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# AA 279 A – Space Mechanics

## Lecture 6: Notes

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# Table of Contents

- Time and Reference systems
  
- Reading for next week (lectures 7 and 8)
  - Bate                      n/a
  - Montenbruck        6.2.2, 9.1.1
  - Vallado                6.8

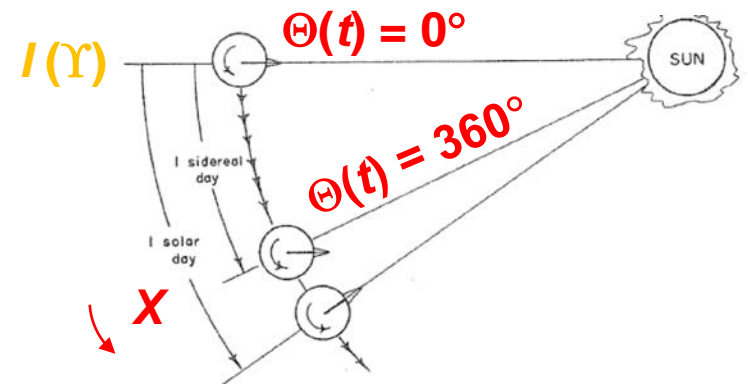
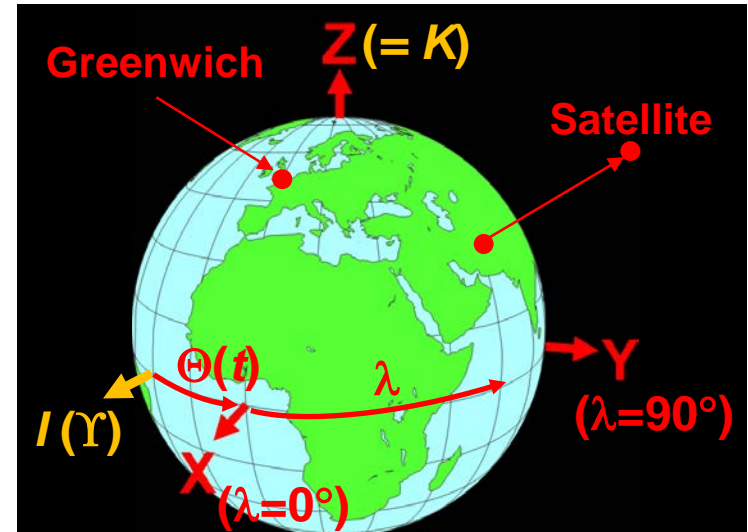
# Earth-Centered Earth-Fixed Coordinate System

- Solar Day ( $d$ ) – Familiar w.r.t. Sun  
Time between two subsequent meridian transits of sun is traditionally measured in solar days of **86400 [solar sec]**
- Sidereal Day – w.r.t. inertial direction  
Time taken for earth to rotate once w.r.t. inertial space, i.e. for  $\Theta$  to span  $360^\circ$ , takes about **86164 [solar sec]**, i.e. about 4 solar minutes shorter than solar day
- Neglecting small secular changes, the *Greenwich mean sidereal time* is given by

