Probing antigravitational effects through CP violation on the Moon

G. M. Piacentino*3, A. Gioiosa^{†1,2}, D. Hajdukovic^{‡4}, A. Palladino §5, V. Testa \P^3 , and G. Venanzoni $\|^6$

¹Università degli Studi del Molise, Campobasso, Italy ²INFN, Sezione di Roma2, Roma, Italy ³INAF, Sezione di Monte Porzio Catone ⁴INFI, Cetinje, Montenegro ⁵Boston University, Boston, USA ⁶INFN, Sezione di Pisa, Pisa, Italy

July 17, 2019

Abstract

The environment on the Moon has numerous features that make it interesting not only for the study of astrophysical phenomena, but also elementary particle physics. In fact, vacuum conditions, low gravity, and exposure to a relatively intense irradiation of cosmic protons covering a large energy spectrum, make the lunar environment attractive for a wide range of particle physics experiments otherwise unworkable on Earth. We suggest one such experiment measuring the difference between the amount of CP violation as measured on the surface of the Earth and on the surface of the Moon, which could indicate quantum gravitational effects.

1 Introduction

The response of antimatter to a gravitational field has not yet been measured in any experimental test. A relatively large number of experiments on the gravitational interaction of anti-matter have been proposed and even started, e.g., ASACUSA [36, 37, 38, 39, 40], ATHENA [27, 33], AEGIS [1], ALPHA [2], ATRAP [3], and GBAR [4], to name only a few. Most of them are related with direct measurements of the interaction of anti-protons and of anti-nuclei and had, or have to, face technical difficulties due to copious electromagnetic noise

and the difficulties of producing and confining antiatoms. All of the ongoing experiments are still in the preliminary step of attempting to confine the anti-atoms in order to perform their measurements [26]..[66]. This paper is focused on a different aspect of the problem; we are looking for a variation in CP violation as a function of gravitational field intensity. Two experiments, CPLEAR [5] and KLOE [6], dealt with CP violation in the neutral kaon system, both looking for a time modulation of CP violation due to tidal contributions of the Moon, Sun, and galaxy to the gravitational field involved in kaon decays. The results were not incompatible but too weak to see any correlation.

In this paper, we propose a new approach to the problem, motivated by recent scientific interest in constructing a research village on the moon [7, 8, 9]. In fact, the European Space Agency's (ESA) Di-

^{*}giovanni.piacentino@unimol.it.it

[†]antonio.gioiosa@gmail.com

 $^{^{\}ddagger} dragan.hajdukovic@cern.ch$

[§]palladin@bu.edu

 $[\]P_{\text{vincenzo.testa@oa-roma.inaf.it}}$

graziano.venanzoni@lnf.infn.it

rector General recently suggested the creation of a human-robotic lunar outpost as a logical next step in human space exploration [10, 11]. Our approach takes advantage of the large difference between the gravitational field on the surface of the Earth and that on the Moon. On the Moon, the gravitational acceleration $g_{\rm Moon}=1.622~{\rm m/s^2}$ is only 16.54% of the corresponding acceleration on Earth. Such a difference should be extremely useful to our investigation on the contribution of gravity to the mixing in the neutral kaon system.

2 Antimatter and Gravitation

Any kind of relativistic and quantum theory of interactions, in the limit of perfectly static interactions, from the point of view of a Galileian inertial frame of reference, must converge to the classical expressions for the electromagnetic and gravitational fields.

$$F_{q,q'} = \frac{1}{4\pi\varepsilon_0} \frac{qq'}{r^2} \widehat{U}_{q,q'} \tag{1}$$

where $\varepsilon_0 = 8.85 \times 10^{12} \; \frac{C^2}{\mathrm{N \cdot m^2}}$, and

$$F_{m,m'} = \frac{1}{4\pi q_0} \frac{mm'}{r^2} \widehat{U}_{m,m'} \tag{2}$$

where $g_0 = -1.1935 \times 10^9 \frac{\text{kg}^2}{\text{N} \cdot \text{m}^3}$.

