

# Parameters Constraints on $\Lambda$ CDM and XCDM using H(z)+BAO data

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## Abstract

Here I reproduce the results of Cao and Ratra [1] for the spatially flat and non-flat  $\Lambda$ CDM and XCDM dark-energy models. Following [1], I first analyze the Old [2] and New [1] Hubble-parameter measurements, and then jointly use them with the Old [2] and New [1] BAO data to constrain the relevant cosmological parameters. The resulting parameter estimates are in close agreement with those reported in the original studies.

## 1 Data and Methodology

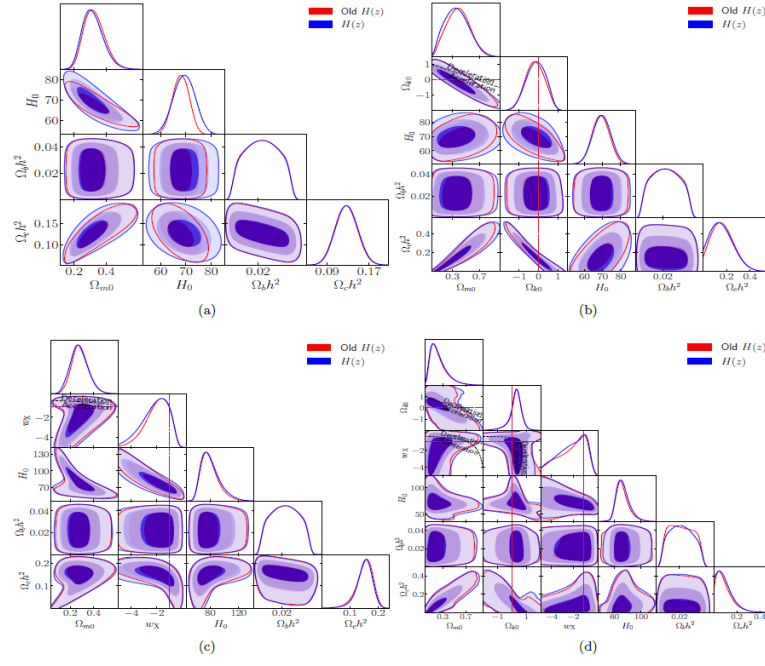
For the old H(z) and old BAO data, I use Table-I and II of [2] and for the new data, I use Table-I(accounting for correlations) and II of [1]. First, I implemented the  $\chi^2$  function and minimized it using `scipy.minimize`. This yielded the minimized value  $\chi^2_{\min}$  and the best-fit estimates of the free parameters  $\Omega_{k0}$ ,  $w_x$ ,  $\Omega_b h^2$ ,  $\Omega_c h^2$ , and  $H_0$ , consistent with those used in the reference paper. From these free parameters one can also compute the best-fit value of the derived matter-density parameter  $\Omega_m = \frac{\Omega_b h^2 + \Omega_c h^2 + \Omega_\nu h^2}{h^2}$  where I use the neutrino-density value  $\Omega_\nu h^2 = 0.06/93.14$  as given in the paper.

I then ran `emcee` to obtain Markov chains for these free parameters. For all parameters I adopted the same priors as used in the paper. The MCMC simulations were run with 200 walkers for 5000-10,000 steps varying accross different models.

