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AA 279 A – Space Mechanics

Lecture 2: Notes

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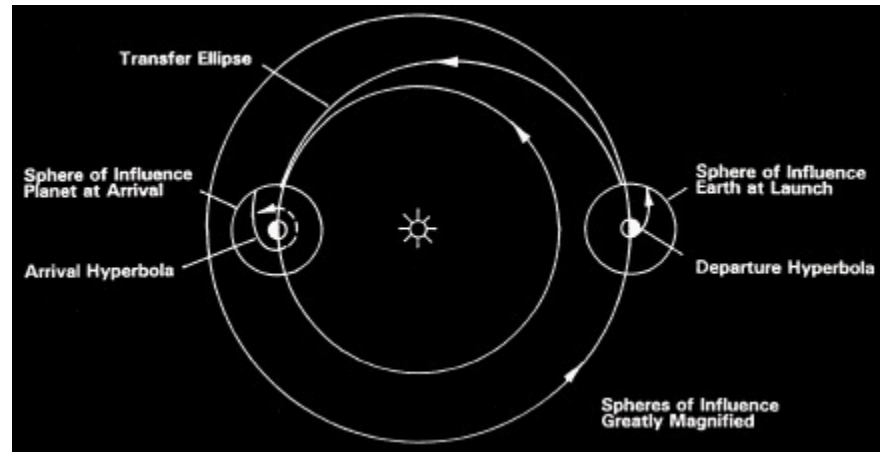
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- Specific mechanical energy
- Specific angular momentum
- Properties of conic sections



- Reading for next week
 - Bate 1.7-1.10, 4.1-4.3, 4.6.4
 - Montenbruck 2.2.1-2.2.2
 - Vallado 1.3.5-1.3.6, 2.2-2.2.3, 2.2.5, 2.3

Gravitational Potential Energy (1)

- The gravity force field is conservative, i.e. its work done along a closed path is zero. This allows describing the force of gravity as the gradient of a scalar potential

Radially symmetric mass distribution

$$\vec{F}_{\text{grav}} = -\vec{\nabla}V(r)$$

$$V(r) = \int_{r_{\text{ref}}}^r -F_{\text{grav}} dr =$$

$$= \int_{r_{\text{ref}}}^r - \underbrace{\frac{m\mu}{r^2}}_{\text{Newton IV.}} dr$$

- The gravitational potential represents the work required to raise a mass m from an arbitrary reference to a radius r in the gravity field

- In astrodynamics the arbitrary constant of the gravitational potential is chosen to be $C = 0$ ($r_{\text{ref}} = \infty$) such that $V < 0$ and $V(\infty) = 0$

$$V(r) = \left[-\frac{m\mu}{r} \right]_{r_{\text{ref}}}^r =$$

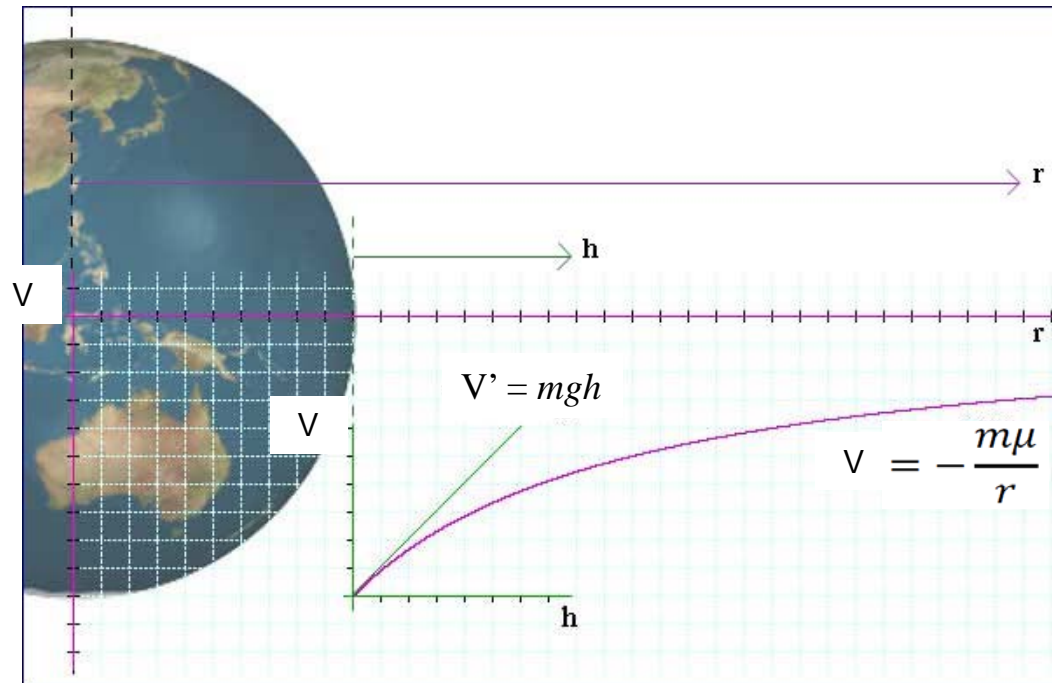
$$= -\frac{m\mu}{r} + \frac{m\mu}{r_{\text{ref}}}$$

