

# The Galactic dynamics revealed by the filamentary structure in the atomic hydrogen (HI) emission

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**ECOgal collaboration:** S. Molinari, R. S. Klessen, P. Hennebelle, S. C. O. Glover, A. Trafficante, E. Schisano, D. Elia, M. Sormani, R., Tress, P. Girichidis, R. J. Smith, T. Colman

## How do the Sun and other stars form?

- What is the Sun made of?

# The composition of the Sun

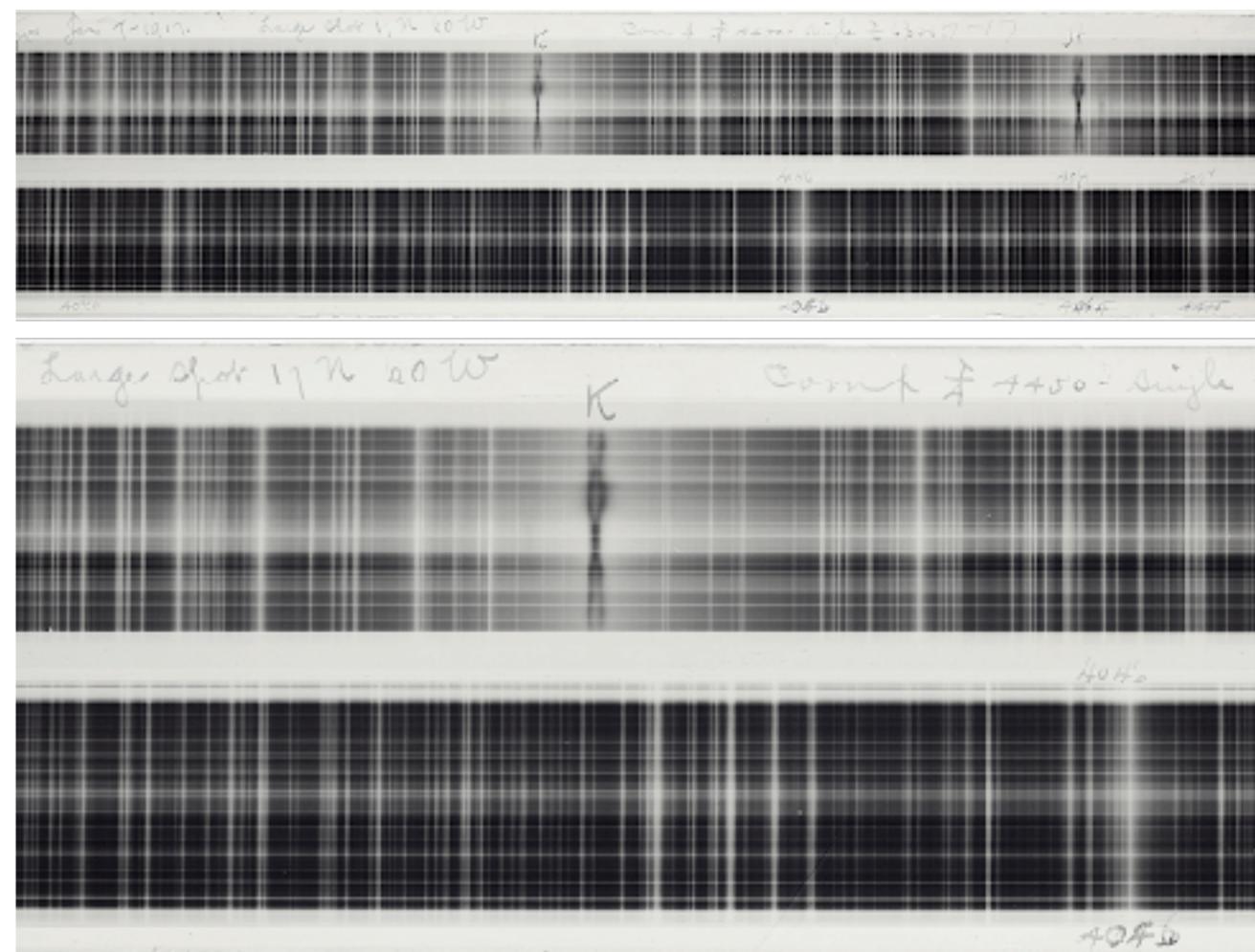
Payne, C. 1925. PhD. Radcliffe College.

## STELLAR ATMOSPHERES

A CONTRIBUTION TO THE OBSERVATIONAL  
STUDY OF HIGH TEMPERATURE IN THE  
REVERSING LAYERS OF STARS

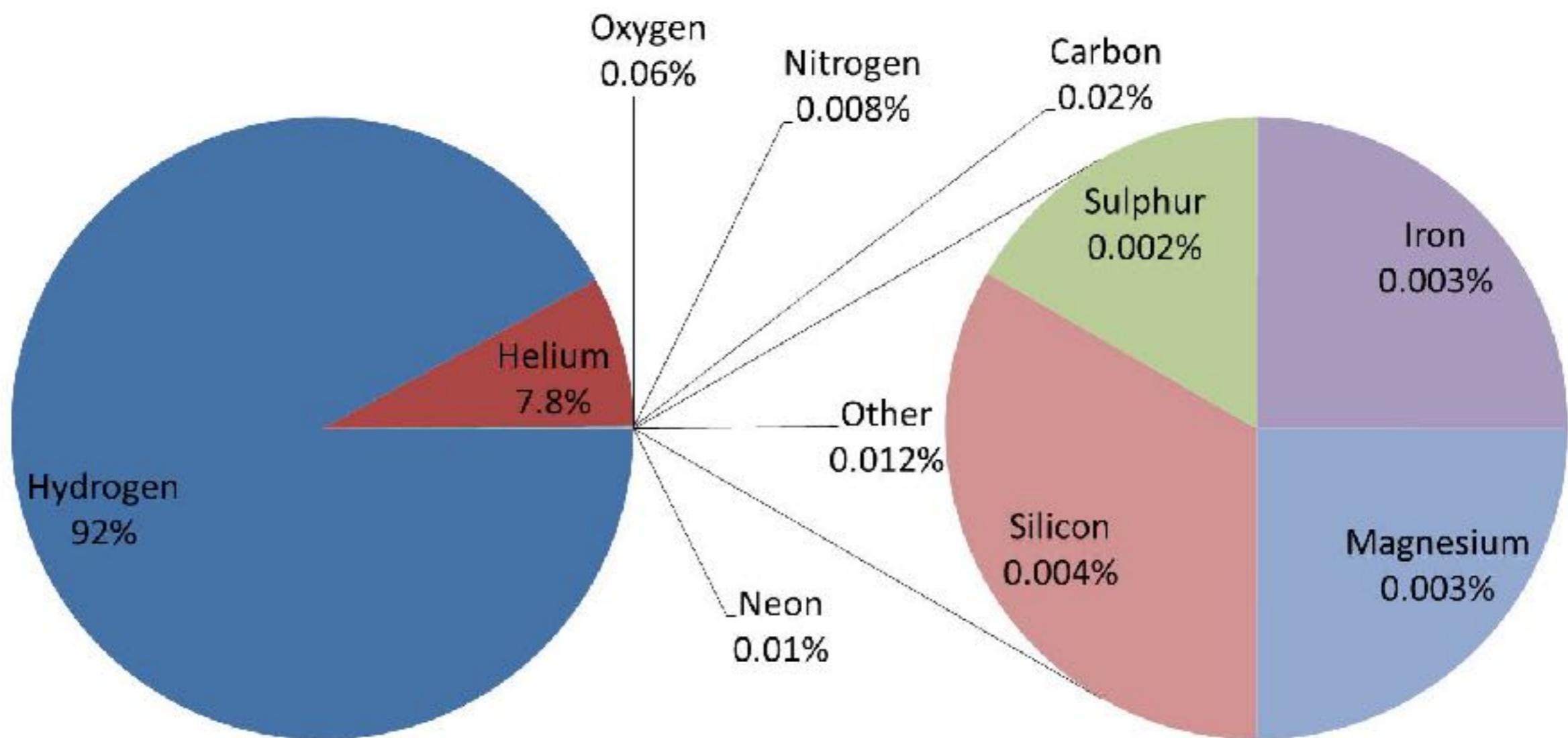
BY

CECILIA H. PAYNE



# The composition of the Sun (by number of atoms)

M. Asplund, A. Amarsi, and N. Grevesse 2021, A&A, 653, A141



## How do the Sun and other stars form?

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Mostly hydrogen ( $n \sim 10^{25} \text{ cm}^{-3}$ )
- Where is the hydrogen before it becomes a star?

# Atomic hydrogen emission

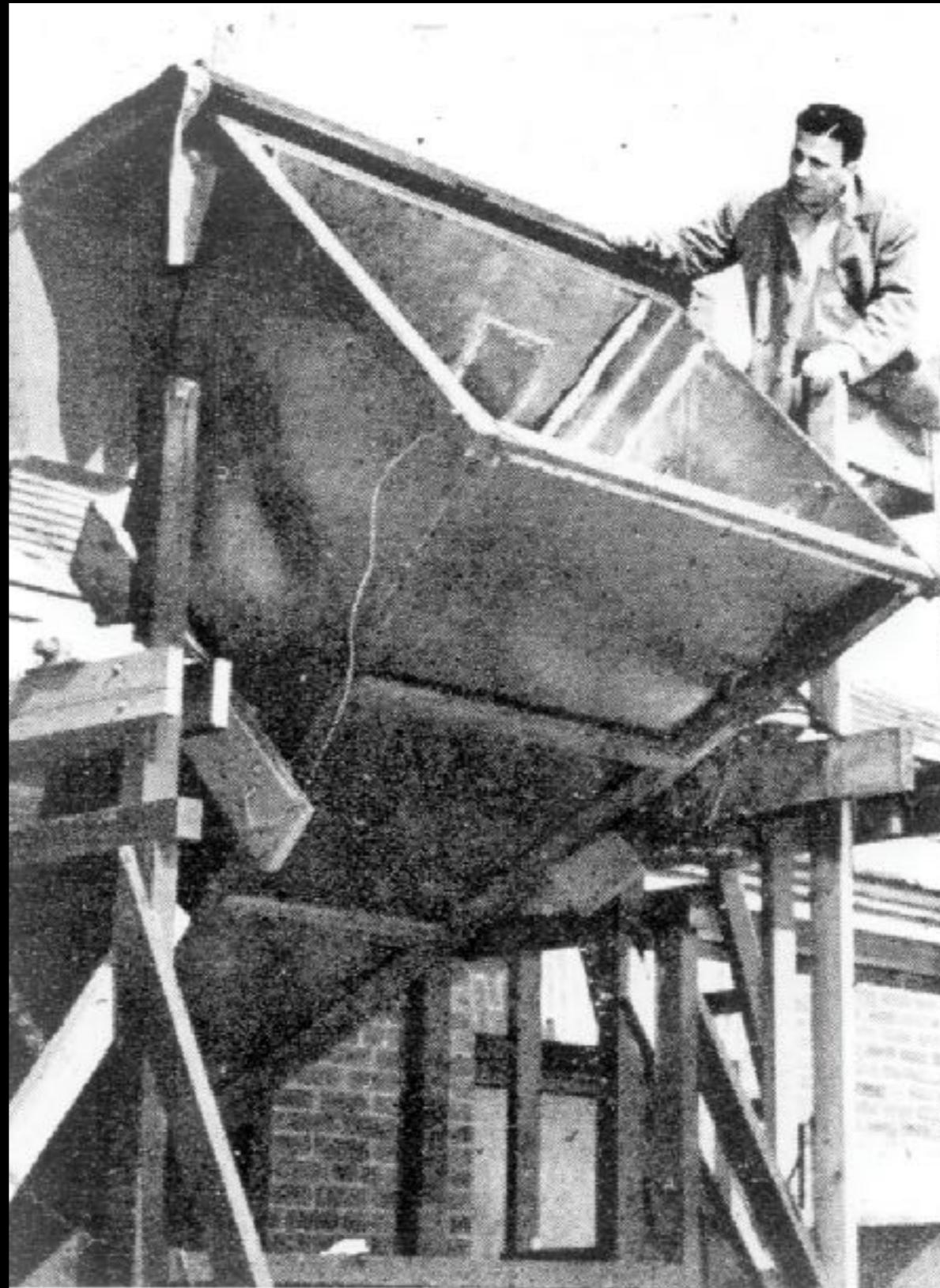
Netherlands. June, 1941.



