

A Road Map to Mitigating Chromatic Timing Variations for NANOGrav

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NANOGrav's Goal: Mitigate Chromatic Effects of the Interstellar Medium on Pulsar Timing

Dispersion

Time delay between two frequency components of a pulse due to free electrons in the ISM:

$$\Delta t_{\text{DM}} \approx 4.15 \times 10^6 \text{ ms} \times \text{DM} \left(\frac{1}{\nu_1^2} - \frac{1}{\nu_2^2} \right)$$

$$\text{DM}(t) = \int_0^{L(t)} n_e(s) \hat{n}(t), t ds$$
[1]

Birefringence

Time delay due to magnetized medium:

$$\pm \Delta t_{\text{RM}} = 28.6 \text{ ps} \times \text{RM} / \nu^3$$

$$\text{RM} = \frac{e^3}{2\pi m_e^2 c^4} \int_0^L n_e B_{||} ds$$
[2]

Pulse Broadening

Multi-path scattering causes pulse broadening with mean delay:

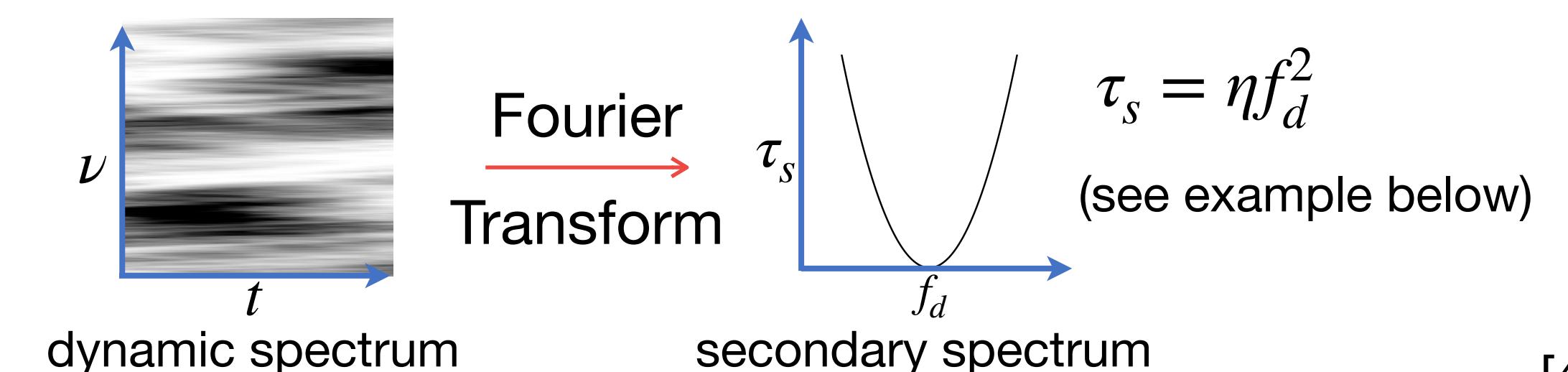
$$\tau_d = 1/2\pi\delta\nu_d \propto \nu^{-4.4} \quad (\text{Kolmogorov})$$

The pulse broadening function is:

