yingxun Zhang†, Phys. Rev. C 109, 054624(2024).

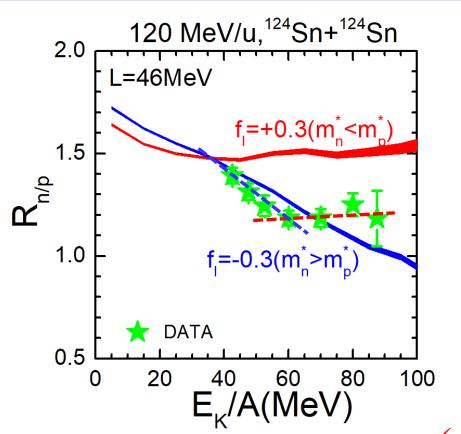
Single-particle potential and symmetric potential:



Junping Yang, Xiang Chen, Ying Cui, Yangyang Liu, Zhuxia Li, and Yingxun Zhang†, Phys. Rev. C 109, 054624(2024).

✓ Compared with the standard MDI, the symmetric potential obtained by the extended MDI flattens out with the increase of kinetic energy.

III. Effective mass splitting from HICs.



$$R_{n/p} = \frac{\mathrm{Y(n)}}{\mathrm{Y(p)}} \propto \exp\left(\frac{\mu_n - \mu_p}{T}\right) = \exp\left(\frac{2V_{asy}\delta}{T}\right)$$
 (1)

$$\Delta m_{np}^* \approx -(\frac{m^*}{m})^2 4m\delta \frac{\partial V_{asy}}{\partial p^2}$$
 (2)

$$R_{n/p} \propto \exp\left[\frac{2(V_{asy}^0 + \frac{\partial V_{asy}}{\partial p^2}p^2 + \cdots)\delta}{T}\right]$$
 (3)

$$\approx \exp\left[\frac{2V_{asy}^0\delta}{T}\right] \exp\left[-\frac{(\frac{m}{m^*})^2\Delta m_{np}^*}{T}E_k'\right].$$

 $E'_k = \lambda E_k / A(E'_k)$ the relative kinetic energy between colliding nucleon pairs)

$$S_{n/p} = \frac{\partial ln R_{n/p}}{\partial E_{k}/A} = -\frac{\lambda}{T} (\frac{m}{m^*})^2 \Delta m_{np}^*$$

 $\checkmark S_{n/p}$ is directly related to the Δm_{np}^* !!

III. Effective mass splitting from HICs.

Junping Yang, Xiang Chen, Ying Cui, Yangyang Liu, Zhuxia Li, and Yingxun Zhang.

Phys. Rev. C, 109:054624, May 2024. 120 MeV/u L=46MeV 0.01 0.00 -0.01¹¹²Sn+¹¹²Sn ¹²⁴Sn+¹²⁴Sn (b) -0.02 L=100MeV 0.01 0.00 -0.01 112Sn+112Sn 124Sn+124Sn -0.02 35-55 55-95 35-55 55-95

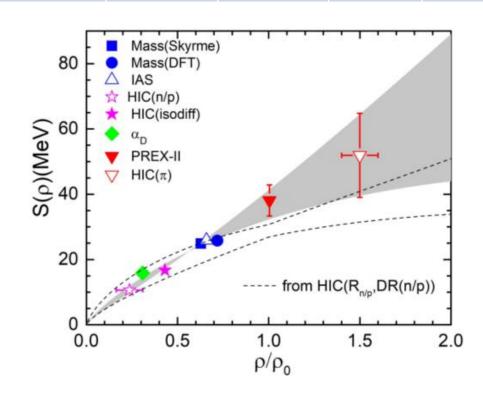
✓ The calculations tend to favor different effective mass splitting at different kinetic energy regions.

 $\Delta E_{L}/A(MeV)$

 $\sqrt{m_n^*} > m_p^* (35 \sim 55 \text{MeV}); m_n^* < m_p^* (55 \sim 95 \text{MeV});$

Determination of multi-dimensional parameter space:

$\rho_0(fm^{-3})$	$E_0(MeV)$	$K_0(MeV)$	m_s^*/m	f_I	$S_0(MeV)$	L(MeV)
0.16	-16	230	[0.6, 1.0]	[-0.5, 0.4]	[32,41]	[54.5,134]



HIC(n/p): W. Lynch, M.B. Tsang, arXiv:1805.10757.

