

Detecting Exoplanet Magnetic Fields

Stevo Bailey

Outline

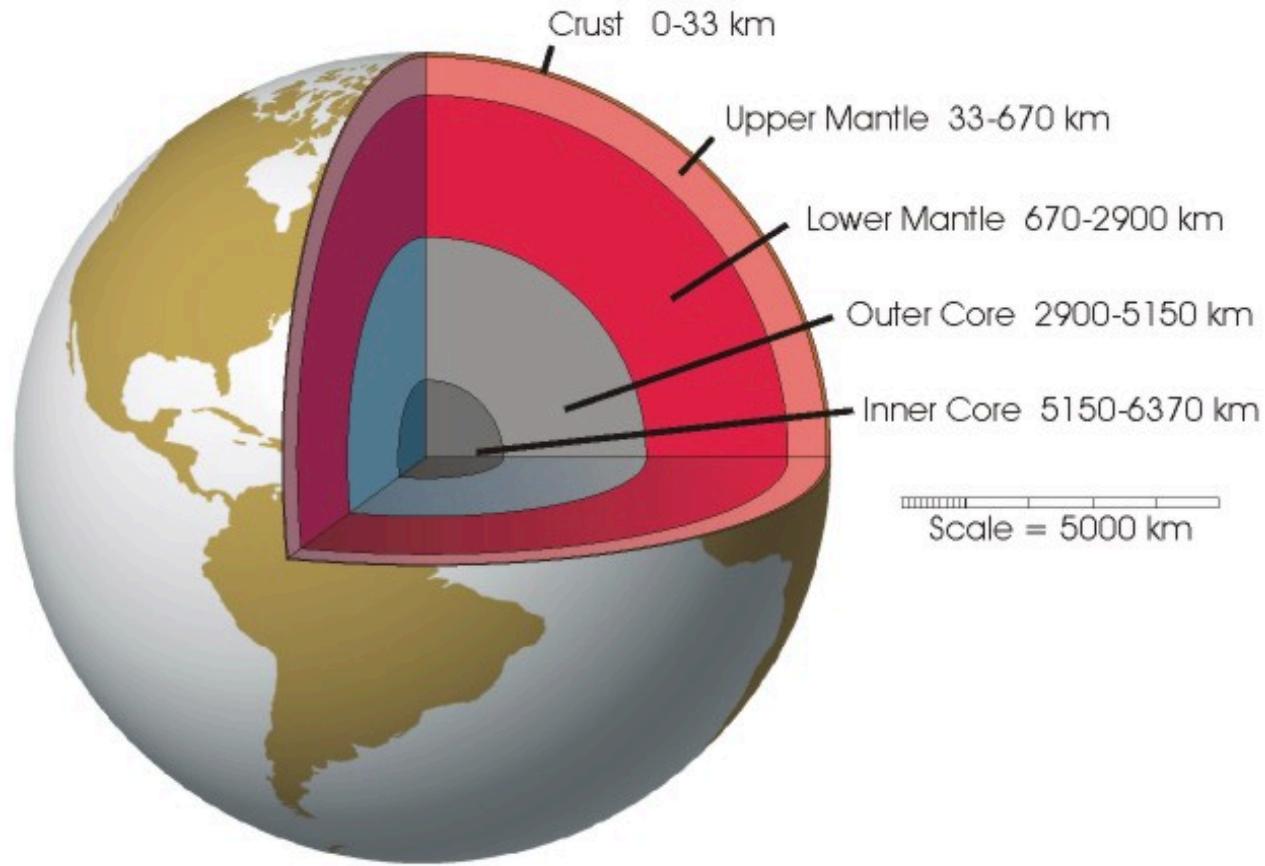
- Motivation
- Exoplanet modeling
- Improving detection
- Limitations
- Conclusions

Motivation

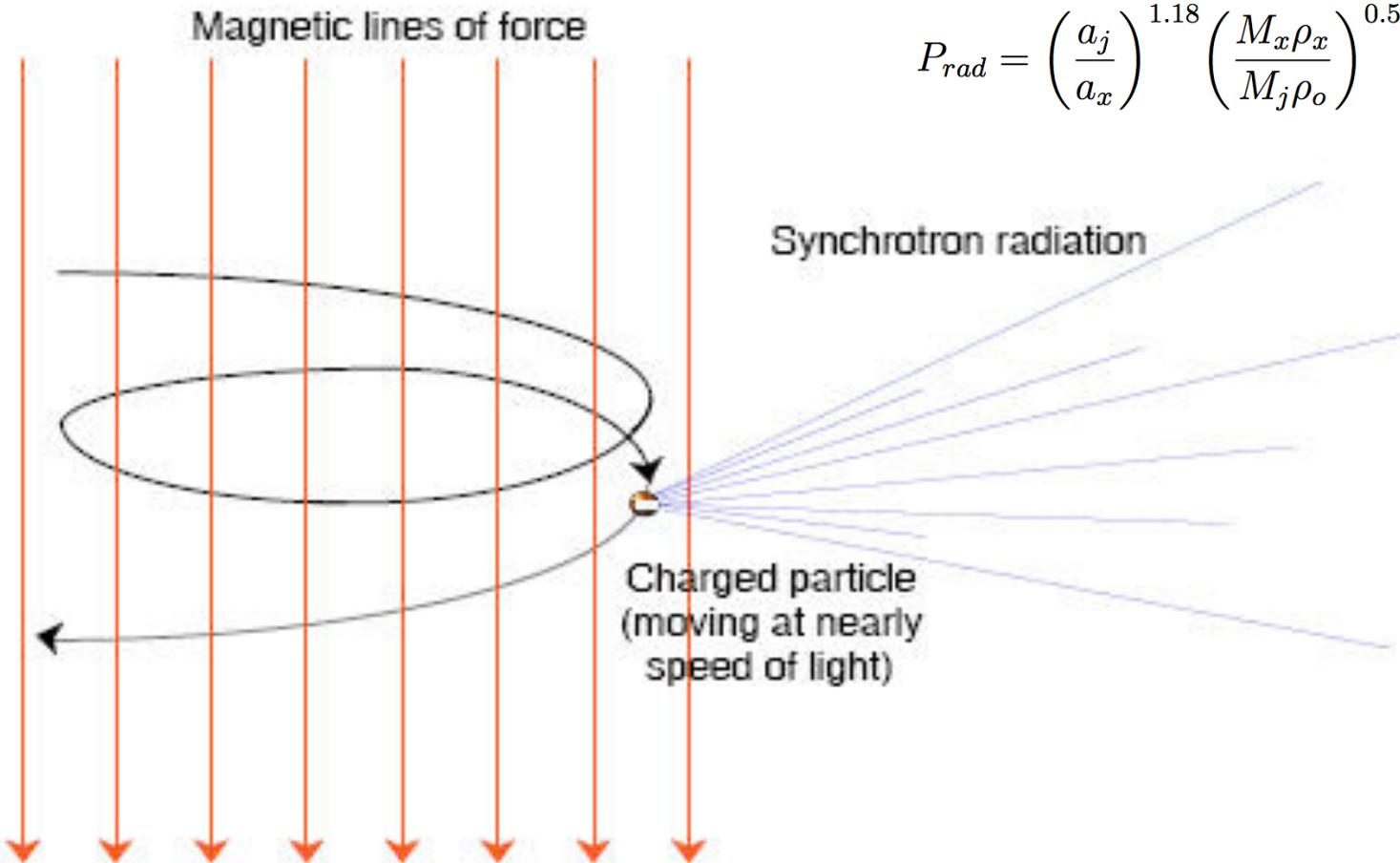


Exoplanet Model

- Differentiated planet
- Earth-like boundary conditions
- Strong convection
- Active dynamo



Cyclotron Emission

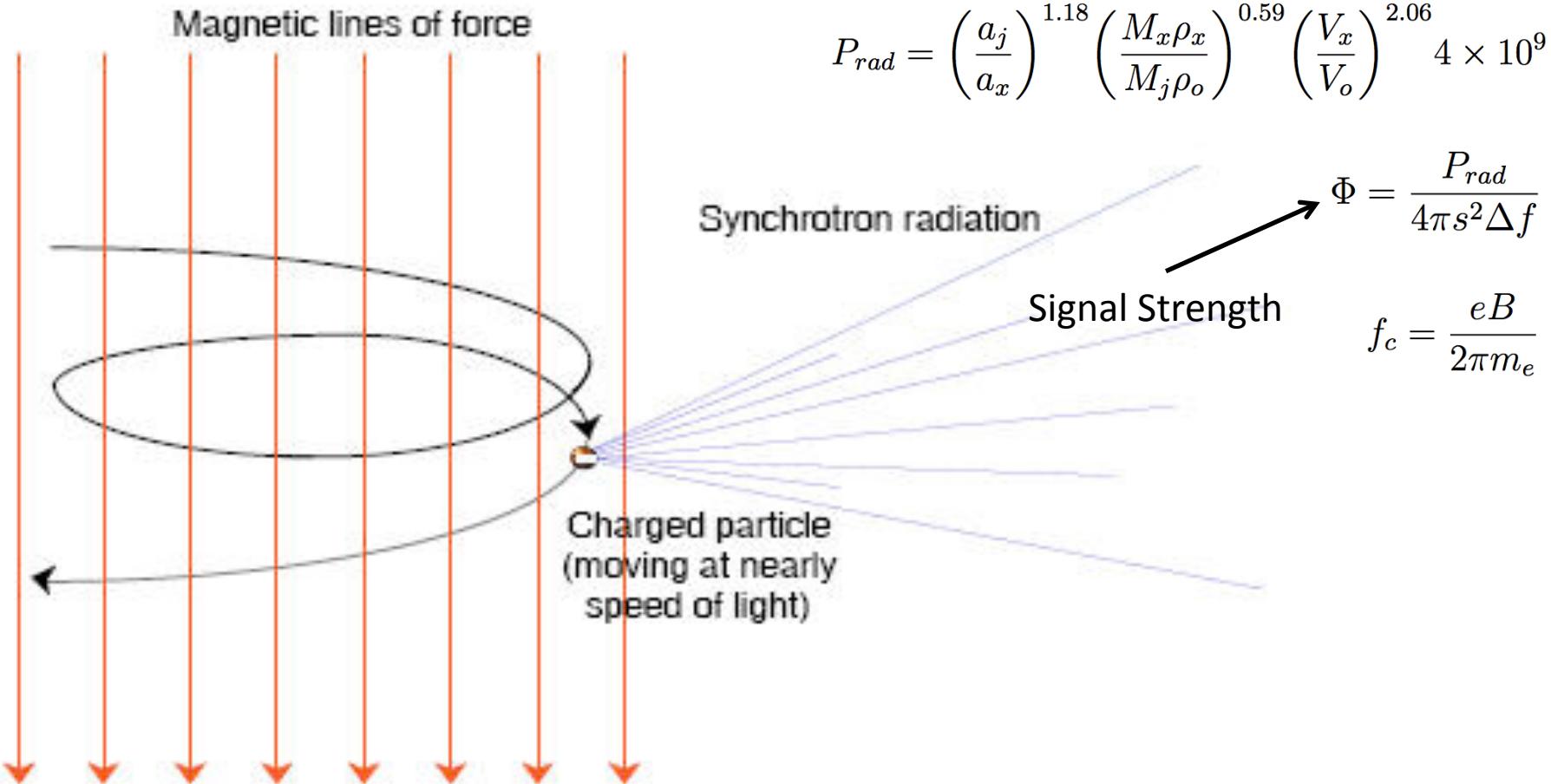


$$P_{rad} = \left(\frac{a_j}{a_x} \right)^{1.18} \left(\frac{M_x \rho_x}{M_j \rho_o} \right)^{0.59} \left(\frac{V_x}{V_o} \right)^{2.06} 4 \times 10^9$$

