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On the asymptotic magnitude of subsets of Euclidean space

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Abstract Magnitude is a canonical invariant of finite metric spaces which has its origins in category theory; it is analogous to cardinality of finite sets. Here, by approximating certain compact subsets of Euclidean space with finite subsets, the magnitudes of line segments, circles and Cantor sets are defined and calculated. It is observed that asymptotically these satisfy the inclusion-exclusion principle, relating them to intrinsic volumes of polyconvex sets.

Keywords Magnitude · Metric space · Euler characteristic · Intrinsic volumes

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Introduction

In [10] one of us introduced the notion of the Euler characteristic of a finite category and showed how it linked together various notions of size in mathematics, including the cardinality of sets and the Euler characteristics of topological spaces, posets and graphs. In [11,13] it was shown how to transfer this to a notion of 'magnitude' of a finite metric space, using the fact that a metric space can be viewed as an enriched category.

One way of viewing magnitude is as the 'effective number of points'. Consider, for example, the n-point metric space in which any two points are a distance d apart. When d is very small, the magnitude is just greater than 1—there is 'effectively only one point'. As d increases, the magnitude increases, and when d is very large, the magnitude is just less

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¹ The terms 'Euler characteristic' and 'cardinality' could have been used here, as in [10] and [11], but we have decided to use a word with less mathematical ambiguity.

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than n—there are 'effectively n points'. The magnitude of a finite metric space actually first appeared in the biodiversity literature [17], under the name 'effective number of species', although its mathematical properties were hardly explored. It should be noted that, contrary to the simple example given above, magnitude can display wild behaviour of various types [13]: when a space is scaled up, its magnitude can sometimes *decrease*, and there are some exceptional finite metric spaces for which the magnitude is not well-defined.

In this paper—which requires no category theory—we consider the notion of magnitude for certain non-finite metric spaces, in particular for certain compact subsets of Euclidean space. This is done by approximating such a subset A with a sequence of finite subsets of A and taking the limit of the corresponding sequence of magnitudes. In the cases we consider here—circles, line-segments and Cantor sets—as the subset is scaled up this answer behaves like a linear combination of 'intrinsic volumes', such as the length and Euler characteristic, which satisfy the inclusion-exclusion principle. This leads us to conjecture that for a subspace A the magnitude |A| decomposes as follows:

$$|A| = P(A) + q(A)$$

where P is a function, defined on some class of subsets of Euclidean space, which satisfies $P(A \cup B) = P(A) + P(B) - P(A \cap B)$ and q(A) tends to zero as A is scaled bigger and bigger. In other words, the magnitude of subsets of Euclidean space asymptotically satisfies the inclusion-exclusion principle. Whilst the proof of this for some examples requires only elementary analysis, the proof for circles requires more subtle asymptotic analysis. Empirical calculations for some subsets in two and three dimensions are consistent with this and appear elsewhere [19].

In an earlier version of this paper we were less definite in some of our assertions. We calculated the limiting value of approximations of each of the spaces and claimed that that *should* be the value of the magnitude of the space in question, and that it should be independent of the choice of approximating sequence. Motivated by our work, Meckes [15] proved that this was indeed the case and that our methods do indeed give unique results (see Sect. 1.3 below). Another development since the time of writing is the use of "weight measures", this development means that several calculations presented here can be done more quickly, see [15] and [20]. However, the methods presented here seem to be more general as it is not known, and seems unlikely, that every compact metric space, or even every compact subset of Euclidean space, admits a weight measure.

There will now follow a more detailed description of the magnitude of metric spaces, the inclusion-exclusion principle and intrinsic volumes.

Asymptotic conjectures. Given a metric space X with n points one can try to associate to it an invariant called the magnitude; the definition is given in Sect. 1.2 and the definition is motivated by category theory in Sect. 1.1. If X is a metric subspace of some Euclidean space then it will have a well-defined magnitude (see Sect. 1.3). However, not every finite metric space has a well-defined magnitude, but every 'sufficiently separated' one does (Theorem 2). In particular, if for t > 0 we define tX to be X scaled by a factor of t, so that it is a metric space with the same points as X but with the metric defined by $d_{tX}(x, x') := td_{X}(x, x')$, then for t sufficiently large the magnitude |tX| is well-defined and $|tX| \to n$ as $t \to \infty$ (Theorem 3). So asymptotically, the magnitude is the number of points in the metric space.

Whilst the magnitude in the case of *finite* metric spaces is interesting, not least for its connections with biodiversity measures (see [12]), in this paper we consider extending this notion of magnitude to non-finite metric spaces, primarily in the form of compact subsets of Euclidean space with the subspace metric. In the cases we consider here one interesting



feature which emerges is that the magnitude seems to 'asymptotically' obey the inclusionexclusion principle, where 'asymptotically' means with regard to the space being scaled up larger and larger.

The inclusion-exclusion principle is embodied in the notion of a valuation. Let \mathcal{D} be some class of subsets of Euclidean space closed under binary intersections. A *valuation* on \mathcal{D} is a real-valued function P on \mathcal{D} satisfying the following:

• for all $A, B \in \mathcal{D}$ such that $A \cup B \in \mathcal{D}$ we have

$$P(A \cup B) = P(A) + P(B) - P(A \cap B);$$

 \bullet $P(\emptyset) = 0.$

Following Klain and Rota [8], we will often take \mathcal{D} to be the collection of *polyconvex sets* in \mathbb{R}^m —that is those subsets of Euclidean space which are finite unions of compact, convex sets. Similarly following Klain and Rota, we define an *invariant valuation* on \mathbb{R}^m to be a valuation on the set of polyconvex subsets of \mathbb{R}^m for which the following axioms hold.

- It is invariant under rigid motions.
- It is continuous on *convex* sets with respect to the Hausdorff topology.

Examples of such things include the Euler characteristic and the m-dimensional volume. Hadwiger's Theorem says that the vector space of invariant valuations on \mathbb{R}^m is (m+1)-dimensional and that there is a canonical basis $\{V_0,\ldots,V_m\}$ where each V_i is scaling-homogeneous in the sense that $V_i(tA) = t^i \ V_i(A)$ for any scaling t>0 and any polyconvex set A, and is normalized on cubes so that $V_i([0,1]^i)=1$. The valuation V_i is called the ith i

It might seem that the *i*th intrinsic volume in \mathbb{R}^m should be denoted by say V_i^m to remove ambiguity: however, they are normalized so that this is unnecessary, namely if $A \subset \mathbb{R}^m$ is a polyconvex set which actually is contained in a subspace $\mathbb{R}^p \subset \mathbb{R}^m$ then $V_i^p(A) = V_i^m(A)$.

In this paper we calculate the magnitude for some subspaces of Euclidean space. The asymptotic behaviour of these is summarized in Table 1. Looking at the data one might think that, restricting to polyconvex sets, the magnitude splits as

$$|A| = P(A) + q(A)$$

where $q(tA) \to 0$ as $t \to \infty$ and $P(A) = V_0(A) + \frac{1}{2}V_1(A)$ + higher order terms. In fact, plausibility arguments and some rough empirical results given in [19], together with some partial results in [13, Section 3.5] suggest

$$P(A) = \sum_{i=0}^{\infty} \frac{V_i(A)}{i! \omega_i} = V_0(A) + \frac{V_1(A)}{2} + \frac{V_2(A)}{2\pi} + \cdots,$$

where ω_i is the volume of the unit *i*-ball in \mathbb{R}^i . It is perhaps worth noting here that the natural normalization of the *i*-dimensional Hausdorff measure (see, for example, [5]) differs from the usual normalization of the *i*-dimensional Lebesgue measure by a factor of ω_i . So it might



² A three-dimensional rectangular box.

subspace A		magnitude $ A $
n points		$n+q_1(A)$
ternary Cantor set of length ℓ		$\int f(\ell)\ell^{\log_3 2} + q_2(\ell)$
closed interval of length ℓ		$\ell/2+1$
circle of circumference ℓ		$\ell/2 + q_3(\ell)$

Table 1 The asymptotic behaviour of the magnitude of some subsets of Euclidean space

Here $q_1(tA), q_2(t), q_3(t) \to 0$ as $t \to \infty$, and f is a 'nearly constant' function: $f(\ell) \simeq 1.205$ for all ℓ . Note that $\log_3 2$ is the Hausdorff dimension of the Cantor set

be useful to think in terms of the Hausdorff measures rather than the V_i , which are defined in terms of Lebesgue measures, especially since we are considering the magnitude of fractal sets as well.

We are then led to speculate about the general behaviour of the magnitude. It is possible to have varying degrees of optimism about this, so here are some possible conjectures. The first one we suspect is too strong to be true.

