

2/13/23

Coordinate Systems

- 2 forms of coordinate systems
  - 1) Inertial / Fixed
  - 2) Body Frame

ECEF (Earth Centered Earth Fixed) "global"

- cartesian ( $x, y, z$ )
- Geocentric Polar ( $r, \theta, \lambda$ )
- spheroidal (ellipsoidal or geodetic) ( $\phi, \lambda, h$ )

East - North - Up "Local"

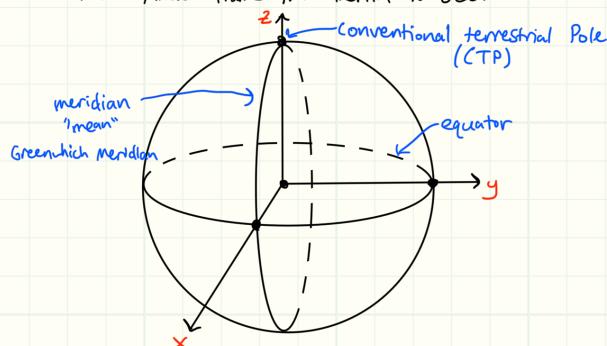
- Plane tangent to the center of the earth
- requires a reference point

>> lladeref  
 >> eceflla  
 >> lladenu  
 >> enudlla

} textbook appendix

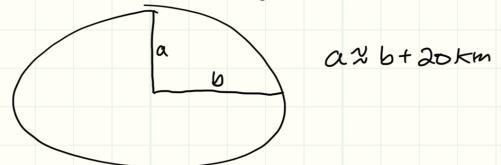
GPS Coordinate System

- needed global coordinate system
  - ECEF  $\rightarrow$  fixed to center of earth
    - $\hookrightarrow$  EARTH IS NOT RIGID (CHANGES SHAPE)
    - $\hookrightarrow$  Not elliptical (bulges at equator, flatter at poles)
    - $\hookrightarrow$  Not uniform
  - Use sea level as an approximate shape
  - Satellites + user to use same coordinate system
    - $\hookrightarrow$  user frame fixed to earth (not moving from user standpoint)
    - $\hookrightarrow$  SV in inertial frame (fixed in space)
      - \* Rotate/move frame from inertial to ECEF

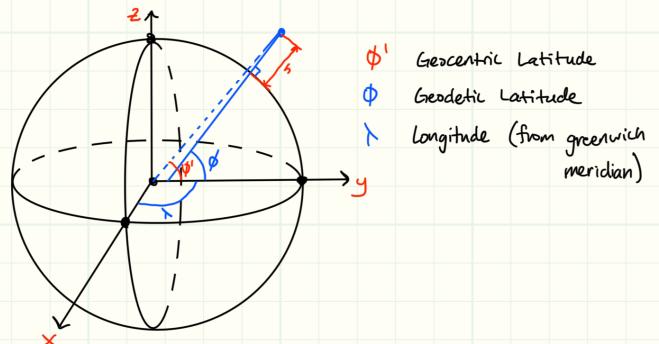
World Geodetic System

- 1984, DoD developed WGS84
  - Not accurate, but globally used
  - Not intuitive and not on earth's surface
- ECEF in Latitude, Longitude, Height / Altitude
  - $\hookrightarrow$  relative to hypothetical surface
- Geoid: surface consisting of the mean sea level (MSL) and under land if reached by small, frictionless channels

- geoid is non-even (mass distribution of earth is not equal)
- difficult to calculate the perfect geoid
- Approximated (curve-fit) by a spheroid or ellipsoid



$$\begin{aligned} \text{- eccentricity } &\Rightarrow e^2 = \frac{a^2 - b^2}{a^2} \\ \text{- flattening } &\Rightarrow f = \frac{a - b}{a} \end{aligned} \quad \left. \begin{array}{l} e^2 = 2f - f^2 \\ \hline \end{array} \right\}$$