The different sign of the two constants, e_0 and g_0 , means that similar charges repel for the electromagnetic interaction but attract for the gravitational interaction. This means that there is an electromagnetic screening effect, i.e., a bound state of electric charges seems to have less charge than the sum of the individual charges, while on the contrary a bound state of massive particles seems to have more mass from the point of view of an external observer.

In the Newtonian formulation the definition of inertial mass is related to the definition of force and acceleration. Therefore this implies a dependence on the definition of time. Negative mass should be possible in the case of negative time. Particles with intrinsic reversed time evolution are classically possible if reinterpreted as classical antimatter or tachyons from the point of view of a Galileian reference frame so that in this frame it is possible to

consider a repulsion between particle and antiparticle mediated by gravitation [17]. The definition of antimatter in quantum field theories was provided by Feynman and Stueckelberg [12, 13]. It implies that matter-antimatter transformations are equivalent to charge conjugation and time reversal in a flat space-time. When General Relativity (GR) is considered, space-time is not in general flat, however the GR equations remain symmetric in the time coordinate even if it is not explicitly concerned about antimatter. Several theories that allow repulsion between matter and antimatter have been proposed []. In fact, this hypothesis may overcome many of the open problems in astrophysics and cosmology:

- *i.* In the observed universe, matter prevails over antimatter (Matter-Antimatter Asymmetry);
- ii. The Cosmic Microwave Background (CMB) is neither anisotropic nor inhomogeneous enough to be compatible with the Big Bang model without the introduction of an unknown interaction (Inflation);
- iii. Given the attractive gravitational interaction, we expect a slowing of the acceleration of the expansion of the universe. On the contrary the acceleration seems to be increasing (Dark Energy); and
- iv. The gravitational field of galaxies, clusters, and even our own solar system seems much stronger than expected due to visible matter (Dark Matter).

Most of them could be addressed by the same hypothesis. In fact, the net force in a universe with matter-antimatter symmetry should be repulsive because in such a case the forces do not necessarily cancel out. In an ionic solid, a very similar physical system, the same number of positive and negative charges are present, but the overall electrostatic force on the crystal is attractive (compensated by the Fermi exclusion principle). If we now visualize each positive ion as a particle, and each negative ion as an antiparticle, and replace the electrostatic interaction by the gravitational potential, we obtain from the Madelung-like model:

$$U_g = \frac{1}{2} N \alpha \frac{m^2}{R},\tag{3}$$

where U_g is the total gravitational energy, N is the total number of particles with mass m,R the separation between nearest antimatter, and α is the Madelung constant. For simplicity, and to keep the analogy with a crystal, m and R are assumed to be the same for the entire matter-antimatter mixture. The Madelung constant has values between 1.6 and 1.8 for most crystal structures. The overall force on the universe $(\mathrm{d}U_g/\mathrm{d}R)$ is repulsive [25]. Therefore, gravitational repulsion between matter and antimatter may have dramatic consequences in astrophysics and cosmology.

Several authors [14, 15, 16] have studied possible consequences of anomalous antimatter gravity if we live in a symmetric universe, i.e., a universe with the same amount of matter and antimatter, with antimatter somehow hidden (for instance in the vacuum). According to a radically different direction of research [17], we live in an asymmetric universe (i.e., matter dominates antimatter) but the gravitational properties of antimatter (negative gravitational mass) may have a major impact because of the quantum vacuum fluctuations interpreted as virtual gravitational dipoles with the potential to explain the nature of what we call dark matter. According to this model of the universe, the only matter-energy content of the universe is the Standard Model matter (i.e., matter composed of quarks and leptons interacting through the exchange of gauge bosons). Thus, the phenomena usually attributed to hypothetical dark matter may be considered as a consequence of the local variation of the gravitational polarization due to the baryonic matter immersed in the quantum vacuum.