# Atomic hydrogen emission

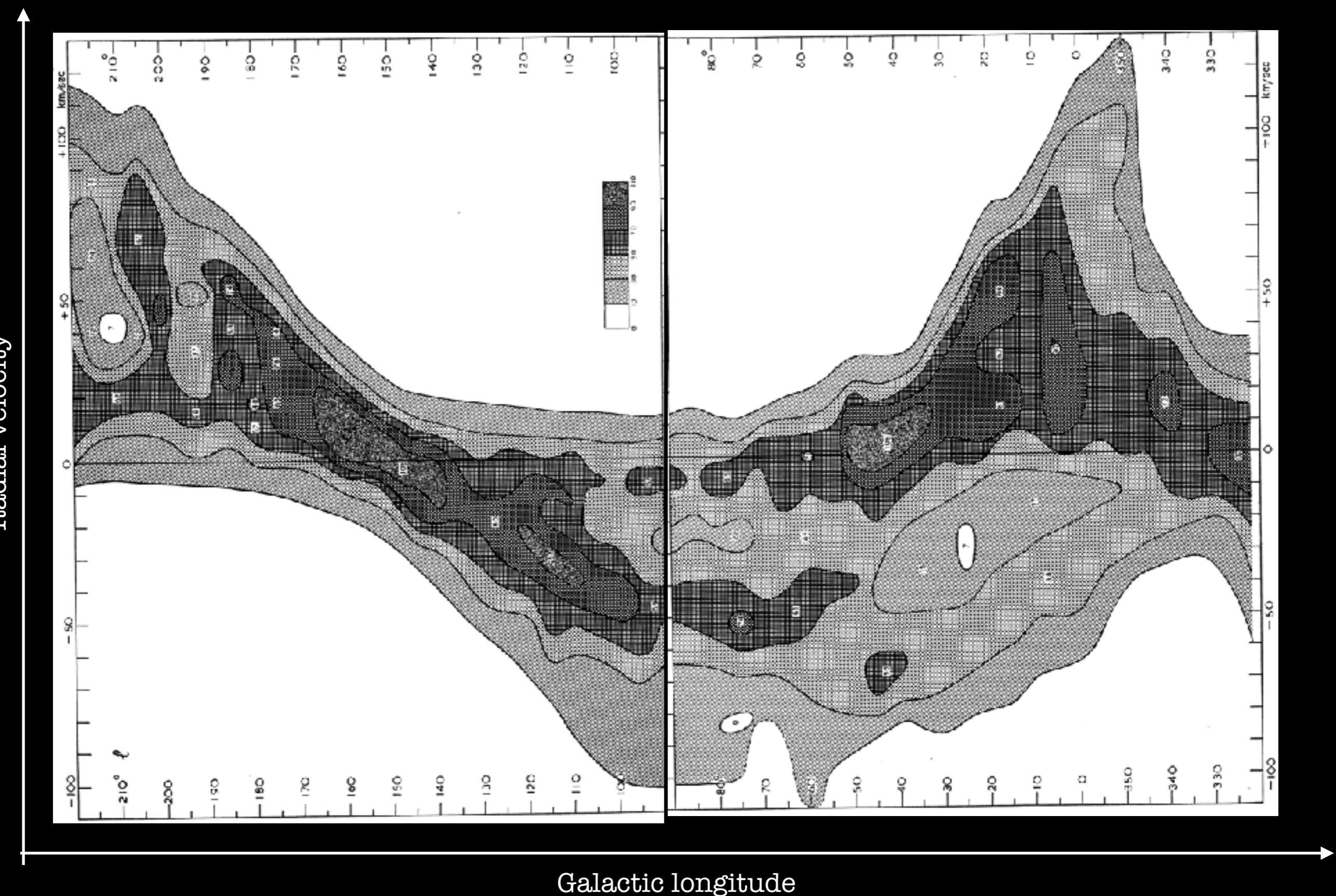
Harold Ewen y Edward Purcell.

Lyman Laboratory - Harvard University. March, 1951



# Atomic hydrogen emission toward the Milky Way

van de Hulst, Muller & Oort (1954).

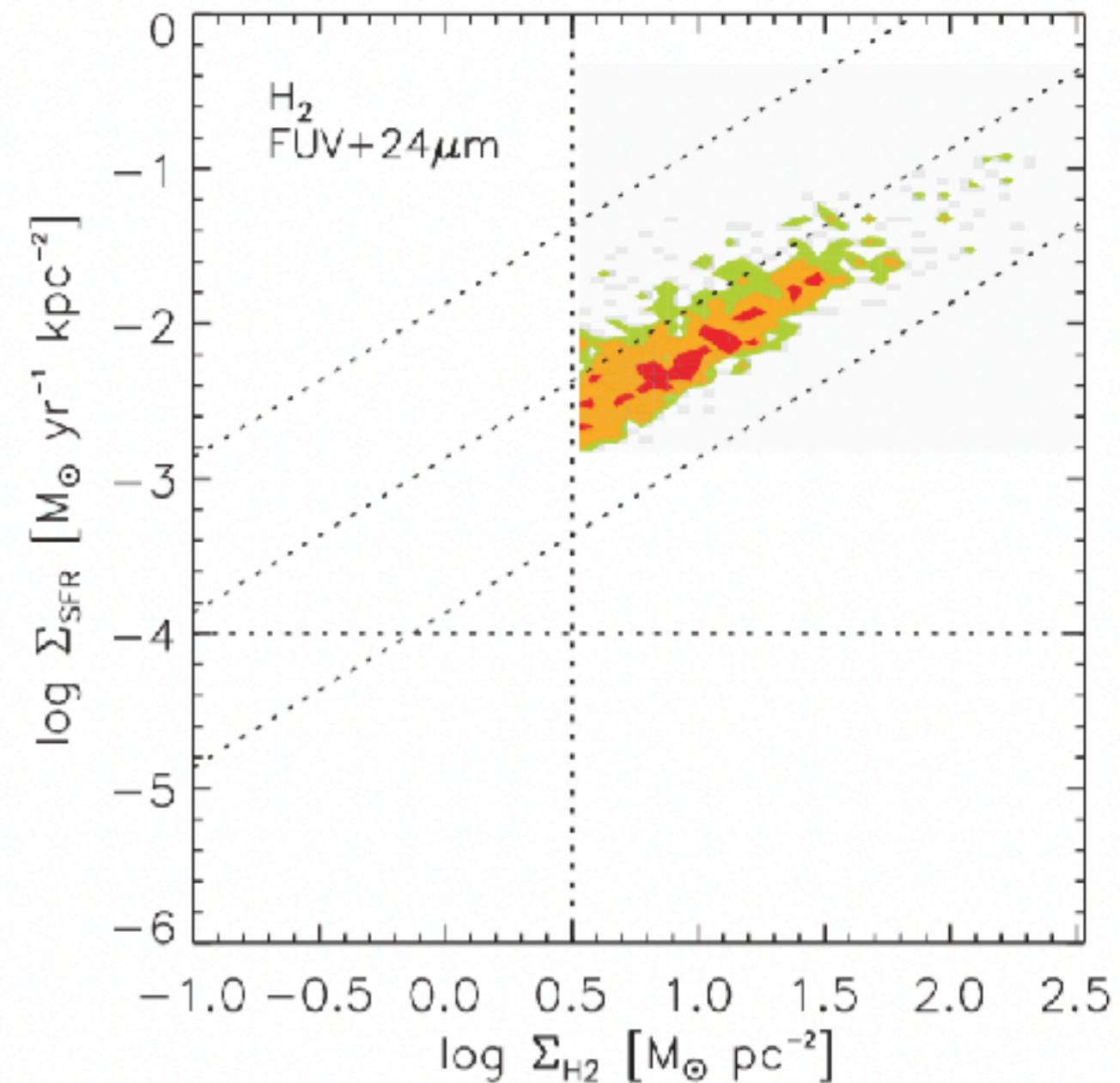
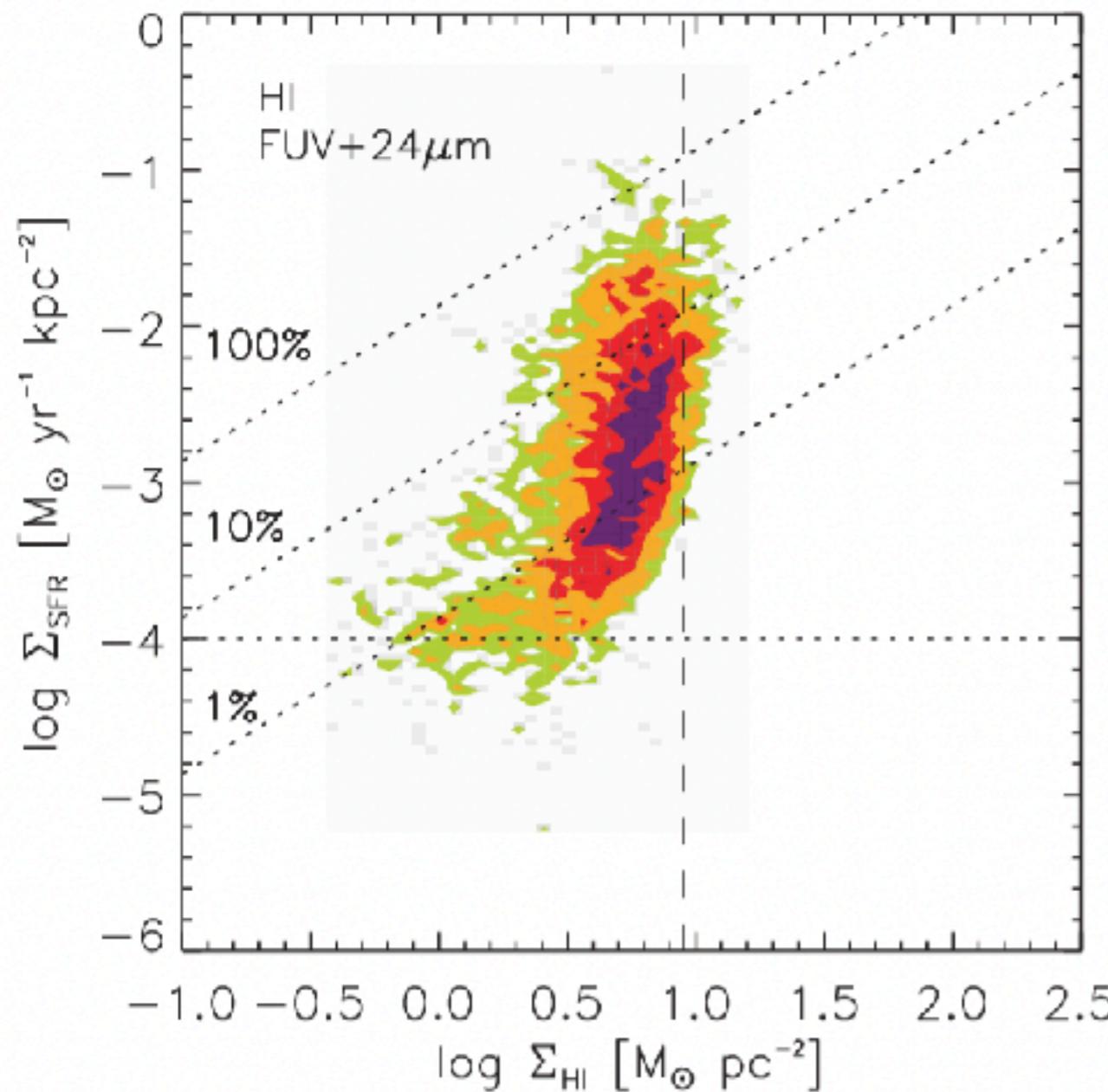


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Almost everywhere in the Milky Way ( $n \sim 1 \text{ cm}^{-3}$ )
- How is the hydrogen becoming stars?