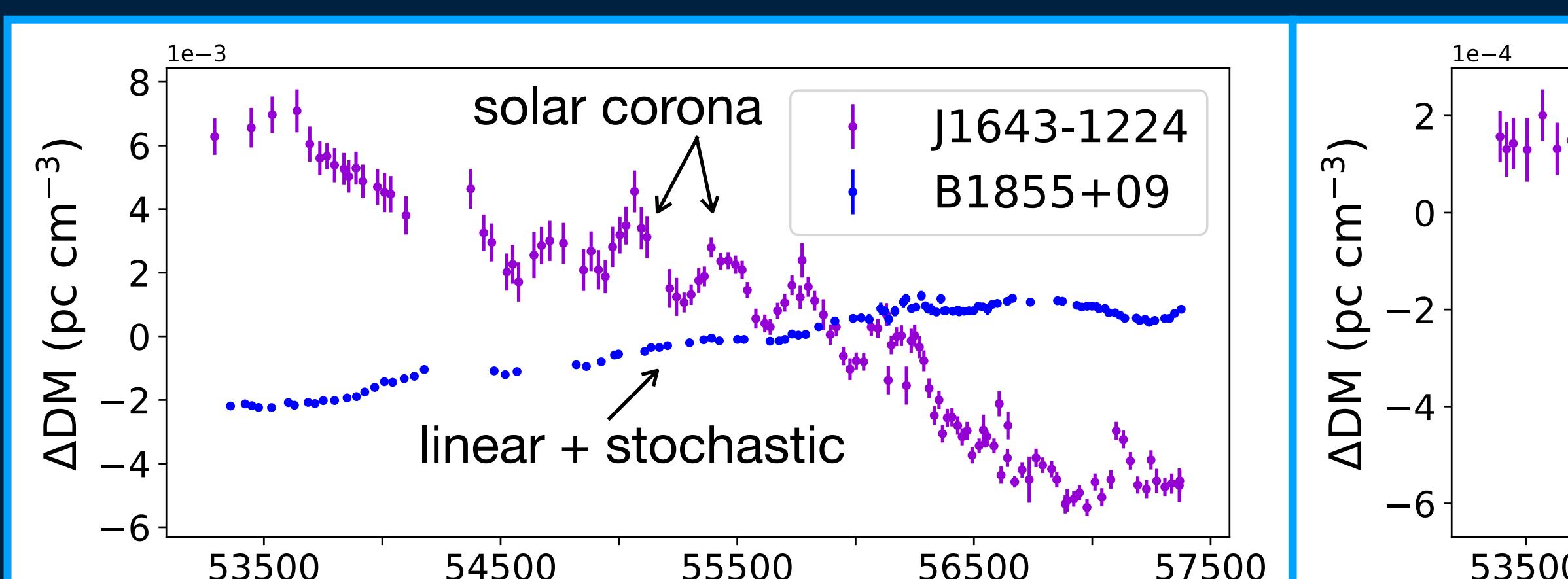
$$\text{PBF}(t) \propto e^{-t/\tau_d}$$
[3]

Scintillation

Scattering induced phase delays cause interference between ray paths. In the Fourier domain the interference pattern is a parabolic arc:

