

Energy - Momentum Tensor (p. 49)

(or Stress - Energy tensor) $g = \text{diag}(-1, 1, 1, 1)$

Given $\varphi: (\mathbb{R}^4, g) \rightarrow \mathbb{R}$, the energy-momentum tensor is defined as

$$T = d\varphi \otimes d\varphi - \frac{1}{2} (g^{ij}(d\varphi, d\varphi) - m^2 \varphi^2) g$$

(It's symmetric)

Its components:

$$T_{\alpha\beta} = T_{\mu\nu}$$

$$T_{\alpha\beta} = \varphi_{,\alpha} \varphi_{,\beta} - \frac{1}{2} g_{\alpha\beta} g^{rs} \varphi_{,r} \varphi_{,s} - \frac{1}{2} m^2 \varphi^2 g_{\alpha\beta}$$

Then $j^\alpha_{\text{energy}} = -T_\epsilon^\alpha \equiv -g^{\alpha\beta} T_{\epsilon\beta}$

{ energy density = $T^{\epsilon\epsilon}$

$$j^\alpha_{\text{momentum}} = -T_i^\alpha \equiv -g^{\alpha\beta} T_{i\beta}, i=1, 2, 3$$

Momentum density in direction i is $-T^{\epsilon i}$

Conservation of energy & momentum is encoded
in an identity:

$$g^{\beta\delta} \partial_\delta T_{\alpha\beta} = \nabla_\alpha (\square - m^2) \varphi$$



where $\square \varphi = g^{\alpha\beta} \partial_\alpha \partial_\beta \varphi$.

So, if $(\square - m^2) \varphi = 0$, then

$$g^{\beta\delta} \partial_\delta T_{\alpha\beta} = 0$$

i.e. The divergence of the energy-momentum tensor vanishes

Note: Change of reference frame mixes up
energy & momentum. So one often says, "Conservation
of energy-momentum", not just one of these.

A symmetry $f: \mathbb{R}^4 \rightarrow \mathbb{R}^4$ is a diff. s.t.
 $f \circ \hat{\varphi} = \varphi \circ f$, then the Lagrangian is preserved.

So if

$$L(x, \dot{x}, \partial\varphi) = L(x, \dot{x}, \partial\varphi)$$

then the Lagrangian is preserved. There
is a more general notion of symmetry
called divergence symmetry.

Since, $\hat{L} = L + (\text{div term})$, then

$\mathcal{E}(\hat{L}) = \mathcal{E}(L)$, unless which change the
Lagrangian by a divergence term are also
considered to be symmetries

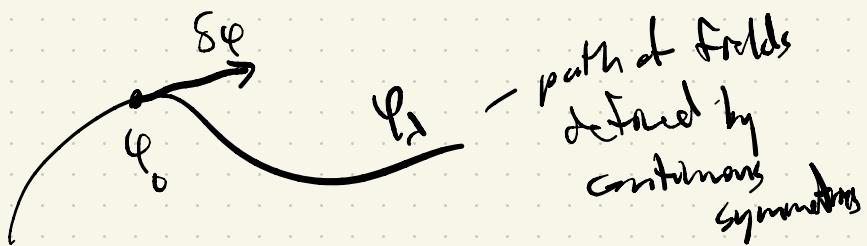
We typically study 1-param families of symmetries,

say $f_\lambda : \mathbb{R}^4 \rightarrow \mathbb{R}^4$ { let $\varphi_\lambda \doteq \varphi \circ f_\lambda$. Then

we want $\frac{dL(\varphi_\lambda, \dot{\varphi}_\lambda)}{d\lambda} = 0.$

Infinitesimal Symmetry (a vector field) p.55

Space of fields



$\delta\varphi \doteq \left. \frac{d\varphi_\lambda}{d\lambda} \right|_{\lambda=0}$. This $\delta\varphi$ is an infinitesimal symmetry.

