

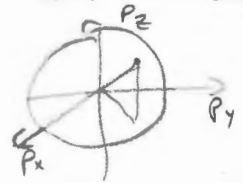
Bose-Einstein Condensation

We've seen what happens when our particles are massless and $\mu=0$ (photons). What happens with massive bosons?

We assume we have spin-0 bosons of mass m that are non-relativistic, i.e.

$$\epsilon = \frac{p^2}{2m} = -\frac{\hbar^2}{2m} \nabla^2 \rightarrow \text{sphere in } p \text{ space}$$

$$= \frac{\hbar^2 k^2}{2m} = \frac{2\pi^2 \hbar^2}{mL^2} (n_x^2 + n_y^2 + n_z^2) \quad n_{x,y,z} = 0, 1, 2, \dots$$



Let's put them in our box of volume $V = (L)^3$. As before for the photons, the number of plane wave states per unit volume in k space is $\frac{V}{8\pi^3}$. So in p space ($\vec{p} = \hbar \vec{k}$),

$$\rho(k) = \frac{V}{(2\pi)^3} \Rightarrow \rho(p) = \frac{V}{(2\pi\hbar)^3}$$

Again making the continuum approximation, the number of plane wave states in a volume of width $d\epsilon$ is

$$g(\epsilon) d\epsilon = (\text{surface area of sphere}) (\text{width of sphere}) \cdot (\text{density of states})$$

$$= (4\pi p^2) \cdot dp \cdot \left(\frac{V}{(2\pi\hbar)^3} \right) = (4\pi p^2) \left(\frac{dp}{d\epsilon} d\epsilon \right) \left(\frac{V}{(2\pi\hbar)^3} \right)$$

$$\epsilon = \frac{p^2}{2m} \Rightarrow d\epsilon = \frac{|\vec{p}|}{m} dp, \quad p = \sqrt{2m\epsilon}$$

$$\therefore \frac{dp}{d\epsilon} = \frac{m}{p} = \sqrt{\frac{m}{2\epsilon}}$$

$$\Rightarrow g(\epsilon) d\epsilon = [4\pi(2m\epsilon)] \left[\sqrt{\frac{m}{2\epsilon}} d\epsilon \right] \left[\frac{V}{(2\pi\hbar)^3} \right]$$

$$g(\epsilon) d\epsilon = \frac{V m^{3/2}}{\sqrt{2} \pi^2 \hbar^3} \sqrt{\epsilon} d\epsilon$$

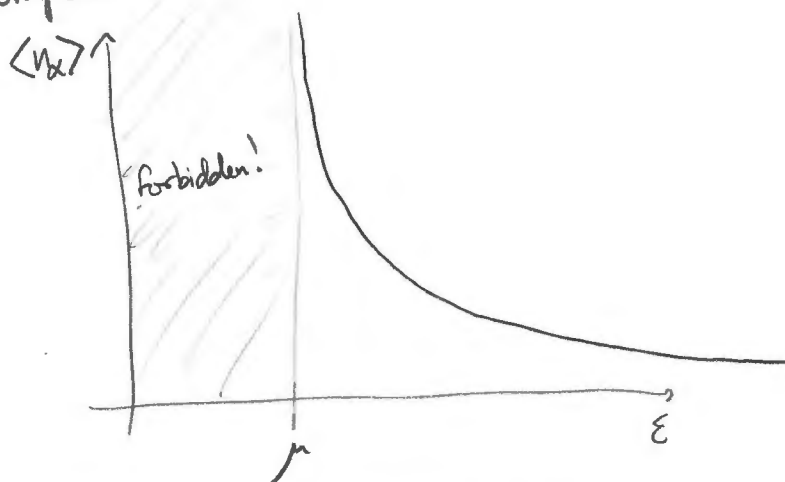
Now we fill the box up with bosons. The total number of particles is

$$\begin{aligned}\langle N \rangle &= \int \langle n_{\epsilon} \rangle g(\epsilon) d\epsilon \\ &= \int \frac{g(\epsilon)}{e^{\beta(\epsilon-\mu)} - 1} d\epsilon\end{aligned}$$

Ok, now an aside. Remember we derived $\langle n_{\alpha}^{BE} \rangle$ as a sum of a geometric series:

$$Z_{\alpha} = \sum_{n_{\alpha}=0}^{\infty} e^{-\beta(\epsilon_{\alpha}-\mu)n_{\alpha}} = \frac{1}{1 - e^{-\beta(\epsilon_{\alpha}-\mu)}}$$

This only converges if $\epsilon_{\alpha} > \mu$, i.e. the lowest energy level ϵ_0 must be higher than μ . Otherwise we can lower the energy of the system by adding more particles and everything collapses.



Now, for a given μ , $\langle N \rangle$ particles will fill the system at equilibrium. If we forcefully add another particle, we must supply energy μ to add it. This effectively raises μ for the next particle, and so on. Also, as we add more particles, the density $\frac{\langle N \rangle}{V}$ goes up

If we keep adding particles until $\mu = \epsilon_0$. Here $\langle n_0 \rangle$ diverges and the density heads to infinity.

This is Bose-Einstein condensation.

Interestingly, our integral for $\langle N \rangle$ (assuming a continuum of states) still converges for $\mu=0$:

$$\begin{aligned}\langle N \rangle &= \int \frac{g(\epsilon)}{e^{\beta\epsilon} - 1} d\epsilon \\&= \frac{V m^{3/2}}{\sqrt{2} \pi^2 \hbar^3} \int_0^\infty d\epsilon \frac{\sqrt{\epsilon}}{e^{\beta\epsilon} - 1} \\&= V \underbrace{\left(\frac{m k_B T}{2\pi \hbar^2} \right)^{3/2}}_{1/\lambda_T^3} \cdot \frac{2}{\sqrt{\pi}} \int_0^\infty \frac{\sqrt{z}}{e^z - 1} dz \quad z = \beta\epsilon\end{aligned}$$

Now,

$$\frac{2}{\sqrt{\pi}} = \frac{1}{(1/2)!} = \frac{1}{\Gamma(\frac{3}{2})}$$

So we have

$$\begin{aligned}\frac{2}{\sqrt{\pi}} \int_0^\infty \frac{z^{1/2}}{e^z - 1} dz &= \frac{1}{\Gamma(\frac{3}{2})} \int_0^\infty \frac{z^{\frac{3}{2}-1}}{e^z - 1} dz \\&\equiv \zeta\left(\frac{3}{2}\right)\end{aligned}$$

So

$$\langle N \rangle = V \lambda_T^{-3} \zeta\left(\frac{3}{2}\right)$$

Hence when the density gets higher than

$$\frac{\langle N \rangle}{V} = \frac{\zeta(\frac{3}{2})}{\lambda_T^3} = \frac{2.612 \text{ particles}}{\text{de Broglie volume}}$$

Usually, the way this is done is not to add particles at fixed temp, but to lower the temp at fixed N . We massage our equation to yield

$$T_c^{\text{BEC}} = \frac{2\pi\hbar^2}{m k_B} \left(\frac{N}{V \zeta(\frac{3}{2})} \right)^{2/3}$$

This is the critical temperature at which BEC occurs.