

LEOPACK



svenspec

Solution **V**ector **E**Nergy **S**PECtrum

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1 svenspec

Solution Vector ENergy SPECTrum

Reads in a solution vector in standard form (i.e. with a `.ints`, `.vecs` and `.xarr` file) and produces somewhat miscellaneous data on the spherical integrals of kinetic and magnetic energies of the solution.

The stand-alone source code version of the program is compiled by typing

```
make svenspec
```

within this directory. Once the executable is created, begin execution by typing

```
svenspec < inputfile
```

The inputs file must have the following format.

```
../../../../EXAMPLES/FUNDAMENTALS/case1.ints
../../../../EXAMPLES/FUNDAMENTALS/case1.xarr
../../../../EXAMPLES/FUNDAMENTALS/case1.vecs
example_aOUTPUT
```

Note that, unlike the input files to many of the other programs, an asterisk (*) can **NOT** be used to “comment out” any part of the input file. Neither can comments be inserted after the variables on each line. The input file must therefore be exactly as the form given.

The arguments are as follows

- **LINE ONE:** name of already existing indices file describing a solution vector. (i.e. a `.ints` file).
- **LINE TWO:** name of already existing radial spacings file describing solution. (i.e. a `.xarr` file).
- **LINE THREE:** name of already existing vector file describing solution. Must contain the same number of radial functions as indicated in the `.ints` file and the same number of radial grid nodes as indicated in the `.xarr` file.
- **LINE FOUR:** Name of output file.

1.1 Subprograms required for svenSpec

SUBS subroutines

vecop.f matop.f hmfrd.f xarrrd.f svfrd.f svfdcf.f
shkeer.f fopen.f fclose.f ldgnmf.f gfdcf.f asvdr.f
fnamer.f

SUBS double precision function

emmult.f sqrl1.f

SUBS integer function

indshc.f indfun.f

BLAS integer function

idamax.f

BLAS subroutines

dgemm.f dtrsm.f dgemv.f dswap.f dger.f dscal.f
dtrmm.f dtrmv.f

LAPACK subroutines

dgetrf.f dgetri.f dgetf2.f dlaswp.f xerbla.f dtrtri.f
dtrti2.f

LAPACK integer function

ilaenv.f

LAPACK logical function

lsame.f

1.2 Run-time limitations

Several parameters are set at the outset which limit the physical size of the problem.

```
INTEGER NRMAX, NHMAX, NDCS, LHMAX, NBN, NCFM, LHLH2M
PARAMETER ( NRMAX = 200, NHMAX = 5000, LHMAX = 156,
1           NDCS = LHMAX + 4, NBN = 3, NCFM = 2*NBN + 1,
2           LHLH2M = LHMAX*(LHMAX+2) )
```

If the values are insufficient, then change them and recompile.

- NRMAX is the maximum number of radial grid nodes allowed.
- NHMAX is the maximum number of spherical harmonic radial functions allowed.
- LHMAX is the highest permitted spherical harmonic degree, l .
- NBN is the number of bounding nodes on either side of the finite difference stencil. NBN = 3 is recommended value.

1.3 Outputs from SVENSPEC

The file whose name is specified on the fourth line of the input file has the first 8 lines as follows:

```

LINE 1:Poloidal kinetic energy    =          164.204596
LINE 2:Toroidal kinetic energy    =          275.099295
LINE 3:Total kinetic energy       =          439.303891
LINE 4:Poloidal magnetic energy   =           26.878899
LINE 5:Toroidal magnetic energy   =           18.156107
LINE 6:Total magnetic energy      =           45.035007
LINE 7:Scaled kinetic energy      =           30.091780
LINE 8:Scaled magnetic energy     =            3.084843

```

The variables output are as follows:-

1. $K_P = 0.5 \int_{\text{Vol}} [\mathbf{u}_P]^2 dV$ where \mathbf{u}_P is the poloidal velocity
2. $K_T = 0.5 \int_{\text{Vol}} [\mathbf{u}_T]^2 dV$ where \mathbf{u}_T is the toroidal velocity
3. $K = K_P + K_T$.
4. $M_P = 0.5 \int_{\text{Vol}} [\mathbf{B}_P]^2 dV$ where \mathbf{B}_P is the poloidal magnetic field
5. $M_T = 0.5 \int_{\text{Vol}} [\mathbf{B}_T]^2 dV$ where \mathbf{B}_T is the toroidal magnetic field
6. $M = M_P + M_T$.
7. $K/\text{Vol.}$.
8. $M/\text{Vol.}$.

Following this, a series of lines are written

```

L=      1   0.0000000D+00   2.7267820D+01   2.1497321D+01   0.0000000D+00
L=      2   4.4378594D-01   0.0000000D+00   0.0000000D+00   3.2058448D+00
etc.

```

After the specification of spherical harmonic degree, l , the four numbers given are, from the left,

1. K_P as above, but with velocity \mathbf{u} restricted to poloidal harmonics of degree l only.
2. K_T as above, but with velocity \mathbf{u} restricted to toroidal harmonics of degree l only.
3. M_P as above, but with magnetic field \mathbf{B} restricted to poloidal harmonics of degree l only.
4. M_T as above, but with magnetic field \mathbf{B} restricted to toroidal harmonics of degree l only.

Following this, a series of lines are written

```
M=      0   1.5550156D+00   3.6112105D+01   2.3689472D+01   9.9976062D+00
M=      4   1.4610767D+02   2.2203166D+02   2.3194027D+00   7.1619426D+00
etc.
```

After the specification of spherical harmonic order, m , the four numbers given are, from the left,

1. K_P as above, but with velocity \mathbf{u} restricted to poloidal harmonics of order m only.
2. K_T as above, but with velocity \mathbf{u} restricted to toroidal harmonics of order m only.
3. M_P as above, but with magnetic field \mathbf{B} restricted to poloidal harmonics of order m only.
4. M_T as above, but with magnetic field \mathbf{B} restricted to toroidal harmonics of order m only.

1.4 Sample runs of svenspec

The directory

```
$LEOPACK_DIR/SAMPLERUNS/SVENSPEC
```

contains example input files and model output. Do not under any circumstances edit these files, as these examples should serve as a control for the correct working of the code. After compiling the program, copy the `.input` files to another directory, run the code and confirm that the output agrees with that in the directory.

1.4.1 Example a

We calculate the energy spectrum of the solution described by the 3 files

```
$LEOPACK_DIR/EXAMPLES/FUNDAMENTALS/case1.ints
```

```
$LEOPACK_DIR/EXAMPLES/FUNDAMENTALS/case1.vecs
```

```
$LEOPACK_DIR/EXAMPLES/FUNDAMENTALS/case1.xarr
```

References