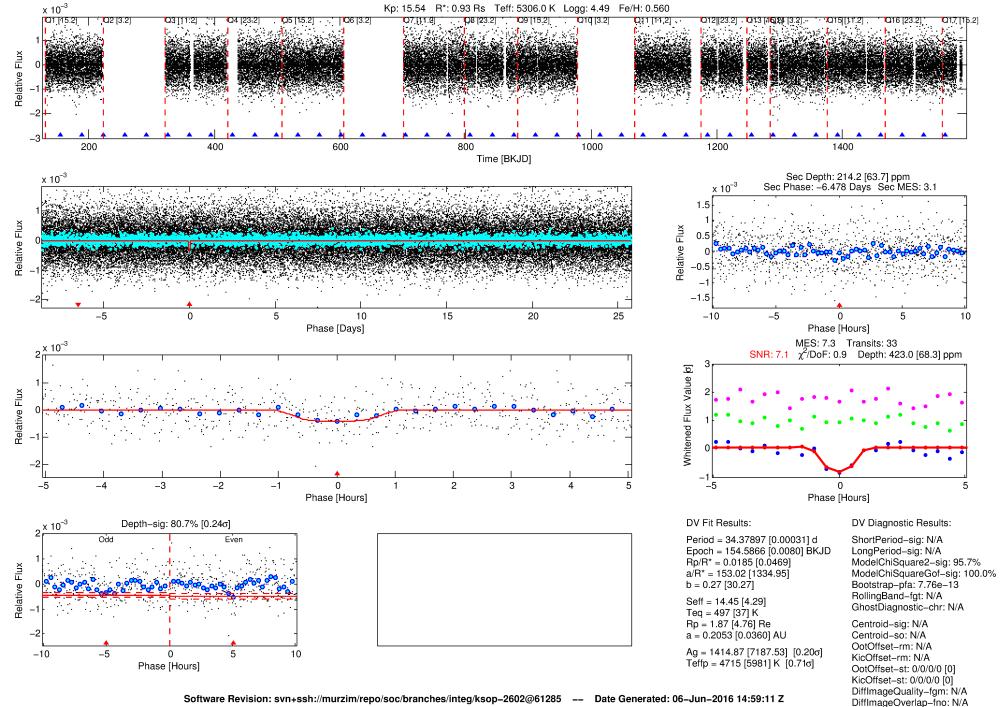
## WARNING: THIS DATA IS SIMULATED, NOT OBSERVED