STRONG ASYMPTOTIC CONJECTURE Let C be the class of all compact subsets of \mathbb{R}^m , for any m. Then there are unique functions $P, q: C \to \mathbb{R}$ with the following properties:

- (1) |A| = P(A) + q(A);
- (2) P is a valuation, i.e. satisfies the inclusion-exclusion principle;
- (3) if $A \in \mathcal{C}$ is polyconvex then $P(A) = \sum_{i \geq 0} \frac{1}{i! \omega_i} V_i(A)$;
- (4) for all $A \in \mathcal{C}$ we have $q(tA) \to 0$ as $t \to \infty$;
- (5) if $K \in \mathcal{C}$ is convex then q(K) = 0.

Perhaps the conjecture that we feel most plausible is essentially part (5) of the Strong Conjecture, namely that for any convex subset K of Euclidean space, the magnitude is given by the valuation P, i.e.,

$$|K| = \sum_{i \ge 0} \frac{V_i(K)}{i! \, \omega_i}.$$

We call this the Convex Magnitude Conjecture. Further evidence for this kind of conjecture is presented in [13,15,19]. A weaker conjecture is the following.

WEAK ASYMPTOTIC CONJECTURE There is a set \mathcal{D} of compact subsets of Euclidean spaces \mathbb{R}^m , which includes the set of finite sets of points, convex sets, circles and Cantor sets, and there is a unique function $P: \mathcal{D} \to \mathbb{R}$ with the following properties:

- (1) if $A \in \mathcal{D}$ then $|tA| P(tA) \to 0$ as $t \to \infty$;
- (2) P is a valuation, i.e. satisfies the inclusion-exclusion principle;
- (3) if $A \in \mathcal{D}$ is polyconvex then $P(A) = \sum_{i \geq 0} \frac{1}{i! \omega_i} V_i(A)$.

A corollary of the Weak Asymptotic Conjecture would be the asymptotic inclusion-exclusion principle:

$$|t(A \cup B)| - |tA| - |tB| + |t(A \cap B)| \to 0$$
 as $t \to \infty$

whenever $A, B, A \cup B, A \cap B \in \mathcal{D}$.



What is in this paper. In the first section the notion of the magnitude of a finite metric space is motivated and defined. It is then proved that for any *n*-point metric space, provided that the points are sufficiently far apart then the magnitude is well-defined and satisfies

$$|X| = n + q_1(X)$$

where $q_1(tX) \to 0$ as $t \to \infty$. So asymptotically the magnitude is just the cardinality of the underlying set or, equivalently, the Euler characteristic of the underlying topological space. Note that for subsets of Euclidean space, the "sufficiently far apart" caveat is not necessary and the magnitude is always well-defined [13].

In the second section we show that for a length ℓ closed line segment, L_{ℓ} , we can calculate the magnitude using any sequence of finite subspaces of \mathbb{R} converging in the Hausdorff topology to the line segment and that the magnitude is

$$|L_{\ell}| = \ell/2 + 1.$$

This is *precisely* half of the length plus the Euler characteristic, so there is no 'asymptotic correction'; this is perhaps related to the fact that a line segment is convex.

In the third section we consider T_{ℓ} , the 'middle third' or 'ternary' Cantor set of length ℓ . We take an obvious sequence of approximations to this and calculate the magnitude of T_{ℓ} as the limit of the corresponding sequence of magnitudes. We find that this is of the form

$$|T_{\ell}| = p(\ell) + q_2(\ell)$$

where $q_2(\ell) \to 0$ as $\ell \to \infty$ and p satisfies the functional equation $2p(\ell) = p(3\ell)$, which is an inclusion-exclusion result as the ternary Cantor set has the self-similarity property $T_\ell \sqcup T_\ell = T_{3\ell}$. This means that the magnitude $|T_\ell|$ grows like $\log_3 2$, the Hausdorff dimension of the Cantor set.

In the fourth section the focus is on circles. Taking C_{ℓ} to be the circle of circumference ℓ embedded in Euclidean space in the canonical circular fashion, and thus equipped with the subspace metric, we calculate the magnitudes of a sequence of symmetric approximation to the circle, the limit of which is the magnitude of the circle. Using some non-trivial classical asymptotic analysis we show that

$$|C_{\ell}| = \ell/2 + q_3(\ell)$$

where $q_3(\ell) \to 0$ as $\ell \to \infty$. So again, asymptotically we get half of the length plus the Euler characteristic, where here the Euler characteristic is zero. We then go on to use the same techniques to look at *other* metrics on the circle. The other obvious metric on the circle is the 'arc-length' metric; this can be viewed as an intrinsic metric on the circle which does not depend on the embedding in Euclidean space, and the magnitude of the circle with this metric is shown to have the same asymptotics. In fact there is a family of metrics interpolating and extrapolating the above two metrics; each of these metrics is obtained by embedding the circle in a constant curvature surface and using the subspace metric. We show that with these metrics we again have the same asymptotic behaviour.

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1 The magnitude of finite and positive definite metric spaces

In the first part of this section a metric space is viewed as an enriched category so that the magnitude can be defined analogously to the Euler characteristic of a finite category, intro-



duced in [1] and [10]. This point of view acts purely as motivation, and an understanding of category theory is not necessary in order to read this paper. After the definition, some basic properties are given including a useful observation of David Speyer on the magnitude of homogeneous metric spaces. Next an overview of work of Mark Meckes on extending magnitude to a bigger class of metric spaces, namely positive definite spaces, is given. Finally in this section, it is shown that any finite metric space with sufficiently separated points has a well-defined magnitude, and that asymptotically the magnitude is simply the number of points.

1.1 Category theoretic motivation

This part can be skipped if desired, but it describes where the seemingly ad hoc definition of magnitude comes from. Those desirous of some category theoretic background can refer to the books of Borceux and Mac Lane [3,4,14].

Recall that a metric space consists of a set X together with a distance $d(x, x') \in [0, \infty)$ defined for each pair of elements $x, x' \in X$. These are required to satisfy the triangle inequality and the zero-self-distance axiom,

$$d(x, x') + d(x', x'') \ge d(x, x'')$$
 and $0 \ge d(x, x)$

(where the latter is usually written 0 = d(x, x) but our reasons for writing it as we have will become clear below), together with the symmetry and separation axioms,

$$d(x, x') = d(x', x)$$
 and $0 \ge d(x, x') \Leftrightarrow x = x'$.

Lawvere observed [9] that the first two of these conditions are analogous to composition of morphisms and the inclusion of the identity in a category:

$$\operatorname{Hom}(x, x') \times \operatorname{Hom}(x', x'') \to \operatorname{Hom}(x, x'')$$
 and $\{*\} \to \operatorname{Hom}(x, x)$.

Lawvere used this observation to interpret metric spaces as categories enriched over the following monoidal category. Take $[0, \infty]$ to be the monoidal category with the non-negative real numbers together with infinity as its objects, and with precisely one morphism from a to a' if $a \ge a'$ and no such morphisms otherwise; the monoidal product is addition + and the unit is 0. This category has categorical products and coproducts: the categorical product of a set of objects is the supremum of the set and the categorical coproduct is the infimum.

A category enriched over $[0, \infty]$ is then a set X with a function $d: X \times X \to [0, \infty]$ which satisfies the triangle inequality and zero-self-distance axiom. (The usual conditions these have to satisfy, namely associativity and the unit axiom, are vacuously satisfied in this case.) The notion of such an enriched category is then a generalization of the usual notion of metric space in that the distance between two points can be infinite, the distance does not have to be symmetric and the distance between two different points can be zero; Lawvere argued convincingly that many "metric spaces" in nature are of this more general form. In this paper the metric spaces will be symmetric and satisfy the separation axiom, but the notion of magnitude is defined in a similar way in the more general setting.

The definition of the Euler characteristic of a finite category can now be adapted to this enriched category situation. Recall briefly the definition of the Euler characteristic of a finite category [10]. If C is a finite category then a *weighting* on C is a choice of real number $w_i \in \mathbb{R}$ for each object $i \in ObC$ such that for every object i

$$\sum_{j \in \mathsf{Ob}\mathcal{C}} \# \left(\mathsf{Hom} \left(i, \, j \right) \right) w_j = 1$$



where # is just the cardinality or number-of-elements function on finite sets. If there exists a weighting on both \mathcal{C} and \mathcal{C}^{op} then the Euler characteristic is defined to be the sum of the weights: $\chi(\mathcal{C}) := \sum_i w_i$. The Euler characteristic is independent of the choice of weighting.