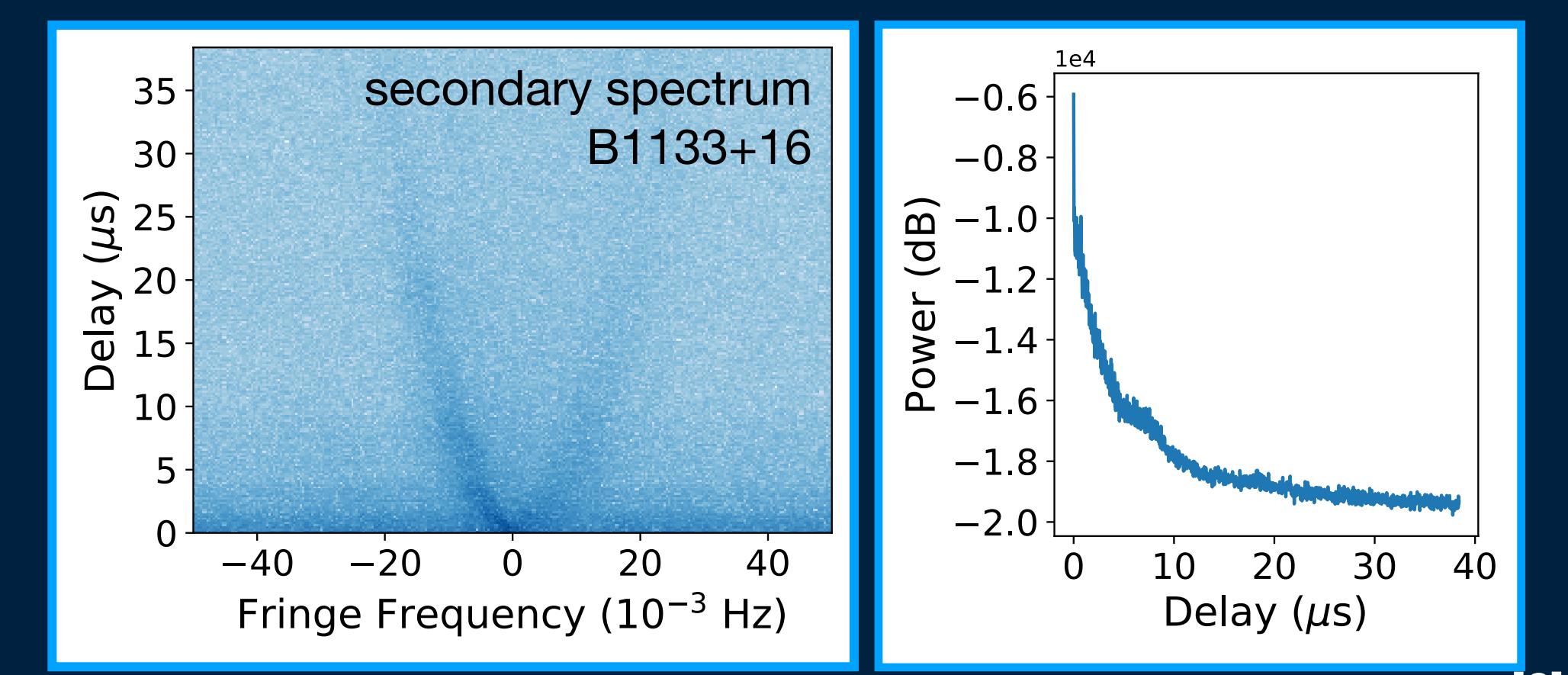


Time-Dependent Dispersion



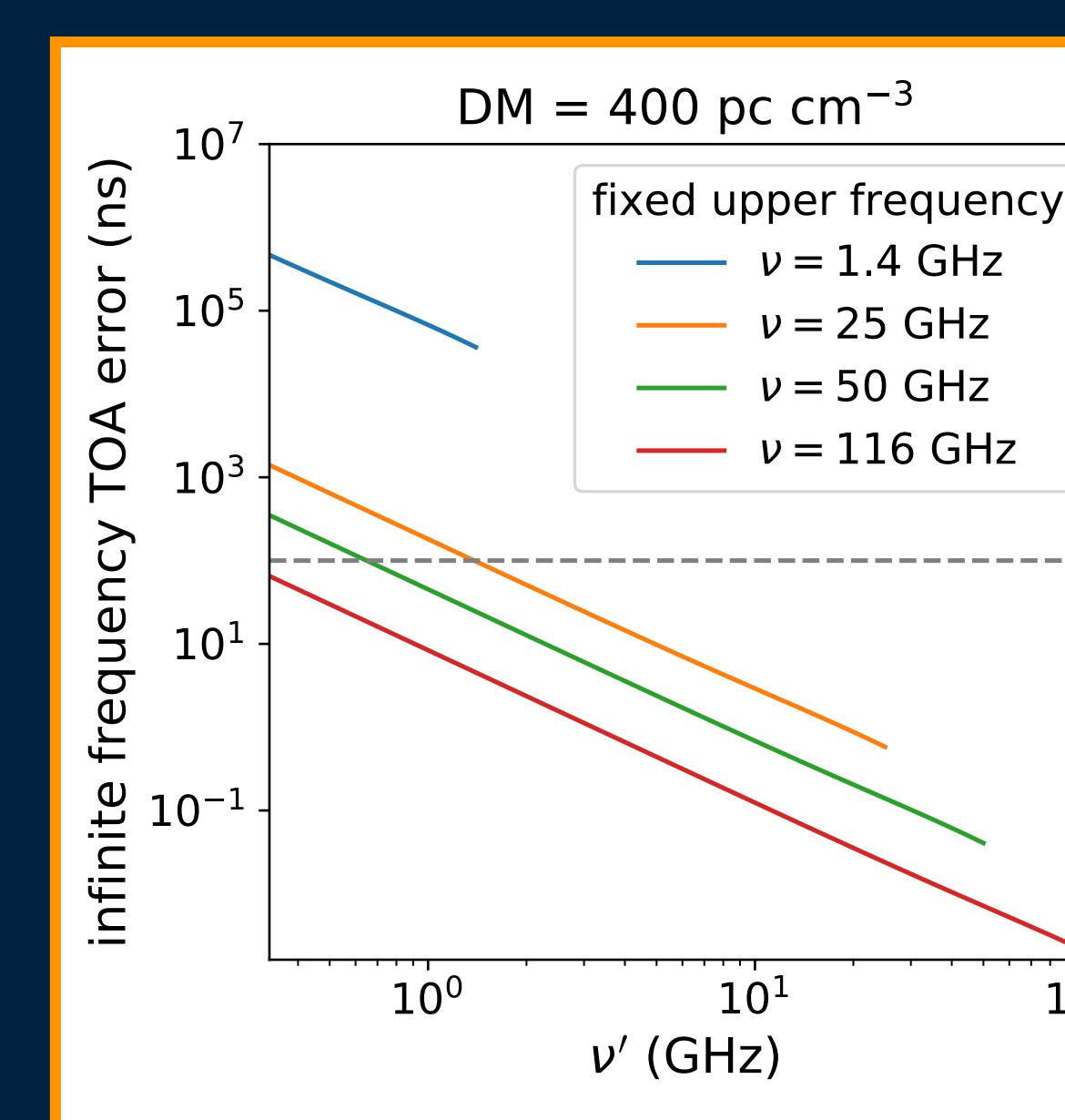
Stochastic Impulse Response Effects

- Observed pulse is a convolution of emitted pulse & scattering impulse response PBF(t).
- Extent of scintillation arc in delay = width of PBF(t).
- Scattering might still be important in observations where pulse broadening seems small.