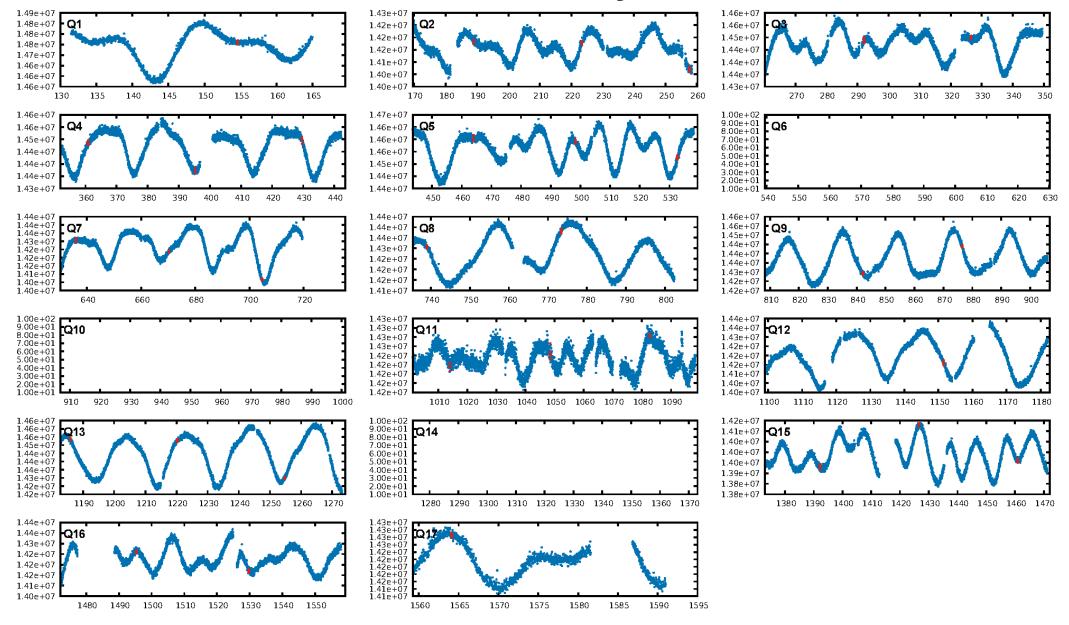
## DV One-Page Summary

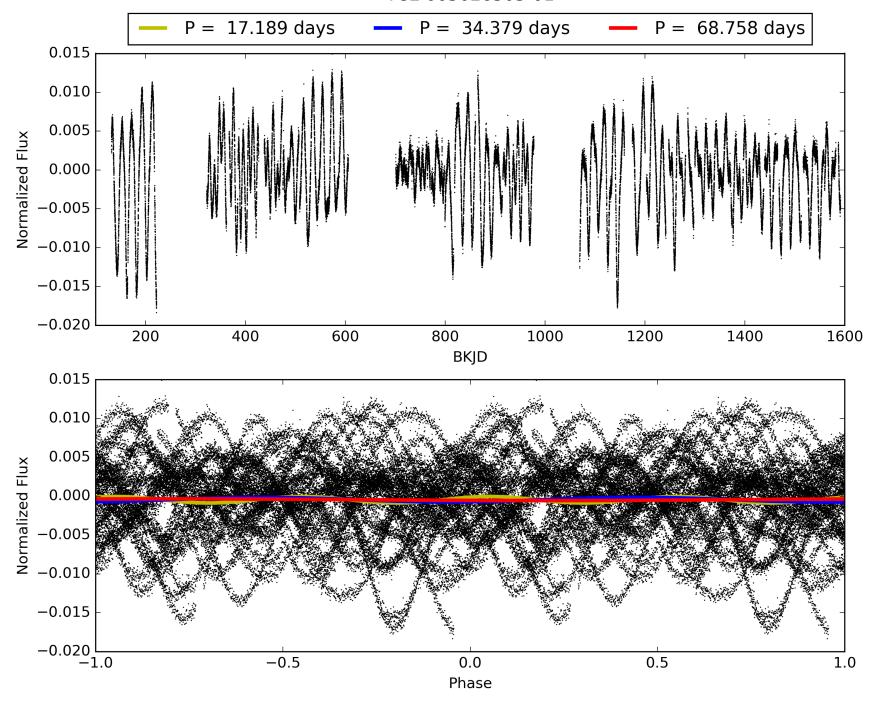