View it as a vector field
in the space of fields

Since Lagrangian densities which differ by a divergence term gives the same Euler-Lagrange eqns, then we expand our definition of symmetry to also include $f: \mathbb{R}^4 \rightarrow \mathbb{R}^4$ ($\hat{\varphi} = \varphi \circ f$) s.t.

$$L(x, \hat{\varphi}, \partial_x \hat{\varphi}) = L(x, \varphi, \partial_x \varphi) + \partial_\alpha V^\alpha.$$

Eg. Consider time translation:

$$\varphi(t, x) \mapsto \varphi(t + \lambda, x)$$

means partial
derivative
wrt $\alpha \in \mathcal{B}$

$$\text{Then } L = -\frac{1}{2} (g^{\alpha\beta} \underbrace{\dot{\varphi}_\alpha \dot{\varphi}_\beta}_{\varphi_\alpha \varphi_\beta} + m^2 \varphi^2)$$

$$g^{\alpha\beta} = \text{diag}(-1, 1, 1, 1)$$

$$\text{Then } \delta L = -L g^{\alpha\beta} \dot{\varphi}_{,\alpha} \overset{\circ}{\varphi}_{,\beta} + m^2 \varphi \overset{\circ}{\varphi}$$

$$= \partial_\alpha \mathcal{L}$$

$$= \partial_\alpha (\delta^\alpha + \mathcal{L}). \quad \leftarrow \begin{array}{l} \text{Divergence} \\ \text{term} \end{array}$$

This means there is no preferred instant of time
in KGB theory

Noether's Thm

let $\mathcal{L} = \mathcal{L}(x, \varphi, \partial\varphi)$. If $S\varphi$ is an infinitesimal
variational symmetry,
then $S\mathcal{L} = 0$.

But also, at any pt in the space of fields

$$S\mathcal{L} = \varepsilon(\mathcal{L}) S\varphi + D_\alpha V^\alpha$$

$$\text{where } \varepsilon(\mathcal{L}) = \frac{\partial \mathcal{L}}{\partial \dot{\varphi}} - D_\alpha \left(\frac{\partial \mathcal{L}}{\partial \varphi_{,\alpha}} \right)$$

holds for
any field
variation

$$\left\{ \begin{array}{l} V^\alpha = \frac{\partial \mathcal{L}}{\partial \dot{\varphi}_{,\alpha}} S\varphi \end{array} \right.$$

$$\text{So if } S\mathcal{L} = 0, \text{ then } D_\alpha V^\alpha = -\mathcal{E}(I) S_q$$

This is exactly the type of identity needed for defining a conserved current V^α : If ψ satisfies the KG eqn, then $\mathcal{E}(I) = 0$ is so

$$\underline{D_\alpha V^\alpha = 0}.$$

$$\text{Or if } S\mathcal{L} = D_\alpha W^\alpha, \text{ then}$$

$$D_\alpha (V^\alpha - W^\alpha) = -\mathcal{E}(I) S_q$$

{ The conserved current is $V^\alpha - W^\alpha$.

Noether's First Thm:

If $S[\varphi(x, \eta, \partial\eta, \dots)]$ is a divergence symmetry of $\mathcal{L}(x, \eta, \partial\eta)$, i.e.

$$\delta \mathcal{L} = D_\alpha W^\alpha,$$

then there exists a conserved current given by

$$j^\alpha = \frac{\partial \mathcal{L}}{\partial \eta_{,\alpha}} \delta \varphi - W^\alpha.$$

Application: the time translation symmetry gives conserved current defining conservation of energy

Recall the time translation is a divergence symmetry:

$$\delta \varphi = \dot{\varphi} \Rightarrow \delta \mathcal{L} = D_\alpha (\delta^\alpha_t \mathcal{L})$$

Recognizing that $\delta q = \dot{q}$, $L' = S_{\epsilon}^{\alpha} L$, i

$$\frac{\partial L}{\partial q_{i,\kappa}} = -g^{\alpha\beta} q_{i,\beta},$$

we have: $j^{\alpha} = -g^{\alpha\beta} q_{i,\beta} q_{i,t} - S_{\epsilon}^{\alpha} L = -T_t^{\alpha}$

This j^{α} is a conserved current. We do the calculation $D_{\alpha} j^{\alpha} = 0$ under mass density conservation of energy.

We sometimes say the conserved current for energy is the Noether current associated to time translation symmetry

This was predictable since L , in the jet space, has no dependence on t , only on q, \dot{q} .

Space translations } Conservation of Momentum

let \hat{n} be a const^{unit} vect field on \mathbb{R}^3 . The space translation is given by $\psi(t, \vec{x}) \mapsto \psi_\lambda(t, \vec{x}) \doteq$

$$\vec{x} = (x, y, z) \quad \quad \quad \vec{x} + \lambda \hat{n}$$

$$\begin{aligned} \text{Then } S\psi &\doteq \frac{d\psi_\lambda}{d\lambda} \Big|_{\lambda=0} = \partial_x \psi_\lambda(t, \vec{x}) \cdot \hat{n}_x \\ &+ \partial_y \psi_\lambda(t, \vec{x}) \cdot \hat{n}_y \\ &+ \partial_z \psi_\lambda(t, \vec{x}) \cdot \hat{n}_z \\ &= \hat{n} \cdot \nabla \psi_\lambda(t, \vec{x}) \Big|_{\lambda=0} \\ &= \hat{n} \cdot \nabla \psi \\ &= n_i \cdot \psi_{,i} \end{aligned}$$

