

**LAB 11 ASSIGNMENT**  
**CE 593 ADVANCED REMOTE SENSING**

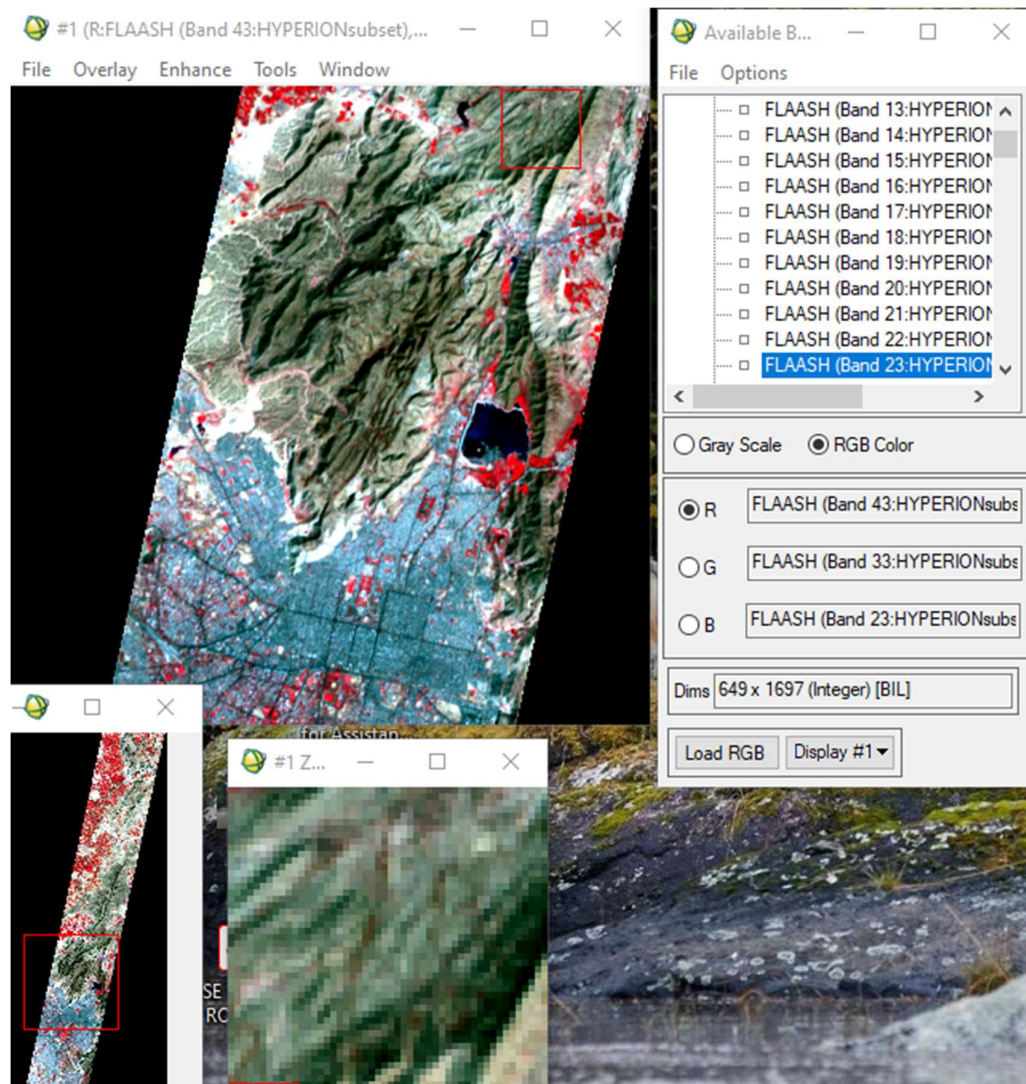
### Spectral Hourglass for Hyperspectral Data Analysis

The hourglass processing flow analyzes hyperspectral data to find spectrally pure pixels (endmembers) and map their locations.

It starts with atmospherically corrected reflectance data.

Users can subset, visualize data in n-D space, and cluster pure pixels into endmembers or input their own.

The flow maps endmember distribution and abundance.



ENVI Spectral Hourglass Wizard

ENVI Spectral Hourglass Wizard

INTRODUCTION

Welcome to the ENVI Spectral Hourglass Wizard. This Wizard leads you through the ENVI hourglass processing flow to automatically find and map image spectral endmembers from hyperspectral or multispectral data. The hourglass processing flow leverages the unique spectrally over determined nature of hyperspectral data, permitting sub-pixel target detections, material identification and unambiguous spectral unmixing. The processing steps performed by this Wizard are:

1. Input/Output File Selection (input data are usually in reflectance)

2. MNF Transformation to reduce spectral dimensions

3. Review of MNF Results

4. Data Dimensionality determination using Spatial Coherence measures

5. Pixel Purity Index (PPI) to reduce spatial dimensions

6. Review of PPI Results

7. n-Dimensional Visualization, including auto-clustering, to select and retrieve individual endmembers

