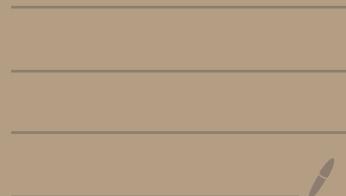


# Classical Field Theory - Charles Torre



A mechanical system is a dynamical system w/  
finitely many degrees of freedom.

A field is also a dynamical system but w/  
infinitely many degrees of freedom.

Mathematically, a field is a section of a fiber  
bundle.

Let  $\varphi: \mathbb{R}^4 \rightarrow \mathbb{R}$  be a scalar field

$$x^\alpha = (t, x, y, z)$$

The Klein-Gordon Equation is

$$\square \varphi - m^2 \varphi = 0$$

$$\text{where } \square = -\partial_t^2 + \partial_x^2 + \partial_y^2 + \partial_z^2$$

a wave operator called the d'Alembertian

When  $m=0$ , we get the wave eqn:

$$\Delta \varphi = \partial_t^2 \varphi$$

↑  
Laplacian

The KG eqn came about as an attempt to give a relativistic Schrödinger eqn but this did not work; after all, Schrödinger is mechanical, i.e. finitely many degrees of freedom

But KG is sort of a classical limit of a quantum field in relativistic settings

How to solve KG? Suppose  $\varphi$  is well behaved, such as

$$\varphi \in L^2(\mathbb{R}^3) \quad \forall t.$$

Taking a Fourier expansion:  $\varphi(t, r) = \left( \frac{1}{2\pi} \right)^{3/2} \int_{\mathbb{R}^3} \hat{\varphi}_k(t) e^{ik \cdot r} d^3 k$

We have  $\hat{\varphi}_{-\mathbf{k}} = \hat{\varphi}_{\mathbf{k}}^*$  since  $\varphi$  is real valued.

If  $\varphi$  satisfies  $\square \varphi = m^2 \varphi$ , then

$$\int -\partial_t^2 \left( \hat{\varphi}_{\mathbf{k}}(\mathbf{r}) \right) e^{i\mathbf{k} \cdot \mathbf{r}} d^3 r$$

$$+ \int \partial_x^2 \left( \hat{\varphi}_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{r}} \right) + \partial_y^2 \left( \hat{\varphi}_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{r}} \right) + \underline{\partial_z^2 \left( \hat{\varphi}_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{r}} \right)} d^3 r \\ = m^2 \int \hat{\varphi}_{\mathbf{k}} \cdot e^{i\mathbf{k} \cdot \mathbf{r}} d^3 r$$

$$\hat{\varphi}_{\mathbf{k}} \cdot \partial_z^2 (e^{i\mathbf{k} \cdot \mathbf{r}}) \\ = -k_z^2 \hat{\varphi}_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{r}}$$

$\Rightarrow$  In the integrands:

$$\left( -\ddot{\hat{\varphi}}_{\mathbf{k}} - \vec{\mathbf{k}} \cdot \vec{\mathbf{k}} \hat{\varphi}_{\mathbf{k}} \right) e^{i\mathbf{k} \cdot \mathbf{r}} = m^2 \hat{\varphi}_{\mathbf{k}} e^{i\mathbf{k} \cdot \mathbf{r}}$$

dot product

$$\Rightarrow \ddot{\hat{\varphi}}_{\mathbf{k}} + (k^2 + m^2) \hat{\varphi}_{\mathbf{k}} = 0.$$

norm square



This is easy to solve; it's just a 2nd order ODE of a simple form.

$$\hat{\psi}_k(t) = a_k e^{i\omega_k t} + b_k e^{-i\omega_k t}$$

$$\omega_k = \sqrt{k^2 + m^2}$$

Since  $\hat{\psi}_{-k} = \hat{\psi}_k^*$ , then  $b_{-k} = a_k^*$ .

$$\text{so } \hat{\psi}(x) = \left( \frac{1}{2\pi} \right)^{3/2} \int_{\mathbb{R}^3} \left( a_k e^{i(k \cdot r - \omega_k t)} + a_k^* e^{-i(k \cdot r - \omega_k t)} \right) d^3 r$$

KG field

Note: The solution of to KG is essentially an infinite collection of uncoupled harmonic oscillators for each  $k \in \mathbb{R}^3$