To check it is a symmetry, compute

$$\begin{aligned} S^2 &= S(\pm (\dot{\psi}^2 - |\nabla \psi|^2 - m^2 \psi^2)) \\ &= \dot{\psi} S\dot{\psi} - \nabla \psi \cdot D S \psi - m^2 \psi \cdot S \psi \\ &= \dot{\psi} (\hat{n} \cdot \nabla \psi) - \nabla \psi \cdot \nabla (\hat{n} \cdot \nabla \psi) - m^2 \psi (\hat{n} \cdot \nabla \psi) \\ &= \hat{n} (\dot{\psi} \cdot \nabla \psi - \nabla \psi \cdot \nabla \psi - m^2 \psi \cdot \nabla \psi) \end{aligned}$$

\hat{n} is
const

On the other hand:

$$\nabla \mathcal{L} = \dot{\varphi} \nabla i - \nabla \varphi \cdot \nabla^2 \varphi - m^2 \varphi \nabla \varphi$$

$$\text{So, } S_L = \hat{n} \cdot \nabla \mathcal{L} = \nabla (\hat{n} \cdot \mathcal{L})$$

$$= D_\alpha (w^\alpha) \quad \text{where}$$

$$\begin{matrix} \text{total} \\ \text{derivative} \end{matrix} \quad w^\alpha = (0, n^i L)$$

So space translation = divergence symmetry.

$$\text{Then, } j^\alpha = (p_i, j^i) = \frac{\partial \mathcal{L}}{\partial q_{i\alpha}} S_q - w^\alpha \quad \text{by} \\ \text{Noether's theorem}$$

$$\text{So } p = \dot{\varphi} S_q - 0 - \dot{\varphi} \hat{n} \cdot \nabla \varphi.$$

$$\text{Since } |\nabla \varphi|^2 = \varphi_{,1}^2 + \varphi_{,2}^2 + \varphi_{,3}^2, \text{ then}$$

$$j^i = -q_i \hat{n} \cdot \nabla \varphi - \frac{n^i}{2} L.$$

$$\rho = \bar{q} \hat{n} \cdot \nabla \varphi, j^i = -q_i \hat{n} \cdot \nabla \varphi + \frac{1}{2} \bar{n} (\|\nabla \varphi\|^2 - \dot{\varphi}^2 + m^2 \varphi^2)$$

Let's verify $\frac{d\rho}{dt} + \nabla \cdot \vec{j} = 0$.

$$\frac{d\rho}{dt} = \bar{q} \hat{n} \cdot \nabla \varphi + \cancel{\bar{q} \hat{n} \cdot \nabla \dot{\varphi}}$$

Once summed over i , these cancel.

$$\partial_i j^i = -q_i \hat{n} \cdot \nabla \varphi - q_i \hat{n} \cdot \nabla \varphi_i + \bar{n} (\nabla \varphi_i \cdot \nabla \varphi - \dot{q}_i \dot{\varphi}_i + m^2 \varphi_i \varphi_i)$$

$$\Rightarrow \nabla \cdot \vec{j} = -\nabla^2 \varphi \cdot \hat{n} \cdot \nabla \varphi - \bar{q} \hat{n} \cdot \nabla \dot{\varphi} + m^2 \varphi \cdot \hat{n} \cdot \nabla \varphi$$

$$\begin{aligned} \frac{d\rho}{dt} + \nabla \cdot \vec{j} &= \bar{q} \hat{n} \cdot \nabla \varphi - \nabla^2 \varphi \cdot \hat{n} \cdot \nabla \varphi + m^2 \varphi \hat{n} \cdot \nabla \varphi \\ &= \underbrace{-(\square - m^2) \varphi}_{0 \text{ since } \varphi \text{ is a KG field.}} \cdot \hat{n} \cdot \nabla \varphi = 0. \end{aligned}$$

Since \hat{n} is arbitrary, we actually have 3 indep conservation laws

for 3 linearly indep charges of \hat{n} .

Angular momentum.

Q: What symmetries conserve angular momentum?

