

# Riemann solver for non-linear Green-Naghdi II Thermal Equations

Suresh Chaudhary

Supervisor - Prof. Thomas Heuzé



Institut de Recherche en Génie Civil et Mécanique (GEM)  
Centrale Nantes - Computational Mechanics

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# State of the art

The previous work, Finite element approach to Green Naghdi model type II been solved but numerical oscillation observed due to hyperbolic equations[4].

We propose the **Riemann solver** to non linear thermal response Green Naghdi-II equations which are hyperbolic and formulated by potential function also describe the constitutive relation and can be solved numerically for FVM (Gudnov method or even higher order).

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# Objective

- To provide a Riemann solution to non linear thermal response Green Naghdi model type-II dissipationless.

# Introduction

- Existence of second sound - heat propagates as thermal wave
- Possibility of second sound in solids
- classical Fourier law of conduction is based on assumption
- the heat flux is an immediate response to a temperature change
- due to parabolic nature the infinite wave speed
- the Fourier law gives a macroscopic description of the microscopic phenomena associated with heat diffusion
- Cattaneo resolve the problem of infinite speed

# Maxwell-Cattaneo theory

- Cattaneo employs a time relaxed relation with the heat flux density and the temperature gradient.
- finite speed of propagation

$$\tau \frac{\partial q}{\partial t} + q = -k \frac{\partial T}{\partial x} \quad (1)$$

eliminating  $q$  leads to second order equation on  $T$  as the damped wave equation

$$\tau \frac{\partial^2 T}{\partial t^2} + \frac{\partial T}{\partial t} = \frac{k}{\rho C} \frac{\partial^2 T}{\partial x^2} \quad (2)$$

- leads to hyperbolic heat equation

# Green Naghdi model

- Thermal displacement - time integral of the temperature  $T$

$$\alpha(\mathbf{X}, t) := \int_{t_0}^t T(\mathbf{X}, \tau) d\tau + \alpha_0 \quad (3)$$

- it holds  $\dot{\alpha} = T$
- heat propagates as a thermal wave with finite speed
- 3 different model known as
  - type I - classical Fourier based
  - type II - dissipationless
  - type III - with dissipation
- this work focus on non-linear thermal response for type II

# Green Naghdi - type II

## Constitutive equations - non linear thermal response

- free energy  $W$  and internal energy  $U$

$$W(T, \beta) = \rho C \left( (T - T_0) - T \ln \left( \frac{T}{T_0} \right) \right) + \frac{1}{2} \frac{\hat{K}}{T_0} \beta^2$$

$$U(S, \beta) = \rho C T_0 \left( e^{\left( \frac{S}{\rho C} \right)} - 1 \right) + \frac{1}{2} \frac{\hat{K}}{T_0} \beta^2$$

- free energy  $W$  and internal energy  $U$  an potential for entropy  $S$  and entropy flux  $h$

$$S = -\frac{\partial W}{\partial T} = \rho C \ln \left( \frac{T}{T_0} \right)$$

$$h = \frac{\partial U}{\partial \beta} = \frac{\hat{K} \beta}{T_0}$$

with  $\beta$  thermal displacement gradient and  $T_0$  reference temperature



# Green Naghdi - II

## non linear thermal response

- The thermodynamic relation between heat flux  $q$  and entropy flux  $h$

$$q = T h$$

- The conservation equation

$$\dot{S} + \operatorname{div} h = \frac{\rho r}{T}$$

with no external source  $r = 0$

$$\boxed{\dot{S} + \operatorname{div} h = 0} \quad (4)$$

$$\boxed{\dot{\beta} + \operatorname{div} T = 0} \quad (5)$$

# Green Naghdi - II

## linear thermal response

- free energy  $W$  and internal energy  $U$  for linear thermal response

$$W(T, \beta) = -\frac{\rho C}{2T_0}(T - T_0)^2 + \frac{1}{2} \frac{\hat{K}}{T_0} \beta^2 \quad (6)$$

$$U(S, \beta) = -\frac{T_0}{2\rho C} S^2 + S_0 + \frac{1}{2} \frac{\hat{K}}{T_0} \beta^2 \quad (7)$$

the difference between linear and non linear thermal response depend on assumption[4].