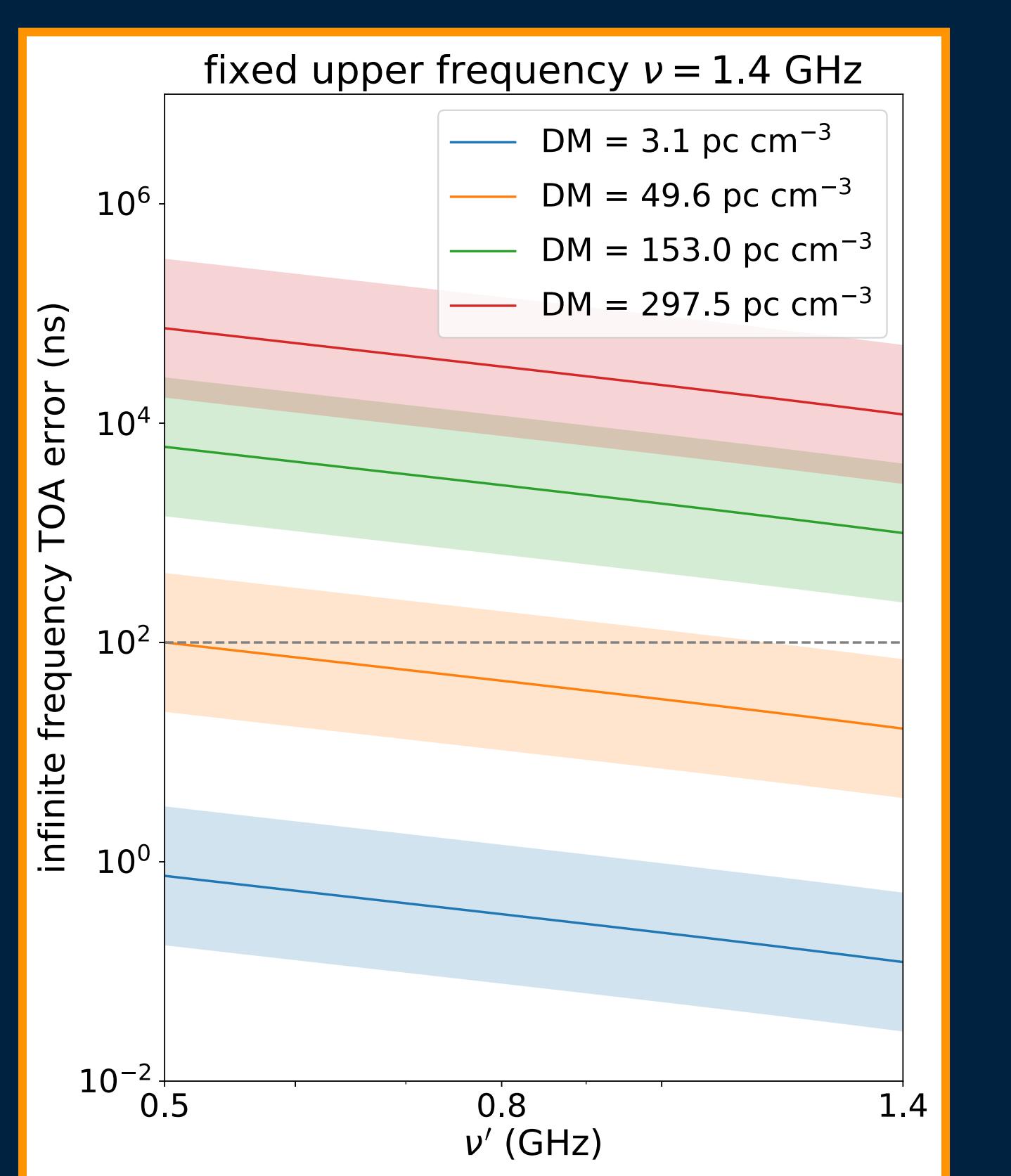


Frequency-Dependent Dispersion

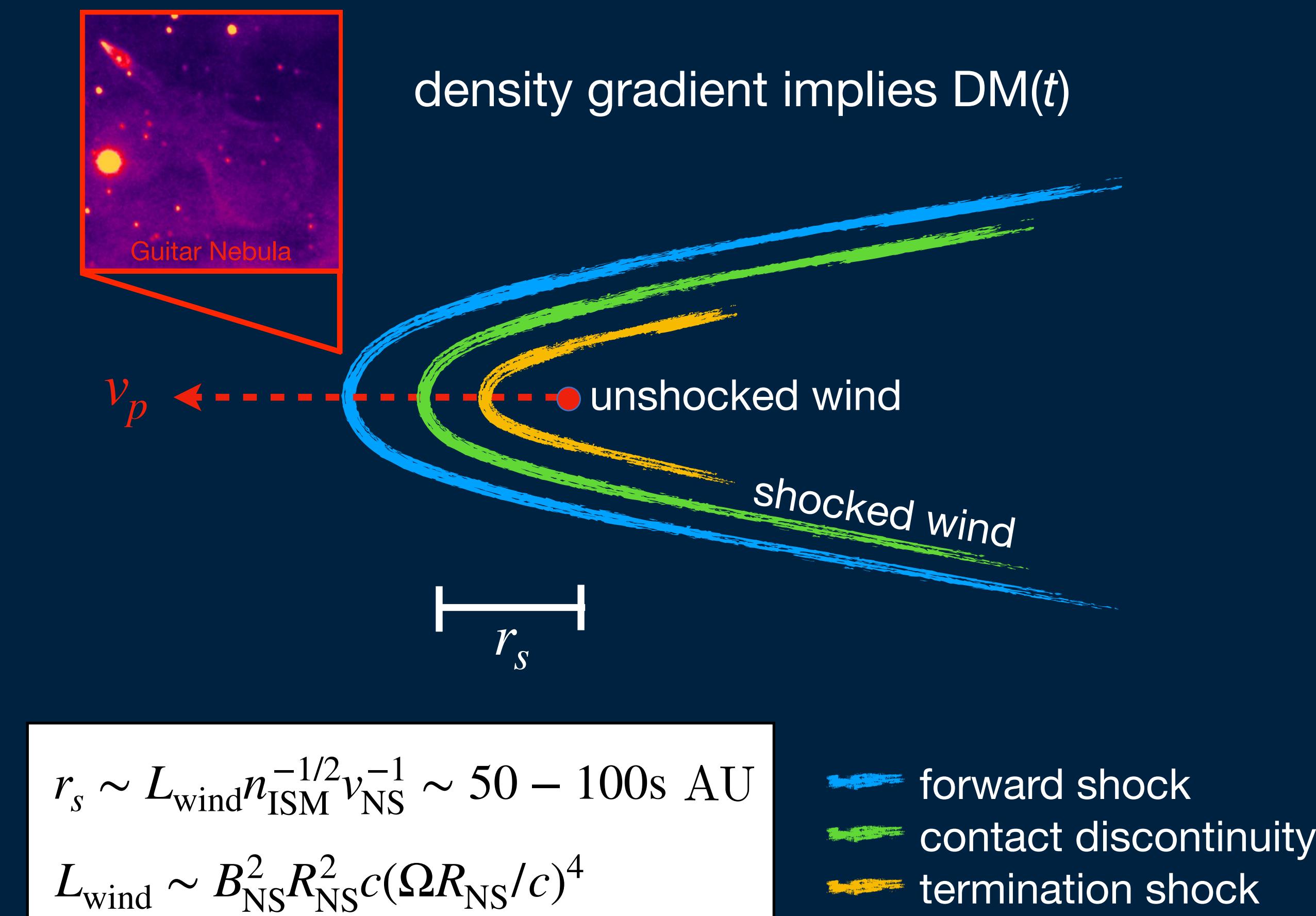
- Multi-path propagation causes frequency-dependent DMs.
- This may ultimately limit the highest precision TOAs.
- Mitigation: higher frequencies (e.g., ngVLA)



