### **Wave Optics Simulator**

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#### Contents

- Theoretical Approach
- Say Hello to the Application
   Single Slit experiment
   Double Slit experiment
- Single Lens Simulations

Point Spread Function

**Abboration Test** 

Best Image Finder

Fourier Operation

Focusing

Perfect Lensing

Far Objects

Holography

Black-Hole weak Gravitational Lensing

Multi Lens Simulations

Telescope and Allignment

Hybrid Images

High-order Lens

Metasurface

## Theoretical Approach

#### Scalar Wave Equation

Helmholtz equation: 
$$\vec{r} + \vec{k} + \vec{k} = 0$$
 Source free

Where  $\vec{V}(\vec{r})$  is the Electro-Magnetic field

and we have the intensity:  $\vec{I}(\vec{r}) = |\psi(\vec{r})|^2$ 
 $i\vec{k} \cdot \vec{r}$ 

Plane waves:  $\psi(\vec{r}) = 0$ 

(a complete solution for the problem)

 $\psi(\vec{r}) = 0$ 

#### **Green Function**

Green function: 
$$\nabla^2 G + k^2 G = \delta (\vec{r} - \vec{r}')$$
,  $G = G (\vec{r}, \vec{r}')$ 

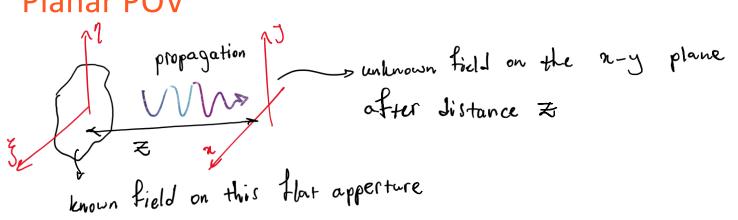
$$G = G (\vec{r}, \vec{r}') = G (\vec{r} - \vec{r}') = \frac{e^{i\vec{k}\cdot(\vec{r} - \vec{r}')}}{4\vec{r} | \vec{r} - \vec{r}'|}$$

isotropic

homogeneous

space

#### Planar POV



#### **Huygens Principle**

Green's 3rd theorem:
$$\int (U \nabla^2 V - V)$$

$$i + i \quad V = G(\dot{\tau}, \dot{\tau}, \dot{\tau})$$

Green's 3" theorem:
$$\int (U \nabla^2 V - V \nabla^2 U) dV = \int (U \nabla V - V \nabla U) dS^{n}$$

$$if: V = G(\dot{\tau}, \dot{\tau}, ) \qquad \nabla^2 G = -\delta(\dot{\tau} - \dot{\tau}_0) - k^2 G$$

$$U = \psi(\vec{r}| = 0 \quad \forall \vec{r} = -k^{2}\psi$$

$$= 0 \quad U(\vec{r}_{1}) = \frac{1}{j\lambda} \quad \text{is } \cos \theta \quad \frac{e^{jk|\vec{r}_{1} - \vec{r}_{0}|}}{|\vec{r}_{1} - \vec{r}_{0}|} \quad \text{is } u(\vec{r}_{0})$$

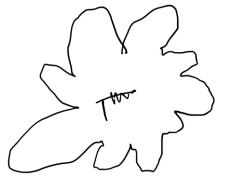
$$= 0 \quad \text{on the hole} \quad \text{each point } radiates \text{ in sphere}$$

#### Near Field to Far Field

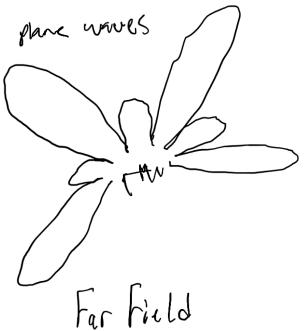
Abtuna vier:



ment field



Fersnel Field



tar tield (Frankater field)