KIC: 5026303 Candidate: 1 of 1 Period: 34.379 d

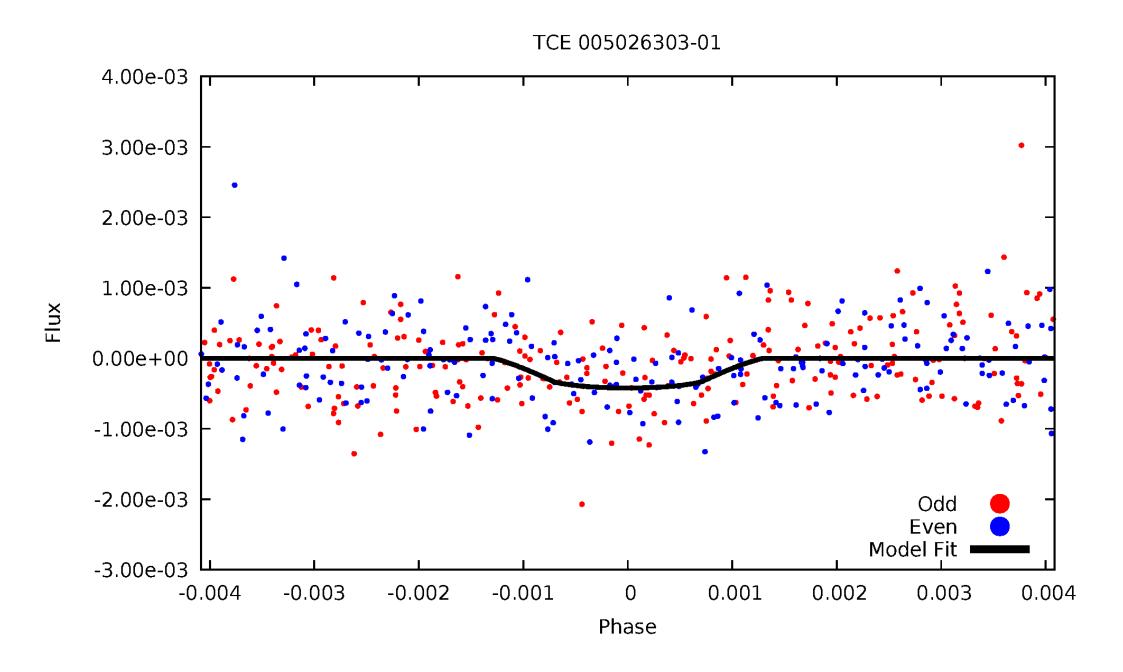
# WARNING: THIS DATA IS SIMULATED, NOT OBSERVED



