

- Scattering, absorption and extinction cross sections
- Vector harmonics
- scattering by a sphere.

1 Formulation of scattering problem:

Input: incident radiation, object: scatterer.

Solving: B.C. on scatterer's surface are imposed.

require continuity of tangential components of
 \vec{E} , \vec{B} field on the object's surface.

Decompose total field into superposition of different contributions:

i) Incident field, defined at each point of space
 as if the obstacle was not there.

Usually taken as a plane wave, but we can
 take it as a cylindrical or Gaussian wave.

ii) Scattered field: value of the field external to
 the object.

iii) Internal field in the obstacle.

These fields are represented as series expansion
 of suitable vector harmonics.

Mathematical tool: Vector harmonics (or wavefunctions)

Vector functions: solutions of the vector wave equation
any \vec{E} , \vec{B} field or potential can be expanded in
series of these harmonics with suitable coefficient

Particular properties of the vector harmonic
allow one to easily compute the different fields
and easily impose boundary conditions.

Using vector harmonics, one can express EM fields in the
same form independently of the scatterer shape.
It will be possible to couple the spherical wave
with cylindrical one.

2. Scattering, Extraction and absorption cross sections

Cross section: quantity with dimension of area
related with EM power scattered or absorbed by the
object interacting with the incident wave.

$$\begin{aligned} \text{Incident: } & \vec{E}_i, \vec{H}_i \\ \text{Scattered: } & \vec{E}_s, \vec{H}_s \end{aligned}$$

$$\begin{aligned} \text{Total Field outside scatterer: } & \vec{E} = \vec{E}_i + \vec{E}_s \\ & \vec{H} = \vec{H}_i + \vec{H}_s \end{aligned}$$

$$\vec{E}_i = \vec{E}_0 e^{i(\vec{k} \cdot \vec{r} - wt)}$$

$$\vec{H}_i = \vec{H}_0 e^{i(\vec{k} \cdot \vec{r} - wt)}$$

E_0, H_0 are electric and magnetic polarization vectors.

Time dependence is omitted throughout the paper (quasi-stationary)

Scattering vector of total field:

$$S = \frac{1}{2} \operatorname{Re} [\vec{E} \times \vec{H}^*]$$

$$S = \frac{1}{2} \operatorname{Re} [E \times H^*]$$

$$= \frac{1}{2} \operatorname{Re} [I(E_i + E_s) \times (H_i^* + H_s^*)]$$

$$= \frac{1}{2} \operatorname{Re} [E_i \times H_i^*] + \frac{1}{2} \operatorname{Re} [E_s \times H_s^*] +$$

$$\frac{1}{2} \operatorname{Re} [E_i \times H_s^* + E_s \times H_i^*]$$

$$S = S_i + S_s + S_e$$

S_i = pointing vector of incident wave.

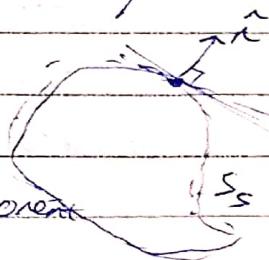
S_s = " scattered wave

S_e = " interaction of incident and scattered fields

Inside the scatterer: \vec{E}_i, \vec{H}_i

Consider the scatterer's surface S_s , the amount of energy absorbed by the object can be computed as:

$$W_a = - \int_{S_s} \hat{n} \cdot S dS$$



(Equality due to continuity of tang. component of E and H field at the surface.)

Applying Green's second vector theorem, it can be proved that this integral is equal to the integral extended to a spherical surface S , centered on the scatterer with large radius

$$W_a = - \int_S \hat{n} \cdot S dS = W_i - W_s + W_e$$

$$\text{where } W_i = - \int_S \hat{n} \cdot S_i dS, W_s = \int_S \hat{n} \cdot S_s dS, W_e = - \int_S \hat{n} \cdot S_e dS$$

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Power associated with incident wave is 0.