The constitutive assumptions and relations leads to

$$\dot{S} + \operatorname{div} h = \frac{\rho r}{T}$$

we have  $\dot{S} = \frac{\rho C \dot{T}}{T}$

$$\frac{\rho C \dot{T}}{T} + \frac{\hat{K}}{T_0} \left( \frac{\partial \beta}{\partial x} \right) = \frac{\rho r}{T} \quad (8)$$

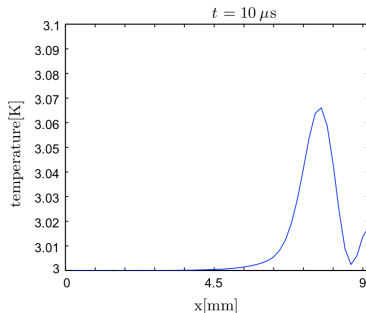
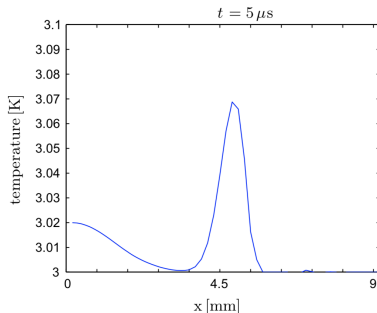
with  $T = \dot{\alpha}$

$$\frac{\rho C \ddot{\alpha}}{\dot{\alpha}} - \frac{\hat{K}}{T_0} \left( \frac{\partial^2 \alpha}{\partial x^2} \right) = \frac{\rho r}{T} \quad (9)$$

$$\frac{\rho C \ddot{\alpha}}{\dot{\alpha}} = \frac{\hat{K}}{T_0} \nabla^2 \alpha + \frac{\rho r}{T} \quad (10)$$

- a hyperbolic wave equation
- thermal displacement travel as a wave with finite speed with no dissipation (undamped wave).

The non linear thermal response Green Naghdi model type II been solved by FEM[10].



- plot shows temperature vs time
- thermal waves at finite speed
- numerical oscillations

# Riemann problem

The Initial Value Problem (IVP) for the conservation laws,

$$q_{,t} + (f(q))_{,x} = 0$$

$q$  as a vector of conserved variables and  $f(q)$  is the vector of fluxes with piecewise constant initial data,

$$q(x, 0) = q^{(0)}(x) = \begin{cases} q_L \\ q_R \end{cases}$$

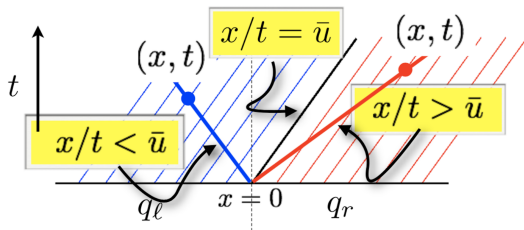
- solution based on characteristic field
- represent the one solution of hyperbolic system
- used in Godunov method based on FVM
- linear Riemann problem easy to solve but non linear complicated

# Linear riemann problem

$$q_t + \bar{u}q_x = 0$$

subject to initial conditions

$$q(x, 0) = \begin{cases} q_\ell & x < 0 \\ q_r & x > 0 \end{cases}$$



Solution :

$$q(x, t) = \begin{cases} q_\ell & x/t < \bar{u} \\ q_r & x/t > \bar{u} \end{cases}$$

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# non linear green naghdi II equations