8. The option to manually add User Supplied Endmembers

9. Spectral mapping using Spectral Angle Mapper (SAM) and/or Mixture-Tuned Matched Filtering (MTMF)

10. Investigation of Mapping Results

SELECT INPUT/OUTPUT FILES

Typically, data that have been converted to reflectance via atmospheric correction, are used as input, particularly if library spectra or other external spectra are used as endmembers in the mapping process. Wavelength values must be assigned in the ENVI header of the input file. If you are not using any external spectra in the processing, then radiance or even uncalibrated data can be used as input. However, unless reflectance data are used, the image-derived endmembers cannot be identified by reference to a spectral library using the Spectral Analyst tool.

Spatial and spectral subsets of the input data can be selected if desired. For example, typical wavelength ranges used in VNIR/SWIR data are 0.4 - 1.3 micrometers for vegetation analysis and/or iron oxide mineral mapping or 2.0 - 2.5 micrometers for mapping of most other geologic materials. A mask can also be selected to exclude selected pixels such as image borders, bad pixels, or specific materials such as water or clouds. Be judicious in your choices of spatial and spectral subsets, avoiding unnecessary complications, data volumes and scene complexity.

Various suffixes are attached to the output root name for output files from the various processes (e.g. \_mnf for MNF Transform results, \_ppi for Pixel Purity Index results, etc).

The choices for this step are:

Select Input File...

Input File: <None Selected>

Select Output Root Name...

Output Root Name: <None Selected>

Cancel

<- Prev

Next ->

Cancel

<- Prev

Next ->

Spectral Hourglass Wizard Input File

Select Input File:

outputfilereflec

File Information:

File: D:\INDIAN INSTITUTE OF TECHNOLOGY, G  
Dims: 649 x 1697 x 196 [BIL]  
Size: [Integer] 431,730,376 bytes.  
File Type : ENVI Standard  
Sensor Type: Unknown  
Byte Order : Host (Intel)  
Projection : UTM, Zone 43 North  
Pixel : 30 Meters  
Datum : WGS-84  
Wavelength : 426.81601 to 2397.530029 Nanomete  
Upper Left Corner: 1,1  
Description: Mosaic Result [Mon Oct  
28 16:54:05 2024]