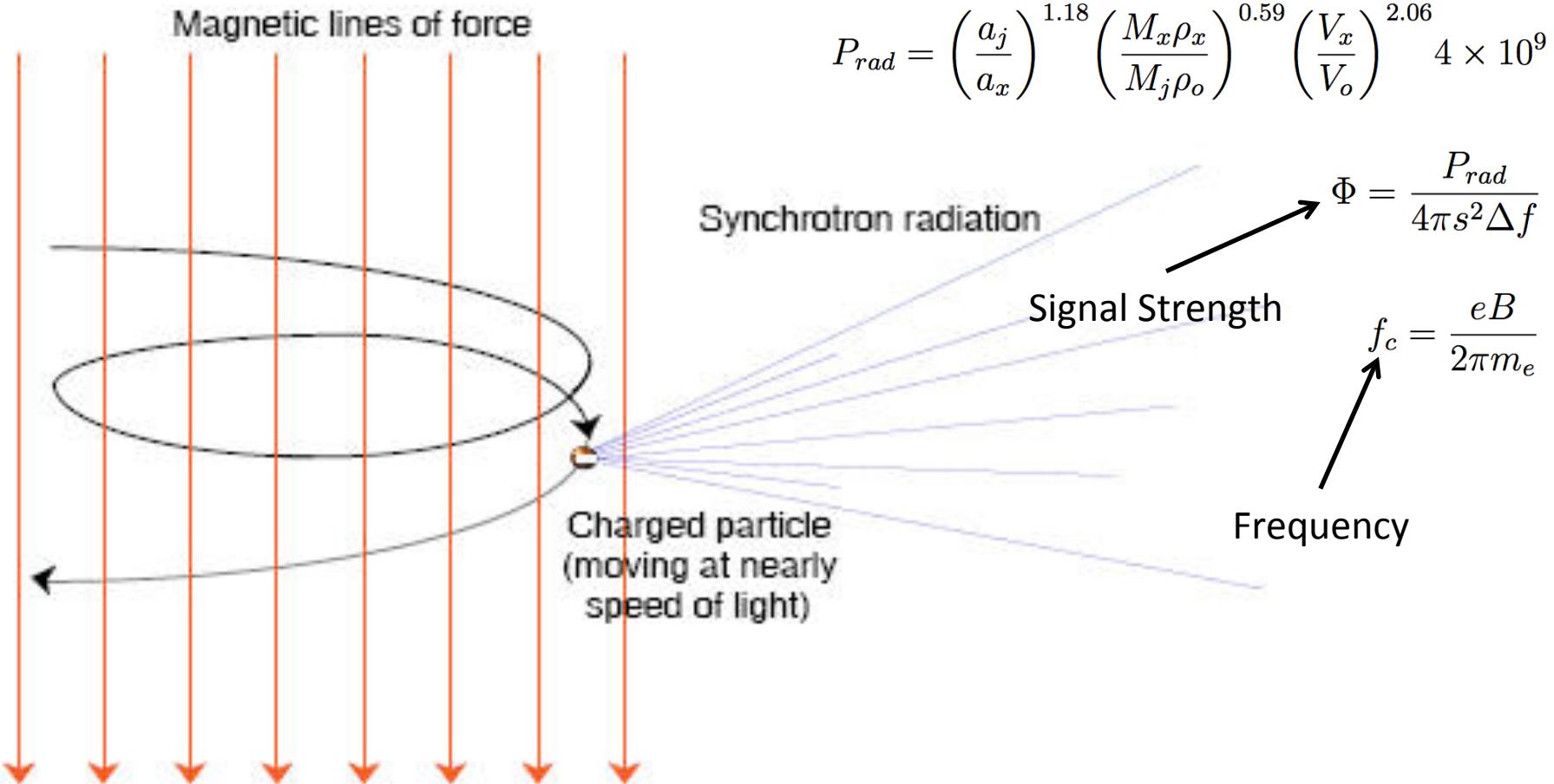
$$\Phi = \frac{P_{rad}}{4\pi s^2 \Delta f}$$