202

distance

#### **Paraxial Approximation**

$$\mathcal{T}(x,y) = \frac{1}{j\lambda} \sum_{k=1}^{\infty} \left[ \nabla(x,y) \frac{e^{jkT_{0j}}}{r_{0j}} \cos dx \right]$$

$$\mathcal{T}(x,y) = \frac{1}{j\lambda} \sum_{k=1}^{\infty} \left[ \nabla(x,y) \frac{e^{jkR}}{R^2} dx dy \right], \quad R^2 = 2^2 + (\alpha - 2)^2 + (\alpha - 1)^2$$

$$\mathcal{T}(x,y) = \frac{e^{jkR}}{J\lambda} + \cdots \quad \left[ |\alpha| < 1 \Rightarrow R = 2 \left( 1 + \left( \frac{x-2}{2} \right)^2 + \left( \frac{y-y}{2} \right)^2 \right) \approx 2 \left( 1 + \frac{y}{2} \frac{(x-2)^2 + (y-y)^2}{2} \right)$$

$$\mathcal{T}(x,y) = \frac{e^{jkR}}{J\lambda} \int_{\mathbb{R}^2} \left[ \nabla(x,y) \exp\left( \frac{jk}{2} \right) (x-y)^2 + (y-y)^2 \right] dx dy$$

$$\mathcal{T}(x,y) = \frac{e^{jkR}}{J\lambda} \int_{\mathbb{R}^2} \left[ \nabla(x,y) \exp\left( \frac{jk}{2} \right) (x-y)^2 + (y-y)^2 \right] dy dy$$

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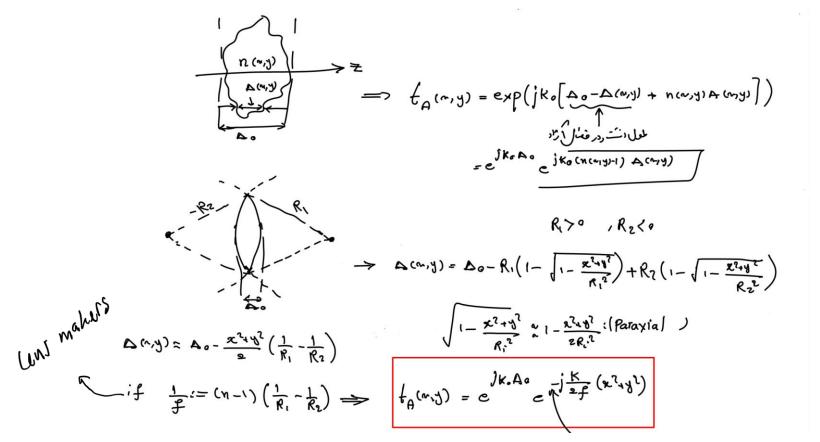
$$\mathcal{T}(x,y) = \frac{e^{jkR}}{J\lambda} \int_{\mathbb{R}^2} \left[ \nabla(x,y) \exp\left( \frac{jk}{2} \right) (x-y)^2 + (y-y)^2 \right] dy dy$$

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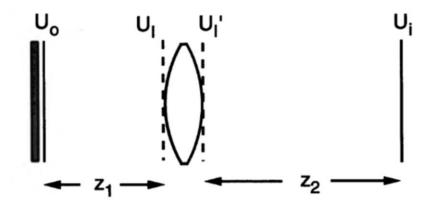
$$\mathcal{T}(x,y) = \frac{e^{jkR}}{J\lambda} \int_{\mathbb{R}^2} \left[ \nabla(x,y) \exp\left( \frac{jk}{2} \right) (x-y)^2 + (y-y)^2 \right] dy dy$$

#### Free Space as a Transfer Function

#### **Lens Transfer Fubction**



#### **Imaging Condition**



For having a point image of an initial point input we should satisfy the condition:

$$\frac{1}{Z_1} + \frac{1}{Z_2} = \frac{1}{P}$$

#### Light as Quantum Photons

Instead of the classical coherent light we can shine the coherent state, although it has uncertaibty in

both amplitude and phase.