The key thing that needs adapting for the enriched case is the function #. We need a corresponding function $\widehat{\#}\colon [0,\infty]\to\mathbb{R}$ on the objects of the enriching category $[0,\infty]$. The function # satisfies $\#(X\times Y)=\#(X)\times \#(Y)$, so we will pick a function that satisfies $\widehat{\#}(a+b)=\widehat{\#}(a)\widehat{\#}(b)$. The obvious choice for such a function is $\widehat{\#}(a)=\alpha^a$ for some nonnegative number α and this is what we shall use, taking $\alpha=e^{-1}$. There is no obvious reason why other functions cannot be used; it is just that our choice gives interesting results. We can now give the definition.

1.2 Definition and basic properties for finite metric spaces

The magnitude of a finite metric space is defined in the following way.

Given a finite metric space X, a weighting is a choice of real number $w_x \in \mathbb{R}$ for each point $x \in X$ such that for each $x \in X$ we have

$$\sum_{x' \in X} e^{-d(x,x')} w_{x'} = 1.$$

If such a weighting exists then the *magnitude* |X| of the metric space X is defined to be the sum of the weights:

$$|X| := \sum_{x} w_x$$
.

There are several things to note from this definition. Firstly, there are finite metric spaces on which no weighting exists (see [11,13]) and the magnitude is then undefined. Secondly, the weightings are not necessarily positive. Thirdly, even if a weighting does exist then it might not be unique, however the magnitude is independent of the choice of weighting, as if w and \bar{w} are two weightings then

$$\sum_{x} w_{x} = \sum_{x} w_{x} \sum_{x'} e^{-d(x,x')} \bar{w}_{x'} = \sum_{x'} \sum_{x} w_{x} e^{-d(x',x)} \bar{w}_{x'} = \sum_{x'} \bar{w}_{x'},$$

where we have used the symmetry of the metric.

In many cases, such as for positive definite spaces (see the next subsection), the following gives a method for calculating the magnitude. Define the matrix Z of exponentiated distances, indexed by the points of X, as follows: $Z_{x,x'} := e^{-d(x,x')}$. If Z is invertible then X has a unique weighting w, and w_x is the sum of entries in the xth row in the inverse Z^{-1} :

$$w_x := \sum_{x'} (Z^{-1})_{x,x'}$$

and thus the sum of all entries of Z^{-1} gives the magnitude:

$$|X| = \sum_{x,x'} (Z^{-1})_{x,x'}.$$

This method can be applied to 'large' metric spaces in a sense made precise in Theorem 2 and, in fact, to 'most' metric spaces in a sense made precise in Proposition 2.2.6(i) of [13]. The following is a useful observation of David Speyer (stated in [18]).



Theorem 1 (The Speyer Formula) If a finite metric space X carries a transitive action by a group of isometries then there is a weighting in which the points all have the same weight and this is given on every point by

$$\frac{1}{\sum_{x' \in X} e^{-d(x,x')}}$$

for any $x \in X$. Thus the magnitude is defined and is given by

$$|X| = \frac{\# X}{\sum_{x' \in X} e^{-d(x,x')}}$$

for any $x \in X$.

Proof The observation is simply that as there is a transitive group action, for each $x \in X$ the set-with-multiplicity of distances $\{d(x,x') \mid x' \in X\}$ is the same, so $\sum_{x'} e^{-d(x,x')}$ is also the same for each x. Thus for every x

$$\sum_{x'} \left(\frac{1}{\sum_{x''} e^{-d(x'', x')}} \right) e^{-d(x, x')} = \frac{\sum_{x'} e^{-d(x, x')}}{\sum_{x''} e^{-d(x'', x')}} = 1$$

and the weighting condition is satisfied.

We can also define the *magnitude function*. For a finite metric space X and a positive number $t \in \mathbb{R}_{>0}$, let tX be the metric space obtained by scaling all of the distances by t, so that it has the same underlying set of points but the metric is given by $d_{tX}(x, x') := td_X(x, x')$. Now define the magnitude function, thought of as a function of t, to be |tX|. This is not necessarily defined for all t, but it *is* defined for all t sufficiently large, see Theorem 2 below. In this paper we will be interested in the asymptotics of this function for large t.

1.3 The magnitude of a compact metric space

In this subsection we will just mention some work of Mark Meckes [15] that followed on from an earlier version of this paper. It transpires that a particularly nice class of spaces is that of positive definite spaces: a metric space X is said to be *positive definite* if for every finite subspace W, the matrix $Z \in \mathbb{R}^{W \times W}$ given by $Z(x, x') := e^{-d(x, x')}$ is positive definite. Finite positive definite spaces always have well-defined, non-negative magnitude. Examples of positive definite spaces include subspaces of Euclidean space and round spheres; all of the infinite spaces considered in this paper are of this nice form.

For a compact positive definite space X the magnitude can be defined in the following two equivalent ways. Firstly, we can define it as

$$|X| := \sup\{|W| : W \subset X \text{ with } W \text{ finite}\};$$

for non-positive definite finite metric spaces this will not necessarily agree with the usual definition of magnitude. Secondly, by [15, Corollary 2.7], if $(X_k)_{k \in \mathbb{N}}$ is a sequence of finite subsets of X with $X_k \to X$ in the Hausdorff metric then $|X_k| \to |X|$. So this could equally be taken as a definition of magnitude. (The definition of the Hausdorff metric is recalled in Sect. 2.)

Let X be a compact positive definite metric space. A weight measure on X is a finite signed Borel measure μ on X such that for all $x \in X$,

$$\int\limits_{Y} e^{-d(x,x')} \, d\mu(x') = 1.$$



For any weight measure μ on X, we have $\mu(X) = |X|$ [15, Theorem 2.3]. This gives a third possible definition of magnitude, but one that only applies to spaces on which a weight measure exists.

Let *X* be a compact metric space such that tX is positive definite for all $t \gg 0$. The magnitude dimension [13, Definition 3.4.5] of *X* can be defined as the growth of the function $t \mapsto |tX|$:

$$\dim(X) = \inf \left\{ \nu \in \mathbb{R} : \frac{|tX|}{t^{\nu}} \text{ is bounded for } t \gg 0 \right\}.$$

For example, it can be shown that if A is a compact subset of \mathbb{R}^n with the Euclidean metric then $\dim(A) < n$, with equality if A has nonzero Lebesgue measure [13, Theorem 3.5.8].

1.4 Asymptotics for finite metric spaces

In this subsection it is shown that every finite metric space with sufficiently well separated points has a well-defined magnitude. Furthermore, asymptotically it is just the number of points; if X is a metric space with n points, then in the notation of the introduction, $|X| = n + q_1(X)$ where $q_1(tX) \to 0$ as $t \to \infty$.

Note that as subsets of Euclidean space are positive definite and finite positive definite spaces have a well-defined magnitude, any finite metric space without a well-defined magnitude is not isometrically embeddable in Euclidean space; however the converse does not hold, as there are plenty of non-Euclidean metric spaces with a well-defined magnitude.

Consider first the two-point space X_d where the two points are a distance d apart.

$$\bullet \stackrel{d}{\longleftrightarrow} \bullet$$

The magnitude $|X_d|$ is easy to calculate and there are many ways to do that. As the space is symmetric, Speyer's Formula, Theorem 1, can be applied to show that the magnitude is $2/(1 + e^{-d})$, which clearly tends to 2 as d tends to infinity. This can be rewritten as

$$|X_d| = 2 - \frac{2}{1 + e^d}$$

which is in the form given above: it is the magnitude of the set of points plus some term which is asymptotically zero as $d \to \infty$.

The general case of *n* points requires a little analysis. The first thing to do is to show that a finite metric space has a well-defined magnitude provided that the points are "sufficiently separated".

Theorem 2 If X is a finite metric space with n points such that the distance between each pair of distinct points is greater than $\ln(n-1)$ then X has a well-defined magnitude.

Note that whilst the bound only makes sense for $n \ge 2$, the condition is satisfied vacuously when n is 0 or 1.

It was pointed out by the referee that this result is a consequence of the Levy–Desplanques Theorem [6, Theorem 6.1.10], and that indeed the proof below can be generalized to give a proof of the Levy-Desplanques Theorem.