$$f_c = \frac{eB}{2\pi m_e}$$

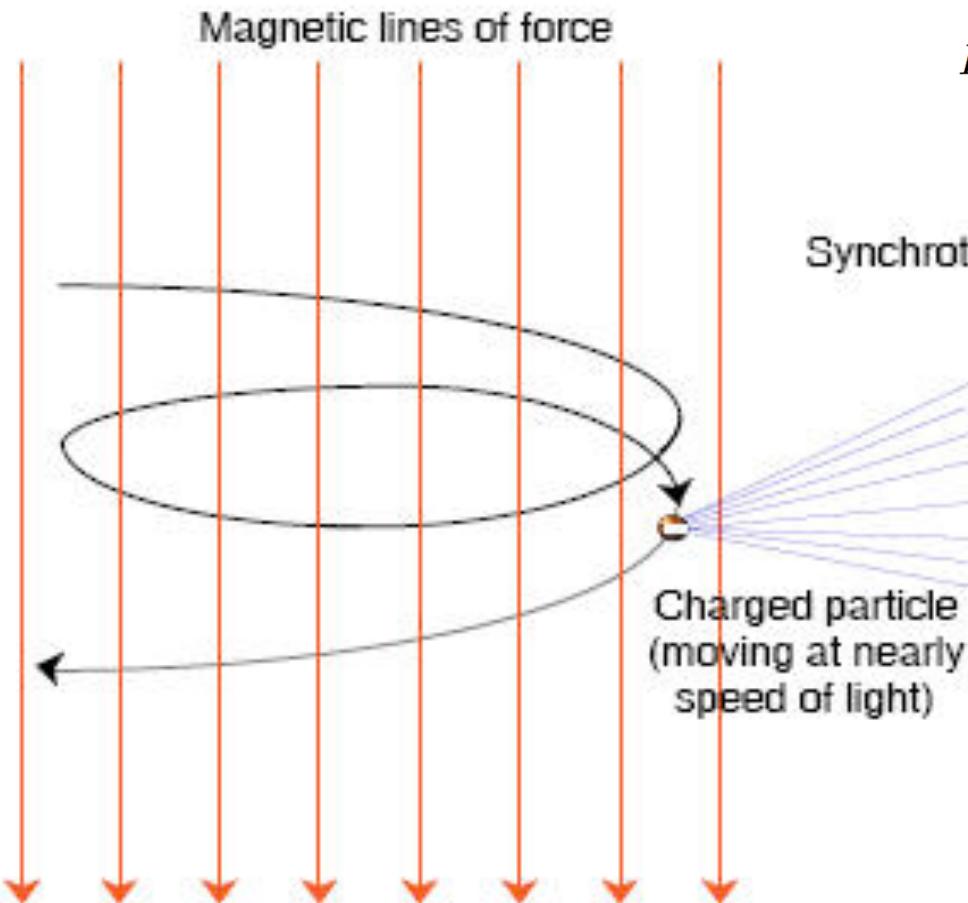
Cyclotron Emission



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Synchrotron radiation

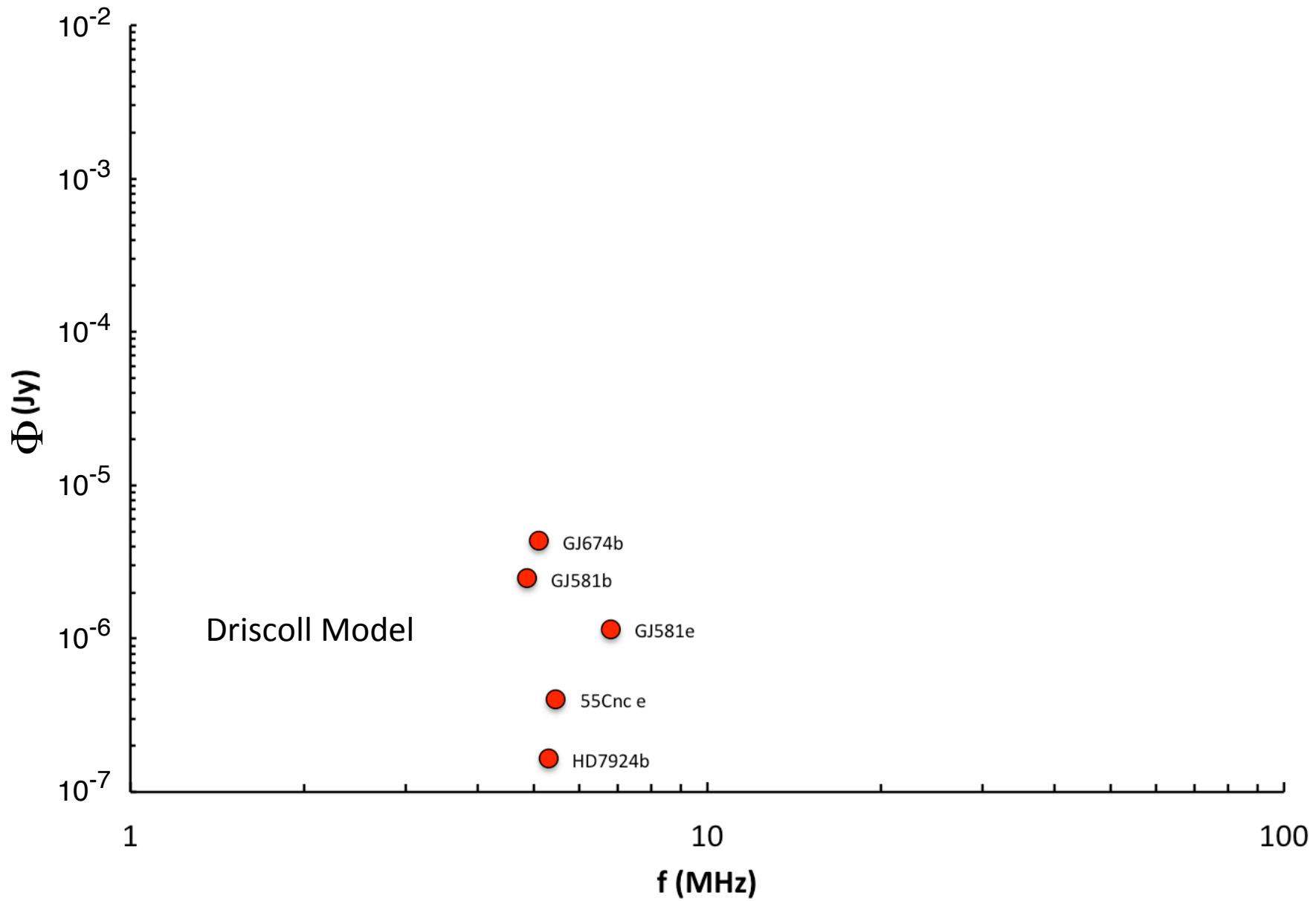
Signal Strength

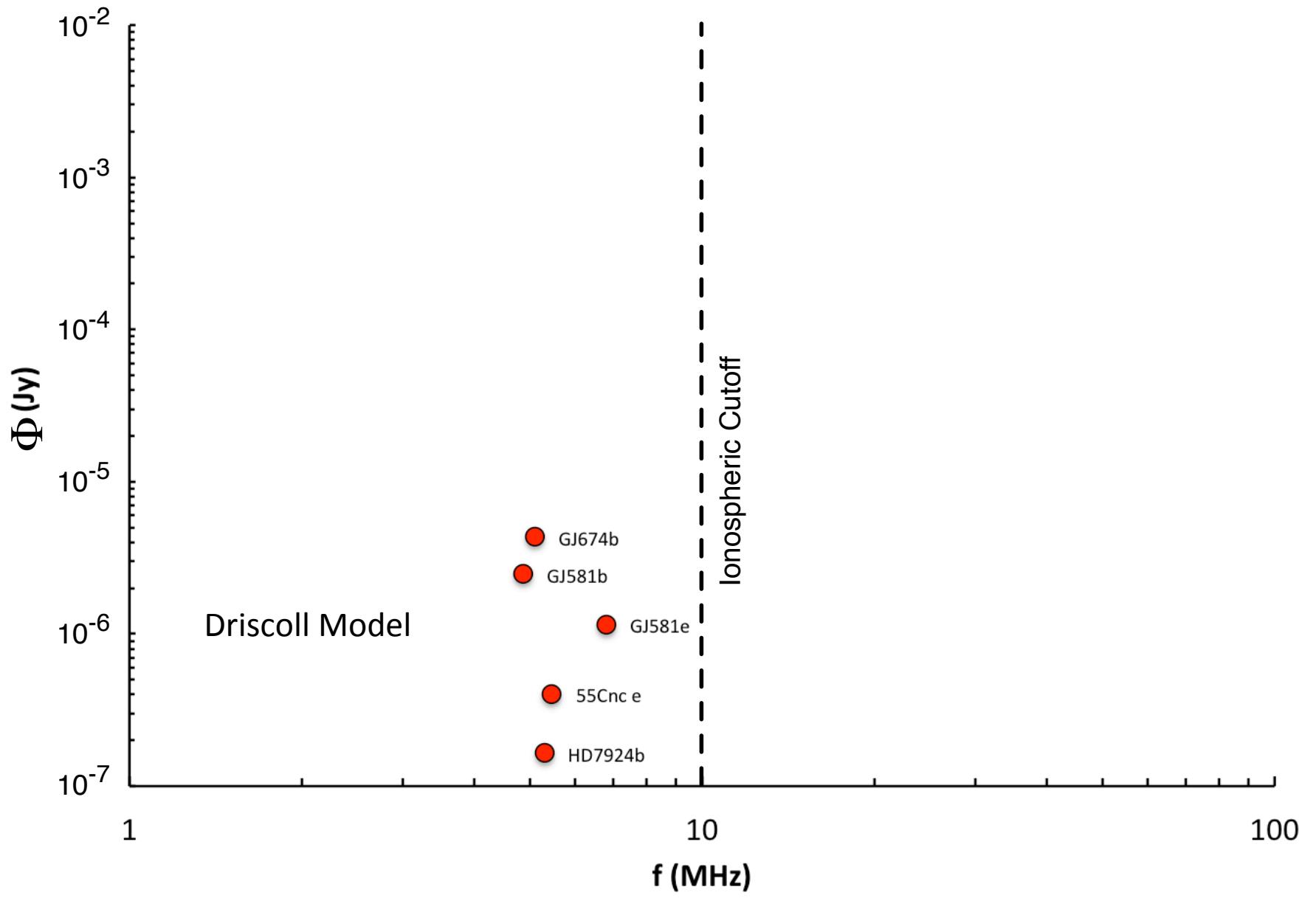
Frequency