3 Antigravity and CP Violation

In 1961, Good [18] calculated, using absolute potentials, that a repulsive gravitational interaction of antimatter should introduce a regeneration of kaons, thus resulting in an anomalously large level of CP violation. At the time of his paper, CP violation in the neutral kaon system was unknown and the argument appeared strong and elegant, but in light of the measurement of CP violation [19], Good's paper has instead turned out to be a strong argument in favor of antigravity. Chardin reformu-

lated Good's argument in terms of relative potentials and showed that the gravitational field on the surface of the Earth is of the right intensity, i.e., of the required order of magnitude to cause CP violation. In particular, the mixing time of the kaon,

$$\Delta \tau = 5.9 \times 10^{10} \text{ s} \simeq 6\tau_{k_a}$$
 (4)

is long enough for the gravitational field of the Earth to attract the matter and repel the antimatter components of the K meson to induce a separation,

$$\Delta \zeta = g \Delta \tau^2 \ , \tag{5}$$

between them inducing regeneration, thus providing a mechanism for indirect CP violation. The amount of such a contribution should be roughly the same order of magnitude as the spatial sepration divided by the Compton wavelength:

$$\chi = \frac{\Delta \zeta}{\Delta L_{\rm K}} = \frac{g\tau_{k_l}}{\frac{\hbar}{m_{\rm K}c}} = \Omega \frac{g\frac{\pi^2 \hbar^2}{\Delta m^2 c^4}}{\frac{\hbar}{m_{\rm K}c}} \qquad (6)$$

$$\chi = \Omega \frac{\pi^2 \hbar g m_{\rm K}}{\Delta m^2 c^3} = \Omega \times 0.88 \times 10^{-3} ,$$

which happens to be the same order of magnitude as the CP violation parameter, ε , as measured on Earth's surface. If we calculate (6) given the gravitational strength on the Moon's surface we would expect the measured ε to be 80% smaller than the ε measured on Earth's surface, assuming gravitational repulsion beween matter and antimatter.

4 Moon Surface Experiment

As mentioned above, we expect only 16% of the CP violation induced by gravity in an experiment performed in the gravitational field on the Moon. Since

$$R = \frac{\Gamma(K_L \to \pi^+ \pi^-)}{\Gamma(K_L \to \pi^+ \pi^- \pi^0)} \tag{7}$$

is quadratic in ε , we expect a very large difference in the value of R when measured on the surface of the Moon. In order to produce the K_L on the Moon we plan to take advantage of the flux of cosmic protons continuously hitting the Moon's surface. A direct measurement of the flux of protons on the lunar surface has not yet been made, but Ackermann

et al. [20] fit the data of the gamma albedo from the Moon surface due to the incoming proton flux finding it to be equal, within a 10% uncertainty, to the proton flux measured by AMS-02 [21] and PAMELA [22]. In the following, we considered the proton flux as determined by AMS-02.

In [23] and [24], we considered an experimental apparatus with cosmic protons incident upon a cylindrical target followed by a detector region consisting of a cylindrical tracking volume (1 meter diameter, 1 meter deep) in a Low Earth Orbit (LEO). On a satellite in LEO, the effect due to gravity on R is $\sim 20\%$. If instead we place the apparatus on the surface of the Moon, we expect a significant improvement in the results, with an effect on R of $\sim 85\%$.

We performed Geant4 simulations using the angular and energy spectrum of the incident cosmic protons as measured by AMS-02 spectrometer. We simulated incident protons with $\theta_{\rm max}=45^{\circ}$ over a target surface corresponding to a $\pi/4$ solid-angle acceptance. We used the same apparatus as in [23] and [24], with the exception of the target. There, we used a cylindrical Tungsten target (1 m diameter, 9 cm deep), while here we considered an active target consisting of alternating layers of Tungsten and scintillating crystals (Stolzite, PbWO₄) for a total depth of 18 cm), to be read with photodiodes.