# Star formation and gas surface density

Bigiel et al. 2008; 18 nearby galaxies ( $2 < d < 15$  Mpc), 750 pc resolution

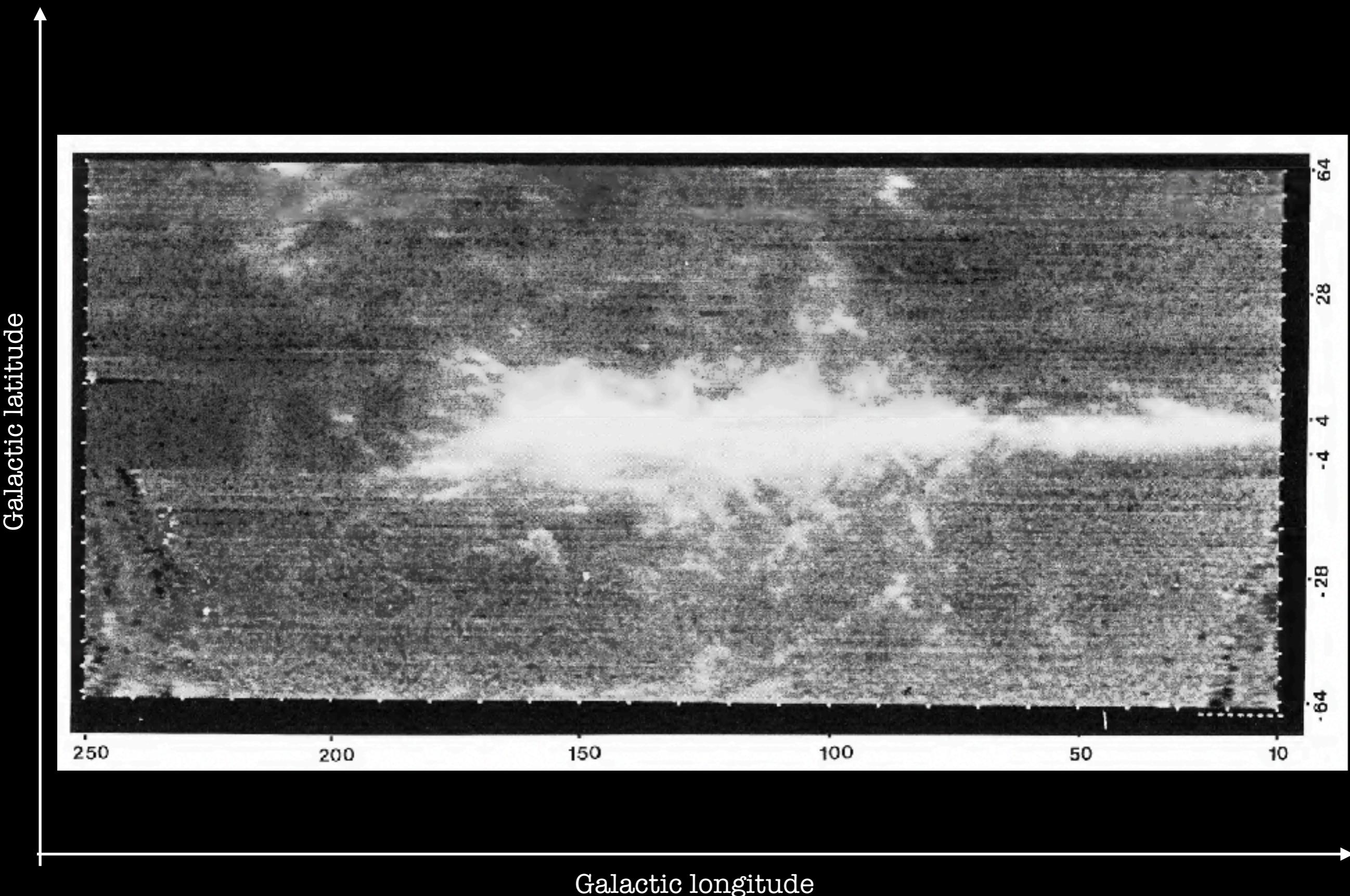


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Hard question.
- How is the hydrogen before becoming stars?

# Atomic hydrogen emission toward the Milky Way

Heiles (1984).



# The HI, OH, and Recombination-line (THOR) survey

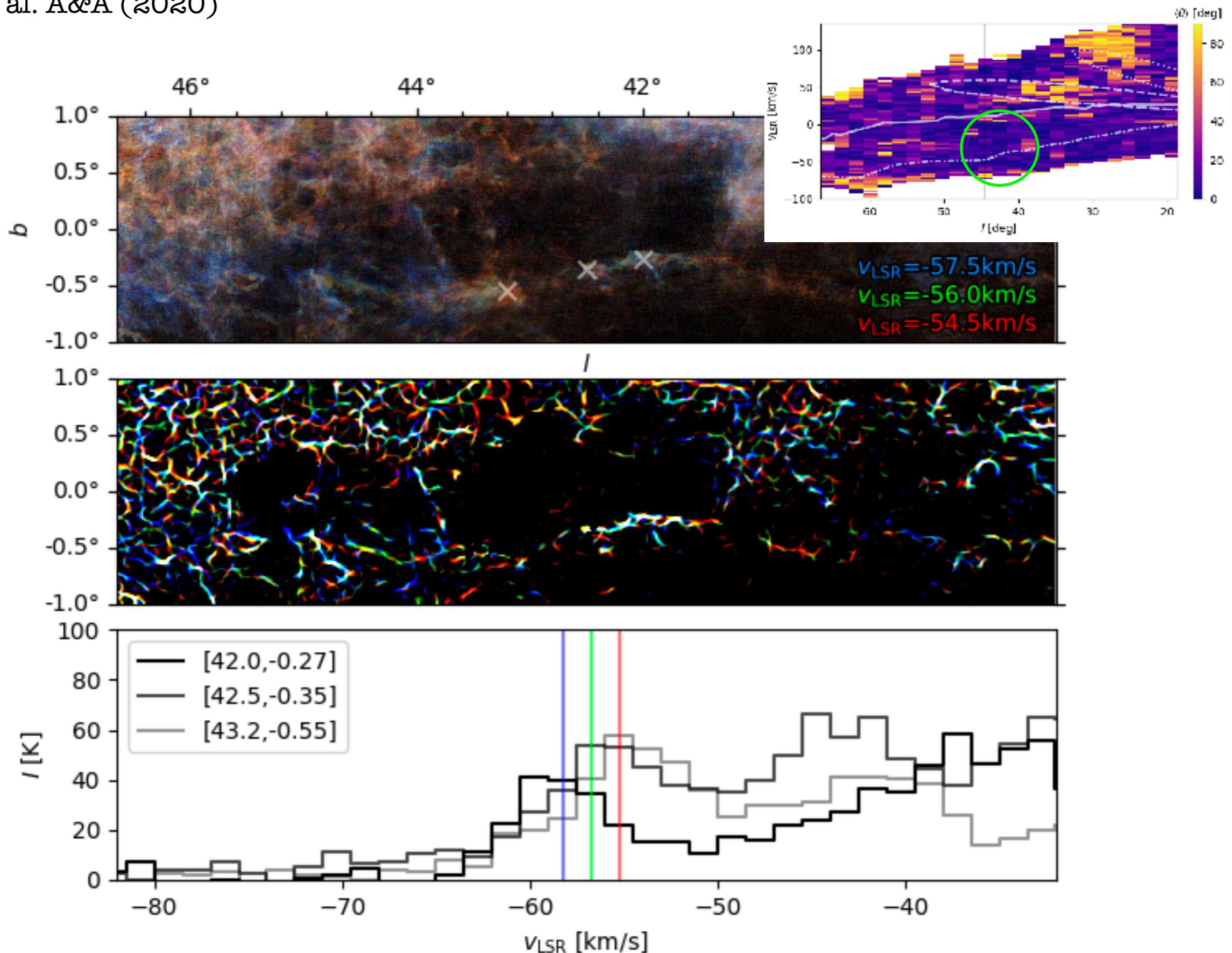
Wang, Y., et al. A&A 2020.



THOR survey area

# The Magdalena filament

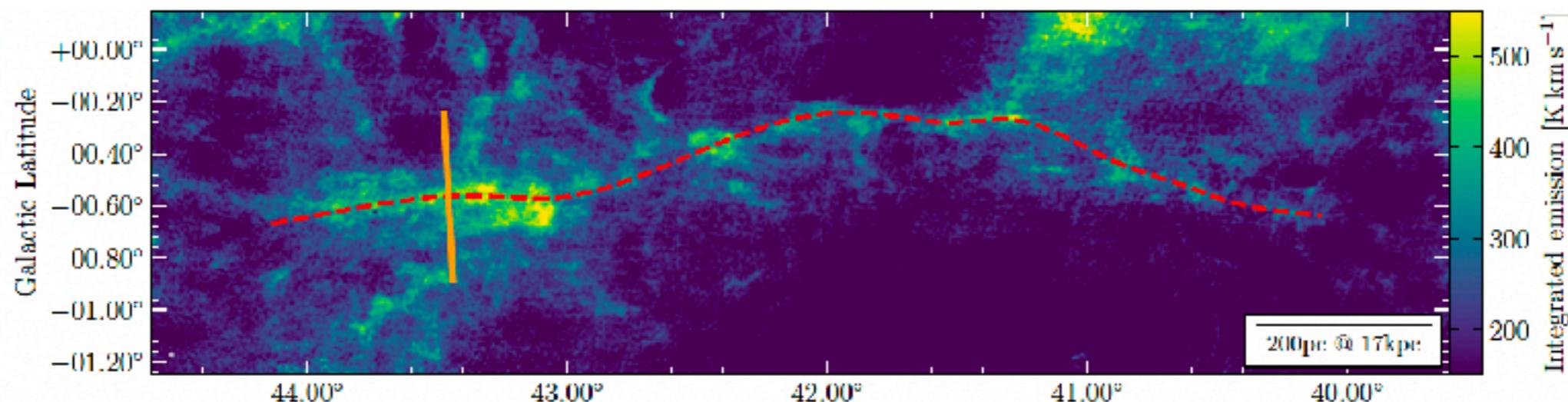
Soler, J.D. et al. A&A (2020)