Arbitrary constant C

Gravitational Potential Energy (2)

- $V(r)$ is negative everywhere and increases with r
- $1/r$ potential gives $1/r^2$ field force
- If we “reset” reference so $V = 0$ at surface of central body, the slope of the potential becomes $V' = mgh$

$$V = -\frac{m\mu}{r}$$



Specific Mechanical Energy

- The specific mechanical energy is defined as kinetic plus potential energy per unit mass

$$\mathcal{E} = \frac{1}{m} \left(\frac{1}{2} m v^2 - \frac{m\mu}{r} \right)$$

Inertial
velocity

v

$$\mathcal{E} = \frac{v^2}{2} - \frac{\mu}{r}$$

**Specific Mechanical
Energy in Gravity Field**

- Remember that potential energy concept is only valid for conservative force fields (e.g., gravity, spring, Coulomb)

- Since gravity is conservative, an object moving in this field does not lose or gain mechanical energy, hence $\mathcal{E} = \text{constant}$

→ This provides a
relationship between
v and r magnitudes

Specific Angular Momentum

- The specific angular momentum is defined as the linear angular momentum per unit mass (with respect to the central body)

$$\vec{h} = \frac{1}{m} (\vec{r} \times m \vec{v})$$

Inertial velocity
Momentum

$$\vec{h} = \vec{r} \times \vec{v}$$

Specific Angular Momentum

- The derivative with respect to time of the specific angular momentum equals

$$\frac{d\vec{h}}{dt} = \frac{d\vec{r}}{dt} \times \vec{v} + \vec{r} \times \frac{d\vec{v}}{dt} = \vec{r} \times \vec{a}$$

0

- Since gravity pulls towards central body (point mass), the gravity force \mathbf{F}_g is anti-parallel to the relative position vector \mathbf{r} , thus no torques are generated ($\mathbf{r} \times \mathbf{a} = 0$), hence $\mathbf{h} = \text{constant}$