We studied the production of K_L that would decay inside the volume of a detector and their distribution on the detector cross section. The particles decaying before z = 18 cm decay inside the PbWO₄ target and are lost, however those decaying inside a downstream cylindrical tracking region could potentially be reconstructed. At this stage in evaluating the the feasibility of performing the experiment, we made no attempt to fully reconstruct the K_L decay products inside the detector. We just took as a crude approximation the reconstruction efficiency as equal to 1 inside the fiducial volume of the tracking region. With protons incident on the target distributed in angle and energy as cosmic protons, we estimate the number of K_L decays per year inside two possible tracking volumes, listed in Table 1, with and without additional kinematical cut on the axial momentum at the decay vertex. Table 2 shows the length of time it would take to record sufficient K_L decays to provide 3σ and 5σ measurements of R.

Volume,	(r < 50 cm)	(r < 50 cm,	(r < 100 cm,
kinematics		$p_z < 0.5 \text{ GeV/c})$	$p_z < 0.5 \; {\rm GeV/c})$
$\frac{N(K_{\text{L}} \text{decays})}{\text{year}}$	3.54×10^{6}	1.50×10^{6}	3.36×10^{6}
$\frac{N(K_{\text{S}} \text{decays})}{N(K_{\text{L}} \text{decays})}$	3.03×10^{-1}	4.15×10^{-5}	

Table 1: Number of K_L that decay within various tracking region volumes 50 cm downstream of the target (50 < z < 150 cm, r < 50 cm or r < 100 cm), with and without a kinematical cut on the axial momentum at the $K_{\rm S,L}$ decay vertex.

	Requirement		Simulation result	
	3σ	5σ	3σ	5σ
N (K _L decays)	$> 2.5 \times 10^4$	$> 7 \times 10^4$	6 days	17 days
$\frac{N(K_{S} \text{ decays})}{N(K_{L} \text{ decays})}$	$< 1 \times 10^{-4}$	$< 5.7 \times 10^{-5}$	4.15×10^{-5}	
$\frac{\delta N(K_L \to \pi \mu \nu)}{N(K_L \to \pi \pi)}$	$< 4 \times 10^{-2}$	$< 2 \times 10^{-2}$	kinematical cuts	

Table 2: Critical parameters necessary for 3σ and 5σ measurements of a gravitational modulation in the level of CP violation (85% change in R) along with the values obtained from our Monte Carlo simulation. The results take into account a basic geometrical event selection of $K_{\rm S,L}$ decay vertices within a 1 m \times 1 m cylindrical tracking volume 50 cm downstream of the target (50 < z < 150 cm, r < 50 cm), and axial momentum at the $K_{\rm S,L}$ decay vertex of p_z < 0.5 GeV/c. These values assume a 100% detection efficiency, 2% (4%) statistical and 2% (4%) systematic fractional uncertainties for 5σ (3 σ).

5 Conclusions

By constructing a detector consisting of a 1 m diameter, 18 cm thick active target and a 1 m diameter \times 1 m deep tracking and particle identification system, and placing it on the surface of the Moon, we could perform a direct measurement of the ratio of the number of K_L decaying to two charged pions to those decaying to three pions in a low-gravity environment. We estimate that it will take 6 days (17 days) to record sufficient K_L decays for a 3σ (5σ) measurement. Any difference between the amount of CP violation on the surface of the Moon compared to the level of CP violation on the surface of Earth would be an indication of a quantum gravitational effect.