Spatial Subset

Full Scene

Spectral Subset

158/196 Bands

Select Mask Band

<None Selected>

Mask Options ▾

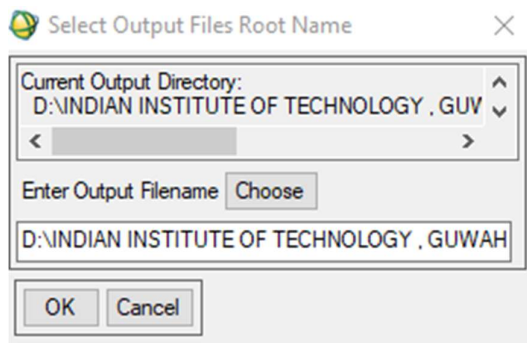
OK

Cancel

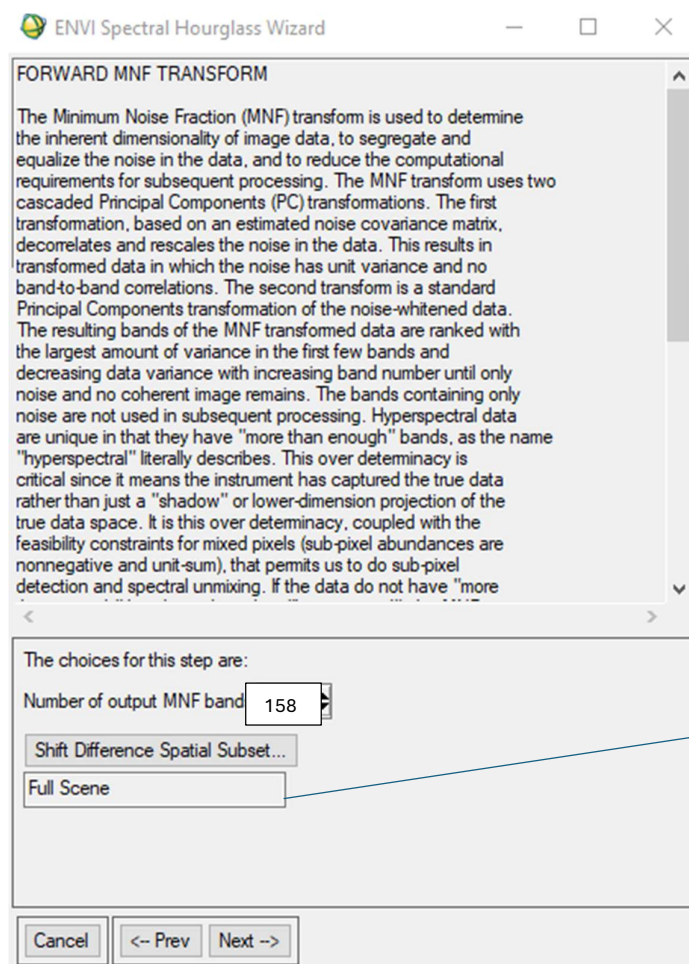
Previous

Open ▾

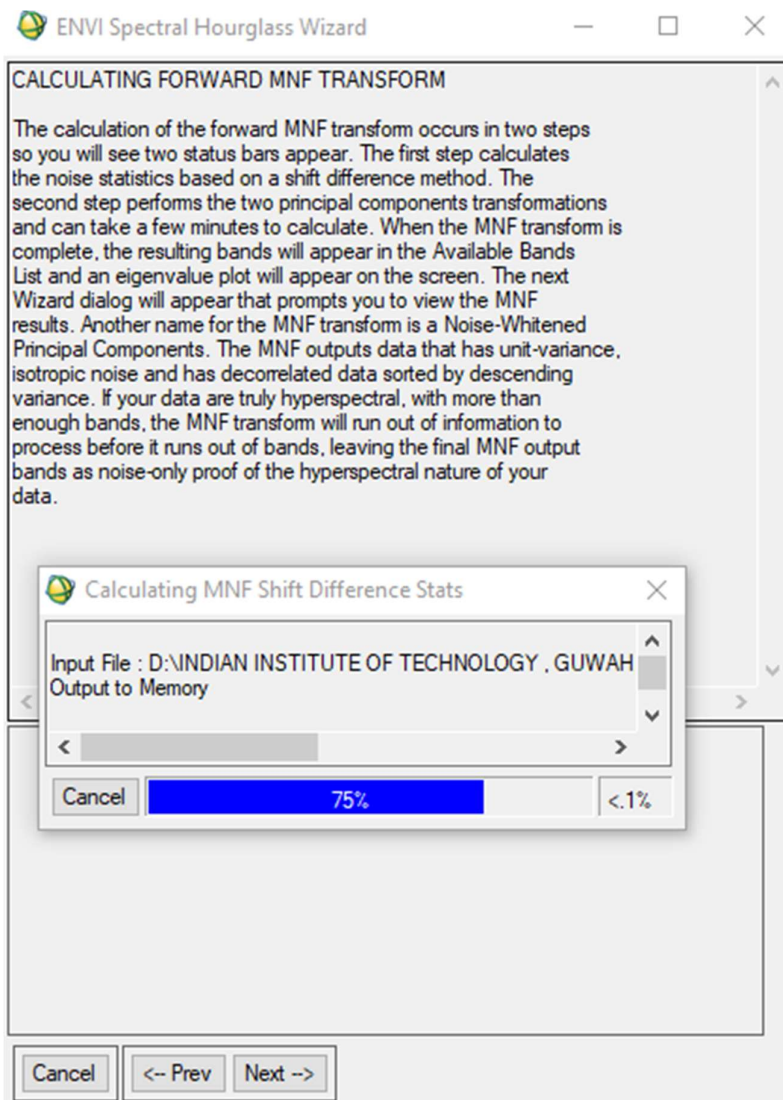
Number of output  
MNF bands



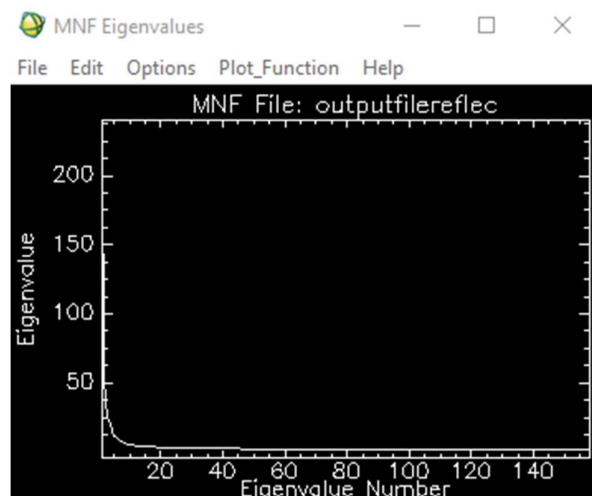
