



FACULTY
OF MATHEMATICS
AND PHYSICS
Charles University

MASTER THESIS

Jan Střeleček

Some funny small things looking for
counterdiabatic elements

Supervisor of the master thesis: prof. RNDr. Pavel Cejnar, Dr.,
DSc.

Study programme: Theoretical physics

Study branch: hmmmmm

Prague 2022

I declare that I carried out this master thesis independently, and only with the cited sources, literature and other professional sources. It has not been used to obtain another or the same degree.

I understand that my work relates to the rights and obligations under the Act No. 121/2000 Sb., the Copyright Act, as amended, in particular the fact that the Charles University has the right to conclude a license agreement on the use of this work as a school work pursuant to Section 60 subsection 1 of the Copyright Act.

In date
Author's signature

Dedication.

Title: Some funny small things looking for counterdiabatic elements

Author: Jan Střeleček

Institute:

Supervisor:

Abstract: Abstract.

Keywords: key words

Contents

Some notes to the notation	3
Introduction	4
1 Mathematical introduction	5
1.1 Fiber bundle	6
1.2 Vector Bundle	6
1.2.1 Connection on vector bundles	7
1.2.2 Metric on vector bundles	7
1.3 Section	7
1.4 Pull-back and push forward	7
1.5 Flow	8
1.6 Covariant derivative and parallel transport	8
1.7 Parallel transport on vector bundles	8
1.8 Antisymmetric tensors and wedge product	8
1.9 Useful definitions	9
1.10 Useful theorems	9
2 Physical introduction	11
2.1 Transporting states	13
2.2 Metric and geometric tensor	15
2.3 Derivation of the geometric tensor	17
2.4 Hamiltonian induced flow on manifold	17
2.5 Adiabaticity	18
2.5.1 Slow transports	18
2.6 Geodetic driving	19
2.7 Gauge potentials	19
2.7.1 Classical gauge potential	20
2.7.2 Quantum gauge potential	20
2.7.3 Adiabatic gauge potential	21
2.8 Counter-diabatic driving	22
2.8.1 Explicit form	23
2.9 Berry phase and curvature	23
2.10 Approximations of adiabatic potentials	24
2.10.1 Variational methods	24
3 Geodesics	25
4 Spin 1/2	26
4.1 Special case	27
5 First order transition Hamiltonian	29
5.1 Hamiltonian analysis	29
5.1.1 Case N=3	30
5.2 Arbitrary N	34
5.3 Limit $N \rightarrow \infty$	34

Conclusion	39
Bibliography	40
List of Figures	41
A Appendix 1	42
B Attachments	43
B.1 First Attachment	43

Some notes to the notation

Symbol	Meaning	Defining formula
\mathcal{A}	Gauge (calibrational) potential	$\mathcal{A}_\mu = i\hbar\partial_\mu$

Mathematical spaces will be denoted *mathcal* and operators with *hat*.

Introduction

1. Mathematical introduction

The modern approach to the closed system dynamics is using differential geometry formalism. Before we get to the quantum mechanics itself, lets define the formalism and recapitulate some definitions of this branch of mathematics. More detailed notes can be found for example in [Fecko, 2006].

Let's have a manifold \mathcal{M} and curves

$$\gamma : \mathbb{R} \xrightarrow{\text{open}} I \rightarrow \mathcal{M} \quad \xi \mapsto \gamma(\xi).$$

The space of functions is $\mathcal{F}(\mathcal{M}) \equiv \{f : \mathcal{M} \rightarrow \mathbb{R}\}$, where

$$f : \mathcal{M} \rightarrow U \xrightarrow{\text{open}} \mathbb{R} \quad x \mapsto f(x).$$

To define *vectors* on \mathcal{M} , we need to make sense of the *direction*. It is defined using curves satisfying

$$\begin{aligned} \gamma_1(0) &= \gamma_2(0) \equiv P \\ \frac{d}{dt} x^i(\gamma_1(t)) \Big|_{t=0} &= \frac{d}{dt} x^i(\gamma_2(t)) \Big|_{t=0}. \end{aligned}$$

Taking the equivalence class created by those two rules, sometimes noted as $[\gamma] = v$, we have element of the tangent space to \mathcal{M} . We will use standard notation for the tangent space to \mathcal{M} in some point xP as $\mathbb{T}_P\mathcal{M}$ and contangent space as $\mathbb{T}_P^*\mathcal{M}$. Unifying all those spaces over all x we get tangent and cotangent bundle, $\mathcal{T}\mathcal{M}$ and $\mathcal{T}^*\mathcal{M}$ respective. To generalize this notation to higher tensors, we denote $\mathbb{T}_P\mathcal{M} \in \mathcal{T}^1\mathcal{M}$, $\mathbb{T}_P^*\mathcal{M} \in \mathcal{T}_1\mathcal{M}$, thus the space of p -times contravariant and q -times covariant tensors is denoted $\mathcal{T}_q^p\mathcal{M}$.

Using the congruence of the curves on \mathcal{M} , the expression

$$\frac{d}{d\xi} f \circ \gamma(\xi) \Big|_{\xi=0} \tag{1.1}$$

has a good meaning and we can define the *derivative* in some $P \in \mathcal{M}$ as

$$\mathbf{v} : \mathcal{F}(\mathcal{M}) \rightarrow \mathbb{R} \quad f \mapsto \mathbf{v}[f] \equiv \frac{df(\gamma(\xi))}{d\xi} \Big|_P \equiv \partial_\xi \Big|_P f. \tag{1.2}$$

It holds, that $\mathbf{v} \in \mathbb{T}_P\mathcal{M}$ and can be expressed as the *derivative in direction*,¹ which can be understood in coordinates as

$$\mathbf{v}[f] = \frac{d}{d\mathbf{v}} f \circ \gamma(\xi) \Big|_{\xi=0} = v^\mu \frac{d}{dx^\mu} f(x) \Big|_P. \tag{1.4}$$

The directionnal derivative will be denoted

$$\nabla_v$$

¹ The direction itself is usually denoted as

$$\frac{D}{d\alpha} \gamma(\xi), \tag{1.3}$$

where the big D notation is used to point out that it's not a classical derivative, but it maps curves to some entirely new space of directions.

and in basis $\mathbf{e}_i \equiv \partial/\partial x^i$ we will denote

$$\nabla = (\mathbf{e}_x, \mathbf{e}_y, \mathbf{e}_z).$$

To get some physical application, we need to define one strong structure on manifolds – differentiable metric tensor $g_{\mu\nu} \in \mathcal{T}_2^0 \mathcal{M}$ – so the covariant derivatives and parallel transport are well-defined everywhere.

1.1 Fiber bundle

We already know tangent and cotangent spaces. Thinking of manifold \mathcal{M} as a fiber, for which at any point x we create new manifold $\mathbb{T}_x \mathcal{M}$, we and unifying all those tangent spaces, we get so-called *vector bundle*.

Definition 1. *Structure $(\mathcal{E}, \mathcal{B}, \pi, \mathcal{F})$, for topological spaces \mathcal{E} (total space), \mathcal{B} (base space), \mathcal{F} (fibre) and a continuous surjection $\pi : \mathcal{E} \rightarrow \mathcal{B}$ satisfying a local triviality² is called a Fiber bundle (projection map). In addition, the \mathcal{B} is assumed to be connected³ and for every $x \in \mathcal{B}$, there is an open neighborhood $\mathcal{U} \subset \mathcal{B}$ (trivializing neighborhood) such that there exists a homeomorphism from \mathcal{U} to so-called product space*

$$\phi : \pi^{-1}(\mathcal{U}) \rightarrow \mathcal{U} \times \mathcal{F},$$

such that $\pi^{-1}\pi(\mathcal{U}) = \mathcal{U}$.

The above mappings can be more clear from figure 1.1. Because projections of products are open maps, $\pi : \mathcal{E} \rightarrow \mathcal{B}$ must be open map. The meaning of

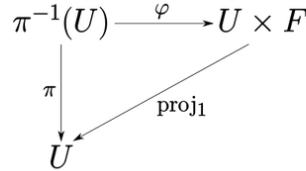


Figure 1.1: Mappings needed for the bundle definition.

definition is, that manifolds at every point $x \in \mathcal{F}$ are all locally diffeomorphic to each other.

1.2 Vector Bundle

Conversely, given a fiber bundle $(\mathcal{E}, \mathcal{X}, \pi, \mathcal{Rk})$ with a $GL(k)$ cocycle acting in the standard way on the fiber \mathcal{Rk} , there is associated a vector bundle. This is sometimes taken as the definition of a vector bundle

²local triviality

³can't be represented as a union of two disjoint sets

1.2.1 Connection on vector bundles

[Loring, 2017][chap. 10.1] Connection maps vector from tangent space to base manifold \mathcal{X} with some element from total space \mathcal{E} to total space

$$\Gamma : \mathcal{X} \times \mathcal{E} \rightarrow \mathcal{E}$$

such, that

- $\Gamma_X s$ is \mathcal{F} -linear in X and \mathbb{R} -linear in s
- Leibniz rule for $f \in \mathcal{C}^\infty$ is satisfied: $\Gamma_X(fs) = (Xf)s + f\Gamma_X s$

1.2.2 Metric on vector bundles

Map

$$g_{\mu\nu} : \mathcal{B} \times \mathcal{B} \rightarrow \mathbb{R}$$

1.3 Section

Section is a function

$$f : \mathcal{B} \rightarrow \mathcal{F},$$

such that $\pi(f(x)) = x$ for $\forall x \in \mathcal{B}$. This defines new manifold cutting throw \mathcal{E} .

Sectioning of fiber bundles creates vector spaces

1.4 Pull-back and push forward

Push-forward and pull-back are used to transport vectors and covectors between manifolds. Let's have two manifolds \mathcal{M}, \mathcal{N} , a smooth mapping ϕ and functions f, \tilde{f} such that

$$\begin{aligned} \phi : \mathcal{M} &\rightarrow \mathcal{N} & x &\mapsto \phi x \\ \tilde{f} : \mathcal{N} &\rightarrow \mathbb{R} & & \end{aligned}$$

Pull-back of the function then defines a new function $f : \mathcal{M} \rightarrow \mathbb{R}$ as

$$\phi^* : \mathcal{F}\mathcal{N} \rightarrow \mathcal{F}\mathcal{M} \quad \tilde{f} \mapsto f = (\phi^*\tilde{f})(x) \equiv \phi^*\tilde{f}(x) = \tilde{f}(\phi x).$$

Push-forward of a vector is defined as

$$\phi_* : \mathbb{T}_x \mathcal{M} \rightarrow \mathbb{T}_{\phi x} \mathcal{N} \quad \phi_* \frac{D\gamma(\xi)}{d\xi} \Big|_x = \frac{D\phi\gamma(\xi)}{d\xi} \Big|_x$$

and *pull-back of a covector* $\tilde{\alpha} \in \mathbb{T}_{\phi x} \mathcal{N}$ is

$$\phi^* : \mathbb{T}_{\phi x} \mathcal{N} \rightarrow \mathbb{T}_x \mathcal{M} \quad (\phi^*\tilde{\alpha})_\mu v^\mu \Big|_x = \tilde{\alpha}_\mu (\phi_* v)^\mu \Big|_{\phi x}.$$

If ϕ has a smooth inversion, i.e. it is a diffeomorphism, we can define pull-back of vectors as

$$\phi^* = \phi_*^{-1} \tag{1.5}$$

and push-forward of covectors

$$\phi_* = (\phi^{-1})^* \tag{1.6}$$

1.5 Flow

1.6 Covariant derivative and parallel transport

this section is probably not needed

Covariant derivative is generally... Metric covariant derivative is...

Parallel transport of vector $v \in \mathbb{T}_p\mathcal{M}$ will be denoted $\text{par}_\gamma v \in \mathbb{T}$

Affine connection can be expressed as

$$\Gamma_{\mu\nu}^\alpha = \frac{1}{2}g^{\alpha\beta}(g_{\beta\mu,\nu} + g_{\nu\beta,\mu} - g_{\mu\nu,\beta}), \quad (1.7)$$

where we used comma notation for the coordinate derivative. The covariant derivative of $a \in \mathbb{T}_P\mathcal{M}$ is then defined

$$\frac{D a^\mu}{dx^\nu} = a^\mu_{,\nu} - \Gamma_{\alpha\beta}^\mu x^\alpha a^\beta \quad (1.8)$$

and for $\alpha \in \mathbb{T}_P^*\mathcal{M}$ it is

$$\frac{D \alpha_\mu}{dx^\nu} = \alpha_{\mu,\nu} - \Gamma_{\mu\beta}^\alpha x^\beta \alpha_\alpha \quad (1.9)$$

The vector $v \in \mathbb{T}_P\mathcal{M}$ is said to be parallel transported along curve $\gamma(\lambda)$, if its covariant derivative

$$\frac{D v^\mu}{d\xi} = 0 \quad (1.10)$$

vanishes along γ .

1.7 Parallel transport on vector bundles

this is what we need Parallel transport of vector V along curve γ will be denoted

$$\text{par}_\gamma V.$$

It is

1.8 Antisymmetric tensors and wedge product

p -form $A \in \mathcal{T}_p\mathcal{M}$ is called *antisymmetric*, if changing the order of the indices has impact only on the sign, symbolically

$$A_{i_1 \dots i_p} = \text{sign}(\sigma) A_{i_{\sigma_1} \dots i_{\sigma_p}},$$

where σ is some permutation. *Antisymmetrisation* is defined as a normalized sum over all permutation

$$A^{[i_1 \dots i_p]} \equiv \frac{1}{p!} \sum_{\sigma} A^{[i_{\sigma_1} \dots i_{\sigma_p}]} \quad (1.11)$$

The *wedge product* of $A \in \mathcal{T}_p\mathcal{M}$ and $B \in \mathcal{T}_q\mathcal{M}$ is antisymmetrisation of the tensor product in the sense

$$A \wedge B \equiv \frac{(p+q)!}{p!q!} A^{[i_1 \dots i_p]} \otimes B^{i_1 \dots i_q} \quad (1.12)$$

1.9 Useful definitions

Definition 2 (Riemannian manifold). *Manifold is called Riemannian, iff its equipped with positive definite metric tensor.*

Definition 3 (Connected manifold). *Manifold is connected, iff the distance between two points is infimum of the lengths of curves joining the two points.*

Definition 4 (Compact manifold). *Manifold is said to be compact if its every open cover has a finite subcover.*

Definition 5 (Geodesical completeness). *A manifold is said to be geodesically complete if its every geodesic can be extended to infinite values of their affine parameter.*

This condition holds if the space does not contain any singularities and its coordinate-independent notion.

Definition 6 (Geodesic maximality). *A manifold is said to be geodesically maximal if its eigter geodesically complete, or every non-complete geodesic (such that cannot be extended to infinite values of their affine parameter) ends in a singularity.*

Geodesic maximality is coordinate dependent notion, if the manifold is geodesically complete.

1.10 Useful theorems

Theorem 1 (Von Neumann-Wigner). (*Von Neumann J, Wigner E. 1929. Physikalische Zeitschrift 30:467-470*) This sometimes called the Non-Crossing Theorem states, that the eigenvalues of Hermitian matrix driven by N continuous real parameters forms at maximum $N - 2$ dimensional submanifold.

Theorem 2 (Hopf-Rinow Theorem). *For connected Riemannian manifold \mathcal{M} with the metric g , following are equivalent:*

- (\mathcal{M}, g) is geodesically complete, i.e. all geodesics are infinite
- (\mathcal{M}, g) is geodesically complete at some point P , i.e. geodesics going through P are infinite
- (\mathcal{M}, g) satisfies the Heine-Borel property, i.e. every closed bounded set is compact
- (\mathcal{M}, g) is complete as a metric space.

