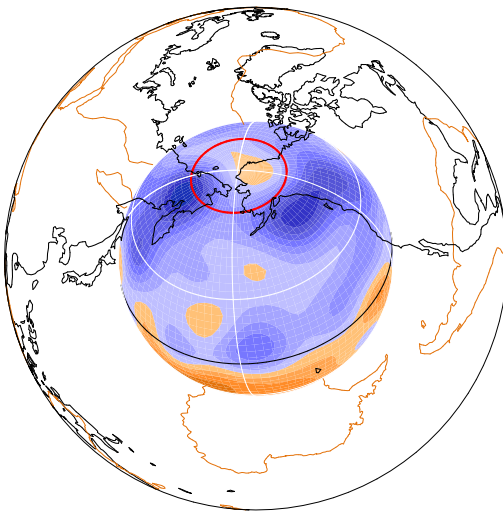


# LEOPACK



**msvip**

**Multiple Solution Vector Ingestion Program**

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# 1 msvip

## Multiple Solution Vector Ingestion Program

Generates an arbitrary - and potentially zero solution vector in standard form (i.e. `.ints`, `.vecs` and `.xarr` files) which satisfies the requested characteristics. To this solution vector, we add scalar multiples of other solution vectors. It is essentially a simple adaption of `rsvfg`.

The primary purpose of this program was to produce single solution vectors from laterally heterogeneous heat-flux convection calculations for plotting. In the time-stepping codes, the homogeneous solution and heterogeneous temperature parts are kept in separate solution vector files: a single file is required for my plotting utilities.

The stand-alone source code version of the program is compiled by typing

```
make msvip
```

within this directory. Once the executable is created, begin execution by typing

```
msvip < inputfile
```

The input file must have the following format

---

```
* input file for msvip
*
plotvec                               : File-name stem
 40      4      4                     : nr  iformf  nnds
0.6666666 1.6666666                   1 0.0 : ri   ro   insf   fac
24      1      0      2      1      12   : lhv isv lhm ism minc mmax
2              2                       : ivelbc ithebc
*
* We now add FAC2 times the following solution vector to
* FAC1 times the existing solution vector.
*
    0.0    0.3                        : FAC1  FAC2
Y22_flux_function.ints
Y22_flux_function.vecs
Y22_flux_function.xarr
*
* We now add FAC2 times the following solution vector to
* FAC1 times the existing solution vector.
*
    1.0    1.0                        : FAC1  FAC2
```

```

Pr0.6_cg200_ch1200_e0.30.ints
Pr0.6_cg200_ch1200_e0.30.vecs
Pr0.6_cg200_ch1200_e0.30.xarr
*
```

Any line in the input file beginning with an asterisk, `*`, is ignored by the program and can thus be used to enter comments and notes.

The fixed set of arguments is as follows

- **filename stem**: First characters in output files to be generated by current run. Running `msvip` with the above input file will create the files `root.ints`, `root.vecs` and `root.xarr`.
- **nr**: number of radial grid nodes requested.
- **iformf**: Flag which chooses the order in which elements in the solution vector are stored. **iformf** can take the values 3 or 4.
- **ri**: Radius of the inner boundary,  $r_i$ .
- **ro**: Radius of the outer boundary,  $r_o$ .
- **insf**: Radial grid node spacings flag.  
`insf = 1` forces evenly spaced grid nodes from ESNAAS and `insf = 2` forces Chebyshev zero spaced nodes from ZCPAAS
- **fac**: Scales the random initial solution. It is probably best to begin with “numerical noise” and hence a low value of **fac**.
- **lhv**: Highest spherical harmonic degree,  $l$  requested for velocity and temperature
- **isv**: Symmetry flag for velocity.
  1. Velocity is equatorially symmetric (ES)
  2. Velocity is equatorially anti-symmetric (EA)
  3. Velocity contains both ES and EA parts
- **lhv**: Highest spherical harmonic degree,  $l$  requested for magnetic field
- **isv**: Symmetry flag for velocity.
  1. Magnetic field is equatorially symmetric (ES)
  2. Magnetic field is equatorially anti-symmetric (EA)

