CITA computing project

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The goal of this project is to simulate two double white dwarf populations with COSMIC¹, a binary population synthesis suite and compare their output.

1 Getting familiar with COSMIC

- 1. Install COSMIC by cloning from the Github repository
- 2. Simulate 3 binaries that produce a WD + WD, NS + NS, and BH + BH system and plot the time evolution of their masses, orbital period, and eccentricity from Zero Age Main Sequence to the formation of the second remnant (WD/NS/BH). This will take some trial and error with the initial conditions.
- 3. Simulate a grid of binaries with an initial orbital period of 50 days, initial eccentricity of 0.5, primary masses ranging from 10 to $100 \,\mathrm{M}_\odot$ in $10 \,\mathrm{M}_\odot$ increments, and 10 secondary masses for each primary mass with a mass ratio ranging from 0.1 to 1.0. Plot the orbital period at the formation of the second stellar remnant as a function of mass ratio for each primary mass.

2 Simulating double white dwarf populations

- 1. Initialize 1,000,000 binaries with the independent initial condition sampler using the following models: primary masses following kroupa01, eccentricity following a uniform distribution, orbital periods following sana12, a burst of star formation 13.7 Gyr in the past with a duration 0.0 yr, all with solar metallicity: Z=0.017.
- 2. Evolve all of these with the standard BSEDict from the COSMIC documentation.
- 3. Select all WD + WD binaries at the formation of the second WD and plot a histogram of their ages, masses, and orbital periods.
- 4. Write two functions based on the equations in Peters (1964) for circular binaries: one to evolve the WD + WD binary semimajor axis as a function of time due to GW emission; and one which calculates the merger time.
- 5. Calculate the time between the ZAMS formation of the WD + WD binaries and their merger times (also called a delay time), and plot a histogram of the delay time distribution for the population.
- 6. Repeat the process above but change the alpha parameter in the BSEDict from 1 to 0.25.
- 7. Compare the delay time distributions and WD + WD formation distributions to describe how this change affects the population

¹https://github.com/COSMIC-PopSynth/COSMIC