(Incident power is both incoming and outgoing from surface (?)).

$$W_e = W_a + W_s$$

↳ sum of energy absorbed by scatterer and of the scattered energy.

i.e.: Amount of energy subtracted from incident wave because of the interaction with scatterer.

∴ called Extraction Power (W_e)

Define cross sections:

$$\tau_s = \frac{W_s}{I_{S_i} l}$$

$$\tau_a = \frac{W_a}{I_{S_i} l}$$

$$\tau_e = \frac{W_e}{I_{S_i} l}$$

$$(dim = \text{Area})$$

$$\tau_e = \tau_a + \tau_s$$

Scattering cross section represents the amount of power scattered by object over amount of power per unit area carried by incident wave.

Coefficients: $\delta_s = \frac{\tau_s}{g}$ $\delta_a = \frac{\tau_a}{g}$ $\delta_e = \frac{\tau_e}{g}$

g is particle cross sectional area projected onto plane to incident wave.

3. Wave vector (Vector wave) Formalism

Hansen first proposed this as "to vector wave eqn".
Maxwell's Eqn:

$$\nabla \times \vec{E} + \frac{\partial \vec{B}}{\partial t} = 0$$

$$\nabla \times \vec{H} + \frac{\partial \vec{B}}{\partial t} = \vec{I}$$

Can describe any field or potential through Helmholtz Eqn. Consider any generic field \vec{C} .

$C = \vec{E}_s + \vec{H}_s \rightarrow \vec{F}_s$. At time t , then such a field in the absence of sources and in a linear, isotropic, homogeneous medium, must satisfy the general homogeneous wave Eqn:

$$\nabla^2 \vec{C} - \mu \epsilon \frac{\partial^2 \vec{C}}{\partial t^2} - \mu \epsilon \frac{\partial \vec{C}}{\partial t} = 0$$

In frequency domain: Helmholtz Eqn

$$\nabla^2 \vec{C} + k^2 \vec{C} = 0$$

$$k^2 = \omega^2 \epsilon \mu + i \nu \omega \quad (\text{omitting } e^{-i \omega t})$$

This vector Eqn can be projected along unit vectors of a generic ref system giving 3 scalar differential Eqns. However, such a system is not easy to define in most coordinate systems.

For scalar is: $\nabla^2 \psi + k^2 \psi = 0$ soln known in all coord. systems.
Define vector harmonics:

$$\vec{L} = \nabla \psi \quad M = \nabla \times (\hat{a} \psi) \quad N = \frac{1}{k} \nabla \times M$$

\hat{a} is typically a const. unit vector called pilot vector ($\nabla \times \hat{a} = 0$ in gen).

check that these satisfy vector Helmholtz Eqn
 $\operatorname{div} \mathbf{A} = 0$ and $\nabla \cdot \mathbf{M} = \nabla \times \mathbf{C} \hat{\alpha} \psi$
 $\therefore \nabla \cdot \mathbf{M} = 0$, using vector identities;

$$\nabla^2 \mathbf{M} + k^2 \mathbf{M} = \nabla \times \left[\hat{\alpha} (\nabla^2 \psi + k^2 \psi) \right] = 0$$

$\therefore \mathbf{M}$ satisfies vector wave equation if ψ is soln to scalar wave eqn $\nabla^2 \psi + k^2 \psi = 0$
 Similarly for \mathbf{N} and \mathbf{L} .

Properties: $\mathbf{M} = \frac{1}{k} \nabla \times \mathbf{N}$ and $\nabla \cdot \mathbf{L} = \nabla^2 \psi = -k^2 \psi$

$$\nabla \times \mathbf{L} = 0 \quad \nabla \cdot \mathbf{M} = 0 \quad \nabla \cdot \mathbf{N} = 0$$

\mathbf{L} is irrotational, \mathbf{M} and \mathbf{N} are solenoidal.