WGS84

- \* Not inertial frame
- \* Datum is ellipsoid to model geoid
- \* origin is earth's center of mass

$$a = 6,378,137.0 \text{ m}$$

$$b = 6,356,752 \text{ m}$$

- \* Different datums produce different coordinates for same point
- \* All points of constant φ are circular
- \* All points of constant λ are elliptical
- \* 1° of λ is different distance depending on φ
  - $\approx 110 \text{ km}$  at equator
  - $\approx 80 \text{ km}$  at  $45^\circ \phi$
- \* Go between ECEF and ellipsoid with model of ellipsoid (i.e. WGS84)

$$a = 6,378,137.0 \text{ m}$$

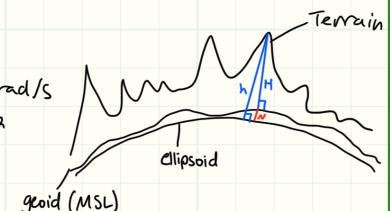
$$1/f = 298.257223563$$

$$\omega_{\text{earth}} = 7.292115 (10^{-5}) \text{ rad/s}$$

$$g_c = 39,860,004.418 \text{ m}^3/\text{s}^2$$

$$C = 299,792,458 \text{ m/s}$$

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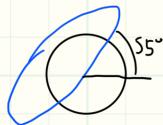
- GPS Provides "h", height above ellipsoid

- WGS84 provides contours of the geoid with respect to ellipsoid (N)

2/15/23

Satellite Orbits

- 24+ SVs (24 minimum spec. for fully functional)
- 6 orbital planes
  - 55° inclination angles (less coverage at poles)
    - optimize user coverage
  - approximately circular
  - 12 hr orbits
    - SV position repeats ~ every 23.56 hours
    - 20,162 km from equator (26,561 from center of the earth)
    - Travel at ~ 2.7 km/s

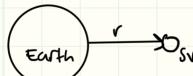


$$\text{Sidereal Time: } 24 \text{ sidereal hours} = 1.002738 \times 24 \text{ solar hours}$$

Newton vs. Kepler

$$M_E = 3,986,004,418 \frac{\text{m}^3}{\text{s}^2}$$

$$G = 6.674 \times 10^{-1} \frac{\text{m}^3}{\text{kg s}^2}$$



$$F = \frac{GM_E m_{SV}}{r^2} = \frac{M_E m_{SV}}{r^2}$$

Using Newton's Laws:  $\sum F = m \ddot{r}$ 

- SV:  $m_{SV} \ddot{r}_S = -\frac{GM_E m_{SV}}{r^2} \cdot \frac{\vec{r}}{r} = -\frac{GM_E m_{SV}}{r^3} \vec{r}$

- EARTH:  $M_E \ddot{r}_E = \frac{GM_E m_{SV}}{r^3} \vec{r}$

Differencing Equation:

$$M_E m_{SV} \ddot{r}_S - M_E m_{SV} \ddot{r}_E = -\frac{GM_E m_{SV}^2}{r^3} \vec{r} - \frac{GM_E^2 m_{SV}}{r^3} \vec{r}$$

$$M_E m_{SV} \ddot{r} = -\frac{G}{r^3} \vec{r} (M_E m_{SV} + M_E^2 m_{SV})$$

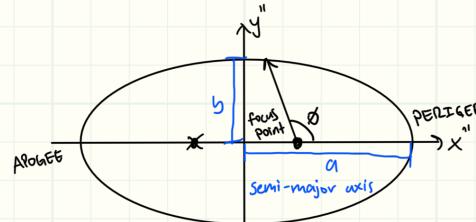
$$\ddot{r} = -\frac{G}{r^3} \vec{r} (M_E + m_{SV})$$

$$\ddot{r} + \frac{GM_{\text{tot}}}{r^3} \vec{r} = 0 \quad \leftarrow 6^{\text{th}} \text{ order non-linear homogeneous diff. eq.}$$

- requires 6 ILS  
 $\vec{r}(0)$  and  $\dot{\vec{r}}(0)$

Solution results in Kepler's 3 Laws of Orbits:

- 1) Elliptical motion
- 2) Motion is faster when closer to orbiting body
- 3)  $\frac{r^2}{t_{\text{orbit}}} = K d_{\text{avg}}^3$   
 $\uparrow$  average orbital distance

Position of SV in Orbital Frame

$$x'' = r \cos(\phi) \quad T_p = \frac{2\pi}{\sqrt{GM_E}} a^{3/2} \quad e = \frac{a^2 - b^2}{a} \quad r = \frac{a(1-e^2)}{1+e \cos(\phi)}$$

$$y'' = r \sin(\phi) \quad \text{Orbital period} \quad b = a \sqrt{1-e^2}$$

SV Position

$$\dot{n} = \frac{2\pi}{T_p} = \frac{GM_E}{a^3} = \frac{6GM}{a^3} \quad |r| = \frac{a(1-e^2)}{1+e \cos(\phi)} \quad \uparrow \text{mean velocity}$$

$$v = \tan^{-1} \left( \frac{\sin(E) \sqrt{1-e^2}}{\cos(E) - e} \right) \quad \uparrow \text{eccentric anomaly}$$

$$M = n(t - t_{\text{PERIGEE}}) \quad \uparrow = E - e \sin(E) \quad x = |r| \cos(v)$$

$$\text{mean anomaly} \quad y = |r| \sin(v) \quad \uparrow \text{true anomaly}$$

\* Must solve for M iteratively until  $\Delta E < 10^{-12}$ 

Taking derivatives:

$$\dot{M} = \dot{E} (1 - \cos(E)) = n$$

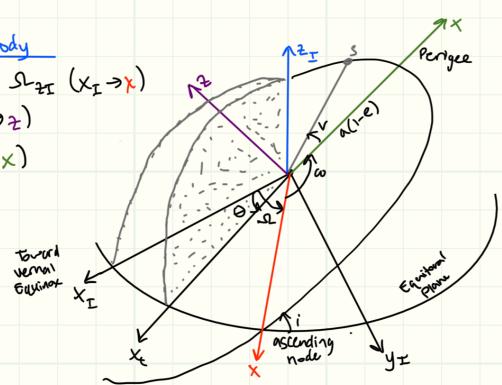
$$\dot{r}_T = \dot{r}_3(\theta) r_i + r_3(\theta) \dot{r}_i$$

$$\dot{r}_X = \frac{-n a \sin(E)}{1 - e \cos(E)} \quad \dot{r}_Y = \frac{n a \cos(E) \sqrt{1-e^2}}{1 - e \cos(E)}$$

$$\ddot{r} = -\frac{GM_{\text{tot}}}{r^3} \vec{r}$$

Rotation Matrices• Orbital Frame  $\rightarrow$  ECEF frame

• Order of rotations is critical!

Inertial to Body1st:  $\theta_{ZI}$  and  $\Omega_{ZI}$  ( $x_I \rightarrow x$ )2nd:  $i_X (z_I \rightarrow z)$ 3rd:  $\omega_z (x \rightarrow x)$ 

# MECH 6970: GPS

## Position Translation with Rotation Matrices

$$\vec{r} = R_3(\omega) R_1(i) R_2(\Omega) \vec{r}_I$$

$\left[ \begin{array}{c} \text{about } z\text{-axis} \\ \omega \end{array} \right] \quad \left[ \begin{array}{c} \text{about } x\text{-axis} \\ i \end{array} \right] \quad \left[ \begin{array}{c} \text{about } z\text{-axis} \\ \Omega \end{array} \right]$