See Petersen [1998]/[p.125].

Theorem 3 (Modified Hopf-Rinow Theorem). *For connected Riemannian manifold \mathcal{M} with the metric g , any two points of \mathcal{M} can be joined with a minimizing geodesic. See Gorodski [2012]/[Chapter 3].*

This generally means, that in a space with singularity exists such points, which cannot be connected with the rest of the manifold using geodesics. Term "event horizon" will be used for such segments of manifold, as is common in General relativity.

Theorem 4. *A compact Riemannian manifold is geodesically complete. See Gorodski [2012]/[Chapter 3].*

2. Physical introduction

Most parts of this chapter are inspired by [Kolodrubetz et al., 2017] and original notes [Berry, 1984], [Berry, 1989], [Berry, 2009]. Now we will assign some physical background to the structure defined in the first chapter. We will see, that the whole space structure is quite complicated, because it's of fiber structure, where every fiber is another fiber bundle. Luckily what we will be using later on are only some much easier Riemannian submanifolds.

Assume parameter $\lambda \in \mathcal{U} \subset \mathbb{R}^n$ controlling some Hamiltonian $\hat{H}(\lambda)$, which is bounded from below and the spectrum is discrete for the first $k > 1$ energies, will be clear later on. From this we can construct fiber bundle, such that at every point of base manifold $\lambda \in \mathcal{U}$, we construct fiber spanning all possible states of $\hat{H}(\lambda)$, thus the fiber structure can be according to section 1.1 written as

$$\left(\mathcal{H}_{full} := \bigcup_{\lambda} \mathcal{H}(\lambda), \mathcal{U} \subset \mathbb{R}^n, \pi, \mathcal{H}(\lambda) := \bigcup_{states} |\psi(\lambda)\rangle \right).$$

The projection is defined as $\pi(\lambda) : |\psi(\lambda)\rangle \mapsto \lambda$ and $\mathcal{H}(\lambda)$ is Hilbert space for all pure states of $\hat{H}(\lambda)$. Geometric intuition is displayed in fig. 2.1

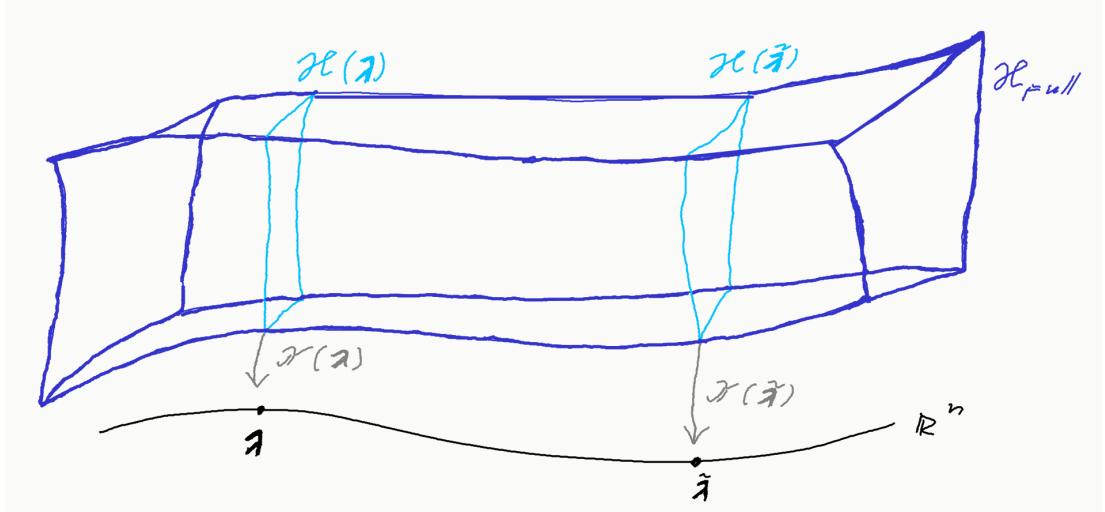


Figure 2.1: Fiber bundle over \mathbb{R}^n with Hilbert spaces $\mathcal{H}(\lambda)$ as individual fibers.

Because we are interested only in discrete part of spectrum¹, it will further on be referred to only as *spectrum*.

The states of the system evolve according to the Schrödinger equation

$$i\hbar d_t |\psi(\lambda(t))\rangle = \hat{H}(\lambda) |\psi(\lambda(t))\rangle, \quad (2.1)$$

which for eigenstates of instantaneous Hamiltonian reads as energy Schrödinger equation

$$\hat{H}(\lambda) |s(\lambda)\rangle = E_s(\lambda) |s(\lambda)\rangle. \quad (2.2)$$

For every $\mathcal{H}(\lambda)$ the first k energies can be sorted from the lowest to create discrete set $\sigma(\hat{H}(\lambda)) \equiv \{E_0, \dots, E_k\}$. Clearly, there exists a bijection between all fibers

¹Spectrum of the operator consists of discrete spectrum, calculable as eigenvalue problem, continuous and residual spectrum.

$\mathcal{H}(\lambda)$, thus we can define *section* \sec_s , mapping eigenstate corresponding to energy E_s to base manifold

$$\sec_s : |o(\lambda)\rangle \mapsto \mathcal{U} \subset \mathbb{R}^n,$$

which is also bijection. For $\forall s \in \{0, k\}$ we can create according to sec. 1.3 new *energy manifolds*

$$\mathcal{M}_s \subset \mathcal{M},$$

with special importance of the *ground state manifold* \mathcal{M}_0 , which will be used later on for adiabatic transports of ground states. Geometrical intuition is drawn on fig. 2.2. Because those manifolds were created by sectioning, they are considered to be vector spaces in a geometrical sense. This was expected, because they contain quantum states.

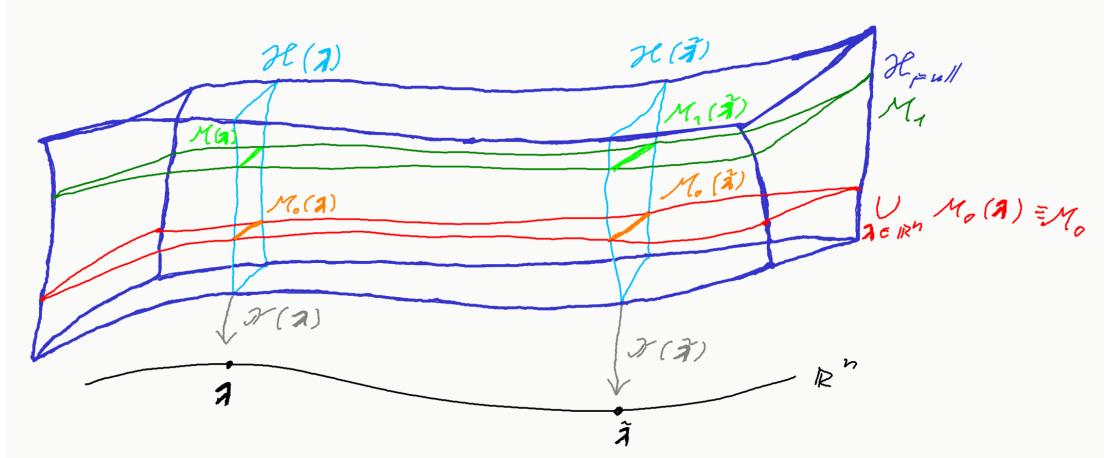


Figure 2.2: Geometrical intuition to transport on fiber manifold sections \mathcal{M}_i .

All Hilbert spaces above are considered to be spaces of *bare states* \mathcal{H} . In quantum mechanics, physical observables are related to the *space of rays*, defined as $\mathcal{PH} := \mathcal{H}/U(1)$, where elements of $U(1)$ are unitary transformations $e^{i\phi}$ for $\phi \in \mathbb{R}$ defining gauge symmetry between quantum states. The geometrical intuition is drawn for any λ on fig. 2.3.

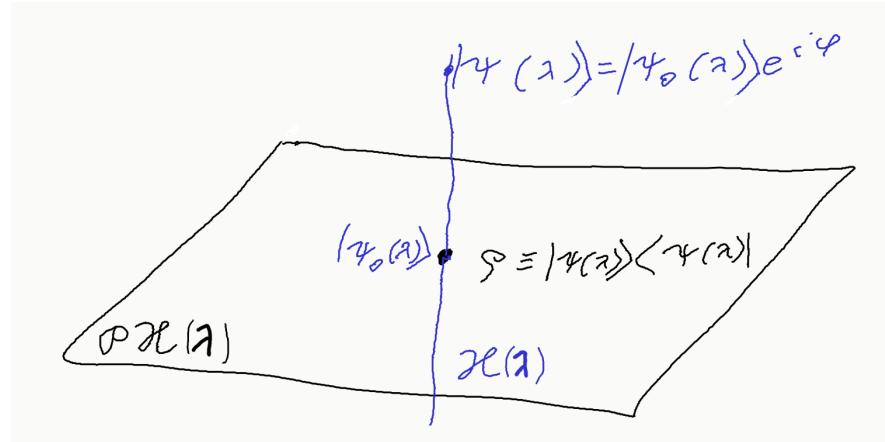


Figure 2.3: Space of bare states and its projection to space of rays $\mathcal{PH}(\lambda) := \mathcal{H}(\lambda)/U(1)$.

This means, that energy manifolds \mathcal{M}_s , generated by bare states, are again of fiber structure, defined as

$$(\mathcal{H}, \mathcal{PH}, \pi_{rays}, \{e^{i\phi} |\psi\rangle \text{ for } \phi \in \mathbb{R}\}),$$

where π_{rays} is just rule setting phase ϕ to zero. Because all \mathcal{M}_s are compact and connected, they are diffeomorphic to each other.

2.1 Transporting states

Let's now focus on decomposition of \mathcal{H}_{full} to different state manifolds \mathcal{M}_s , as displayed on figure 2.4.

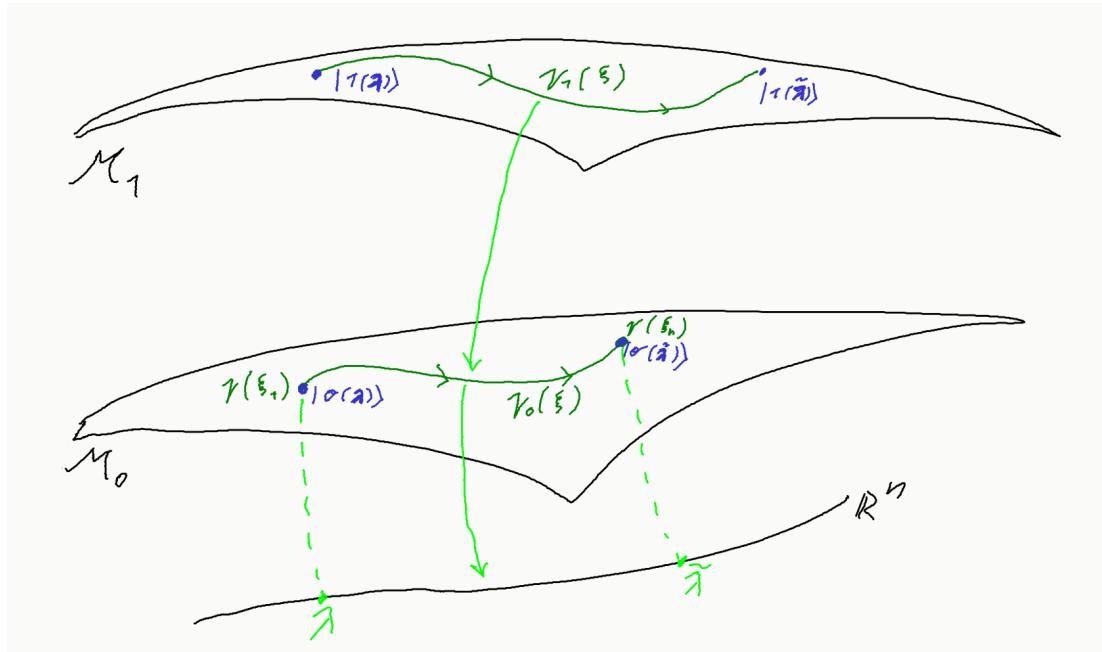


Figure 2.4: Geometrical intuition to transport on fiber manifold sections \mathcal{M}_s .

Changing state from eigenstate $|s(\lambda)\rangle$ to any pure state $|\psi(\tilde{\lambda})\rangle$, during some time period, is unitary transformation and can be thought of as *parallel transport on fiber bundle* between two vectors. Assuming the transport goes along curve $\{\gamma_s(\xi) | \xi \in [\xi_1, \xi_2] \subset \mathbb{R}\} \subset \mathcal{M}_s$ This can be written as

$$|\psi(t)\rangle \equiv \text{par}_{\gamma_s} |s(\lambda)\rangle = \exp\left(-\frac{i}{\hbar} \int_0^t E_s(\tau) d\tau\right) \exp(i\gamma_s(\xi)) |s(\lambda)\rangle. \quad (2.3)$$

The first exponential, the *dynamical phase*, is well known solution to energy Schrödinger equation 2.2 with $\lambda = \text{const}$ and depends only on time and energy of states during the transport. The second exponential is called *geometrical phase*. This phase is generally non-integrable, meaning it cannot be written simply as $\gamma_s(\lambda)$ and for some closed curve on \mathcal{M}

$$C = \{\lambda(\xi) | \xi \in [0, \Xi], \text{ such that } \lambda(0) = \lambda(\Xi)\} \quad (2.4)$$

we generally get $\text{par}_C |\psi(\lambda)\rangle \neq |\psi(\lambda)\rangle$. This property is sometimes more generally called an *anholonomy* and geometric intuition can be seen on fig. 2.5.

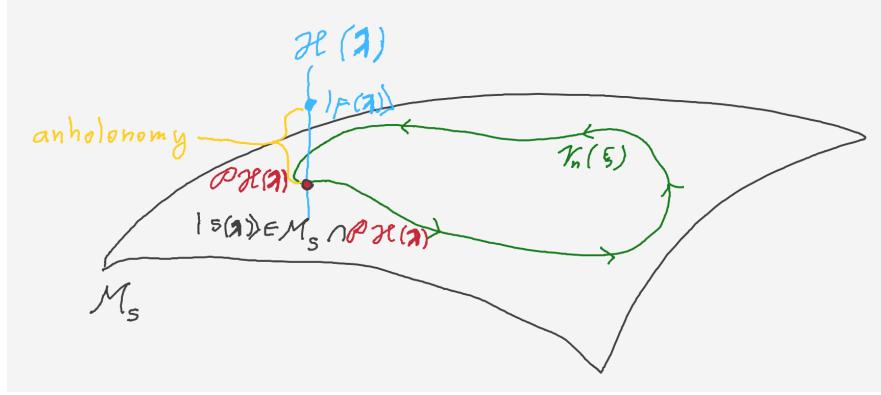


Figure 2.5: Parallel transporting around some closed curve C with anholonomy drawn in yellow.

For quantum states, the anholonomy can be measured as a non-zero angle between $|V\rangle$ and $\text{par}_C|V\rangle$, meaning

$$\langle V | \text{par}_C | V \rangle \neq 0.$$

Substituting general solution 2.3 to eq. 2.1 yields (see [Berry, 1984])

$$d_t \gamma(\lambda) = i \langle s(\lambda) | \partial_\mu s(\lambda) \rangle d_t \lambda^\mu(\lambda). \quad (2.5)$$

Integrating this equation around some closed curve C and assuming the dynamical phase to be zero, thus not exciting the system, we get

$$\gamma_n(C) = i \oint_C \langle s(\lambda) | \partial_\mu s(\lambda) \rangle d\lambda^\mu. \quad (2.6)$$

We see, that the geometric phase does not depend on energy or time, only on the sequence of Hamiltonians, which means it depends only on the path itself.