## Forward MNF Transform



Shift difference method  
that uses local pixel  
variance to estimate noise



## MNF Eigenvalues plot

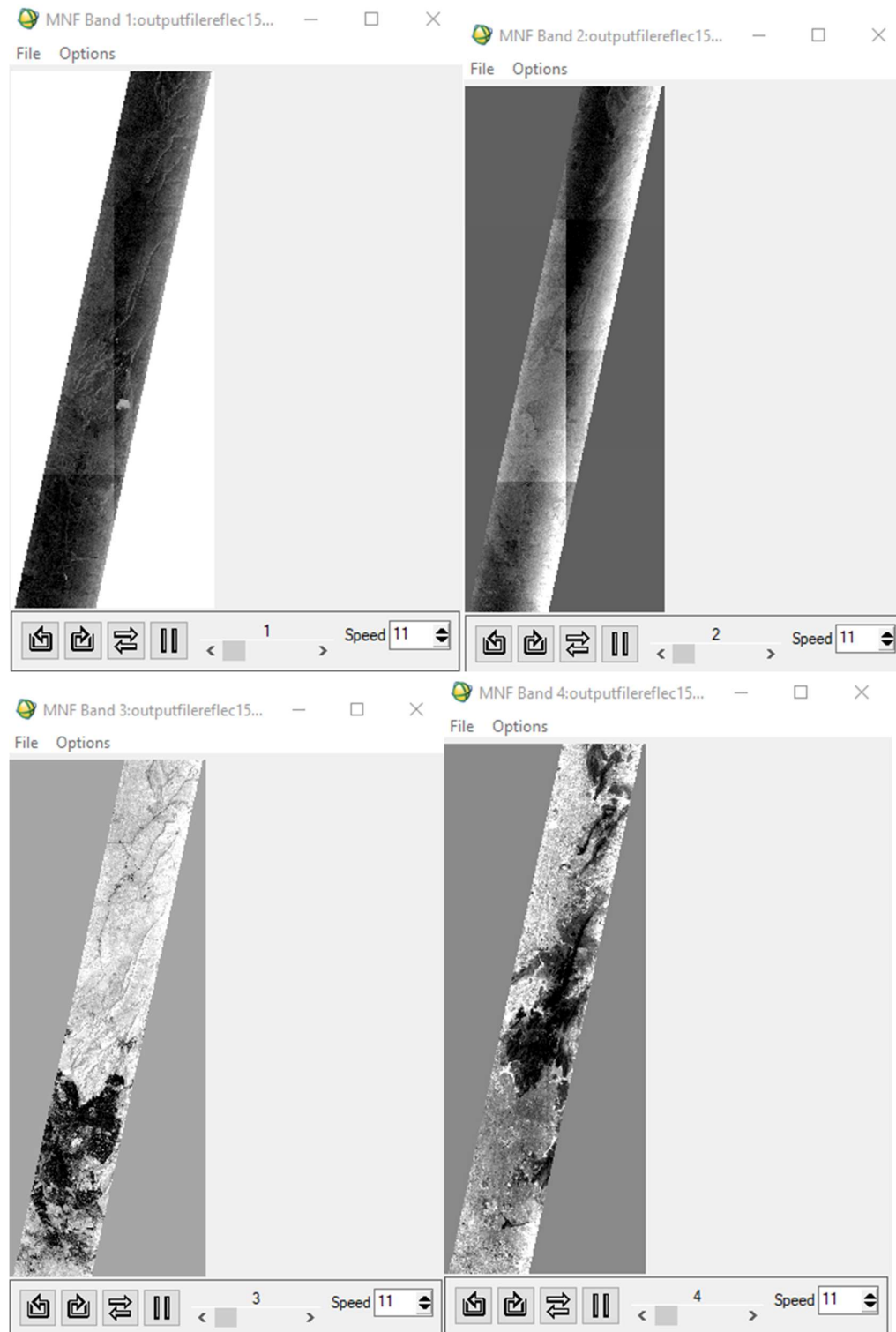


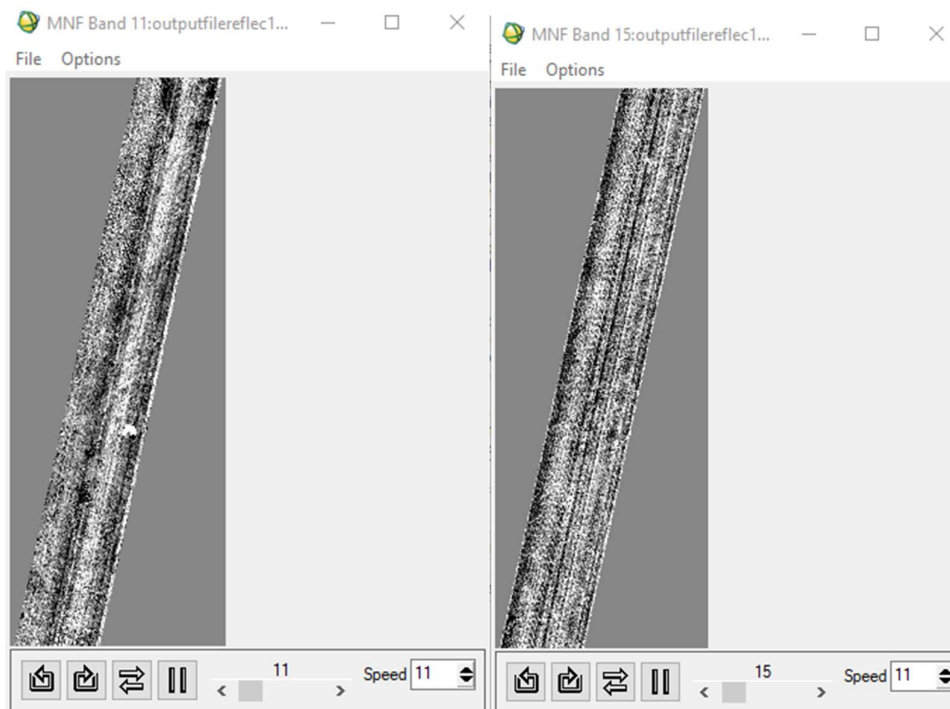
Larger eigenvalues indicate higher data variance in the transformed band and may help to indicate data dimensionality.