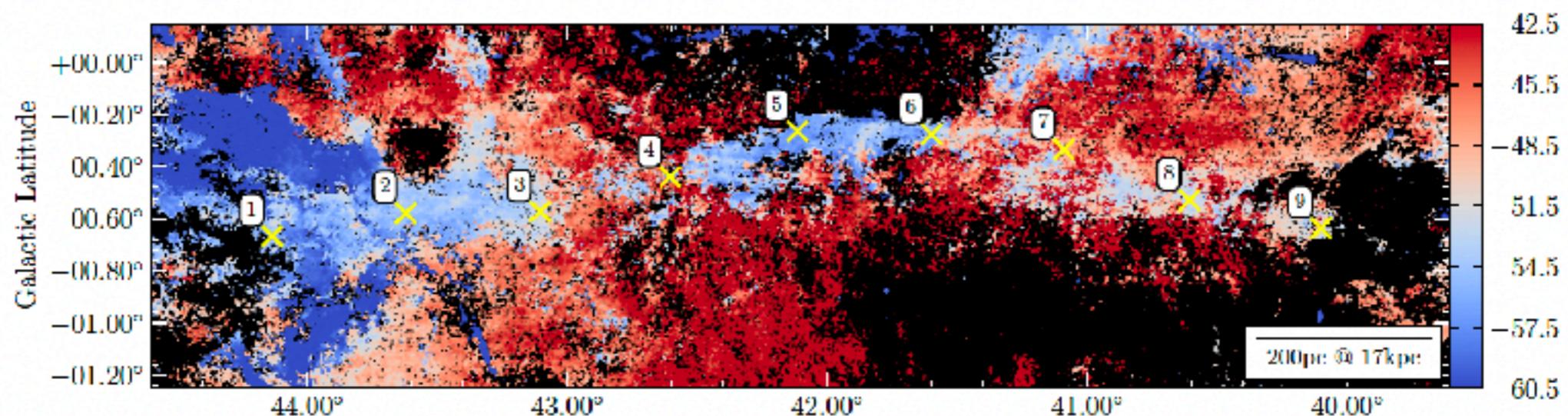
# The Magdalena filament

Syed, J., et al. (2022)

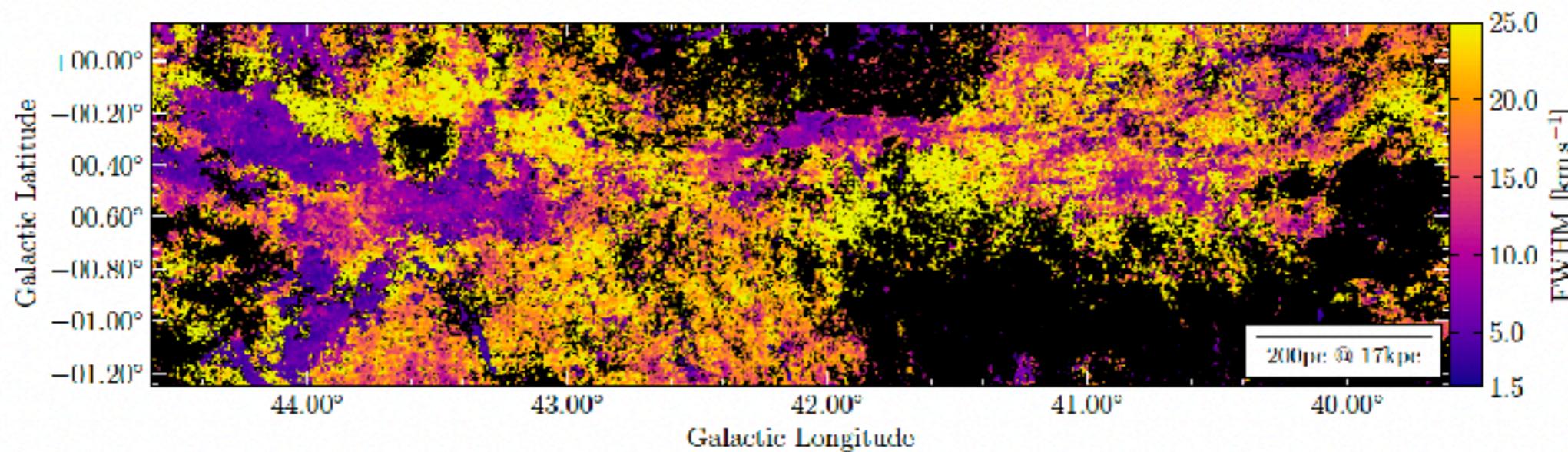
$N_{\text{H}}$



$\langle v \rangle$



$\sigma_v$



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Static properties: density, temperature, velocity\*.  
Dynamics properties: ?

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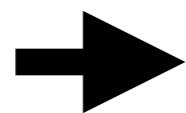
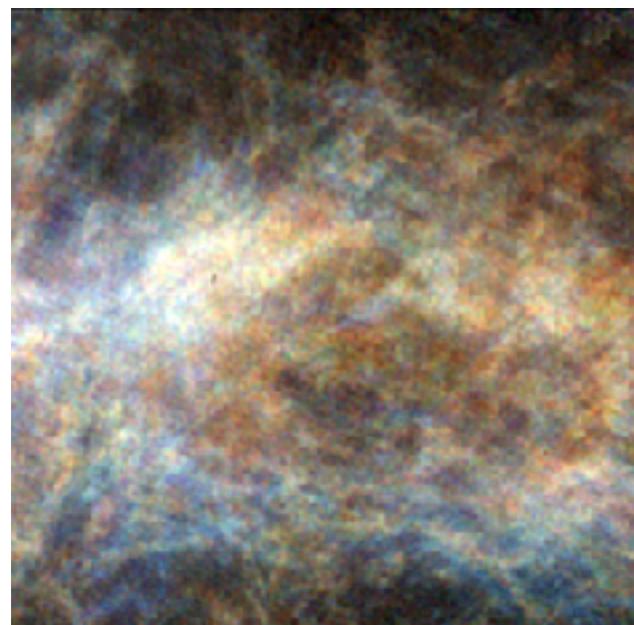
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With two eyes

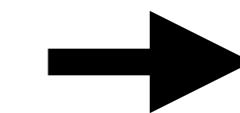
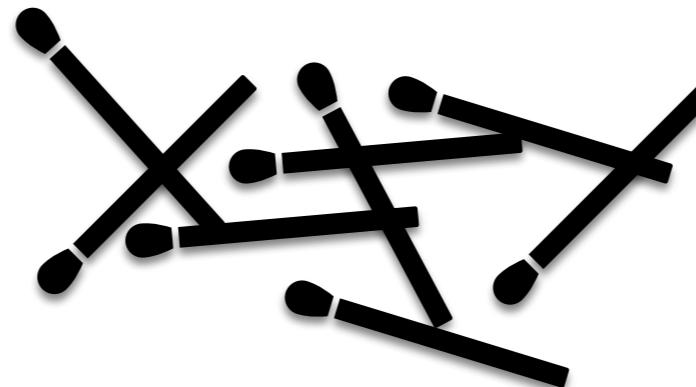
# Epistemology tab

Soler & Nicoglou. Société d'histoire et d'épistémologie des sciences. 2022

Observations



Characteristics



Knowledge



# Why sticks?

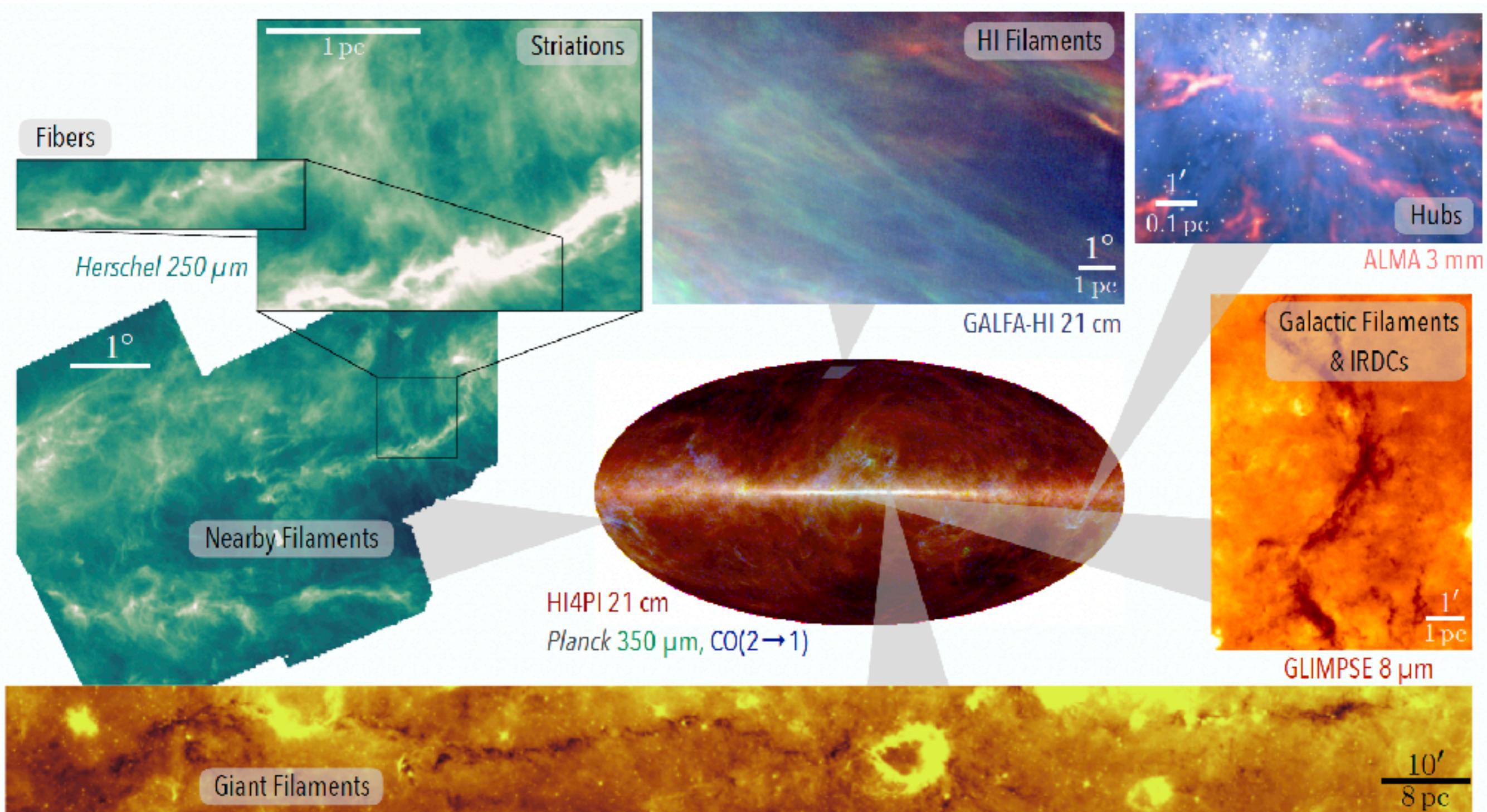
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Hacar et al. 2022. Protostars and Planets VII

“To a worm in horseradish  
the world is horseradish.”

# Why sticks?

Hacar et al. 2022. Protostars and Planets VII



# Atomic hydrogen emission

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GALFA-HI survey; Peek et al. 2018

## **Observations: The Galactic Arecibo L-band Feed Array HI (GALFA-HI) Survey**

Data set obtained with the Arecibo L-band Feed Array (ALFA) on the Arecibo 305m telescope.

