#### 1. WHO5 SYSTEM OPERATION

Below, we overview how the WHO5 Secondary Parameter Determination System works. We will list the needed files and system setup, define the program parameters in header files, walk through a basic data run, and explain how the results are packaged.

This system was built on the openSUSE operating system, using tesh scripts and the Portland Group Fortran 90 compiler (pgf90). The compiler can be changed by modifying the scripts below.

The input files are provided based on the CSP observations of the UBV colors and models, see <a href="https://doi.org/10.48550/arXiv.2311.03473">https://doi.org/10.48550/arXiv.2311.03473</a> for the analysis.

Different filters and model eigenfunctions can be used by modifying list of files (see below) within this tool set.

#### 2. RUNNING THE SYSTEM

1. Before the run, the file 'in.dat\_all' should be created which lists the names of the photometry files. This file should have x entries, with each SN on a new line:

2004ef.v.fin 2004eo.v.fin 2004gu.v.fin

- ... The first few names (presently eight, can be modified by the future user) of the "in.dat\_all" file must be the "messenger" SNe Ia, these eight SNe must be present in all runs.
- 2. File "redshifts25.dat\_all". This must have the names and the redshifts (in floating point) of the same SNe names mentioned in "in.dat\_all". The names may or may not be in the same order as "in.dat\_all". Example of "redshifts25.dat\_all" file:

2004ef.v.raw 0.029771 2004gu.v.raw 0.046897 2005iq.v.raw 0.032929

The names of the SNe are in the first column and the corresponding redshifts are in the second column. The redshift values all have six digits after the decimal.

Afterward, the system is run by the following single command:

## WHO5 [name] [color] [PC choice]

where

[name] is the name given to the run.

[color] is the color to be analyzed; valid choices are [u,b,v].

[PC choice] determines which type of PC analysis to run; valid choices are:

PCd for differential analysis.

PCr for residual analysis.

Examples:

# WHO5 Y35 v PCr

#### WHO5 D40 b PCd

will create a run named "Y35" analyzing all V-band data and will perform a PC analysis on the residuals. will create a run named "D40" analyzing all B-band data and will perform a PC analysis on the differentials.

#### 3. SYSTEM CONTENTS

The working directory should have these components present:

ALPHA11 script which guides the primary-parameter conditioning (shifting and stretching) of the raw data.

alpha.h parameter file for ALPHA11 and its dependencies.

ATEST script which guides the primary-parameter determination based on V-band light curves.

ATEST1\_tcsh script which calls ATEST and recalls ATEST using augmented data (from NARC) if LCs with large gaps around maximum light are found.

**BANDER2v** script which handles finding the times of maximum and maximum apparent magnitude of bands other than (V).

**BANDER2u** is same as BANDER2v, but special for (U) band.

**CDAT** folder containing temporary light curve photometry that has been "conditioned," i.e. been stretched and shifted by ALPHA11.

**CODES** folder containing all source code.

**CONSOLID** script which combines "in.dat.opt" files when augmented data is needed for some SNe (when NARC is called).

**DATACHECK\_NEW1** script that checks if there is LC data in region II for all SNe Ia, except for CSP-I. If an object does not have data in that region, then its name is written in a file named "BadSNeAuto". All other usable object names are written in the "in.dat" file, and their redshifts in the "redshifts25.dat" file. The numbers of the usable SNe are written in the "input.par" file.

**DATACHECK\_NEW2** script checks if the LC data covers up to +25 days. If not, then the name of the object is written in the "BadSNeAuto2" file. It also rewrites the "in.dat", "redshifts25.dat", and "input.par" files.

**DIFL** folder containing temporary differential light curves for all SNe pairs created by SADIE.

EPCSOLV3v solves the overdetermined system of principal component coefficients, for use when PCs of residuals are sought.