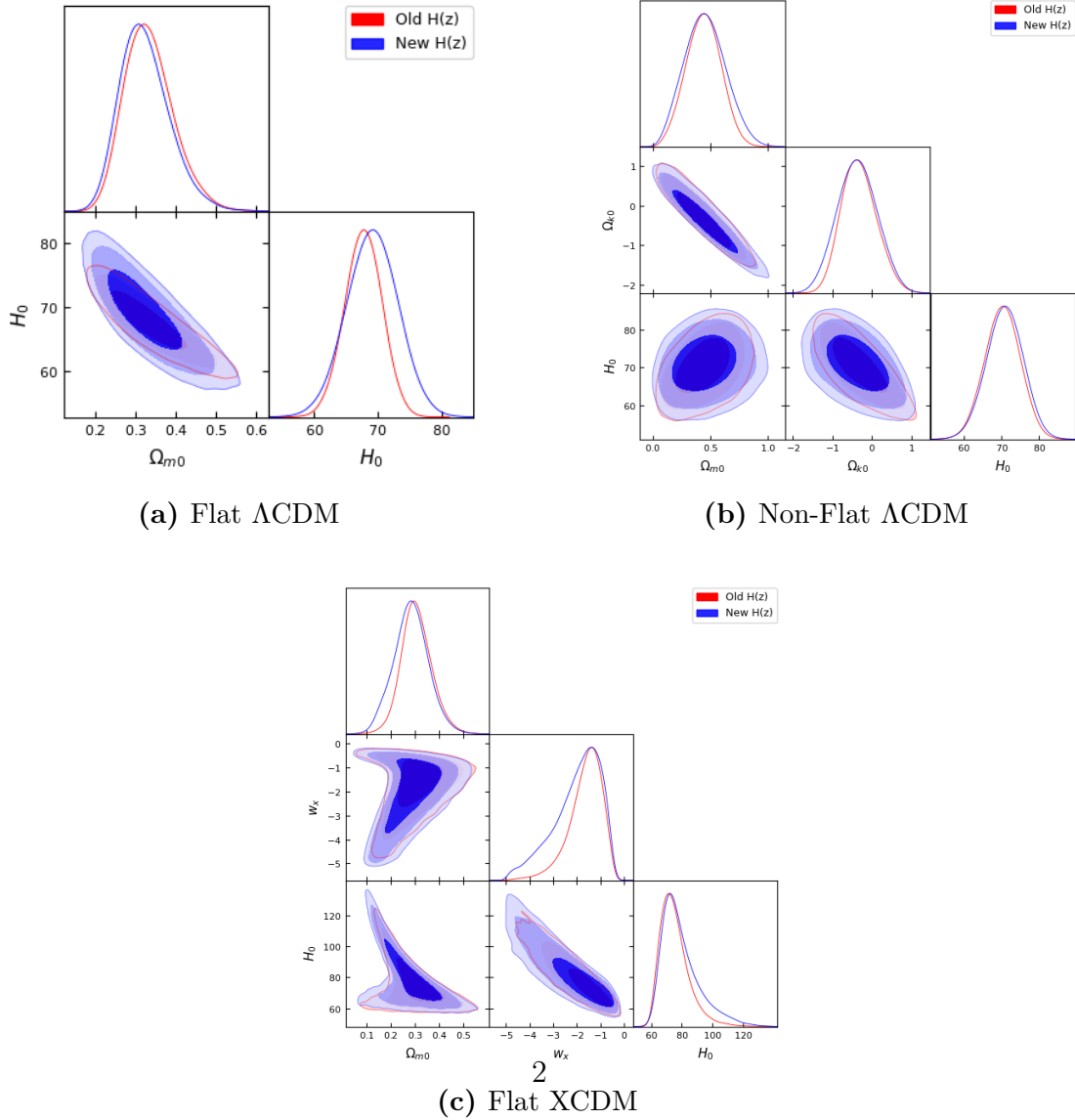
## 2 Comparison of Results

### 2.1 Plots

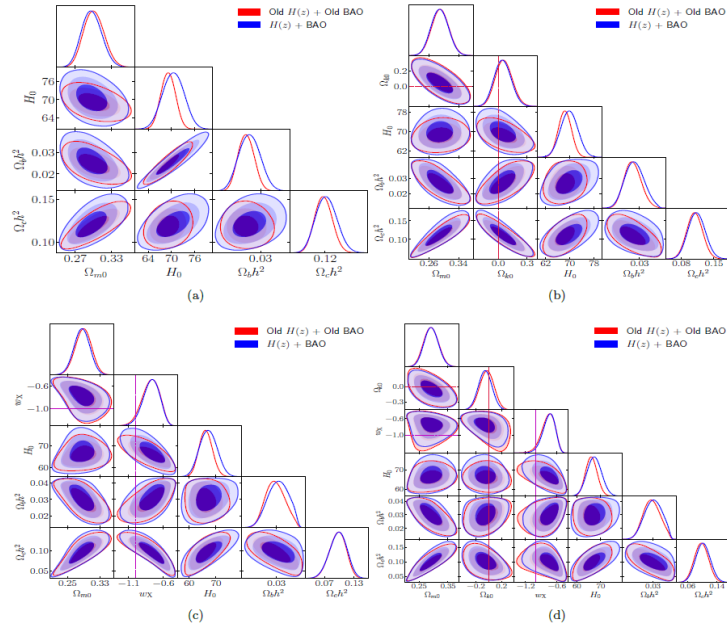
Below I report the plots I obtained, at first using only H(z) data and then combining with the BAO data and performing a joint analysis. I compare them with the original plots in [1].