Displacement

Displacement

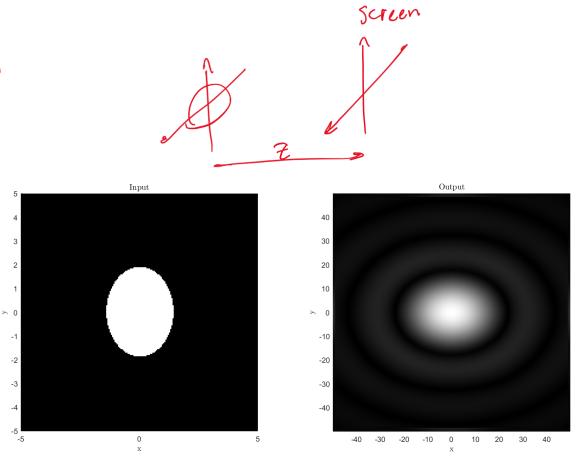
$$X_1$$
 $X_2$ 
 $X_3$ 
 $X_4$ 
 $X_4$ 
 $X_5$ 
 $X_6$ 
 $X_6$ 
 $X_7$ 
 $X_8$ 
 $X_8$ 

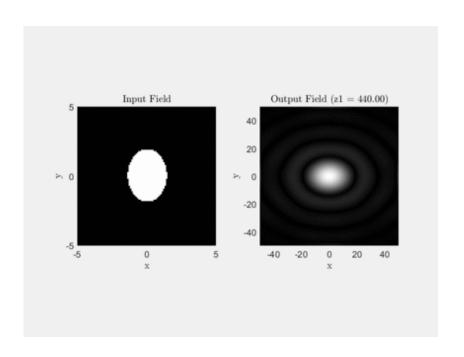
#### Light as Quantum Photons

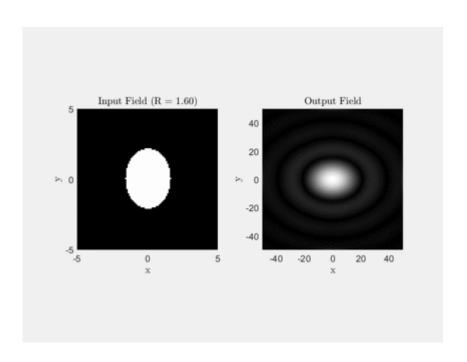
For different modes summing we will have the same wavefront for the coherent state which changes in the linear optical system.

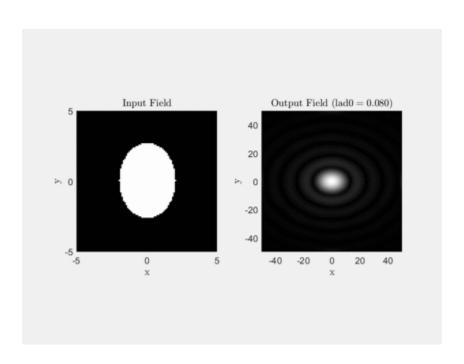
Set of 
$$a_{k}^{\dagger}$$
,  $[a_{k}^{\dagger}, a_{k}^{\dagger}] = \delta_{k,k'}$  in moneyton set if  $a_{k}^{\dagger}$ ,  $[a_{k}^{\dagger}, a_{k'}] = f(k-k')$  in position  $a_{k'}^{\dagger}$  of  $a_{k'}^{$ 