**EPCSOLV3**u same as EPCSOLV3, but for use with the (U) band.

**EPCSOLV4v** same as EPCSOLV3, but for use when PCs of differentials are sought.

**EPCSOLV4u** same as EPCSOLV4, but for use with the (U) band.

**EPSILv** script guiding the principal component finder.

**EPSILu** same as EPSIL but for the (U) band.

**GRID** folder for temporary storage of secondary-parameter  $\chi^2$  data produced by SADIE.

**INCLUDE** folder containing all the header files.

in.dat\_all file contains a list of all filenames of SNe photometry data.

LCP3v\_d1, LCP3v\_d2, LCP3v\_d3, LCP3v\_d4 scripts generate overlay plots of stretched and shifted light curves.

LCP3u same as LCP3v but for use with (U)-band data.

 $\mathbf{LCP5}$  script which plots overlaid scaled PC signals. NARC script which handles LCs with large gaps by filling them with template LC data.

**names\_CSPI8.dat** - this file lists the names of the eight messenger SNe Ia from CSPI. This file must be present in the folder named "Peter".

**NEWCHECK** removes the duplicate or same-day data points from LC data.

**Peter** - this folder contains the source codes for the projection of the generic parameters to the physical space, as well as the theoretical relationships to project them ("mrho.data" file). It also contains the names of the "messenger" SNe.

redshifts25.dat\_all file containing redshift information. The format of each line is the six-character SN name followed by the floating-point redshift value.

**RDAT** folder containing all raw (unstretched, unshifted) data. This data should be in files with names 12 characters long; the nominal file format is:

"xxxxxx.x.fin", where the first six characters "xxxxxx" contain the SN identifier (i.e., "2004ef") while the second field "x" is the band of the data (valid options are u,v,b,g,r,i). The last three characters must be "fin". The data files need to have a three-line header and be in a three-column format:

(epoch) (apparent magnitude) (photometric uncertainty)

The data files can have any length necessary.

**RSDL** folder containing all temporary residuals created by SADIE.

**RUNS** folder into which all run tarballs are placed.

SADIE script which guides the secondary-parameter determination from the shifted and stretched light curves.

SADIE013 script which calculates the second-stage primary parameter modifications.

SDAT folder containing temporary data secondary-parameter light curve signals.

SSIG folder containing temporary data used to construct residual plots.

ssptv script which creates plots of residuals and weighted signals.

ssptu is similar to ssptv, but for U-band analysis.

**TDAT** folder containing all templates, for both interpolation and s-correction normalization. This directory must contain the templates:

ntemp.ti

btemp.ti6

utemp.ti6

temp.dat data file listing the names and dm\_15's of templates in TDAT/.

**UDEV** script which computes (U-V) color.

WHO5 master script.

wscrp script which creates the secondary-parameter covariance plots.

#### 4. THE ROUGH ANATOMY OF A RUN

- 1. WHO5 runs "DATACHECK\_NEW1", which checks if there is LC data in region II, and writes "in.dat" and "redshifts25.dat" files.
- 2. "NEWCHECK" removes duplicate LC data points.
- 3. WHO5 looks for the presence of a pre-existing primary parameter file, "in.dat.opt.".
- 4. if the file is not found, ATEST1\_tcsh is called, which uses V-band data to produce the primary parameters and store them in a new file called "in.dat.opt" by running a subscript ATEST. It then calls the NARC script which searches for large gaps in LC data and fills them with template points if necessary. If additional points are indeed needed, ATEST1\_tcsh will run ATEST again to get primary parameters from the augmented LC photometry and store them in "in.dat.opt".
- 5. "DATACHECK\_NEW2" checks if LC data covers up to at least +25 days, if not, then discards the SNe, and rewrites "in.dat" and "redshifts25.dat" files.
- 6. The second-stage primary parameter calculator SADIE013 is called. It modifies the primary parameters in "in.dat.opt". The old "in.dat.opt" of first-stage parameters is saved as "in.dat.opt.1".
- 7. The old or newly-created "in.dat.opt" file is read by the ALPHA11 script which takes the photometry data in RDAT/ and stretches and shifts the data and places the results in the temporary folder CDAT/.
- 8. The file "in.dat.opt" is renamed "in.dat.name" where name is the run name.
- 9. For B or V bands, "LCP3v\_d1", "LCP3v\_d2", "LCP3v\_d3", "LCP3v\_d4" plotting scripts are called. They create LC overlay plots by breaking the sample into four groups (for a large number of SNe, the overlay is hard to see if all LCs are overlaid together in one single plot). "LCP3v\_in" is called next. This script plots the individual stretched and shifted LCs separately. In the case of u-band, these are replaced with "LCP3u".