The problem of expressions above lies in $\partial_\lambda s(\lambda)$, which locally requires knowledge of single-valued basis $\{|0\rangle, \dots, |k\rangle\}$. This can be avoided in 3-dimensions using Stokes's theorem for S as the surface with boundary $\partial S = C$, for coordinate gradient ∇

$$\begin{aligned} \gamma_s(C) &= -\text{Im} \iint_C dS \cdot \nabla \times \langle s(\lambda) | \nabla n(\lambda) \rangle \\ &= -\text{Im} \iint_C dS \cdot \langle \nabla s(\lambda) | \times | \nabla s(\lambda) \rangle \\ &= -\text{Im} \iint_C dS \cdot \sum_{m \neq s} \langle \nabla s(\lambda) | m(\lambda) \rangle \times \langle m(\lambda) | \nabla s(\lambda) \rangle \\ &= -\iint_C dS \cdot V_s(\lambda) \end{aligned} \quad (2.7)$$

for

$$V_s(\lambda) = \text{Im} \frac{\langle s(\lambda) | \nabla_\lambda \hat{H}(\lambda) | m(\lambda) \rangle \times \langle m(\lambda) | \nabla_\lambda \hat{H}(\lambda) | s(\lambda) \rangle}{(E_m(\lambda) - E_s(\lambda))^2} \quad (2.8)$$

where the element of summation $m = s$ in third step of derivation is real, therefore has no influence on γ_s and can be omitted. The last equivalence holds, because

if we differentiate the Schrödinger equation 2.2, we get for any $|s\rangle, |m\rangle \in \mathcal{M}$

$$\begin{aligned} \nabla \hat{H} |s(\boldsymbol{\lambda})\rangle + \hat{H} |\nabla s(\boldsymbol{\lambda})\rangle &= E_s |\nabla s(\boldsymbol{\lambda})\rangle \\ \langle m(\boldsymbol{\lambda}) | \nabla \hat{H} |s(\boldsymbol{\lambda})\rangle + \langle m(\boldsymbol{\lambda}) | E_m |\nabla s(\boldsymbol{\lambda})\rangle &= \langle m(\boldsymbol{\lambda}) | \nabla \hat{H} |s(\boldsymbol{\lambda})\rangle \\ \langle m(\boldsymbol{\lambda}) | \nabla s(\boldsymbol{\lambda})\rangle &= \frac{\langle m(\boldsymbol{\lambda}) | \nabla \hat{H} |s(\boldsymbol{\lambda})\rangle}{E_m(\boldsymbol{\lambda}) - E_s(\boldsymbol{\lambda})}, \quad s \neq m, \end{aligned} \quad (2.9)$$

where we used $|\nabla s\rangle \equiv \nabla |s\rangle$. Comparing the first expression in eq. 2.7 with its last one and extending it to real numbers, we get

$$V_s(\boldsymbol{\lambda}) = \nabla \times \langle s(\boldsymbol{\lambda}) | \nabla m(\boldsymbol{\lambda}) \rangle, \quad (2.10)$$

defining vector potential of $V_s(\boldsymbol{\lambda})$. In addition, it extends our definition from single valued basis to any solution of 2.2, thus instead of ground state manifold, we can use any \mathcal{M}_s .

As was mentioned, the above procedure from eq. 2.6 was performed only for three-dimensional space. Proper generalization to n-dimensional space would yield, see [Berry, 1984],

$$\gamma_s(C) = - \iint_C dS \cdot \text{Im} \frac{\langle s(\boldsymbol{\lambda}) | d\hat{H}(\boldsymbol{\lambda}) | m(\boldsymbol{\lambda}) \rangle \wedge \langle m(\boldsymbol{\lambda}) | d\hat{H}(\boldsymbol{\lambda}) | s(\boldsymbol{\lambda}) \rangle}{(E_m(\boldsymbol{\lambda}) - E_s(\boldsymbol{\lambda}))^2}, \quad (2.11)$$

which will not be needed in this thesis.

2.2 Metric and geometric tensor

As a playground for this chapter, we will choose the ground state manifold \mathcal{M}_0 , but it can be easily generalized to any energy states manifold \mathcal{M}_s . From now on we will use natural units, so $\hbar = 1$.

Let's first look at \mathcal{PM}_s , which is needed to be *gauge independent*. Gauge dependence in quantum mechanics means, that the change in phase factor ϕ of some state $|o(\boldsymbol{\lambda})\rangle \in \mathbb{T}_{\boldsymbol{\lambda}}\mathcal{M}$ induces the change

$$|o(\boldsymbol{\lambda})\rangle \mapsto e^{i\phi(\boldsymbol{\lambda})} |o(\boldsymbol{\lambda})\rangle \implies \langle o(\boldsymbol{\lambda}) | \nabla o(\boldsymbol{\lambda}) \rangle \mapsto \langle o(\boldsymbol{\lambda}) | \nabla o(\boldsymbol{\lambda}) \rangle + i \nabla \phi(\boldsymbol{\lambda}) \quad (2.12)$$

For $\phi(\boldsymbol{\lambda}) \in \mathcal{C}^2$ we see from eq. 2.10, that gauge independent choice would be for infinitesimal change for example

$$f = \langle o(\boldsymbol{\lambda} + \delta\boldsymbol{\lambda}) | o(\boldsymbol{\lambda}) \rangle, \quad (2.13)$$

sometimes referred to as the *fidelity of a ground state*². We can see it's physical meaning imagining *quantum quench* (rapid change of some Hamiltonian parameters), in which case f^2 is the probability that system will remain in the new

²Generalization for any mixed state is for two density matrices σ, ρ as

$$f := \left(\text{Tr} \sqrt{\sqrt{\rho} \sigma \sqrt{\rho}} \right)$$

ground state. $1 - f^2$ is therefore probability of exciting the system during this quench, which leads to the definition of *distance on \mathcal{M}_0* ³

$$ds^2 \equiv 1 - f^2 = 1 - |\langle o(\lambda + \delta\lambda) | o(\lambda) \rangle|^2. \quad (2.14)$$

We can easily check, that the axioms of metric defined using distance as bilinear form s between elements $|\psi\rangle, |\phi\rangle \in \mathcal{M}_0$ are for **closed systems** satisfied:

- identity of indiscernibles $s(|\psi\rangle, e^{i\alpha}|\psi\rangle) = 0 \Leftrightarrow |\psi\rangle = |\phi\rangle, \alpha \in \mathbb{R}$,
- symmetry for any two states $|\psi\rangle, |\phi\rangle$ is implied by $|\langle \psi | \phi \rangle| = |\langle \phi | \psi \rangle|$
- triangle inequality: $s(|\psi\rangle, |\psi_2\rangle) < s(|\psi\rangle, |\psi\rangle_1) + s(|\psi_1\rangle, |\psi_2\rangle)$ for any $|\psi_1\rangle$.

Because $1 - f^2 > 0$, the first term of Taylor expansion is zero, thus we have for a metric tensor

$$ds^2 = g_{\mu\nu} d\lambda^\mu d\lambda^\nu + \mathcal{O}(\lambda^3). \quad (2.15)$$

Let's define the metric on the space of states \mathcal{M}_0 and then see, how it corresponds to fidelity. This metric is called the *Geometric tensor*, and can be expressed as

$$\chi_{\mu\nu} := \langle \partial_\mu o | \partial_\nu o \rangle_c \equiv \langle \partial_\mu o | \partial_\nu o \rangle - \langle \partial_\mu o | o \rangle \langle o | \partial_\nu o \rangle, \quad (2.16)$$

where shortened notation $\partial_\nu := \frac{\partial}{\partial \lambda^\nu}$ was used.

Proof for Geometric tensor expression. In one-dimensional case we get from eq. 2.13 using Taylor expansion and shortened notation $o := o(\lambda)$

$$f = 1 + \langle \partial_\lambda o | o \rangle \delta\lambda - \frac{1}{2} \langle \partial_\lambda^2 o | o \rangle \delta\lambda^2 + \mathcal{O}(\lambda^3), \quad (2.17)$$

which plugged into eq. 2.14 gives

$$ds^2 = 1 - \bar{f}f = \left[-\langle o | \partial_\lambda o \rangle \langle \partial_\lambda o | o \rangle + \frac{1}{2} \langle \partial_\lambda^2 o | o \rangle + \frac{1}{2} \langle o | \partial_\lambda^2 o \rangle \right] \delta\lambda^2. \quad (2.18)$$

Generalizing for $\lambda \in \mathbb{R}^n$, we get⁴

$$ds^2 = \left[-\langle o | \partial_\mu o \rangle \langle \partial_\nu o | o \rangle + \underbrace{\frac{1}{2} \langle o | \partial_\mu \partial_\nu o | o \rangle + \frac{1}{2} \langle o | \partial_\nu \partial_\mu o | o \rangle}_{\Re \langle \partial_m u o | \partial_\nu o \rangle} \right] \delta\lambda^\mu \delta\lambda^\nu. \quad (2.19)$$

Because only symmetric part contributes in the sum in eq. 2.15, this proves symmetric part of equation 2.16 and shows, that some more general tensor (*geometric tensor*) exists. \square

³d notation is in differential geometry assumed to be an exterior differential. On functions, it acts as $d : \mathcal{F}\mathcal{M} \rightarrow \mathcal{T}_1\mathcal{M}$ and intuitively corresponds to total differential from functional analysis.

⁴assuming Hermiticity of the derivative operator everywhere on the ground state manifold

It is practical to decompose it the geometric tensor

$$\chi_{\mu\nu} \equiv g_{\mu\nu} - i\frac{1}{2}\nu_{\mu\nu}, \quad (2.20)$$

where the *Fubini-Study tensor*⁵, as it's called, is metric on \mathcal{PM}_0 and can be expressed as

$$g_{\mu\nu} = \frac{\chi_{\mu\nu} + \chi_{\nu\mu}}{2} = \Re \langle \partial_\mu o | \partial_\nu o \rangle_c = \Re \sum_{o \neq j} \frac{\langle o | \frac{\partial \mathcal{H}}{\partial \lambda^\mu} | j \rangle \langle j | \frac{\partial \mathcal{H}}{\partial \lambda^\nu} | o \rangle}{(E_o - E_j)^2}, \quad (2.21)$$

and the *curvature tensor* a.k.a. *Berry curvature* is

$$\nu_{\mu\nu} = i(\chi_{\mu\nu} - \chi_{\nu\mu}) = \text{Im} \langle o | [\overleftarrow{\partial}_\nu, \partial_\mu] | o \rangle_c = -2\text{Im} \sum_{o \neq j} \frac{\langle o | \frac{\partial \mathcal{H}}{\partial \lambda^\mu} | j \rangle \langle j | \frac{\partial \mathcal{H}}{\partial \lambda^\nu} | o \rangle}{(E_o - E_j)^2}, \quad (2.22)$$

where $\overleftarrow{\partial}_\nu$ affects the covector on the left.

2.3 Derivation of the geometric tensor

To prove the correspondence of geometric tensor, described by eq. 2.16, to distance on \mathcal{M}_0 , see eq. 2.14, we start with eigenstate $|o(\boldsymbol{\lambda})\rangle \in \mathcal{M}_0 \cap \mathcal{H}(\boldsymbol{\lambda})$. Changing parameters $\boldsymbol{\lambda}$ to $\boldsymbol{\lambda} + \delta\boldsymbol{\lambda}$ results in Hamiltonian \hat{H}_f with eigenstates $|s(\boldsymbol{\lambda} + \delta\boldsymbol{\lambda})\rangle \in \mathcal{M}_s \cap \mathcal{H}(\boldsymbol{\lambda} + \delta\boldsymbol{\lambda})$, meaning it can be excited. Probability amplitude of going to such state is

$$\begin{aligned} a_s &= \langle s(\boldsymbol{\lambda} + \delta\boldsymbol{\lambda}) | o(\boldsymbol{\lambda}) \rangle \approx \delta\lambda^\mu \langle \partial_\mu s(\boldsymbol{\lambda}) | o(\boldsymbol{\lambda}) \rangle \\ &= -\delta\lambda^\mu \langle s(\boldsymbol{\lambda}) | \partial_\mu | o(\boldsymbol{\lambda}) \rangle. \end{aligned} \quad (2.23)$$

If we introduce the *gauge potential*, aka *calibration potential*, as⁶

$$\hat{\mathcal{A}}_\mu \equiv i\partial_\mu, \quad (2.24)$$

the probability amplitude can be expressed as

$$a = i \langle s(\boldsymbol{\lambda}) | \hat{\mathcal{A}}_\mu | o(\boldsymbol{\lambda}) \rangle \delta\lambda^\mu, \quad (2.25)$$

which has meaning of matrix elements of the gauge potential. Probability of the excitation i.e. transition to any state $s > 0$ from ground state is then (omitting the $\boldsymbol{\lambda}$ dependence in notation)

$$\begin{aligned} \sum_{s \neq 0} |a_s|^2 &= \sum_{s \neq 0} \delta\lambda^\mu \delta\lambda^\nu \langle o | \hat{\mathcal{A}}_\mu | s \rangle \langle s | \hat{\mathcal{A}}_\nu | o \rangle + \mathcal{O}(|\delta\lambda^3|) \\ &= \delta\lambda^\mu \delta\lambda^\nu \langle o | \hat{\mathcal{A}}_\mu \hat{\mathcal{A}}_\nu | o \rangle_c =: \delta\lambda^\mu \delta\lambda^\nu \chi_{\mu\nu} + \mathcal{O}(|\delta\lambda^3|), \end{aligned} \quad (2.26)$$

where last term defines the geometric tensor.

The gauge transformation means, that if the fidelity $f < 1$, this potential can be subtracted from the Hamiltonian leading to $f = 1$. This will be discussed later on in chapter 2.6.

2.4 Hamiltonian induced flow on manifold

Hamiltonian induces flow on manifold according to Schrödinger equation....

⁵In some literature, this is called Geometric tensor

⁶In SI units, the gauge potential is $\hat{\mathcal{A}}_\mu \equiv i\hbar\partial_\mu$

2.5 Adiabaticity

Adiabatic transformation is such a transformation from \mathcal{M} to \mathcal{M} , which does not excite the system, meaning the fidelity $f = 1$. Generally it can be achieved by two ways – infinitely slow transformation of states, or adding some *counter-diabatic elements* to the Hamiltonian to counter the excitation.

In this chapter, we will be dealing with the system described by the finite-dimensional Hamiltonian $\hat{H}(\boldsymbol{\lambda})$ which drives the system according to Schrödinger equation from some initial state $|s(\boldsymbol{\lambda})\rangle$ to $|s(\tilde{\boldsymbol{\lambda}})\rangle$ along some path. Again we reduce the dimension from \mathbb{R}^n to 1 prescribing the curve $\gamma(\lambda)$ parametrized by time t , so we get $\hat{H} = \hat{H}(\lambda)$. Before going through the details of adiabatic transformations, let's define its meaning properly.

Definition 7 (Adiabaticity). *Slow change of parameters driving Hamiltonian in a sense, that it does not excite the system and allows the system to return to the same energetic state after circulation around any closed path on the manifold with fidelity $f = 1$. For more see Theorem 5.*

2.5.1 Slow transports

[Kolodrubetz et al., 2017][chap. 2.3] As was mentioned in the introduction to this chapter, one way to change the system parameters without exciting it is to change the driving parameter slowly enough. The meaning of the word "slow" clears up next theorem.