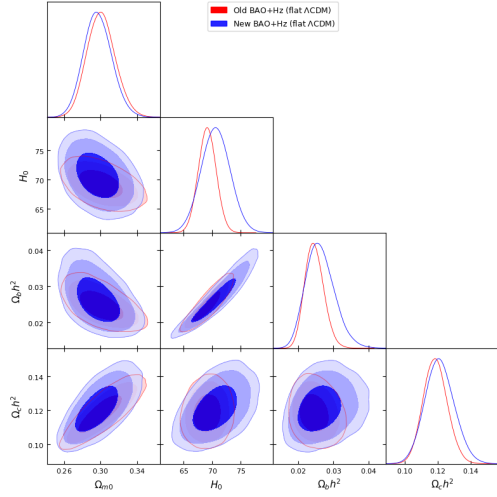
**Figure 1:** Original Plots in Cao and Ratra[1]. a) Flat  $\Lambda$ CDM, b) Non-Flat  $\Lambda$ CDM, c) Flat XCDM, d) Non-Flat XCDM



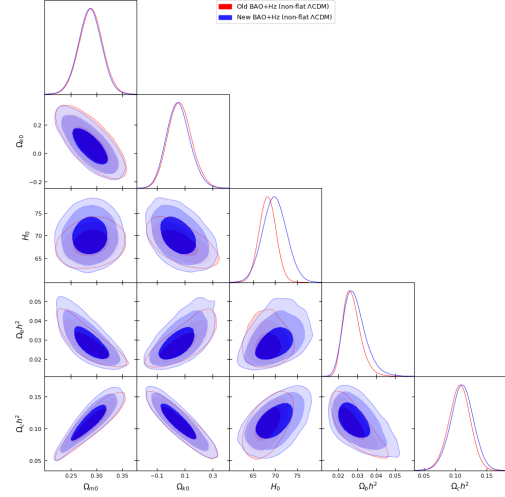
**Figure 2:** Plots generated using emcee and scipy.



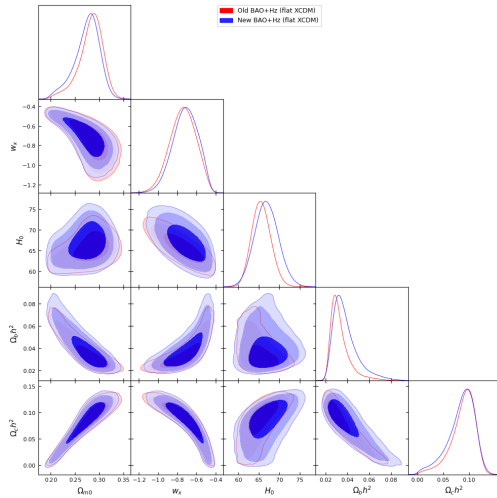
Original Plots in Cao and Ratra [1]. a) Flat  $\Lambda$ CDM, b) Non-Flat  $\Lambda$ CDM, c) Flat XCDM, d) Non-Flat XCDM