- 10. The WHO5 script searches for all data in CDAT/ to create an input file, called "sadie.dat," for the secondary-parameter determination program SADIE.
  - Also, the SDAT/ folder is searched for all signals of the user-chosen color.
- 11. The SADIE script is run, which determines the secondary signal presence, creates differentials and puts them into DIFL/, creates residuals and puts them in RSDL/, creates parameter covariance data, and places them in GRID/.
- 12. The script ssptv is run, which takes data from DIFL/, RSDL/, and SSIG/ to create residual plots. If the band analyzed is (U), then the variant ssptu is run. Later these plots are placed in the folder "nameRES" where "name" is the user-defined run name.
- 13. The script wscrp is run, which creates the signal coefficient covariance plots. Later these plots are placed in the folder named "nameTST" where "name" is the run name.
- 14. The UDEV script is run if the analysis color is (U). This script calculates the average (U-V) color around the maximum, and outputs several files of (U-V) data:
  - udev0.dat: the average differential magnitude in the [0:5] day interval.
  - udev1.dat: the individual magnitude of (U-V) color for each SN, the solution of the overdetermined system of the data in 'udev0.dat'.
  - the average (V-U) color taken directly from the data in the udev2.dat: [-3:5] day interval.
- 15. Next, the codes in the folder named "Peter" are run. They translate the generic progenitor parameters into physical secondary parameters by using theoretical relationships. Also, Monte Carlo is done to determine the possible physical range covered by the limited number of SNe in the sample. Important things to note:
  - If using only the CSP-I SNe, then it should be a "reference" run.
  - A file called "switch.dat" must be present in the folder Peter, this must have two integers, "switch" and "jump". The only accepted values for these parameters are 0 or 1. "switch" should be 1 for a reference run (CSP-I), and 1 for all non-reference runs (mix of CSP-I and II, only CSP-II, or any other data, except for just CSP-I). The parameter jump should always be 1 unless the Monte-Carlo method is wished to be avoided.
  - A file named "plotdataPrimaryBen" should be present, this is for use in the reference run only. This file should not be modified, it contains the generic and physical secondary parameter values.
  - If it is a reference run, then the generic and physical secondary parameter values from the run are stored in a file called "plotdataPrimaryBig", used as a standard in all non-reference runs.
  - The "messenger" SNe names are stored in a file called "names\_CSPI8.dat". This file is used for all runs.
  - The results of this part of WHO5 are stored in files named "order.dat" and "plotdataPrimary". This "plotdataPrimary" file would be required for plotting all the secondary parameter distributions later.
- 16. The script EPSILv (or EPSILu, if the (U) band is being analyzed) is run, which initiates the PC analyzer.
- 17. The EPCSOLV3v, EPCSOLV4v, EPCSOLV3u, and EPCSOLV4u scripts are run, depending on the desired results and band analyzed. This script solves the overdetermined PC coefficient system and plots all residuals or differentials with their optimal scaled PCs overlaid.
- 18. The LCP5 script is run, which plots all scaled PCs overlaid upon each other.
- 19. Temporary data is deleted (or stored if necessary, examples are ".sm1", and ".sm2" files, which can be used for re-plotting the  $\chi^2$  maps later if any changes are needed. This can be easily switched off in WHO5) from the working directory, and the results are packaged into a tarball called "RUNname.tar," and all raw data, stretched and shifted data, templates, signals, differentials, residuals, parameter files, scripts, and source code are saved and bundled into the tarball.

- 20. The run tarball is placed into the directory RUNS/ and the script exits, leaving the working directory unchanged.
- 21. "epstojpg\_all" command, where used, converts the ".eps" files to ".png". This command is switched off inside WHO5 currently and can be switched on whenever needed.
- 22. The run folder is placed in the main directory. WHO5 is complete.