References

- [1] A. Kellerbauer, et al. (AEgIS Collaboration), Proposed antimatter gravity measurement with an antihydrogen beam, Nucl. Instrum. Methods Phys. Res. B **266** (2008) 351.
- [2] A.E. Charman, et al. (ALPHA Collaboration), Description and first application of a new technique to measure the gravitational mass of antihydrogen, Nature Comm. 4 (2013) 1785.
- [3] G. Gabrielse, et al. (ATRAP Collaboration), Trapped antihydrogen in its ground state, Phys. Rev. Lett. 108 (2012) 113002.
- [4] G. Chardin, P. Grandemange, D. Lunney, et al., Proposal to Measure the Gravitational Behaviour of Antihydrogen at Rest, Tech. Rep. CERN-SPSC-2011- 029, SPSC-P-342.
- [5] CPLEAR Collaboration, Tests of the Equivalence Principle with Neutral Kaons, Phys. Lett. B **452** (1999) 425–433. (arXiv:hep-ex/9903005) CERN-EP/99-22
- [6] G. Mambriani, L. Trentadue, Testing CP Conservation at KLOE. (2000) (arXiv:hep-ex/0007004).
- [7] G.J. Taylor, The need for a lunar base: answering basic questions about planetary science. In: Lunar Bases and Space Activities of the 21st Century (ed. W.W. Mendell), Houston, TX, (1985) 189–197.
- [8] I.A. Crawford, et al., Back to the Moon: the scientific rationale for resuming lunar surface exploration, Planet. Space Sci., 74, (2012) 3–14.
- [9] C. McKay, The case for a NASA research base on the Moon. New Space, 1, 162–166, (2013).
- [10] J. Woerner and B. Foing, The "Moon Village" concept and initiative, Annual Meeting of the Lunar Exploration Analysis Group, 1–3 November 2016, Columbia, Maryland. LPI Contribution No. 1960, id.5084, (2016).
- [11] I.A. Crawford, Science enabled by a Moon Village, International Academy of Astronautics, 10th IAA Symposium on the Future of

- Space Exploration: Towards the Moon Village and Beyond, 27–29 June 2017, Torino, Italy, (2017). https://arxiv.org/abs/1706.06698
- [12] R. P. Feynman, The reason for antiparticles in R. .P. Feynman and S. Weinberg, The 1986 Dirac memorial lectures. Cambridge University Press. ISBN 0-521-34000-4.
- [13] E. C. G. Stueckelberg, La mecanique du point materiel en theorie del relativite et en theorie des quanta, Helv. Phys. Act. 15, 23–37 (1942).
- [14] A. Benoit-Levy and G. Chardin, *Introducing the Dirac-Milne universe*, Astron. Astrophys. **537**, A78 (2012).
- [15] G. Chardin, *CP violation and antigravity (revisited)*, Nuclear Physics A **558** (1993) 477c.
- [16] M. Villata, *CPT symmetry and antimatter gravity in general relativity*, EPL **94**, 2, 20001 (2011).
- [17] D. Hajdukovic. Quantum vacuum and virtual gravitational dipoles: the solution to the dark energy problem?, Astrophysics and Space Science. Volume **339**, Issue 1, (2012) 1–5.
- [18] M. L. Good, K_2^0 and the Equivalence Principle, Phys.Rev. **121** (1961) 311–313.
- [19] J. H. Christenson, J. W. Cronin, V. L. Fitch and R. Turlay, *Evidence for the 2\pi Decay of the K_2^0 Meson*, Phys. Rev. Lett. 13, **138** (1964).
- [20] M. Ackerman et al. [Fermi-LAT Collaboration], Measurement of the high-energy gamma-ray emission from the Moon with the Fermi Large Area Telescope, Phys. Rev. D 93 8, 082001 (2016).
- [21] M. Aguilar et al., Phys. Rev. Lett 110 (2013) 141102.
 - K. Luebelsmeyer et al., Nucl. Instr. Meth. A **654** (2011) 639.
 - B. Alpat et al., Nucl. Instr. Meth. A **613** (2010) 207.
 - A. Basili et al., Nucl. Instr. Meth. A 707 (2013) 99; V. Bindi et al., Nucl. Instr. Meth. A 623 (2010) 968.