## No Lens Test

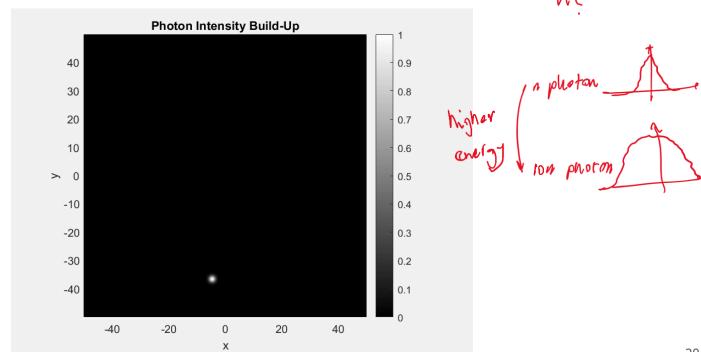




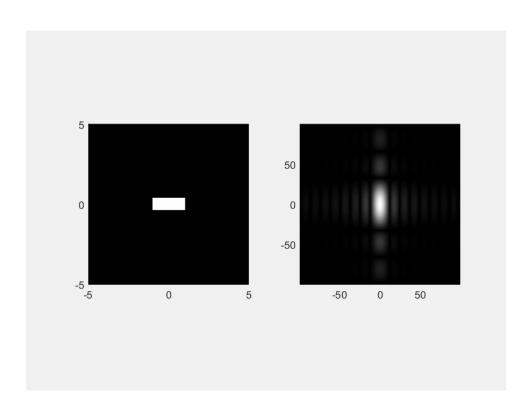




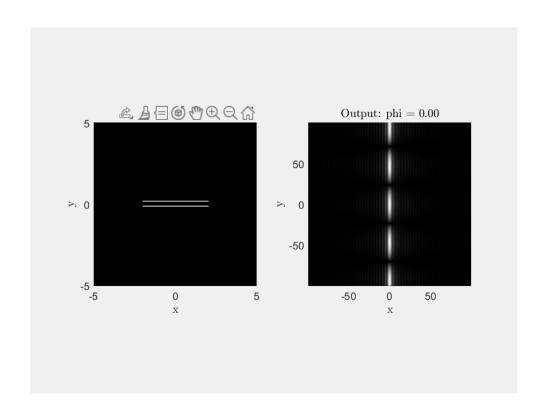
#### **Photons Sampling**



#### Single Slit

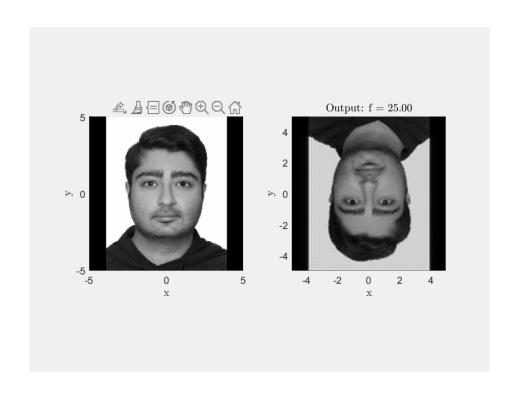


#### **Double Slit**

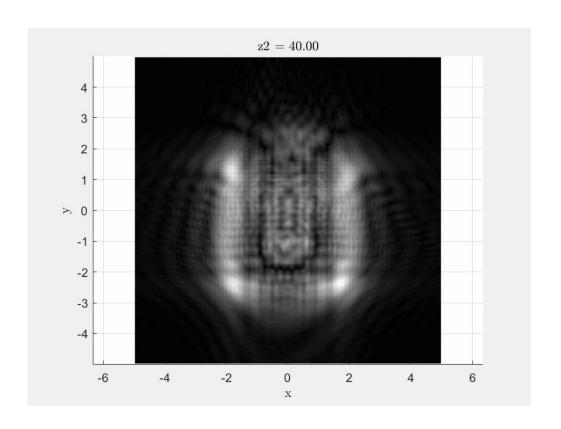


# Single Lens Simulations

#### Single Lens

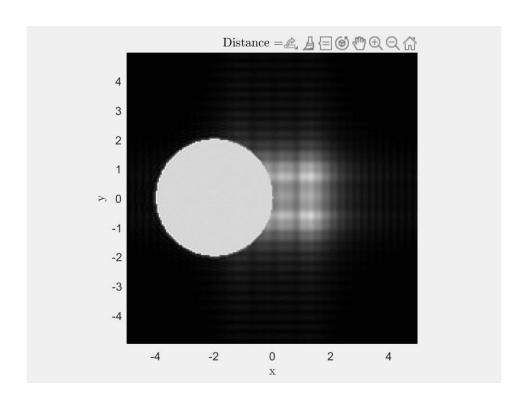


#### Single Lens

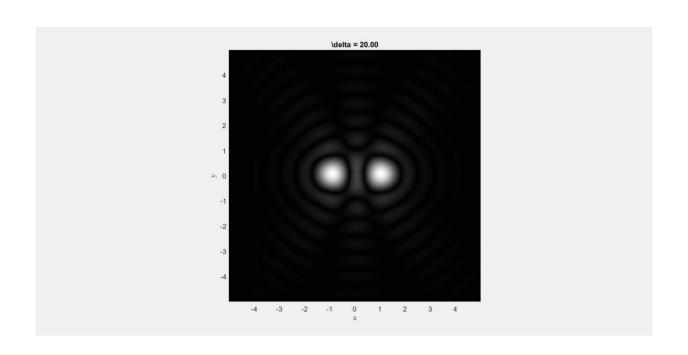


#### 2f System

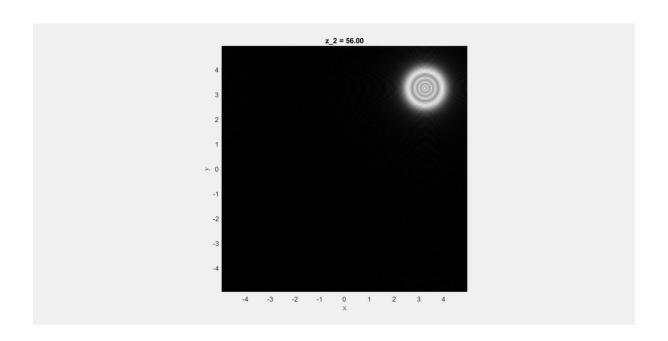
#### Focusing



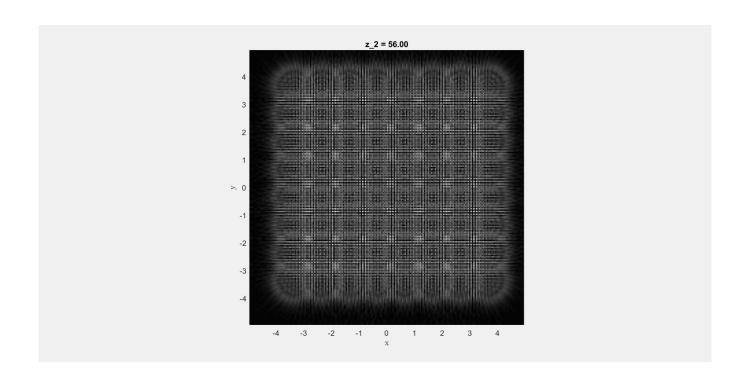