# 5. PROGRAM PARAMETERS

In this section, we will define the program parameters as they are found in the header files.

Header file	Variable Name	Variable Type	Description
alpha.h	inpath	CHARACTER*5	Path to raw data, default = "RDAT/".
	outpath	CHARACTER*5	Path to conditioned data, default = "CDAT/".
	mrec	INTEGER	Maximum number of records per raw photometry file, default = 200.
	msne	INTEGER	Maximum number of SNe, $default = 275$ .
	fnlen	INTEGER	Character length of photometry file names plus path length, default = 19.
	$\operatorname{tmax}$	$\operatorname{REAL}$	Deprecated.
	$\operatorname{tmin}$	$\operatorname{REAL}$	Deprecated.
	nstep	$\operatorname{REAL}$	Deprecated.
	nlen	INTEGER	Character length of photometry file names, $default = 14.$
	ierror	INTEGER	Deprecated.
	stand	REAL	Inverse theoretical characteristic signal stretch parameter, default = 1.087.
atest.h	ec	$\operatorname{REAL}$	Deprecated.
	tmax	$\operatorname{REAL}$	Deprecated
	$_{ m tmin}$	REAL	Deprecated
	$dsmin\_p$	REAL	Minimum grid stretch value, default $= -0.05$ .
	$dsmax\_p$	REAL	Minimum grid stretch value, default = $0.05$ .
	$dsmin_p$	REAL	Deprecated
	$\operatorname{dtminp}$	REAL	Deprecated
	$\operatorname{dtmaxp}$	REAL	Deprecated
	chi2min	REAL	Deprecated
	$\operatorname{tfmin}$	REAL	Minimum of flat window interval, default $= -4.0$ .
	tfmax	REAL	Maximum of flat window interval, default $= 15.0$ .
	qcon	REAL	Deprecated

H	leader file	Variable Name	Variable Type	Description
		slack	REAL	Value in days; if no photometry points are found within
				[tfmin:(tfmin+slack)] then 'tfmin' will be extended to 'tfmin-slack/2.0'. Similarly, if no photometry points are
				found within [(tfmax-slack):tfmax] then 'tfmax' will be extended to 'tfmax+slack'. Default = 3.0.
		snorm	REAL	Deprecated
		gaplim	$\operatorname{REAL}$	Deprecated
		sf1	REAL	Signal factor 1. This prevents the "over-flattening" of the LC in the flat window. this factor is multiplied by
				characteristic signal 1, and its value is interpolated at the time of photometry points being fit in the flat window. The interpolated value is added to the
				photometry point's uncertainty. Default = 1.0.
		sf2	REAL	Same as 'sf1' but for characteristic signal 2.  Default = 1.0.
		hlevel	REAL	Deprecated
		qlevel	REAL	Deprecated
		kw	REAL	Deprecated
		eps	REAL	Deprecated
		$\operatorname{seglim}$	$\operatorname{REAL}$	Deprecated
		dmmin	$\operatorname{REAL}$	Deprecated
		dmmax	$\operatorname{REAL}$	Deprecated
		mwinmin	$\operatorname{REAL}$	Deprecated
		mwinmax	$\operatorname{REAL}$	Deprecated
		exten	REAL	Deprecated
		malpha	INTEGER	Defined in parameter file as $msne^{**}2/2-msne/2$ .
		ieo	INTEGER	Deprecated
		mrec	INTEGER	Maximum number of photometry records per object. Default $= 200$ .
		nstep	INTEGER	Deprecated
		msne	INTEGER	Maximum number of SNe. Default $= 275$ .
		fnlen	INTEGER	Number of characters of photometry file names plus path name. $default = 19$ .
		nlen	INTEGER	The number of characters in the photometry file names. Default = 14.