Proof Firstly, 0- and 1-point spaces have magnitude 0 and 1 respectively, so we may assume that $n \ge 2$. We wish to show that if $d(x, x') \ge \ln(n-1)$ for all $x \ne x'$ then the exponentiated distance matrix Z is invertible, so it suffices to show that if Z is an $n \times n$ real matrix with $Z_{ii} = 1$ for all i and $0 \le Z_{ij} < 1/(n-1)$ for all $i \ne j$ then Z is invertible. In fact we



show that *Z* is positive definite: $x^t Z x \ge 0$ for all $x \in \mathbb{R}^n$, with equality if and only if x = 0. Indeed,

$$x^{t}Zx = \sum_{i} x_{i}^{2} + \sum_{i \neq j} Z_{ij}x_{i}x_{j} \ge \sum_{i} x_{i}^{2} - \frac{1}{n-1} \sum_{i \neq j} |x_{i}||x_{j}|$$
$$= \frac{1}{2(n-1)} \sum_{i \neq j} (|x_{i}| - |x_{j}|)^{2} \ge 0.$$

If $x^t Z x = 0$ then all of the inequalities must be equalities and so $|x_1| = \cdots = |x_n| = \alpha$, say. However, as there is the *strict* inequality $Z_{ij} < 1/(n-1)$ it must be that $\alpha = 0$. Thus Z is positive definite and hence invertible, as required.

By the above theorem, if X is a finite metric space, then for $t \in \mathbb{R}$ sufficiently large, the scaled-up version of X has a well-defined magnitude |tX|, so it makes sense to talk of the asymptotic behaviour of |tX| even if the magnitude of X itself is not defined. The fundamental result can now be stated.

Theorem 3 If X is a finite metric space with n points then $|tX| \to n$ as $t \to \infty$.

Proof Let \mathcal{M}_n be the space of real $n \times n$ matrices. On the subspace $GL_n(\mathbb{R})$ of *invertible* matrices, define the real-valued function f by

$$f(Z)$$
:=(sum of the entries of Z^{-1})
= (sum of the entries of $adj(Z)$)/ $det(Z)$.

where adj(Z) is the adjugate matrix of Z. Then f(Z) is a rational function of the entries of Z, so f is continuous.

Writing Z_Y for the exponentiated distance matrix of a finite metric space Y, if Z_Y is invertible then $|Y| = f(Z_Y)$. We have $\lim_{t\to\infty} (Z_{tX}) = I$. As $t\mapsto Z_{tX}$ is continuous, $I\in \mathrm{GL}_n(\mathbb{R})$ and $\mathrm{GL}_n(\mathbb{R})$ is open in \mathcal{M}_n , then for $t\gg 0$ we have $Z_{tX}\in \mathrm{GL}_n(\mathbb{R})$ and

$$\lim_{t \to \infty} |tX| = \lim_{t \to \infty} f(Z_{tX}) = f\left(\lim_{t \to \infty} Z_{tX}\right) = f(I) = n.$$

In other words, in the language of the introduction, if X is finite metric space with n points which has a magnitude then we may define $q_1(X)$ by

$$|X| = n + q_1(X),$$

and then, for X any finite metric space, $q_1(tX) \to 0$ as $t \to \infty$.

2 The magnitude of a straight line segment

In this section we approximate L_ℓ , a closed straight line segment of length ℓ , by a sequence of finite metric spaces consisting of points lying in a line. We show that no matter which approximating sequence of this type is chosen, the sequence of magnitudes always converges to the same value, namely $|L_\ell|$. In fact $|L_\ell| = \ell/2 + 1$, which is *exactly* the conjectured valuation; for this space there is no need to make an asymptotic statement.

Note that in [20] the magnitude of a straight line segment is calculated using a weight measure. This takes advantage of the fact that L_{ℓ} has a weight measure; we do not expect L_{ℓ}^{m} ,



the *m*-dimensional cube of length ℓ , to have a weight measure for any m > 1. The method given here is more general and demonstrates some elementary aspects of the theory.

Start by considering finite metric spaces consisting of points arranged in a line; we call these *linear metric spaces*. For an (n-1)-tuple $\mathbf{d} = (d_1, \dots, d_{n-1})$ of strictly positive real numbers, define $X_{\mathbf{d}}$ to consist of n points with the distance between consecutive points being given by the d_i s, as in the following picture.

$$X_{\mathbf{d}}: \bullet \stackrel{d_1}{\longleftrightarrow} \bullet \stackrel{d_2}{\longleftrightarrow} \bullet \cdots \bullet \stackrel{d_{n-1}}{\longleftrightarrow} \bullet$$

This metric space has a weighting on it with the property that the weight of a point only depends on the distance to its nearest neighbours, giving rise to a simple expression for the magnitude.

Theorem 4 Suppose X_d is a linear metric space as above. Then there is a weighting on it such that weight of the ith point is

$$\frac{1}{2} \left(\tanh \left(d_{i-1}/2 \right) + \tanh \left(d_{i}/2 \right) \right),$$

where, for convenience, we write $d_0 = d_n = \infty$. Thus the magnitude of the linear space $X_{\mathbf{d}}$ is given by:

$$|X_{\mathbf{d}}| = 1 + \sum_{i=1}^{n-1} \tanh(d_i/2).$$

Proof The distance d_{ij} between the ith and j points, for i < j, is given by $\sum_{s=i}^{j-1} d_s$, so $e^{-d_{ij}} = \prod_s e^{-d_s}$. Thus writing $a_i := e^{-d_i}$ the exponentiated distance matrix of $X_{\mathbf{d}}$ is given by the following matrix, which we write out for the case n = 5, as the general pattern should be clear from this:

$$\begin{pmatrix} 1 & a_1 & a_1a_2 & a_1a_2a_3 & a_1a_2a_3a_4 \\ a_1 & 1 & a_2 & a_2a_3 & a_2a_3a_4 \\ a_1a_2 & a_2 & 1 & a_3 & a_3a_4 \\ a_1a_2a_3 & a_2a_3 & a_3 & 1 & a_4 \\ a_1a_2a_3a_4 & a_2a_3a_4 & a_3a_4 & a_4 & 1 \end{pmatrix}.$$

It is easy to verify that the inverse of such a matrix is

$$\begin{pmatrix} \frac{1}{1-a_1^2} & \frac{-a_1}{1-a_1^2} & 0 & 0 & 0\\ \frac{-a_1}{1-a_1^2} & \frac{1+a_1^2}{2(1-a_1^2)} + \frac{1+a_2^2}{2(1-a_2^2)} & \frac{-a_2}{1-a_2^2} & 0 & 0\\ 0 & \frac{-a_2}{1-a_2^2} & \frac{1+a_2^2}{2(1-a_2^2)} + \frac{1+a_3^2}{2(1-a_3^2)} & \frac{-a_3}{1-a_3^2} & 0\\ 0 & 0 & \frac{-a_3}{1-a_3^2} & \frac{1+a_3^2}{2(1-a_3^2)} + \frac{1+a_4^2}{2(1-a_4^2)} & \frac{-a_4}{1-a_4^2}\\ 0 & 0 & 0 & \frac{-a_3}{1-a_3^2} & \frac{1+a_3^2}{2(1-a_3^2)} + \frac{1+a_4^2}{1-a_4^2} & \frac{1-a_4^2}{1-a_4^2} \end{pmatrix}.$$

The weight of the *i*th point is the sum of the entries in the *i*th row, so for $i \neq 1$, n the weight is

$$\frac{1}{2} \left(\frac{1 - a_{i-1}}{1 + a_{i-1}} + \frac{1 - a_i}{1 + a_i} \right)$$



and as

$$\frac{1-a_i}{1+a_i} = \frac{a_i^{-1/2} - a_i^{1/2}}{a_i^{-1/2} + a_i^{1/2}} = \tanh(d_i/2)$$

it follows that the weight is just $\frac{1}{2} \left(\tanh(d_{i-1}/2) + \tanh(d_i/2) \right)$. At the endpoints the weight is simply

$$\frac{1}{1+a_m} = \frac{1}{2} \left(\frac{1-a_m}{1+a_m} + 1 \right) = \frac{1}{2} \left(\tanh(d_m/2) + 1 \right)$$

for m = 1, n - 1 respectively.

We wish to approximate a straight line interval by a sequence of such finite linear spaces, where 'approximate' means in the sense of the *Hausdorff metric*. Recall that the Hausdorff metric can be defined as follows (see, for example, [8]). If X' is a compact subset of a metric space X then for $\varepsilon \geq 0$, the ε -expansion $E(X', \varepsilon)$ of X' consists of all the points in X of distance at most ε from a point in X'. The Hausdorff distance between two compact subsets $X', X'' \subseteq X$ is defined to be the least $\varepsilon \geq 0$ such that each subset is contained within the ε -expansion of the other:

$$d(X', X'') := \inf\{\varepsilon \mid X' \subseteq E(X'', \varepsilon) \text{ and } X'' \subseteq E(X', \varepsilon)\}.$$

The following is straightforward from the definitions.