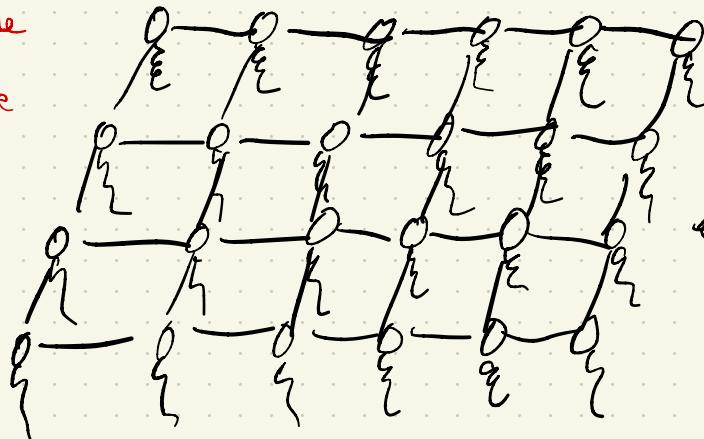
So there are infinite deg of freedom

Also, the KG field is often called the free or non-interacting

field b of the uncoupled nature of the oscillators

So picture an array of springs as a visual of a field

Discrete  
picture



$\psi$  satisfying  
 $\nabla^2 \psi = 0$

← nudge it here  
to "excite" the  
array. Then

a wave will  
propagate outwards.

This excitation which leads to a wave is what  
we call a particle. Particles are waves in QM.

$$\text{Let } L = \frac{1}{2} \int_R \left( \dot{\varphi}^2 - |\nabla \varphi|^2 - m^2 \varphi^2 \right) dx \quad (\text{Lagrangian})$$

$= \int \text{(Lagrangian density)}$

on space

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$$\{ S[\varphi] = \int_{t_1}^{t_2} L dt, \text{ (action)}, \text{ let } R = [t_1, t_2] \times \mathbb{R}^3$$

indude time now

Variation of  $S$ .

Let  $\lambda \in \mathbb{R}$  be a parameter  $\{ \varphi_\lambda \}$  a 1-parameter family of fields.

w/  $\varphi_0 = \varphi$ .

$$\delta S = \frac{dS[\varphi_\lambda]}{d\lambda} \Big|_{\lambda=0}. \quad \text{Also } \delta \varphi \equiv \frac{d\varphi_\lambda}{d\lambda} \Big|_{\lambda=0},$$

Then

$$\delta S = \int_R (\dot{\varphi} \delta \dot{\varphi} - \nabla \varphi \cdot \nabla \delta \varphi - m^2 \varphi \delta \varphi) dx.$$

Note:  $\nabla \cdot (\nabla \varphi \cdot \delta \varphi) = \nabla^2 \varphi \delta \varphi + \nabla \varphi \cdot \nabla \delta \varphi.$

Using integration by parts, note:

$$\int \vec{q} \cdot \vec{\delta q} = \vec{q} \cdot \vec{\delta q} - \int \vec{\delta q} \cdot \vec{q}.$$

$$SS = \int_R (-\vec{u} + \nabla^2 \varphi - m^2 \varphi) \delta q \, d^4x + \left[ \int_R \vec{q} \cdot \vec{\delta q} \right]_{t_1}^{t_2}$$

$$-\int_R \nabla \cdot (\nabla \varphi \delta q) \, d^4x.$$

$$= - \int_{t_1}^{t_2} dt \int_{r \rightarrow \infty} n \cdot \nabla \varphi \delta q \, d^3A \quad \text{(Divergence Theorem)}$$

If  $\varphi$  has cpt support or  $\rightarrow 0$  faster than  $\frac{1}{r^2}$ ,  
then the last term (boundary term) vanishes.  
Other boundary conditions, such as the endpoints  $\vec{q}_{t_1}$   
are fixed will make  $\delta q|_{t_1} = \delta q|_{t_2} = 0$ .