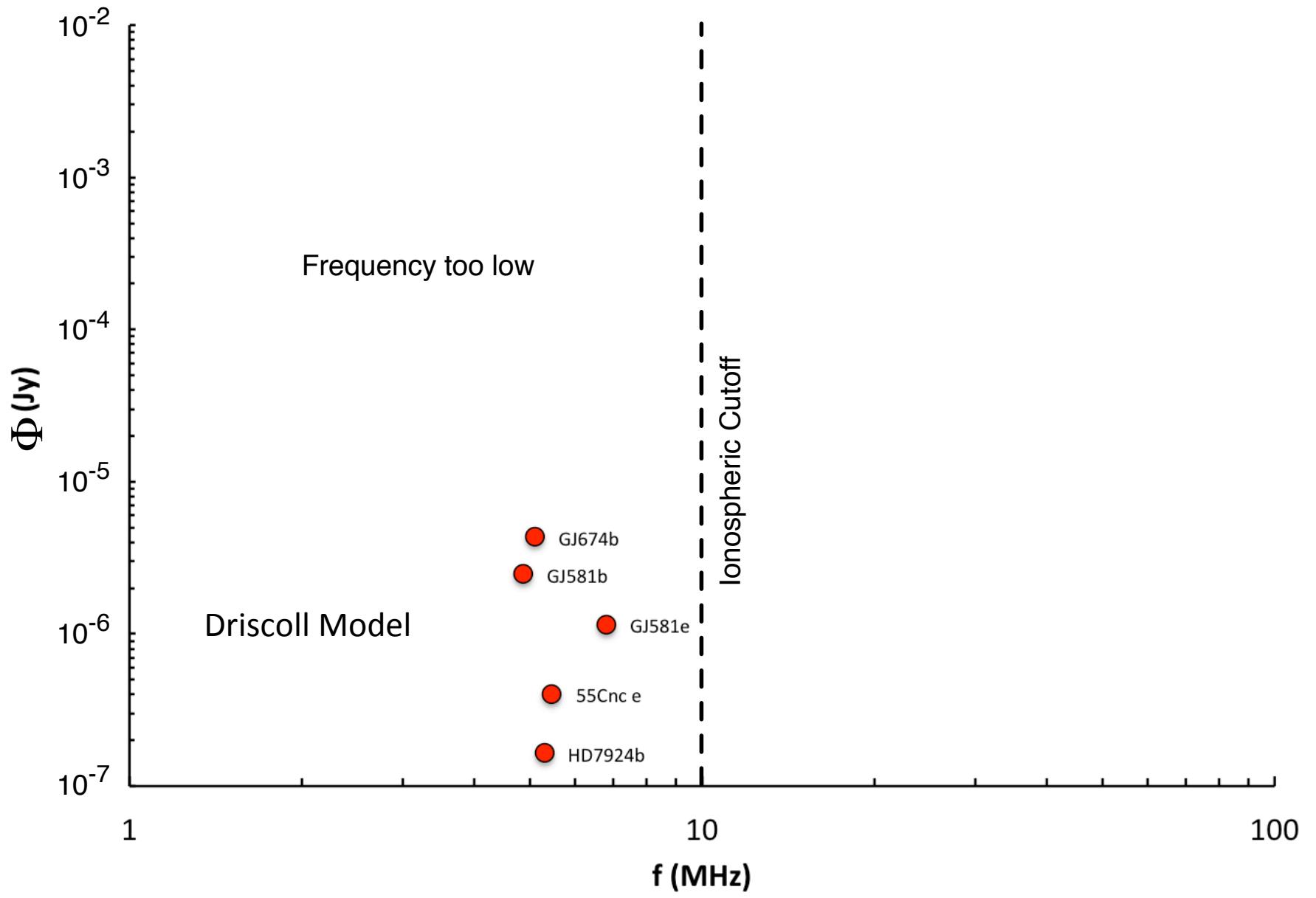
Solar Wind Properties

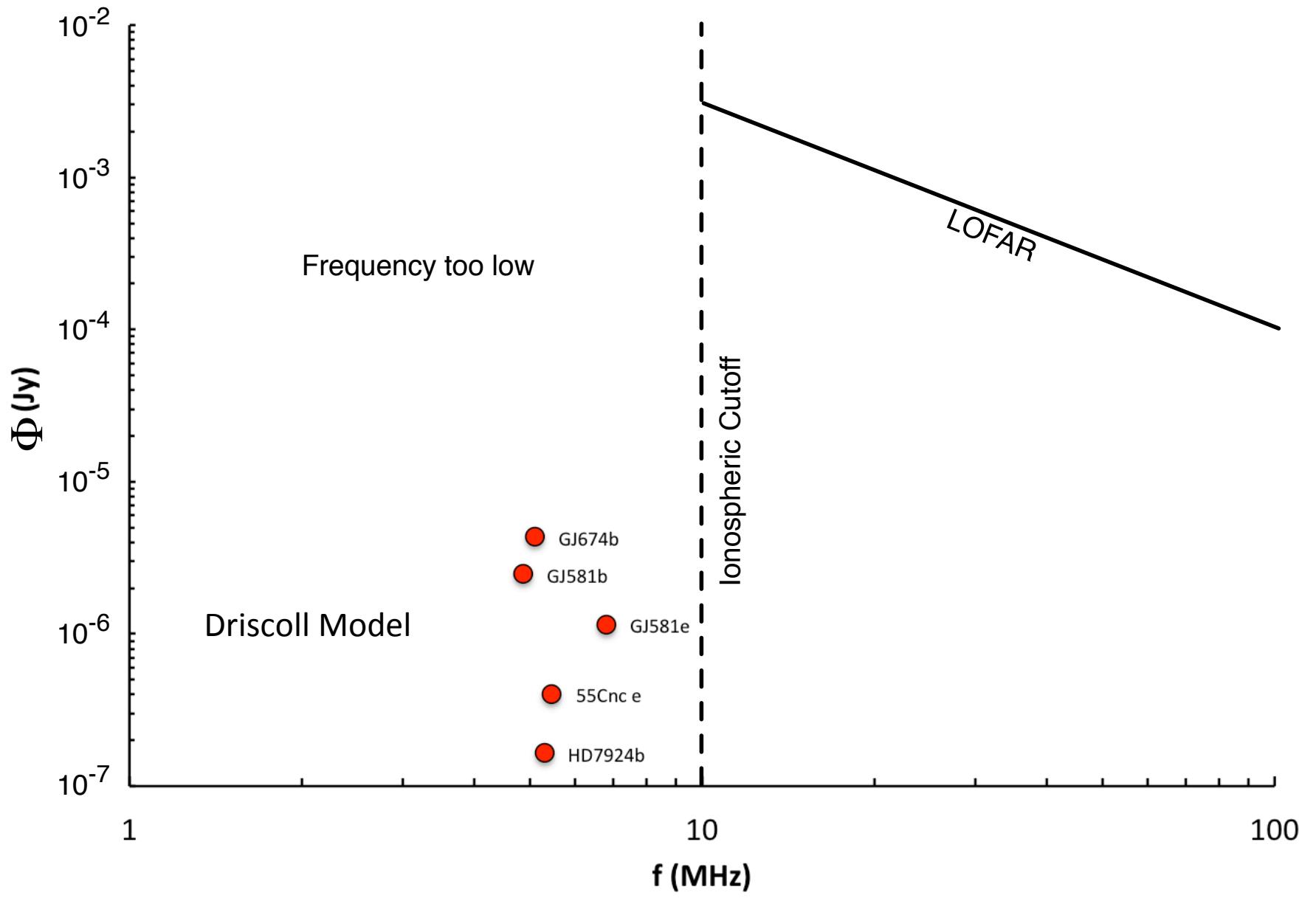
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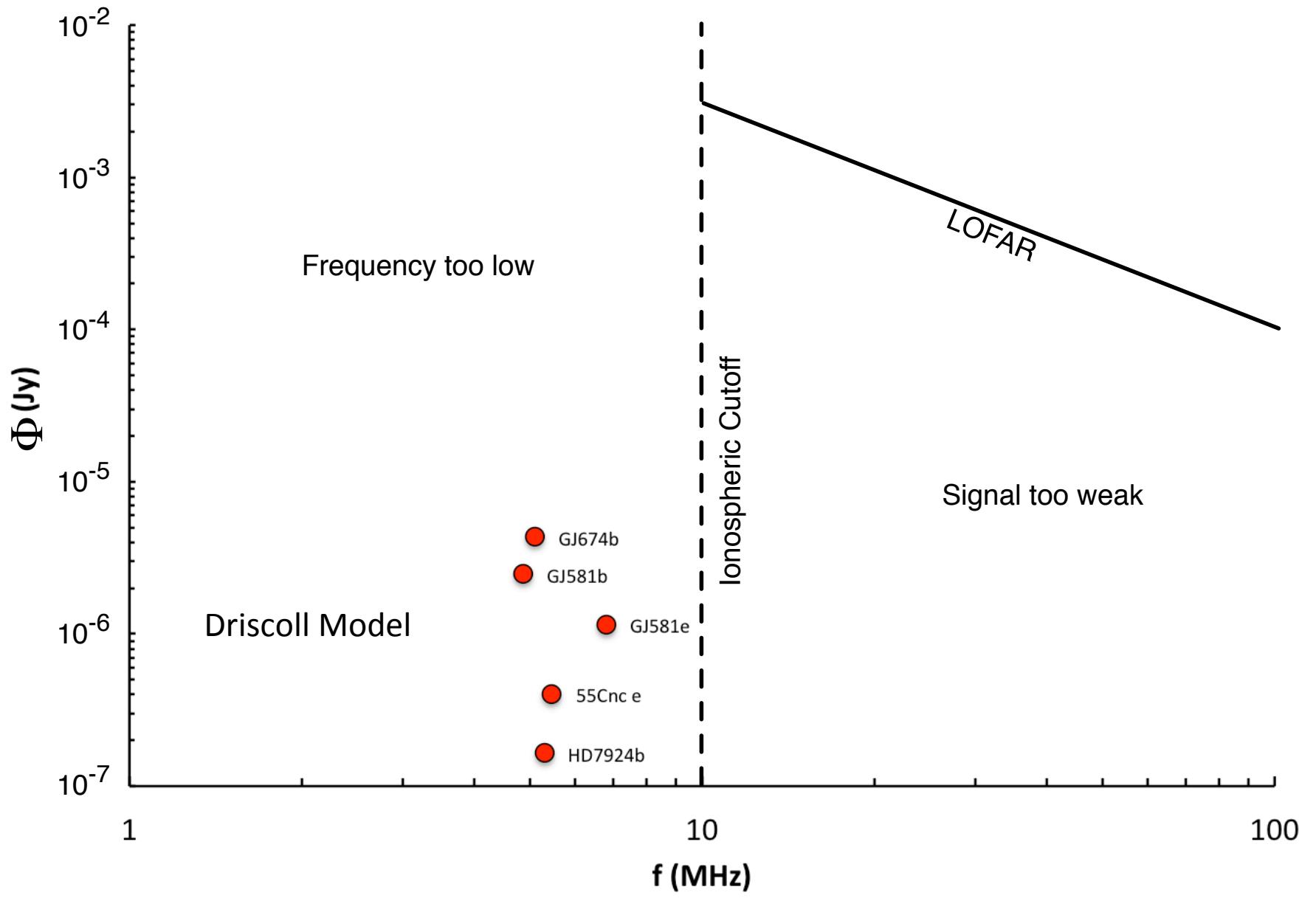
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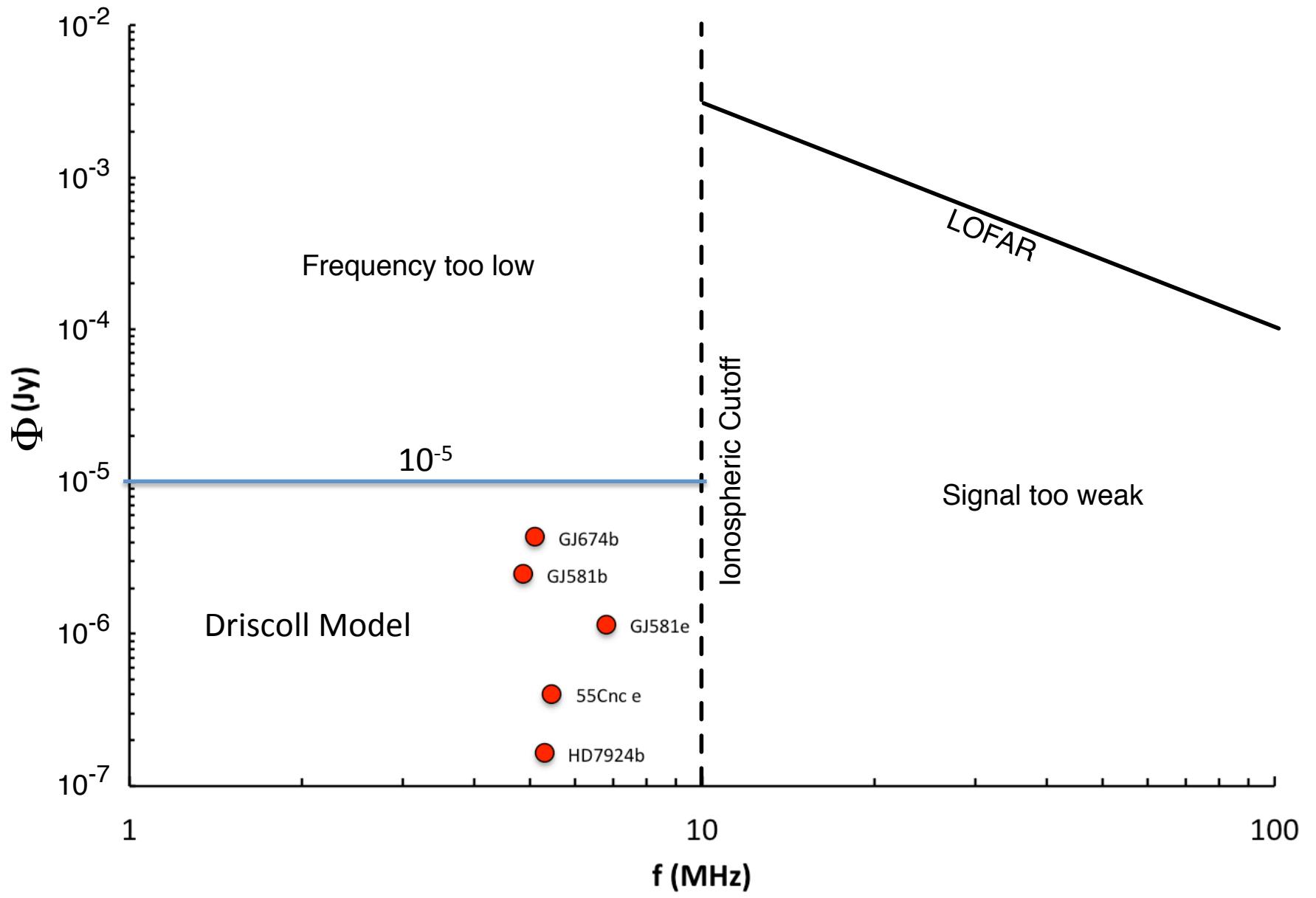