$$\Theta(d) = 280.4606^\circ + 360.9856473^\circ \cdot d$$

**Solar time in days since January 1, 2000 at 12h**

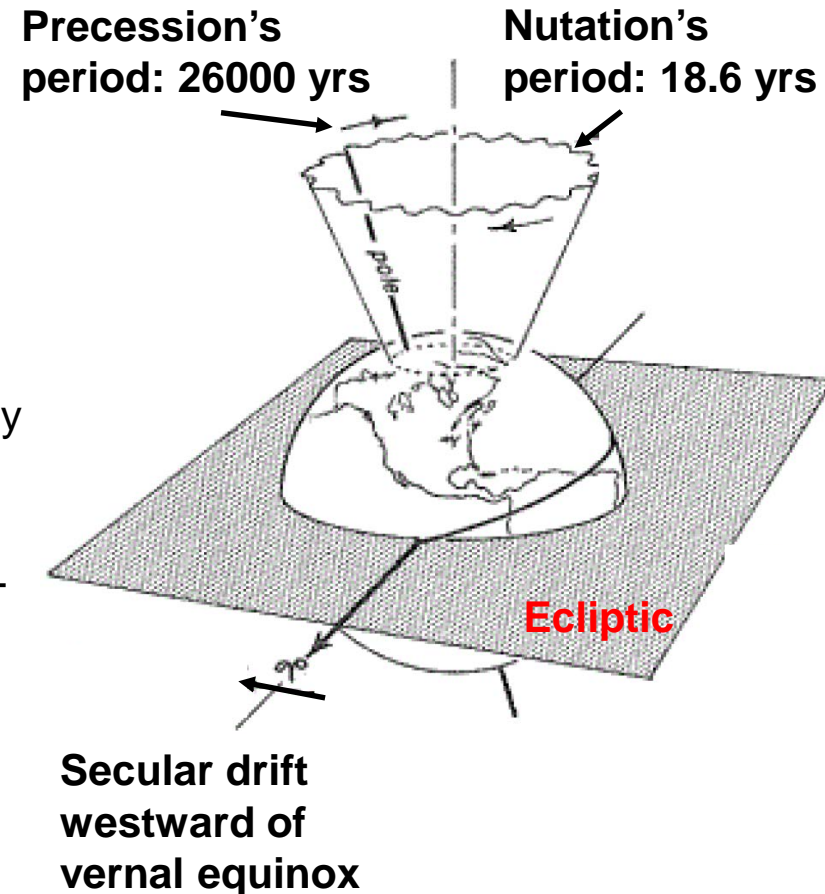
## Ideal configuration ECI/ECEF



# Time and Reference Systems

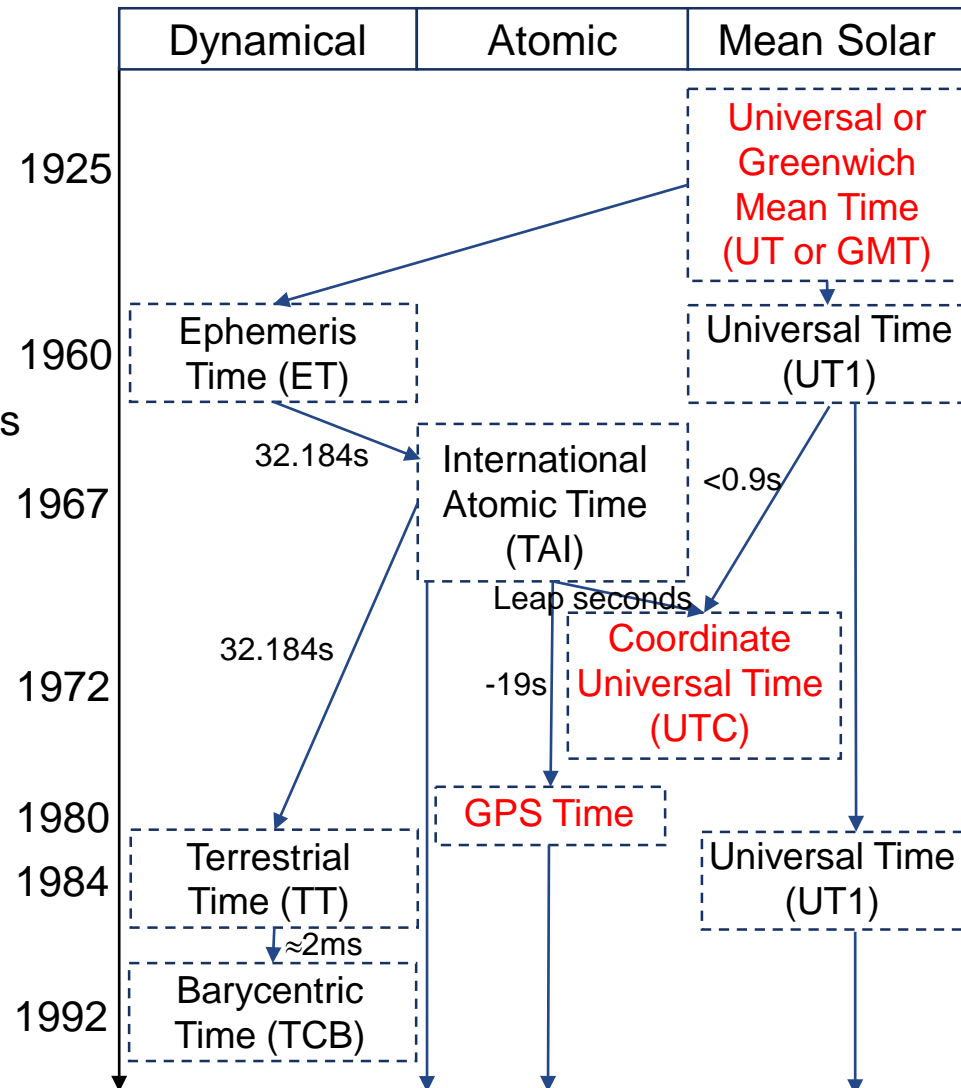
- So far we have tacitly assumed the availability of a unique time and reference system for the equation of motion (inertial)
- Traditionally, time and reference systems have been tied to the Earth's rotation and its revolution around the Sun
- In practice the Earth's orbit is slightly eccentric ( $e = 0.0167$ ). In addition we don't live in a 2-body system and the rotational dynamics of the Earth is complex (perturbations are covered later)
- The Moon, Sun, and other planets act on a non-spherical Earth and produce torques on top of the purely radial attractive force
- This causes secular, periodic and short-term variations of the vernal equinox, equator, and Earth's rotation axis