The non linear Green naghdi II equations, given by

$$\frac{\partial \mathcal{S}}{\partial t} + \frac{\partial h}{\partial x} = 0 \quad (11)$$

$$\frac{\partial \beta}{\partial t} + \frac{\partial T}{\partial x} = 0 \quad (12)$$

is an example of a system of equations written in conservative form. More generally, we can write PDEs in conservative form as

$$q_{,t} + (f(q))_{,x} = 0$$

$q$  denotes a vector of the considered state variables,  $f(q)$  are flux function vector. The subscript,  $x$  denotes a partial differentiation with respect to  $x$

$$q_t + f(q)_x = 0$$

where

$$q = \begin{bmatrix} S \\ \beta \end{bmatrix} \quad f(q) = \begin{bmatrix} h \\ T \end{bmatrix} = \begin{bmatrix} \frac{\partial U}{\partial \beta} \\ \frac{\partial U}{\partial S} \end{bmatrix}$$

for smooth solutions, this can also be written as

$$q_t + f'(q) q_x = 0$$

where

$$f'(q) = \begin{bmatrix} \frac{\partial h}{\partial S} & \frac{\partial h}{\partial \beta} \\ \frac{\partial T}{\partial S} & \frac{\partial T}{\partial \beta} \end{bmatrix} = \begin{bmatrix} \frac{\partial^2 U}{\partial S \partial \beta} & \frac{\partial^2 U}{\partial \beta^2} \\ \frac{\partial^2 U}{\partial \beta \partial S} & \frac{\partial^2 U}{\partial S^2} \end{bmatrix}$$

is the flux Jacobian matrix.



Eigenvalues and eigenvectors of the flux Jacobian  $f'(q)$

$$\lambda_2^1 = \pm \sqrt{\frac{\widehat{K}}{\rho C}} e^{\left(\frac{s}{\rho C}\right)}$$

The eigenvectors are given by,

$$K^1 = \begin{bmatrix} \xi \\ \alpha \end{bmatrix} \quad K^2 = \begin{bmatrix} -\xi \\ \alpha \end{bmatrix}$$

with  $\xi = \frac{\sqrt{\widehat{K}\rho C}}{T_0 e^{\left(\frac{s}{2\rho C}\right)}}$

Eigenvalues and eigenvectors depend on  $q$  (i.e.  $S$ )

# Characteristic Field

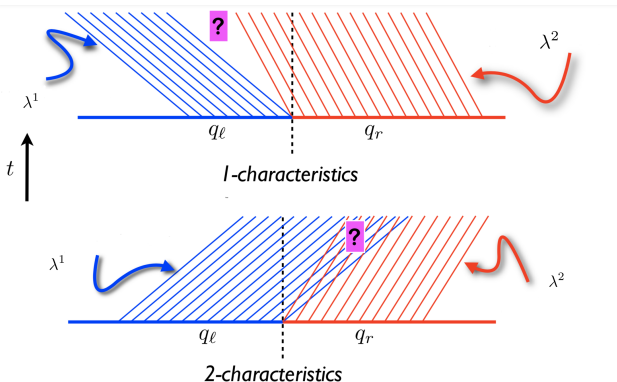
Characteristic Field- A characteristic field is defined by both the characteristic speed  $\lambda$  and the eigenvector  $K$  associated with it.

**linearly degenerate fields** -  $\nabla \lambda_i(q) \cdot K^{(i)}(q) = 0, \quad \forall q \in \mathbb{R}^m$   
**genuinely nonlinear fields** -  $\nabla \lambda_i(q) \cdot K^{(i)}(q) \neq 0, \quad \forall q \in \mathbb{R}^m$   
 where  $\mathbb{R}^m$  is the set of real-valued vectors of  $m$  components

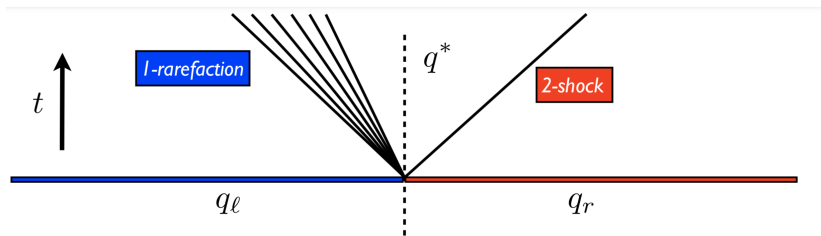
$$\nabla \lambda_1(q) \cdot K^{(1)}(q) = \nabla \lambda_2(q) \cdot K^{(2)}(q) = \frac{\hat{K}}{2\rho CT_0} \neq 0$$