HIC(isodiff): P. Morfouace, C. Tsang, Y. Zhang, W.

Lynch, et al., Phys. Lett. B 799, 135045 (2019)

 α_D (electric dipole polarization): A. Tamii, et al., Phys.

Rev. Lett. 107 (2011) 062502

MASS(Skyrme): B. A. Brown, Phys. Rev. Lett. 111, 232502 (2013).

MASS(DFT): M. Kortelainen, J. McDonnell, W.

Nazarewicz, et al., Phys. Rev. C. 85, 024304 (2012).

IAS(Isobaric analog state): P. Danielewicz, P. Singh, and

J. Lee, Nucl. Phys. A 958, 147 (2017).

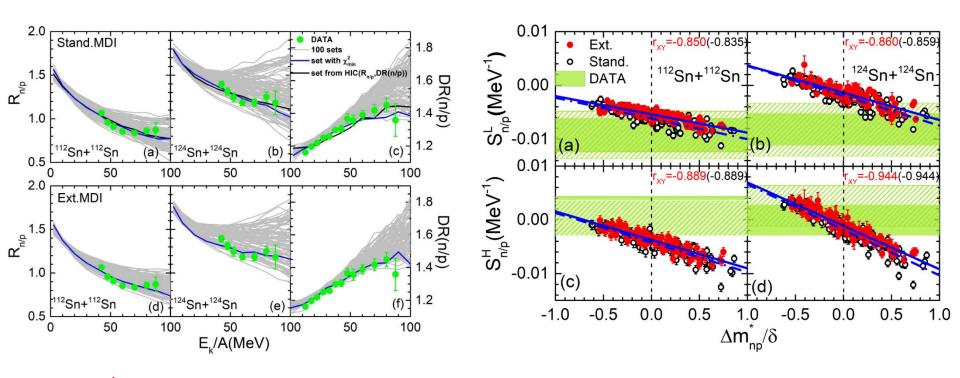
PREX-II: D. Adhikari, H. Albataineh, D. Androic, K.

Aniol, et al., Phys. Rev. Lett. 126, 172502 (2021).

 $HIC(\pi)$: J. Estee, W. G. Lynch, C. Y. Tsang, et al., Phys.

Rev. Lett. 126, 162701 (2021).

$R_{n/p}$, $DR_{n/p}$ and $S_{n/p}$ calculated by $100\ parameter\ sets$:



- $\sqrt{m_n^*} > m_p^*$ (low energy region); $m_n^* < m_p^*$ (high energy region);
- \checkmark $S_{n/p}$ is strongly correlated to the neutron-proton effective mass splitting!

IV. Summary and Outlook

- 1) The slope of lnRn/p as a function of Ek/A, i.e., $S_{n/p}$ is directly related to the Δm_{np}^* ($r_{XY}=0.944$ for neutron-rich system)
- 2) mn*>mp*(low energy region); mn*<mp*(high energy region).

Future plan:

- 1) The form of symmetry potential could first decrease and then increase with momentum. (the finite range Gogny force D1S and D250)
- 2) The effects of high-momentum tail in the initial nucleus could also influence the energy spectra of the emitted nucleons.



Thank you for your attention!

$R_{n/p}$ and $DR_{n/p}$ calculated by 100 parameter sets:

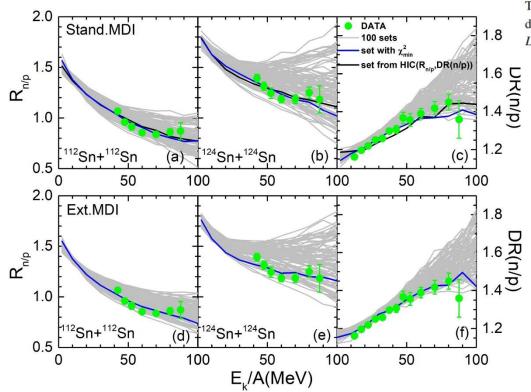
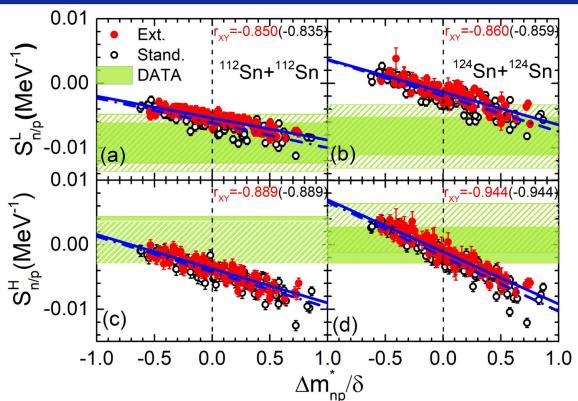