A: Lorentz symmetries.

def. A Lorentz symmetry of \mathbb{R}^4 is a linear transf

$$x^\alpha \mapsto S_\beta^\alpha x^\beta$$

s.t. the quadratic form $g_{\alpha\beta} x^\alpha x^\beta = -t^2 + x_1^2 + x_2^2 + x_3^2$
is invariant.

$$\text{s.t. } S_\gamma^\alpha S_\delta^\beta g_{\alpha\beta} = g_{\gamma\delta} \quad (\#)$$

let $S(\lambda)$ be a 1-param family of Lorentz symm.

$$\text{s.t. } S_\beta^\alpha(0) = S_\beta^\alpha, \quad \omega_\beta^\alpha \doteq \left(\frac{\partial S_\beta^\alpha}{\partial \lambda} \right)_{\lambda=0}$$

Differentiating: $\omega_\gamma^\alpha g_{\alpha\beta} + \omega_\beta^\alpha g_{\alpha\beta} = 0$
(#) wrt λ : $\omega_\gamma^\alpha g_{\alpha\beta} + \omega_\beta^\alpha g_{\alpha\beta} = 0$
neglect

Define $\omega_{\alpha\beta} = g_{\beta\gamma} \omega_\alpha^\gamma$. Then Lorentz transf are generated by ω iff $\omega_{\alpha\beta} = -\omega_{\beta\alpha}$.

So the Lie algebra consists of 4×4 matrices X

s.t. $g X g = -X^T$ ($g = \text{diag}(-1, 1, 1, 1)$).

$$\text{Then } \delta \varphi = (\omega_\beta^\alpha x^\beta) \varphi_{,\alpha}$$

$$\delta I = D_\alpha (\omega_\beta^\alpha x^\beta I)$$

$$\left\{ \bar{j}^\alpha = \omega_{\beta\gamma} M^{\alpha(\gamma)}(x) \right.$$

$\underbrace{\phantom{\bar{j}^\alpha = \omega_{\beta\gamma} M^{\alpha(\gamma)}(x)}}$
conserved currents associated to
relativistic angular momentum

Clarification:

All continuous isometries of flat spacetime are contained in

The Poincaré group = diffeomorphisms generated by Lorentz transforms $\{$ spacetime translations,

Internal Symmetries

There are symmetries on the space of fields, not on spacetime. E.g. $\varphi \rightarrow -\varphi$.

Only unless $m=0$, there are no interesting continuous internal symmetries.

$$\text{e.g. } m=0, \psi_\lambda = \psi + \lambda$$

$\stackrel{0}{\parallel}$

$$\text{Then } S_\psi = 1 \quad ; \quad S_L = \dot{\psi} \delta \dot{\psi} - \nabla \psi \cdot \nabla S_\psi$$

$$\delta \dot{\psi} = \frac{d\dot{\psi}_\lambda}{d\lambda} \Big|_{\lambda=0} = 0$$

$$S_0 \quad j^\alpha = \frac{\partial L}{\partial \dot{\psi}_\alpha} \quad S_\psi - W^\alpha \quad \text{const?}$$

$D_\alpha W^\alpha$

$$\text{Seems } \rho = \dot{\psi} \Rightarrow \frac{d\rho}{dt} + \nabla \cdot \vec{j} = -\square \psi = 0$$

$$\vec{j} = -\psi_i$$

Since $m \neq 0$;

$$\therefore \square \psi = 0.$$

Charged KG Field & its internal symmetries.

A charged KG field is a \mathbb{C} -valued f^n

$$\psi: M \rightarrow \mathbb{C} \quad g_{\alpha\beta} = g^{\alpha\beta} = \text{diag}(-1, 1, 1, 1)$$

$$\text{The Lagrangian is now } \mathcal{L} = -g^{\alpha\beta} (\bar{\psi}_{,\alpha} \psi^*_{,\beta} + m^2 |\psi|^2)$$

We can write this Lagrangian as the sum of two \mathbb{R} -valued Lagrangians w/ masses $m_1 = m_2 = m$
} fields ψ_1, ψ_2 .

Then the charged field can be written as

$$\psi = \frac{1}{\sqrt{2}} (\psi_1 + i\psi_2).$$

So really, we can take the real & imaginary parts if we want: $\psi \& \psi^*$

The Euler-Lagrange eqns are now:

$$\mathcal{E}_\psi(L) = (\square - m^2)\psi^* = 0$$

$$\mathcal{E}_{\psi^*}(L) = (\square - m^2)\psi = 0$$

Working on \mathbb{C} gives new deg of freedom which allows us to introduce conserved electric charges.

In QFT, we get anti-particles.