Properties of Rot. Mat.:

$$1) R^{-1} = R^T$$

$$\hookrightarrow R^{-1}(\theta) = R(-\theta) = R^T(\theta)$$

$$\vec{r} = R_3(-\Omega) R_1(-i) R_2(-\omega) \vec{r}_T$$

$$\vec{r}_T = R_2(\Omega) \vec{r}_I$$

$\uparrow$  about same axis as RAAN, combined in GPS

## GPS SV Position Calculation

• GPS calculates position from ascending node:

$$\phi = \omega + v$$

- ∵ no final rotation about z-axis

• Position is calculated as:

$$x = r \cos(\phi)$$

$$y = r \sin(\phi)$$

• Rotating the position from orbital to inertial:

$$\vec{r}_I = R_3(-\Omega) R_1(-i) \vec{r}$$

• GPS uses Longitude of Ascending Node (LAN) which combines the Right Ascension of Ascending Node (RAAN) and the Greenwich Apparent Sidereal Time (GAST) rotations as:

$$\Omega = \Omega_{\text{LAN}}(t) = \Omega_{\text{RAAN}} - \Theta_{\text{GAST}}(t)$$

- Makes it easier to go to ECEF

$$R_3 = \begin{bmatrix} \cos(\Omega) & -\sin(\Omega) & 0 \\ \sin(\Omega) & \cos(\Omega) & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$R_1 = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos(i) & -\sin(i) \\ 0 & \sin(i) & \cos(i) \end{bmatrix}$$

$$\begin{bmatrix} x \\ y \\ z \end{bmatrix} = R_3(-\Omega) R_1(-i) \begin{bmatrix} x' \\ y' \\ z' \end{bmatrix}$$

$$= \begin{bmatrix} x' \cos(\Omega) - y' \sin(\Omega) \cos(i) \\ x' \sin(\Omega) + y' \cos(\Omega) \sin(i) \\ y' \sin(i) \end{bmatrix}$$

$$\begin{aligned} x' &= r \cos(\omega + v) \\ y' &= r \sin(\omega + v) \\ z' &= 0 \end{aligned}$$

$$\Omega = \Omega_{\text{LAN}}$$

## Orbit Perturbations

- Rockets firing intervals

- Non-central (uniform) gravitational force field

- Equatorial bulge

• Produced torque on SV

• Harmonic perturbations (twice per orbit)

- Gravity of sun and moon

- Solar radiation pressure

$$\ddot{\vec{r}} = -\frac{GM_{\text{tot}}}{r^3} \vec{r} + \vec{F}_{\text{dist}}(r, \dot{r}, t)$$

$$\frac{GM_{\text{tot}}}{r^3} \vec{r} \gg \vec{F}_{\text{dist}}(r, \dot{r}, t)$$

• GPS does not broadcast its position, but instead ephemeris correction terms (curve fits) to calculate position (Kepler Orbital mechanics)

## GPS Ephemeris

• Ephemeris: orbit data

• Ephemerides: individual parameters of orbit

• You provide "t"

• GPS provides "t<sub>oe</sub>"

- Nominal Ephemerides:  $\begin{aligned} - \sqrt{a} & - \omega_0 \\ - e & - i_0 \\ - M_0 & - \Omega_0 \end{aligned}$

- Perturbation Effects:  $\begin{aligned} - \Delta n, (\text{IDOT}), \dot{\Omega} & \uparrow \text{secular perturbation} \\ - C_{\text{u}\cos}, C_{\text{u}\sin} & \\ - C_{\text{e}\cos}, C_{\text{e}\sin} & \uparrow \text{Harmonic Perturbation} \\ - C_{\text{i}\cos}, C_{\text{i}\sin} & \end{aligned}$

1) Non-spherical earth

2) Tidal effects

3) Solar radiation pressure

## SV Position Subtleties

• "t" is transmission time

↳ must be corrected by transit time (approx  $\frac{\text{range}}{c}$  or  $\frac{\text{pseudo range}}{c}$ )

- can solve iteratively with user position (exact)

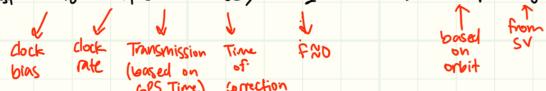
- may, account for earth's rotation during transit time

• check for  $a$  vs  $\sqrt{a}$

## SV Clock Data Corrections

$$\bullet B^k = B^k_{\text{broadcast}} + dB^k$$

$$\bullet B^k_{\text{broadcast}} = af_0 + af_1(T_{\text{tr}} - T_{\text{oc}}) + af_2(T_{\text{tr}} - T_{\text{oc}})^2 + \Delta t_{\text{rel}} + T_{\text{gd}}$$



Ephemeris Updates

- Ephemeris are updated every 2 hours
- Issue of Data Ephemeris (IODE)
  - change of IODE indicates an update to the ephemeris
- Ephemeris are good for 4 hours (maintains spec)
- Ephemeris is group of "ephemerides"
- Must check for clock rollover of t-tse at

How to Get full set of Ephemerides?