Peek, J.E.G., et al. 2018 ApJS, 234:2

**Music:  
Gabriela Supelano**

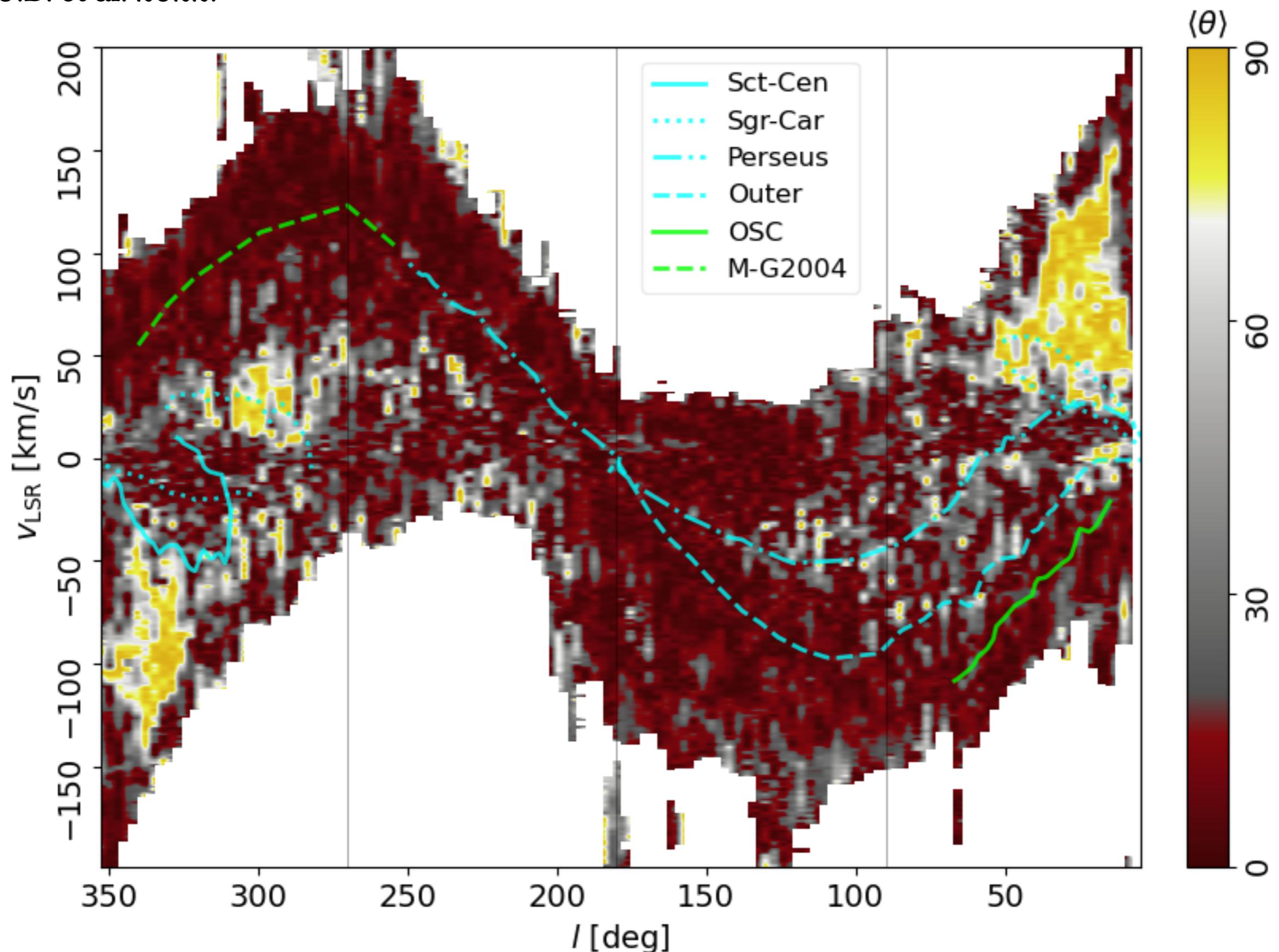
**Movie:  
Juan Diego Soler / IAPS-INAF**

## How do the Sun and other stars form?

- What is happening to the hydrogen before becoming stars?  
It looks like a bunch of sticks.
- What is important about the sticks?
  - They are features in a 2D scalar field.
  - Properties:
    - Width
    - Length
    - Orientation ←

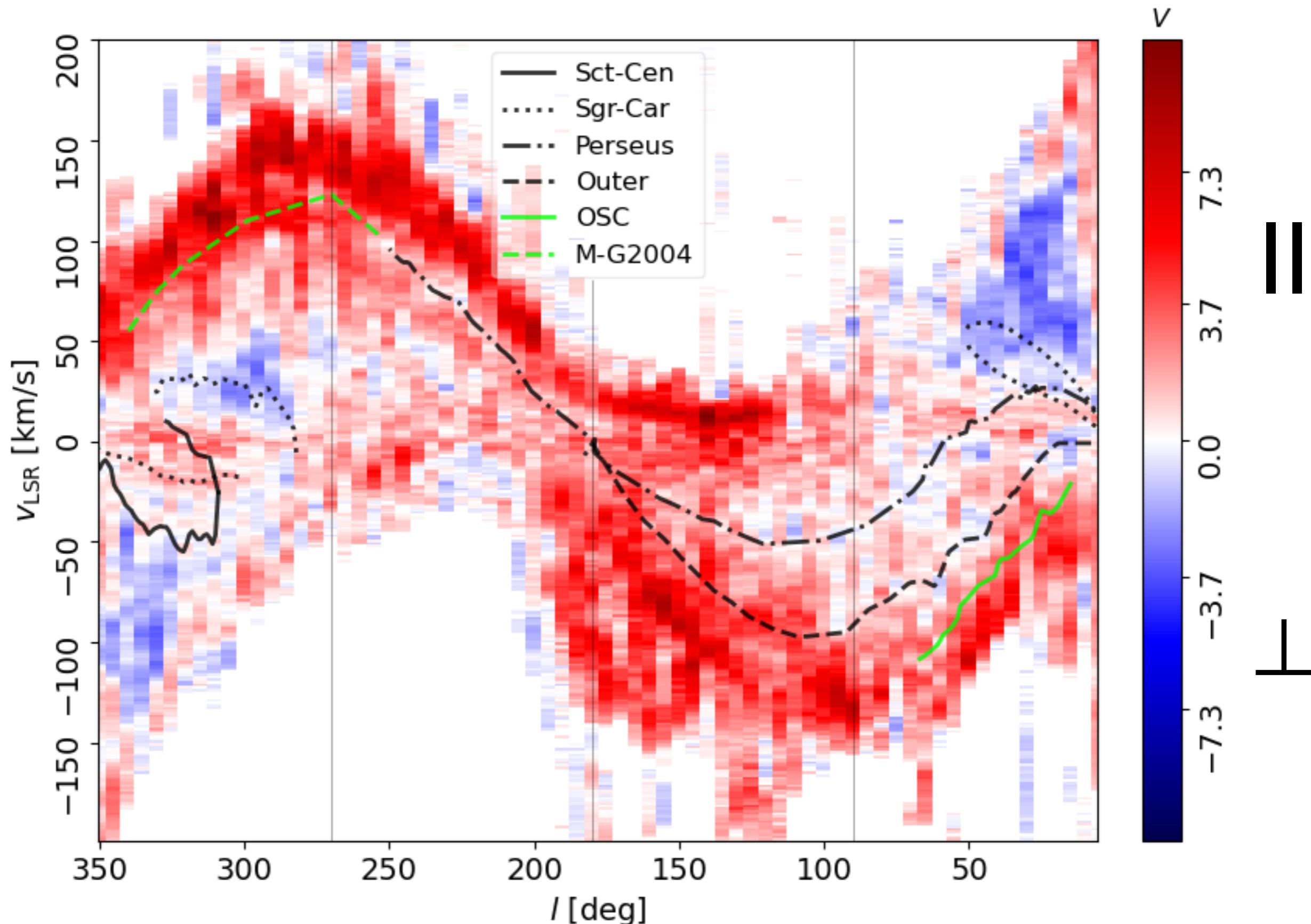
# Orientation of atomic filaments

Soler, J.D. et al. 2022.



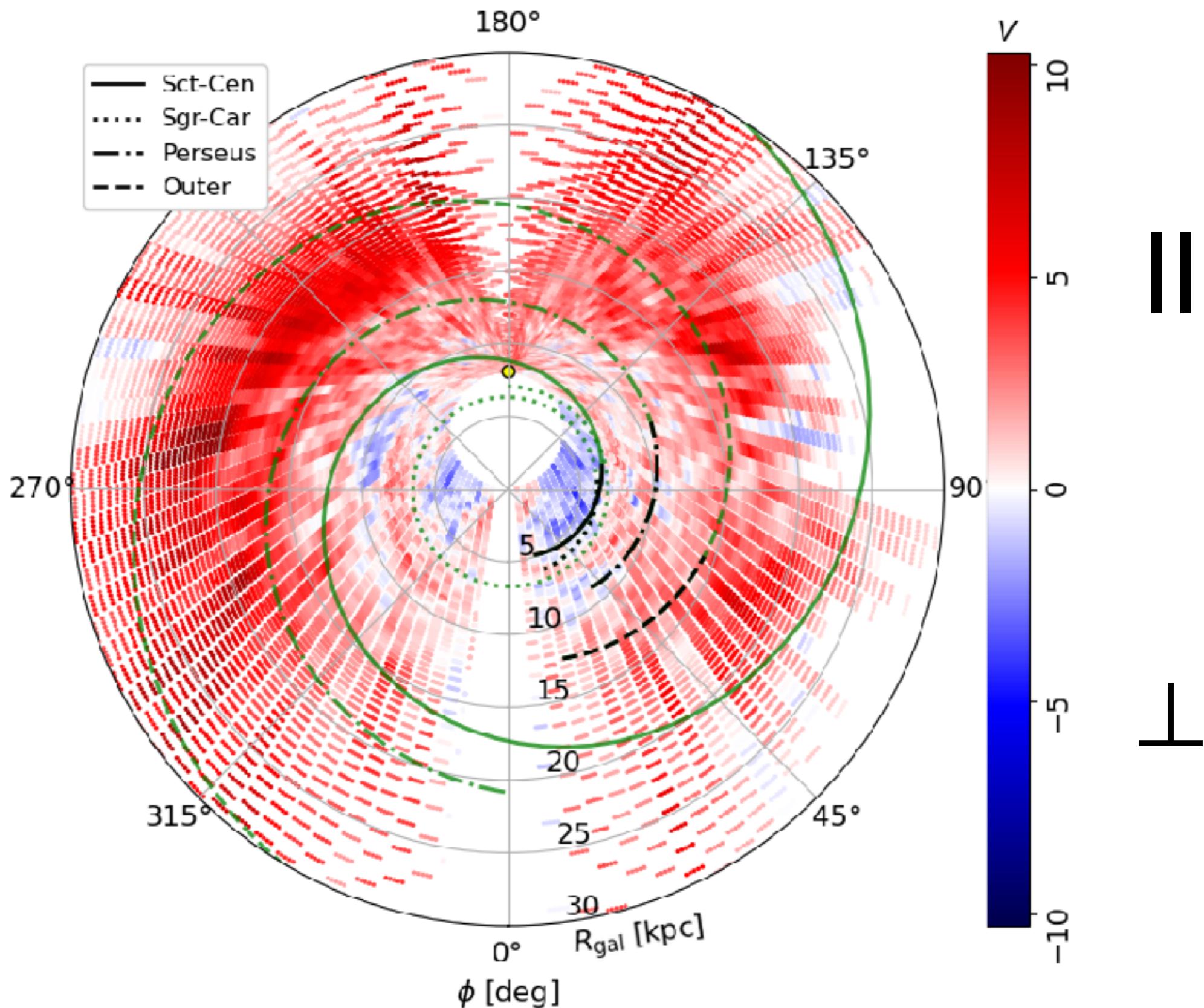
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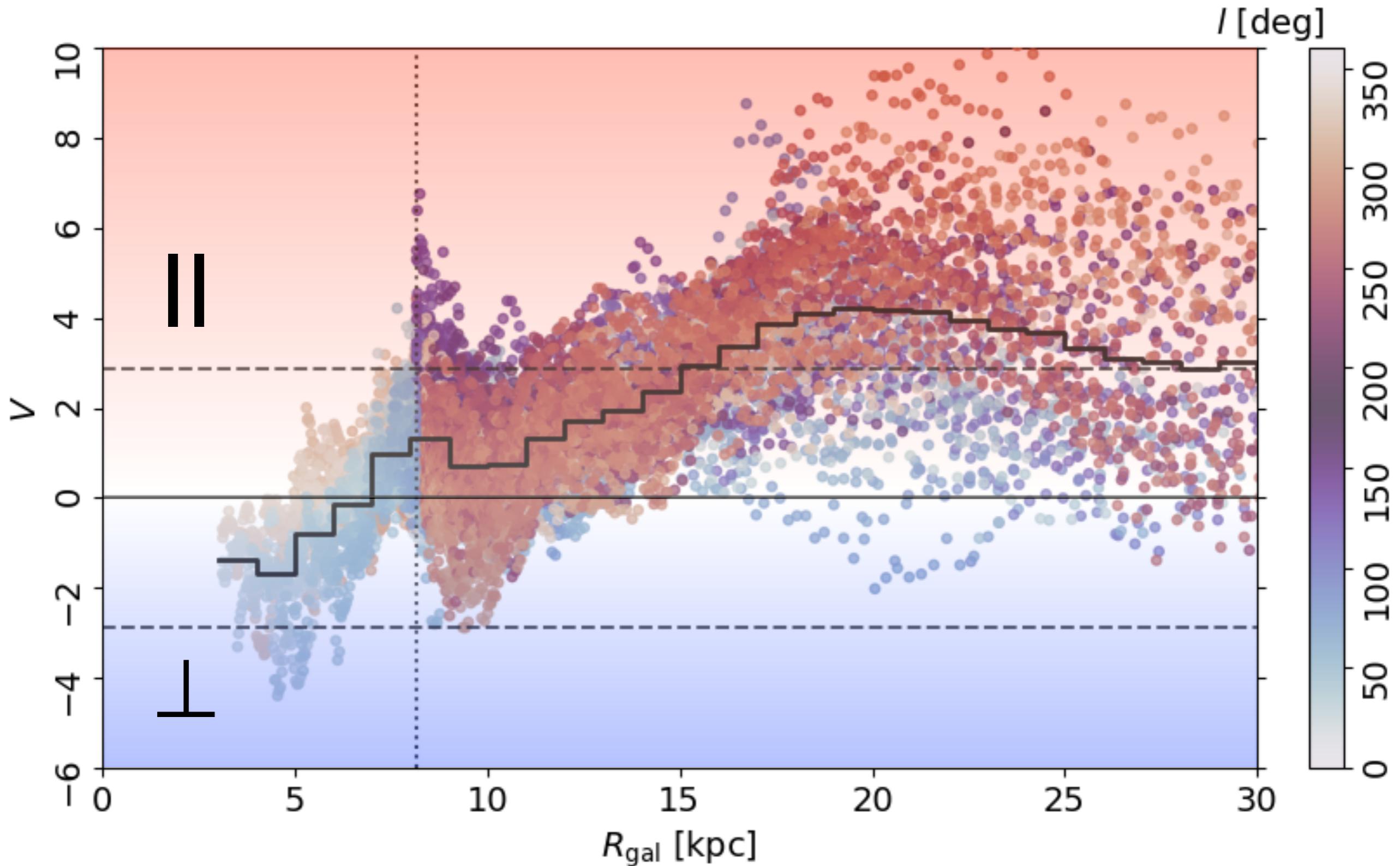
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Soler, J.D. et al. 2022. A&A in press.



