

On the uniform, achromatic rotation of CMB polarization

Stefan Hamann^{a,1}, Alessandro Rizzo^{b,2,*}

^aIndependent Researcher, Ahaus, Germany

^bIndependent Researcher, Brescia, Italy

Abstract

Polarization angles are sharp probes of spacetime geometry. On a metric-compatible Riemann–Cartan background, a single variational structure yields a uniform, frequency-independent rotation β of CMB linear polarization without post hoc fitting. A Fujikawa computation fixes the abelian vertex $g_{a\gamma\gamma} = -1/(2\pi)$; geometric optics gives $d\psi/d\eta = \frac{1}{2} g_{a\gamma\gamma} da/d\eta$; a conformal trace stationarity at fixed connection, normalized in the torsionless GR limit, fixes $|\Delta a| = \varphi_0 = 0.0533 \pm 0.0007$. Fit-free with respect to CMB data: after anomaly normalization and the GR renormalization condition, no parameters are tuned to observations; the resulting prediction is $\beta_{\text{th}} = \varphi_0/(4\pi) = 0.243^\circ \pm 0.003^\circ$ (degrees). Planck PR4 and ACT DR6 are consistent with this level and show no ν^{-2} slope. Angles are quoted in our convention $\beta \equiv -\beta_{\text{exp}}$. Index-level steps, robustness to non-axial torsion and minimal deformations, and calibration-aware likelihood details are provided in the Supplementary material (online).

Keywords: cosmic birefringence, CMB polarization, trace anomaly, axial anomaly, Riemann–Cartan torsion, geometric optics

1. Setup and conventions

Heaviside–Lorentz units with $c = \hbar = 1$, signature $(+, -, -, -)$, and Levi–Civita symbol $\varepsilon_{0123} = +1$ are used; the dual is $\tilde{F}^{\mu\nu} = \frac{1}{2} \varepsilon^{\mu\nu\rho\sigma} F_{\rho\sigma}$. Hats denote torsionless curvature, e.g. $\hat{R}_{\mu\nu\rho\sigma}$; ∇ is the full Riemann–Cartan covariant derivative. Angles in equations are in radians. The polarization-angle convention matches TB/EB analyses via $\beta \equiv -\beta_{\text{exp}}$. The field a is a *dimensionless* axial phase (not the cosmological scale factor). To avoid dimensional ambiguities in the trace equation, the gravitational coupling is understood in the standard renormalized form $\bar{\kappa} := \mu \kappa_{\text{phys}}$; the bar is dropped below. Observables derived here are RG-stationary at one loop once the GR renormalization condition is imposed.

Result (no fit). In the setting above, the chiral Jacobian in Fujikawa regularization fixes $g_{a\gamma\gamma} = -1/(2\pi)$ with the adopted dual and interaction conventions. The modified Maxwell equations imply $d\psi/d\eta = \frac{1}{2} g_{a\gamma\gamma} da/d\eta$, so $\beta = \frac{1}{2} g_{a\gamma\gamma} \Delta a$ along the light path. The trace channel of the single action at fixed connection yields a monotone condition $I(\kappa, L) = \kappa^3 - 2c_3^2 \kappa^2 - 8b_1 c_3^6 L = 0$ with $L := \ln(1/\varphi_0)$, $\partial I/\partial L < 0$. Imposing in the torsionless limit the renormalization condition $I(\kappa_p, L^*) = 0$ at the GR point fixes L^* uniquely and therefore $\varphi_0 = e^{-L^*}$; no CMB data enter this fixation. With the physical orientation $\Delta a = -\varphi_0$, the prediction is $\beta_{\text{th}} = \varphi_0/(4\pi) = 0.243^\circ \pm 0.003^\circ$ (degrees).

2. Single action, field equations, and trace stationarity

A metric-compatible Riemann–Cartan geometry with torsion $T^\rho_{\mu\nu}$ and contorsion $K^\rho_{\mu\nu}$ is considered; axial torsion is $S_\mu = \frac{1}{6} \varepsilon_{\mu\nu\rho\sigma} T^{\nu\rho\sigma}$. The bulk action reads

$$S = \int d^4x \sqrt{-g} \left[\frac{1}{2\kappa^2} (\hat{R} - \nabla_\lambda K^\lambda_{\mu}{}^\mu + K_{\mu\nu\lambda} K^{\mu\nu\lambda}) - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{4} g_{a\gamma\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{\lambda}{2} \partial_\mu a S^\mu + \frac{1}{3} U(\kappa, L) \right] + S_{\partial\mathcal{M}}, \quad (1)$$