Header file	Variable Name	Variable Type	Description
1100001 1110	nlarge	INTEGER	Maximum number of
		11.11.0110	differential photometry
			points. Default $= 300$ .
	nsstep	INTEGER	The number of discrete
			steps in stretch parameter
			used during stretch primary
			parameter scan. Default = $32$ .
	ntstep	INTEGER	The number of discrete
	P		steps in the time
			parameter, used during the
			time primary parameter
			scan. Default = $32$ .
	nsteptemp	INTEGER	Deprecated
	ntemplate	INTEGER	Deprecated
	nwinmin	INTEGER	Deprecated
	nmstep	INTEGER	Number of discrete steps in
			magnitude parameter, used
			during magnitude primary parameter scan.
	ntrace	INTEGER	Deprecated
	itest1	INTEGER	Deprecated
	itest2	INTEGER	Deprecated
	itertest	INTEGER	Deprecated
	$iter\_max$	INTEGER	Deprecated
	ndmstep	INTEGER	Deprecated
	alph2	INTEGER	Deprecated
	natr	INTEGER	Number of records in
			template data. Default $=$ 191.
	iflatness	INTEGER	Deprecated
	srec	INTEGER	Number of records in the
			characteristic signal data. Default $= 245$ .
	inpath	CHARACTER*5	Path to raw data files. Default = 'RDAT/'.
	outpath	CHARACTER	Deprecated
	tpath	CHARACTER*5	Path to template files.  Default = 'TDAT/'.
epcs.h	malpha	INTEGER	The largest number of differentials allowed. automatically set as ([msne]**2-[msne])/2.
	msne	INTEGER	Maximum number of SNe used; default = 275.
	mrec	INTEGER	Array initializer. Default = 200.
	nstep	INTEGER	Number of uniformly interpolated time bins. Default $= 61$ .
	scalsigpath	CHARACTER*5	Path to the temporary directory holding scaled signals. Default = "ESCL/".

Header file	Variable Name	Variable Type	Description
epsil.h	nlarge	INTEGER	Array initializer. Default = 300.
	nstep	INTEGER	Number of uniformly interpolated time bins for PC analysis. Default = 61.
	msne	INTEGER	Maximum number of SNe. Default $= 50$ .
	malpha	INTEGER	Maximum number of pairs. Automatically = $(msne^{**}2-msne)/2$ .
narc.h	iflatness	INTEGER	Deprecated
	erradd	REAL	Uncertainty added to template-augmented photometry points. Default = 0.005.
	gap	REAL	Value in days; half the limit for photometry gaps in the flat window. larger gaps than (2*gap) will be filled with temporary template points. Default = 2.5.
	tfmax	REAL	Maximum day value of flat window. same as in 'atest.h'. Default = 15.0.
	tfmin	REAL	Minimum day value of flat window. same as in 'atest.h'. Default = -4.0.
	mrec	INTEGER	Maximum number of photometry records per SN. Default $= 200$ .
	natr	INTEGER	Number of records in the template file. same as in 'atest.h'. Default = 191.
	nlarge	INTEGER	Array initializer. Default $=$ 300.
	msne	INTEGER	Maximum number of SNe. Default = $275$ .
	one	INTEGER	$\begin{array}{c} \text{Integer constant. Default} = \\ 1. \end{array}$
	inpath	CHARACTER*5	Path to raw photometry data. Default = 'RDAT/'.
	outpath	CHARACTER*5	Path to place temporary augmented data. Default = 'RDAT/'.
	echar	CHARACTER*10	Diagnostic character. Default = 'narcxxxxxx'.
sadie.h	header	INTEGER	Deprecated
	nsig	INTEGER	Number of characteristic signals to analyze. Default $= 2$ .
	msne	INTEGER	Maximum number of SNe. Default $= 275$ .
	malpha	INTEGER	Maximum number of pairs, automatically = (msne**2-msne)/2).