# Atomic filament orientation

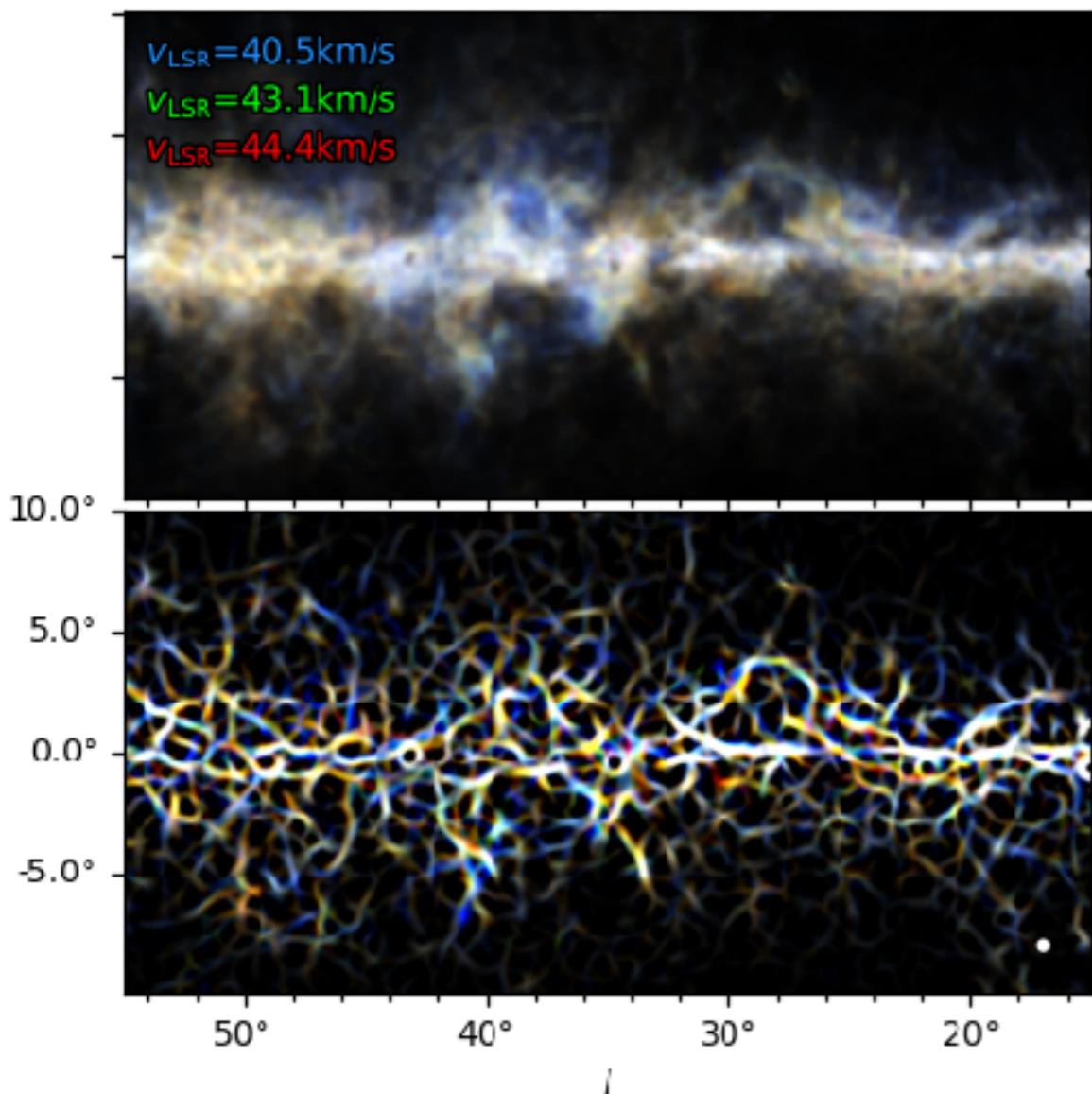
Soler, J.D. et al. 2022. A&A in press.



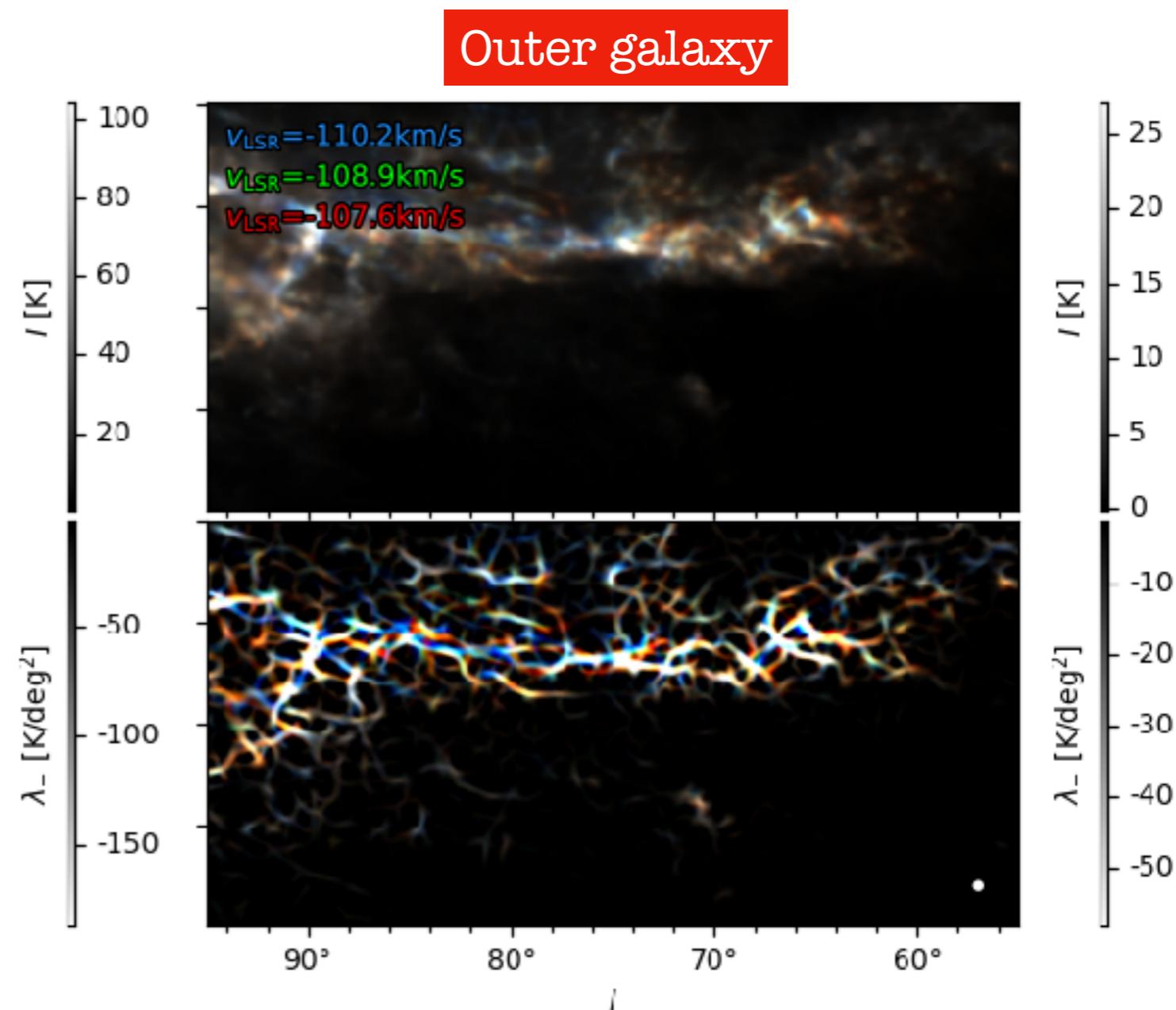
# Atomic filament orientation

Soler, J.D. et al. 2022. A&A in press.