#### Two Stars Reileigh Criterion



#### **Comma Abboration**

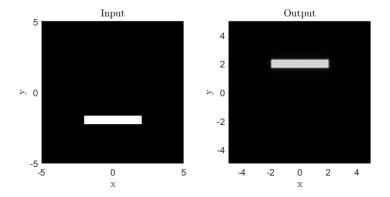


#### **Point Spread Function**

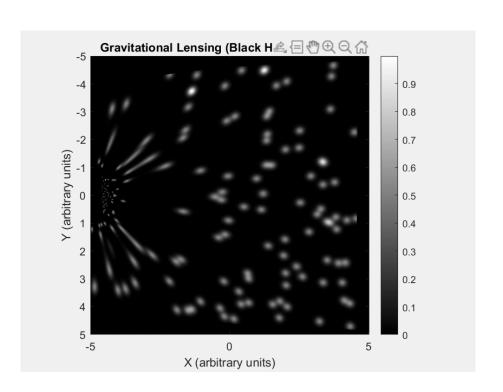


#### Perfect Lens

#### **Perfect Lens**

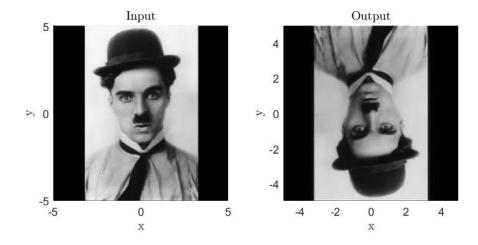


#### Black Hole as a Lens



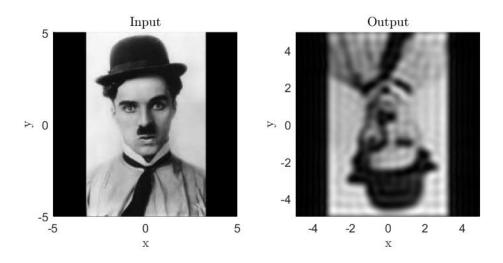
# Multi Lens Simulations

#### 4f System

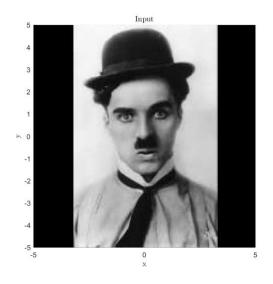


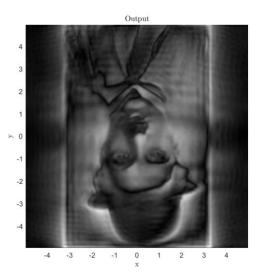
35

#### **Low Pass Filter**

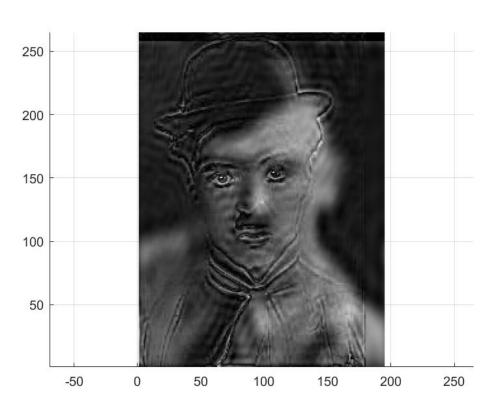