## Actual Earth's rotation dynamics



# Evolution of Time Scales

- Solar time
  - Tied to motion of Earth/Sun
  - Based on *mean* vernal equinox
  - Non-uniform
- Dynamical time
  - Tied to motion of solar system bodies
  - Based on dynamics model/theory
  - Uniform (theoretical)
- Atomic time
  - Tied to periodic properties of atoms
  - Based on laboratory measurements
  - Uniform (practical)
- Conversions between time scales are needed depending on adopted Earth's rotation model



# Simplified Calendrical Calculations

- Earth's rotation is typically expressed through Greenwich mean sidereal time  $\Theta$
- Spacecraft time is typically expressed through more familiar  $t_{\text{UTC}}$  (or  $t_{\text{GPS}}$ ) time
- Intermediate steps are required to convert between these time scales
- Since the civilian calendar is not well suited to handle time calculations, a Modified Julian Date (MJD) is used
- The algorithm to compute MJD from a calendar date is given in “Montenbruck” on page 321 (A.1.1)

- 1) Obtain necessary time differences  $t_{\text{UT1}} - t_{\text{UTC}}$  (and  $t_{\text{UTC}} - t_{\text{GPS}}$ ) from IERS bulletin
- 2) Compute universal time  $t_{\text{UT1}}$
- 3) Compute Modified Julian Date (MJD) corresponding to  $t_{\text{UT1}}$
- 4) Compute approximate Greenwich mean sidereal time in [deg]

$$\Theta(d) = 280.4606^\circ + 360.9856473^\circ \cdot d$$

$$d = \text{MJD} - 51544.5$$

# Conventional Reference Systems

## ➤ Celestial Reference Frame (CRF)

- Quasi-inertial
- Origin at Earth's center of mass
- I-axis aligned with mean equinox at J2000 epoch
- K-axis normal to mean equator's plane at J2000 epoch
- Determined from a set of precise coordinates of extragalactic radio sources
- Mean equator and equinox of J2000 were defined by the International Astronomical Union (IAU)
- *Earth Mean Equator and Equinox of J2000 (EME2000)* is a practical implementation based on the FK5 star catalogue

## ➤ Terrestrial Reference Frame (TRF)

- Co-rotating with Earth but slightly moving w.r.t. Earth's crust
- Origin at Earth's center of mass
- Z-axis aligned with International Earth Rotation Service (IERS) pole
- X-axis defined by intersection of plane normal to Z and Greenwich mean meridian
- Determined from a set of precise coordinates of ground stations and updated every year
- *World Geodetic System 1984 (WGS84)* is another popular implementation based on the Global Positioning System (GPS)



# Coordinate Transformation

- The transformation between CRF and TRF is accomplished by models for

- Precession

*Secular change in orientation of Earth's rotation axis and equinox*

- Nutation

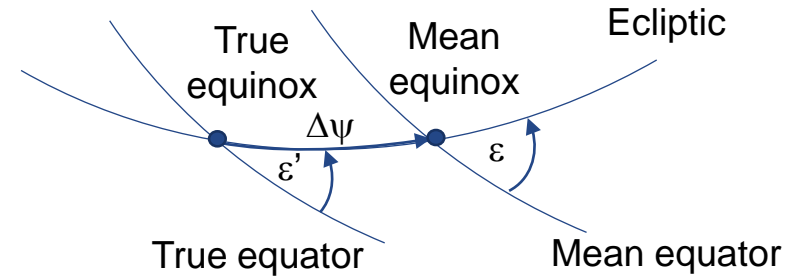
*Periodic and short-term variation of equator and equinox*

- Sidereal Time (w.r.t. UT1)

*Earth's rotation about its axis*

- These models are supplemented by the IERS Earth Observation Parameters (EOP), comprising

- Observations of UT1-TAI difference
- Measured coordinates of rotation axis relative to IERS pole



$$\vec{r}_{\text{MeanOfDate}} = \vec{R}_{\text{Pre}}(t) \vec{r}_{\text{CRS}}$$

$\approx 50''/\text{year}$  (1.54km/year)

$$\vec{r}_{\text{TrueOfDate}} = \vec{R}_{\text{Nut}}(t) \vec{r}_{\text{MeanOfDate}}$$

$\approx 20''$  (618m)

$$\vec{r}_{\text{TRS}'} = \vec{R}_{\text{Rot}}(t) \vec{r}_{\text{TrueOfDate}} = \vec{R}_z(\Theta + \Delta\psi \cos \epsilon) \vec{r}_{\text{TrueOfDate}}$$

Equation of Equinoxes:  $\approx 20''$  (618m)

UT1-UTC:  $\approx 15''$  (464m)

$$\vec{r}_{\text{TRS}} = \vec{R}_{\text{Pol}}(t) \vec{r}_{\text{TRS}'}$$

$\approx 0.3''$  (10m)

$$\vec{r}_{\text{TRS}} = \vec{R}_{\text{Pol}}(t) \vec{R}_{\text{Rot}}(t) \vec{R}_{\text{Nut}}(t) \vec{R}_{\text{Pre}}(t) \vec{r}_{\text{CRS}}$$