Table 2: The constraints on S_0 , L, m_s^*/m and Δm_{np}^* obtained by simultaneously describing $R_{n/p}^{112}$, $R_{n/p}^{124}$, and $R_{n/p}^{124}$, and $R_{n/p}^{124}$ are in dimensionless, $R_{n/p}^{124}$, and $R_{n/p}^{124}$ are in MeV.

Para.	Ext.	Stand.	Ref.[28]
S_0	34.9±2.2	35.5±2.2	28.8±1.9
L	105.1±22.5	102.0 ± 22.3	49.6±13.7
m_s^*/m	0.80 ± 0.11	0.76 ± 0.12	0.67 ± 0.03
Δm_{np}^*	$(0.31\pm0.46)\delta$	$(0.18\pm0.44)\delta$	$(-0.05\pm0.09)\delta$

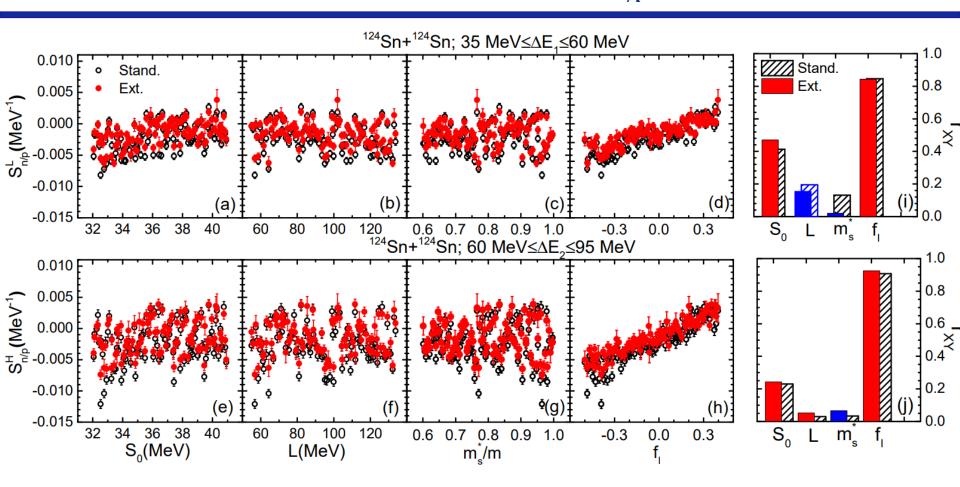
$$S_0 \approx 35 MeV$$
 $L \approx 104 MeV$
 $m_s^*/m \approx 0.78$
 $\Delta m_{np}^* = 0.18 - 0.31$

III. Effective mass splitting from HICs.



- \checkmark m_n^{*} > m_p^{*} (low energy region); m_n^{*} < m_p^{*} (high energy region);
- ✓ $S_{n/p}$ is strongly correlated to the neutron-proton effective mass splitting!

The sensitivity of the nuclear matter parameters to $S_{n/p}$:



Extend Skyrme MDI(N³LO pseudo-potential):

$$\begin{split} V_{Ext} &= t_0^{(0)} (1 + x_0^{(0)} P_\sigma) \delta(r) \\ &+ \frac{1}{2} t_1^{(2)} \left(1 + x_1^{(2)} P_\sigma \right) \delta(r) [k'^2 + k^2] \\ &+ t_2^{(2)} \left(1 + x_2^{(2)} P_\sigma \right) [k' \cdot \delta(r) k] \\ &+ \frac{1}{4} t_1^{(4)} \left(1 + x_1^{(4)} P_\sigma \right) \delta(r) [(k'^2 + k^2)^2 + 4(k' \cdot k)^2] \\ &+ t_2^{(4)} \left(1 + x_2^{(4)} P_\sigma \right) \delta(r) (k' \cdot k) \left(k'^2 + k^2 \right) \\ &+ \frac{1}{4} t_1^{(6)} \left(1 + x_1^{(6)} P_\sigma \right) \delta(r) (k'^2 + k^2) [(k'^2 + k^2)^2 + 12(k' \cdot k)^2] \\ &+ t_2^{(6)} \left(1 + x_2^{(6)} P_\sigma \right) \delta(r) (k' \cdot k) \left[3(k'^2 + k^2)^2 + 4(k' \cdot k)^2 \right] \\ &+ \frac{1}{6} t_3 (1 + x_3 P_\sigma) [\rho(R)]^\sigma \delta(r) \\ &+ i W_0 \sigma \cdot [k' \times \delta(r) k] \end{split}$$

