Physics 106a — Classical Mechanics

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Lecture 8

Unbound Orbits in Central Potentials: Scattering

Unbound Orbits in Central Potentials: Scattering

- Review of planetary orbits
- \blacksquare Repulsive 1/r potential: Rutherford scattering
- Scattering problems: what do we want to know?
- Particle trajectory for repulsive 1/r potential
- Rutherford scattering cross section
- Effect of finite target mass: kinematics and dynamics

Review of planetary orbits

Rotational symmetry

- Eliminate center of mass motion
- Angular momentum \vec{l} is conserved
- Orbit lies in plane perpendicular to \vec{l} : specify by (r, ϕ)
- Lagrangian for 2d motion

$$\mathcal{L} = \frac{1}{2}\mu(\dot{r}^2 + r^2\dot{\phi}^2) - V(r)$$

Constant of the motion

$$l = \mu r^2 \dot{\phi}$$

— Kepler's second law (constant rate of sweeping out area)

Hamiltonian

Hamiltonian is total energy and is constant

$$E = \frac{1}{2}\mu \dot{r}^2 + V_{\rm eff}(r) \quad {\rm with} \quad V_{\rm eff}(r) = \frac{l^2}{2\mu r^2} + V(r)$$

■ For $V(r) = -\frac{k}{r}$, introduce $u = \frac{1}{r}$, and use then $\dot{r} = \dot{\phi} \frac{dr}{d\phi} = -\frac{l}{\mu} \frac{du}{d\phi}$

$$E = \frac{l^2}{2\mu} \left[\left(\frac{du}{d\phi} \right)^2 + \left(u - \frac{1}{p} \right)^2 - \frac{1}{p^2} \right] \quad \text{with } p = \frac{l^2}{\mu k}.$$

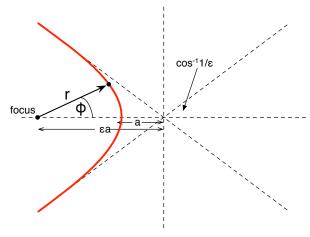
 $\mathbf{u}(\phi)$ is sinusoidal about 1/p

$$\frac{1}{r} = \frac{1}{p} + \frac{\epsilon}{p} \cos \phi$$

■ Conic section with eccentricity ϵ and $E = \frac{l^2}{2\mu p^2} (\epsilon^2 - 1)$

For
$$\epsilon > 1$$
 the orbits are hyperbolas

$$\frac{1}{r} = \frac{1}{p} + \frac{\epsilon}{p} \cos \phi$$



Rutherford Scattering

Geiger and Marsden, 1909

Scattering of alpha particles off thin gold film [Ernest Rutherford]

It was quite the most incredible event that has ever happened to me in my life. It was almost as incredible as if you fired a 15-inch shell at a piece of tissue paper and it came back and hit you.

On consideration, I realized that this scattering backward must be the result of a single collision, and when I made calculations I saw that it was impossible to get anything of that order of magnitude unless you took a system in which the greater part of the mass of the atom was concentrated in a minute nucleus. It was then that I had the idea of an atom with a minute massive center, carrying a charge.

Scattering Problems

What do we want to know?

Repulsive k/r potential

Hamiltonian is total energy and is constant

$$E = \frac{1}{2}\mu\dot{r}^2 + V_{\rm eff}(r) \quad {\rm with} \quad V_{\rm eff}(r) = \frac{l^2}{2\mu r^2} + V(r)$$

■ For $V(r) = \frac{k}{r}$ (with k > 0), introduce $u = \frac{1}{r}$ so that $\dot{r} = \dot{\phi} \frac{dr}{d\phi} = -\frac{l}{\mu} \frac{du}{d\phi}$

$$E = \frac{l^2}{2\mu} \left[\left(\frac{du}{d\phi} \right)^2 + \left(u + \frac{1}{p} \right)^2 - \frac{1}{p^2} \right] \quad \text{with } p = \frac{l^2}{\mu k} > 0$$

■ $u(\phi)$ is sinusoidal about -1/p

$$\frac{1}{r} = -\frac{1}{p} + \frac{\epsilon}{p}\cos\phi$$

Physical solutions r > 0 for $\cos \phi > 1/\epsilon$

■ Conic section with eccentricity $\epsilon > 1$ (hyperbola) and $E = \frac{l^2}{2\mu p^2}(\epsilon^2 - 1)$

Hyperbolic orbits: $\epsilon > 1$

Attractive and repulsive interactions

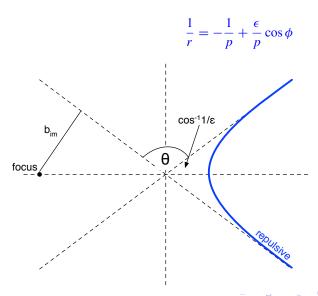
$$\frac{1}{r} = \frac{1}{p} + \frac{\epsilon}{p} \cos \phi \qquad \frac{1}{r} = -\frac{1}{p} + \frac{\epsilon}{p} \cos \phi$$

$$\cos^{-1}1/\epsilon$$

$$\epsilon a \qquad \epsilon a \qquad \epsilon a$$

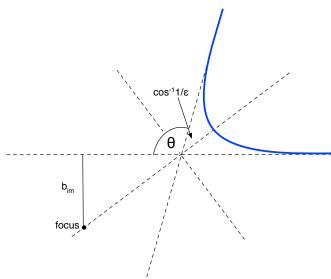
Hyperbolic orbits

Scattering in repulsive 1/r potential



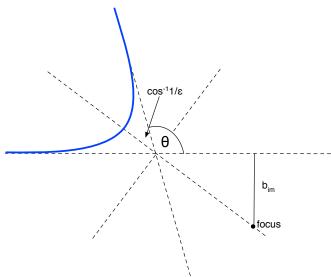
Hyperbolic orbits

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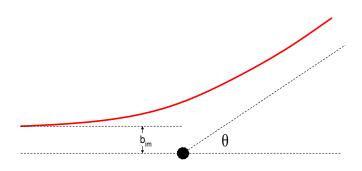


Hyperbolic orbits

Scattering in repulsive 1/r potential

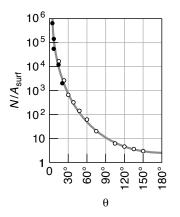


Scattering Geometry



Experimental Results

Geiger and Marsden 1909+



From *The Atomic Nucleus* by R. D. Evans

The closest distance of approach was 30 fm (3 \times 10⁻¹⁴m) for 7.78 MeV alpha particle at 150 deg scattering

Finite Target Mass

 $p = 1, \epsilon = 1.2 \Rightarrow \theta_{\rm cm} = 146^{\circ}$

Center of mass frame