→ This provides a relationship between \mathbf{v} and \mathbf{r} directions

Two Connections for All Conic Sections

- Angular momentum and semi-parameter completely determine each other

$$h = \sqrt{\mu p} ; p = \frac{h^2}{\mu}$$

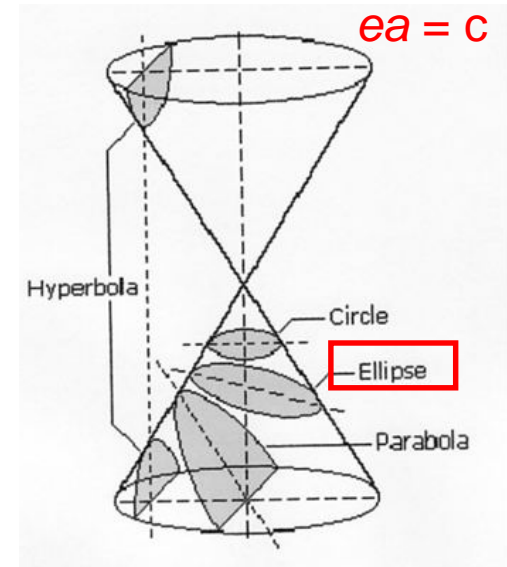
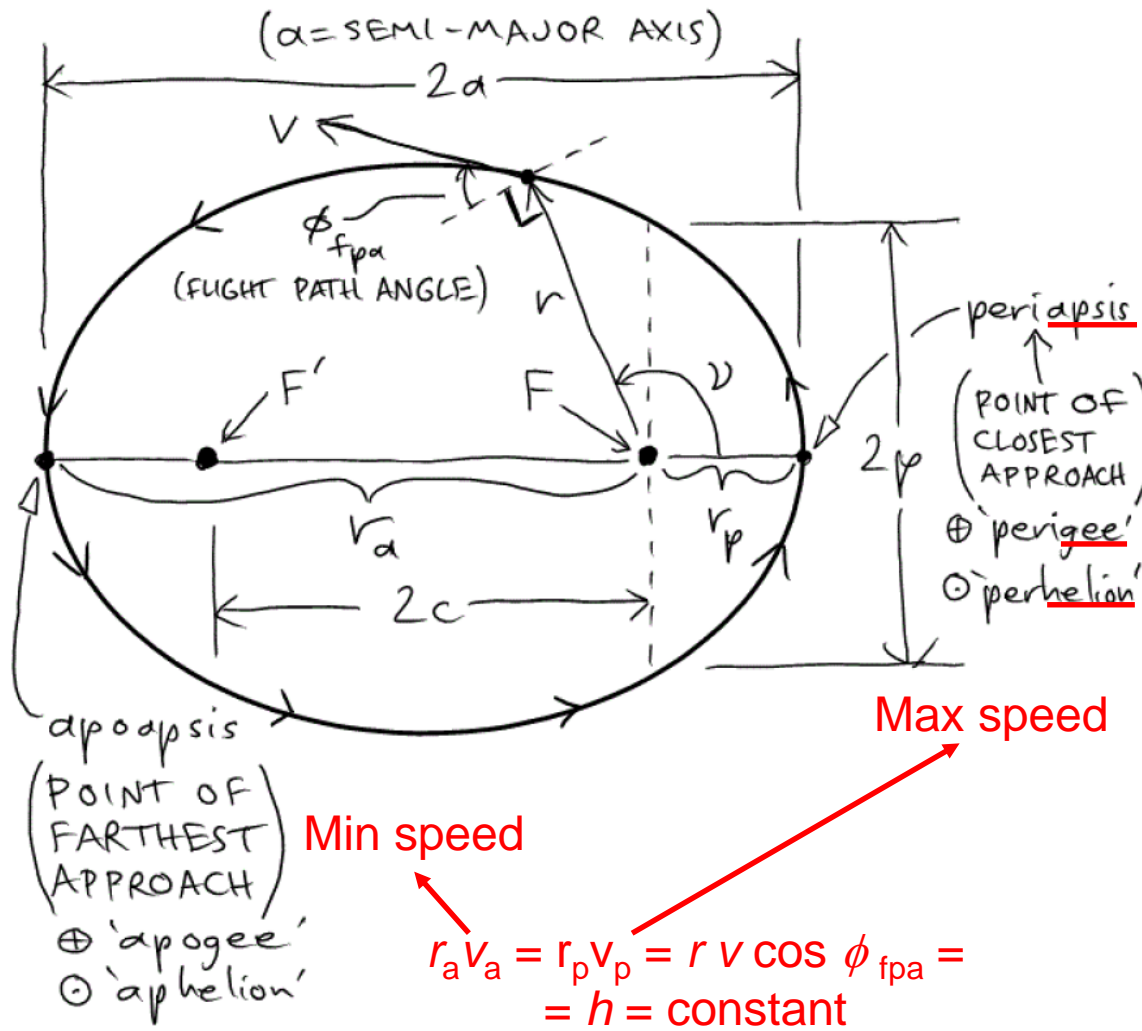
- Mechanical energy and semi-major axis completely determine each other, so that energy is solely determined by orbit “size”

$$\mathcal{E} = -\frac{\mu}{2a} ; a = -\frac{\mu}{2\mathcal{E}}$$

Semi-major axis
of conic section

- Various combinations of only two quantities (e.g., h & \mathcal{E} , p & a , h & a , p & \mathcal{E} , etc.) can tell a lot about the orbit shape. Simple energy calculations can be used to derive orbit shape...