#### High Pass Filter

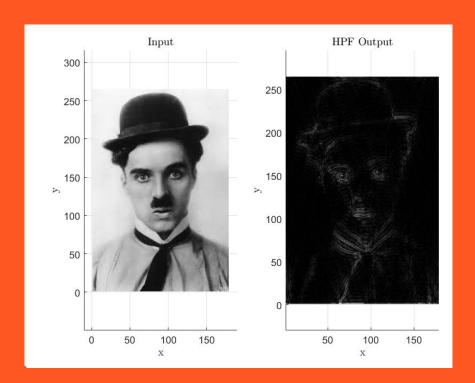


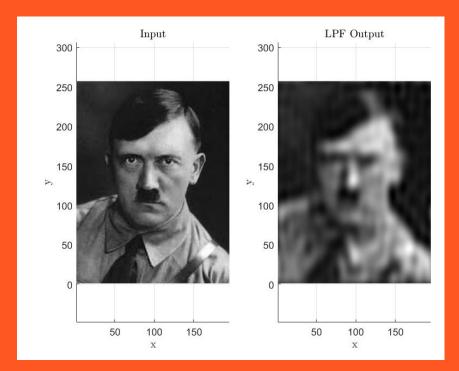


#### **Hybrid Image**



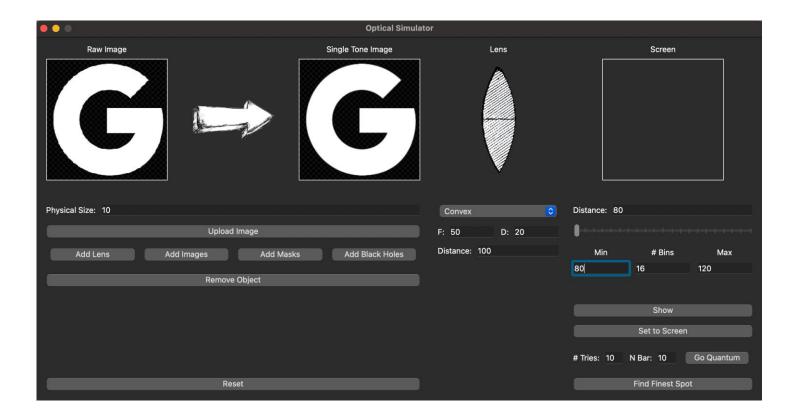
#### Hybrid Image





# Say Hello to the Application

#### **User Interface**



# Appendix

## Any question?

#### Refrences

- <u>J. Goodman</u>. Introduction to Fourier optics, 3rd ed., by JW Goodman. Englewood, CO: Roberts & Co. Publishers, 2005, (2005)
- M. Rezai and J. A. Salehi, "Quantum fourier optics (QFO)," 2023

#### Materials

Found around the house!

- 2 drinking glasses
- Table salt
- 2 eggs
- Water