3. Magnetic field contains both ES and EA parts

- **minc**: Increment of wavenumber.
- **mmax**: Maximum wavenumber. For example, **minc** = 2 and **mmax** = 6 will include the wavenumbers {0, 2, 4, 6}.
- **ivelbc**: Velocity boundary condition:
  1. Rigid (no-slip) boundary condition
  2. Stress-free (free-slip) boundary condition
- **ithebc**: Temperature boundary condition:
  1. Fixed temperature at inner and outer boundaries
  2. Fixed temperature at inner boundary and fixed heat-flux at outer boundary
  3. Fixed heat-flux at inner boundary and fixed temperature at outer boundary

After this, arguments are entered in sets of four uncommented lines. The first of these lines contains two real numbers; **FAC1** and **FAC2**. This adds **FAC2** multiplied by the solution vector described the following three lines to **FAC1** multiplied by the solution vector which is currently stored. In the example above, we first zero the new vector (**FAC1** = 0.0) and then add 0.3 multiplied by the solution vector specified by the files

```
Y22_flux_function.ints
Y22_flux_function.vecs
Y22_flux_function.xarr
```

## 1.1 Subprograms required for msvip

### SUBS subroutines

```
esnaas.f zcpaas.f hminda.f shvecf.f bihfrd.f svfrd.f
xarrrd.f assvar.f hmfwt.f svfwt.f xarrwt.f fopen.f
fclose.f svrint.f smitsl.f pmfabf.f fnamer.f gfdcfd.f
mnewtr.f matop.f
```

### SUBS double precision function

```
pvnsrf.f tvsfrf.f pvsfrf.f pmfirf.f chorch.f chorcf.f
chorsf.f pmfsef.f
```

## **SUBS integer function**

`indfun.f`

## **BLAS integer function**

`idamax.f`

## **BLAS subroutines**

`dgemv.f dgemm.f dtrsm.f dswap.f dger.f dscal.f`  
`dtrmm.f dtrmv.f`

## **LAPACK subroutines**

`xerbla.f dgetrf.f dgetri.f dgetrs.f dgetf2.f dlaswp.f`  
`dtrtri.f dtrti2.f`

## **LAPACK integer function**

`ilaenv.f`

## **LAPACK logical function**

`lsame.f`

## **1.2 Run-time limitations**

Several parameters are set at the outset which limit the physical size of the problem.

```
INTEGER NRMAX, NHMAX, LHMAX, NDCS, ISVMAX
PARAMETER ( NRMAX = 300, NHMAX = 6000, LHMAX = 64,
1           NDCS  = LHMAX + 6, ISVMAX = NRMAX*NHMAX )
```

If the values are insufficient, then change them and recompile.

- NRMAX is the maximum number of radial grid nodes allowed.
- NHMAX is the maximum number of spherical harmonic radial functions allowed.
- LHMAX is the highest permitted spherical harmonic degree,  $l$ .

## **1.3 Outputs from MSVIP**

If the filename stem “root” was specified in the input file, the files `root.ints`, `root.vecs` and `root.xarr` will all be created.

## 1.4 Sample runs of msvip

### 1.4.1 Example a

The files `Pr0.6_cg200_ch1200_e0.30.*` describe the homogeneous boundary part of a non-magnetic convection solution which is subjected to a laterally varying heat-flux at the outer boundary. The heat-flux heterogeneity is imposed by the temperature vector stored within the files `Y22_flux_function.*`. The input file `example_a.input` causes `msvip` to produce a single solution vector, described by the files `example_aOUTPUT.*`. This comprises of the full time-dependent part of the convection solution with 30% of the temperature heterogeneity solution added - the scaling which was applied in the `o2ibtctsc3` program used to calculate the solution.