## Mathematical Physics (Individual Contest)

**Prob. 1** The Hellmann-Feynman theorem: Given a Hamiltonian H which has discrete energy levels and smoothly depends on a coupling parameter  $\gamma$ , define  $\mathcal{A} = \frac{\partial H}{\partial \gamma}$ .

- For every eigenstate  $|\Psi_j\rangle$  of H with its energy eigenvalue  $E_j$ , prove that  $\langle \Psi_j | \mathcal{A} | \Psi_j \rangle = \frac{\partial E_j}{\partial \gamma}$ .
- Taking as an example the one-dimensional quantum oscillator described by the Hamiltonian  $H=-\frac{\partial^2}{\partial x^2}+\frac{\omega^2 x^2}{4}, \quad \omega \in R \quad \text{and } \omega>0$ , show that

$$\langle \Psi_j | x^2 | \Psi_j \rangle = \frac{(2j+1)}{\omega}.$$

**Prob. 2** You are given the weak-field Newtonian limit of space-time as

$$ds^{2} = -(1+2\phi)dt^{2} + dx^{2} + dy^{2} + dz^{2},$$
(1)

and the Newtonian gravitational potential  $\phi$  satisfying  $\vec{\nabla}^2 \phi = 4\pi G \rho$  (i.e.,  $\vec{\nabla}^2 g_{00} = -8\pi G \rho$ ) with  $\vec{\nabla}^2 = \partial_x^2 + \partial_y^2 + \partial_z^2$ , G the Newton constant and  $\rho$  the matter density. Based on this information and the principle of general covariance, derive the complete Einstein field equation. You are given

$$R^{\rho}_{\mu\sigma\nu} = \partial_{\sigma}\Gamma^{\rho}_{\mu\nu} - \partial_{\nu}\Gamma^{\rho}_{\sigma\mu} + \Gamma\Gamma - \text{terms}$$

$$\Gamma^{\rho}_{\mu\nu} = \frac{1}{2}g^{\rho\sigma} \left(\partial_{\mu}g_{\sigma\nu} + \partial_{\nu}g_{\mu\sigma} - \partial_{\sigma}g_{\mu\nu}\right). \tag{2}$$

Prob.3 Consider a 2-form field in 6-dimensional spacetime

$$B_{(2)} = \frac{1}{2} B_{\mu\nu} dx^{\mu} \wedge dx^{\nu}, \qquad \mu, \nu = 0, 1, \dots, 5.$$
 (3)

The free theory action is

$$S = -\int d^6x \left( \frac{1}{12} H^{\mu\nu\rho} H_{\mu\nu\rho} + \frac{1}{4} m^2 B^{\mu\nu} B_{\mu\nu} \right), \tag{4}$$

where H is the corresponding 3-form field strength associated with the 2-form  $B_{(2)}$ 

$$H_{(3)} = \frac{1}{6} H_{\mu\nu\rho} dx^{\mu} \wedge dx^{\nu} \wedge dx^{\rho} = \frac{1}{2} \partial_{[\mu} B_{\nu\rho]} dx^{\mu} \wedge dx^{\nu} \wedge dx^{\rho} = dB_{(2)}.$$
 (5)

- (a)In the massive case, count the on-shell propagating degrees of freedom and describe the corresponding representation of Poincaré group.
- (b) In the massless case, count the physical on-shell propagating degrees of freedom by eliminating the gauge redundancies, and describe the corresponding representation of Poincaré group.