Conic Sections: Ellipse

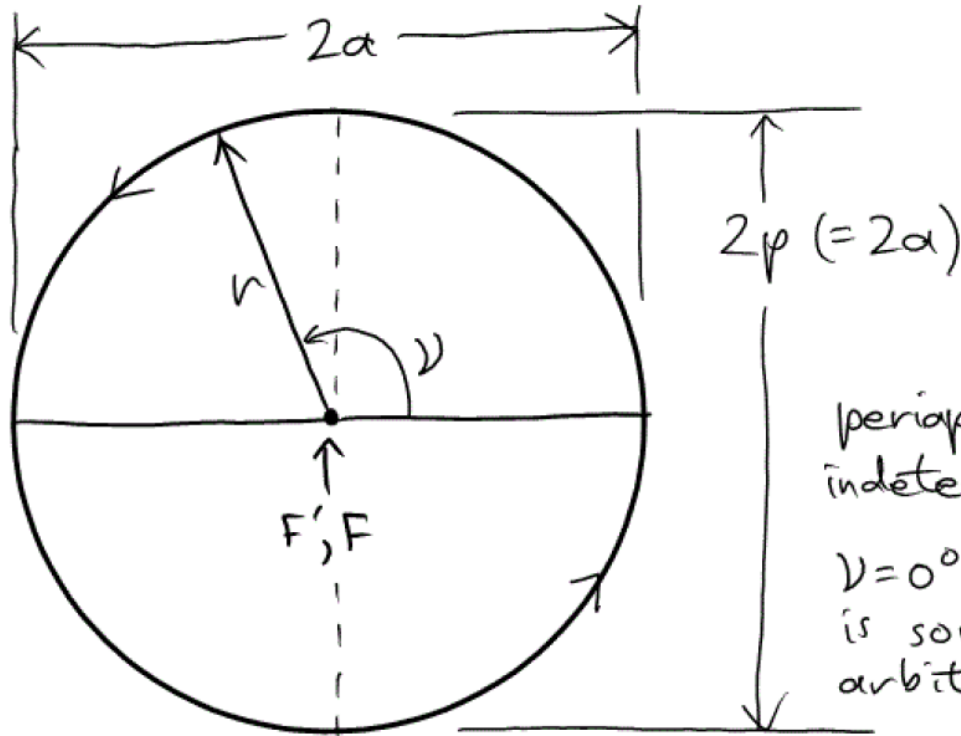


$$0 < e < 1$$

$$\varepsilon < 0$$

$$T = 2\pi \sqrt{\frac{a^3}{\mu}}$$

Conic Sections: Circle



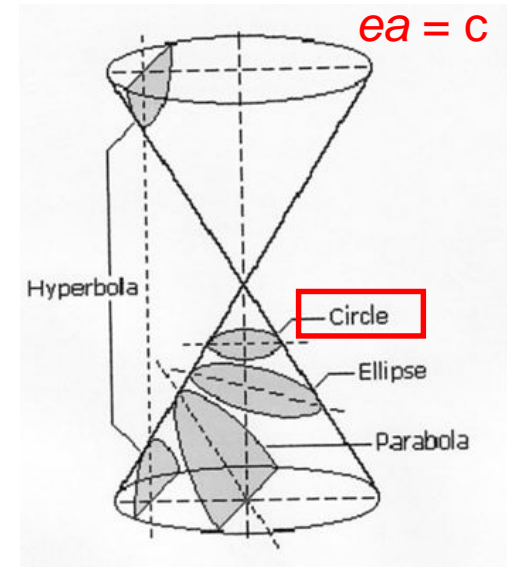
periapsis
indeterminate

$\nu = 0^\circ$ reference
is somewhat
arbitrary

Constant speed

$$r_a V_a = r_p V_p = r V = h = \text{constant}$$

$$v = \sqrt{\frac{\mu}{r}}$$

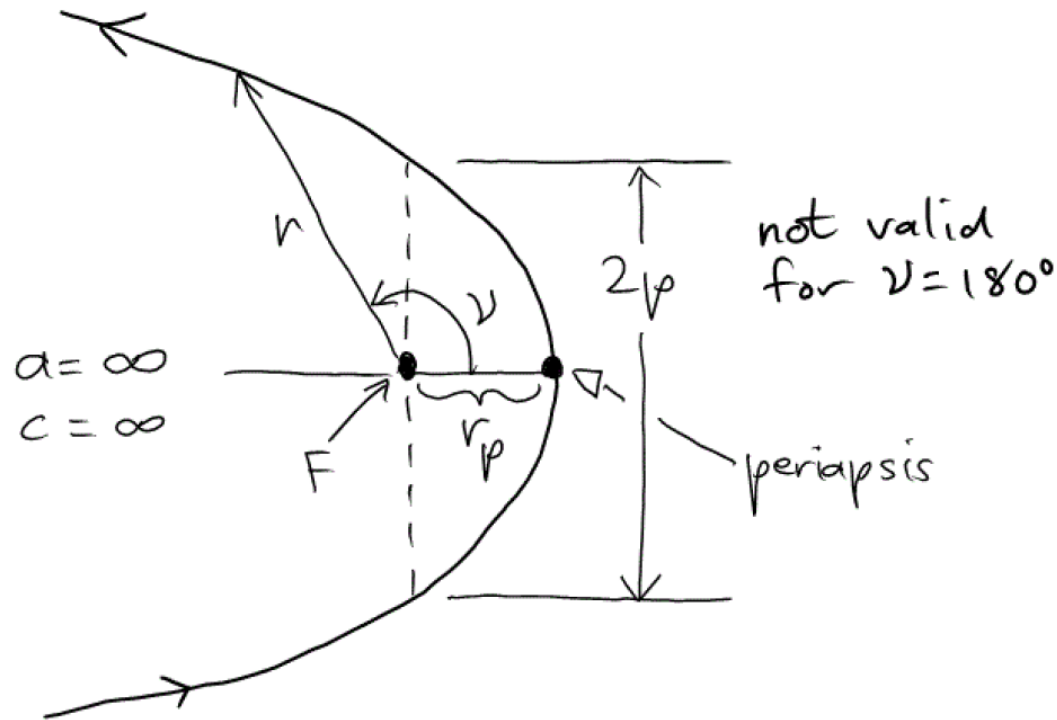


$$e = 0$$

$$\varepsilon < 0$$

$$T = 2\pi \sqrt{\frac{a^3}{\mu}}$$

Conic Sections: Parabola



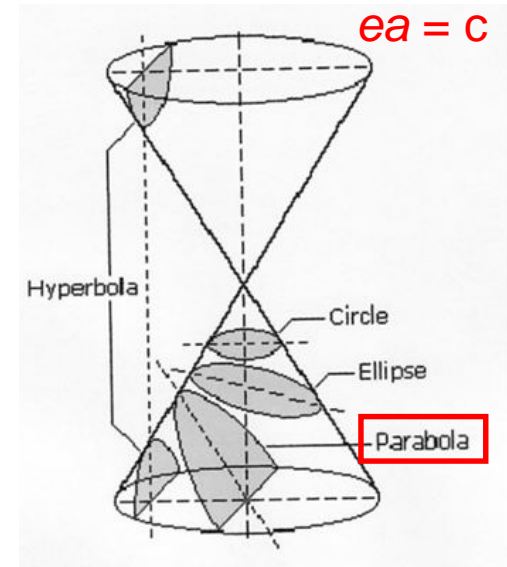
$$\mathcal{E} = 0$$



Minimum energy to escape

$$v = 0 \text{ when } r = \infty$$

Escape speed is defined as speed on parabola

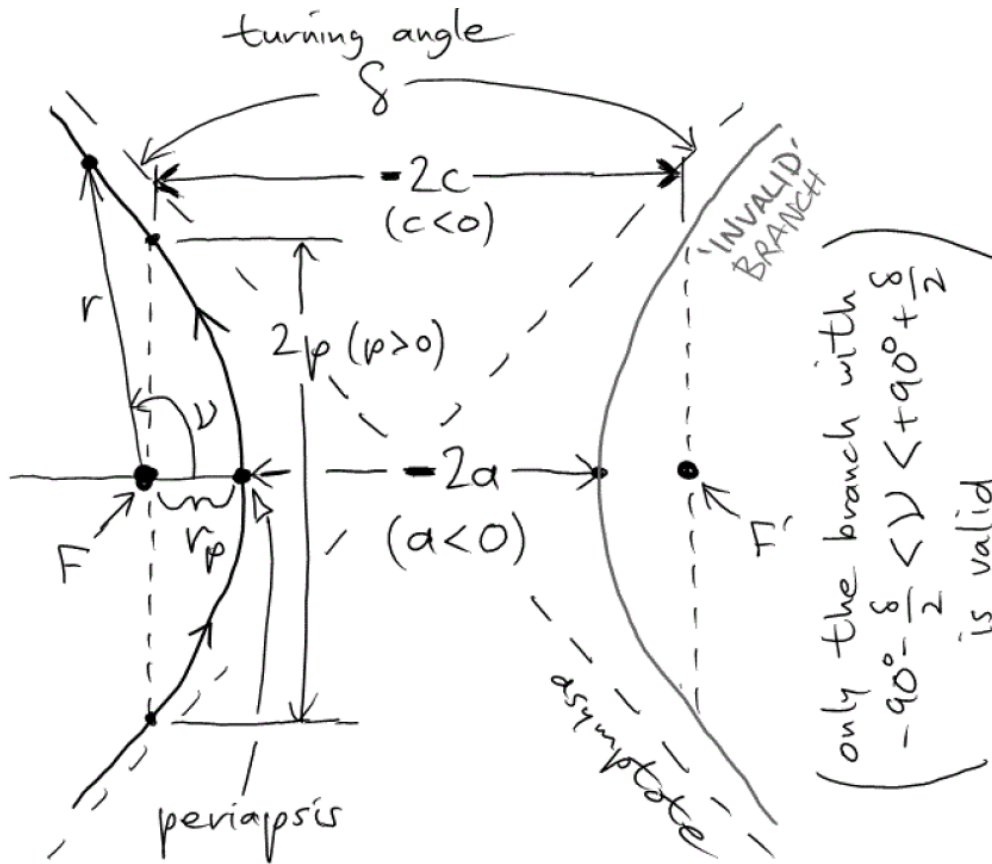


$$e = 1$$

$$\mathcal{E} = 0$$

T Undefined

Conic Sections: Hyperbola



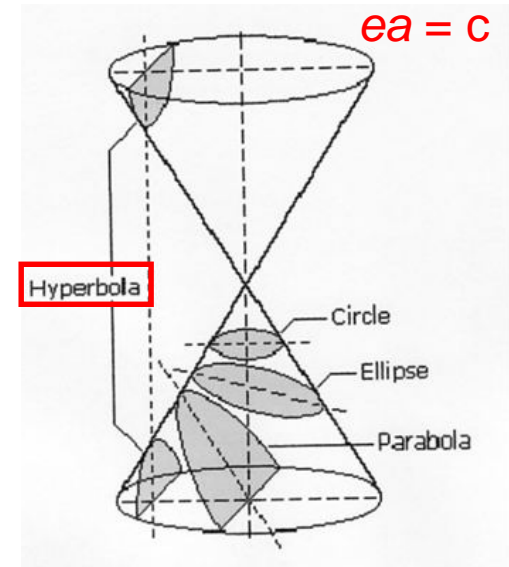