In our work, we assume a phenomenological momentum-dependent interaction as:

$$g(\mathbf{p} - \mathbf{p}') = \delta(\mathbf{r} - \mathbf{r}') \sum_{I=0}^{N} b_I (\mathbf{p} - \mathbf{p}')^{2I}$$

■ Appendix B. Determination of the expansion number N and coefficient bI.

The MDI potential energy density is taken as,

$$u_{md} = \tilde{C}_0 \int d^3p d^3p' f(\boldsymbol{r}, \boldsymbol{p}) f(\boldsymbol{r}, \boldsymbol{p}') g(\boldsymbol{p} - \boldsymbol{p}')$$

$$+ \tilde{D}_0 \int d^3p d^3p' f_n(\boldsymbol{r}, \boldsymbol{p}) f_n(\boldsymbol{r}, \boldsymbol{p}') g(\boldsymbol{p} - \boldsymbol{p}')$$

$$+ \tilde{D}_0 \int d^3p d^3p' f_p(\boldsymbol{r}, \boldsymbol{p}) f_p(\boldsymbol{r}, \boldsymbol{p}') g(\boldsymbol{p} - \boldsymbol{p}').$$

The nonlocal part of single-particle potential of nucleon in cold uniform nuclear matter is,

$$V_q^{md}(\rho, \delta, p) = \frac{\delta u_{md}}{\delta f_q} = 2\tilde{C}_0 \sum_{\tau=n,p} \int_0^{p_{Fq}} d^3 p' f_{\tau}(\boldsymbol{r}, \boldsymbol{p}') g(\boldsymbol{p} - \boldsymbol{p}')$$

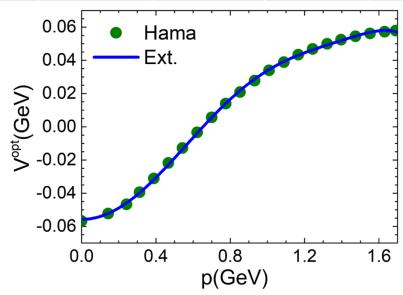
$$f_q = \frac{2}{(2\pi\hbar)^3} \Theta(\boldsymbol{p} - \boldsymbol{p}_{Fq}), q = \boldsymbol{n}, \boldsymbol{p}$$

$$+2\tilde{D}_0 \int_0^{p_{Fq}} d^3 p' f_q(\boldsymbol{r}, \boldsymbol{p}') g(\boldsymbol{p} - \boldsymbol{p}')$$

$$= V_{md}^{0} \pm V_{md}^{asy} \delta + \cdots$$
Fitting Hama data

By fitting the Hama data we can get the values of the parameters b_0 , b_1 , b_2 , b_3 , b_4 .

$b_0(GeV^2)$	$b_1(GeV^0)$	$b_2(GeV^{-2})$	$b_3(GeV^{-4})$	$b_4(GeV^{-4})$	$\widetilde{C}_0(fm^3/GeV)$	$\widetilde{D}_0(fm^3/GeV)$
-1.105	3.649	-2.608	0.826	-0.093	0.182	0.00



The experimental datas from:

S. Hama, B. C. Clark, E. D. Cooper, H. S. Sherif, and R. L. Mercer, Phys. Rev. C 41, 2737 (1990).

■ Appendix C. Nuclear matter parameters and its relation to the interaction parameters

Given the values of nuclear matter parameters at normal density, i.e., S_0 , L, K_0 , E_0 and ρ_0 , and the effective mass at normal density and fermi momentum,

(1)
$$P = \rho_0^2 \frac{dE/A(\rho_0, \delta=0)}{d\rho} = 0;$$

(2)
$$E_0 = E/A(\rho_0)$$
;