with $L = \ln(1/\varphi_0)$. Variation in S^μ gives

$$\partial_\mu a = \lambda S_\mu, \quad (2)$$

and for purely axial torsion

$$K_{\mu\nu\rho} = \frac{1}{2} \varepsilon_{\mu\nu\rho\sigma} S^\sigma = \frac{1}{2\lambda} \varepsilon_{\mu\nu\rho\sigma} \partial^\sigma a, \quad (3)$$

which makes the leading eikonal transport achromatic. Metric variation at fixed connection yields

$$Q_{\mu\nu} - \frac{1}{2} g_{\mu\nu} Q = \kappa^2 T_{\mu\nu}^{(a)} + \kappa^2 T_{\mu\nu}^{(\text{EM})} + \frac{1}{3} g_{\mu\nu} U, \quad (4)$$

with

$$Q_{\mu\nu} = \hat{R}_{\mu\nu} - \nabla_\lambda K^\lambda_{\mu\nu} + K_{\mu\alpha\beta} K_\nu{}^{\alpha\beta}, \quad Q = g^{\mu\nu} Q_{\mu\nu}. \quad (5)$$

A conformal variation $g_{\mu\nu} \rightarrow e^{2\sigma} g_{\mu\nu}$ at fixed connection isolates the trace channel. On the geometric-optics vacuum, matter enters only through the renormalization of U ; normalizing

$$\frac{\partial U}{\partial L} = \frac{8}{3} b_1 c_3^6. \quad (6)$$

where $c_3 = 1/(8\pi)$ by the abelian anomaly (Sec. 3).

*Corresponding author.

Email addresses: sh@sh-future.de (Stefan Hamann), a.rizzo@physiks.net (Alessandro Rizzo)

¹ORCID: 0000-0002-0983-2589.

²ORCID: 0000-0002-8030-3540.

Stationarity $\delta S/\delta\sigma = 0$ gives

$$I(\kappa, L) := \kappa^3 - 2c_3^2\kappa^2 - 8b_1c_3^6L = 0. \quad (7)$$

Since $\partial I/\partial L = -8b_1c_3^6 < 0$, for fixed κ there is a unique root L^* , hence $\varphi_0 = e^{-L^*} \in (0, 1)$.

$$\mathcal{V}_{\mu\nu} := Q_{\mu\nu}^{\text{TF}} + g_{\mu\nu} I(\kappa, L). \quad (8)$$

The full variation is equivalent to $\mathcal{V}_{\mu\nu} = 0$.

Fixed-connection conformal stationarity. “Metric variation at fixed connection” means that in the conformal variation $g_{\mu\nu} \rightarrow e^{2\sigma} g_{\mu\nu}$ the affine structure Γ is held fixed (Palatini logic with metric-compatibility imposed), so the trace channel is isolated without altering the transport sector. This is a renormalization prescription, not a gauge choice; the Einstein–Hilbert boundary term ensures a well-posed variational principle for Dirichlet (g, Γ) data.

3. Anomaly normalization and achromatic transport

A local chiral rotation $\psi \rightarrow e^{\frac{i}{2}a(x)\gamma_5}\psi$ produces the Fujikawa Jacobian

$$\Delta\mathcal{L}_{\text{anom}} = +\frac{a(x)}{8\pi}\partial_\mu K^\mu, \quad K^\mu = \varepsilon^{\mu\nu\rho\sigma}A_\nu\partial_\rho A_\sigma, \quad (9)$$

and since $F_{\mu\nu}\tilde{F}^{\mu\nu} = 2\partial_\mu K^\mu$, matching to $-\frac{1}{4}g_{a\gamma\gamma}aF_{\mu\nu}\tilde{F}^{\mu\nu}$ fixes

$$g_{a\gamma\gamma} = -\frac{1}{2\pi}. \quad (10)$$

Equivalently, matching $-\frac{1}{4}g_{a\gamma\gamma}aF\tilde{F} = +c_3aF\tilde{F}$ gives $g_{a\gamma\gamma} = -4c_3$ and therefore $c_3 = 1/(8\pi)$ in our conventions. The modified Maxwell equations read

$$\nabla_\mu F^{\mu\nu} + \frac{1}{2}g_{a\gamma\gamma}\partial_\mu a\tilde{F}^{\mu\nu} = 0, \quad (11)$$

and in geometric optics

$$\frac{d\psi}{d\eta} = \frac{1}{2}g_{a\gamma\gamma}\frac{da}{d\eta} \Rightarrow \beta = \frac{1}{2}g_{a\gamma\gamma}\Delta a. \quad (12)$$

Sign check. With $\varepsilon_{0123} = +1$ and the above dual, Eq. (9) fixes the overall sign in (10). Since $d\psi/d\eta = \frac{1}{2}g_{a\gamma\gamma}da/d\eta$ and $\Delta a = -\varphi_0 < 0$, one has $\beta_{\text{th}} = \varphi_0/(4\pi) > 0$, identical in sign to the angles reported by experiments after the $\beta \equiv -\beta_{\text{exp}}$ conversion.