This graph shows MNF Band value gets poorer from BAND 15 and onwards .

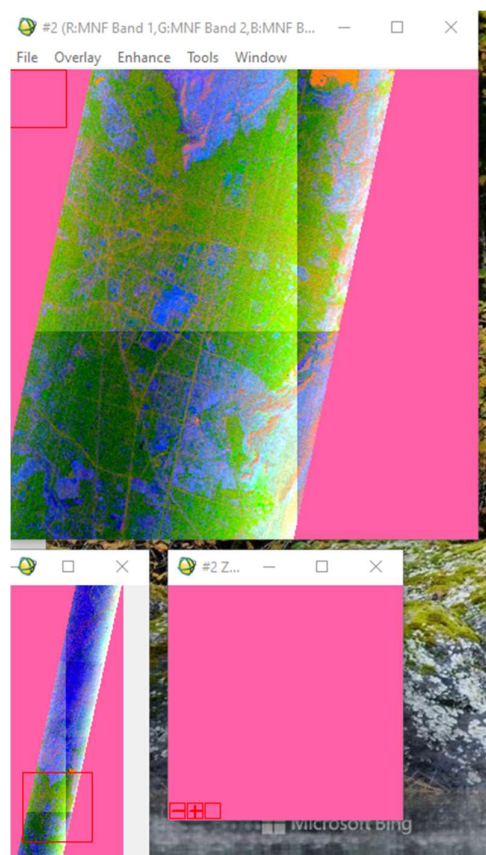


## MNF result as a gray scale animation

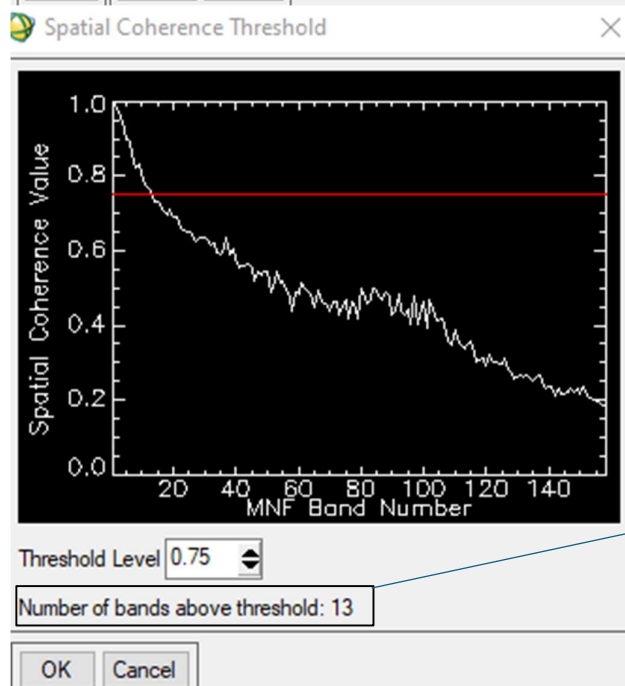
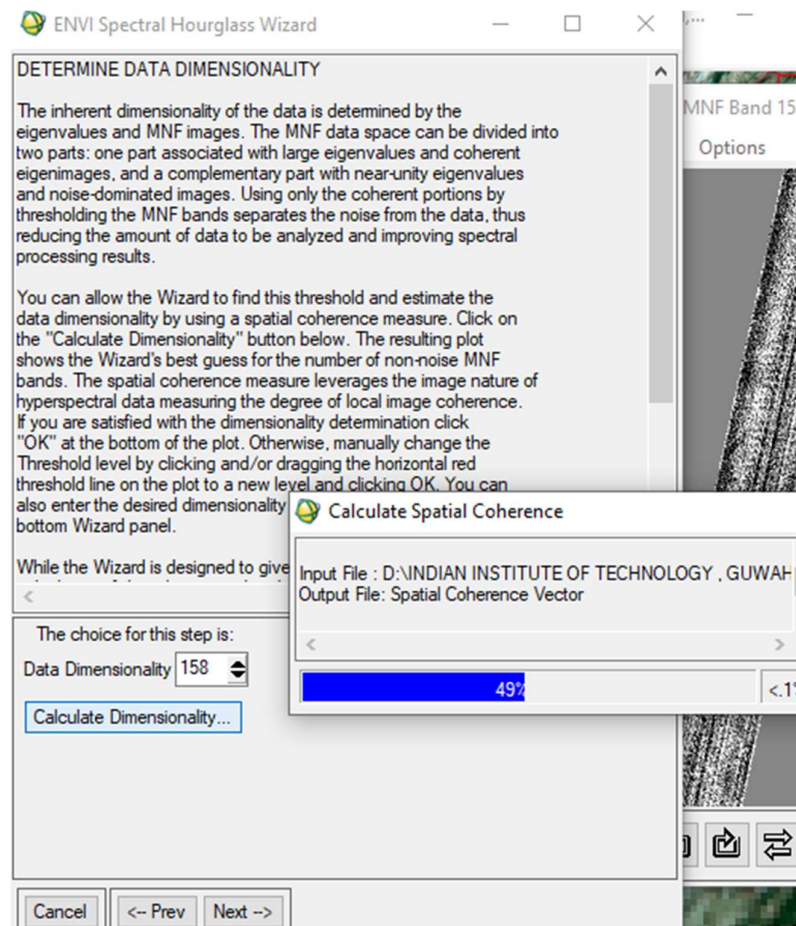