$$\mathcal{E} > 0$$



Enough energy to escape

$v \neq 0$ when $r = \infty$

Hyperbolic excess speed



$$e > 1$$

$$\mathcal{E} > 0$$

T Undefined

Conic Sections: Conservation Properties

Specific Mechanical Energy in Gravity Field

$$\mathcal{E} = \frac{v^2}{2} - \frac{\mu}{r}$$

$$\mathcal{E} = -\frac{\mu}{2a} ; a = -\frac{\mu}{2\mathcal{E}}$$

Velocity vs. distance

$$v = \sqrt{\frac{2\mu}{r} - \frac{\mu}{a}}$$

Vis-Viva Law

$a = \infty$



Escape velocity

$$v_0 = \sqrt{\frac{2\mu}{r_0}}$$

Regardless of velocity direction

Specific Angular Momentum

$$\vec{h} = \vec{r} \times \vec{v}$$

$$h = rvc\cos\phi_{\text{fpa}}$$

$$h = \sqrt{\mu p} ; p = \frac{h^2}{\mu}$$

$$\frac{p}{a} = -\frac{2h^2\mathcal{E}}{\mu^2}$$

$$e = \sqrt{1 - \frac{p}{a}}$$

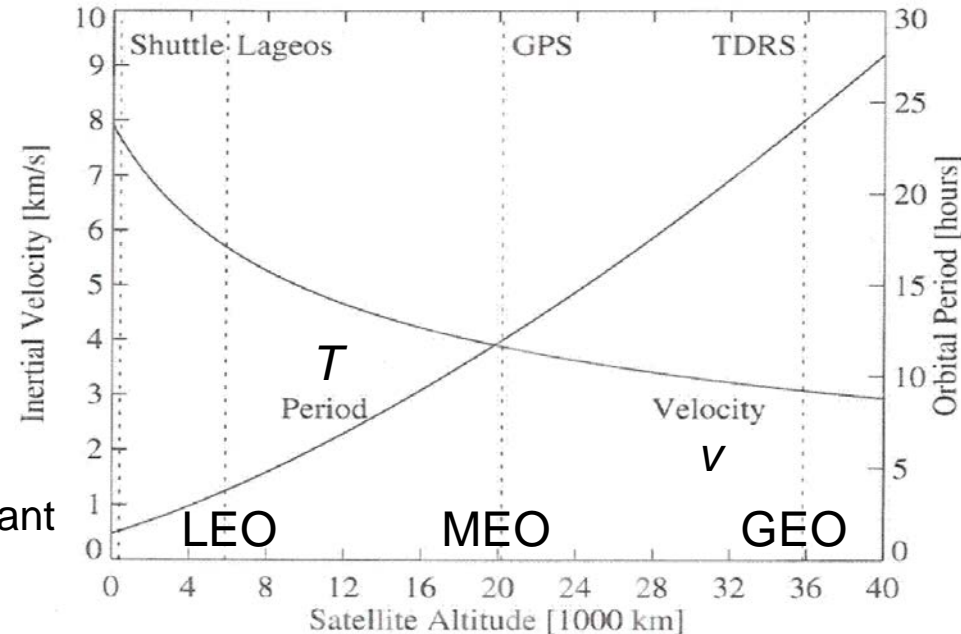
Eccentricity vs. energy and momentum