Inner galaxy

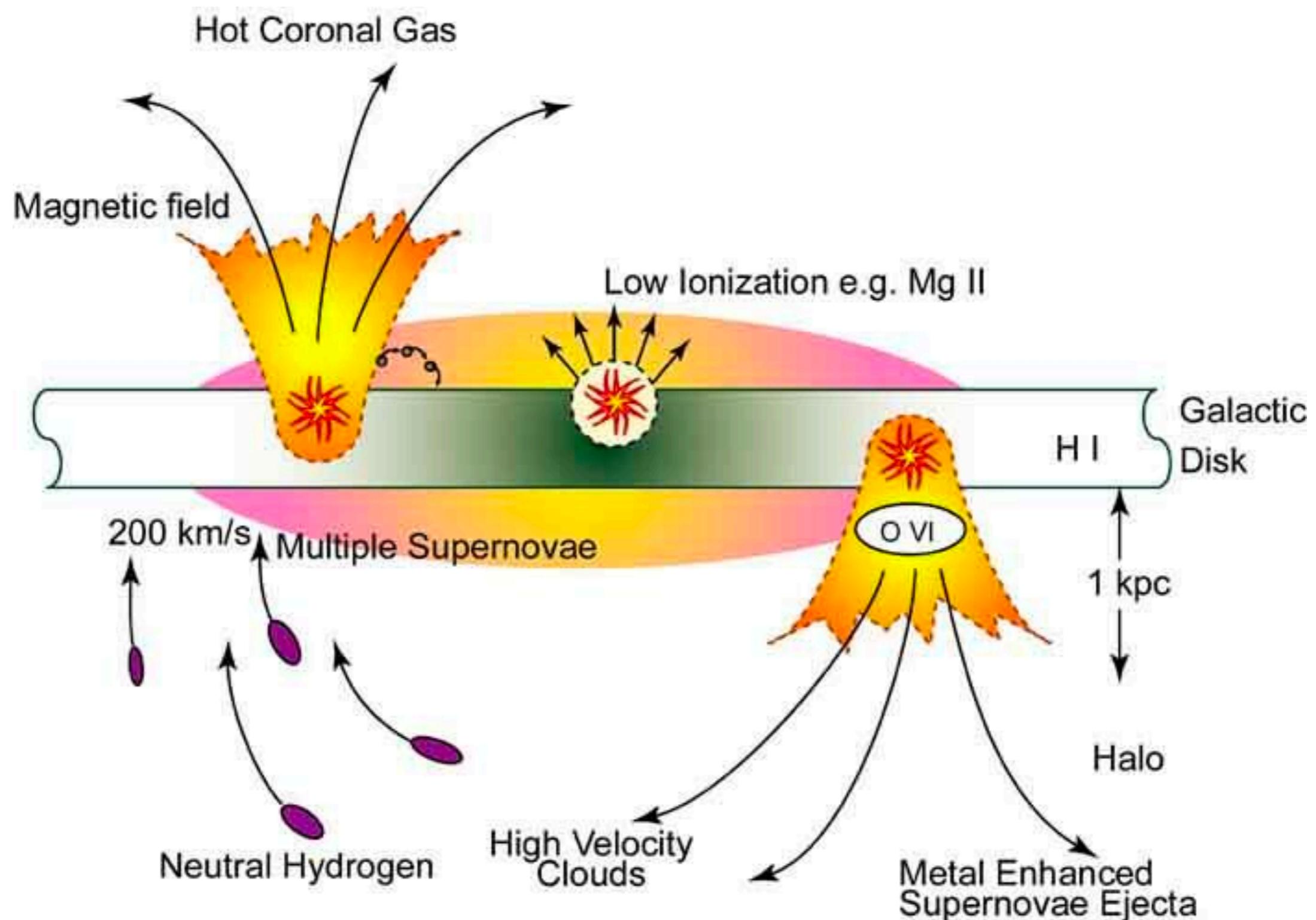


Outer galaxy



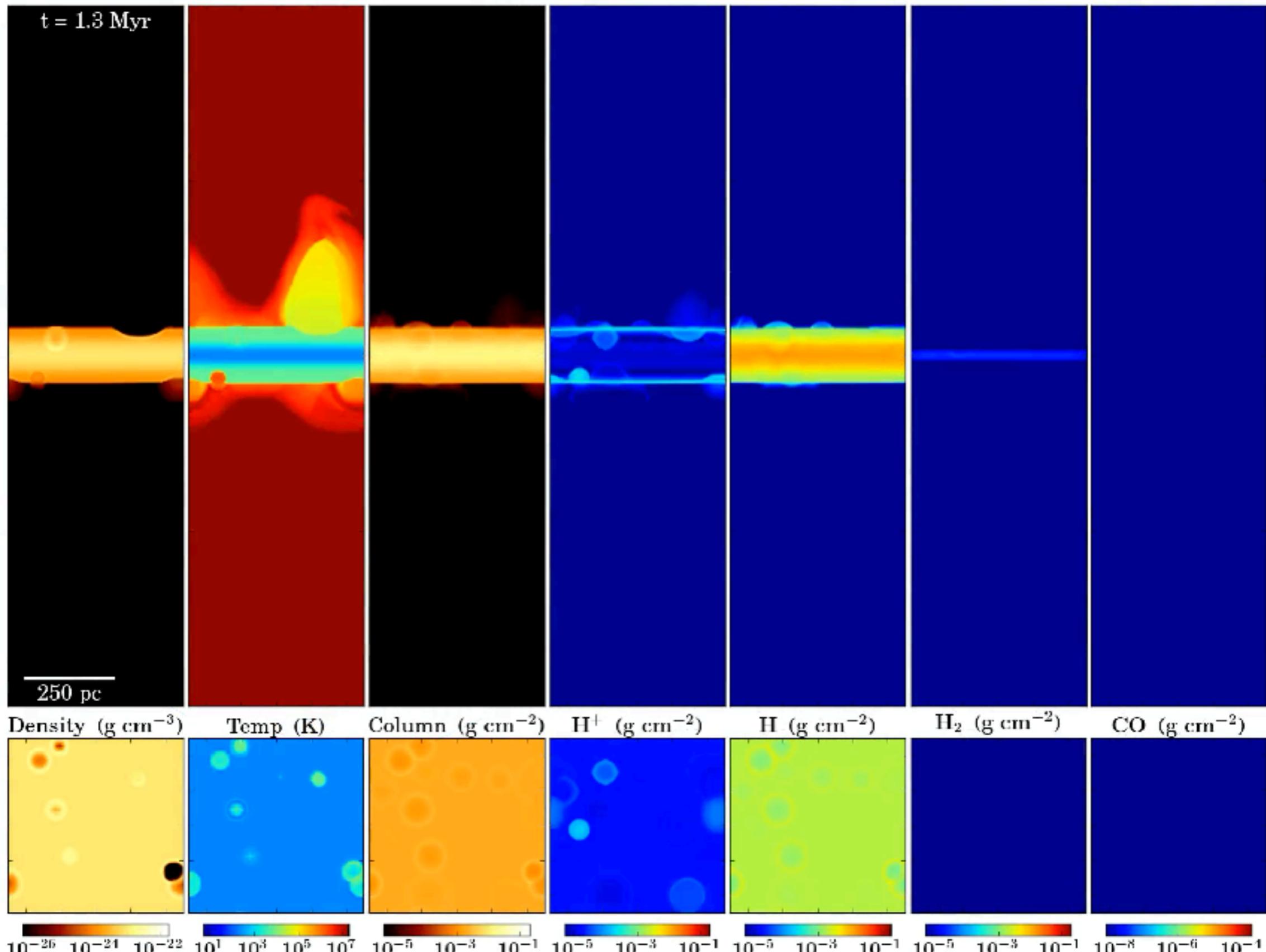
# Atomic worms and chimneys

Heiles, 1994



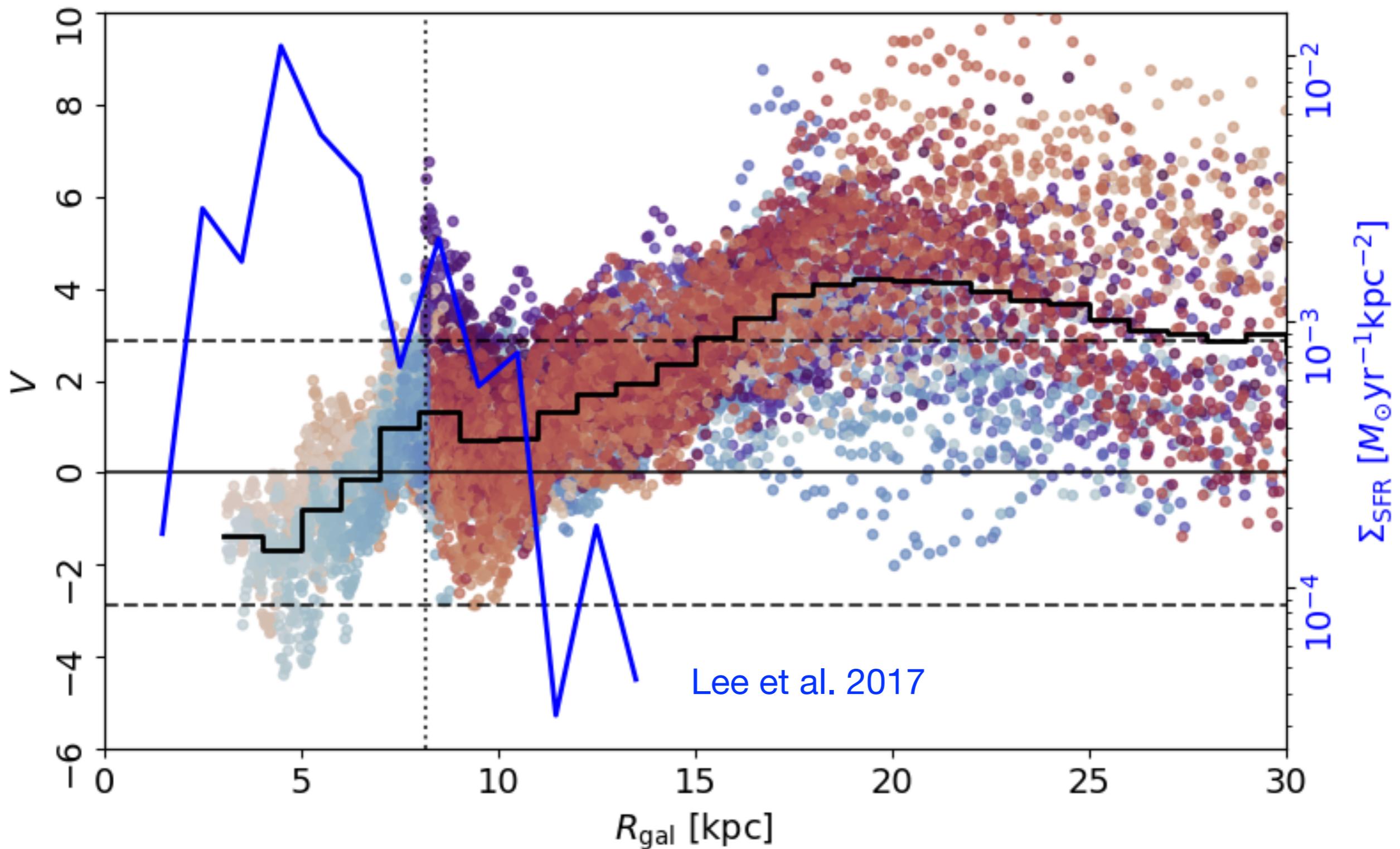
# Atomic filament orientation and HI bubbles

Girichidis et al. MNRAS 2021. SILCC: Simulating the LifeCycle of molecular Clouds