$\therefore \mathbf{EM}$ field in the absence of sources can be expressed only by a superposition of \mathbf{M} and \mathbf{N} , while \mathbf{L} must necessarily be taken into account only when EM source

generic soln ψ to Helmholtz Eqn is a set of solutions ψ_n (eigenfunctions that) form a vector space basis L^2 of square integrable funcns (Hilbert space)

Set of vectors \mathbf{L}_n , \mathbf{M}_n and \mathbf{N}_n are associated with this set of scalar solutions.

Example: Consider magnetic potential \vec{A} to be soln of vector Helmholtz then:

$$\vec{A} = \frac{1}{i\omega} \sum_{n=0}^{\infty} (a_n \mathbf{M}_n + b_n \mathbf{N}_n + c_n \mathbf{L}_n)$$

Applying Maxwell's eqns, $\mathbf{E} = \sum_{n=0}^{\infty} (a_n \mathbf{M}_n + b_n \mathbf{N}_n)$

$$\mathbf{H} = \frac{k}{i\omega \mu} \sum_{n=0}^{\infty} (a_n \mathbf{N}_n + b_n \mathbf{M}_n)$$

In cartesian coordinates, solⁿ of scalar Helmholtz Eqⁿ:

(unitary amplitudes):

$$\psi_{RS} = e^{i\vec{k} \cdot \vec{r}}$$

Then

$$(24) \quad L = i\psi \vec{k} \quad M = i\psi \vec{k} \times \hat{a} \quad N = \frac{i\psi (kx_0)}{k} \times k$$

These allow plane-wave expansions of a generic field

$$\psi_{RS} = \iint_{-\infty}^{\infty} g(kn, by) e^{ik \cdot r} dx dy$$

$$(25) \quad \psi_{RS} = \iint g(\alpha, \beta) e^{ik \cdot r} d\alpha d\beta \quad | \begin{array}{l} \text{d angles of} \\ \text{k with axes} \end{array}$$

$g(\alpha, \beta)$: plane wave angular spectrum

Insert (25) in (24):

$$\vec{L} = i \iint g(\alpha, \beta) \vec{k}(\alpha, \beta) e^{ik \cdot r} d\alpha d\beta$$

$$\vec{M} = i \iint g(\alpha, \beta) \vec{k}(\alpha, \beta) \times \hat{a} e^{ik \cdot r} d\beta d\alpha$$

$$\vec{N} = \frac{1}{k} \iint g(\alpha, \beta) [k(\alpha, \beta) \times \hat{a}] \times k(\alpha, \beta) e^{ik \cdot r} d\beta d\alpha$$

5.

Scattering by a spherical object: 3D

Step 1. determine the expression of spherical harmonics
(spherical vector wavefunctions) M_{mnS} and N_{mnS}

Scalar Helmholtz Eqⁿ in spherical coordinates:

$$\frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \psi}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial \psi}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 \psi}{\partial \phi^2} + k^2 \psi = 0$$

Well known solution:

$$\psi_{RS} = A z_n(kr) P_m^m(\cos \theta) e^{im\phi}$$

$$\psi_{RS} = \pm z_n(kr) J_n^m(\cos\theta) e^{im\phi}$$

\pm is complex constant. $z_n(kr)$ represents a spherical Bessel function of 1st, 2nd, third or fourth kind.

$$z_n(kr) = \{ j_n(kr), g_n(kr), h_n^{(1)}(kr), h_n^{(2)}(kr) \}$$

$J_n^m(\cos\theta)$ is the associated Legendre function.

It is not easy to recognize a pilot vector \hat{a} to insert $(\hat{r}, \hat{\theta}, \hat{\phi})$ are not constant. We need \hat{a} to be $\nabla \times \hat{a} = 0$: $\hat{r} \hat{r}$ is radial \therefore curl is 0.