Header file	Variable Name	Variable Type	Description
	nstep	INTEGER	Deprecated
	nlarge	INTEGER	Array initializer. Default $=$ 300.
	ndmstep	INTEGER	Deprecated
	$\operatorname{ngrid}$	INTEGER	Number of discrete grid elements in covariance analysis. Default $= 100$ .
	inpath	CHARACTER*5	Path to conditioned data.  Default = 'CDAT/'.
	spath	CHARACTER*5	Path to characteristic signals. Default = 'SDAT/'.
	rsidpath	CHARACTER*5	Path where residual data is placed. Default = 'RSDL/'.
	dpath	CHARACTER*5	Path where differentials are placed. Default = 'DIFL/'.
	scalsigpath	CHARACTER*5	Path to the temporary placement of scaled characteristic signals.  Default = 'SSIG/'.
	$\operatorname{tmin}$	$\operatorname{REAL}$	Deprecated
	$\operatorname{tmax}$	$\operatorname{REAL}$	Deprecated
	ec	REAL	Exponential constant used to determine differential point interpolated uncertainty. default = 0.098.
	gaplim	REAL	Largest day gap between photometry points of pair SNe in which differential point will still be created.  Default = 3.5.
	$\operatorname{gapsig}$	REAL	The smallest day gap allowed between two differential data points. Default $= 0.5$ .
	meiminp	$\operatorname{REAL}$	Deprecated
	meimaxp	$\operatorname{REAL}$	Deprecated
	gmax1	REAL	Maximum of coefficient covariance grid used in covariance analysis;
	${ m gmax}2$	REAL	dimension 1. Default = 5.0.  Maximum of coefficient covariance grid used in covariance analysis; dimension 2. Default = 5.0.
	gmin1	REAL	Minimum of coefficient covariance grid used in covariance analysis; dimension 1. Default = -5.0.
	gmin2	REAL	Minimum of coefficient covariance grid used in covariance analysis; dimension 2. Default = -5.0.

Header file	Variable Name	Variable Type	Description
Treader me	ftol	REAL	Fractional tolerance of
	1001	102112	simplex minimizer. default
			= 1.E-30.
	cofsummax	$\operatorname{REAL}$	Deprecated
	moffmin	$\operatorname{REAL}$	Deprecated
	moffmax	$\operatorname{REAL}$	Deprecated
	$\operatorname{mzlim}$	$\operatorname{REAL}$	Deprecated
	mzratio	$\operatorname{REAL}$	Deprecated
	mpoints	INTEGER	Deprecated
	nrand	INTEGER	Deprecated
	iseed	INTEGER	Deprecated
	mgrid	INTEGER	Number of grid points for high-tolerance covariance analysis. Automatically = ngrid**nsig.
sadie0.h	elpoints	INTEGER	Deprecated
	iflatness	INTEGER	Deprecated
	rpath	CHARACTER*5	Path to raw photometry data. Default = 'RDAT/'.
	msne	INTEGER	Maximum number of SNe. Default = $275$ .
	malpha	INTEGER	Maximum number of pairs. Automatically = $(msne^{**}2-msne)/2$ .
	nlarge	INTEGER	Array initializer. Default = 300.
	nmstep	INTEGER	Deprecated
	nsig	INTEGER	Number of characteristic signals analyzed.
	nsrec	INTEGER	Number of records in the characteristic signal files. Default $= 245$ .
	nsstep	INTEGER	The number of discrete steps of the stretch parameter in the grid scan. Default $= 15$ .
	ntstep	INTEGER	The number of discrete steps of time parameter in grid scan. Default = 15.
	itermax	INTEGER	Maximum number of iterations. Default $= 10$ .
	ftol	INTEGER	Simplex minimization fractional tolerance. Default $= 1.0E-5$ .
	$\operatorname{stand}$	REAL	Reciprocal of characteristic signals' stretch. Default = 1.087.
	xmax	REAL	Maximum delta-stretch grid scan value. Default $= 0.10$ .
	xmin	REAL	Minimum delta-stretch grid scan value. Default $= -0.10$ .
	ymax	REAL	The maximum delta-time grid scan value. Default $=$ 3.0.

Header file	Variable Name	Variable Type	Description
	ymin	$\operatorname{REAL}$	The minimum delta-time grid scan value. Default = -3.0.
	zmax	$\operatorname{REAL}$	Deprecated
	zmin	$\operatorname{REAL}$	Deprecated

## 6. PLOTTING ROUTINES

Descriptions of the plotting routines in the system and instructions for using them are given below. All plotting routines are in the main folder SN-DATA-RUN unless otherwise mentioned.