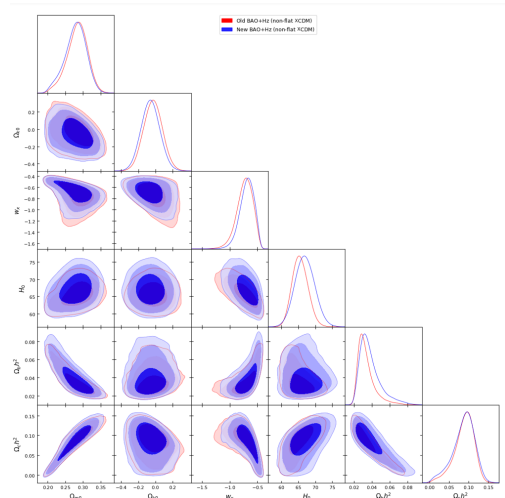
(a) Flat  $\Lambda$ CDM



(b) Non-Flat  $\Lambda$ CDM



(c) Flat XCDM



(d) Non-Flat XCDM

**Figure 3:** Plots generated using emcee and scipy.

**Table 1:** Comparison of Best-Fitting Cosmological Parameters: My Results vs. Cao and Ratra (2023)[1]

Model	Dataset	My Results						Cao and Ratra [Source: 528]					
		$\Omega_b h^2$	$\Omega_c h^2$	$\Omega_{m0}$	$\Omega_{k0}$	$w_X$	$H_0$	$\Omega_b h^2$	$\Omega_c h^2$	$\Omega_{m0}$	$\Omega_{k0}$	$w_X$	$H_0$
Flat $\Lambda$ CDM	Old $H(z)$	—	—	0.3197	—	—	68.116	0.0273	0.1201	0.319	—	—	68.16
	$H(z)$	—	—	0.3097	—	—	69.428	0.0244	0.1181	0.301	—	—	69.43
	Old $H(z)$ +Old BAO	0.024	0.118	0.299	—	—	68.953	0.0254	0.1200	0.309	—	—	68.98
	$H(z)$ +BAO	0.025	0.120	0.295	—	—	70.134	0.0260	0.1200	0.297	—	—	70.12
Non-flat $\Lambda$ CDM	Old $H(z)$	—	—	0.354	-0.11	—	68.93	0.0205	0.1515	0.362	-0.136	—	69.09
	$H(z)$	—	—	0.315	-0.014	—	69.5	0.0180	0.1328	0.314	-0.012	—	69.47
	Old $H(z)$ +Old BAO	0.026	0.110	0.291	0.048	—	68.329	0.0260	0.1098	0.292	0.048	—	68.35
	$H(z)$ +BAO	0.027	0.113	0.289	0.038	—	69.635	0.0269	0.1128	0.289	0.041	—	69.61
Flat XCDM	Old $H(z)$	—	—	0.3226	—	-1.185	70.241	0.0376	0.1236	0.321	—	-1.261	70.95
	$H(z)$	—	—	0.3158	—	-1.133	70.569	0.0106	0.1464	0.316	—	-1.140	70.63
	Old $H(z)$ +Old BAO	0.030	0.094	0.288	—	-0.755	65.615	0.0951	0.0296	0.290	—	-0.754	70.79
	$H(z)$ +BAO	0.033	0.092	0.281	—	-0.731	66.686	0.0938	0.0318	0.283	—	-0.734	66.67
Non-flat XCDM	Old $H(z)$	—	—	—	—	—	—	0.0223	0.0736	0.172	0.324	-2.272	75.05
	$H(z)$	—	—	—	—	—	—	0.0316	0.0530	0.378	0.151	-2.278	75.06
	Old $H(z)$ +Old BAO	0.029	0.098	0.294	-0.054	-0.730	65.682	0.0289	0.0985	0.296	-0.053	-0.730	65.76
	$H(z)$ +BAO	0.032	0.098	0.291	-0.081	-0.699	66.859	0.0305	0.0998	0.293	-0.084	-0.703	66.79

**Table 2:** Comparison of One-dimensional posterior mean parameter values and uncertainties: My Results vs. Cao and Ratra (2023)[1]