Lemma 5 Let $(X^k)_{k=1}^{\infty}$ be a sequence of finite subspaces of the length ℓ line segment L_{ℓ} so that for each k we have $X^k \cong X_{\mathbf{d}^k}$ for some tuple $\mathbf{d}^k = (d_1, \ldots, d_{n_k-1})$. Then $X^k \to L_{\ell}$ as $k \to \infty$ if and only if $\sum_i d_i^k \to \ell$ and $\max_i (d_i^k) \to 0$ as $k \to \infty$.

The reader unfamiliar with the Hausdorff metric can take this as the definition of convergence. We can now see that the limiting magnitude of such spaces is well-defined.

Proposition 6 If $(X^k)_{k=1}^{\infty}$ is a sequence of finite subsets of L_{ℓ} , a straight line segment of length ℓ , which converges to L_{ℓ} then the sequence of magnitudes converges:

$$X^k \to L_\ell \ as \ k \to \infty \ \Rightarrow \ |X^k| \to \ell/2 + 1 \ as \ k \to \infty.$$

Proof By the above lemma we can associate a sequence of tuples (\mathbf{d}^k) such that $\sum_i d_i^k \to \ell$ and $\max_i (d_i^k) \to 0$. Theorem 4 implies $|X^k| = 1 + \sum_i \tanh(d_i^k/2)$. Thus it suffices to prove that $\sum_i \tanh(d_i^k/2) \to \ell/2$ as $k \to \infty$. This requires a small amount of analysis.

We will first show that $|\tanh c - c| \le c^2$ for c > 0. By Lagrange's form for the remainder in the Taylor series, there exists $\xi \in (0, c)$ such that

$$\tanh(c) = \tanh(0) + \tanh'(0)c + \frac{1}{2}\tanh''(\xi)c^2.$$

As tanh(0) = 0, tanh'(0) = 1 and $tanh''(\xi) = 2(tanh^2(\xi) - 1) tanh(\xi)$ we get

$$|\tanh(c) - c| = \left| (\tanh^2(\xi) - 1) \tanh(\xi) \right| c^2 < c^2,$$

as required; the last inequality due to the fact that $|\tanh(\xi)| < 1$.



Now we can see

$$\begin{split} \left| \sum_{i} \tanh(d_{i}^{k}/2) - \ell/2 \right| &= \left| \sum_{i} \tanh(d_{i}^{k}/2) - \sum_{i} d_{i}^{k}/2 + \sum_{i} d_{i}^{k}/2 - \ell/2 \right| \\ &\leq \left| \sum_{i} \left(\tanh(d_{i}^{k}/2) - d_{i}^{k}/2 \right) \right| + \frac{1}{2} \left| \sum_{i} d_{i}^{k} - \ell \right| \\ &\leq \sum_{i} (d_{i}^{k}/2)^{2} + \frac{1}{2} \left| \sum_{i} d_{i}^{k} - \ell \right| \\ &\leq (\max_{i} d_{i}^{k}) \sum_{i} d_{i}^{k}/4 + \frac{1}{2} \left| \sum_{i} d_{i}^{k} - \ell \right| \\ &\to 0 \cdot \ell/4 + \frac{1}{2} \cdot 0 = 0 \quad \text{as } k \to \infty, \end{split}$$

so
$$\sum_{i} \tanh(d_i^k/2) \to \ell/2$$
 and $|X^k| \to \ell/2 + 1$ as required.

This proposition together with [15, Corollary 2.7] gives us the magnitude of any straight line segment.

Theorem 7 The magnitude of the straight line segment of length ℓ has the following form.

$$|L_{\ell}| := \ell/2 + 1.$$

This is precisely the conjectured form, even non-asymptotically, thus supporting part (5) of the Strong Asymptotic Conjecture in the introduction.

3 The magnitude of a ternary cantor set

Here we consider T_{ℓ} , the ternary Cantor set of length ℓ , where $\ell > 0$. We will calculate the magnitude $|T_{\ell}|$ of this Cantor set as a limit and show that the result is consistent with the belief that asymptotically the magnitude satisfies the inclusion-exclusion principle. A different approach to this calculation, using weight measures, is given in [20].

The ternary Cantor set T_{ℓ} is constructed by starting with a closed straight line segment of length ℓ , removing the open middle third, then removing the middle thirds of the two remaining components and continuing like this *ad infinitum*. This process means that $T_{3\ell}$, the Cantor set of length 3ℓ , can be decomposed into two copies of T_{ℓ} , so

$$T_{3\ell} = T_{\ell} \sqcup T_{\ell}$$
.

Thus if P is any valuation, i.e. satisfies the inclusion-exclusion principle, which is defined on some collection of sets including these Cantor sets, then

$$P(T_{3\ell}) = 2P(T_{\ell}).$$

Writing $p(\ell) := P(T_{\ell})$ we get the functional equation

$$p(3\ell) = 2p(\ell)$$
.

We can characterize the functions satisfying this functional equation in the following way.

Lemma 8 Suppose $p: \mathbb{R}_{>0} \to \mathbb{R}$ is a function defined on the positive real numbers; then p satisfies the functional equation

$$p(3\ell) = 2p(\ell)$$
 for all $\ell > 0$

if and only if p is of the form

$$p(\ell) = f(\ell)\ell^{\log_3(2)}$$

where $f: \mathbb{R}_{>0} \to \mathbb{R}$ is some multiplicatively periodic function in the sense that $f(3\ell) = f(\ell)$ for all $\ell > 0$.



Whilst such multiplicatively periodic functions are less frequently encountered than their additive counterparts, they are no less common: to obtain such a multiplicatively periodic function f, pick a period-one ordinarily periodic function $g: \mathbb{R} \to \mathbb{R}$ and define $f(\ell) := g(\log_3(\ell))$.

Proof Suppose first that p is a solution of the functional equation, then define $f(\ell) := p(\ell)\ell^{-\log_3(2)}$ for $\ell > 0$. As $3^{\log_3(2)} = 2$ we have

$$f(3\ell) = p(3\ell)(3\ell)^{-\log_3(2)} = 2p(\ell)2^{-1}\ell^{-\log_3(2)} = f(\ell),$$

and so p has the required form.

Conversely, if f is a function satisfying $f(3\ell) = f(\ell)$, then defining $p(\ell) := f(\ell)\ell^{\log_3(2)}$ is easily seen to give a function satisfying the functional equation:

$$p(3\ell) = f(3\ell)(3\ell)^{\log_3(2)} = f(\ell)2\ell^{\log_3(2)} = 2p(\ell),$$

as required.

The appearance of $\log_3(2)$ here is not outrageous as it is the Hausdorff dimension of the Cantor set T_ℓ for every ℓ . We will show that the magnitude $|T_\ell|$ is of the form $p(\ell) + q_2(\ell)$ where p satisfies the above functional equation and $q_2(\ell) \to 0$ as $\ell \to \infty$. Hence the magnitude dimension (see Sect. 1.3) of the Cantor set is equal to its Hausdorff dimension.

We will use a more constructive definition of the Cantor set T_ℓ . We start with the zeroth approximation T_ℓ^0 which consists of two points on the real line a distance ℓ apart. Let ψ_1 and ψ_2 be the two scalings of the real line by a factor of 1/3 with the two points of T_ℓ^0 as their respective fixed points. Define T_ℓ^k , the kth approximation to the Cantor set, inductively by $T_\ell^k := \psi_1(T_\ell^{k-1}) \cup \psi_2(T_\ell^{k-1})$ for $k \ge 1$. Then by the work of Hutchinson (see [7]) the length ℓ Cantor set is the limit of these sets, $T_\ell = \bigcup_k T_\ell^k$, and it is the unique non-empty compact subset satisfying $T_\ell = \psi_1(T_\ell) \cup \psi_2(T_\ell)$. Using the formula for the magnitude of a set of points in a line given above, it is easy to calculate the magnitudes of these finite approximations to the Cantor set.

Theorem 9 The magnitude of the kth approximation to the Cantor set of length ℓ is

$$\left|T_{\ell}^{k}\right| = 1 + 2^{k} \tanh\left(\frac{\ell}{2 \cdot 3^{k}}\right) + \frac{1}{2} \sum_{i=1}^{k} 2^{i} \tanh\left(\frac{\ell}{2 \cdot 3^{i}}\right).$$

Proof We use the formula of Theorem 4 for the magnitude of a linear metric space in terms of the distances between neighbouring points. Since the (k+1)th approximation, T_ℓ^{k+1} , is two copies of $T_{\ell/3}^k$ a distance $\ell/3$ apart, we know that for a pair of neighbouring points in T_ℓ^{k+1} , either both are in the same copy of $T_{\ell/3}^k$ or else they are in different copies and are a distance $\ell/3$ apart. So by Theorem 4, for $k \ge 0$,

$$\left|T_{\ell}^{k+1}\right| = 1 + 2\left(\left|T_{\ell/3}^{k}\right| - 1\right) + \tanh\left(\frac{\ell}{2\cdot 3}\right).$$

As we know that $\left|T_{\ell}^{0}\right|=1+\tanh(\ell/2)$, the result follows from a straightforward induction argument. \Box

We compute the magnitude of the Cantor set as the limit of these magnitudes.