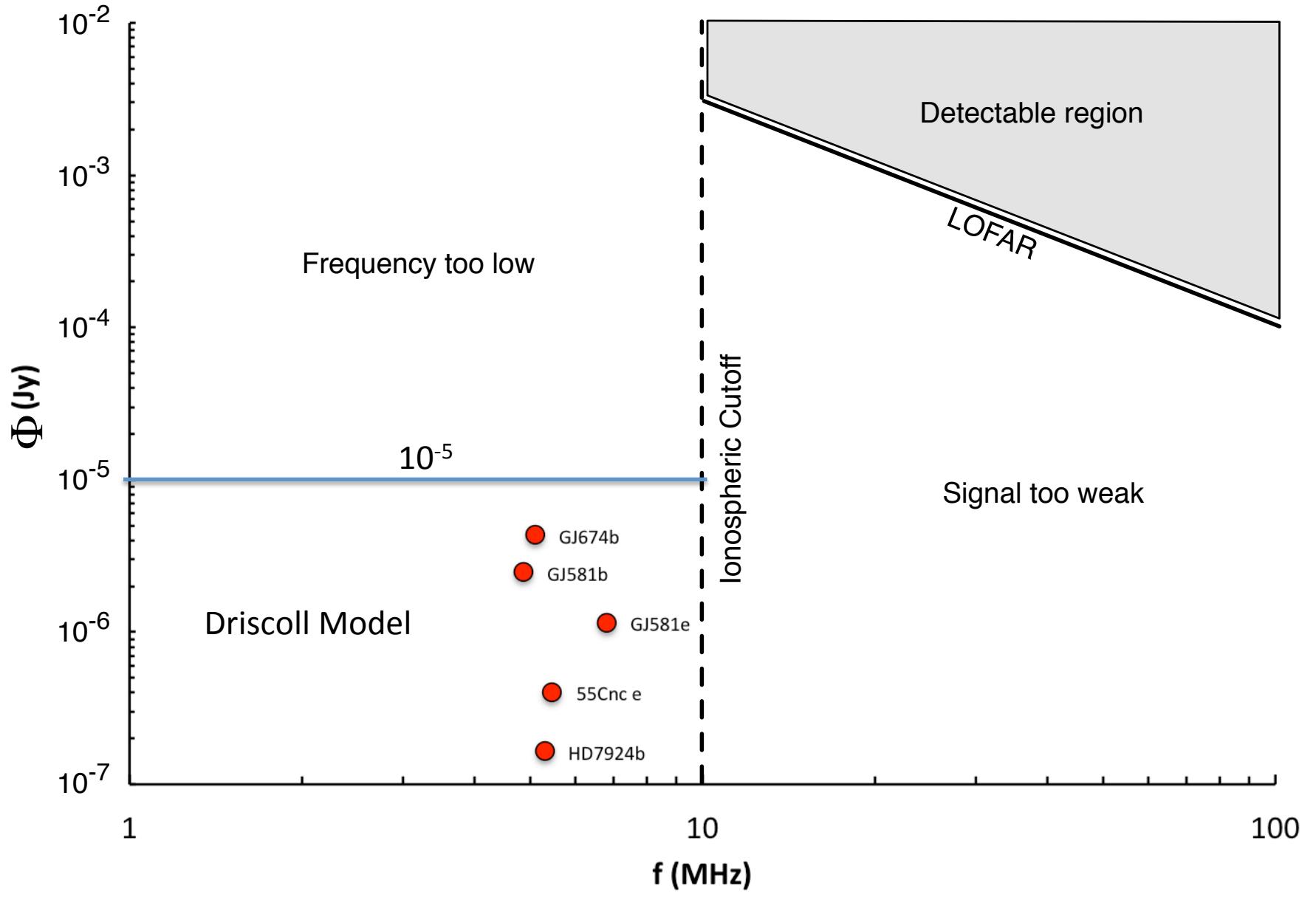












Improving detection

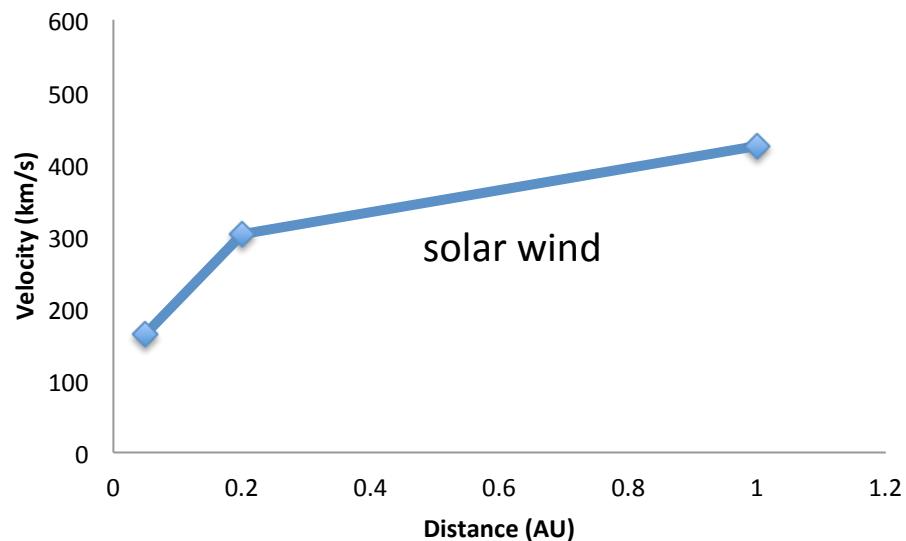
1. Boosting signal strength
2. Better detection range

Improving detection

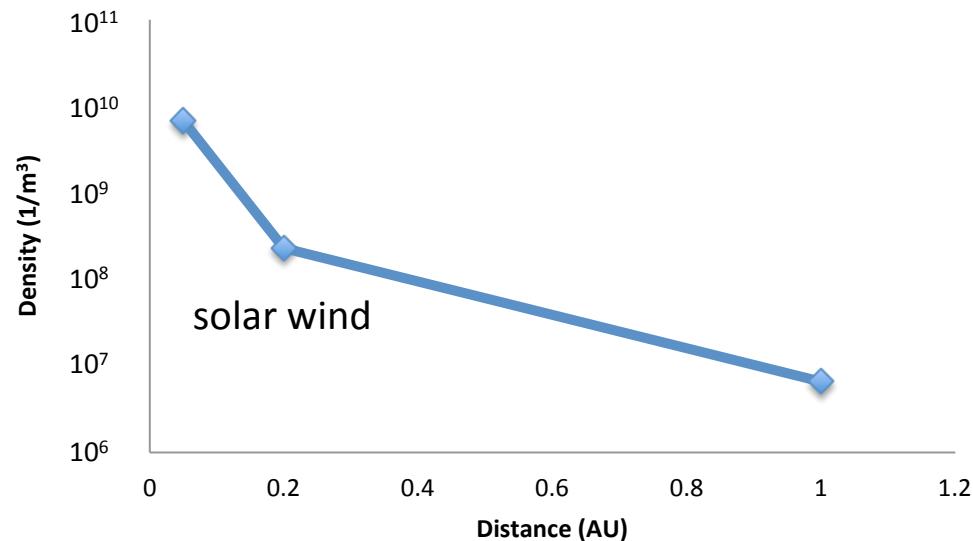
- 1. Boosting signal strength**
2. Better detection range

Solar Wind

Particle Velocity



Particle Density

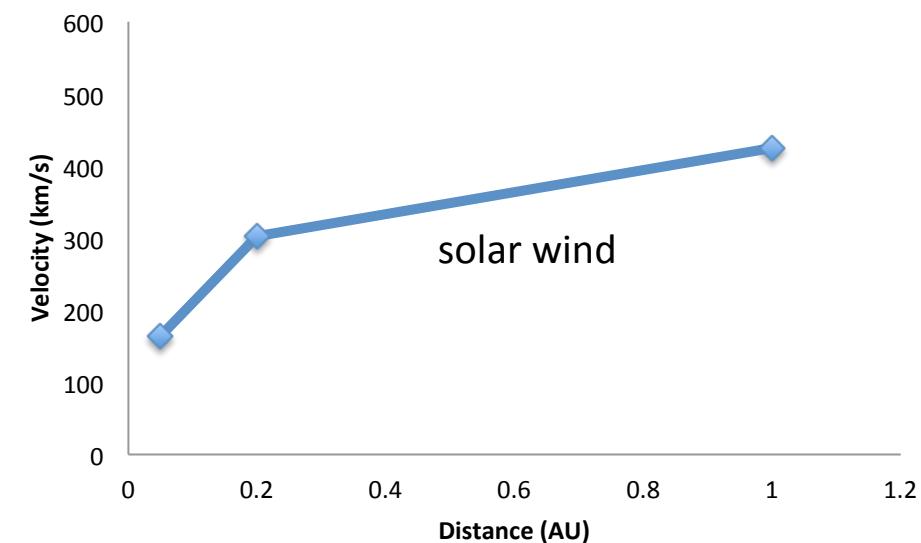


Coronal Mass Ejections!