This will force the middle term to vanish.

i.e. w/ all these conditions,

$$SS = 0 \text{ when } \int_{\mathbb{R}} (\ddot{\varphi} + \nabla^2 \varphi - m^2 \varphi) S_\varphi \, dx = 0$$

$$\Rightarrow \square \varphi - m^2 \varphi = 0 \text{ everywhere in } \mathbb{R}.$$

So,  $S$  is the correct action for KG eqn

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Another way is via Euler-Lagrange eqn.

$$L = \frac{1}{2} \left( \dot{\varphi}^2 - |\nabla \varphi|^2 - m^2 \varphi^2 \right)$$

Treating  $x, \varphi, \varphi_x$  as formal variables (so  $L$  is in the  
(<sup>1<sup>st</sup></sup> Jet space))

$$\text{defn } E(L) = \underline{\frac{\partial L}{\partial \varphi}} - D_\alpha \underline{\frac{\partial L}{\partial \varphi_\alpha}}$$

total derivative

$$\text{Then } \frac{\delta S}{\delta \varphi} = \mathcal{E}(L) \Big|_{\varphi = \varphi(x)} = \square \varphi - m^2 \varphi.$$

In fact, if  $\hat{L} = \frac{1}{2} \varphi (\square - m^2 \varphi)$ , one can show

$$\mathcal{E}(L) = \mathcal{E}(\hat{L}) \quad \{ \quad L = \hat{L} + \text{divergence term}$$

↑  
sum of all the 1<sup>st</sup> partial derivatives of something.

Some Generalizations of KG.

$$\text{Let } \mathcal{L} = \frac{1}{2} (\dot{\varphi}^2 - \nabla \varphi / \square - m^2 \varphi^2) - \bar{j} \varphi$$

where  $\bar{j}: \mathbb{R}^4 \rightarrow \mathbb{R}$  is the "source" (in electromagnetism,  $j$  could be electric charge or current)

$$\text{Then } \mathcal{E}(L) = (\square - m^2) \varphi - \bar{j}.$$

In QFT, the presence of a source leads to particle creation/annihilation via transfer of energy-momentum from the field to its source.

The KG eqns are linear  $\{$  so the solutions are non-interacting.

If we modify KG to be non-linear, we introduce self-interaction.

$$\text{Eq. } \mathcal{L} = \frac{1}{2} (\dot{\varphi}^2 - |\nabla \varphi|^2 - m^2 \varphi^2) - V(\varphi)$$

$$\mathcal{E}(L) = (\square - m^2)\varphi - V'(\varphi) = 0.$$

So long as  $V$  is  $\nabla$  quadratic, this becomes non-linear.

One potential of interest:  $V(\varphi) = -\frac{1}{2} a^2 \varphi^2 + \frac{1}{4} b^2 \varphi^4$ .

If  $V$  is quad, then  $V'(\varphi) = a\varphi - b$

$$\{ \text{ so } (\square - (m^2 + a)\varphi) = b$$

Seems like, you just increase the mass... what does this mean, physically?

Coordinate free description: Let  $(M, g)$  be a Lorentzian mfd.

$$\text{Then } \mathcal{L} = -\frac{1}{2} \left( g^{-1}(d\varphi, d\varphi) + m^2 \varphi^2 \right) \varepsilon(g)$$

$\mathcal{L}$  is the Lagrangian density.

volume form  
of  $g$ .

This is called minimally coupled

Let  $\xi$  be a parameter of  $R(g)$  = scalar curvature. Then

$$\mathcal{L} = -\frac{1}{2} \left[ g^{ab}(d\varphi, d\varphi) + (m^2 + \xi R(g)) \varphi^2 \right] \varepsilon(g) \text{ gives}$$

Curvature coupled KG theory.

These theories are not "diffeomorphism invariant" or "generally covariant"; i.e.  $f: M \xrightarrow{\text{diffeo}} M$   $\Rightarrow \tilde{g} = f^* g$ , then

$$\tilde{\mathcal{L}} = -\frac{1}{2} \left( \tilde{g}^{-1}(d\varphi, d\varphi) + m^2 \varphi^2 \right) \varepsilon(\tilde{g}) \Rightarrow \text{a new}$$

Lagrangian density in general unless  $f$  is a symmetry.

If we allow  $\varphi$  to vary, we get 11 coupled non-linear field eqns instead of 1 linear field eqn.