Theorem 10 The magnitude of T_{ℓ} , the length ℓ ternary Cantor set, is given by

$$|T_{\ell}| = 1 + \frac{1}{2} \sum_{i=1}^{\infty} 2^{i} \tanh\left(\frac{\ell}{2 \cdot 3^{i}}\right).$$

Proof Since $|\tanh c| \le c$ for all $c \ge 0$, the sum $\sum_{i=1}^{\infty} 2^i \tanh\left(\ell/2 \cdot 3^i\right)$ converges and $2^k \tanh(\ell/2 \cdot 3^k) \to 0$ as $k \to \infty$. Hence $\lim_{k \to \infty} |T_{\ell}^k|$ exists. We know by [15, Corollary 2.7] that the magnitude of the length ℓ Cantor set is this limit, and so the result follows.

This magnitude can be decomposed as promised above. First define

$$p(\ell) := \frac{1}{2} \sum_{i=-\infty}^{\infty} 2^i \tanh\left(\frac{\ell}{2 \cdot 3^i}\right); \quad q_2(\ell) := 1 - \frac{1}{2} \sum_{i=0}^{\infty} \frac{1}{2^i} \tanh\left(\frac{3^i \ell}{2}\right).$$

Note the doubly infinite summation in the definition of p: the part of the sum indexed by negative i and the sum in the definition of q_2 both converge because tanh is bounded.

The promised result is almost immediate.

Theorem 11 The magnitude of the length ℓ Cantor set decomposes uniquely as a piece satisfying the functional equation and a piece which is asymptotically zero:

- (1) $|T_{\ell}| = p(\ell) + q_2(\ell);$
- (2) $p(3\ell) = 2p(\ell);$
- (3) $q_2(\ell) \to 0 \text{ as } \ell \to \infty$;
- (4) the functions p and q_2 are uniquely determined by the above three properties.

Proof (1) This follows by definition.

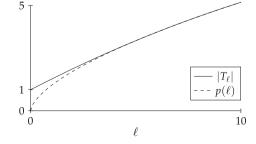
- (2) This follows immediately by substitution.
- (3) This follows from the fact that $\tanh d \to 1$ as $d \to \infty$.
- (4) Suppose \tilde{p} and \tilde{q}_2 form another decomposition of $|T_\ell|$ with these properties. Then defining $s := p \tilde{p} = \tilde{q}_2 q_2$ gives a function which satisfies both $s(3\ell) = 2s(\ell)$ and $s(\ell) \to 0$ as $\ell \to \infty$, from which it follows that s is identically zero, thus $\tilde{p} = p$ and $\tilde{q}_2 = q_2$.

The functions $|T_{\ell}|$ and $p(\ell)$ are plotted in Fig. 1.

We can now get $p(\ell)$ into the form $f(\ell)\ell^{\log_3 2}$ where f is multiplicatively periodic as in Lemma 8. Define

$$f(\ell) := \ell^{-\log_3(2)} \frac{1}{2} \sum_{i=-\infty}^{\infty} 2^i \tanh\left(\frac{\ell}{2 \cdot 3^i}\right)$$

Fig. 1 The magnitude $|T_\ell|$ of the length ℓ Cantor set is asymptotically the same as the valuation-like function $p(\ell)$, and the latter satisfies $1.205 \ \ell^{\log_3(2)} < p(\ell) < 1.206 \ \ell^{\log_3(2)}$ (according to maple)





so that $p(\ell) = f(\ell)\ell^{\log_3(2)}$. The multiplicative periodicity of f is clear when it is written in terms of the base-three logarithm of ℓ :

$$f(3^x) = \frac{1}{2} \sum_{i=-\infty}^{\infty} 2^{i-x} \tanh\left(3^{x-i}/2\right).$$

It can be seen numerically that f is not far from being a constant function; indeed, using maple, we can calculate the Fourier expansion which we see has rapidly decaying coefficients:

$$f(3^x) \simeq 1.2054 + 2.48 \times 10^{-4} \sin(2\pi x + \theta_1) + 3.36 \times 10^{-8} \sin(4\pi x + \theta_2) + \cdots$$

where θ_1 and θ_2 are some constants. Intriguingly, similar near-constant functions arising from the same functional equation were studied in [2]. From these numerics we get the following bounds for the valuation-like function p:

$$1.205 \ \ell^{\log_3(2)} < p(\ell) < 1.206 \ \ell^{\log_3(2)}$$

In conclusion, Theorem 11 tells us that although the magnitude of the Cantor set $|T_{\ell}|$ does *not* satisfy the functional equation, so in general $|T_{3\ell}| \neq 2|T_{\ell}|$, it *is* asymptotically equal to a function $p(\ell)$ that does satisfy the functional equation. Moreover, $|T_{\ell}|$ has growth $\log_3 2$, which is the Hausdorff dimension of the Cantor set.

4 The magnitude of a circle

In this section we consider the magnitude of circles. There are actually several metrics that can be put on a circle. In this paper we are primarily interested in subsets of Euclidean space with the subspace metric, so we will first consider the subspace metric on circles which are embedded in the natural 'round' way in \mathbb{R}^2 . We will see that the magnitude of C_ℓ , a circle of circumference ℓ , is of the form $|C_\ell| = \ell/2 + q_3(\ell)$, where $q_3(\ell) \to 0$ as $\ell \to \infty$. This requires some non-trivial asymptotic analysis. After doing this we will move away from subspaces of Euclidean space and show that the same result holds for other natural metrics on the circle. Despite considering non-Euclidean supspaces, we can apply the same techniques since, by Theorem 3.6(6) of [15], these metrics are positive definite. As in [20], the magnitude calculations performed here can also be done with weight measures.

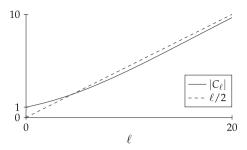
4.1 The subspace metric on the circle

In this section we calculate the magnitude of a circle with the subspace-of-Euclidean-space metric, and then use Laplace's method from asymptotic analysis to show that as the length increases the magnitude becomes close to half the length.

First consider a circle C_ℓ of length, or circumference, ℓ as a subset of \mathbb{R}^2 with the induced metric. This means that the distance between points p_1 and p_2 on the circle which subtend an angle θ at the origin is given by

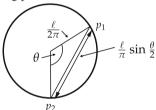


Fig. 2 The magnitude $|C_\ell|$ of the length ℓ circle with the subspace metric is asymptotically the same as half of the length



$$d(p_1, p_2) = \frac{\ell}{\pi} \sin \frac{\theta}{2},$$

as can be seen from the following picture.



Now we will approximate the circle by a finite set of points. We define K_ℓ^n to be a set of n points equally spaced around the circle, equipped with the subspace metric. This finite metric space is homogeneous, as it carries a transitive group action of the cyclic group of order n, so we can apply Speyer's Formula (Theorem 1) to see that the magnitude of this finite approximation to the circle is given by

$$\left|K_{\ell}^{n}\right| = \frac{n}{\sum_{j=1}^{n} \exp\left(\frac{-\ell}{\pi} \sin\frac{\pi j}{n}\right)} = \frac{1}{\sum_{j=1}^{n} \exp\left(\frac{-\ell}{\pi} \sin\frac{\pi j}{n}\right) \frac{1}{n}}.$$

We can take the limit as the number of points tends to infinity and see that the denominator just consists of Riemann sums, so tends to an integral:

$$|K_{\ell}^n| \to \frac{1}{\int_0^1 \exp\left(\frac{-\ell}{\pi}\sin(\pi s)\right) ds} \text{ as } n \to \infty.$$

As these subsets converge to the circle, we know, by [15, Corollary 2.7], that the magnitude of the circle is just this limit; we have therefore proved the following.

Theorem 12 For C_{ℓ} the length ℓ circle equipped with the Euclidean subspace metric, the magnitude is given by

$$|C_{\ell}| = \left(\int_{0}^{1} \exp\left(-\frac{\ell}{\pi}\sin(\pi s)\right) ds\right)^{-1}.$$

This magnitude is plotted for some values in Fig. 2, and it is seen that the magnitude appears to be approaching half of the length; some classical asymptotic analysis shows that this is indeed the case.



