

Solutions to Problems in Perkins,  
*Introduction to High Energy Physics*,  
Third Edition

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## Chapter 1

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### Problem 1.1

An electron of energy 20 GeV is deflected through an angle of  $5^\circ$  in an elastic collision with a stationary proton. What is the value of the square of the 4-momentum transfer,  $q^2$ , and down to what approximate distance does such a collision probe the internal structure of the proton? (The mass of the electron can be neglected compared with the energies involved. The proton mass  $Mc^2$  is 0.938 GeV.)

First let's define terms:  $p_i = 20 \text{ GeV}/c$  is the magnitude of the electron's initial momentum.  $p_e$  and  $p_p$  are the final electron and proton momenta.  $\theta_e = 5^\circ$  and  $\theta_p$  are the electron and proton scattering angles.

Neglecting the electron mass, energy conservation requires

$$p_i c + M_p c^2 = p_e c + [p_p^2 c^2 + M_p^2 c^4]^{1/2}$$

which can be rearranged to give

$$p_p^2 = q^2 + 2qM_p c, \quad q = p_i - p_e. \quad (1)$$

Momentum conservation requires

$$p_i = p_e \cos \theta_e + p_p \cos \theta_p \quad (2)$$

$$0 = p_e \sin \theta_e + p_p \sin \theta_p. \quad (3)$$

From the second of these we obtain

$$\sin \theta_p = -\frac{p_e}{p_p} \sin \theta_e$$

so

$$\cos \theta_p = \left[ 1 - \frac{p_e^2}{p_p^2} \sin^2 \theta_e \right]^{1/2}.$$

With this we can eliminate  $\theta_p$  from (1):

$$p_i = p_e \cos \theta_e + p_p \left[ 1 - \frac{p_e^2}{p_p^2} \sin^2 \theta_e \right]^{1/2}$$

or

$$\begin{aligned} (p_i - p_e \cos \theta_e)^2 &= p_p^2 - p_e^2 \sin^2 \theta_e \\ p_i^2 + p_e^2 - 2p_i p_e \cos \theta_e &= p_p^2. \end{aligned}$$

Plugging (1) into this gives

$$\begin{aligned} p_i^2 + p_e^2 - 2p_i p_e \cos \theta_e &= q^2 + 2qM_p c \\ &= p_i^2 + p_e^2 - 2p_i p_e + 2p_i M_p c - 2p_e M_p c \end{aligned}$$

so

$$2p_i p_e [1 - \cos \theta_e] + 2p_e M_p c = 2p_i M_p c.$$

With this we find

$$\begin{aligned} p_e &= \frac{p_i M_p c}{p_i [1 - \cos \theta_e] + M_p c} \\ &= \frac{(20 \text{ GeV}/c) \cdot (0.938 \text{ GeV}/c)}{(20 \text{ GeV}/c)[1 - \cos 5^\circ] + (0.938 \text{ GeV}/c)} \\ &= 18.5 \text{ GeV}/c. \end{aligned}$$

This gives a  $q$  value (three-momentum transfer) of  $20 - 18.5 = 1.5 \text{ GeV}/c$ . The four-momentum transfer is the square root of the difference between this and the energy transfer, which is just  $p_p c$ . Using (1) we have

$$\begin{aligned} \text{square of four-momentum transfer} &= p_p^2 c^2 - q^2 c^2 \\ &= 2qM_p c^3 \\ &= 2 \cdot (1.5 \text{ GeV}/c) \cdot c \cdot (0.938 \text{ GeV}) = 2.81 \text{ GeV}^2. \end{aligned}$$

To turn an energy into a length you divide  $\hbar c$  by it:

$$\begin{aligned} \text{distance probed} &\approx \frac{\hbar c}{\text{four-momentum transfer}} \\ &= \frac{(6.6 \cdot 10^{-25} \text{ GeV s}) \cdot (3 \cdot 10^{23} \text{ fm/s})}{1.68 \text{ GeV}} \\ &= .12 \text{ fm} . \end{aligned}$$

This does not agree with Perkins' answer. What am I doing wrong?

## Problem 1.2

The flux of primary cosmic rays averaged over the Earth's surface is approximately  $1 \text{ cm}^{-2} \text{ s}^{-1}$ , and their average kinetic energy is 3 GeV. Show that the power transferred to the Earth from cosmic rays is about 2.5 gigawatt. (Earth radius=6400 km).

$$\text{Earth surface area} = 4\pi(6400 \text{ km})^2 = 5.1 \cdot 10^{18} \text{ cm}^2$$

Cosmic ray particle flux  $\approx 1$  per  $\text{cm}^2$  per second

$\rightarrow \approx 5.1 \cdot 10^{18}$  cosmic ray particles per second impinging on Earth's surface

Average particle energy 3 GeV

$\rightarrow \approx 1.5 \cdot 10^{28} \text{ eV/s}$  power delivered to earth surface

$$1.5 \cdot 10^{28} \text{ eV/s} \cdot \frac{1.6 \cdot 10^{-19} \text{ Joule}}{1 \text{ eV}} = 2.4 \cdot 10^9 \text{ Joule/s} \rightarrow 2.4 \text{ Gw}$$