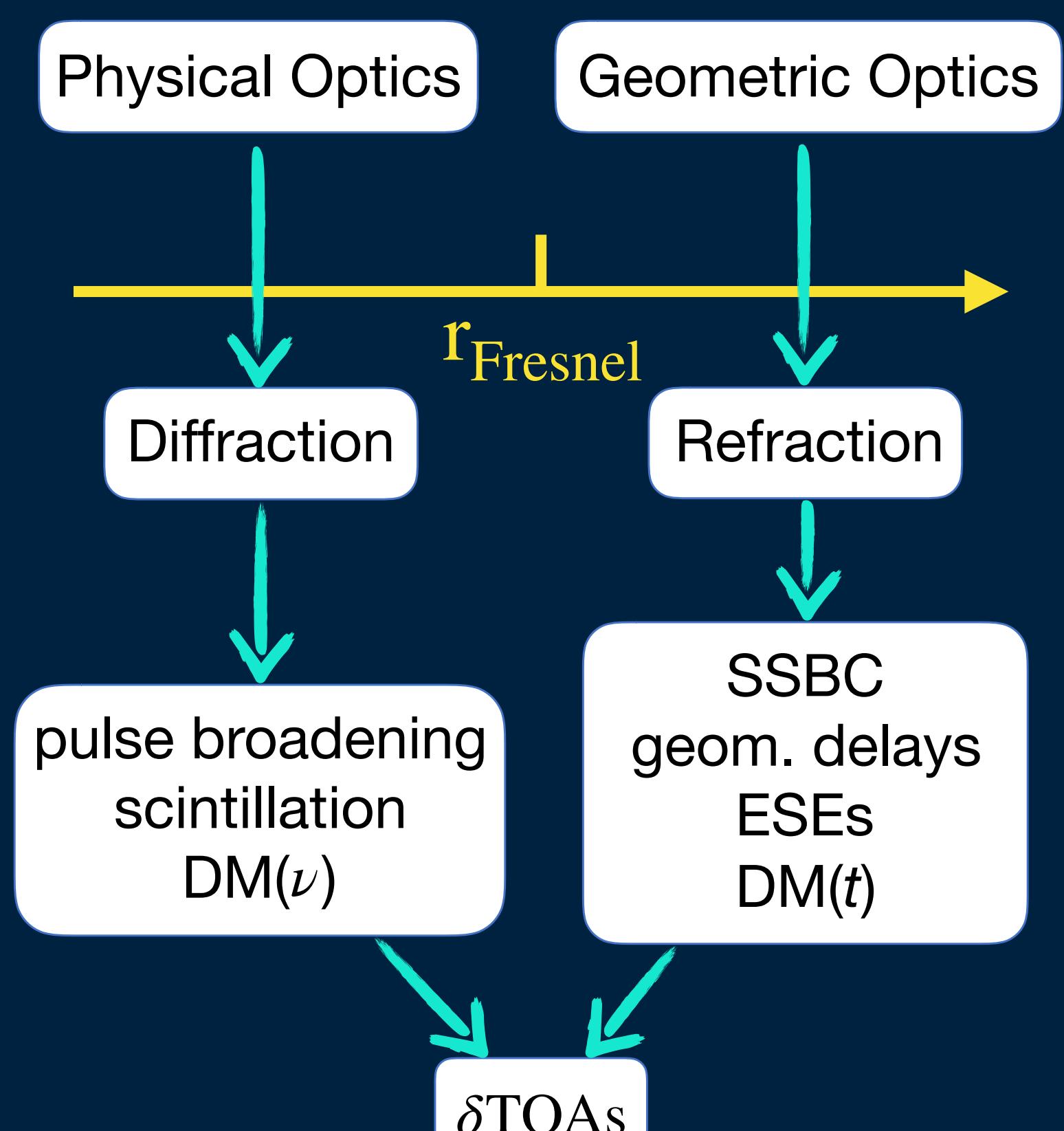
TOA error is lower for high upper frequencies and low DMs



Pulsar Bow Shocks & Wind Nebulae



Propagation Regimes



References: [1] Lam et al. 2016, ApJ, 821, 66 [2] Suresh & Cordes 2019, ApJ, 870, 29 [3] Rickett 1990, ARAA, 28, 561-605 [4] Stinebring et al. 2001, ApJ, 549, p. L97-L100 [5] Jones et al. 2017, ApJ, 841, 125 [6] Hemberger & Stinebring 2008, ApJ, 674 [7] Cordes et al. 2016, ApJ, 817, 16 [8] Barkov et al. 2019, MNRAS, 484, p.4760-4784

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