Conservation Laws are fundamental & give info about complicated dynamics. Also, conservation laws are related to symmetries by Noether's theorem.

def: Let  $j^\alpha = j^\alpha(x, \varphi, \partial\varphi, \dots, \partial^k\varphi) \in \mathcal{J}^k$  be a vector field constructed as a local  $F^n$ .  $j^\alpha$  is a conserved current or defines a conservation law if the divergence of  $j^\alpha = 0$  when  $\varphi$  satisfies the field eqn.

e.g.  $D_\alpha j^\alpha = 0$  when  $(\square - m^2)\varphi = 0$ .

Explicitly, if  $\varphi(x)$  is a solution, take

$$\vec{j}^x(x) \in \vec{j}^x\left(x, \varphi(x), \frac{\partial \varphi(x)}{\partial x}, \dots, \frac{\partial^k \varphi(x)}{\partial x^k}\right) \text{ s.t.}$$

$$\frac{\partial}{\partial x^\alpha} \vec{j}^\alpha = 0.$$

If  $\vec{j}^\alpha = (j^0, j^1, j^2, j^3)$ , call  $\rho \stackrel{!}{=} j^0$  the density  
 $\vec{j} = (j^1, j^2, j^3)$  the

$$\hookrightarrow \Rightarrow \frac{\partial \rho}{\partial t} + \nabla \cdot \vec{j} = 0. \quad \begin{matrix} \text{current} \\ \text{density} \end{matrix}$$

The point of writing this is:

Let  $Q_V(t) = \int_V \rho(t, \vec{x}) d^3x$  be the total charge in region  $V$ .

$$\text{Then } \frac{d}{dt} Q_V(t) = - \int_V \nabla \cdot \vec{j} = - \int_{\partial V} \vec{j} \cdot \hat{n} dS$$

net flux

we say  $Q_V$  is conserved since we can see how it changes over time by purely in terms of the <sup>net</sup> flux, a fixed value. So there's no creation nor destruction of charge; it just moves around.

If we place boundary conditions, such as  $V = \mathbb{R}^3$ ; the field vanishes rapidly enough at  $\infty$ , then  $\frac{d}{dt} Q_V(t) = 0$  ; so the total charge is constant

### Conservation of Energy

$$\text{let } \tilde{j}^0 = \frac{1}{2} (\dot{\varphi}^2 + |\nabla \varphi|^2 + m^2 \varphi^2)$$

$$\tilde{j}^i = -\dot{\varphi}(\nabla \varphi)_i$$

$$\partial_0 \tilde{j}^0 = \dot{\varphi} \ddot{\varphi} + \nabla \varphi \cdot \nabla \dot{\varphi} + m^2 \varphi \cdot \dot{\varphi} \quad \Rightarrow \quad \partial_\alpha \tilde{j}^\alpha =$$

$$\partial_i \tilde{j}^i = -\nabla \dot{\varphi} \cdot \nabla \varphi - \dot{\varphi} \nabla^2 \varphi \quad \Rightarrow \quad -\dot{\varphi} (\square \varphi - m^2 \varphi)$$

Sum the Einstein notation

So, if  $\psi$  is a solution to KG,  $\partial_\alpha j^\alpha = 0$ .

$$\text{Let } E_V = \frac{1}{2} \int_V (\dot{\psi}^2 + |\nabla \psi|^2 + m^2 \psi^2) d^3x$$

$$\text{total energy} = T + U$$

$$\frac{1}{2} \int_V \dot{\psi}^2 d^3x \quad \frac{1}{2} \int_V (|\nabla \psi|^2 + m^2 \psi^2) d^3x$$

kinetic energy                      potential energy

$$\text{Then } \rho = j^0 \quad \vec{j} = (j^1, j^2, j^3) \quad \boxed{\nabla \cdot \vec{j} = (\nabla \cdot \vec{j})}$$

$$\frac{d}{dt} E_V = \int \partial_t \rho = - \int \nabla \cdot \vec{j} = - \int_V (\nabla \cdot \nabla \psi + i \dot{\psi} \nabla^2 \psi) d^3x$$

$$= - \int_V i \dot{\psi} \nabla \cdot \nabla \psi dV.$$

If  $\dot{\psi} = 0$  or the normal component  
of  $\nabla \psi$  to  $\partial V$  vanishes,

then  $\frac{d}{dt} E_V = 0$ . { Energy is conserved.