- [22] O. Adriani, et al., Ten Years of PAMELA in Space, Rivista del Nuovo Cimento, 10, (2017) 473–522.
- [23] G.M. Piacentino, A. Palladino, G. Venanzoni, Measuring gravitational effects on antimatter in space, Physics of the Dark Universe, 13, (2016) 162–165.
- [24] G.M. Piacentino, A. Gioiosa, A. Palladino, G. Venanzoni, Measuring gravitational effects on antimatter in space, Advances in Dark Matter and Particle Physics, EPJ Web of Conferences 142, 01023 (2017).
- [25] J. M. Ripalda "Time reversal and negative energies in general relativity" arXiv:gr-qc/9906012.
- [26] W A Bertsche et al 2014 J. Phys. B: At. Mol. Opt. Phys. 48 232001.
- [27] ATHENA Collaboration 1996 CERN-SPSLC-96-47; SPSLCP-302 (CERN) Antihydrogen production and precision experiments http://cds.cern.ch/record/314017.
- [28] ATRAP Collaboration 1997 CERN-SPSC-97-8; SPSC-P-306 (CERN) Proposal presented to the SPSLC: the production and study of cold antihydrogen https://cds.cern.ch/record/324128.
- [29] ASACUSA Collaboration 1997 CERN-SPSC-97-19; SPSCP-307 (CERN) Atomic spectroscopy and collisions using slow antiprotons ttp://cds.cern.ch/record/622250.
- [30] ALPHA Collaboration 2005 CERN-SPSC-2005-006; SPSCP-325 (CERN) ALPHA Proposal http://cds.cern.ch/record/814351.
- 2007 CERN-[31] AEgIS Collaboration SPSC-P-334; SPSC-2007-017 (CERN) Proposal for the AEGIS experiment the CERN antiproton decelerator http://cds.cern.ch/record/1037532.
- [32] GBAR Collaboration 2011 CERN-SPSC-2011-029; SPSC-P-342 (CERN) Proposal tomeasure the gravitationalbehaviour of antihydrogen at rest https://cds.cern.ch/record/1386684.

- [33] C. Amsler et al., "The ATHENA experiment for the study of antihydrogen, Int. J. Mod. Phys. A 29, no. 20, 1430035 (2014).
- [34] R. McConnell *et al.* [ATRAP Collaboration], "Large numbers of cold positronium atoms created in laser-selected Rydberg states using resonant charge exchange; J. Phys. B **49**, no. 6, 064002 (2016).
- [35] M. Zieliński et al. [ATRAP Collaboration], "Studies on Antihydrogen Atoms with the ATRAP Experiment at CERN; Acta Phys. Polon. Supp. 6, no. 4, 1093 (2013).
- [36] C. Malbrunot, "A hydrogen beam to characterize the ASACUSA antihydrogen hyperfine spectrometer", arXiv:1812.06736v1 [physics.atom-ph] 17 Dec 2018.
- [37] E. Widmann, "Hyperfine spectroscopy of hydrogen and antihydrogen in ASACUSA", arXiv:1809.00875v2 [physics.atom-ph] 16 Dec 2018.
- [38] B. Kolbinger *et al.*, "Recent Developments from ASACUSA on Antihydrogen Detection", EPJ Web Conf. **181**, 01003 (2018).
- [39] C. Malbrunot, "The ASACUSA antihydrogen and hydrogen program: results and prospects", arXiv:1710.03288v1 [physics.atomph] 9 Oct 2017.
- [40] M. Diermaier, "An atomic hydrogen beam to test ASACUSA's apparatus for anti-hydrogen spectroscopy", arXiv:1501.02635v1 [physics.atom-ph] 12 Jan 2015.
- [41] S. Aghion *et al.* "Antiproton tagging and vertex fitting in a Timepix3 detector AEgIS Collaboration", 21 pp. Published in JINST 13 (2018) no.06.
- [42] S. Aghion *et al.*, "Compression of a mixed antiproton and electron non-neutral plasma to high densities", 11 pp. Eur.Phys.J. D72 (2018) no.4, 76.
- [43] M. Doser et al., "AEgIS at ELENA: outlook for physics with a pulsed cold antihydrogen beam", Phil. Trans. Roy. Soc. Lond. A376 (2018) no. 2116, 20170274.