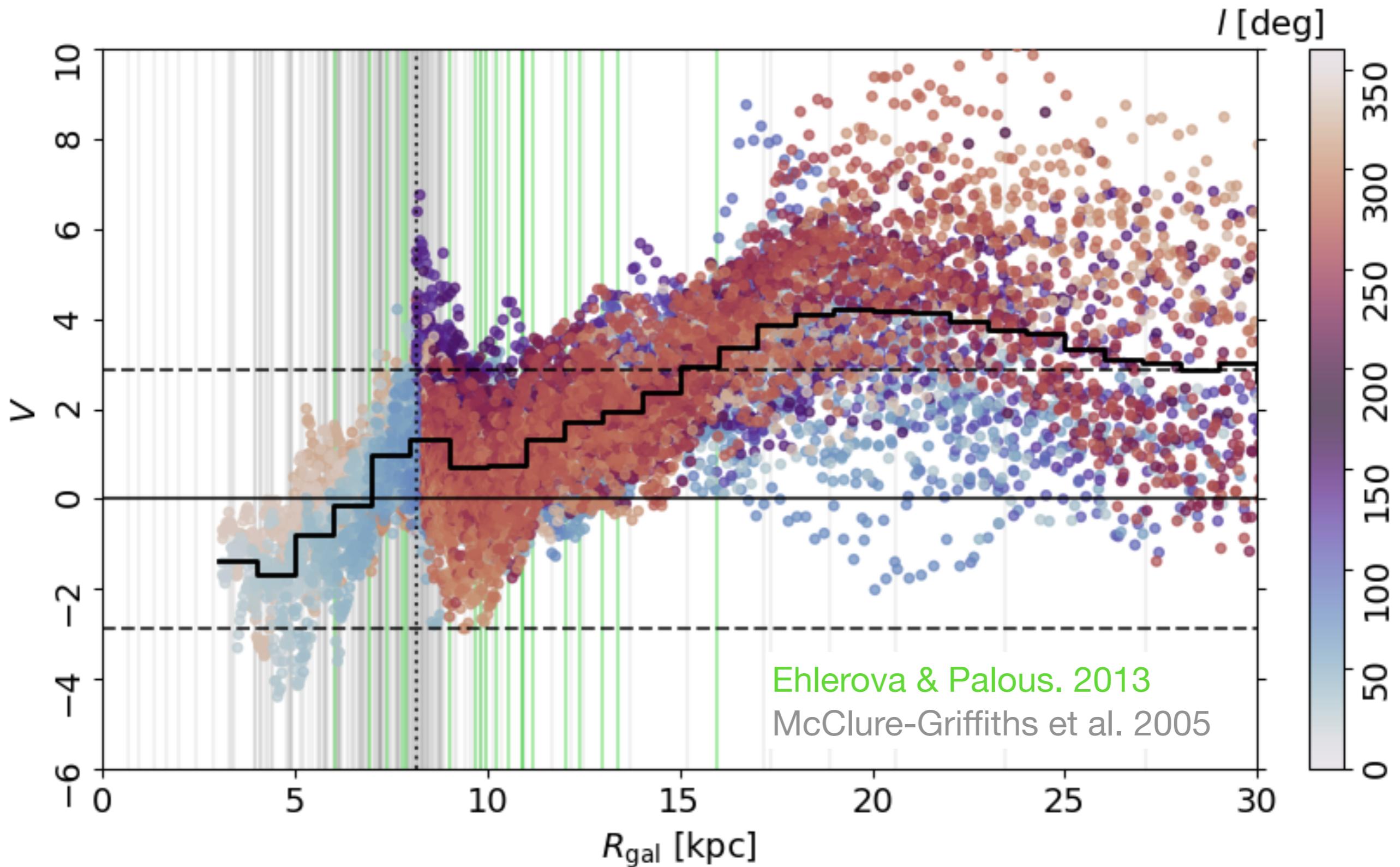
# Atomic filament orientation and star formation

Soler, J.D. et al. 2022. A&A in press.



# Atomic filament orientation and HI bubbles

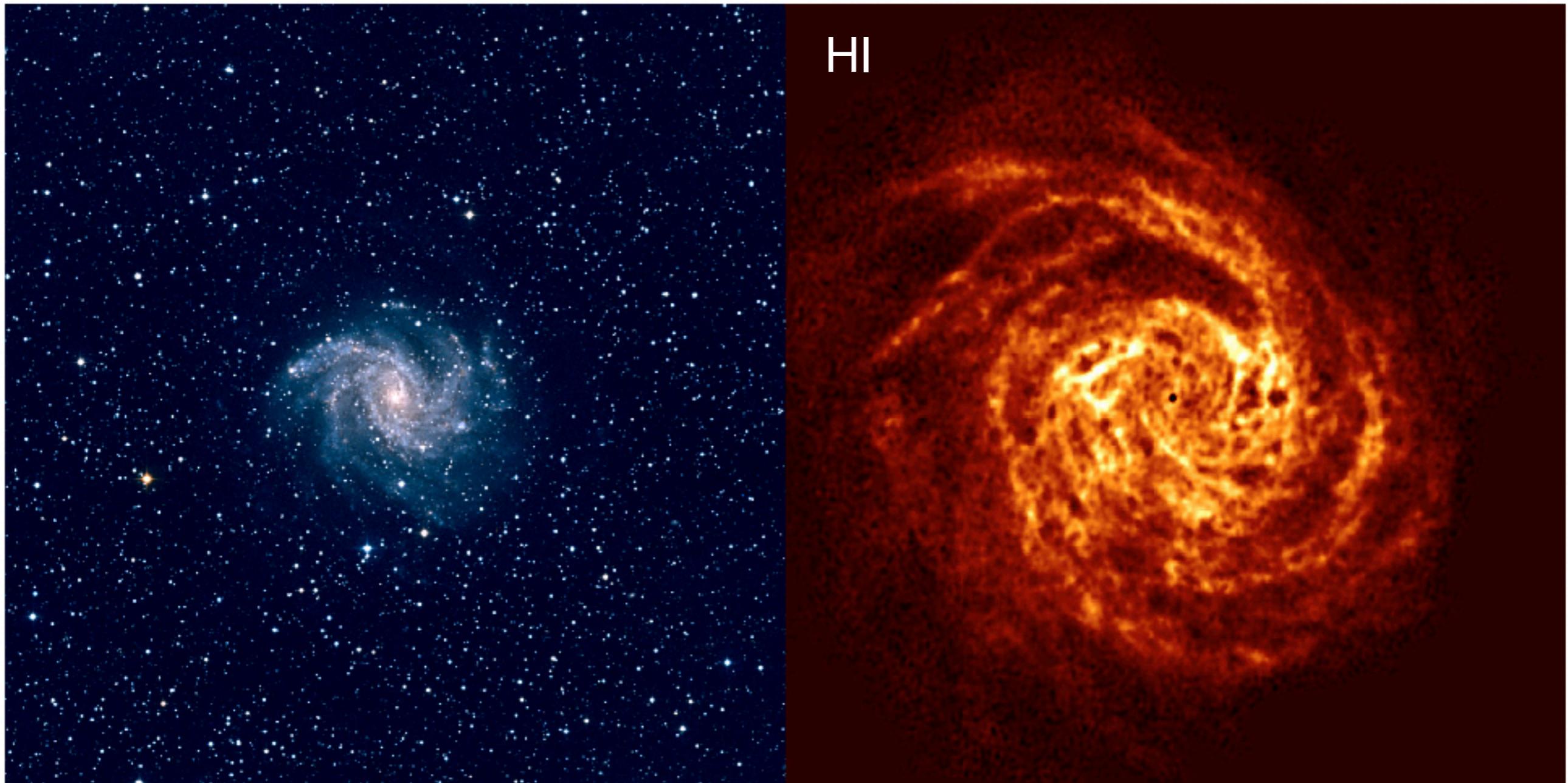
Soler, J.D. et al. 2022. A&A in press.



# HI holes in the spiral galaxy NGC6946

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Boomsma et al. A&A 2008

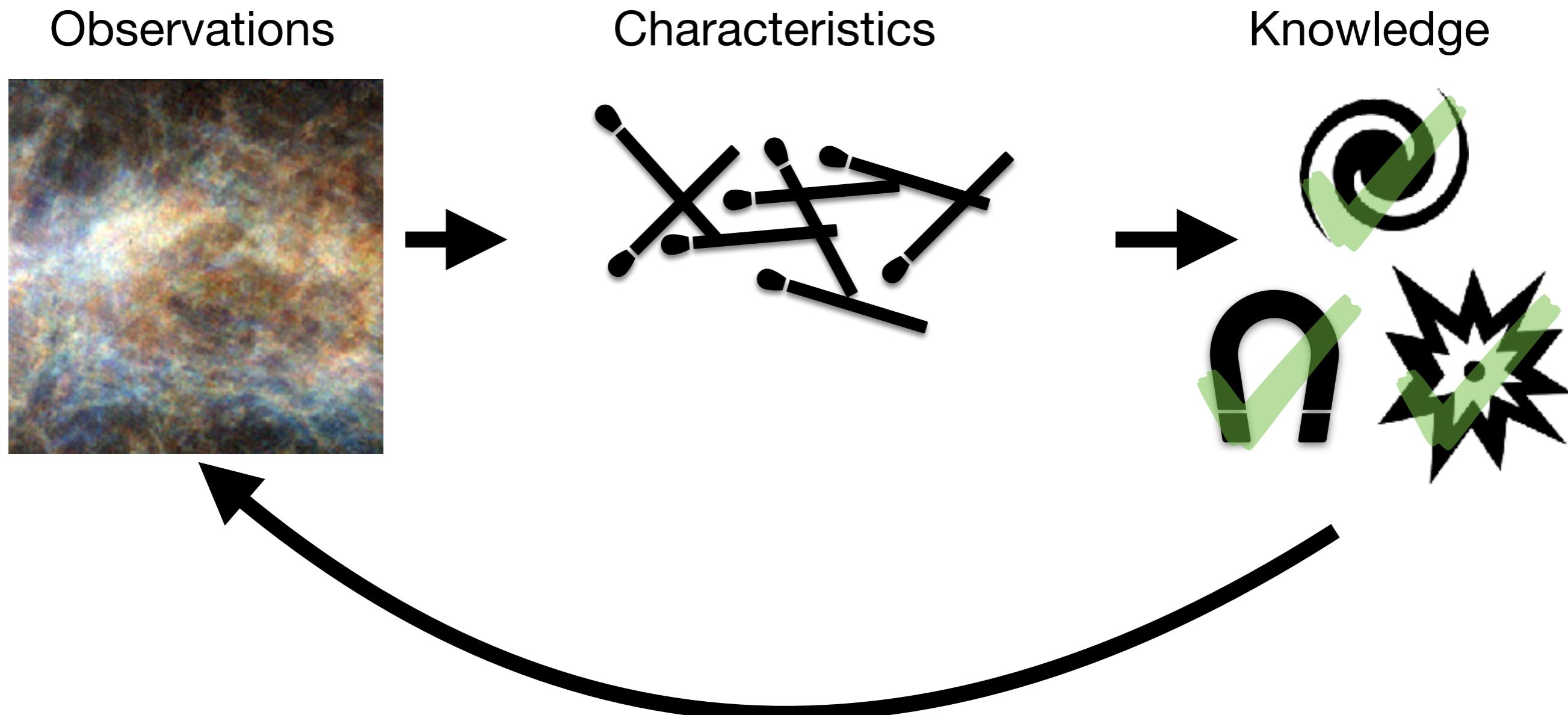


# Data-driven approach

Soler, J.D., et al. 2020. A&A

Soler, J.D., et al. 2021. A&A

Soler, J.D., et al. 2022. A&A



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Soler & Nicoglou. Société d'histoire et d'épistémologie des sciences. 2022

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What is shaping the gas?

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How do the Sun and other stars form?



What is shaping the gas?

# The Galactic dynamics revealed by HI emission



We found that the **HI filament orientation** changes from no preferential orientation to mainly parallel to the Galactic plane with increasing distance from the Galactic center.



The change in the **HI filament orientation** is most likely due to the energy and moment input from **supernova feedback**.



In general, the **HI filament orientation** is **not inherited** by the **CO filaments**, which may be the result of stellar feedback and magnetic fields.

Soler, J.D. et al. 2020. A&A  
Soler, J.D. et al. 2021. A&A  
Soler, J.D. et al. 2022. A&A

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