$$e = \sqrt{1 + \frac{2h^2\mathcal{E}}{\mu^2}}$$

Summary for Conics

- Two-body problem yields conic section solutions with focus at central body
- We have explained terminology and properties of conic sections
- h remains constant
 - Orbital motion lies in plane fixed in inertial space
 - Flight path angle ϕ continually changes to keep $h = r v \cos \phi = \text{constant}$
- \mathcal{E} remains constant
 - Kinetic energy (speed) and potential energy (distance) are exchanged
- We have introduced many formulas and information with no proofs...
- and we cannot describe anything as a function of time...

Satellite in circular Earth orbit



$$v = \sqrt{\frac{\mu}{r}}$$

First cosmic speed
(circular)

$$v_0 = \sqrt{\frac{2\mu}{r_0}}$$

Second cosmic speed
(parabola, escape)

Canonical Units (Different from Astronomical Units)

- Astronomers have introduced a normalized system of length/time units to
 - Scale calculations to natural dimensions of the problem
 - Cope with inaccurate known physical constants (e.g., distance and mass)
- For a given central body CB, define
 - 1 Distance Unit as radius of body

$$1 \text{ DU}_{\text{CB}} = R_{\text{CB}}$$
 - 1 Time Unit as time to travel 1 radian at 1 DU radius circular orbit

$$1 \text{ TU}_{\text{CB}} = \sqrt{\frac{R_{\text{CB}}^3}{\mu_{\text{CB}}}}$$
 - As a consequence by definition

$$\left\{ \begin{array}{l} v_{\text{circ}}(R_{\text{CB}}) = 1 \left[\frac{\text{DU}_{\text{CB}}}{\text{TU}_{\text{CB}}} \right] \\ \mu_{\text{CB}} = 1 \left[\frac{\text{DU}_{\text{CB}}^3}{\text{TU}_{\text{CB}}^2} \right] \end{array} \right.$$
 - DUs and TUs are units of measure (nondimensional) strictly linked to the central body

Backup