4. Trace anomaly and fixation of b_1

On metric-compatible Riemann–Cartan backgrounds with minimal couplings and purely axial torsion, the Maxwell operator is unchanged at one loop. The dilatation anomaly takes the form

$$\langle T^\mu{}_\mu \rangle = \frac{b_1}{16\pi^2} F_{\mu\nu}F^{\mu\nu} + \dots, \quad (13)$$

and b_1 equals the abelian trace coefficient of the Standard Model, $b_1 = 41/10$ in GUT normalization. Axial torsion redistributes purely geometric terms in the DeWitt–Seeley coefficient a_2 , while the $F_{\mu\nu}F^{\mu\nu}$ coefficient retains its torsionless value at one loop under minimal couplings; non-axial torsion and non-minimal operators do not mix into F^2 at this order [6, 7].

EWSB mapping (explicit). With $g_1^2 = \frac{5}{3}g'^2$ and $A_\mu = \sin\theta_W B_\mu + \cos\theta_W W_\mu^3$, kinetic diagonalization removes $F_{\mu\nu}Z^{\mu\nu}$ at one loop under minimal couplings; the abelian trace piece projects onto the canonically normalized photon operator $F_{\mu\nu}F^{\mu\nu}$ with unit weight. Hence the coefficient in front of F^2 remains $b_1 = 41/10$ after EWSB at this order.

5. Renormalization condition at the GR point and prediction

The torsionless limit must reproduce the Einstein–Hilbert sector $(2\kappa_{\text{P}}^2)^{-1}R$ at the chosen renormalization scale; this is a renormalization condition, not a fit. Imposing

$$I(\kappa_{\text{P}}, L^*) = 0 \quad (14)$$

fixes L^* uniquely because $\partial I/\partial L = -8b_1c_3^6 < 0$, hence $\varphi_0 = e^{-L^*}$ is determined without reference to CMB data. Using $c_3 = 1/(8\pi)$ from (10) and the Standard Model value $b_1 = 41/10$, one finds numerically

$$\varphi_0 = 0.0533 \pm 0.0007, \quad \beta_{\text{th}} = \frac{\varphi_0}{4\pi} = 0.243^\circ \pm 0.003^\circ.$$

The quoted uncertainty is the direct propagation of $\delta\varphi_0$ through $\beta_{\text{th}} = \varphi_0/(4\pi)$ (angles converted to degrees); scheme shifts in U cancel against the GR renormalization condition, and higher-order geometric corrections are $\ll 10^{-2}$ deg across CMB bands and are not included.

RG stationarity. Differentiating $I(\kappa_{\text{P}}, L^*) = 0$ with respect to $\ln\mu$ shows that the change $U \rightarrow U + \frac{8}{3}b_1c_3^6\delta\ell$ under $\mu \rightarrow \mu e^{\delta\ell}$ shifts I by a constant that is removed by (14), so $dL^*/d\ln\mu = 0$ and $d\beta_{\text{th}}/d\ln\mu = 0$ at one loop.

6. Consistency with current measurements

All angles below are quoted in our convention $\beta \equiv -\beta_{\text{exp}}$. Planck PR4 reports $\beta = 0.300^\circ \pm 0.110^\circ$ (68% C.L.) and a dedicated ACT DR6 analysis with absolute-angle priors finds $\beta = 0.215^\circ \pm 0.074^\circ$ (68% C.L.), both compatible with our prediction $\beta_{\text{th}} = \varphi_0/(4\pi)$. Because Planck and ACT rely on independent absolute-angle calibrations with different priors and systematics, a common positive β in our sign convention cannot be produced by miscalibration alone; the joint ν^{-2} slope fit consistent with zero further disfavors Faraday-like leakage. A joint multi-band fit of a Faraday-like slope $\beta(\nu) = \beta_0 + m(\nu/\text{GHz})^{-2}$ yields m consistent with zero, as expected for achromatic transport.

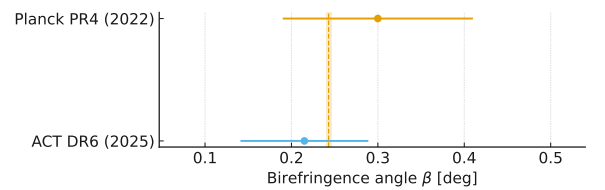


Figure 1: Uniform birefringence angle. Horizontal band: $\beta_{\text{th}} \pm \sigma_{\text{th}}$. Points: Planck PR4 and ACT DR6 (68% C.L.), in our convention $\beta = -\beta_{\text{exp}}$.