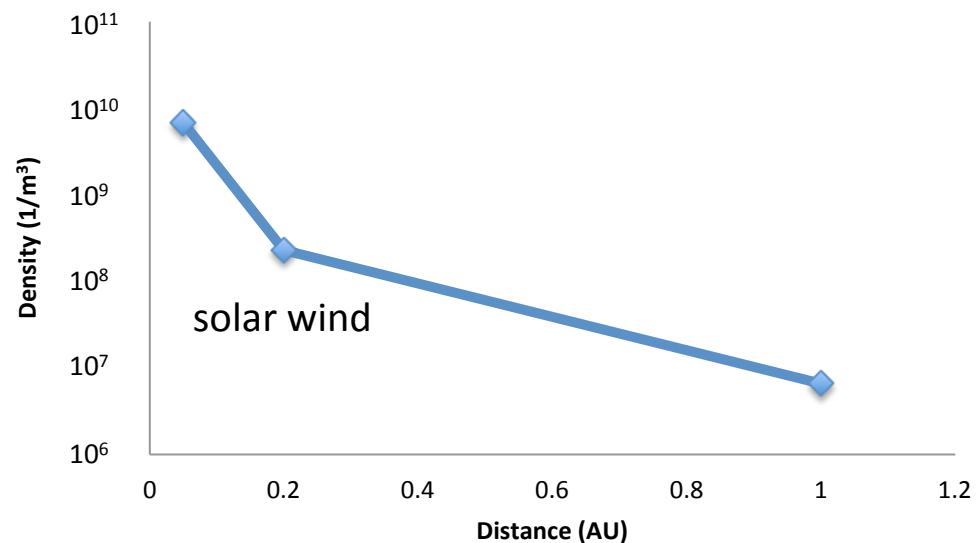


Solar Wind

Particle Velocity

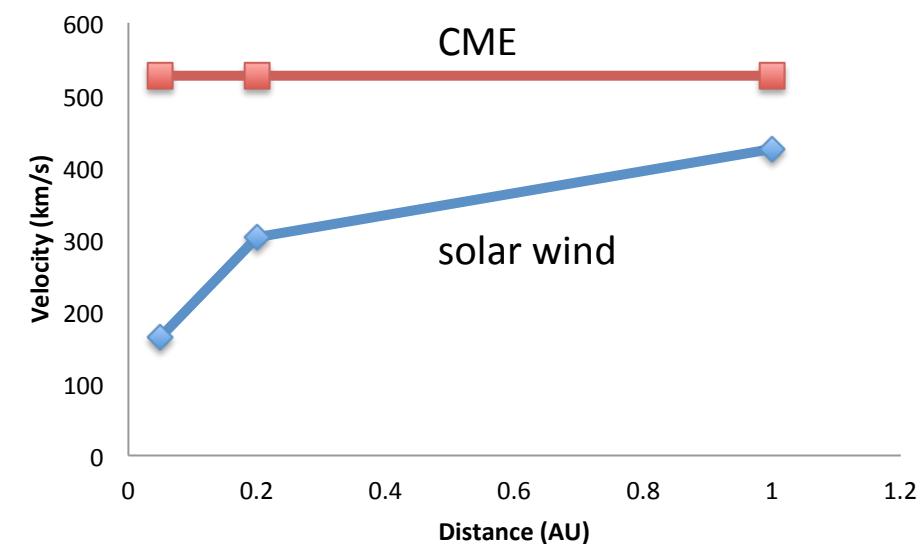


Particle Density

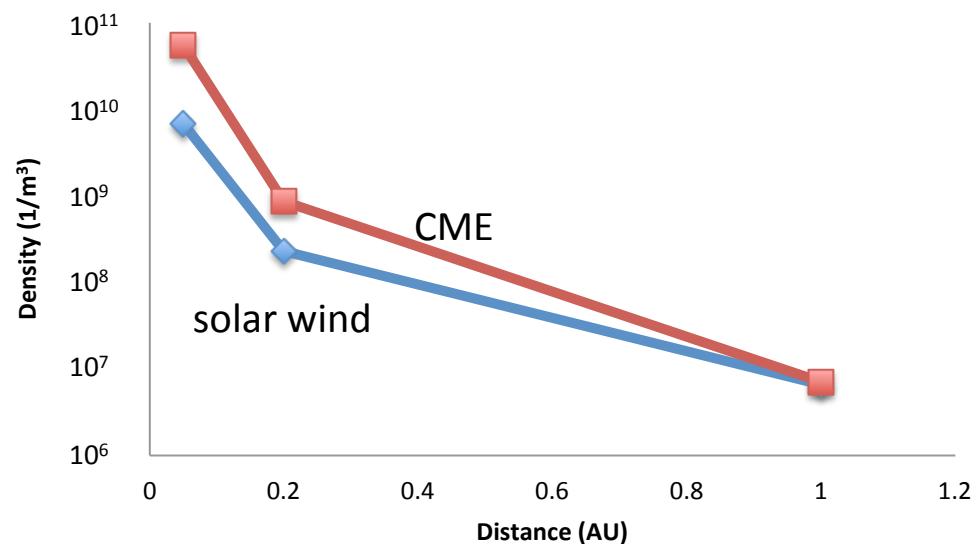


Coronal Mass Ejections (CMEs)