(3)
$$K_0 = 9\rho_0^2 \frac{\partial^2 E/A}{\partial \rho^2}|_{\rho_0};$$

(4)
$$\frac{m}{m_s^*}(\rho_0, p_F);$$

(5)
$$f_I(\rho_0, p_F)$$
;

(6)
$$S_0 = S(\rho_0)$$
;

(7)
$$L = 3\rho_0 \frac{\partial S(\rho)}{\partial \rho}|_{\rho_0}.$$



 $\alpha, \beta, \gamma, A_{sym}, B_{sym}, \widetilde{C}_0 \text{ and } \widetilde{D}_0$

Nuclear matter parameters and its relation to the interaction parameters

The saturation density ρ_0 is obtained by

$$P = \rho_0^2 \frac{dE/A(\rho_0, \delta=0)}{d\rho} = 0; \quad \frac{2}{5} \epsilon_F^0 + \frac{\alpha}{2} + \frac{\beta}{\gamma+1} \gamma + \sum_{I=1}^N \tilde{g}_{md}^I \left(\frac{2I}{3} + 1\right) \rho_0^{2I/3+1} = 0. \tag{1}$$

The binding energy E_0 is,

$$E_0 = E/A(\rho_0) = \frac{3}{5}\epsilon_F^0 + \frac{\alpha}{2} + \frac{\beta}{\nu+1} + \sum_{I=1}^N \tilde{g}_{md}^I \rho_0^{2I/3+1}.$$
 (2)

The incompressibility K_0 is,

$$K_0 = 9\rho_0^2 \frac{\partial^2 E/A}{\partial \rho^2}|_{\rho_0} = -\frac{6}{5}\epsilon_F^0 + 9\frac{\beta}{\gamma+1}\gamma(\gamma-1) + 6\sum_{I=1}^N \tilde{g}_{md}^I \left(\frac{2I}{3} + 1\right)I\rho_0^{2I/3+1}.$$
 (3)

The neutron/proton effective mass is,

$$\frac{m}{m_q^*}(\rho, p) = 1 + \frac{m}{p} \frac{\partial V_{md}^q}{\partial p}, q = n, p.
= 1 + 2\tilde{C}_0 m \left[\sum_{I=1}^N b_I \sum_{k=0, k \in even}^{2I} \tilde{\mathcal{A}}_{Ik} \sum_q \rho_q^{(2I-k+3)/3} k \times p^{k-2} \right]
+ 2\tilde{D}_0 m \left[\sum_{I=1}^N b_I \sum_{k=0, k \in even}^{2I} \tilde{\mathcal{A}}_{Ik} \rho_q^{(2I-k+3)/3} k \times p^{k-2} \right].$$

The isoscalar effective mass m_s^* can be obtained at $\rho_q = \frac{\rho}{2}$

$$\frac{m}{m_c^*}(\rho_0, p_F) = 1 + 4(\tilde{C}_0 + \frac{\tilde{D}_0}{2})m \left[\sum_{l=1}^N b_l \sum_{k=0, k \in even}^{2l} \tilde{\mathcal{A}}_{lk} (\frac{\rho_0}{2})^{(2l-k+3)/3} k \times p_F^{k-2} \right]. \tag{4}$$

Definition of m_s^* and m_v^* from:

E Chabanat, et, al, NPA 635(1998) 231-256

The isovector effective mass m_{ν}^* can be obtained at $\rho_a = 0$

$$\frac{m}{m_n^*}(\rho_0, p_F) = 1 + 4\tilde{C}_0 m \left[\sum_{I=1}^N b_I \sum_{k=0, k \in even}^{2I} \tilde{\mathcal{A}}_{Ik} (\frac{\rho_0}{2})^{(2I-k+3)/3} k \times p_F^{k-2} \right].$$

we define a quantity f_I ,

$$f_{I}(\rho_{0}, p_{F}) = \frac{m}{m_{s}^{*}} - \frac{m}{m_{v}^{*}} = 2\widetilde{D}_{0}m \left[\sum_{I=1}^{N} b_{I} \sum_{k=0, k \in even}^{2I} \widetilde{\mathcal{A}}_{Ik} \left(\frac{\rho_{0}}{2} \right)^{(2I-k+3)/3} k \times p_{F}^{k-2} \right]$$
 (5)