The vector harmonic : $H_{mn}(\theta, \phi) = \nabla \times (\hat{r} \hat{r} Y_{mn})$ is a solution of vector Helmholtz equation.

$$H_{mn}(\theta, \phi) = \frac{i_m}{\sin\theta} z_n(kr) J_n^m(\cos\theta) e^{im\phi} \hat{\theta}$$

$$- z_n(kr) \frac{\partial J_n^m(\cos\theta)}{\partial \theta} e^{im\phi} \hat{\phi}$$

$$N_{mn}(\theta, \phi) = \frac{z_n(kr)}{kr} n(n+1) J_n^m(\cos\theta) e^{im\phi} \hat{r}$$

$$+ \frac{1}{kr} \frac{\partial (r z_n(kr))}{\partial r} \frac{\partial J_n^m(\cos\theta)}{\partial \theta} e^{im\phi} \hat{\theta}$$

$$+ \frac{1}{kr} \frac{\partial (r z_n(kr))}{\partial r} \frac{i_m}{\sin\theta} J_n^m(\cos\theta) e^{im\phi} \hat{\phi}$$

These can be simplified by introducing two scalar functions: scalar Tesselal functions, related to associated Legendre.

$$T_{mn}(\theta) = \frac{n}{i_m} \frac{J_n^m(\cos\theta)}{\sin\theta}$$

$$Z_{mn}(\theta) = \frac{\partial J_n^m(\cos\theta)}{\partial \theta}$$

$$\text{Then } M_{mn} = Z_n(kr) \sum i \pi m (\cos \theta - \sin \theta \cos \phi) \hat{e}^{im\phi}$$

$$N_{mn} = \left\{ n! r^n + \frac{1}{kr} \frac{d}{dr} [Z_n(kr) \sum i \pi m (\cos \theta - \sin \theta \cos \phi)] \right\} e^{im\phi}$$

Introduce three vector functions:

vector	$S_{mn}(\theta, \phi) = e^{im\phi} \sum i \pi m (\cos \theta \hat{\theta} - \sin \theta \cos \phi \hat{\phi})$
Tesselal	$n_{mn}(\theta, \phi) = e^{im\phi} \sum i \pi m (\cos \theta \hat{\theta} + i \sin \theta \cos \phi \hat{\phi})$
functions	$f_{mn}(\theta, \phi) = e^{im\phi} n! r^n + \frac{1}{r} \frac{d}{dr} [Z_n(kr) \sum i \pi m (\cos \theta \hat{\theta} + i \sin \theta \cos \phi \hat{\phi})]$

These functions obey orthogonality relations.

They are independent of the particular spherical tessell function involved in vector harmonics.

$$\text{Final form: } M_{mn}(r, \theta, \phi) = Z_n(kr) S_{mn}(\theta, \phi)$$

$$N_{mn}(r, \theta, \phi) = \frac{Z_n(kr)}{r} f_{mn}(\theta, \phi) + \frac{i}{r} \frac{d}{dr} [Z_n(kr) I_{mn}(\theta, \phi)]$$

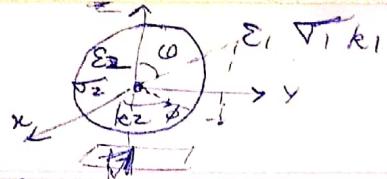
Start with a dielectric sphere radius a , in a linear homogeneous isotropic medium with dielectric permittivity ϵ_r , wave number k_1 . Consider elliptically polarized plane wave incident on the sphere.

$$E_i(r) = E_{pol} e^{ik_1 \cdot r} = (E_{\theta i} \hat{\theta}_i + E_{\phi i} \hat{\phi}_i) e^{ik_1 \cdot r}$$

$\hat{\theta}_i, \hat{\phi}_i$ are the unit vectors of the local spherical coordinate frame w.r.t. wave vectors of the plane wave.

$$k_i = k_1 \hat{k}_i = k_1 (\sin \theta_i \cos \phi_i \hat{x} + \sin \theta_i \sin \phi_i \hat{y} + \cos \theta_i \hat{z})$$

$$\hat{\phi}_i = \frac{\hat{z} \times \hat{k}_i}{|\hat{z} \times \hat{k}_i|} = -\sin \phi_i \hat{x} + \cos \phi_i \hat{y}$$



$$\mathbf{E}_i = \hat{\varphi} \times \hat{k}_i = \cos \theta_i \cos \varphi_i \hat{x} + \cos \theta_i \sin \varphi_i \hat{y} - \sin \theta_i \hat{z}$$