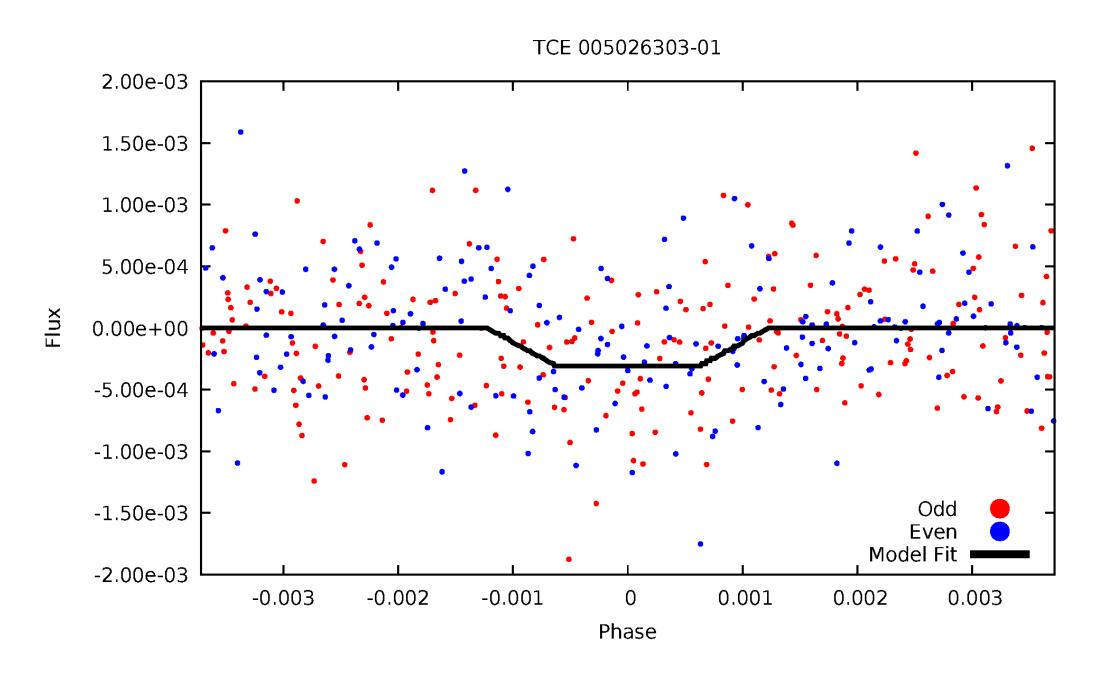
#### TCE 005026303-01, PDC Light Curves



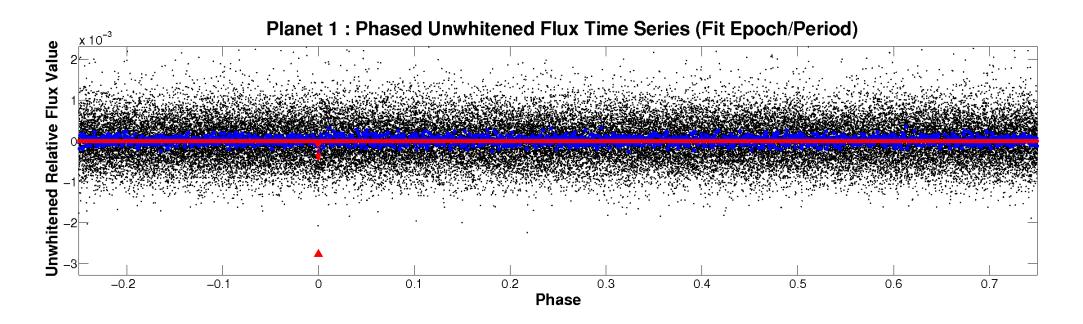


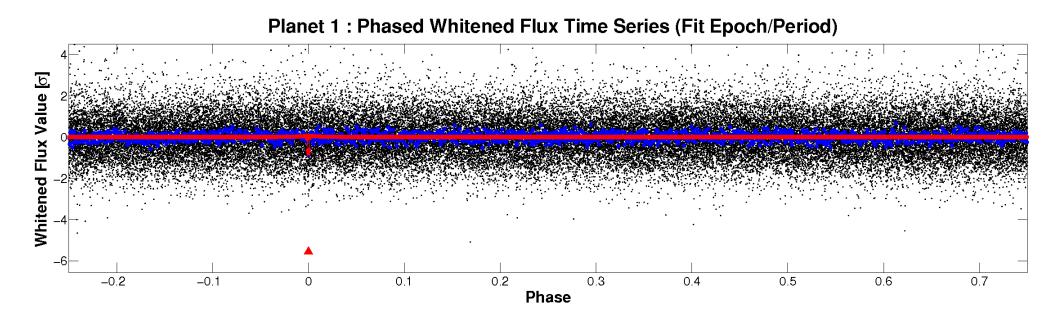


# ALT Odd/Even



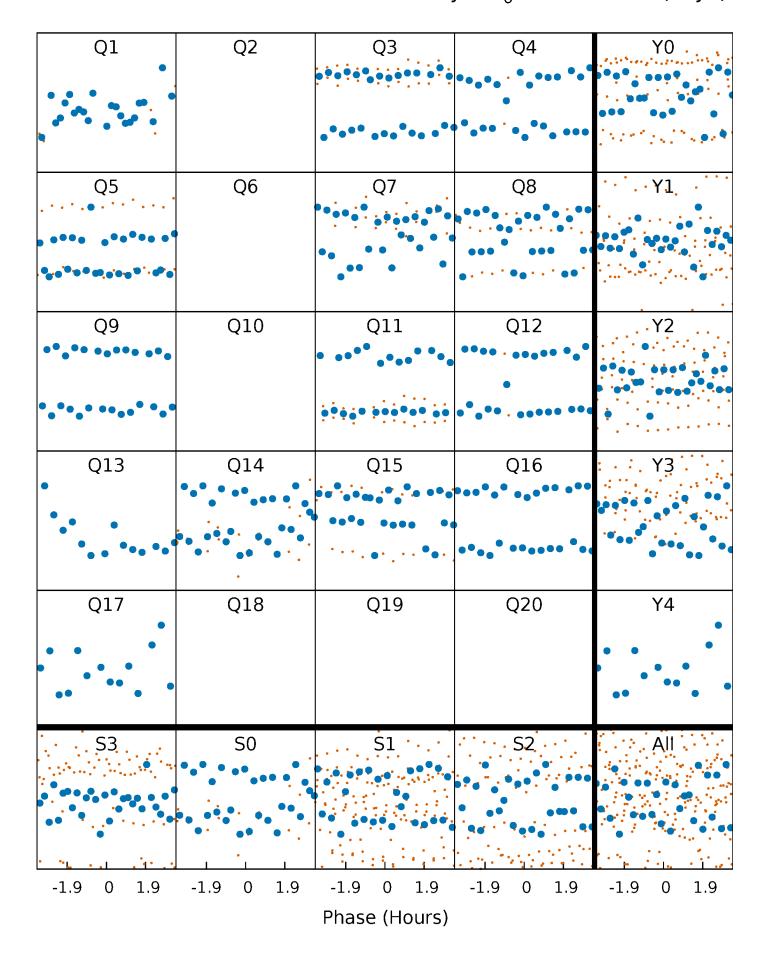