The symmetry energy coefficient S_0 is,

$$S_0 = S(\rho_0) = \frac{1}{3}\epsilon_F^0 + A_{sym} + B_{sym} + \sum_{I=1}^N \tilde{C}_{sym}^I \rho_0^{2I/3+1}.$$
 (6)

The slope of the symmetry energy L is,

$$L = 3\rho_0 \frac{\partial S(\rho)}{\partial \rho}|_{\rho_0} = \frac{2}{9} \epsilon_F^0 + 3A_{sym} + 3B_{sym} \gamma + 3\sum_{I=1}^N \tilde{C}_{sym}^I \left(\frac{2I}{3} + 1\right) \rho_0^{2I/3+1}$$
 (7)

Given the values of nuclear matter parameters at normal density, i.e., S_0 , L, K_0 , E_0 and ρ_0 , and the effective mass at normal density and fermi momentum,

the coefficients α , β , γ , A_{sym} , B_{sym} , \widetilde{C}_0 and \widetilde{D}_0 can be obtained as follows.

Parameters in Calculations:

TABLE I: Four sets of nuclear matter parameters used in this work. $f_I(\rho, p) = \frac{m}{m_s^*}$

· (, , _)				9755			
	f_{I}	m_s^*/m	L(MeV)	$S_0({ m MeV})$	$K_0({ m MeV})$	$E_0(MeV)$	$\rho_0(\mathrm{fm}^{-3})$
$(m_n^* < m_p^*)$	0.3	0.77	46,100	32	230	-16	0.16
$(m_n^* > m_p^*)$	-0.3	0.77	46,100	32	230	-16	0.16

TABLE II. The parameters used in the calculations, i.e., $K_0 = 230 \text{ MeV}$, $m_s^*/m = 0.77$, $S_0 = 32 \text{ MeV}$, and varies both L and f_I . The parameters α , β , A_{sym} , B_{sym} are in MeV. \tilde{C}_0 and \tilde{D}_0 are fm³GeV⁻¹.

Para.	$(L=46, f_I=0.3)$	$(L=46, f_I=-0.3)$	$(L=100, f_I=0.3)$	$(L=100, f_I=-0.3)$
α		-236.58	(-265.78)	
\boldsymbol{eta}		163.95	(194.93)	
γ		1.26	(1.22)	
A_{sym}	83.65 (108.44)	58.57 (62.73)	14.41 (25.32)	-10.67 (-20.40)
B_{sym}	-79.48 (-103.69)	-30.52 (-35.38)	-10.25 (-20.34)	38.72 (47.96)
$ ilde{C_0}$	-79.48 (-103.69) $-7.92 \times 10^{-4} (-2.08 \times 10^{-3})$	0.37 (1.00)	$-7.92 \times 10^{-4} \ (-2.08 \times 10^{-3})$	0.37 (1.00)
$ ilde{D}_0$	0.37 (1.00)	-0.37 (-1.00)	0.37 (1.00)	-0.37 (-1.00)

■ Appendix D. Relation between the effective mass splitting and symmetry potential

Relation between the effective mass splitting and symmetry potential

$$\frac{m}{m_q^*} = 1 + \frac{m}{p} \frac{\partial V_q}{\partial p}, \quad q = n, p.$$

According to the definition of neutron/proton effective mass,

$$\frac{m}{m_n^*} - \frac{m}{m_n^*} = 2m\left(\frac{\partial V_n}{\partial p^2} - \frac{\partial V_p}{\partial p^2}\right) = 4m\delta \frac{\partial V_{asy}}{\partial p^2}.$$

In addition,

$$\frac{m}{m_n^*} - \frac{m}{m_n^*} = -\frac{m(m_n^* - m_p^*)}{m_n^* m_n^*} \approx -(\frac{m}{m^*})^2 \Delta m_{np}^*.$$

The approximation comes from the $\frac{m_q^*}{m} = \frac{m^*}{m} \pm \frac{\delta m^*}{m}$, and the product of $\frac{m_n^* m_p^*}{m^2} = (\frac{m^*}{m})^2 - (\frac{\delta m^*}{m})^2 \approx (\frac{m^*}{m})^2$.

$$\Delta m_{np}^* \approx -(\frac{m^*}{m})^2 4m\delta \frac{\partial V_{asy}}{\partial p^2},$$