## Output MNF Band



## Data dimensionality



The lower MNF bands are expected to have spatial structure and contain most of the information, while higher MNF bands are expected to have little spatial structure and contain most of the noise.

By retaining only the coherent MNF bands and discarding those that are indistinguishable from the noise, the data set is reduced to its inherent dimensionality. This should improve spectral processing results.

ENVI Spectral Hourglass Wizard

### DERIVE OR SELECT ENDMEMBERS

You can choose to have the Wizard derive the endmembers used to classify or unmix the image directly from the input data or you can enter endmember spectra from another source such as a spectral library, regions of interest, ASCII file, or statistics file. Calculating or using endmembers from the image data directly usually produces better results than selecting library endmembers because the image spectra accurately account for any errors in calibration or atmospheric correction, the scales of mixing that occur in your data and sensor response effects.

The choice for this step is:

Derive Endmembers from Image ? ☒ Yes ☐ No

Cancel <-- Prev Next -->

This step helps derive the end member from the Input Image)

## Pixel Purity Index

ENVI Spectral Hourglass Wizard

to find more extreme pixels but they are less likely to be "pure" endmembers.

The PPI is highly CPU intensive and can take a long time to run, with the size of the data cube and the number of iterations being the main controlling factors. Reducing the number of PPI iterations can also save time, however, a minimum of 5000 iterations is usually required to produce useful results. The Maximum Memory Use must correspond to a value smaller than your available system memory (RAM).

The PPI process exploits convex geometry concepts in the n-dimensional data space of the MNF-processed hyperspectral data. Just as a ternary mixing diagram in Chemistry 101 shows the three pure endmembers at the corners of the feasible mixing space, in our n-dimensional data with many endmembers, the purest pixels still must be on the extreme corners or edges of the data cloud. While n-dimensional data may be hard to visualize and imagine, the PPI process is relatively insensitive to dimensionality and acts as an n-dimensional "rock tumbler" or "clothes dryer", tumbling the n-d cloud of data points and counting how many times each pixel is "hit" in this tumbling process. The purer the pixel, the more convex the data cloud is at that location and as a result it will be hit more often and receive a PPI score higher than a less-pure pixel. This convexity concept is based on the assumption

The choices for this step are:

Number of PPI Iterations: 25000

PPI Threshold Value: 2.500

PPI Maximum Memory use: 150 Mb

Cancel <-- Prev Next -->

The PPI is highly CPU-intensive and can take a long time to run, with the size of the data cube and the number of iterations being the main controlling factors. A minimum of 5,000 iterations is usually required to produce useful results

The PPI Threshold Value is used to define how many pixels are marked as extreme at the ends of the projected vector.



ENVI Spectral Hourglass Wizard

### CALCULATING PIXEL PURITY INDEX

The Pixel Purity Index is calculated for the number of iterations set on the previous panel. A Pixel Purity Index Plot appears showing which iteration you are on and the cumulative number of pixels that have been found to be extreme. The curve in this plot usually starts steeply, as new pixels are found in each iteration, and should flatten out as all the extreme pixels have been found. When the set number of iterations have been completed you can return to the previous panel and add more iterations if the plot has not flattened to near horizontal. While it is difficult or impossible to precisely say how many iterations are "enough" it is also true that you can never run "too many" iterations. Having many, many iterations (say 20,000) will give a PPI result with increased dynamic range and thus the ability to find subtle, poorly expressed endmembers that might be undetected if fewer iterations were completed.

When the specified iterations are completed, a PPI image is created in which the value of each pixel corresponds to the number of times that pixel was recorded as extreme. Bright pixels in the PPI image generally are image endmembers. This image is listed in the Available Bands List.