# Non-Whitened Vs. Whitened Light Curve





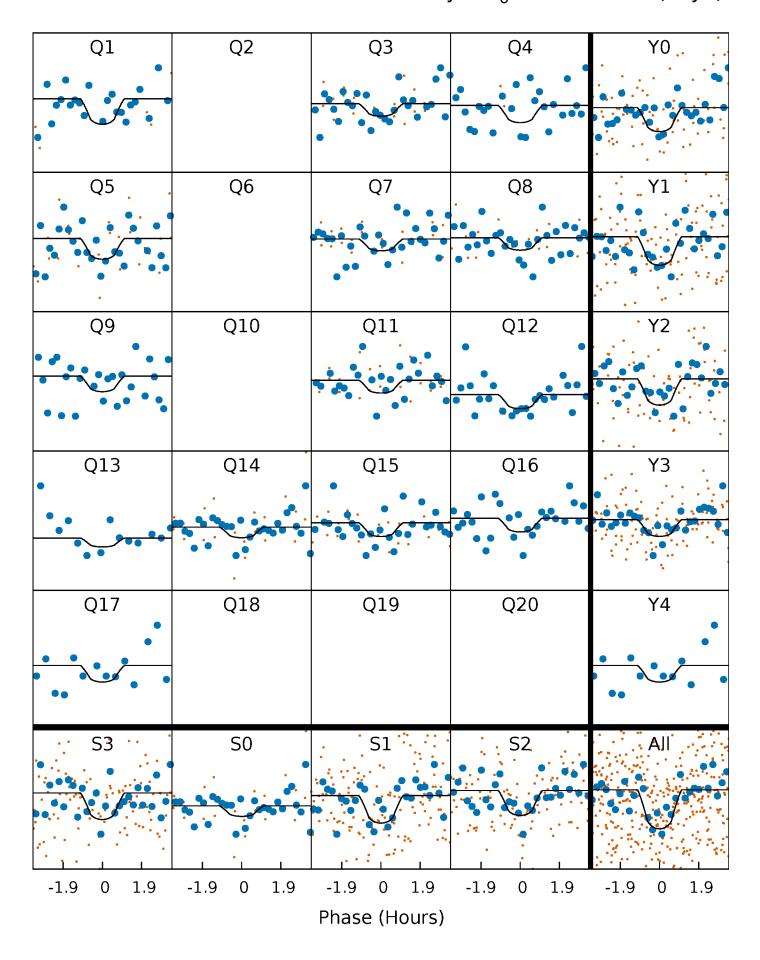
# PDC Quarter-Phased Transit Curves

TCE 005026303-01  $P= 34.378965 Days T_0=154.586571 (BKJD)$ 



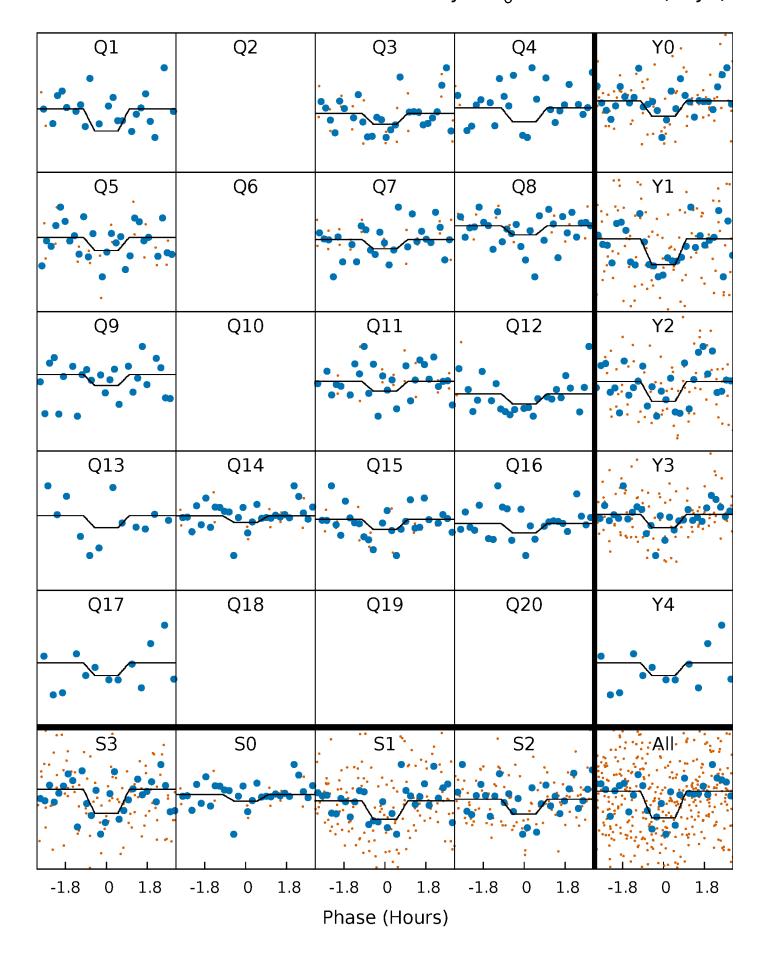