Theorem 5 (Adiabatic theorem). *For Hamiltonian \hat{H} varying in the time range T , the solution of the Schrödinger equation*

$$\hat{H}(\lambda) |\psi_n(\lambda)\rangle = E_n(\lambda) |\psi_n(\lambda)\rangle$$

with initial condition in x -representation $\langle x|\psi(t=0\rangle = \psi_n(x,0)$ can be approximated as

$$||\psi(\lambda) - \psi_{ad}(\lambda)|| \approx o\left(\frac{1}{T}\right) \quad (2.27)$$

for adiabatic state

$$|\psi_{ad}\rangle = e^{\theta_n(\lambda)} e^{\gamma_n(\lambda)} |\psi(\lambda)\rangle, \quad (2.28)$$

where we define nongeometrical phase induced by energy transitions,

$$\theta_n(\lambda) \equiv -\frac{1}{\hbar} \int_0^t E_n(\tau) d\tau$$

and geometrical phase, also called Berry phase

$$\gamma_n(\lambda) \equiv \int_0^t \underbrace{i \langle \psi_n(\tau) | \partial_t \psi_n(\tau) \rangle}_{\nu_n(\tau)} d\tau.$$

Proof. TBD (na wiki je)

□

Assume differentiable and non-singular Hamiltonian $\hat{H}(\boldsymbol{\lambda})$ with degenerate basis

$\{|m, \boldsymbol{\lambda}\rangle\}_m$ called the *adiabatic basis*. This is generally the family of adiabatically connected eigenstates⁷ The transition amplitude between states for adiabatic change is

$$0 = \langle m(\boldsymbol{\lambda}) | \hat{H} | n(\tilde{\boldsymbol{\lambda}}) \rangle \quad \text{for } n \neq m, \forall \boldsymbol{\lambda}, \forall \tilde{\boldsymbol{\lambda}}. \quad (2.29)$$

This can be driven along some curve $\gamma(\lambda)$, i.e. differentiated by ∂_t :

$$\begin{aligned} 0 &= \langle \partial_t m(\boldsymbol{\lambda}) | \hat{H}(\tilde{\boldsymbol{\lambda}}) | n(\tilde{\boldsymbol{\lambda}}) \rangle + \langle m(\boldsymbol{\lambda}) | \overbrace{\partial_t \hat{H}(\tilde{\boldsymbol{\lambda}})}^{\approx \partial_t \hat{H}(\boldsymbol{\lambda})} | n(\tilde{\boldsymbol{\lambda}}) \rangle + \langle m(\boldsymbol{\lambda}) | \hat{H}(\tilde{\boldsymbol{\lambda}}) | \partial_t n(\tilde{\boldsymbol{\lambda}}) \rangle \\ &= E_n(\lambda) \langle \partial_t m(\boldsymbol{\lambda}) | n(\tilde{\boldsymbol{\lambda}}) \rangle + E_m(\lambda) \langle m(\boldsymbol{\lambda}) | \partial_t n(\tilde{\boldsymbol{\lambda}}) \rangle + \langle m(\boldsymbol{\lambda}) | \partial_t \hat{H}(\tilde{\boldsymbol{\lambda}}) | n(\tilde{\boldsymbol{\lambda}}) \rangle \\ &= (E_m(\lambda) - E_n(\lambda)) \underbrace{\langle m | \partial_t(\tilde{\boldsymbol{\lambda}}) \rangle}_{-\frac{i}{\hbar} \langle m | \hat{\mathcal{A}}_t | n(\tilde{\boldsymbol{\lambda}}) \rangle} + \langle m | \partial_t \hat{H} | n(\tilde{\boldsymbol{\lambda}}) \rangle. \end{aligned} \quad (2.30)$$

In matrix form, we can rewrite this equation as

$$i\hbar \partial_t \hat{H} = [\hat{\mathcal{A}}_t, \hat{H}] - i\hbar \hat{M}_t \quad \text{for } \hat{M}_t \equiv - \sum_n \frac{\partial E_n(\lambda)}{\partial t} |n(\lambda)\rangle \langle n(\lambda)|. \quad (2.31)$$

\hat{M} is diagonal in energetic basis and it's elements has meaning of *generalized force*, which correspond to corresponding energetic states. We can easily see that $[\hat{H}, \hat{M}] = 0$, implying

$$[\hat{H}, i\hbar \partial_t \hat{H} - [\hat{\mathcal{A}}_t, \hat{H}]] = 0. \quad (2.32)$$

This can be used as the definition for *counter-diabatic potential* $\hat{\mathcal{A}}_t$. The strength of this equation lies in the fact, that it finds counter-diabatic potential without the need of Hamiltonian diagonalization.

2.6 Geodetic driving

2.7 Gauge potentials

In section 2.3 we introduced the gauge potential without proving its gauge meaning, but only stating its correspondence to transition probability, see eq. 2.25. *Gauge transformations*, in classical mechanics called *canonical*, can be defined such, that they preserve Lagrangian of the system under local transformations from some Lie group. This implies, that gauge transformed Hamiltonian $\mathcal{H}(\boldsymbol{\lambda})$ and $\mathcal{H}(\boldsymbol{\lambda} + d\boldsymbol{\lambda})$ commutes with its canonically transformed version⁸

$$[\hat{H}(\boldsymbol{\lambda}), \hat{H}(\boldsymbol{\lambda} + \delta\boldsymbol{\lambda})] = 0. \quad (2.33)$$

To understand the meaning of gauge symmetries, let's first consider classical system and then move to quantum mechanics.

⁷In the case of energy level crossing, the eigenstates are not unified, because transition between them is not adiabatical.

⁸This can be easily reformulated to the world of classical physics, where the commutator is replaced by Poisson bracket.

2.7.1 Classical gauge potential

In the Hamiltonian classical mechanics, we assume the manifold \mathcal{M} part of the phase space defined using Hamiltonian $H = H(p_i, q_i)$, where momentum p_i and position q_i are assumed to form the orthogonal basis of the phase space

$$\{q^i, p_j\} = \delta_j^i, \quad (2.34)$$

which also defines *calibration freedom* in their choice. *Canonical transformations* then by definition preserve this formula. Using the *Poisson bracket*, defined as

$$\{A, B\} := \frac{\partial A}{\partial q^j} \frac{\partial B}{\partial p_j} - \frac{\partial B}{\partial q^j} \frac{\partial A}{\partial p_j}, \quad (2.35)$$

we will examine continuous canonical transformations generated by gauge potential \mathcal{A}_λ

$$q^j(\lambda + \delta\lambda) = q^j(\lambda) - \frac{\partial \mathcal{A}_\lambda(\mathbf{p}, \mathbf{q})}{\partial p_j} \delta\lambda \Rightarrow \frac{\partial q^j}{\partial \lambda} = -\frac{\partial \mathcal{A}_\lambda}{\partial p_j} = \{\mathcal{A}_\lambda, q^j\} \quad (2.36)$$

$$p_j(\lambda + \delta\lambda) = p_j(\lambda) - \frac{\partial \mathcal{A}_\lambda(\mathbf{p}, \mathbf{q})}{\partial q^j} \delta\lambda \Rightarrow \frac{\partial p_j}{\partial \lambda} = -\frac{\partial \mathcal{A}_\lambda}{\partial q^j} = \{\mathcal{A}_\lambda, p_j\}. \quad (2.37)$$

Substituting this to relations of orthogonality 2.34, we get

$$\{q^j(\lambda + \delta\lambda), p_j(\lambda + \delta\lambda)\} = \delta_j^i + \mathcal{O}(\delta\lambda^2). \quad (2.38)$$

If λ is time parameter and $\mathcal{A}_t = -H$, equations 2.36,2.37 are identical to the Hamilton equations

$$\begin{aligned} \dot{q}^j &= -\{H, q^j\} = \frac{\partial H}{\partial p_j} \\ \dot{p}_j &= -\{H, p_j\} = -\frac{\partial H}{\partial q^j}. \end{aligned} \quad (2.39)$$

Because the Hamiltonian is generator of the movement in the phase space (\mathbf{q}, \mathbf{p}) , we can interpret \mathcal{A}_t as the generators of the movement on \mathcal{M} . Other specific choice might be $\lambda = X^i$, which gives us the momentum components $\mathcal{A}_{X^i} = p_i$.

Generally every gauge symmetry is generated by its gauge potential and corresponds to some conserved property, as theorem of Emma Nöether states.

2.7.2 Quantum gauge potential

[Kolodrubetz et al., 2017][chap. 2.2] The role of Poisson brackets in quantum mechanics is taken by commutators, canonical transformations are called *unitary transformations* and calibration freedom is hidden in the choice of basis. Now let's find some special basis transformations \hat{U} between initial system S and the transformed \tilde{S} . Both of them describe the system with Hamiltonian $\hat{H}(\boldsymbol{\lambda})$ with eigenstates $|n(\boldsymbol{\lambda})\rangle$ and eigenstate manifolds $\mathcal{M}_n \equiv \cup_{\boldsymbol{\lambda}} \{|n(\boldsymbol{\lambda})\rangle\}$.

From fiber structure goes⁹, that any state of $\hat{H}(\boldsymbol{\lambda})$ for $\forall \boldsymbol{\lambda} \in U \subset \mathbb{R}^n$ can be decomposed as

$$|\psi(\boldsymbol{\lambda})\rangle \equiv \sum_n \psi_n(\boldsymbol{\lambda}) |n\rangle \quad (2.40)$$

⁹especially from the fact, that all spaces $\hat{H}(\boldsymbol{\lambda})$ are isomorphic to each other

for some coordinate independent basis $\{|n\rangle\}_n$. Then there exist unitary transformation

$$\hat{U}(\boldsymbol{\lambda}) : \tilde{S} \rightarrow S, \quad \hat{U}(\boldsymbol{\lambda}) |m(\boldsymbol{\lambda})\rangle = |n\rangle. \quad (2.41)$$

The wave function $|\psi\rangle$ in S can be decomposed using Schmidt decomposition¹⁰

$$|\psi(\boldsymbol{\lambda})\rangle = \sum_{m,n} \psi_n(\boldsymbol{\lambda}) |m(\boldsymbol{\lambda})\rangle \underbrace{\langle m(\boldsymbol{\lambda})|n\rangle}_{U_{mn}(\boldsymbol{\lambda})} = \sum_m \underbrace{\tilde{\psi}_m(\boldsymbol{\lambda})}_{\langle m(\boldsymbol{\lambda})|} |m(\boldsymbol{\lambda})\rangle, \quad (2.42)$$

where $U_{mn}(\boldsymbol{\lambda})$ are matrix elements of unitary transformation $\hat{U}(\boldsymbol{\lambda})$. In this work, we will be interested only in the gauge transformations preserving energy of the system.

2.7.3 Adiabatic gauge potential

Adiabatic gauge potentials, sometimes just *adiabatic potentials*, are generators of unitary transformations, so we can define them analogically to the classical case

$$i\hbar\partial_\lambda |\tilde{\psi}(\boldsymbol{\lambda})\rangle = i\hbar\partial_\lambda \left(\hat{U}^+(\boldsymbol{\lambda}) |\psi\rangle \right) = \underbrace{i\hbar \left(\partial_\lambda \hat{U}^+(\boldsymbol{\lambda}) \right)}_{-\tilde{\mathcal{A}}_\lambda} \hat{U}(\boldsymbol{\lambda}) |\tilde{\psi}(\boldsymbol{\lambda})\rangle. \quad (2.43)$$

The adiabatic potential $\tilde{\mathcal{A}}_\lambda$ can be transformed to non-tilde system as

$$\begin{aligned} \hat{\mathcal{A}}_\lambda &= \hat{U}(\boldsymbol{\lambda}) \tilde{\mathcal{A}}_\lambda \hat{U}^+(\boldsymbol{\lambda}) = -i\hbar \hat{U}(\boldsymbol{\lambda}) \left(\partial_\lambda \hat{U}^+(\boldsymbol{\lambda}) \right) = \\ &= -i\hbar \partial_\lambda \left(\underbrace{U^+(\boldsymbol{\lambda}) U(\boldsymbol{\lambda})}_1 \right) - \left(\partial_\lambda U(\boldsymbol{\lambda}) \right) U^+(\boldsymbol{\lambda}) = i\hbar \left(\partial_\lambda U(\boldsymbol{\lambda}) \right) U^+(\boldsymbol{\lambda}). \end{aligned} \quad (2.44)$$

From this we get the equations for adiabatic potential in two systems

$$\hat{\mathcal{A}}_\lambda = i\hbar \left(\partial_\lambda U(\boldsymbol{\lambda}) \right) U^+(\boldsymbol{\lambda}) \quad (2.45)$$

$$\tilde{\mathcal{A}}_\lambda = -i\hbar \left(\partial_\lambda \hat{U}^+(\boldsymbol{\lambda}) \right) \hat{U}(\boldsymbol{\lambda}) \quad (2.46)$$

which can be shown to be Hermitian

$$\tilde{\mathcal{A}}_\lambda^+ = i\hbar U(\boldsymbol{\lambda})^+ \left(\partial_\lambda \hat{U}(\boldsymbol{\lambda}) \right) = -i\hbar \left(\partial_\lambda \hat{U}(\boldsymbol{\lambda})^+ \right) \hat{U}(\boldsymbol{\lambda}) = \tilde{\mathcal{A}}_\lambda, \quad (2.47)$$

analogically for non-tilde potential. Using the eigenbasis of \hat{H} , the matrix elements are

$$\langle n | \tilde{\mathcal{A}}_\lambda | m \rangle = i\hbar \langle n | \hat{U}(\boldsymbol{\lambda})^+ \partial_\lambda \hat{U}(\boldsymbol{\lambda}) | m \rangle = i\hbar \langle n(\boldsymbol{\lambda}) | \partial_\lambda | m(\boldsymbol{\lambda}) \rangle. \quad (2.48)$$

and because

$$\langle n(\boldsymbol{\lambda}) | \hat{\mathcal{A}}_\lambda | m(\boldsymbol{\lambda}) \rangle = \langle n | \tilde{\mathcal{A}}_\lambda | m \rangle, \quad (2.49)$$

we get

$$\hat{\mathcal{A}}_\lambda = i\hbar\partial_\lambda. \quad (2.50)$$

¹⁰The Schmidt decomposition can be performed in finite dimension, or if the Hamiltonian is compact, which is not automatic in quantum mechanics. What's more, the Hamiltonian is usually not even bounded. Anyway, for simple systems with bounded energy we can assume so.

Adiabatic gauge transformations are class of gauge transformations with fidelity $f = 1$. This means, that if the system is driven by Hamiltonian $\hat{H}(\boldsymbol{\lambda})$ with fidelity $f < 1$, there exists such adiabatic potential \mathcal{A}_λ , that driving of the same system using $\hat{H} - \mathcal{A}_\lambda$ has fidelity $f = 1$.

The adiabatic gauge potentials can then be understood as affine connections defining the parallel transport on fiber bundle, if we define covariant derivative as

$$D_\mu = \partial_\mu + i\hat{\mathcal{A}}_\mu, \quad (2.51)$$

which yields $D_\mu |\psi_n\rangle = 0$ for every eigenstate, which yields, that the transport of eigenvalues on \mathcal{M}_0 is parallel. $\hat{\mathcal{A}}_\mu$ is generally defined 2.45, which generally gives non-zero covariant derivative for states not belonging to \mathcal{M}_0 .

Finding of those potentials has many practical applications, so let's introduce one analytical procedure of finding them.