- 1. LCP3v\_d1: Modified from LCP3v script of . Creates overlay plots of the stretched and shifted LCs batch 1.
  - LCP3v\_d2: Modified from LCP3v script of . Creates overlay plots of the stretched and shifted LCs batch 2.
  - $\bullet$  LCP3v\_d3: Modified from LCP3v script of . Creates overlay plots of the stretched and shifted LCs batch 3
  - LCP3v\_d4: Modified from LCP3v script of . Creates overlay plots of the stretched and shifted LCs batch 4.

These scripts are called from within WHO5.

- 2. LCP3v\_in: Creates plots of the stretched and shifted LCs for each individual SN in V-band.
- 3. **LCP3v\_pairOverlay:** Creates overlay plots of the stretched and shifted LCs for each individual pair of SNe in V-band. Commented out currently in WHO5, but can be used to investigate the LC pair further.
- 4. LCP3u: creates overlay plots of the stretched and shifted LCs for u-band.
- 5. ssptv: creates differential plots (residuals and weighted signals) for V-band LCs.
- 6. **ssptu:** creates differential plots (residuals and weighted signals) for u-band LCs.
- 7. wscrp: creates secondary parameter covariance plots.
- 8. LCP5: , plots overlaid scaled PC signals.
- 9. **comparemmscdnfitWOerrpl:** In folder "subsamples". Plots the comparison between the physical secondary parameter values of SNe between the full sample and the subsamples.
- 10. reddeningfreqpl: Plots the reddening distribution histogram. In folder "REDDENING".
- 11. **reddenningpl:** Plots the relationship between  $M_{MS}$  and  $\rho_c$  values and reddening of the SNe in the sample, marks the average  $M_{MS}$  and  $\rho_c$  values of the low, medium, and high reddening groups separately. In folder "REDDENING".
- 12. **reddenning99aa91Tpl:** Same as "reddenningpl", but only for 99aa-likes and 91T-likes. In folder "REDDEN-ING".
- 13. mmscdncorrplSpiralNOerror: Plots  $M_{MS}$  and  $\rho_c$  correlation for all SNe, marks normal-bright, underluminous, 91T-like, and 99aa-like objects separately. In folder "MMSCDNcorr".
- 14. **plHostdm15corrAll:** Plots the correlation between  $\Delta m_{15}(B)$ , s, and host-types for all SNe used in our sample, and the outliers. Normal-bright, 91bg-like, and outliers are marked separately. In folder "HostDm15SNtypeRelation\_all".

- 15. **mmsfreqLpl:** Plots  $M_{MS}$  distribution with normal-bright, transitional, and subluminous SNe marked with patterns. In folder "NEWerrplots".
- 16. **cdnfreqLpl:** Same as "mmsfreqLpl", but for  $\rho_c$ . In folder "NEWerrplots".
- 17.  $\mathbf{mmsfreqYoungplot:}$  In folders "PlusMinusDwarfE", "PlusMinusDwarfEbulge", and "PlusMinusDwarfEbulgeS0". Plots the  $M_{MS}$  distribution for SNe in the spiral, S0, irregular, and small E hosts in folder "PlusMinusDwarfEbulgeS0", this script excludes the SNe in spiral bulges, and in folder "PlusMinusDwarfEbulgeS0", it excludes SNe in S0 galaxies. Run FORTRAN code "plotml.f" (placed in folder "CODES") to prepare the data file required for the plotting routine to work. The make file for this code is "Makefile.plotml", placed in "SN-DATA-RUN" (the main directory containing WHO5). Run "make -f Makefile.plotml" and then "xplotml" to run the code. Input files for the code that must be present in the main directory are "hostmass130.dat", "in.dat.opt", and "plotdataPrimary". The "hostmass130.dat" file contains the host type information for all SNe. "plotml.f" prepares "cdnfreqhistEllipticalBatchX.dat" or "cdnfreqhistEllipticalBatchX.dat" files, on which the distribution scripts can run.For example, a part of the "hostmass130.dat" file is shown below:

```
AS15db
         12.8824955
                          bulge
                      3
         25.7039578
                      1
                          large E
AS15dd
AS15eb
         91.2010839
                          S0
                      2
                          S0
AS15ga
         23.9883292
         0.87096359
                          spirals
AS15gr
                      4
```