# DV Quarter-Phased Transit Curves

TCE 005026303-01  $P= 34.378965 Days T_0=154.586571 (BKJD)$ 



## Alt. Detrend Quarter-Phased Transit Curves

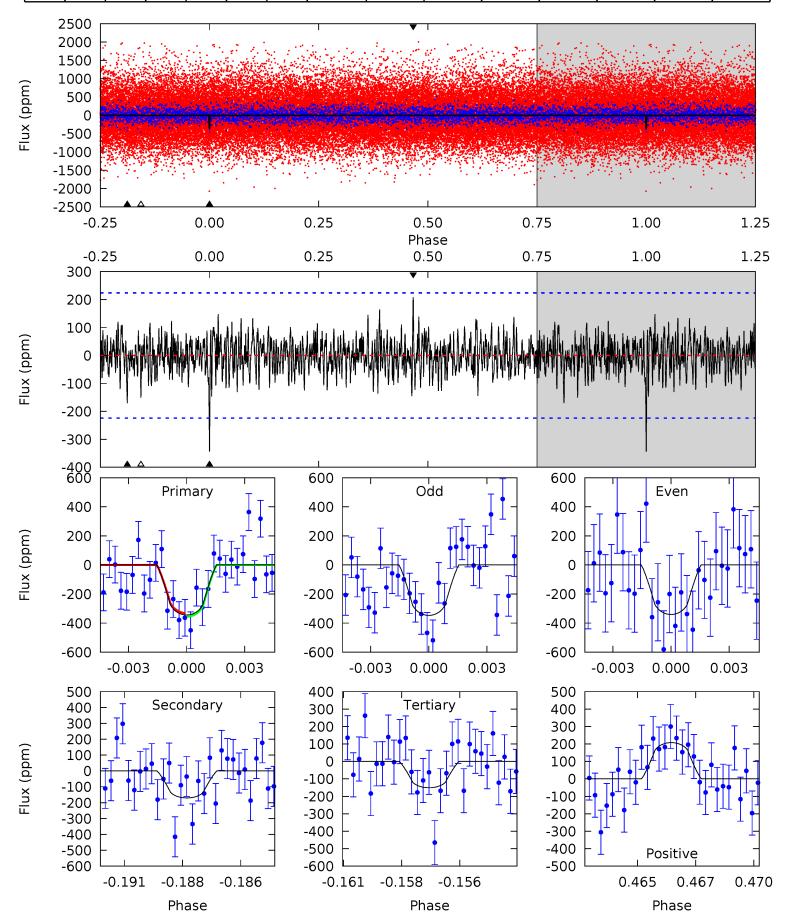
TCE 005026303-01 P= 34.378884 Days  $T_0=154.591821$  (BKJD)