2.8 Counter-diabatic driving

[Kolodrubetz et al., 2017][page 15–17] The main idea of a counter-diabatic driving is, that any excitation of the system can be countered by adding so called *counter-diabatic potential* to the Hamiltonian. Consider again any eigenstate $|\psi(t)\rangle$ of the Hamiltonian $\hat{H} = \hat{H}(\lambda)$ driven along the curve $\gamma(\lambda(t))$ on \mathcal{M}_0 depending on time t , during which the fidelity $f \neq 0$. Because the system is not measured during the trip, it can't be stated if or if not it was excited, but the main goal here is to make fidelity zero, which is iff $\tilde{\hat{H}}$ is diagonal. For diagonalizable Hamiltonian, there exist a transformation, see eq. 2.41, for which the fidelity will be zero. Such a transformation does not have to be unique, but we can choose any one of them. This can be seen more clearly from direct transformation of the Schrödinger equation.

The Schrödinger equation

$$i\hbar \frac{d}{dt} |\psi(\lambda)\rangle = \hat{H}(\lambda) |\psi(\lambda)\rangle \quad (2.52)$$

can be transformed using

$$\hat{U}(\lambda)^+ |\psi(\lambda)\rangle = |\tilde{\psi}(\tilde{\boldsymbol{\lambda}})\rangle, \quad (2.53)$$

for which $\tilde{\hat{H}} := \hat{U}^+ \hat{H} \hat{U}$ is diagonal, leading to

$$i\hbar \frac{d}{dt} (\hat{U}(\tilde{\boldsymbol{\lambda}}) |\tilde{\psi}(\tilde{\boldsymbol{\lambda}})\rangle) = \hat{H}(\lambda) \hat{U}(\tilde{\boldsymbol{\lambda}}) |\tilde{\psi}(\tilde{\boldsymbol{\lambda}})\rangle \quad (2.54)$$

$$i\hbar \frac{d\lambda}{dt} \partial_\lambda \hat{U}(\tilde{\boldsymbol{\lambda}}) |\tilde{\psi}(\tilde{\boldsymbol{\lambda}})\rangle + i\hbar \hat{U}(\tilde{\boldsymbol{\lambda}}) \frac{d}{dt} |\tilde{\psi}(\tilde{\boldsymbol{\lambda}})\rangle = \hat{H}(\lambda) \hat{U}(\tilde{\boldsymbol{\lambda}}) |\tilde{\psi}(\tilde{\boldsymbol{\lambda}})\rangle. \quad (2.55)$$

This can be rewritten using adiabatic potential from eq. 2.50, using *dot* notation for time derivatives and omitting the points in which the objects are evaluated, as

$$i\hbar \frac{d}{dt} |\tilde{\psi}\rangle = \left[\hat{U}^+ \hat{H} \hat{U} - \dot{\lambda} \tilde{\hat{\mathcal{A}}}_\lambda \right] |\tilde{\psi}\rangle = \left[\tilde{\hat{H}} - \dot{\lambda} \tilde{\hat{\mathcal{A}}}_\lambda \right] |\tilde{\psi}\rangle =: \tilde{\hat{H}}_m |\tilde{\psi}\rangle, \quad (2.56)$$

where the term $-\dot{\lambda}\tilde{\mathcal{A}}_\lambda$ is called *Galilean* and \tilde{H}_m is the Hamiltonian in transformed system. Because \tilde{H} is diagonal, it drives $|\tilde{\psi}\rangle$ with fidelity $f = 1$. This means that for any driving defined by $\hat{H}(\lambda(t))$, which defines the unitary transformation \hat{U} , there exists such *counter-diabatic potential* $\hat{\mathcal{A}}_\lambda$, that $\hat{H}_m + \dot{\lambda}\hat{\mathcal{A}}_\lambda$ has $f = 1$.

This procedure does not directly tell us how to calculate the counter-diabatic potential, only states its existence. For many simple cases the calculation can be done analytically, but most often some approximation methods are needed.

2.8.1 Explicit form

If we now consider the parametrization with time $t := t$, \hat{U} can be explicitly expressed according to [ansatz](#) in eq. 2.3 as

$$\hat{U}(t) = \sum_n \exp\left(\frac{i}{\hbar} E_n(\tau) d\tau - \int_0^t \langle s(\tau) | \partial_\tau n(\tau) \rangle d\tau\right) |s(t)\rangle \langle s(0)|. \quad (2.57)$$

Inserting to eq. ??, we get explicit form of the Hamiltonian, which can be decomposed into the diagonal form of the original Hamiltonian and a counter-diabatic potential

$$\hat{H}(t) = \sum_n |n\rangle E_n \langle n| + i\hbar \sum_n |\partial_\lambda n\rangle \langle n| - \langle n | \partial_\lambda n \rangle |n\rangle \langle n| =: \hat{H}_0(t) + \hat{H}_1(t), \quad (2.58)$$

for shortened notation $|n\rangle \equiv |n(t)\rangle$, analogically for bras. Using

$$\hat{H}_0(t) |n\rangle = E_n |n\rangle \quad \Rightarrow \quad \langle m | \partial_\lambda n \rangle = \frac{\langle m | \partial_\lambda \hat{H}_0 | n \rangle}{E_n - E_m} \quad (2.59)$$

we have explicit formula

$$\hat{H}_1(t) = i\hbar \sum_{m \neq n} \frac{|m\rangle \langle m | \partial_\lambda \hat{H}_0 | n \rangle \langle n|}{E_n - E_m} \quad (2.60)$$

2.9 Berry phase and curvature

Let's now consider only the ground state manifold, because it will be used later on. Thought every step can be easily generalized to any state manifold \mathcal{M}_m .

On the ground state manifold $\mathcal{M}_0 \equiv \cup_\lambda \{|o(\lambda)\rangle\}$, the *Berry connection* is defined as

$$A_\mu(\lambda) \equiv \langle o(\lambda) | \hat{\mathcal{A}}_\mu | o(\lambda) \rangle = -i\hbar \langle o(\lambda) | \partial_\mu | o(\lambda) \rangle, \quad (2.61)$$

which uses the decomposition of λ to some basis, thus it empowers us to take derivatives in any direction in the base manifold \mathbb{R}^n and thus the geometric tensor can be written as

$$\chi_{\mu\nu}(\lambda) = \partial_\mu A_\nu(\lambda) - \partial_\nu A_\mu(\lambda) \quad (2.62)$$

and the *Berry phase*¹¹ as the integral of a connection along some closed curve \mathcal{C}

$$\phi_B \equiv - \oint_{\mathcal{C}} A_\mu(\lambda) d\lambda^\mu = \int_{\mathcal{S}} \chi_{\mu\nu}(\lambda) d\lambda^\mu \wedge d\lambda^\nu, \quad (2.65)$$

where we used the Stokes theorem for some area \mathcal{S} with boundary $\partial\mathcal{S} = \mathcal{C}$.

¹¹ The reasonability of this definition can be seen, if we assume the ground state of a free

2.10 Approximations of adiabatic potentials

Adiabatic potentials can be calculated from the principal of minimal action, which leads to variational method.

If the difference between eigenstates of \hat{H} is small, or generalized force between some states is zero, the computation of the adiabatic potential is numerically unstable. The knowledge of exact adiabatic potential would allow to maintain the system in the ground state thus not exciting it, as the Eigenstate thermalization hypotheses states.

Hypotheses 1 (Eigenstate thermalization hypotheses). *For the difference between eigenstates of \hat{H} and extensive thermodynamic entropy S , it holds that*

$$E_n - E_m \propto \exp\left(\frac{S}{2}\right). \quad (2.66)$$

If the states are close, better approximation would be $E_n - E_m \propto \exp(S)$. For matrix elements it holds, that they vanish exponentially with the characteristic scale of the system a , i.e.

$$\langle m | \hat{\mathcal{A}}_\lambda | n \rangle = i\hbar \frac{\langle m | \partial_\lambda \hat{H} | n \rangle}{E_m - E_n} \propto \exp(-a). \quad (2.67)$$

Fortunately in the limit "number of particles" $\rightarrow \infty$ the expression in eq. 2.67 converges.

2.10.1 Variational methods

In the case of simple systems, the adiabatic potentials can be found analytically, but for more complicated Hamiltonians we will be forced to use approximations, or some perturbational and variational methods.

particle $\langle \mathbf{x} | i(\boldsymbol{\lambda}) \rangle = i(\mathbf{x}, \boldsymbol{\lambda}) = |i(\mathbf{x})| e^{i\phi(\boldsymbol{\lambda})}$, then the Berry connection is

$$A_\mu = - \int d\mathbf{x} |i(\mathbf{x}, \boldsymbol{\lambda})|^2 \partial_\mu \phi(\boldsymbol{\lambda}) = -\partial_\mu \phi(\boldsymbol{\lambda}) \quad (2.63)$$

and Berry phase

$$\phi_B = \oint_C \partial_\mu \phi d\lambda^\mu, \quad (2.64)$$

which represents total phase accumulated by the wave function. It is really the analogy for Berry phase in classical mechanics, which for example in the case Foucault pendulum on one trip around the Sun makes $\phi_B = 2\pi$

3. Geodesics

Of special importance in theory of the ground state manifold are geodesics. Its not yet clear what role they have in general, but there are known cases where they play a role of "maximum fidelity at any time" transport. This means, that if we compare geodesic $\mathcal{G}(t)$ with some curve $\gamma(t)$ spanning between points P_i and P_f during some time t_f , i.e. $\mathcal{G}(0) = \gamma(0) = P_i$, $\mathcal{G}(t_f) = \gamma(t_f) = P_f$, at any given time t the fidelity on corresponding point of \mathcal{G} will be less than of γ . This does not necesarilly mean, that the overall fidelity will be smaller for geodesics.
WHYYYYY?????

4. Spin 1/2

The Hamiltonian for a single spin in a magnetic field \mathbf{B} is

$$H = -\mu \mathbf{B} \cdot \mathbf{S} \quad (4.1)$$

for spin 1/2 we have using sigma matrices

$$S = \frac{\hbar}{2} (\sigma_x, \sigma_y, \sigma_z) \equiv \frac{\hbar}{2} \left(\begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \right) \quad (4.2)$$

The coordinate system can be oriented such that $\mathbf{B} = (0, 0, B_z)$ and eigenfunctions of the spin are thus in spherical coordinates θ, ϕ and x-representation

$$u_1 = \begin{pmatrix} \cos \frac{\theta}{2} e^{-i\phi} \\ \sin \frac{\theta}{2} \end{pmatrix}; \quad u_2 = \begin{pmatrix} -\sin \frac{\theta}{2} e^{-i\phi} \\ \cos \frac{\theta}{2} \end{pmatrix} \quad (4.3)$$

For both of these states the Berry connection can be defined using 2.61

$$A^{(u_1)} = \begin{pmatrix} \langle u_1 | \frac{du_1}{d\theta} \rangle \\ \langle u_1 | \frac{du_1}{d\phi} \rangle \end{pmatrix}; \quad A^{(u_2)} = \begin{pmatrix} \langle u_2 | \frac{du_2}{d\theta} \rangle \\ \langle u_2 | \frac{du_2}{d\phi} \rangle \end{pmatrix} \quad (4.4)$$

and geometric tensor using 2.16 we have

$$\chi_{ij}^{(u_k)} = \langle \partial_i u_k | \partial_j u_k \rangle - \langle \partial_i u_k | u_k \rangle \langle u_k | \partial_j u_k \rangle \quad (4.5)$$

implying

$$\chi^{(u_1)} = \begin{pmatrix} \frac{1}{4} & \frac{i}{4} \sin \theta \\ -\frac{i}{4} \sin \theta & \frac{\sin^2 \theta}{4} \end{pmatrix}; \quad \chi^{(u_2)} = \begin{pmatrix} \frac{1}{4} & -\frac{i}{4} \sin \theta \\ \frac{i}{4} \sin \theta & \frac{\sin^2 \theta}{4} \end{pmatrix}. \quad (4.6)$$

It's real part is according to 2.21 the metric tensor

$$g^{(u_1)} = g^{(u_2)} = \begin{pmatrix} \frac{1}{4} & 0 \\ 0 & \frac{\sin^2 \theta}{4} \end{pmatrix} \quad (4.7)$$

For two state system, the metric tensor is always the same for both states. (just guessing now :). Berry curvature can be easily calculated as minus imaginary part of the geometric tensor.

From the metric the Ricci curvature and Gaussian curvature are

$$R = \begin{pmatrix} 1 & 0 \\ 0 & \frac{1}{4} \sin^2 \theta \end{pmatrix} \quad (4.8)$$

$$K = \frac{1}{8}(1 + \sin^4 \theta) \quad (4.9)$$

If the magnetic field is positive, u_1 is the ground state. The metric on \mathcal{M}_0 is then $g^{(u_1)}$ calculated above. During some arbitrary change of magnetic field whilst not measuring the state of the system, it can be transported from $|\psi_i\rangle := |u_1\rangle$ to some new ground state by changing the magnetic field direction, as shown on fig. 4.1

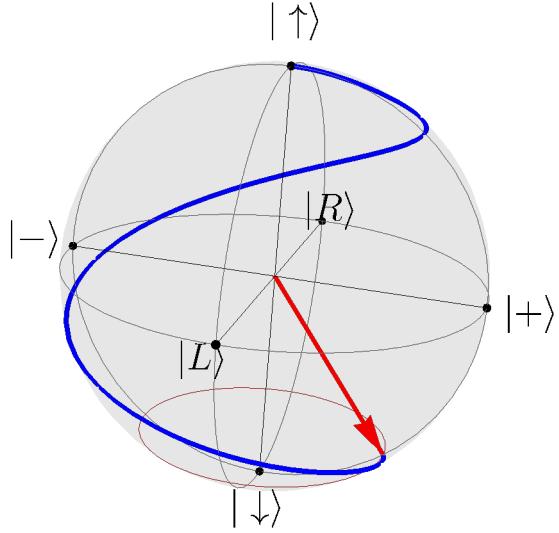


Figure 4.1: Bloch sphere for the eigenstate u_1 with the initial direction of magnetic field $\mathbf{B} = (0, 0, B_z)$ transported along $\gamma = \{(\phi, \sqrt{\phi}) | \phi \in [0, 2\pi]\}$ (blue). The final orientation of the magnetic field is marked red and the states on the Bloch sphere correspond to initial eigenstate.

It's pedagogical to point out the correspondence to figure 2.4, where $\mathcal{M}_0 = \cup_{\lambda} \{|u_1\rangle\}$, $\mathcal{M}_1 = \cup_{\lambda} \{|u_2\rangle\}$, $|o(\lambda)\rangle = u_1(\mathbf{B} = B(0, 0))$ and $|o(\tilde{\lambda})\rangle = u_1(\mathbf{B} = B(2\pi, \sqrt{2\pi}))$, written in spherical coordinates.

If the transport is made infinitesimally slow with respect to time t , after every step $d\theta$ the eigensystem has to be calculated, and we find out that the system collapses into instantaneous $u_1(t)$ state of $\mathcal{H}(t)$, so the fidelity at any θ is $f_0(\theta) = 1$ and $ds^2(\theta) = 1 - f_0^2 = 0$. Opposite to the adiabatic transport would be the quantum quench, meaning the orientation of \mathbf{B} is changed instantly resulting in a final state

$$|f\rangle = \alpha |u_1\rangle + \beta |u_2\rangle, \quad \text{for } \frac{\alpha}{\beta} = e^{-i\phi} \tan \frac{\theta}{2}, \quad \alpha^2 + \beta^2 = 1. \quad (4.10)$$

This yields probability of staying in the state u_1 after quench is $f_q = \langle u_1 | f \rangle = \alpha$ and $s = 1 - \alpha$.

4.1 Special case

Assume again transport of $|u_1\rangle$ by rotating the magnetic field along the path $\gamma = \{(0, \theta) | \theta \in [0, \pi/2]\}$, as drawn on figure 4.2. The fidelity for infinitesimal speed is at every point on the path $f_0 = 1$ and for quantum quench $f_q = 1/2$.

