

HOMEWORK 4

1. (a) With $E = \mu \cdot B$ and considering the spin up and spin down cases, we have:

$$\begin{aligned}
 \Delta E &= \mu_B B \cos(0) - \mu_B B \cos(\pi) \\
 &= 2\mu_B B \\
 &= 2 \frac{e\hbar}{2m_e} B \\
 &= 2.24 \times 10^{-23} \text{ J} = 1.40 \times 10^{-4} \text{ eV}
 \end{aligned}$$

- (b) The energy required for the electron to flip spin is equal to the ΔE found in the last question.

Using the equation $E = \frac{hc}{\lambda}$, we have:

$$\begin{aligned}
 \lambda &= \frac{hc}{\Delta E} \\
 &= 8.94 \times 10^{-3} \text{ m}
 \end{aligned}$$

The wavelength which causes the electrons' spin to flip is 8.94 mm

2. We have:

$$\Delta E = 2\mu_B B = \frac{hc}{\lambda}$$

Thus,

$$B = \frac{hc}{\lambda} \frac{1}{2\mu_B} = .0511 \text{ T}$$

3. For the states with $\ell = 4$ and $s = 1/2$, The state with the largest possible j and largest possible m_j has $j = 9/2$ and $m_j = 9/2$.

- (a) We have:

$$\begin{aligned}
 \vec{J} &= \vec{L} + \vec{S} \\
 |\vec{J}| &= \sqrt{j(j+1)}\hbar = \sqrt{99}\hbar/2 \\
 |\vec{L}| &= \sqrt{\ell(\ell+1)}\hbar = \sqrt{120}\hbar \\
 |\vec{S}| &= \sqrt{s(s+1)}\hbar = \sqrt{3}\hbar/2
 \end{aligned}$$

Now, apply the law of cosines

$$\begin{aligned}
 J^2 &= L^2 + S^2 - 2LS \cos(\pi - \theta) \\
 \frac{99}{4} &= 20 + \frac{3}{4} + 2\sqrt{15} \cos \theta \\
 \theta &= \cos^{-1} \left(\frac{2}{\sqrt{15}} \right) = 1.03 \text{ rad or } 58.9^\circ
 \end{aligned}$$

- (b) $\vec{\mu}_\ell$ is antiparallel to \vec{L} and $\vec{\mu}_s$ is antiparallel to \vec{S} , thus the angle is the same as in part (a), $\theta = 58.9^\circ$.

(c)

$$\cos \theta = \frac{J_z}{|\vec{J}|} = \frac{9}{\sqrt{99}}$$

Thus, $\theta = 25.24^\circ$.

4. (a) For the $3d$ level, we know that $n = 3$, $\ell = 2$, and $s = 1/2$.

Therefore, $j = \frac{3}{2}, \frac{5}{2}$ and $m_j = -\frac{5}{2}, -\frac{3}{2}, -\frac{1}{2}, \frac{1}{2}, \frac{3}{2}, \frac{5}{2}$

For $j = \frac{3}{2}$, we have $m_j = -\frac{3}{2}, -\frac{1}{2}, \frac{1}{2}, \frac{3}{2}$, i.e., 4 degenerate states. Also, the spectroscopic notation:

$$3^2D_{3/2}$$

For $j = \frac{5}{2}$, we have $m_j = -\frac{5}{2}, -\frac{3}{2}, -\frac{1}{2}, \frac{1}{2}, \frac{3}{2}, \frac{5}{2}$, i.e., 6 degenerate states. Also, the spectroscopic notation:

$$3^2D_{5/2}$$

- (b) It is impossible to have this state because the maximum possible value for ℓ is $\ell = n - 1$, and since $n = 2$ in this state, it is impossible for $\ell = 2$, i.e., the D subshell.
- (c) The ground state of the H atom is not split into two sub-levels by spin-orbit coupling because $\ell = 0$, and there is only one electron, thus there is no internal magnetic field present to interact with the electron's magnetic moment.