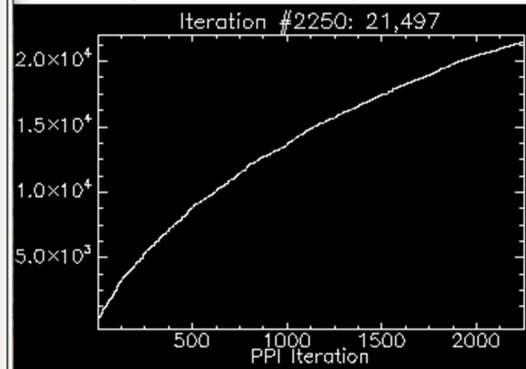
Cancel

<-- Prev

Next -->

Pixel Purity Index Plot

File Edit Options Plot\_Function Help



Fast Pixel Purity Index Calculation

Input File : D:\INDIAN INSTITUTE OF TECHNOLOGY, GUWAHATI  
Output File: D:\INDIAN INSTITUTE OF TECHNOLOGY, GUWAHATI

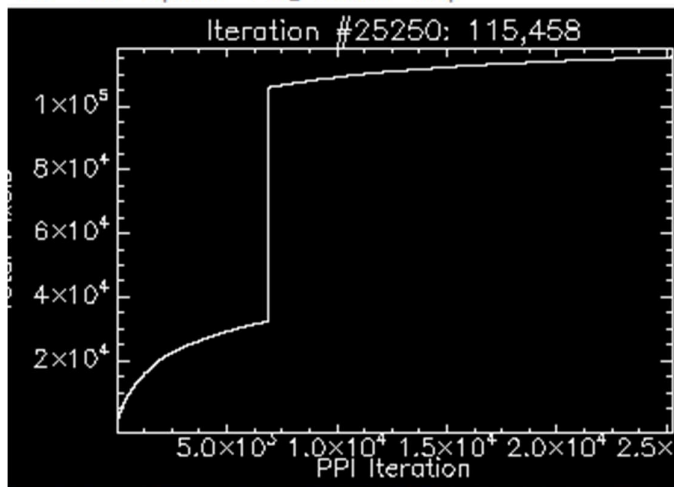
Cancel

9%

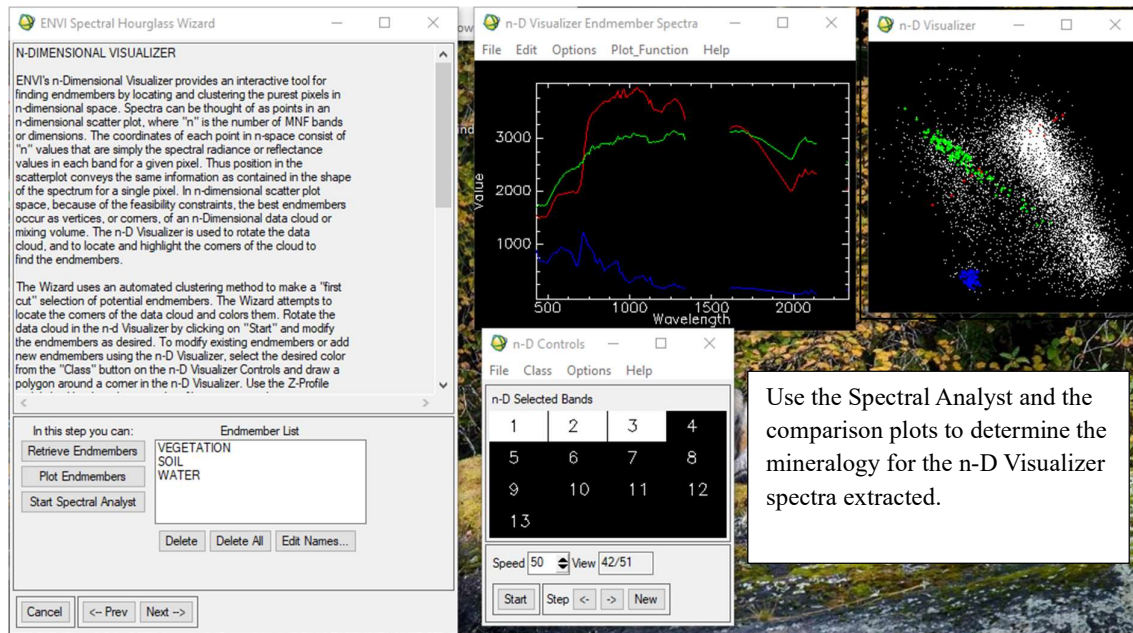
<.1%

Pixel Purity Index Plot

File Edit Options Plot\_Function Help



The curve in this plot usually starts steeply, as new pixels are found in each iteration