Now let's calculate the fidelity for some finite time transport.

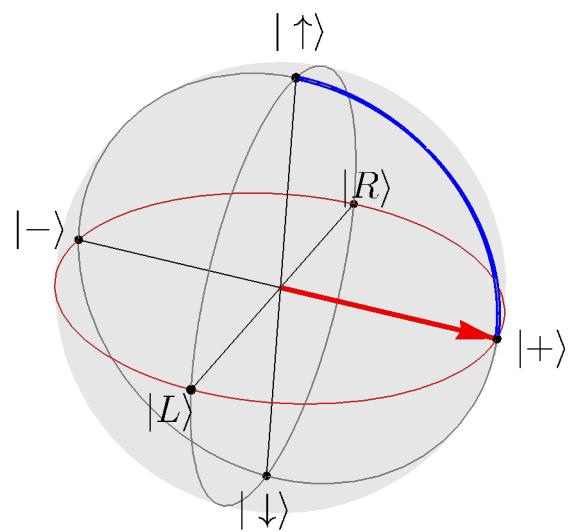


Figure 4.2: Bloch sphere for the eigenstate u_1 with the initial direction of magnetic field $\mathbf{B} = (0, 0, B_z)$ transported along $\gamma = \{(0, \theta) | \theta \in [0, \pi/2]\}$ (blue). The final orientation of the magnetic field is marked red and the states on the Bloch sphere correspond to initial eigenstate.

5. First order transition Hamiltonian

First order transition Hamiltonian can be written in a form

$$\hat{H} = \hat{J}_3 - \frac{\tilde{\lambda}(t)}{2j} (\hat{J}_1 + \tilde{\chi}(\hat{J}_3 + j\mathbb{1}))^2 \quad (5.1)$$

for $\hat{J} = (\hat{J}_1, \hat{J}_2, \hat{J}_3)$ angular momentum operator, $\tilde{\chi} \in \mathbb{R}$ and $\tilde{\lambda}(t)$ some real time dependent parameter. The aim is to diagonalize this Hamiltonian to find the spectrum. To make driving parameters less entwined, we will instead use the Hamiltonian

$$\hat{H} = \hat{J}_3 + \lambda \hat{V}_1 + \chi \hat{V}_2 + \chi^2 \hat{V}_3, \quad (5.2)$$

for

$$\hat{V}_1 = -\frac{1}{2j} \hat{J}_1^2 \quad (5.3)$$

$$\hat{V}_2 = -\frac{1}{2j} [\hat{J}_1(\hat{J}_3 + j\mathbb{1}) + (\hat{J}_3 + j\mathbb{1})\hat{J}_1] \quad (5.4)$$

$$\hat{V}_3 = -\frac{1}{2j} (\hat{J}_3 + j\mathbb{1})^2 \quad (5.5)$$

Using the eigenbasis $\{|j, m\rangle\}$ for j angular momentum quantum number and m its projection to the direction \hat{J}_3 and defining

$$\hat{J}_{\pm} := \frac{1}{2}(\hat{J}_1 \pm i\hat{J}_2), \quad (5.6)$$

we get matrix elements

$$\langle j'm' | \hat{J}^2 | jm \rangle = j(j+1)\delta_{j'j}\delta_{m'm} \quad (5.7)$$

$$\langle j'm' | \hat{J}_3 | jm \rangle = m\delta_{j'j}\delta_{m'm} \quad (5.8)$$

$$\langle j'm' | \hat{J}_{\pm} | jm \rangle = \sqrt{(j \mp m)(j \pm m + 1)}\delta_{j'j}\delta_{m'm\pm1}, \quad (5.9)$$

where $\delta_{a'b}$ is Kronecker delta. Using

$$(\hat{J}_1 + \chi(\hat{J}_3 + j))^2 = \hat{J}_1^2 + \chi^2(\hat{J}_3^2 + j^2\mathbb{1} + 2j\hat{J}_3) + \chi(\hat{J}_1\hat{J}_3 + \hat{J}_3\hat{J}_1) + 2\chi j \hat{J}_1 \quad (5.10)$$

$$\hat{J}_1^2 = \frac{1}{4}(\hat{J}_+ + \hat{J}_-)^2 = \frac{1}{4}(\hat{J}_+^2 + \hat{J}_-^2 + \hat{J}_+\hat{J}_- + \hat{J}_-\hat{J}_+) \quad (5.11)$$

$$\hat{J}_{\pm}\hat{J}_{\mp} = \hat{J}^2 - \hat{J}_3^2 \mp \hat{J}_3 \quad (5.12)$$

$$\hat{J}_1 = (\hat{J}_+ + \hat{J}_-), \quad (5.13)$$

we get pentadiagonal matrix representation of \hat{H} .

5.1 Hamiltonian analysis

The behavior of cases $N = 1$ and $N = 2$ is not characteristic for our Hamiltonian and has no physical meaning in this case, therefore the properties of $N = 3$ case will be discussed at first and afterwards generalized to arbitrary N along with the limit $N \rightarrow \infty$. Due to complexity of chosen Hamiltonian, only some aspects are possible to calculate analytically, but mostly the numerical analysis is needed.

5.1.1 Case N=3

The lowest dimension with characteristics common to higher N is 3 with Hamiltonian in eigenbasis

$$\hat{H} = \begin{pmatrix} -\frac{\lambda+6}{4} & -\frac{\chi}{2\sqrt{3}} & -\frac{\lambda}{2\sqrt{3}} & 0 \\ -\frac{\chi}{2\sqrt{3}} & \frac{(-7\lambda-4\chi^2-6)}{12} & -\chi & -\frac{\lambda}{2\sqrt{3}} \\ -\frac{\lambda}{2\sqrt{3}} & -\chi & \frac{(-7\lambda-16\chi^2+6)}{12} & -\frac{5\chi}{2\sqrt{3}} \\ 0 & -\frac{\lambda}{2\sqrt{3}} & -\frac{5\chi}{2\sqrt{3}} & -\frac{\lambda}{4} - 3\chi^2 + \frac{3}{2} \end{pmatrix}. \quad (5.14)$$

Spectrum of this Hamiltonian can be written in analytically using some substitutions A, B, C, D , see Appendix A, as

$$E_0 = \frac{1}{12} \left(G - F - \frac{\sqrt{D - E}}{2} \right) \quad (5.15)$$

$$E_1 = \frac{1}{12} \left(G - F + \frac{\sqrt{D - E}}{2} \right) \quad (5.16)$$

$$E_2 = \frac{1}{12} \left(G + F - \frac{\sqrt{D + E}}{2} \right) \quad (5.17)$$

$$E_3 = \frac{1}{12} \left(G + F + \frac{\sqrt{D + E}}{2} \right). \quad (5.18)$$

Eigenvectors can also be written analytically, but due to their compexity its better to analize them only numerically. Sections $\lambda = 1$ and $\chi = 1$ are shown in figures 5.1,5.2

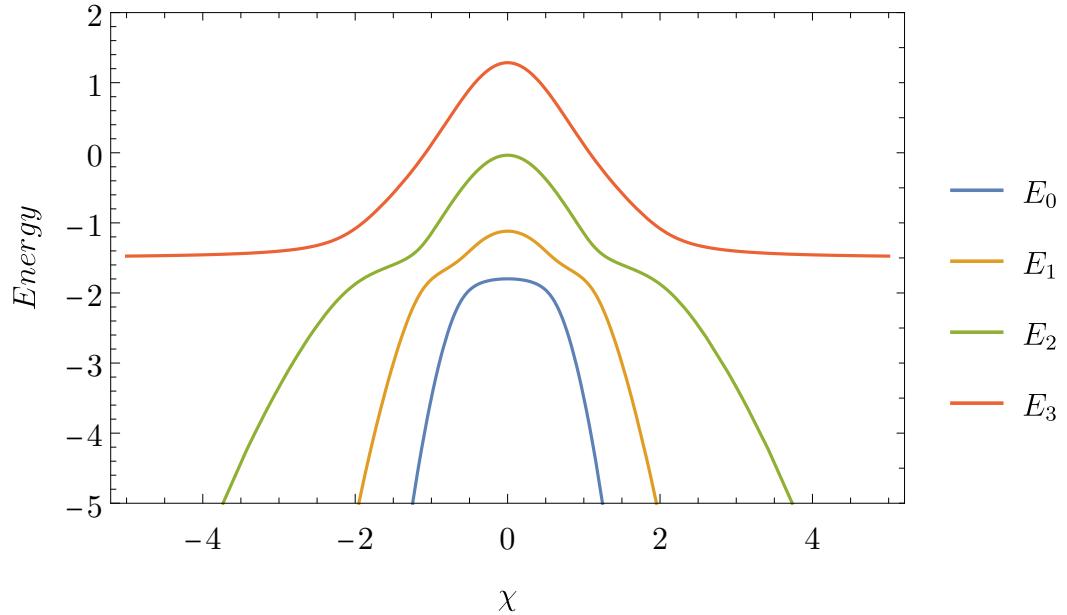


Figure 5.1: Energy for the case $N = 3$ Hamiltonian, section $\lambda = 1$

From equations 5.15, 5.16 can be seen that $E_0 = E_1$ for $D = E$, which for real values λ, χ has two solutions

$$(\lambda_d, \pm \chi_d) = \left(-\frac{1}{2}, \pm \sqrt{\frac{3}{5}} \right).$$

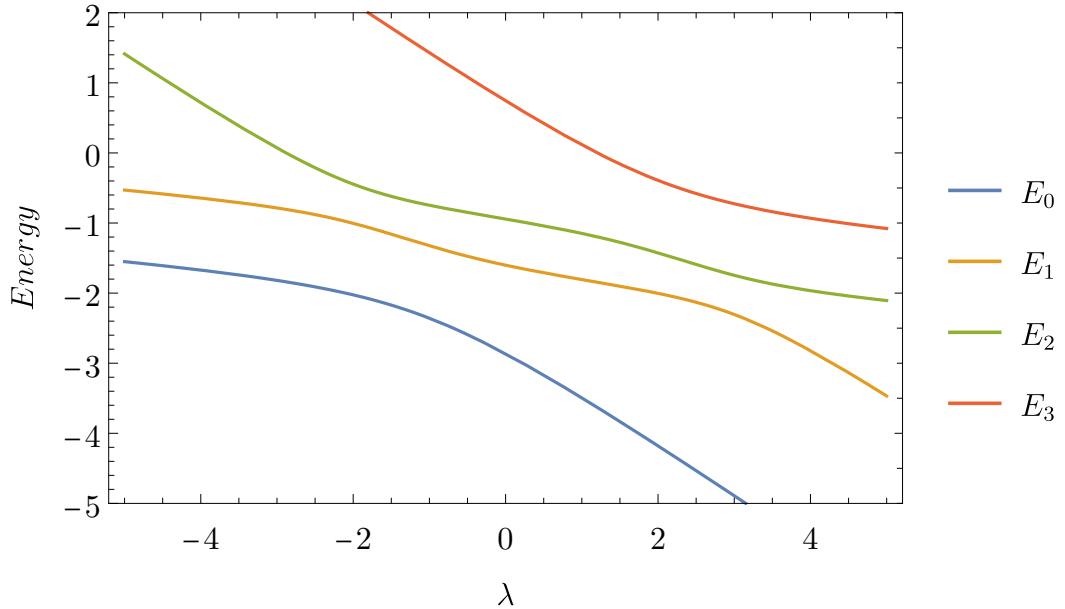


Figure 5.2: Energy for the case $N = 3$ Hamiltonian, section $\chi = 1$

Point-like characteristics corresponds to Theorem 1, which states that Hamiltonian driven by two real parameters can be degenerated only on 0-dimensional manifolds.

If the energy spectrum is degenerate and metric tensor diverges, see individual elements on fig. 5.3, its determinant also diverges, as shown on fig. 5.4, along with Christoffel symbols elements, see fig. 5.5. Note that metric tensor determinant is positive definite, thus the manifold is Riemannian. Further on it reflects symmetry $\chi \leftrightarrow -\chi$, except for off-diagonal elements, i.e. $g_{12}, \Gamma_{121}, \Gamma_{211}, \Gamma_{222}$, which switches their sign.

Due to metric tensor degeneracy, the space is not geodesically complete and according to 3 there exist some "event horizon" around singularities $(\lambda_d, \pm\chi_d)$. This can be seen more clearly from geodesics, i.e. by solving initial value problem with conditions

$$(\lambda(t_i); \chi(t_i)) = (\lambda_i; \chi_i)$$

$$\left(\frac{d\lambda(t)}{dt}; \frac{d\chi(t)}{dt} \right) \Big|_{t_i} = (\lambda'_i; \chi'_i).$$

Results for these geodesics starting at $(\lambda; \chi) = (0; 0), (\lambda', \chi') = (\cos \theta; \sin \theta)$ for $\theta \in [-0.63; 0.63]$ with step 0.01. Above $\theta = 0.63$ the geodesics get too close to singularity making calculations numerically unstable. The same holds for interval $\theta \in [2.43; 2\pi - 2.43]$, see fig. 5.6. We see, that geodesics tend to get closer to the singularity. This is well known fact from the General Relativity (GR) with one important difference, that our "test particle" is repulsed from the singularity, not attracted to it. Therefore our intuition about black holes in GR will be broken, as can be seen in fig. 5.7. Instead, we can think of it as two oppositely charged monopoles in the theory of electromagnetism. **Because Hamiltonian consists of more elements, we can check, if they all contribute repulsively, or some of them are attractive and their relative strength to other elements just drops close to the**