Model	Dataset	My Results						Cao and Ratra [Source: 539]					
		$\Omega_b h^2$	$\Omega_c h^2$	$\Omega_{m0}$	$\Omega_{k0}$	$w_X$	$H_0$	$\Omega_b h^2$	$\Omega_c h^2$	$\Omega_{m0}$	$\Omega_{k0}$	$w_X$	$H_0$
Flat $\Lambda$ CDM	Old $H(z)$	—	—	0.334 <sup>+0.009</sup> <sub>-0.070</sub>	—	—	67.7 ± 3.0	0.0225 ± 0.0108	0.1264 ± 0.0207	0.328 <sup>+0.002</sup> <sub>-0.055</sub>	—	—	67.98 ± 3.24
	$H(z)$	—	—	0.324 <sup>+0.049</sup> <sub>-0.073</sub>	—	—	69.1 ± 4.1	0.0225 ± 0.0107	0.1275 ± 0.0208	0.319 ± 0.050	—	—	69.31 ± 4.25
	Old $H(z)$ +Old BAO	0.0246 <sup>+0.0024</sup> <sub>-0.0029</sub>	0.1184 ± 0.0073	0.301 <sup>+0.014</sup> <sub>-0.016</sub>	—	—	69.1 ± 1.6	0.0247 ± 0.0030	0.1186 <sup>+0.0076</sup> <sub>-0.0075</sub>	0.301 <sup>+0.016</sup> <sub>-0.018</sub>	—	—	69.14 ± 1.85
	$H(z)$ +BAO	0.0263 <sup>+0.0035</sup> <sub>-0.0045</sub>	0.1211 <sup>+0.0062</sup> <sub>-0.016</sub>	0.297 <sup>+0.015</sup> <sub>-0.016</sub>	—	—	70.6 ± 2.5	0.0260 ± 0.0040	0.1212 <sup>+0.0091</sup> <sub>-0.0101</sub>	0.297 <sup>+0.015</sup> <sub>-0.016</sub>	—	—	70.49 ± 2.74
Non-flat $\Lambda$ CDM	Old $H(z)$	—	—	0.44 ± 0.15	-0.32 <sup>+0.40</sup> <sub>-0.49</sub>	—	70.3 ± 4.7	0.0223 <sup>+0.0109</sup> <sub>-0.0108</sub>	0.1685 <sup>+0.130</sup> <sub>-0.113</sub>	0.390 <sup>+0.172</sup> <sub>-0.155</sub>	-0.174 <sup>+0.491</sup> <sub>-0.494</sub>	—	69.09 <sup>+1.67</sup> <sub>-1.67</sub>
	$H(z)$	—	—	0.45 <sup>+0.17</sup> <sub>-0.19</sub>	-0.37 ± 0.50	—	70.9 ± 4.9	0.0222 ± 0.0108	0.1612 <sup>+0.0691</sup> <sub>-0.1064</sub>	0.374 <sup>+0.210</sup> <sub>-0.210</sub>	-0.136 <sup>+0.477</sup> <sub>-0.481</sub>	—	69.56 <sup>+1.88</sup> <sub>-1.88</sub>
	Old $H(z)$ +Old BAO	0.0274 <sup>+0.0034</sup> <sub>-0.0054</sub>	0.107 ± 0.017	0.289 ± 0.023	0.064 <sup>+0.079</sup> <sub>-0.091</sub>	—	68.3 ± 1.9	0.0266 <sup>+0.0029</sup> <sub>-0.0045</sub>	0.1088 ± 0.0166	0.291 ± 0.023	0.059 <sup>+0.081</sup> <sub>-0.091</sub>	—	68.37 ± 2.10
	$H(z)$ +BAO	0.0288 <sup>+0.0042</sup> <sub>-0.0067</sub>	0.111 ± 0.018	0.287 ± 0.022	0.054 <sup>+0.077</sup> <sub>-0.090</sub>	—	69.9 ± 2.7	0.0275 <sup>+0.0046</sup> <sub>-0.0051</sub>	0.1131 <sup>+0.0186</sup> <sub>-0.0204</sub>	0.289 ± 0.023	0.047 <sup>+0.082</sup> <sub>-0.093</sub>	—	69.81 ± 2.80
Flat XCDM	Old $H(z)$	—	—	0.306 <sup>+0.067</sup> <sub>-0.066</sub>	—	-1.64 <sup>+0.84</sup> <sub>-0.43</sub>	75.4 <sup>+11.4</sup> <sub>-11.4</sub>	0.0225 ± 0.0107	0.1505 <sup>+0.0617</sup> <sub>-0.0617</sub>	0.285 <sup>+0.117</sup> <sub>-0.117</sub>	—	-1.972 <sup>+1.164</sup> <sub>-1.164</sub>	79.55 <sup>+1.70</sup> <sub>-1.70</sub>