Theorem 13 For C_{ℓ} the length ℓ circle equipped with the Euclidean subspace metric, the magnitude satisfies

$$|C_{\ell}| = \ell/2 + q_3(\ell),$$

where $q_3(\ell) \to 0$ as $\ell \to \infty$.

Proof We will in fact prove that $|C_{\ell}| = \ell/2 + O(\ell^{-1})$ as $\ell \to \infty$; this follows by taking the reciprocal of both sides of the assertion

$$\int_{0}^{1} e^{-\frac{\ell}{\pi}\sin(\pi s)} ds = \frac{2}{\ell} + O(\ell^{-3}) \text{ as } \ell \to \infty,$$

which is what we will now derive using the classical asymptotic analysis technique known as Laplace's method (see, for example, [16]).

Define $D_0(s) := \frac{1}{\pi} \sin(\pi s)$ for $s \in [0, 1]$; this is plotted in Fig. 4. Thus D_0 is a function that takes minimum value zero precisely at its endpoints, $D_0(0) = 0 = D_0(1)$, and is infinitely differentiable with $D_0'(0)$, $D_0'(1) \neq 0$. We wish to consider the asymptotic behaviour, as $\ell \to \infty$, of the integral

$$F(\ell) := \int_0^1 e^{-\ell D_0(s)} \mathrm{d}s.$$

We break the integral into two pieces:

$$F(\ell) = \int_{0}^{1/2} e^{-\ell D_0(s)} ds + \int_{1/2}^{1} e^{-\ell D_0(s)} ds$$

=: $F_{\text{left}}(\ell) + F_{\text{right}}(\ell)$.

Concentrating on the asymptotics of F_{left} , as D_0 is one-to-one on [0, 1/2], it has an inverse there, so we can use a change of variable and Watson's Lemma (see [16, Chapter 3.3] for the details) to obtain

$$F_{\rm left}(\ell) = \frac{1}{D_0'(0)\ell} - \frac{D_0''(0)}{D_0'(0)^3\ell^2} + O(\ell^{-3}) \text{ as } \ell \to \infty.$$

As $D'_0(0) = 1$ and $D''_0(0) = 0$, we obtain

$$F_{\text{left}}(\ell) = \frac{1}{\ell} + O(\ell^{-3}).$$

By symmetry F_{right} gives the same contribution; adding these together gives the asymptotics of the integral:

$$\int_{0}^{1} e^{-\frac{\ell}{\pi}\sin(\pi s)} dx = \frac{2}{\ell} + O(\ell^{-3}) \text{ as } \ell \to \infty$$

which suffices to prove the theorem.

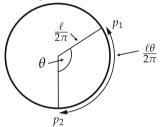


The analysis in the proof can be easily extended to show that the magnitude function asymptotically looks like $\ell/2 - \pi^2/2\ell + \cdots$.

Similarly, one can try to do a Taylor expansion of the magnitude function around $\ell=0$. For instance, the derivative of $|C_{\ell}|$ at $\ell=0$ is $\frac{\ell}{\pi}\int_0^1\sin(\pi s)\mathrm{d}s=2\ell/\pi^2$. It is not clear if the $\pi^2/2$ there is related to the one in the preceding paragraph.

4.2 The intrinsic metric

Another obvious, perhaps more obvious, choice of metric on a length ℓ circle is the 'arclength' metric so that the distance between two points p_1 and p_2 on the circle which subtend an angle $\theta \in [0, \pi]$ at the centre is $\ell\theta/2\pi$.



We will denote the length ℓ circle with this metric by \bar{C}_{ℓ} . On the one hand this metric can be viewed as 'the' intrinsic metric on the circle which exists independent of any embedding and corresponds to the natural Riemannian structure on the circle. On the other hand this can be seen as another way to put a 'subspace metric' on a subset of a metric space: on a metric space one can define the arc-length of a path in the space and define a metric on any subspace by taking the distance between two points to be the length of the shortest path between them which lies in the subspace.

To calculate the magnitude of this metric space we again approximate the circle by a sequence of subsets of evenly spaced points. Note that although this is not a metric inherited from Euclidean space, the metric is still positive definite [15, Theorem 3.6(6)] and so the approximation approach still works. So, let \bar{K}_{ℓ}^{n} be the metric space with n points labelled $1, \ldots, n$ such that the distance between the ith and jth points is given by

$$d_{i,j} = \frac{\ell}{n} \min(|i-j|, n-|i-j|).$$

Again this is a homogeneous space as the cyclic group of order n acts transitively on it, so we can apply Speyer's Formula (Theorem 1) to obtain the magnitude of this finite approximation to the length ℓ circle with the intrinsic metric:

$$\left| \bar{K}_{\ell}^{n} \right| = \frac{n}{\sum_{j=1}^{n} e^{-d_{1,j}}}.$$

Now we can find the magnitude of the actual circle by letting the number of points tend to infinity whilst keeping the length fixed.

Theorem 14 For \bar{C}_{ℓ} the length ℓ circle with its intrinsic metric, the magnitude is given by

$$\left|\bar{C}_{\ell}\right| = \frac{\ell/2}{1 - e^{-\ell/2}}.$$

Proof We know by [15, Corollary 2.7] that $|\bar{K}_{\ell}^n| \to |\bar{C}_{\ell}|$, so we just need to prove $|\bar{K}_{\ell}^n| \to \frac{\ell/2}{1-e^{-\ell/2}}$. There are at least two ways to prove this. One can express the denominator in the



formula for $|\bar{K}^n_\ell|$ as a geometric progression and then take the limit, or one can express the limit of the denominator as an integral. The second method is the analogue of the method in the last subsection and so that is the method given here.

As the limit of the sequence of magnitudes $|\bar{K}_{\ell}^1|$, $|\bar{K}_{\ell}^2|$, $|\bar{K}_{\ell}^3|$, ... exists it is equal to the limit of the 'odd' subsequence of magnitudes $|\bar{K}_{\ell}^1|$, $|\bar{K}_{\ell}^3|$, $|\bar{K}_{\ell}^5|$, ... We have

$$\left| \bar{K}_{\ell}^{2m+1} \right| = \frac{2m+1}{2\sum_{k=1}^{m} e^{-\ell k/(2m+1)} + 1} = \frac{\ell}{2\sum_{k=1}^{m} e^{-\ell k/(2m+1)} \frac{\ell}{2m+1} + \frac{\ell}{2m+1}}.$$

Now observe that

$$\sum_{k=1}^{m} e^{-\ell k/(2m+1)} \frac{\ell}{2m+1} \longrightarrow \int_{0}^{\ell/2} e^{-x} dx = 1 - e^{-\ell/2} \text{ as } m \to \infty.$$

Thus

$$\left| \bar{K}_{\ell}^{2m+1} \right| \to \frac{\ell/2}{1 - e^{-\ell/2}} \quad \text{as } m \to \infty,$$

and the proof is completed.

In other words we have

$$|\bar{C}_{\ell}| = \ell/2 + \frac{\ell/2}{e^{\ell/2} - 1},$$

and we immediately see that this has the suggested behaviour of being half of the length plus an asymptotically zero piece. This case is even better behaved than the subspace metric case considered above, in that this required far less analysis to derive and the difference between the magnitude and half of the length is exponentially small in the length—the magnitude and half of the length are asymptotically the same to all orders.

4.3 Other round metrics on the circle

As with the last subsection, this subsection is slightly away from the main theme of the paper in that it deals with metric spaces which are not subspaces of Euclidean space, but are subspaces of manifolds with the subspace metric; however, these are still positive definite spaces and this should be of some interest anyway. In the above two subsections we saw that there are two obvious metrics on the circle, the Euclidean subspace metric and the intrinsic metric. Here we will see that they are part of a family of metrics on the circle; each member of the family is a subspace metric got by considering the circle of length ℓ as the locus of points equidistant from some fixed point on a homogeneous surface of specified curvature. Think, for instance, of a circle of fixed latitude on the Earth as points equidistant from the North Pole. In particular the intrinsic metric on a circle comes from embedding it as an equator in a sphere. We will see that the asymptotic relationship to half of the length of the circle holds in an appropriate sense.

As shown in Fig. 3 a circle of length ℓ can be embedded 'roundly', by which we will mean as a locus of points equidistant from some fixed point, on any sphere of radius at least $\ell/2\pi$. We will consider the plane to be a sphere of infinite radius and as being the limit of arbitrarily large spheres. The curvature of a sphere of radius R is $1/R^2$, so a length ℓ circle

³ This terminology is just to distinguish a 'round' circle from a topological circle.