Table 1: Prediction vs. measurements (degrees; 68% C.L.).

Entry	Value	Note
Prediction β_{th}	0.243 ± 0.003	from $\varphi_0/(4\pi)$
Planck PR4	0.300 ± 0.110	nearly full-sky
ACT DR6	0.215 ± 0.074	absolute-angle priors

7. Robustness

Non-axial torsion modifies $K^\rho_{\mu\nu}$ but leaves $g_{\alpha\gamma}$ and the F^2 anomaly coefficient unchanged in the minimal theory [6, 7]. The uniform rotation depends on $\partial_\mu a$ via (2) and is achromatic at leading eikonal order [1, 8]. Small non-minimal operators organize as higher-derivative corrections to the constitutive tensor and would induce suppressed frequency dependence; the joint multi-band slope consistent with zero constrains such effects at the $\lesssim 10^{-2}^\circ$ level across CMB bands. Being a total derivative, the Nieh–Yan density renormalizes L -independent geometric counterterms and does not renormalize the $aF\tilde{F}$ vertex or the F^2 coefficient in the minimal, metric-compatible setup.

8. Falsifiability

A single-experiment 5σ rejection of β_{th} against $\beta = 0$ requires a total polarization-angle uncertainty $\sigma_{\text{tot}} \lesssim 0.049^\circ$; for 3σ , $\sigma_{\text{tot}} \lesssim 0.081^\circ$. These thresholds set the target for absolute polarization-angle calibration; once reached, the prediction $\beta_{\text{th}} = \varphi_0/(4\pi)$ is immediately testable.

9. Summary

A single variational structure on metric-compatible Riemann–Cartan backgrounds predicts a uniform, achromatic rotation of CMB polarization with no tunable parameters. The anomaly fixes $g_{\alpha\gamma}$, geometric optics fixes the transport law, and the trace equation with the GR renormalization condition fixes $|\Delta a| = \varphi_0$, yielding $\beta_{\text{th}} = \varphi_0/(4\pi) = 0.243^\circ \pm 0.003^\circ$. Present data are consistent with a constant rotation at this level and show no ν^{-2} slope. Algebraic steps, RG details, and calibration-aware likelihoods are provided in the Supplementary material (online).

References

- [1] F. W. Hehl, Y. N. Obukhov, *Foundations of Classical Electrodynamics*, Birkhäuser (2003).
- [2] W.-T. Ni, A nonmetric theory of gravity, *Phys. Rev. D* **7**, 2880 (1973).
- [3] S. M. Carroll, G. B. Field, R. Jackiw, Limits on a Lorentz- and parity-violating modification of electrodynamics, *Phys. Rev. D* **41**, 1231 (1990).
- [4] A. Lue, L. Wang, M. Kamionkowski, Cosmological signature of new parity-violating interactions, *Phys. Rev. Lett.* **83**, 1506 (1999).
- [5] Y. Minami, E. Komatsu, Simultaneous determination of cosmic birefringence and miscalibrated polarization angles, *Phys. Rev. Lett.* **125**, 221301 (2020).
- [6] I. L. Shapiro, Physical aspects of the space-time torsion, *Phys. Rept.* **357**, 113–213 (2002).
- [7] I. L. Buchbinder, S. D. Odintsov, I. L. Shapiro, *Effective Action in Quantum Gravity*, IOP Publishing (1992), Ch. 10.
- [8] Y. N. Obukhov, G. F. Rubilar, Fresnel analysis of wave propagation in media with general linear constitutive law, *Phys. Rev. D* **66**, 024042 (2002).
- [9] P. Diego-Palazuelos *et al.*, Cosmic Birefringence from Planck PR4, *Phys. Rev. Lett.* **128**, 091302 (2022).
- [10] P. Diego-Palazuelos *et al.*, Cosmic Birefringence from Planck Public Release 4, arXiv:2203.04830 [astro-ph.CO] (2022).
- [11] P. Diego-Palazuelos, E. Komatsu, Cosmic Birefringence from the Atacama Cosmology Telescope Data Release 6, arXiv:2509.13654 [astro-ph.CO] (2025).
- [12] J. Aumont *et al.*, Absolute calibration of the polarisation angle for future CMB B-mode experiments, *Astron. Astrophys.* **634**, A100 (2020). doi:10.1051/0004-6361/201935785