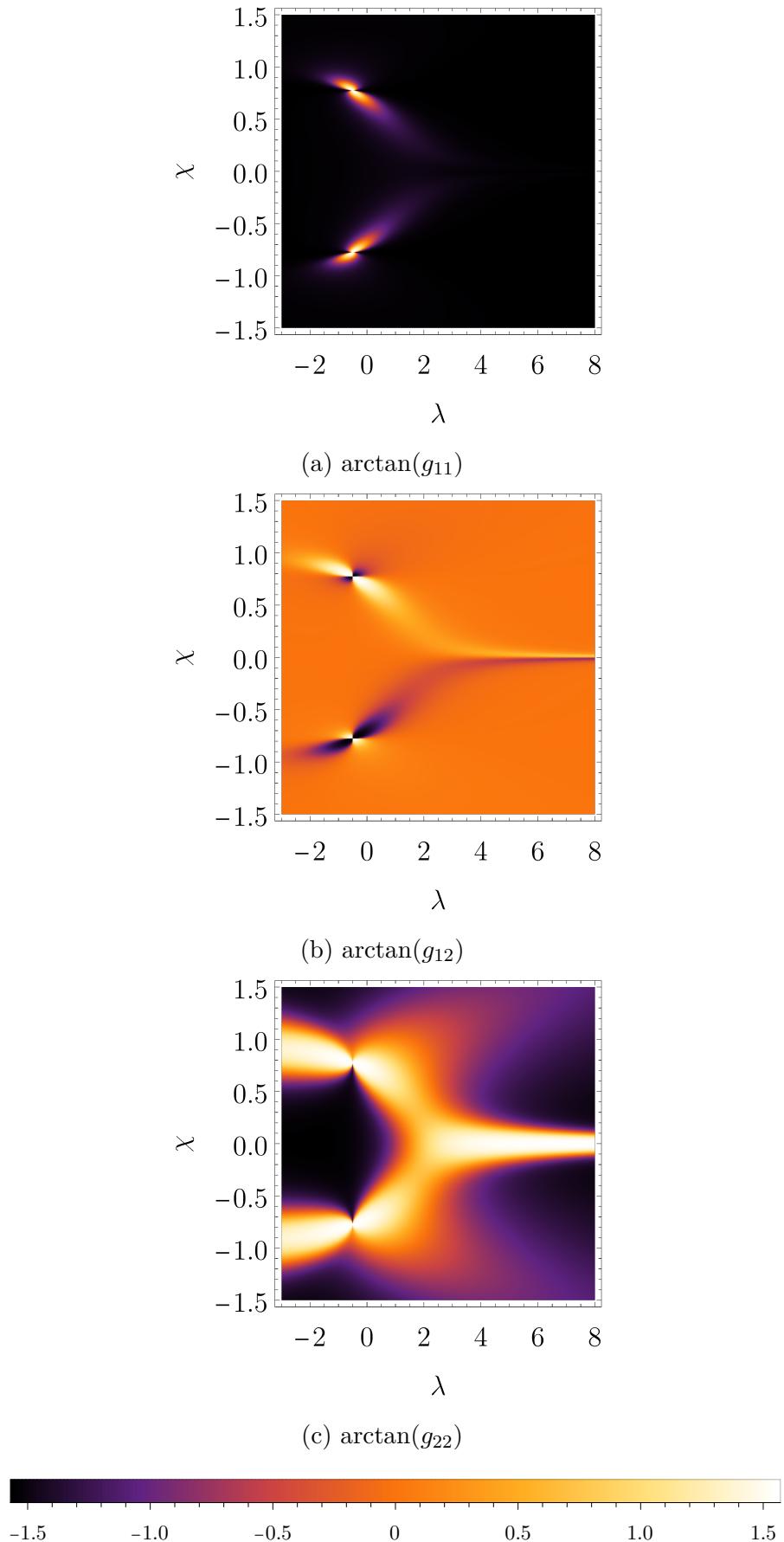


Figure 5.3: Arcustangens of metric tensor elements for the case $N = 3$

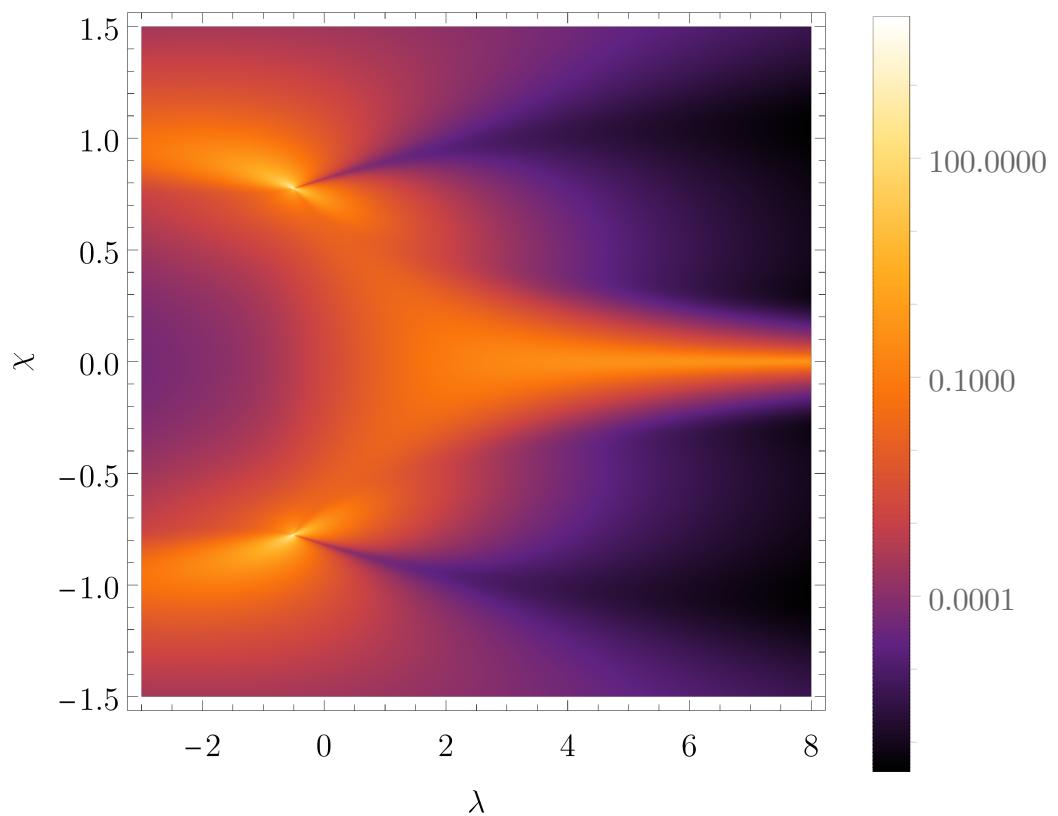


Figure 5.4: Metric determinant in a parameter space for $N = 3$.

singularity.

5.2 Arbitrary N

Looking at the energy spectrum of $N = 10$ case (see section for $\lambda = 1$ on fig. 5.8 and section $\chi = 1$ in fig. 5.9) we see the characteristic structure of the system. Between every two energy levels is at least one avoided crossing. Another interesting aspect are metric tensor singularities, which are moving along [line](#), which was using linear regression from datapoints in fig. 5.10 approximated as

$$\chi = -0.14\lambda + 0.71 \quad (5.19)$$

5.3 Limit $N \rightarrow \infty$

In the limit $N \rightarrow \infty$, the first order transition lies in curve

$$\chi^2 = \frac{\lambda - 1}{\lambda - 2} \quad (5.20)$$

and χ -axis to the right from the point of second order transition, which lies in $(\lambda, \chi) = (1, 0)$. On figure 5.11 can be seen the transition curve with comparison to minimum between the ground state and first excited state for $N = 3$ case. With increasing N , the line is slowly converging to the transition line.

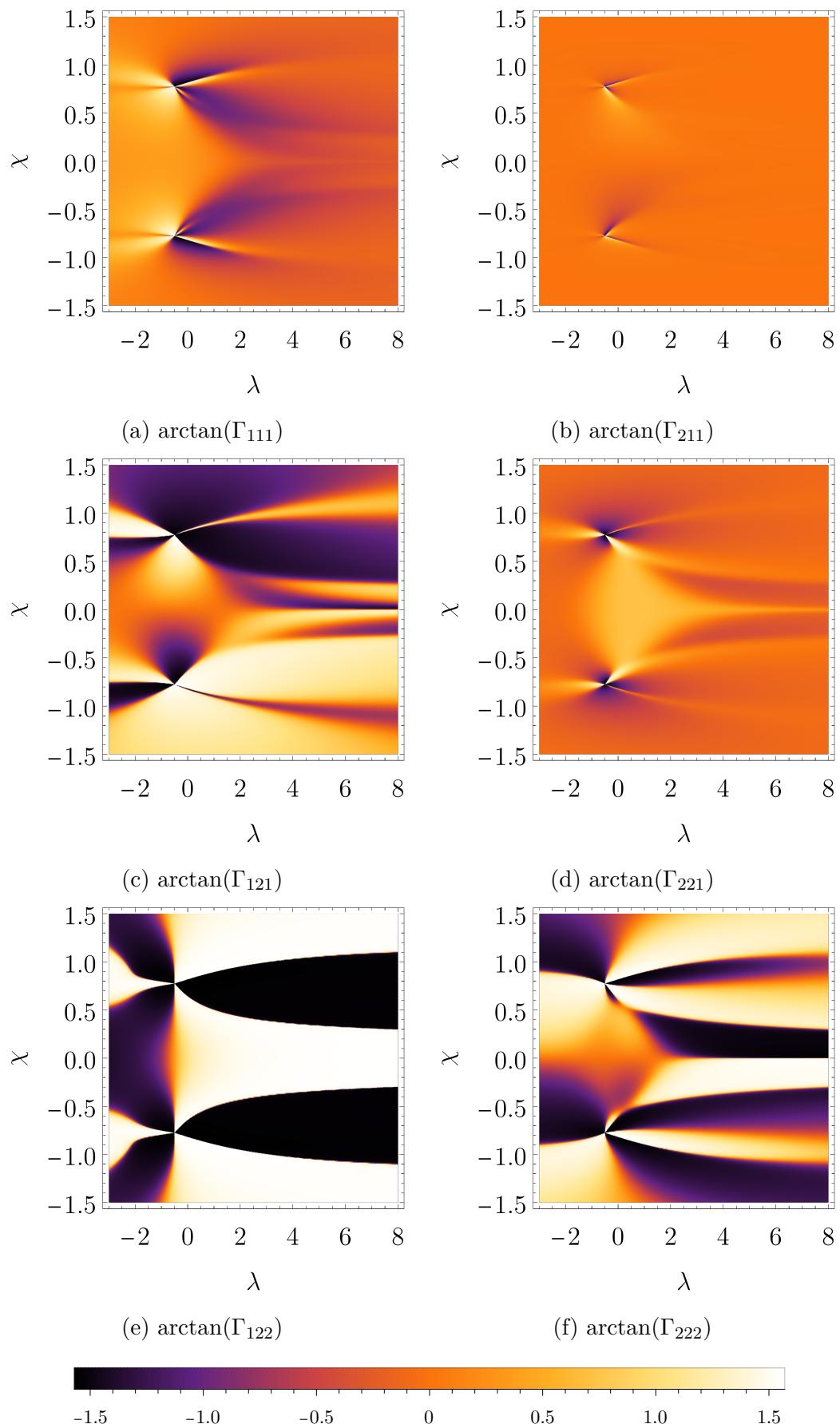


Figure 5.5: Arcustangens of Christoffel symbols for the case $N = 3$

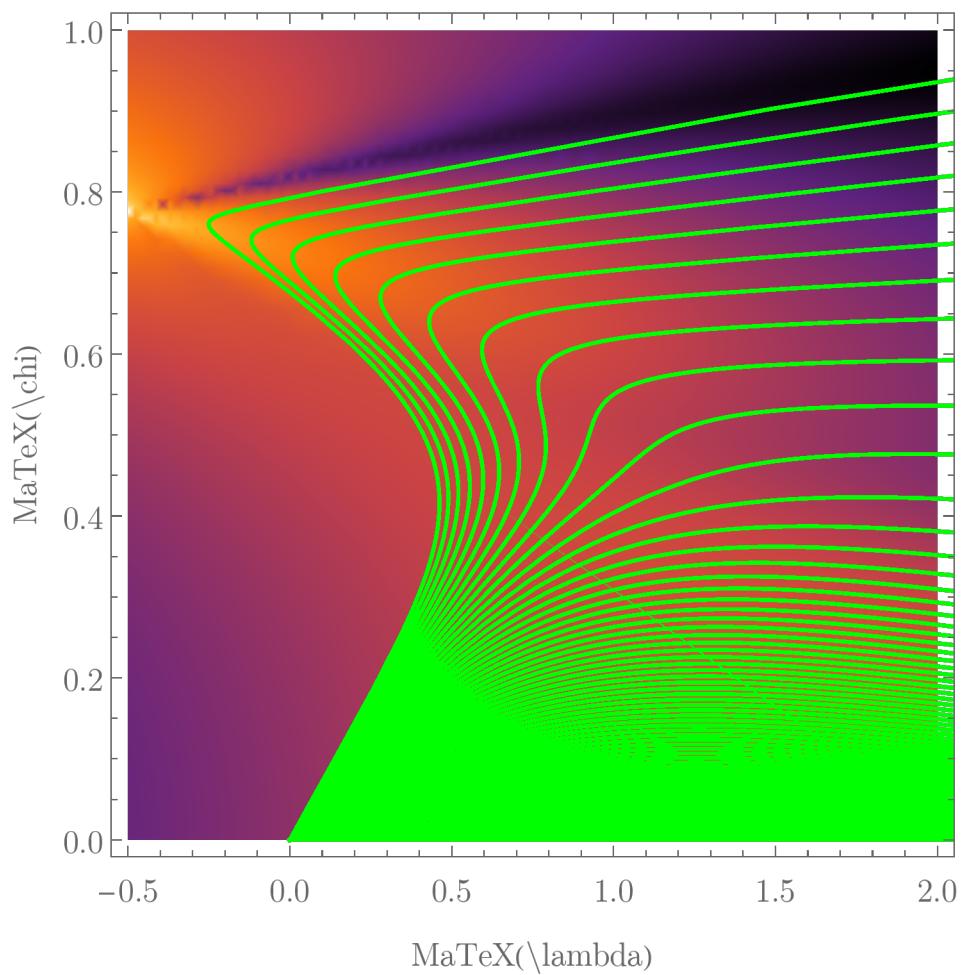


Figure 5.6: Geodesics starting from $(\lambda_i; \chi_i) = (0; 0)$ with $(\lambda'_i; \chi'_i) = (\cos \theta; \sin \theta)$
 $N = 3$.

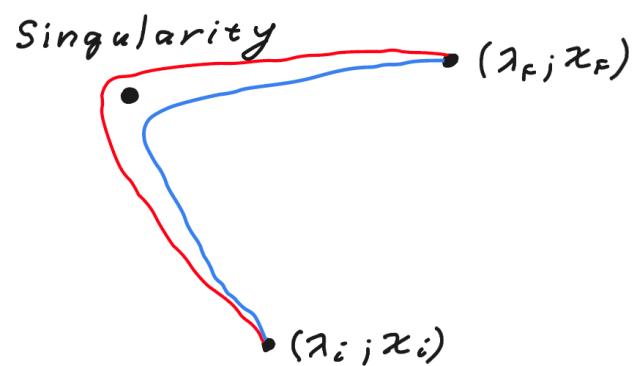


Figure 5.7: Comparing geodesics with repulsing (blue) and attracting (red) metric tensor divergence in the spherical symmetrical space.

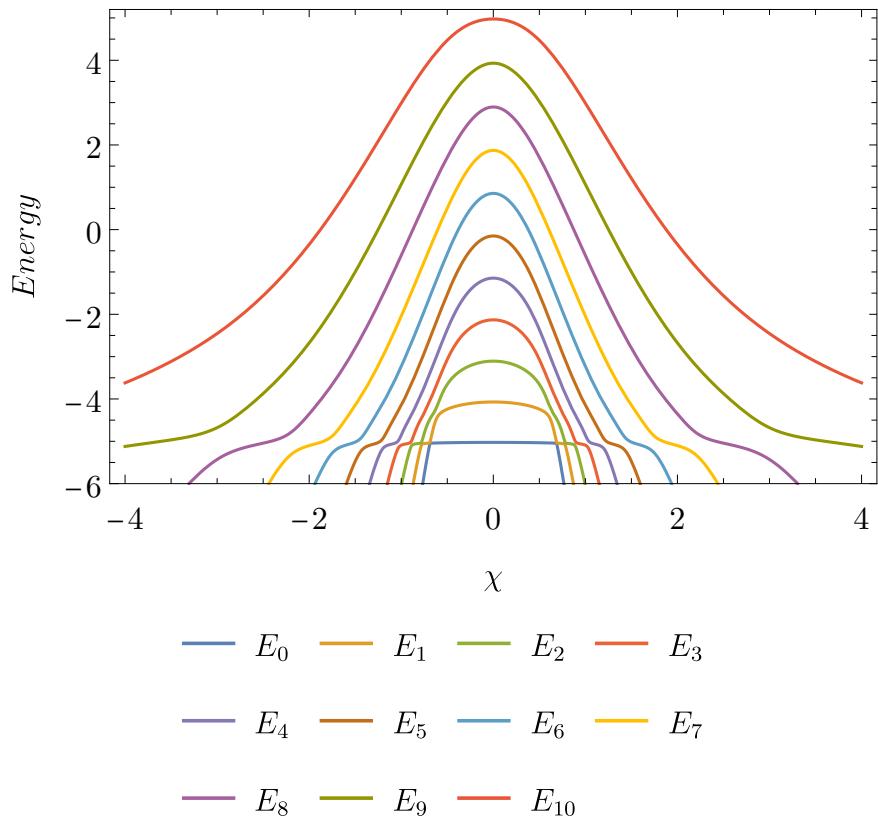


Figure 5.8: Energy spectrum as function of χ , for $\lambda = 1$ and $N = 10$.