	$H(z)$	—	—	0.283 ± 0.073	—	-1.98 <sup>+1.3</sup> <sub>-0.53</sub>	79.4 <sup>+15.9</sup> <sub>-15.9</sub>	0.0225 ± 0.0108	0.1505 <sup>+0.0617</sup> <sub>-0.0617</sub>	0.278 <sup>+0.081</sup> <sub>-0.081</sub>	—	-2.127 <sup>+1.335</sup> <sub>-1.335</sub>	80.96 ± 16.10
	Old $H(z)$ +Old BAO	0.0337 <sup>+0.0032</sup> <sub>-0.010</sub>	0.087 <sup>+0.028</sup> <sub>-0.014</sub>	0.281 <sup>+0.029</sup> <sub>-0.017</sub>	—	-0.74 <sup>+0.14</sup> <sub>-0.13</sub>	65.6 <sup>+2.1</sup> <sub>-2.5</sub>	0.0308 <sup>+0.0038</sup> <sub>-0.0046</sub>	0.0969 <sup>+0.0178</sup> <sub>-0.0174</sub>	0.289 ± 0.020	—	-0.784 <sup>+0.140</sup> <sub>-0.107</sub>	66.22 <sup>+2.23</sup> <sub>-2.54</sub>
	$H(z)$ +BAO	0.0382 <sup>+0.0050</sup> <sub>-0.013</sub>	0.085 <sup>+0.031</sup> <sub>-0.016</sub>	0.274 <sup>+0.030</sup> <sub>-0.019</sub>	—	-0.72 <sup>+0.15</sup> <sub>-0.12</sub>	67.0 ± 2.9	0.0303 <sup>+0.0059</sup> <sub>-0.0047</sub>	0.0978 <sup>+0.0181</sup> <sub>-0.0175</sub>	0.285 ± 0.019	—	-0.776 <sup>+0.135</sup> <sub>-0.100</sub>	67.18 <sup>+2.28</sup> <sub>-2.28</sub>
Non-flat XCDM	Old $H(z)$	—	—	—	—	—	—	0.0218 <sup>+0.0082</sup> <sub>-0.0082</sub>	0.0927 ± 0.0217	0.228 <sup>+0.095</sup> <sub>-0.175</sub>	0.241 <sup>+0.354</sup> <sub>-0.201</sub>	-2.148 <sup>+1.826</sup> <sub>-1.826</sub>	71.98 <sup>+2.88</sup> <sub>-2.88</sub>
	$H(z)$	—	—	—	—	—	—	0.0218 ± 0.0093	0.0927 ± 0.0217	0.228 <sup>+0.054</sup> <sub>-0.175</sub>	0.228 <sup>+0.356</sup> <sub>-0.201</sub>	-2.149 <sup>+1.873</sup> <sub>-1.873</sub>	73.06 ± 6.61
	Old $H(z)$ +Old BAO	0.0337 <sup>+0.0038</sup> <sub>-0.011</sub>	0.088 <sup>+0.029</sup> <sub>-0.018</sub>	0.283 <sup>+0.032</sup> <sub>-0.025</sub>	-0.03 ± 0.11	-0.73 <sup>+0.17</sup> <sub>-0.11</sub>	65.6 <sup>+2.1</sup> <sub>-2.5</sub>	0.0294 <sup>+0.0047</sup> <sub>-0.0045</sub>	0.0980 ± 0.0186	0.292 ± 0.025	-0.027 ± 0.109	-0.770 <sup>+0.158</sup> <sub>-0.109</sub>	66.13 <sup>+2.35</sup> <sub>-2.35</sub>
	$H(z)$ +BAO	0.0382 <sup>+0.0053</sup> <sub>-0.014</sub>	0.087 <sup>+0.033</sup> <sub>-0.020</sub>	0.278 <sup>+0.034</sup> <sub>-0.025</sub>	-0.06 ± 0.11	-0.69 <sup>+0.15</sup> <sub>-0.10</sub>	67.0 ± 2.9	0.0303 <sup>+0.0048</sup> <sub>-0.0051</sub>	0.1021 ± 0.0193	0.292 ± 0.024	-0.054 ± 0.103	-0.757 <sup>+0.135</sup> <sub>-0.093</sub>	67.33 ± 2.96

## References

- [1] Shulei Cao, Bharat Ratra, " $H_0 = 69.8 \pm 1.3 \text{ km s}^{-1}\text{Mpc}^{-1}$ ,  $\Omega_{m_0} = 0.288 \pm 0.017$ , and other constraints from lower-redshift, non-CMB, expansion-rate data", arXiv:2302.14203 [astro-ph.CO].
- [2] Shulei Cao, Bharat Ratra, "Using lower-redshift, non-CMB, data to constrain the Hubble constant and other cosmological parameters", arXiv:2203.10825 [astro-ph.CO].