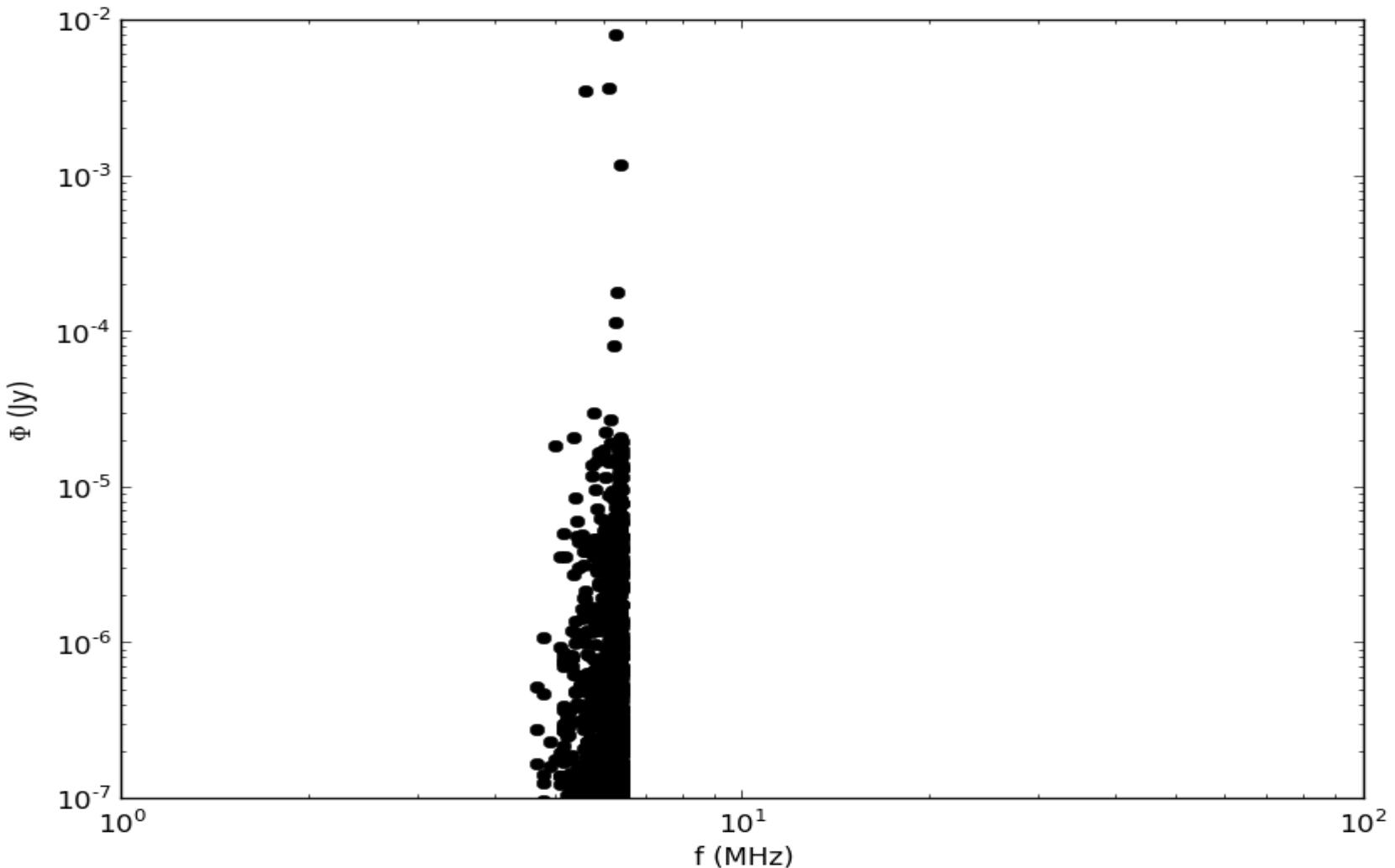
Particle Velocity



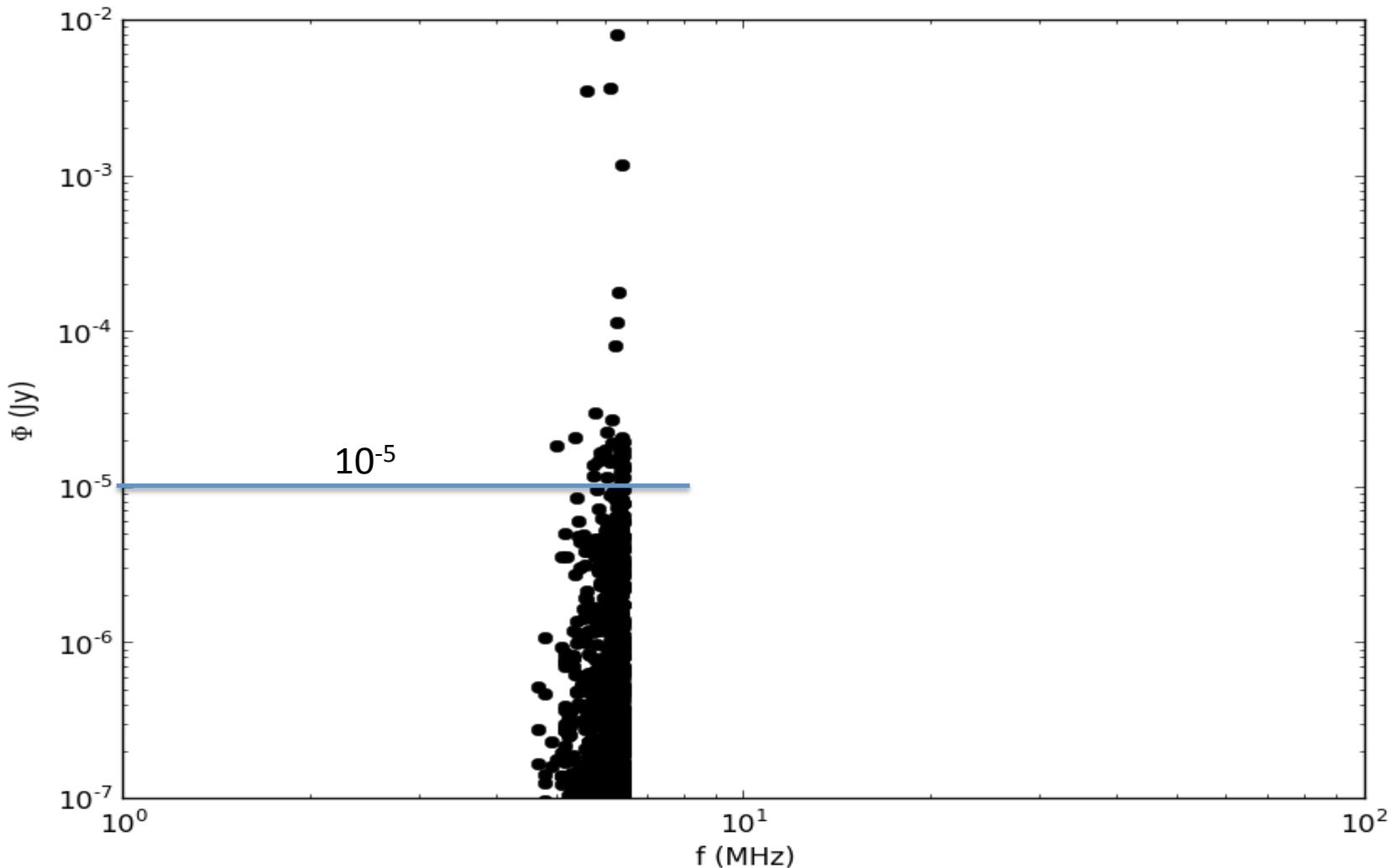
Particle Density



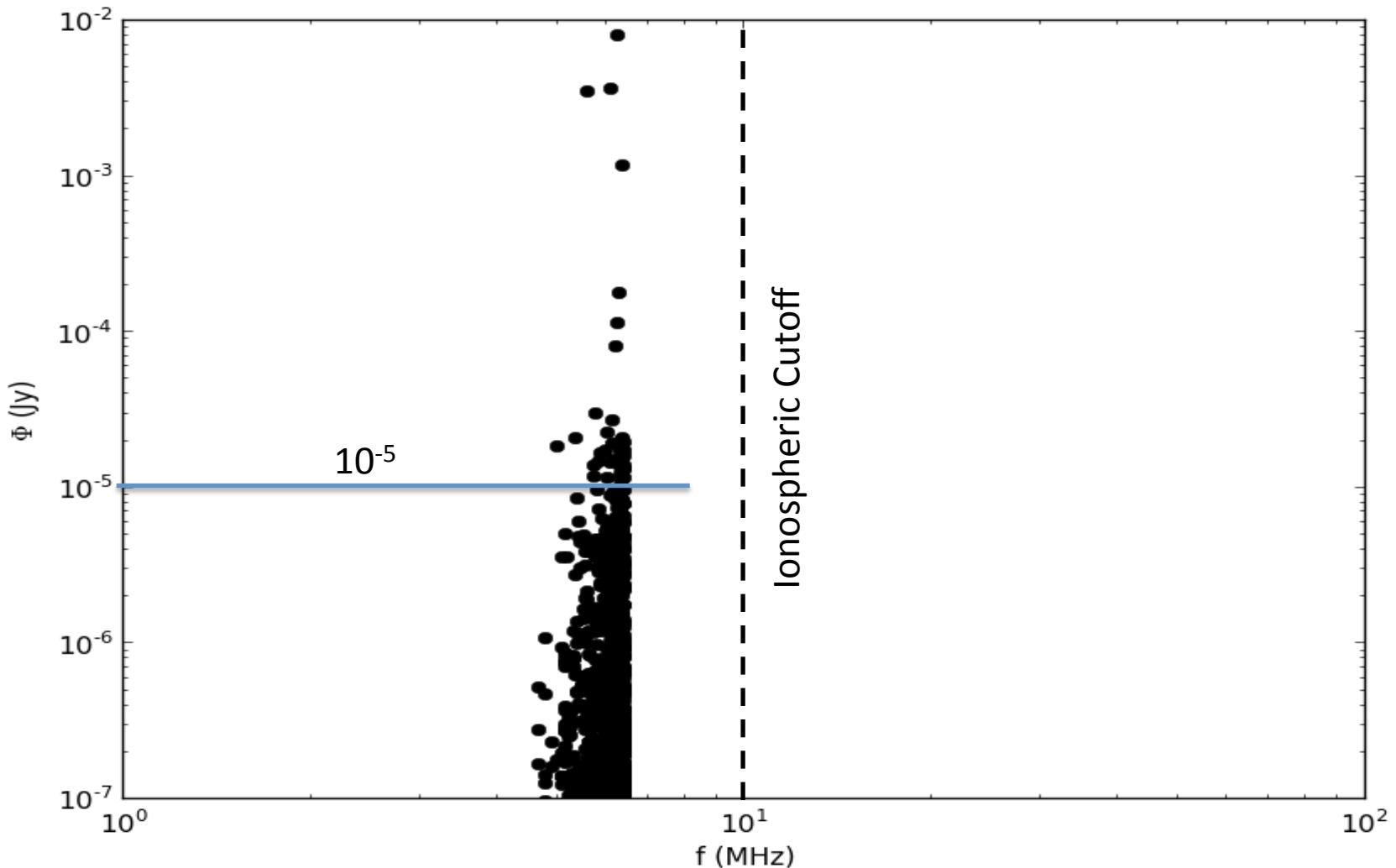
CME Model Results



CME Model Results



CME Model Results



Improving detection

1. Boosting signal strength
2. Better detection range

Improving detection

1. Boosting signal strength
2. **Better detection range**

Radio astronomy on the Moon!

LOLA

LOw frequency Lunar Array (LOLA)

LOLA vs. LOFAR:

- Same resolution requires a bigger array for lower signal frequencies
- Better sensitivity: lower noise on Moon, better technology

Cost of LOLA



\$15B total estimate

\$7.5B to transport antennas:

- Size: 3x3x1.7m
- Weight: 100lbs
- Amount: 20,000 antennas
- Falcon Heavy = 40,000lb payload
- 50 trips x \$150M/trip

\$7.5B estimated other overhead:

- Base station, deployment, construction, power, and maintenance

BONUS:

- Reuse hardware designs

Cost of LOLA



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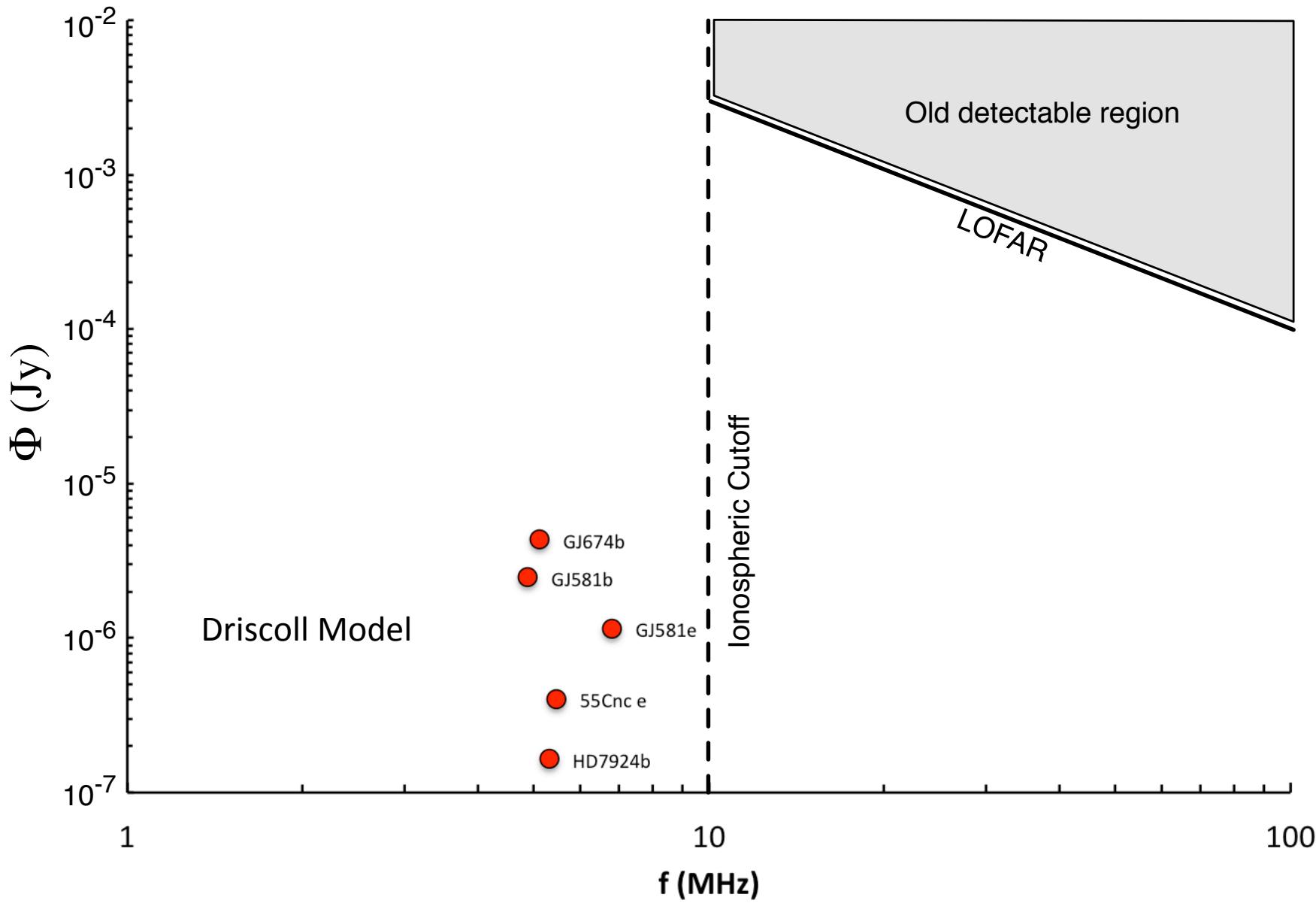
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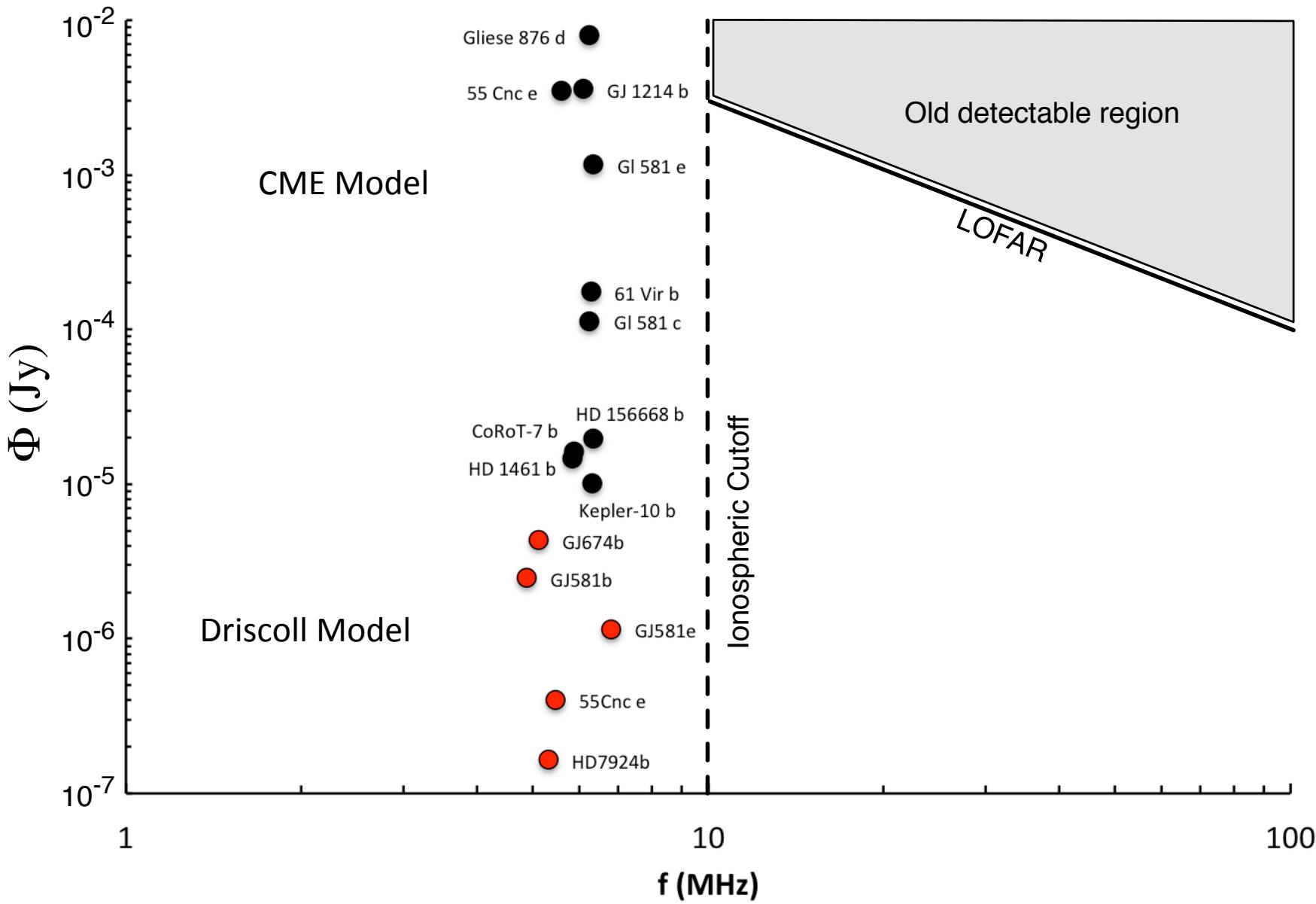
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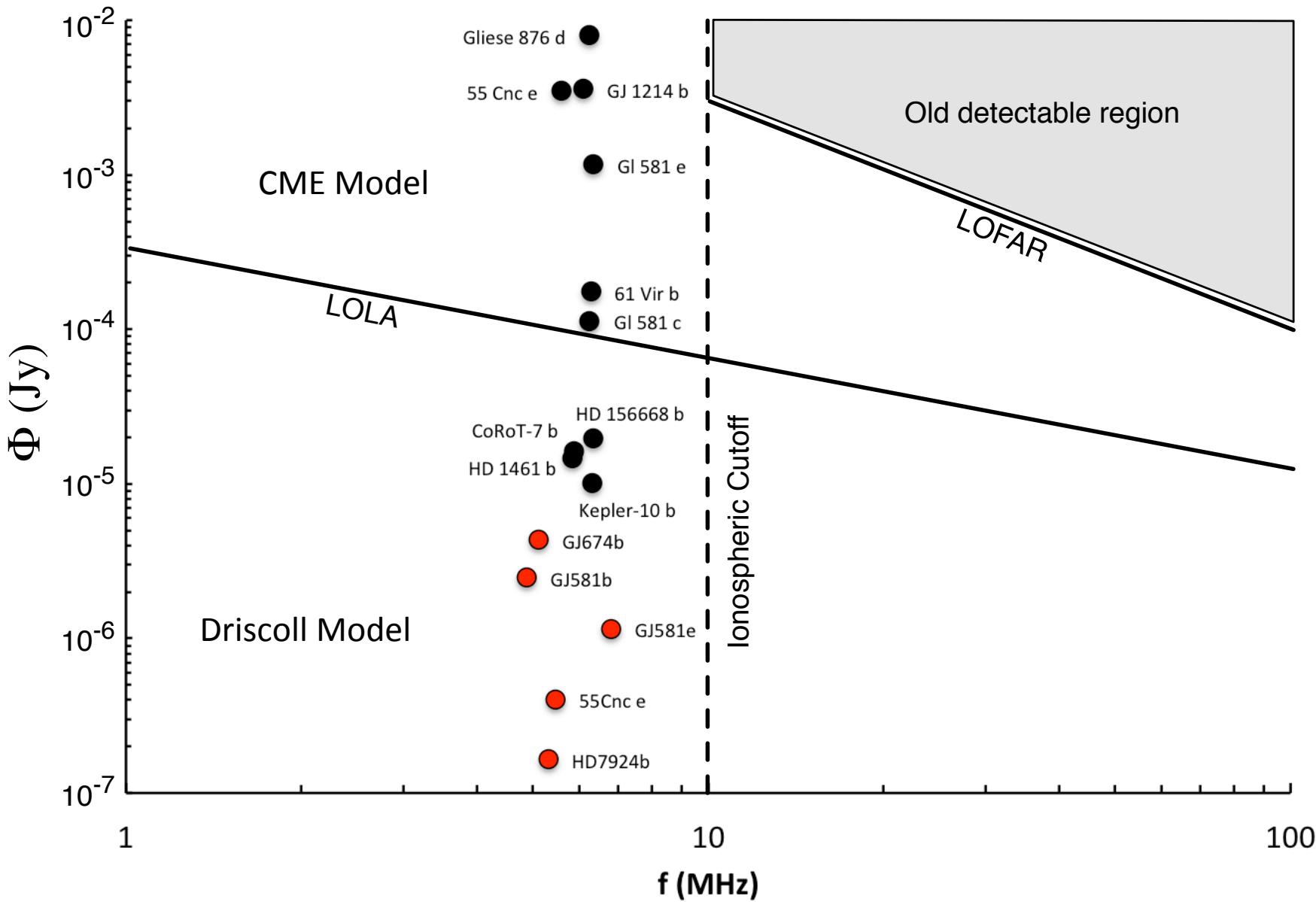
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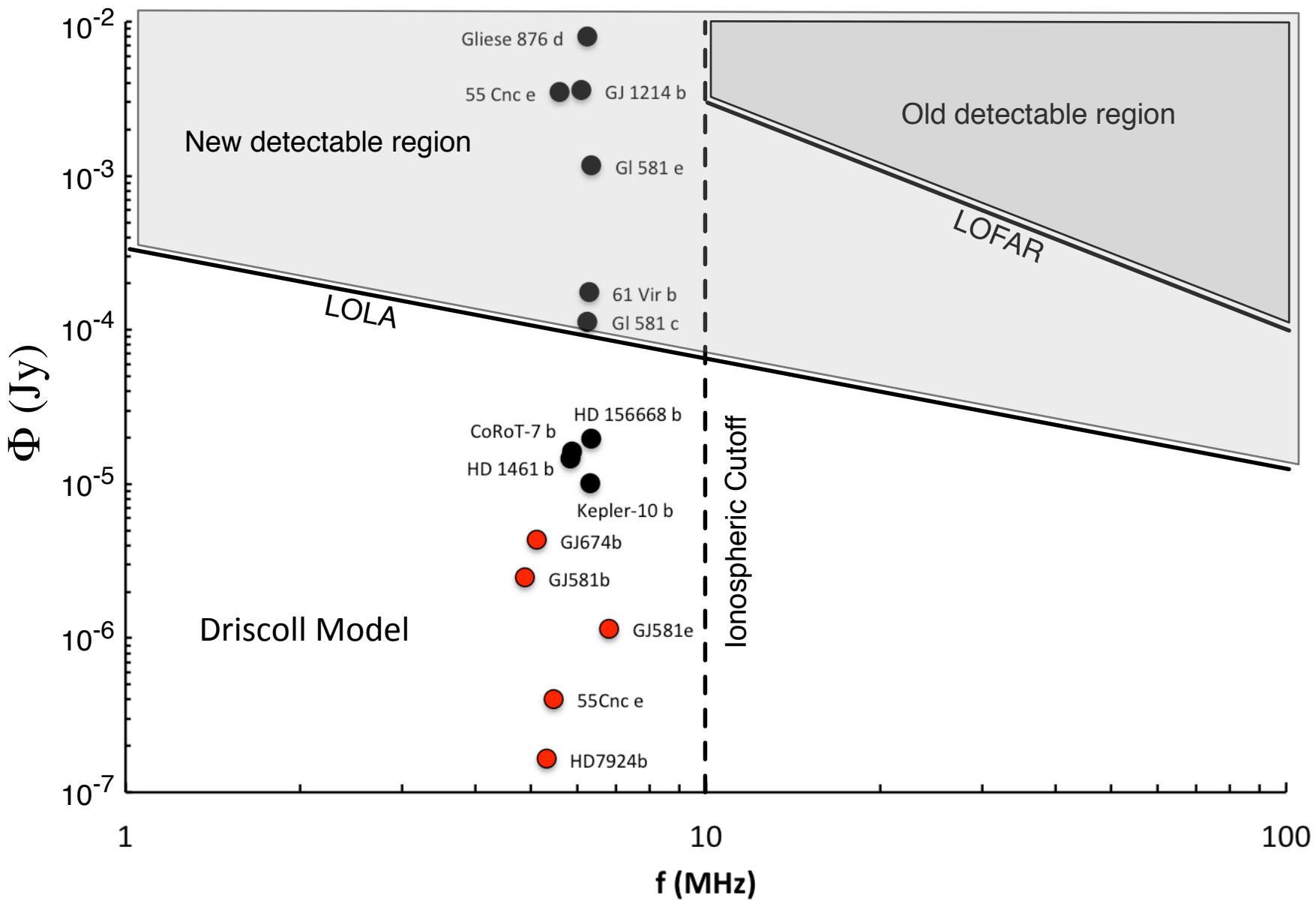
BONUS:

- Reuse hardware designs









Limitations

- Exoplanet Model
 - Optimistic dynamo
 - Stellar magnetic fields
- CMEs
 - Infrequent
 - Temporal
 - Localized
- LOLA
 - Expensive
 - Construction time

Conclusions

- Exoplanet magnetic field signals currently undetectable
 - too weak
 - too low frequency
- Timely coronal mass ejection events may boost signal strength significantly
- An expensive, low-frequency telescope on the Moon could detect these signals

Target exoplanets:

- Gliese876 d
- 55Cnc e
- GJ1214 b
- GI581 e
- 61Vir b
- GI581 c