- Navigation Message (data broadcast from each SV)
  - can only get data from SV's being tracked
- Overlaid on GPS code "chips"
- GPS C/A code repeats 20 times per bit
- 50 bits/sec
- 1500 bits = 1 frame, 1 frame = 30 sec
  - ↳ IT TAKES 30 SEC TO RECEIVE ALL EPHEMERIS TO COMPUTE THE SV POSITIONS

Subframes

#	10, 30 bit words forming 6s subframe		
	5 subframes form 30s frame (1500 bits)		
1	TLM	HOW	Block 1: Clock Correction
2	TLM	HOW	Block 2: Ephemeris
3	TLM	HOW	Block 3: Ephemeris
4	TLM	HOW	Block 4: Message
5	TLM	HOW	Block 5: Almanac

↑ 25 frames required for complete almanac  
- used to determine when to start tracking different SV

TLM + HOW

- TLM begins with an 8 bit sync pattern
  - 10001011 (0x8b)
  - occurs every 6 sec
- HOW is 17 most significant bits (MSB) of the 19 bit Time of week (Tow) count
  - 6 seconds of resolution
- GPS Time is 29 bits
  - 10 bits for week (1024)
  - 19 bits for Tow (1.5 sec increments)

Overview

- Subframe = 6 sec = 300 bits
- 5 subframes per frame
- Subframes 1-3 repeat every 30 sec

Subframes 4-5

- 25 "pages" for each, repeating after page number 25
- Pages increment every 30 sec
- $25 \times 30 \text{ seconds} = 12.5 \text{ minutes}$  to guarantee reception of all 25 pages

2/20/23

- Orbit is model / diff EQ (6th Order)
- Actual orbit is not perfect
  - Report SV PVT
  - Corrected orbital parameters
    - ↳ provide "curve fit adjustments" to nominal orbit params

$$t_{\text{TRANS}} = t_{\text{RECV}} - \frac{c}{c} \quad ? \quad ?$$

SV pos  
↳ my pos estimate  
 $t_{\text{REC}}, \Delta n, \sqrt{A}, \dots$

GPS Time

- GPS uses its own reference
- GPS → UTC ≈ 18 seconds
- GPST = # weeks + # second in week (12:00 am sunday)
- GPS weeks is Modulo 1024
  - Midnight Jan. 6, 1980
  - Midnight Aug. 22, 1999
  - Midnight Apr. 6, 2019
- 604,800 seconds in a GPS week
- $b^j = t^j - t_{\text{GPS}} = af_0 + af_1(t - t_{\text{rec}}) + af_2(t - t_{\text{rec}})^2 + \Delta t_R - T_{\text{AD}}$

Oscillator Stability

$$\frac{f - f_0}{f_0} = \frac{\Delta f}{f_0} = \frac{\Delta t}{T}$$

$$f(t) = f_0 + \Delta f + \dot{f}(t - t_0) + \eta$$

$$\Delta t = \int f(t) dt = \Delta t(t_0) + \frac{\Delta f}{f_0}(t - t_0) + \frac{\dot{f}}{2f_0}(t - t_0)^2 + \int \eta dt$$

- GPS time is maintained to ~10 ns

Stability Analysis of Clocks

- ↳ Allan variance
- $\sigma_y^2(T) = \frac{1}{N} \sum_{i=1}^{N-1} \frac{1}{2} (y_{i+1} - y_i)^2$
- $N = \frac{T}{T_s}$