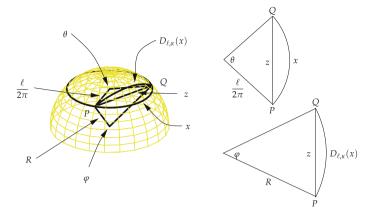


Fig. 3 Calculating the spherical distance between points P and Q on a length ℓ circle embedded on a radius R sphere: x is the circular distance between the points; z is the Euclidean distance between the points; and $D_{\ell,K}(x)$ is the spherical distance between the points

can be roundly embedded in any sphere of curvature from $(2\pi/\ell)^2$ down to 0; and it will be convenient to define the *relative curvature* of a radius R sphere to be

$$\kappa := \frac{1/R^2}{1/(\ell/2\pi)^2} = \frac{\ell^2}{4\pi^2 R^2}.$$

Thus a circle of length ℓ can be embedded in spheres of relative curvature from 1 down to 0. We can equip the length ℓ circle with the subspace metric, call this the κ -metric and denote the resulting metric space by $C_{\ell,\kappa}$. Thus $C_{\ell,1}$ is just \bar{C}_{ℓ} , the circle with the intrinsic metric, and $C_{\ell,0}$ is just C_{ℓ} , the circle with the Euclidean metric.

We will now give a formula for the κ -metric. First it is necessary to decide how to parametrize the length ℓ circle; we will parametrize it by arc-length, whereas in the previous two subsections we have parametrized by angle subtended at the centre. Define $D_{\ell,\kappa}:[0,\ell]\to\mathbb{R}$ by taking $D_{\ell,\kappa}(x)$ to be the subspace distance between two points a distance x away on the length ℓ circle which is roundly embedded in the sphere (or plane) of relative curvature κ , as in Fig. 3. In the case $\kappa=0$, by the formula in Sect. 4.1, we have

$$D_{\ell,0}(x) = \frac{\ell}{\pi} \sin \frac{\pi x}{\ell},$$

whereas for $0 < \kappa \le 1$ we have the following.

Theorem 15 The subspace metric on the length ℓ circle embedded roundly in the 2-sphere of radius R and of relative curvature $\kappa = \ell^2/4\pi^2 R^2 > 0$ is given by

$$D_{\ell,\kappa}(x) = \frac{\ell}{\sqrt{\kappa}\pi} \sin^{-1}\left(\sqrt{\kappa}\sin\frac{\pi x}{\ell}\right),\,$$

where \sin^{-1} is interpreted as a function $[0, 1] \rightarrow [0, \pi/2]$.

Proof The essential ideas of the proof are contained in Fig. 3. Suppose that we have a round, length ℓ circle on a sphere of radius R. We will make things easier for ourselves by embedding the 2-sphere in the standard way into Euclidean 3-space as in Fig. 3. Suppose further that P and Q are two points on the circle such that the arc of the circle between them has length x.



We wish to find $D_{\ell,\kappa}(x)$ which is the spherical distance between them, that is to say the the length of the arc of a great circle from P to Q, where a great circle is a circle on the sphere whose centre is at the origin.

The points P and Q are the endpoints of segments of two circles, the circle of length ℓ and the great circle passing through P and Q. These two segments are pictured in Fig. 3 with θ written for the angle subtended by P and Q at the centre of the length ℓ circle, φ written for the angle subtended by P and Q at the centre of the sphere, and Z written for the Euclidean distance between P and Q. From the two segments it is seen that

$$z = \frac{\ell}{\pi} \sin \frac{\theta}{2}$$
 and $z = 2R \sin \frac{\varphi}{2}$,

but the angles θ and φ are seen to be given by

$$\theta = \frac{2\pi x}{\ell}$$
 and $\varphi = \frac{D_{\ell,\kappa}(x)}{R}$,

so equating the two expressions for z we get

$$\frac{\ell}{\pi}\sin\frac{\pi x}{\ell} = 2R\sin\frac{D_{\ell,\kappa}(x)}{2R}$$

whence

$$D_{\ell,\kappa}(x) = 2R\sin^{-1}\left(\frac{\ell}{2\pi R}\sin\frac{\pi x}{\ell}\right) = \frac{\ell}{\sqrt{\kappa}\pi}\sin^{-1}\left(\sqrt{\kappa}\sin\frac{\pi x}{\ell}\right),\,$$

which is what was required.

Note that when $\kappa=1$ this recovers the intrinsic metric and that $D_{\ell,\kappa}\to D_{\ell,0}$ as $\kappa\to 0$; this corresponds to the idea that locally a large sphere looks metrically like a patch of the Euclidean plane. Thus this gives a family of metrics interpolating between the Euclidean and intrinsic ones.

This family of metrics can actually be extended by taking the relative curvature κ to be in $(-\infty, 1]$. The formula for $D_{\ell,\kappa}$ is valid for negative κ and corresponds to a metric on the length ℓ circle induced by embedding it roundly in a suitably curved hyperbolic space; this can be proved either by geometric means akin to those in Theorem 15, or by the standard algebraic trick of considering a hyperbolic plane as a sphere of imaginary radius. If the reader is unhappy with the imaginary quantities in the expression for $D_{\ell,\kappa}$ when κ is negative then they could perhaps be reassured by rewriting the expression for $D_{\ell,\kappa}$ in terms of hyperbolic sines:

$$D_{\ell,\kappa}(x) = \frac{\ell}{\sqrt{-\kappa}\pi} \sinh^{-1}\left(\sqrt{-\kappa}\sin\frac{\pi x}{\ell}\right) \quad \text{for } x \in [0,\ell].$$

Observe that $D_{\ell,\kappa}$ has the scaling property

$$D_{\ell \kappa}(\ell s) = \ell D_{1 \kappa}(s) \text{ for } s \in [0, 1],$$

which is why we chose to parametrize by the relative curvature κ . This means that we can concentrate on $D_{1,\kappa}$ which we will write as D_{κ} .

The graph of D_{κ} for some values of κ is plotted in Fig. 4. Note that if κ is taken to be sufficiently negative then the diameter of the length 1 circle can be made arbitrarily small. Note as well that for all values of κ

$$D'_{\kappa}(0) = +1$$
, $D'_{\kappa}(1) = -1$, and $D''_{\kappa}(0) = 0 = D''_{\kappa}(1)$.



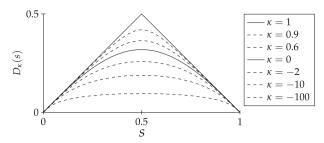


Fig. 4 The graph of D_{κ} for some values of κ from 1 at the *top* to -100 at the *bottom*. The *two solid lines* are $\kappa = 1$ and $\kappa = 0$ corresponding to the intrinsic and Euclidean metrics on the circle respectively

These are the key facts which will be used below and they are saying that all of these metrics are infinitesimally the same to second order. This is related to the fact that they all correspond to the standard *Riemannian* metric on the circle.

Just as in Sect. 4.1 we can approximate $C_{\ell,\kappa}$, a circle of length l with the κ -metric for $\kappa \in (-\infty, 1]$, by using a set of n points equally spaced on the circle and equipped with the subspace metric. Exactly the same argument leads us to the magnitude of the length l circle with the κ metric.

Theorem 16 For any $\kappa \in (-\infty, 1]$ the magnitude of the the κ -metric length ℓ circle satisfies

$$|C_{\ell,\kappa}| := \left(\int\limits_0^1 e^{-\ell D_{\kappa}(s)} ds\right)^{-1}.$$

and asymptotically it behaves as follows

$$|C_{\ell,\kappa}| = \ell/2 + O(\ell^{-1})$$
 as $\ell \to \infty$.

Proof The calculation of the magnitude is the same as in Sect. 4.1. Similarly, the asymptotic calculation requires the same argument as used in the proof of Theorem 13. The relevant facts about D_{κ} used are the following.

$$\begin{split} D_{\kappa}(0) &= 0 = D_{\kappa}(1); \quad D_{\kappa}(x) > 0 \text{ for } x \in (0, 1); \\ D_{\kappa}'(0) &= 1 = -D_{\kappa}'(1); \quad D_{\kappa}''(0) = 0 = D_{\kappa}''(1). \end{split}$$

Just as with Theorem 13 the analysis in the proof can be easily extended to show that the magnitude function asymptotically looks like $\ell/2 + \pi^2(\kappa - 1)/2\ell + \cdots$.

To conclude, we note that whilst the magnitudes associated to these metrics on a circle are all different, they all have the same asymptotic behaviour. These ideas have subsequently been developed further in [20] where the magnitude of higher dimensional spheres with their intrinsic metric are calculated and the asymptotics of the magnitudes of homogenous Riemannian manifolds are given. In that paper the calculations are done using weight measures. Here we have confined ourselves to the more elementary, and perhaps more widely applicable, method of finite approximation.

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