5. (a)

$$a_0 = \frac{4\pi\epsilon_0\hbar^2}{m_e e^2}$$

$$\begin{aligned} E_I &= \frac{e^2}{4\pi\epsilon_0} \frac{1}{2a_0} \\ &= \frac{e^2}{4\pi\epsilon_0} \frac{m_e^2 e^2}{8\pi\epsilon_0 \hbar^2} \\ &= \frac{1}{2} \frac{e^4}{16\pi^2 \epsilon_0^2 \hbar^2 c^2} m_e^2 c^2 \\ &= \frac{1}{2} \left(\frac{e^2}{4\pi\epsilon_0 \hbar c} \right)^2 m_e^2 c^2 \\ &= \frac{1}{2} \alpha^2 m_e^2 c^2 \end{aligned}$$

- (b) We have the relativistic energy:

$$E = \sqrt{(\mu c^2)^2 + (pc)^2}$$

And we know that the rest energy, E_0 :

$$E_0 = \mu c^2$$

Thus we can find Kinetic energy by subtracting the rest energy from the total relativistic energy.

$$\begin{aligned}
 \text{KE} &= \sqrt{(\mu c^2)^2 + (pc)^2} - \mu c^2 \\
 &= \sqrt{\mu^2 c^4 + \frac{p^2 c^2 \mu^2 c^4}{\mu^2 c^4}} - \mu c^2 \\
 &= \mu c^2 \sqrt{1 + \left(\frac{pc}{\mu c^2}\right)^2} - \mu c^2 \\
 &\approx \mu c^2 \left[1 + \frac{1}{2} \left(\frac{pc}{\mu c^2}\right)^2 - \frac{1}{8} \left(\frac{pc}{\mu c^2}\right)^4 \right] - \mu c^2 \quad (\text{using a Taylor series expansion}) \\
 &= \mu c^2 \left[\frac{1}{2} \left(\frac{pc}{\mu c^2}\right)^2 - \frac{1}{8} \left(\frac{pc}{\mu c^2}\right)^4 \right] \\
 &= \frac{p^2}{2\mu} - \frac{1}{8} \frac{p^4}{\mu^3 c^2} \\
 &= -\frac{p^4}{8\mu^3 c^2} \quad (\text{subtracting classical } KE = p^2/2\mu)
 \end{aligned}$$

Thus, we have that the lowest order relativistic correction to the classical kinetic energy is

$$\Delta \text{KE} = \frac{p^4}{8\mu^3 c^2}.$$

(c) Since $j = \ell + s$, there are two possibilities for j :

$$j = \ell - \frac{1}{2} \quad \text{and} \quad j = \ell + \frac{1}{2}$$

Thus,

$$\ell = j + \frac{1}{2} \quad \text{and} \quad \ell = j - \frac{1}{2}$$

Start with the $\ell = j - 1/2$ case:

$$\begin{aligned}
 f(j, \ell) &= \frac{j(j+1) - 3\ell(\ell+1) - \frac{3}{4}}{\ell(\ell + \frac{1}{2})(\ell+1)} \\
 &= \frac{j(j+1) - 3(j - \frac{1}{2})(j + \frac{1}{2}) - \frac{3}{4}}{(j - \frac{1}{2})(j)(j + \frac{1}{2})} \\
 &= \frac{j^2 + j - 3j^2}{(j - \frac{1}{2})(j)(j + \frac{1}{2})} \\
 &= \frac{j(1 - 2j)}{(j - \frac{1}{2})(j)(j + \frac{1}{2})} \\
 &= \frac{1 - 2j}{(j - \frac{1}{2})(j + \frac{1}{2})} \\
 &= \frac{-2(j - \frac{1}{2})}{(j - \frac{1}{2})(j + \frac{1}{2})} \\
 &= -\frac{2}{(j + \frac{1}{2})}
 \end{aligned}$$

Now, start with the $\ell = j + 1/2$ case:

$$\begin{aligned}
 f(j, \ell) &= \frac{j(j+1) - 3\ell(\ell+1) - \frac{3}{4}}{\ell(\ell + \frac{1}{2})(\ell+1)} \\
 &= \frac{j(j+1) - 3(j + \frac{1}{2})(j + \frac{3}{2}) - \frac{3}{4}}{(j + \frac{1}{2})(j+1)(j + \frac{3}{2})} \\
 &= \frac{j(j+1) - 3j^2 - 6j - 3}{(j + \frac{1}{2})(j+1)(j + \frac{3}{2})} \\
 &= \frac{-2j^2 - 5j - 3}{(j + \frac{1}{2})(j+1)(j + \frac{3}{2})} \\
 &= \frac{-2(j^2 + \frac{5}{2}j + \frac{3}{2})}{(j + \frac{1}{2})(j^2 + \frac{5}{2}j + \frac{3}{2})} \\
 &= -\frac{2}{j + \frac{1}{2}}
 \end{aligned}$$

Hence, we have seen for both case that $f(j, \ell)$ is actually independent of ℓ and equal to

$$f(n, \ell) = -\frac{2}{j + \frac{1}{2}}.$$