The charged Lagrangian now admits the internal continuous symmetry $\psi_t = e^{it}\psi$, $\psi_t^* = e^{-it}\psi^*$.

These are sometimes called phase transformations or rigid $U(1)$ transformations.

See p. 66-7 for the conserved current - the outcome is

$$\tilde{j}^\alpha = -ig^{\alpha\beta} \left(\psi^* \psi_\beta - \psi \psi_{,\beta}^* \right)$$

Let $V \subset \mathbb{R}^3$ at a fixed time. Then the total $U(1)$ charge

$$Q = i \int_V (\psi^* \dot{\psi} - \psi \dot{\psi}^*) d^3x$$

Use this for modeling electric charge in electrodynamics
(can also use it to model charge-density currents w/ neutral currents in electroweak theory).

Of course, we can do all this for general groups w/ representations on vector space V .

Consider $\varphi: M \rightarrow V$ w/ internal symmetries given by $r: G \rightarrow \text{Aut}(V)$.

Or even, let V be a vector bundle & $\varphi: M \rightarrow V$ a section

Let $G = \text{SU}(2)$, $V = \mathbb{C}^2$.

An element $U \in \text{SU}(2)$ can be written as

$$U(\theta, n) = \cos\left(\frac{\theta}{2}\right) I + i \sin\left(\frac{\theta}{2}\right) n^i \sigma_i$$

$$n = (n^1, n^2, n^3), \quad \sum_{j=1}^3 (n^j)^2 = 1 \quad \begin{cases} & \\ & \end{cases}$$

Pauli
matrices

$$\sigma_1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma_2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma_3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

If $U(\lambda)$ is any 1-param family in $\text{SU}(2)$, and $U(0) = I$

$$\therefore \Psi_\lambda = U(\lambda)\varphi, \text{ then}$$

$$S\varphi = i\tau\varphi, \text{ where } \tau \text{ is a Hermitian, traceless}$$

2×2
matrix

$$\text{defined by } \tau = \frac{i}{2} \left(\frac{dU}{d\lambda} \right)_{\lambda=0}$$

$$\text{So } i\tau \in \text{SU}(2).$$

Of course, \mathcal{L} is a linear comb. of the ∂_i .

Hermitian inner prod on \mathbb{C}^n : $(\varphi_1, \varphi_2) = \varphi_1^* \varphi_2$

Of course, it is $U(n)$ invariant. So we build a lag.

$$\mathcal{L} = -[g^{\alpha\beta}(\varphi_{,\alpha}, \varphi_{,\beta}) + m^2(\varphi, \varphi)]$$

—
Claim: for symmetry $\delta\varphi = i\epsilon\delta$, the assoc conserved

$$\text{currents } j^\alpha = i g^{\alpha\beta} (\varphi_{,\beta}^* \tau \varphi - \varphi^* \tau \varphi_{,\beta}).$$

There are 3 indep conserved currents here b/c of τ .

The 3 conserved charges assoc w/ $SU(2)$ symm are

called ISO Spins

Claim: The converse to what we wrote as
Noether's 1st Thm is true:

For each conservation law for a system of
Euler-Lagrange eqns, there is a corresponding
symmetry of the Lagrangian.

For many types of field theories, we also have
1-1 correspondence b/w conservation laws &
symmetries of the Lagrangian.
(including KG field theory)

let $M = (\mathbb{R}^{2n}, \omega)$ (phase space) $\rightsquigarrow H: \mathbb{R}^{2n} \rightarrow \mathbb{R}$
 a Hamiltonian. Let $\gamma: [0, 1] \rightarrow \mathbb{R}^{2n}$ be a classical traj.

Then for any function F , $\{F, H\} = \frac{d}{dt} F(\gamma(t))$.

This is b/c we define $\{F, H\} := -(\nabla F)^T J_0 \nabla H$

$$= (\nabla F)^T \dot{\gamma}$$

$$= \frac{d}{dt} (F \circ \gamma)$$

More generally, since $\{, \}: C^\infty(M) \times C^\infty(M) \rightarrow C^\infty(M)$,

take another function G ; we can think of G as the infinitesimal generator of a transformation:

$$\delta F = \{F, G\}$$

Then, if $\delta H = \{H, G\} = 0$, we call G a symmetry of the Hamiltonian system.

$$\text{But } \dot{G} = \{G, H\} = -\{H, G\} = 0$$

So G is also a conserved quantity along class. traj.
i.e. A f^n on phase space generates a
symmetry iff it is conserved.