Therefore both the characteristic field are **genuinely nonlinear**

# Problem with non linear



# Solving the non linear riemann problem



find a state  $q^*$  such that  $q_l$  is connected to  $q^*$  by physically correct 1-shock wave or 1-rarefaction, and  $q_r$  is connected to  $q^*$  by a physically correct 2-shock or 2-rarefaction.

find a state  $q^*$  that simultaneously satisfies both conditions. use a nonlinear root-finder to find the middle state  $q^*$ .

# Shock wave

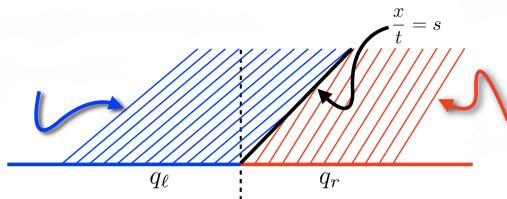
Consider the case where  $\lambda^p(q_l) > \lambda^p(q_r)$  :

- the two characteristic field will collide each other (i.e. shock)

This is the required jump condition across shocks. More generally we can write this condition as **Rankie Hugoniot jump condition**

$$f(q_r) - f(q_l) = s(q_r - q_l) \quad (13)$$

where  $s$  is the shock speed



# shock

The solution within the **left shock** hence satisfies :

$$\beta^* = \beta_L - \frac{2T_0}{\sqrt{K}} \sqrt{S_L - S^*} \sqrt{\exp\left(\frac{S_L}{\rho C}\right) - \exp\left(\frac{S^*}{\rho C}\right)} \quad (14)$$

The solution within the **right shock** hence satisfies :

$$\beta^* = \beta_R + \frac{2T_0}{\sqrt{K}} \sqrt{S_R - S^*} \sqrt{\exp\left(\frac{S_R}{\rho C}\right) - \exp\left(\frac{S^*}{\rho C}\right)} \quad (15)$$

# Rarefaction

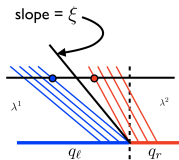
Consider the case where  $\lambda^p(q_l) < \lambda^p(q_r)$  :

- connects the left and right state through a smooth transition.
- this smooth transition can be computed using the Riemann invariants
- this invariants constant along any ray of the rarefaction.

$$\frac{dw_1}{k_1^i} = \frac{dw_2}{k_2^i} = \dots = \frac{dw_m}{k_m^i}$$

- gives the set of ODEs which are integrated between two states connected by the rarefaction to determine the states spanned within the rarefaction

slope  $\xi = \frac{x}{t}$  of the characteristic.  
we need to find  $q(\xi)$



# Rarefaction

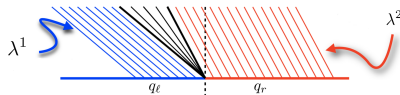
## 1-Riemann Invariant

The 1-Riemann invariant associated with the characteristic field.  
the ODE given,

$$\frac{d\left(\frac{S}{\xi} + \beta\right)}{-\xi} = \frac{d\left(-\frac{S}{\xi} + \beta\right)}{\alpha} \quad (16)$$

Integrating the above equation between star region and left region we get,

$$\beta^* = \beta_L - \int_{\left(\frac{S}{\xi}\right)_*}^{\left(\frac{S}{\xi}\right)_L} \left(\frac{\alpha - \xi}{\alpha + \xi}\right) d\left(\frac{S}{\xi}\right) \quad (17)$$





# Rarefaction

## 2-Riemann Invariants

similarly, 2-Riemann invariant ODE,

$$\frac{d\left(\frac{S}{\xi} + \beta\right)}{\xi} = \frac{d\left(-\frac{S}{\xi} + \beta\right)}{\alpha} \quad (18)$$