Incident plane wave can be expanded in spherical harmonics as follows:

$$E_i(r_s) = \sum_{n=1}^{\infty} \sum_{m=-n}^n [a_{mn} J_{mn}(kr_s) + b_{mn} N_{mn}(kr_s)]$$

(45) Boundary condition: tangential component continuous

$$a_{mn} = E_1 S^{(n)} \frac{2n+1}{n(n+1)} \underbrace{J_{mn}(r_i)}_{(r_i+m)!} \cdot \text{epol} \cdot n_{mn}^*(\theta_i, \varphi_i)$$

$$b_{mn} = E_1 S^{(n)} \frac{2n+1}{n(n+1)} \underbrace{J_{mn}(r_i)}_{(r_i-m)!} \cdot \text{epol} \cdot n_{mn}^*(\theta_i, \varphi_i)$$

$J^{(1)}$ indicated spherical Bessel funcⁿ of 1st kind
typical for stationary waves

Scattered field can be expanded in spherical vector wave functions as:

$$E_S(r_s) = \sum_{n=1}^{\infty} \sum_{m=-n}^n [c_{mn} M_{mn}(kr_s) + d_{mn} N_{mn}^{(3)}(kr_s)]$$

(3) in vector harmonics : Bessel of 3rd kind ie. spherical Hankel func of 1st type: typical for progressive waves.

c, d are unknowns to be obtained by B.C.

$$(\mathbf{E}_i + E_S) \times \hat{r} = 0 \quad \text{for } r = a$$

make explicit:

$$\mathbf{E}_i = \sum_{n=1}^{\infty} \sum_{m=-n}^n [a_{mn} j_{mn}(\theta_i, \varphi_i) \sin(k_r r_s) + b_{mn} n_{mn}(\theta_i, \varphi_i) j_{mn}'(k_r r_s)]$$

$$E_S = \sum_{n=1}^{\infty} \sum_{m=-n}^n [c_{mn} m_{mn}(\theta_i, \varphi_i) h_n^{(1)}(k_r r_s) + d_{mn} n_{mn}(\theta_i, \varphi_i) h_n^{(3)}(k_r r_s)]$$

Then $a_{mn}(\theta, \phi) \times \hat{r} = -n a_{mn}(\theta, \phi)$
using

$$a_{mn}(\theta, \phi) \times \hat{r} = n a_{mn}(\theta, \phi)$$

$$a_{mn}(\theta, \phi) \times \hat{r} = 0$$

substitute in E.C.:

$$\sum_{n=1}^{\infty} \sum_{m=-n}^{n} [a_{mn}(\theta, \phi) \sum_{k} b_{mk} j_n(k) a_k + d_{mn} h_n^{(1)}(k) a_k] \\ - n a_{mn}(\theta, \phi) \sum_{k} a_{mk} j_n(k) a_k + c_{mn} h_n^{(1)}(k) a_k = 0$$

Using orthogonality of vector spherical harmonics
two scalar equations can be obtained:
is taken dot product with $n^* m^* \sin\theta$ and integrating
in θ between $[0, \pi]$ and $[0, 2\pi]$ resp.
and latter by $n^* m^* \sin\theta$ with same integration.

