## PROBLEM 1

Sphere *floating* in water, i. e. the sphere is still.

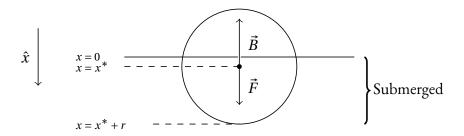


FIGURE 1: Sphere in equilibrium.

The sphere has a density of 0.6 times that of water. Archimedes' principle states that the buoyant force exerted og the sphere is equal (in magnitude) to the *weight* of the displaced fluid. Using density  $\rho = 1$  for water, the weight of the displaced fluid is

$$-\vec{B} = m_{\text{fluid}} \cdot g = v_{\text{fluid}} \cdot g,$$

and we know that the weight of the sphere itself is

$$\vec{F} = m_{\text{sphere}} \cdot g = 0.6 \cdot \nu_{\text{sphere}} \cdot g.$$

At equilibrium, these balance out:

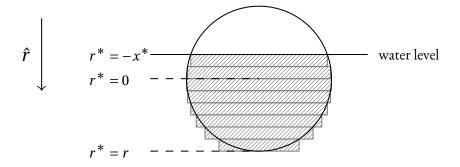
$$0 = \sum_{\vec{f} \text{ force}} \vec{f} = \vec{F} - \vec{B} = g(0.6\nu_{\text{sphere}} - \nu_{\text{fluid}})$$

Obviously, the volume of displaced fluid is equal to the volume of the part of the sphere that is submerged. We divide the equation by g, and thus

$$v_{\text{submerged}} = v_{\text{fluid}} = 0.6 \cdot v_{\text{sphere}}$$
 (1)

we also know that

$$v_{\text{sphere}} = \frac{4\pi r^3}{3} \tag{2}$$



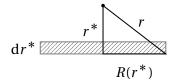


FIGURE 2: Volume of the submerged part of the sphere.

and with some analysis, we can work out the volume of the submerged part of the sphere. Our strategy is to integrate over cylinder-shaped volume elements.

Each volume element has a volume of  $\pi \cdot R(r^*) \cdot dr^*$ , and we need to integrate from water level and downward in a coordinate system on the sphere. Thus,

$$v_{\text{submerged}} = \pi \int_{r^* = -x^*}^{r^* = r} R(r^*)^2 dr^* \quad \text{where} \quad R(r^*) = \sqrt{r^2 - (r^*)^2}.$$
 (3)

We now have our mathematical model. It remains to find a solution  $x^*$  such that

$$I(x^*) := \frac{v_{\text{submerged}}}{v_{\text{sphere}}} = \frac{3}{4\pi r^3} \pi \int_{r^* = -x^*}^{r^* = r} R(r^*)^2 dr^* = 0.6$$

Given an algorithm to compute definite integrals, this can probably already be solved quite easily, but we don't know any such algorithms yet, so our strategy is to evaluate the integral by hand, and try to create a

root-finding problem.

All that remains is to solve the equation I(x\*) = 0.6, or equivalently

$$I(x^*) - 0.6 = \frac{x^*}{4r} \left( 3 - \left(\frac{x^*}{r}\right)^2 \right) + \frac{1}{2} - 0.6 = 0$$

Our analysis reveals that  $I(x^*)$  is a polynomial of degree three, which is good because its image is all of **R**, so a root like this is guaranteed to exist. *Unfortunately, three roots exist*. We need to define a criteria to determine which solution is the correct one.

#### RECALL

 $x^*$  is the position of the center of the sphere along the  $\hat{x}$ -axis. Clearly,  $x^* < -r$  is not a solution that can correspond to anything in the real world, as this means the sphere is hovering in mid-air. Similarly,  $x^* > r$  is not a very good solution, since this is possible only if the sphere has density  $\rho \ge 1$ , and by assumtion  $\rho = 0.6$ , hence a physical solution must exist in the interval [-r, r].

#### Approximating Solutions

We know that if the model corresponds to a physical situation, then a solution must exist in the interval [-r, r]. We also know that there exists exactly three solutions in all of **R**. If we can find two impossible solutions *outside of out interval*, we can be sure that the solution is unique in our interval!

If we factor our equation, we see that

$$I(x^*) - 0.6 = \frac{x^*}{4r} \left( 3 - \left( \frac{x^*}{r} \right)^2 \right) + \frac{1}{2} - 0.6$$

now, we look for solutions to the equation  $I(x^*) - 0.5 = 0$ . These solutions will be quite close to the solutions of  $I(x^*) - 0.6$ , and are super easy to find.

$$I(x^*) - \frac{1}{2} = \frac{x^*}{4r} \left( 3 - \left( \frac{x^*}{r} \right)^2 \right) = 0 \implies x^* = 0 \quad \text{or} \quad 3 - \left( \frac{x^*}{r} \right)^2 = 0.$$

This yields two solutions,

$$x^* = \pm \sqrt{3}r$$

which are both far outside the interval [-r, r]. Not only does this mean that it is safe to search for the original solution on this interval by bisection, but the last solution  $(x^* = 0)$  of the simplified problem tells us that it is going to be smart to start searching close to zero, which may helps us approximating the solution faster.

This does *not* prove that Newtons method will converge on the *entire* interval, but starting at  $x_0 = 0$ , it will, and we will find the solution very quickly.

#### PROBLEM 2

a. Want to use the Taylor series of f around the point a,  $T_a\{f\}$  (x), to show that

$$\frac{\mathrm{d}f}{\mathrm{d}x} = \frac{f(x + \Delta x) - f(x)}{\Delta x} + \mathcal{O}(\Delta x),$$

in other words, that the error we expect if we compute

$$\frac{\mathrm{d}f}{\mathrm{d}x} \approx \frac{f(x + \Delta x) - f(x)}{\Delta x},$$

is proportional to  $\Delta x$ , i. e. the error goes to zero as  $\Delta x$  approaches zero.

Fix x and  $\Delta x$ . We expand the Taylor series to two terms, also including the error term, around the point x.

$$f(x^*) = T_x \{f\} (x^*)$$
  
=  $f(x) + f'(x)(x - x^*) + \text{error}(2)$ 

Which yields

$$f'(x) = \frac{f(x^*) - f(x) - \text{error}(2)}{x - x^*}$$

By evaluating the series at  $x^* = x + \Delta x$ 

$$f'(x) = \frac{f(x + \Delta x) - f(x)}{\Delta x} - \frac{\text{error(2)}}{\Delta x}$$

To show that the error term is  $\mathcal{O}(\Delta x)$  is straight forward:

$$E(\Delta x) := \frac{\text{error}(2)}{\Delta x} = \frac{1}{\Delta x} \frac{f''(\tilde{x})}{2!} (x^* - x)^2$$

for some  $\tilde{x} \in [x, x^*]$ . Again, we evaluate the error term at  $x^* = x + \Delta x$ .

$$= \frac{f''(\widetilde{x})}{2\Delta x} (\Delta x)^2$$
$$= \frac{f''(\widetilde{x})}{2} \Delta x$$

As long as f'' does not blow up on  $I = [x, x + \Delta x]$ , because og the fact that I is bounded and closed, f'' attains its maximum and minimum on I. Thus we can let

$$L = \operatorname{diam} f(I) = \max_{x, x' \in I} |f''(x) - f''(x')|$$

Which makes

$$|E(\Delta x)| \le L\Delta x$$

### PROBLEM 3

# PROBLEM 4