Integrating the above equation between right region and star region we get,

$$\int_{\left(\frac{S}{\xi}\right)_R}^{\left(\frac{S}{\xi}\right)_*} \left(\frac{\alpha + \xi}{\alpha - \xi}\right) d\left(\frac{S}{\xi}\right) + \int_{\left(\frac{S}{\xi}\right)_R}^{\left(\frac{S}{\xi}\right)_*} d\beta = 0 \quad (19)$$

$$\beta^* = \beta_R - \int_{\left(\frac{S}{\xi}\right)_R}^{\left(\frac{S}{\xi}\right)_*} \left(\frac{\alpha + \xi}{\alpha - \xi}\right) d\left(\frac{S}{\xi}\right) \quad (20)$$

it leads to,

# Rarefaction

The solution within the **left rarefaction** hence satisfies :

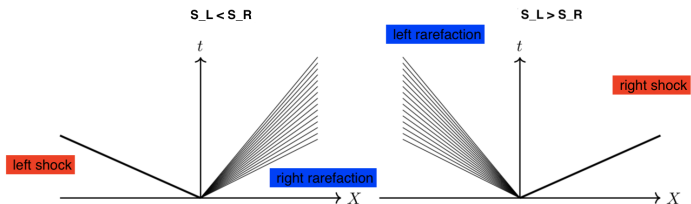
$$\begin{aligned}\beta^* = \beta_L - \frac{a}{b} & \left[ \frac{2b}{\alpha} \ln \left| \frac{\alpha + b \exp\left(\frac{S^*}{a}\right)}{\alpha + b \exp\left(\frac{S_L}{a}\right)} \right| - \frac{1}{\exp\left(\frac{S^*}{a}\right)} + \frac{1}{\exp\left(\frac{S_L}{a}\right)} + \frac{2b}{\alpha} (S^* - S_L) \right] \\ & + \frac{2a}{\alpha} \left[ \ln \left| \alpha + b \exp\left(\frac{S^*}{a}\right) \right| \frac{S^*}{a} - \ln \left| \alpha + b \exp\left(\frac{S_L}{a}\right) \right| \frac{S_L}{a} + \text{Li}_2 \left( -b \exp\left(\frac{S^*}{a}\right) \right) \right. \\ & \left. - \text{Li}_2 \left( -b \exp\left(\frac{S_L}{a}\right) \right) + \frac{1}{b} \left( \exp\left(\frac{-S^*}{a}\right) (a - S^*) - \exp\left(\frac{-S_L}{a}\right) (a - S_L) \right) - \left( \frac{(S^*)^2 - (S_L)^2}{a\alpha} \right) \right]\end{aligned}$$

The solution within the **right rarefaction** hence satisfies :

$$\begin{aligned}\beta^* = \beta_R - \frac{a}{b} & \left[ \frac{-2b}{\alpha} \ln \left| \frac{\alpha - b \exp\left(\frac{S^*}{a}\right)}{\alpha - b \exp\left(\frac{S_R}{a}\right)} \right| - \frac{1}{\exp\left(\frac{S^*}{a}\right)} + \frac{1}{\exp\left(\frac{S_R}{a}\right)} + \frac{2b}{\alpha a} (S^* - S_R) \right] \\ & - \frac{2a}{\alpha} \left[ \ln \left| \alpha - b \exp\left(\frac{S^*}{a}\right) \right| \frac{S^*}{a} - \ln \left| \alpha - b \exp\left(\frac{S_R}{a}\right) \right| \frac{S_R}{a} + \text{Li}_2 \left( b \exp\left(\frac{S^*}{a}\right) \right) \right. \\ & \left. - \text{Li}_2 \left( b \exp\left(\frac{S_R}{a}\right) \right) + \frac{1}{b} \left( \exp\left(\frac{-S^*}{a}\right) (a - S^*) - \exp\left(\frac{-S_R}{a}\right) (a - S_R) \right) + \left( \frac{(S^*)^2 - (S_R)^2}{a\alpha} \right) \right]\end{aligned}$$