**ENVI Spectral Hourglass Wizard**

**USER SUPPLIED ENDMEMBERS**

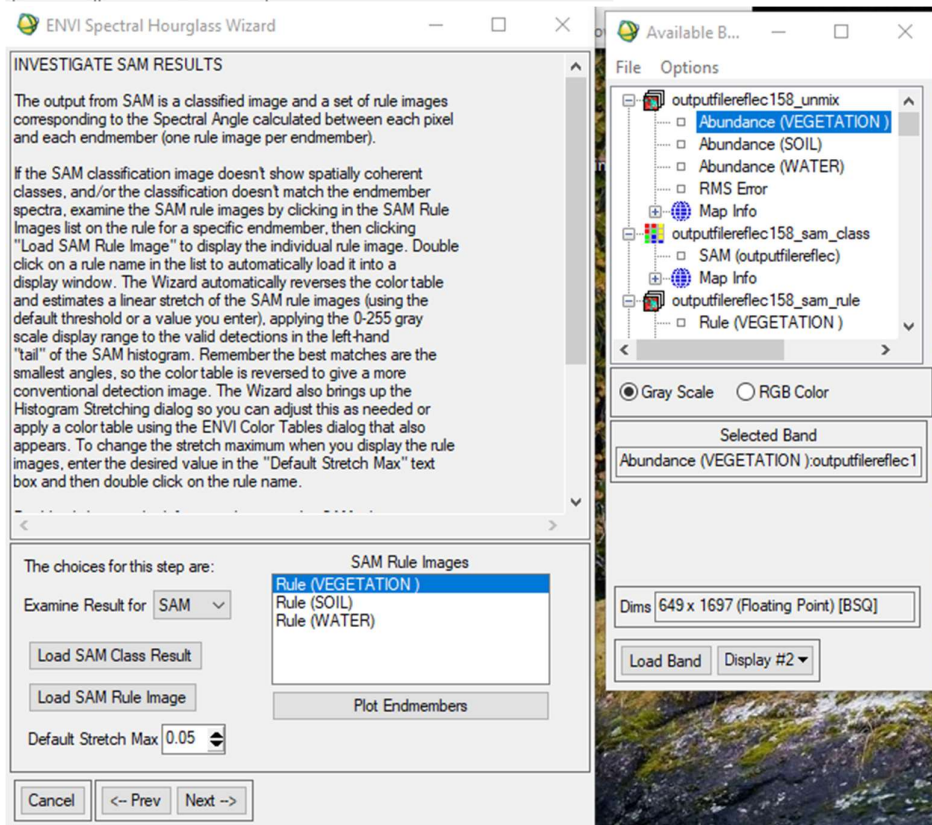
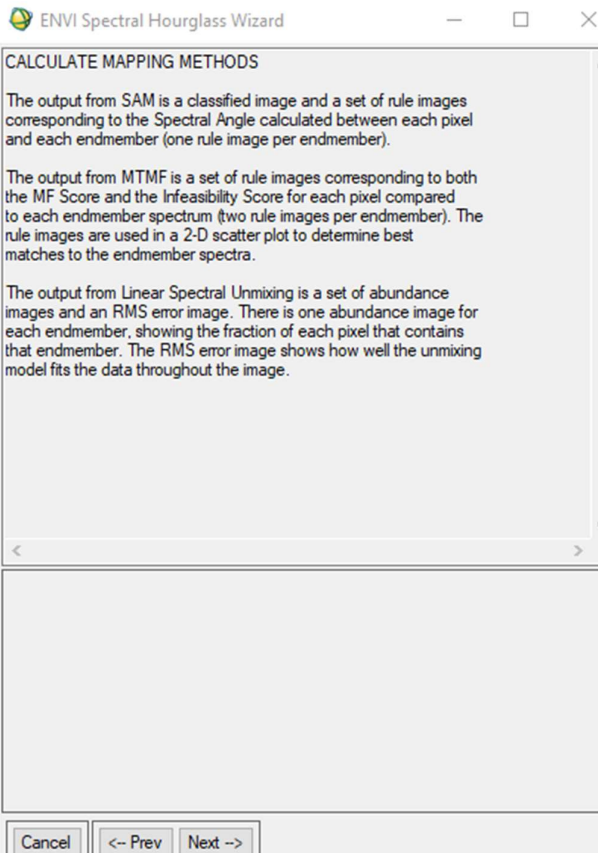
You have the option to manually enter endmembers to be used in the mapping process. The Endmember Collection dialog allows you to add endmembers from spectral libraries, individual spectral plots, text files, Regions of Interest, or Statistics files. All endmember spectra must be in the same units, and use the same scaling, as the image data (e.g. reflectance vs radiance).

The choice for this step is:

User supplied Endmembers ? No ☐

Cancel <-- Prev Next -->

**N-D visualizer** and control helps to locate, identify, and cluster the purest pixels and most extreme spectral responses in a data set (PPI image)





## Results

Red – Vegetation

Green – Barren Land

Blue - Water