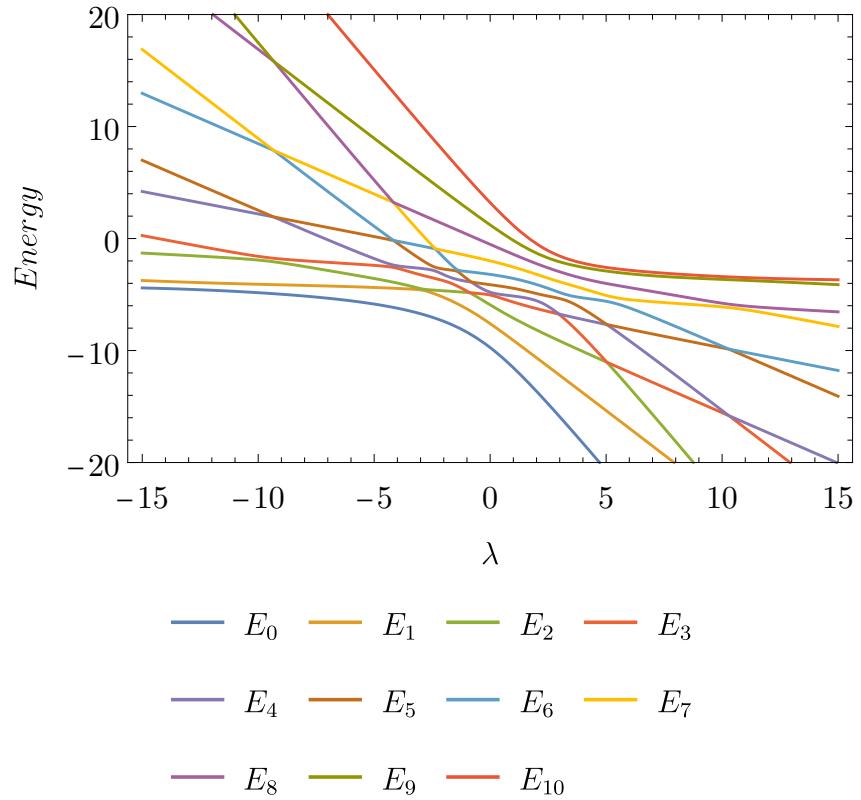


Figure 5.9: Energy spectrum as function of λ , for $\chi = 1$ and $N = 10$.

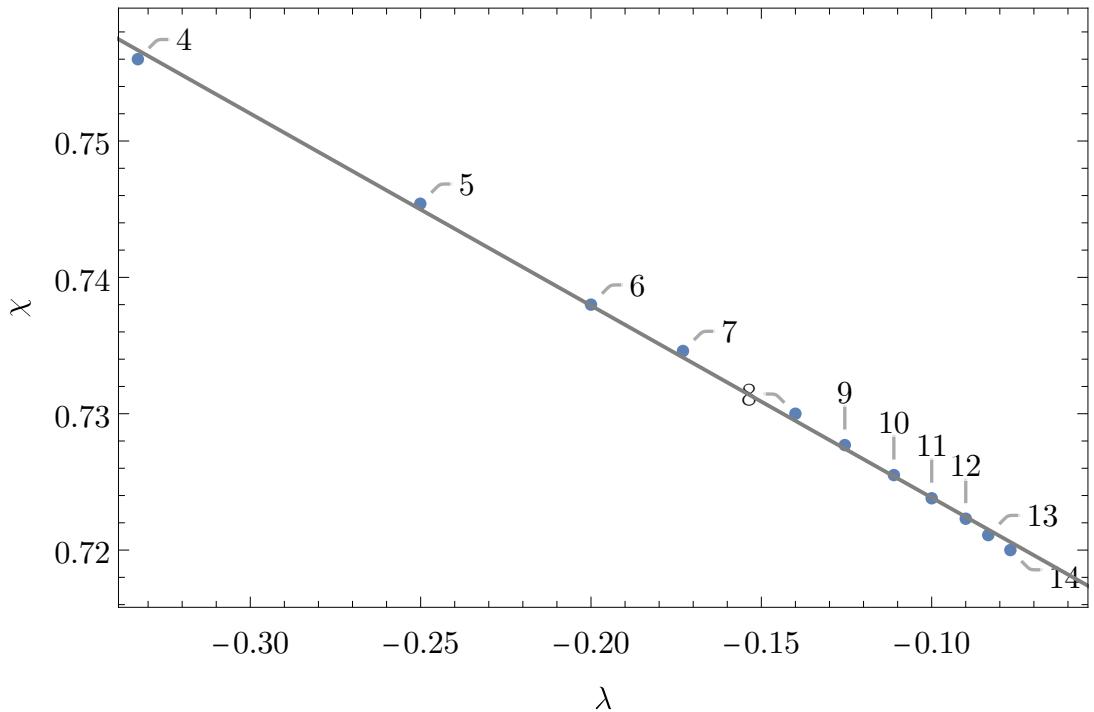


Figure 5.10: Singularities in the metric tensor for different values of N with linear regression according to eq. 5.19. Green points mark minimums.

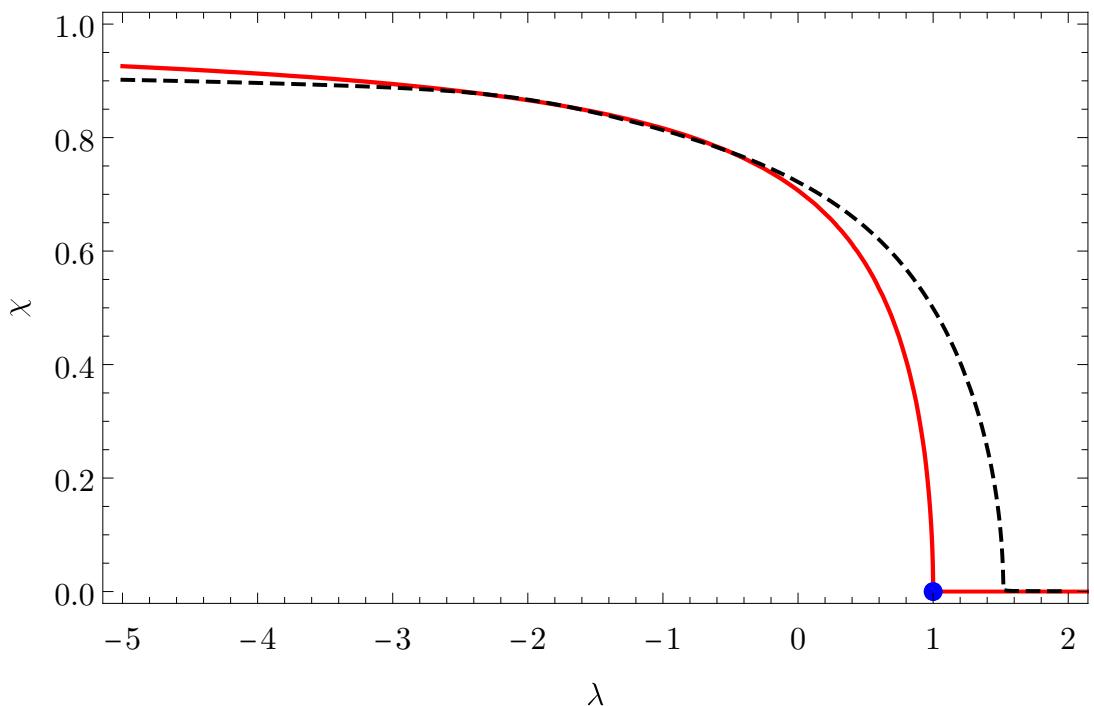


Figure 5.11: First order transition (red), second order transition (blue point), minimum between the ground state and first excited state in $N=3$ case (black, dashed).

Conclusion

Bibliography

- M. V. Berry. Quantal phase factors accompanying adiabatic changes. *Proc. R. Soc. Lond. A*, 392(1802):45–57, March 1984. URL <http://rspa.royalsocietypublishing.org/content/392/1802/45>.
- M V Berry. Transitionless quantum driving. *Journal of Physics A: Mathematical and Theoretical*, 42(36):365303, aug 2009. doi: 10.1088/1751-8113/42/36/365303. URL <https://doi.org/10.1088/1751-8113/42/36/365303>.
- Michael V. Berry. THE QUANTUM PHASE, FIVE YEARS AFTER. 1989.
- Marián Fecko. *Differential Geometry and Lie Groups for Physicists*. Cambridge University Press, 2006. doi: 10.1017/CBO9780511755590.
- C. Gorodski. *An introduction to Riemannian geometry*. 2012. Preliminary version.
- Michael Kolodrubetz, Dries Sels, Pankaj Mehta, and Anatoli Polkovnikov. Geometry and non-adiabatic response in quantum and classical systems. *Physics Reports*, 697:1–87, 2017. ISSN 0370-1573. doi: <https://doi.org/10.1016/j.physrep.2017.07.001>. URL <https://www.sciencedirect.com/science/article/pii/S0370157317301989>. Geometry and non-adiabatic response in quantum and classical systems.
- W. Tu Loring. *Differential Geometry*. Springer International Publishing, 2017. doi: 10.1007/978-3-319-55084-8.
- P. Petersen. *Riemannian Geometry*. Springer Verlag, 1998. ISBN 978-0-387-29403-2.

List of Figures

1.1	Mappings needed for the bundle definition.	6
2.1	Fiber bundle over \mathbb{R}^n with Hilbert spaces $\mathcal{H}(\lambda)$ as individual fibers.	11
2.2	Geometrical intuition to transport on fiber manifold sections \mathcal{M}_i .	12
2.3	Space of bare states and its projection to space of rays $\mathcal{PH}(\lambda) := \mathcal{H}(\lambda/U(1))$	12
2.4	Geometrical intuition to transport on fiber manifold sections \mathcal{M}_s .	13
2.5	Parallel transporting around some closed curve C with anholonomy drawn in yellow.	14
4.1	Bloch sphere for the eigenstate u_1 with the initial direction of magnetic field $\mathbf{B} = (0, 0, B_z)$ transported along $\gamma = \{(\phi, \sqrt{\phi}) \phi \in [0, 2\pi]\}$ (blue). The final orientation of the magnetic field is marked red and the states on the Bloch sphere correspond to initial eigenstate.	27
4.2	Bloch sphere for the eigenstate u_1 with the initial direction of magnetic field $\mathbf{B} = (0, 0, B_z)$ transported along $\gamma = \{(0, \theta) \theta \in [0, \pi/2]\}$ (blue). The final orientation of the magnetic field is marked red and the states on the Bloch sphere correspond to initial eigenstate.	28
5.1	Energy for the case $N = 3$ Hamiltonian, section $\lambda = 1$	30
5.2	Energy for the case $N = 3$ Hamiltonian, section $\chi = 1$	31
5.3	Arcustangens of metric tensor elements for the case $N = 3$	32
5.4	Metric determinant in a parameter space for $N = 3$	33
5.5	Arcustangens of Christoffel symbols for the case $N = 3$	35
5.6	Geodesics starting from $(\lambda_i; \chi_i) = (0; 0)$ with $(\lambda'_i; \chi'_i) = (\cos \theta; \sin \theta)$ $N = 3$	36
5.7	Comparing geodesics with repulsing (blue) and attracting (red) metric tensor divergence in the spherical symmetrical space.	36
5.8	Energy spectrum as function of χ , for $\lambda = 1$ and $N = 10$	37
5.9	Energy spectrum as function of λ , for $\chi = 1$ and $N = 10$	37
5.10	Singularities in the metric tensor for different values of N with linear regression according to eq. 5.19. Green points mark minimums.	38
5.11	First order transition (red), second order transition (blue point), minimum between the ground state and first excited state in $N=3$ case (black, dashed).	38

A. Appendix 1

$$\begin{aligned}
A = & 16\sqrt[3]{2} \left(64\lambda^6 - 192\lambda^5\chi^2 + 24\lambda^4 (52\chi^4 - 93\chi^2 + 36) \right. \\
& - 8\lambda^3 (29\chi^4 + 414\chi^2 - 513) \chi^2 \\
& + 6\lambda^2 (1225\chi^8 - 10053\chi^6 + 17595\chi^4 - 10557\chi^2 + 1377) \\
& + \left((64\lambda^6 + 864\lambda^4 + 8262\lambda^2 + 3(818\lambda - 27285)\chi^{10} \right. \\
& + 6(\lambda(1225\lambda - 3198) + 27108)\chi^8 \\
& - (2\lambda(\lambda(116\lambda + 30159) - 89073) + 326727)\chi^6 \\
& + 6(\lambda(\lambda(8\lambda(26\lambda - 69) + 17595) - 42660) + 51516)\chi^4 \\
& - 3(2\lambda(\lambda(4\lambda(\lambda(8\lambda + 93) - 171) + 10557) - 16119) + 36207)\chi^2 \quad (A.1) \\
& + 24013\chi^{12} + 25515) \\
& - \left. (16\lambda^4 + 144\lambda^2 + 2(74\lambda - 1185)\chi^6 + 3(4\lambda(16\lambda - 77) + 1329)\chi^4 \right. \\
& - 2(2\lambda(\lambda(8\lambda + 93) - 207) + 1161)\chi^2 + 889\chi^8 + 1053) \right)^{1/2} \\
& + 6\lambda(409\chi^8 - 3198\chi^6 + 29691\chi^4 - 42660\chi^2 + 16119)\chi^2 + 24013\chi^{12} \\
& - 81855\chi^{10} + 162648\chi^8 - 326727\chi^6 + 309096\chi^4 \\
& \left. - 108621\chi^2 + 25515 \right)^{1/3}
\end{aligned}$$

$$\begin{aligned}
B = & \frac{256\sqrt[3]{2}}{3A} \left(16\lambda^4 + 144\lambda^2 + 2(74\lambda - 1185)\chi^6 + 3(4\lambda(16\lambda - 77) \right. \quad (A.2) \\
& \left. + 1329)\chi^4 - 2(2\lambda(\lambda(8\lambda + 93) - 207) + 1161)\chi^2 + 889\chi^8 + 1053 \right)
\end{aligned}$$

$$C = 4\sqrt{\frac{A}{3\sqrt[3]{2}} + B + \frac{16}{3}(4\lambda^2 - (4\lambda + 33)\chi^2 + 49\chi^4 + 45)} \quad (A.3)$$

$$\begin{aligned}
D = & -\frac{A}{3\sqrt[3]{2}} - B - \frac{8}{3}(59\lambda^2 + 436\lambda\chi^2 + 392\chi^4 + 132\chi^2 - 180) \quad (A.4) \\
& + 8(5\lambda + 14\chi^2)^2
\end{aligned}$$

$$\begin{aligned}
E = & \frac{9216}{C} ((\lambda - 1)\chi^4 - 4(\lambda - 1)\chi^2 + \lambda - 2\chi^6) \\
F = & \frac{1}{2}\sqrt{\frac{A}{3\sqrt[3]{2}} + B + \frac{16}{3}(4\lambda^2 - (4\lambda + 33)\chi^2 + 49\chi^4 + 45)} \quad (A.5) \\
G = & -5\lambda - 14\chi^2
\end{aligned}$$

B. Attachments

B.1 First Attachment