As a result:

$$a_{mn} j_n(k) a_k + c_{mn} h_n^{(1)}(k) a_k = 0$$

$$b_{mn} j_n(k) a_k + d_{mn} h_n^{(1)}(k) a_k = 0$$

obtaining two simple expressions for scattering coeffs:

$$c_{mn} = -a_{mn} \frac{j_n(k) a_k}{h_n^{(1)}(k) a_k} \quad \left. \begin{array}{l} \text{mic scattering} \\ \text{coefficients} \end{array} \right\} \text{for a}$$

$$d_{mn} = -b_{mn} \frac{j_n(k) a_k}{h_n^{(1)}(k) a_k} \quad \left. \begin{array}{l} \text{AEQ sphere} \end{array} \right\}$$

Scattering by Dielectric Sphere

$\epsilon_2 > \epsilon_1$. Consider internal field to the sphere:

$$E_p(\vec{r}) = \sum_{n=1}^{\infty} \sum_{m=-n}^n [\epsilon_{mn} \hat{m}_{ns}^{(1)} + d_{mn} \hat{d}_{ns}^{(1)}]$$

B.C.: $(E_i + E_s - E_p) \times \hat{x} = 0 \text{ for } r=a \quad (1)$

$E \nabla \times (E_i + E_s - E_p) \times \hat{x} = 0 \text{ for } r=a \quad (2)$

Substituting: two vector equations are:

$$\begin{aligned} \sum_{n=1}^{\infty} \sum_{m=-n}^n \epsilon_{mn} (\partial_r \hat{e}_s) \hat{x} [b_{mn} j_n(k_r a) + d_{mn} h_n^{(1)}(k_r a) - f_{mn} j_n(k_r a)] \\ - n_{mn} (\partial_r \hat{e}_s) [a_{mn} j_n(k_r a) + c_{mn} h_n^{(1)}(k_r a) - e_{mn} j_n(k_r a)] = 0 \end{aligned}$$

and $\begin{aligned} \sum_{n=1}^{\infty} \sum_{m=-n}^n \{ \epsilon_{mn} (\partial_r \hat{e}_s) \hat{x} [k_1 a_{mn} j_n(k_r a) + k_1 c_{mn} h_n^{(1)}(k_r a) - k_1 e_{mn} j_n(k_r a)] \\ - n_{mn} (\partial_r \hat{e}_s) \hat{x} [k_1 b_{mn} j_n(k_r a) + k_1 d_{mn} h_n^{(1)}(k_r a) - k_1 f_{mn} j_n(k_r a)] \} = 0 \end{aligned}$

Orthogonal: take dot product with a_{mn}^* , c_{mn}^* , e_{mn}^* , b_{mn}^* , d_{mn}^* , f_{mn}^* and get same

$$a_{mn} j_n(k_r a) + c_{mn} h_n^{(1)}(k_r a) - e_{mn} j_n(k_r a) = 0$$

$$b_{mn} j_n(k_r a) + d_{mn} h_n^{(1)}(k_r a) - f_{mn} j_n(k_r a) = 0$$

$$k_1 a_{mn} j_n(k_r a) + k_1 c_{mn} h_n^{(1)}(k_r a) - k_1 e_{mn} j_n(k_r a) = 0$$

$$k_1 b_{mn} j_n(k_r a) + k_1 d_{mn} h_n^{(1)}(k_r a) - k_1 f_{mn} j_n(k_r a) = 0$$

Introduce $\chi = \frac{k_2}{k_1}$ (dielectric contrast) and solving

$$c_{mn} = -a_{mn} j_n(k_r a) j_n(k_r a) - \chi j_n(k_r a) j_n(k_r a)$$

$$f_{mn}^{(1)}(k_r a) j_n(k_r a) - \chi h_n^{(1)}(k_r a) j_n(k_r a)$$

and $d_{mn} = -b_{mn} j_n(k_r a) j_n(k_r a) - \chi j_n(k_r a) j_n(k_r a)$

$\chi^{(1)}(k_r a) j_n(k_r a) \chi h_n^{(1)}(k_r a) j_n(k_r a)$
- Mie scattering coefficients for dielectric sphere

Compute cross sections as function of scattering coefficients:

$$cs = \frac{2\pi}{k_1^2} \sum_{n=1}^{\infty} (2n+1) \left(\frac{|c_{in}|^2}{a_{in}} + \frac{|d_{in}|^2}{b_{in}} \right)$$