Only the names of the SNe with six letters have been used in the first column. In the second column, the host type codes 1, 2, 3, and 4 have been used for large ellipticals, S0, spiral bulge, and spiral hosts, respectively. Host type code 5 is used when no host mass information is available.

- 18. **cdnfreqYoungplot:** In folders "PlusMinusDwarfE", "PlusMinusDwarfEbulge", and "PlusMinusDwarfEbulgeS0". Same as "mmsfreqYoungplot", but for  $\rho_c$ . Running instructions are the same as "mmsfreqYoungplot".
- 19. mmsfreqOldplot: In folders "PlusMinusDwarfE", "PlusMinusDwarfEbulge", and "PlusMinusDwarfEbulgeS0". Same as "mmsfreqYoungplot", but for SNe in large E hosts in folder "PlusMinusDwarfE". In folder "PlusMinusDwarfEbulge", this script includes the SNe in spiral bulges, and in folder "PlusMinusDwarfEbulgeS0", it includes SNe in S0 galaxies. Running instructions are the same as "mmsfreqYoungplot".
- 20. **cdnfreqOldplot:** In folders "PlusMinusDwarfE", "PlusMinusDwarfEbulge", and "PlusMinusDwarfEbulgeS0". Same as "mmsfreqOldplot", but for  $\rho_c$ . Running instructions are the same as "mmsfreqYoungplot".

# 7. READING RESULTS

The run results are stored in a directory called "RUNname", produced by the run, where "name" is the user-given run name. In the "RUNname" directory, there are several subdirectories:

- The "DATA" folder contains copies of raw data in "RDAT", stretched data in "CDAT", templates in "TDAT", and signals in "SDAT". Modification in this work: the run directory also contains a sub-folder named "DATA\_TST" which has the necessary files for plotting the signal coefficient covariance plots (just in case any modifications are needed).
- Folder "nameRES" contains all differential plots and scaled signals. "name" is the run name.
- Folder "nameTST" contains signal coefficient covariance plots. "name" is the run name.
- The folder "LC" contains all individual and overlaid stretched LC plots.
- File "runlog.name" contains a detailed log of the run.
- Folder "UNCERTAINTY" contains files containing information to determine  $t_{cor}$  and  $f_{cor}$  if needed, which is in the case of a new data set being used.

- Folder "DPAR" contains "cof.dat" signal coefficients file, "epccof.dat" PC coefficients file, "primary.dat" preflatness primary parameter file, "epc.dat", and "avg.dat" which is the error-weighted average of the differentials (if the PCd option was used) or the residuals (if the PCr option was used).
- Folder "EPJPG" contains differentials or residuals with the optimally scaled PC signal overlaid.
- Folder "PPAR" contains all header files along with copies of data files describing templates used "temp.dat" and signals "sig.dat".
- Folder "SRC" contains copies of all source codes and scripts, along with WHO5.
- Along with the subdirectories, the run directory will also contain:

avg.eps a plot of the average differential (if the PCd option was used) or average residual (if the PCr option was used).

cdn.eps a histogram of central density signal coefficients.

chi2.eps a histogram of chi2 fit residuals.

epc.eps a plot of the principal component; this is the most prominent factor in the differentials (if the PCd option was set) or the residuals (if the PCr option was used).

epchist.eps a histogram of optimal PC coefficients.

[name]LC.eps the stretched and shifted light-curve overlay plot.

[name]EPCscaled.eps

a plot of all optimally scaled PCs overlaid.

runlog.[name]

a log containing each stage's status output, along with the start and end times of the run.

mms.eps

a histogram of main-sequence mass signal coefficients.