- [44] P. Yzombard (LAC, Orsay) et al., "Overview of recent work on laser excitation of positronium for the formation of antihydrogen", 11 pp. JPS Conf.Proc. 18 (2017) 011026.
- [45] R. Caravita et al., "Advances in Ps Manipulations and Laser Studies in the AEgIS Experiment", Acta Phys.Polon. B48 (2017) 1583.
- [46] G. Consolati et al., "Positronium for Antihydrogen Production in the AEGIS Experiment", 7 pp. Acta Phys.Polon. A132 (2017) no.5, 1443-1449.
- [47] S. Mariazzi *et al.*, "Positron Manipulation and Positronium Laser Excitation in AEgIS", Defect and Diffusion Forum 373 (2017) 11-16.
- [48] R.S. Brusa *et al.*, "The AEgIS experiment at CERN: measuring antihydrogen free-fall in earth's gravitational field to test WEP with antimatter", J.Phys.Conf.Ser. 791 (2017) no.1, 012014.
- [49] P. Lansonneur *et al.*, "AEgIS Experiment: Status and Outlook", Conference: C17-03-25, p.87-92.
- [50] P. Lansonneur, "Moiré Deflectometry with a Low-Energy Ion Beam for the AEGIS Experiment Pierre", 2017LYSE1191.
- [51] A. Kellerbauer *et al.* (AEGIS Collaboration), "Probing antimatter gravity The AEGIS experiment at CERN", EPJ Web Conf. 126 (2016) 02016.
- [52] S. Aghion et al. (AEgIS Collaboration), "Laser excitation of the n=3 level of positronium for antihydrogen production", Phys.Rev. A94 (2016) no.1, 012507. D. Pagano et al., "The Weak Equivalence Principle With Antimatter: The AEgIS Experiment At CERN", Conference: C16-03-19, p.161-164.
- [53] A. Gligorova (AEGIS Collaboration), "Antimatter annihilation detection with AEgIS", PoS VERTEX2015 (2015) 024.
- [54] M. Kimura et al.(AEgIS Collaboration), "Testing the Weak Equivalence Principle with an antimatter beam at CERN", J.Phys.Conf.Ser. 631 (2015) no.1, 012047.

- [55] G. Consolati et al., "Experiments with lowenergy antimatter", EPJ Web Conf. 96 (2015) 01007.
- [56] C. Pistillo *et al.*, "Emulsion detectors for the antihydrogen detection in AEgIS", Hyperfine Interact. 233 (2015) no.1-3, 29-34.
- [57] B. Mansoulié (GBAR Collaboration), "Status of the GBAR experiment at CERN", Hyperfine Interact. 240 (2019) no.1, 11.
- [58] G. Chardin, G. Manfredi, "Gravity, antimatter and the Dirac-Milne universe Hyperfine Interact. 239 (2018) no.1, 45.
- [59] D.A. Cooke, A. Husson, D. Lunney, P. Crivelli, "Proposal for enhancement of antihydrogen ion production in the GBAR experiment", EPJ Web Conf. 181 (2018) 01002.
- [60] P. Lotrus, G. Durand, A. Gaget, A. Gomes, Y. Le Noa, J.F. Lecointe, J.Y. Roussé, "Status of the GBAR control project at CERN", Conference: C17-10-08, p.TUPHA059, 2018.
- [61] P. Crivelli, N. Kolachevsky, "Optical trapping of anti-hydrogen towards an atomic anticlock", e-Print: arXiv:1707.02214.
- [62] P. Pérez *et al.*, "The GBAR antimatter gravity experiment", Hyperfine Interact. 233 (2015) no.1-3, 21-27.
- [63] P. Perez, "GBAR status and goals", Conference: C15-03-21, p.53-58.
- [64] G. Dufour, "Prospects for studies of the free fall and gravitational quantum states of antimatter", Adv. High Energy Phys. 2015 (2015) 379642.
- [65] Y. Sacquin, "The GBAR experiment: Gravitational behaviour of antihydrogen at rest", Eur.Phys.J. D68 (2014) 31.
- [66] D. P. Van Der Werf (GBAR Collaboration), "The GBAR experiment",J.Mod.Phys.Conf.Ser. 30 (2014) 1460263.