## DV Model-Shift Uniqueness Test

#### 005026303-01, P = 34.378965 Days, E = 120.207606 Days

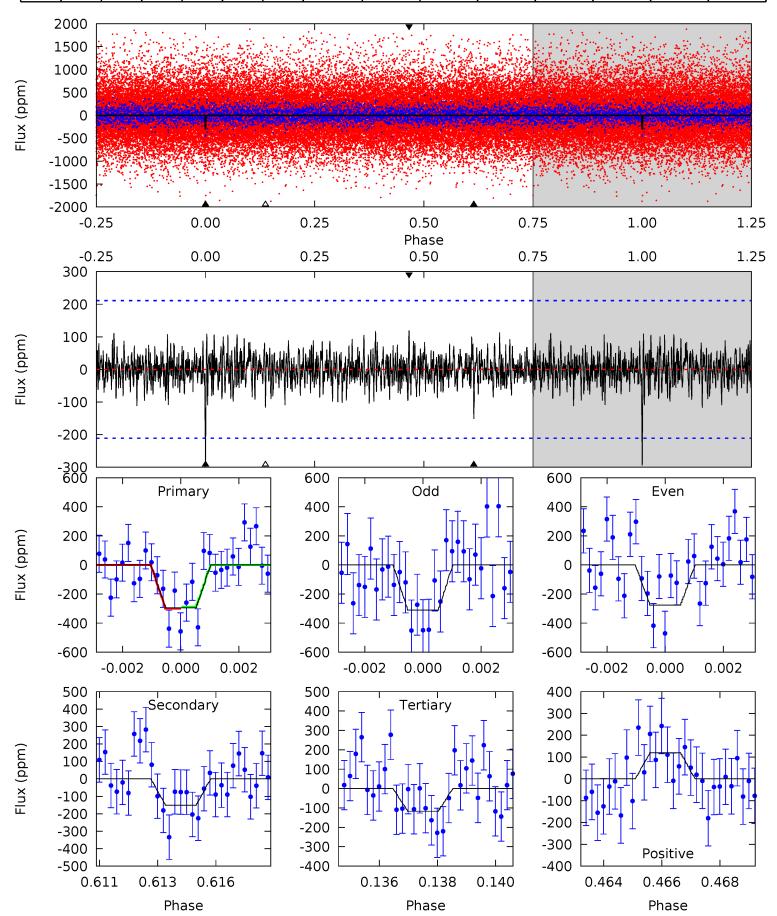
Pri	Sec	Ter	Pos	FA <sub>1</sub>	FA <sub>2</sub>	F <sub>Red</sub>	Pri-Ter	Pri-Pos	Sec-Ter	Sec-Pos	Odd-Evn	DMM	Shape	TAT
8.11	4.02	3.56	4.90	5.29	3.02	1.20	4.56	3.21	0.46	-0.88	0.08	1.27	0.38	0.28



## Alt Model-Shift Uniqueness Test

#### 005026303-01, P = 34.378884 Days, E = 120.212937 Days

Pri	Sec	Ter	Pos	FA <sub>1</sub>	FA <sub>2</sub>	F <sub>Red</sub>	Pri-Ter	Pri-Pos	Sec-Ter	Sec-Pos	Odd-Evn	DMM	Shape	TAT
7.40	3.80	2.93	3.01	5.30	3.04	0.93	4.47	4.39	0.87	0.79	0.44	1.04	0.29	0.15



#### Stellar Parameters For KIC 005026303

	$T_{\rm eff}(K)$	$\log(g)$	[Fe/H]	$R\left(\mathrm{R}_{\odot}\right)$	$M(\mathrm{M}_{\odot})$	$p_{\star} (\text{g} \cdot \text{cm}^{-3})$
	$5306^{+175}_{-159}$	$4.494^{+0.050}_{-0.150}$	$0.560^{+0.050}_{-0.300}$	$0.926^{+0.184}_{-0.085}$	$0.974^{+0.055}_{-0.086}$	$1.729^{+0.418}_{-0.726}$
	+3%/-3%	+1%/-3%	+9%/-54%	+20%/-9%	+6%/-9%	+24%/-42%
Source	PHO1	KIC0	KIC0		DSEP	

KIC = Kepler Input Catalog; PHO = Photometry; SPE = Spectroscopy; AST = Asteroseismology TRA = Transits; DESP = Dartmouth Models; MULT = Multiple Models

#### Secondary Eclipse Parameters for KIC 005026303-01 / KOI

Detrend	Depth (ppm)	$R_p(R_{\bigoplus})$	$T_{max}(K)$	$T_{obs}(K)$	$A_{obs}$
DV	$-170 \pm 42$	$3.85^{+4.08}_{-2.55}$	$705_{-29}^{+38}$	$3569^{+1796}_{-679}$	$249^{+1941}_{-190}$
Alt.	-151±40	$4.11^{+4.22}_{-2.71}$	$707^{+40}_{-29}$	$3489^{+1663}_{-670}$	$216^{+1452}_{-163}$

 $T_{max}$  = Theoretical Maximum Planetary Temperature  $T_{obs}$  = Observed Planetary Temperature (Assuming A=0.3)  $A_{obs}$  = Observed Albedo (Assuming T=0)

If a secondary eclipse is present, the system is likely an EB if  $T_{obs} \gg T_{max}$  AND  $A_{obs} \gg 1.0$ 

# UKIRT Image