## Riemann solution for GN-II

- the characteristic speed are monotone functions of the entropy  $S$
- case I - left shock and right rarefaction (1-shock 2-rarefaction)
- case II - left rarefaction and right shock (1-rarefaction 2-shock )



# Case I : $S_L > S_R$

## left rarefaction and right shock

The 1-rarefaction 2-shock solution one then seeks a state  $q^*$  that is connected to  $q_L$  and  $q_R$  through a shock wave and a rarefaction wave respectively. Hence,  $q$  must satisfy equations,

$$\begin{aligned}\beta^* &= \beta_R + \frac{2T_0}{\sqrt{K}} \sqrt{S_R - S^*} \sqrt{\exp\left(\frac{S_R}{\rho C}\right) - \exp\left(\frac{S^*}{\rho C}\right)} \\ \beta^* &= \beta_L - \frac{a}{b} \left[ \frac{2b}{\alpha} \ln \left| \frac{\alpha + b \exp\left(\frac{S^*}{a}\right)}{\alpha + b \exp\left(\frac{S_L}{a}\right)} \right| - \frac{1}{\exp\left(\frac{S^*}{a}\right)} + \frac{1}{\exp\left(\frac{S_L}{a}\right)} - \frac{2b}{\alpha} (S^* - S_L) \right] \\ &\quad + \frac{2a}{\alpha} \left[ \ln \left| \alpha + b \exp\left(\frac{S^*}{a}\right) \right| \frac{S^*}{a} - \ln \left| \alpha + b \exp\left(\frac{S_L}{a}\right) \right| \frac{S_L}{a} + \text{Li}_2 \left( -b \exp\left(\frac{S^*}{a}\right) \right) \right. \\ &\quad \left. - \text{Li}_2 \left( -b \exp\left(\frac{S_L}{a}\right) \right) \right] + \frac{1}{b} \left( \exp\left(\frac{-S^*}{a}\right) (a - S^*) - \exp\left(\frac{-S_L}{a}\right) (a - S_L) \right) - \left( \frac{(S^*)^2 - (S_R)^2}{a\alpha} \right)\end{aligned}$$

## Case II : $S_L < S_R$

left shock and right rarefaction

similarly, for case II  $q$  must satisfy equation

$$\begin{aligned}\beta^* &= \beta_L - \frac{2T_0}{\sqrt{K}} \sqrt{S_L - S^*} \sqrt{\exp\left(\frac{S_L}{\rho C}\right) - \exp\left(\frac{S^*}{\rho C}\right)} \\ \beta^* &= \beta_R - \frac{a}{b} \left[ \frac{-2b}{\alpha} \ln \left| \frac{\alpha - b \exp\left(\frac{S^*}{a}\right)}{\alpha - b \exp\left(\frac{S_R}{a}\right)} \right| - \frac{1}{\exp\left(\frac{S^*}{a}\right)} + \frac{1}{\exp\left(\frac{S_R}{a}\right)} + \frac{2b}{\alpha a} (S^* - S_R) \right] \\ &\quad - \frac{2a}{\alpha} \left[ \ln \left| \alpha - b \exp\left(\frac{S^*}{a}\right) \right| \frac{S^*}{a} - \ln \left| \alpha - b \exp\left(\frac{S_R}{a}\right) \right| \frac{S_R}{a} + \text{Li}_2 \left( b \exp\left(\frac{S^*}{a}\right) \right) \right. \\ &\quad \left. - \text{Li}_2 \left( b \exp\left(\frac{S_R}{a}\right) \right) \right] + \frac{1}{b} \left( \exp\left(\frac{-S^*}{a}\right) (a - S^*) - \exp\left(\frac{-S_R}{a}\right) (a - S_R) \right) + \left( \frac{(S^*)^2 - (S_R)^2}{a\alpha} \right)\end{aligned}$$

- The above equations represent the solution for non linear Green Naghdi type-II dissipationless
- These two equations to show the non linear scalar equation whose root is sought
- The problem is solved by optimization using SLSQP method
- Newton Raphson method failed due to constraint
- The Scipy library that solves constrained minimization problems
- material parameters extracted from paper[8]

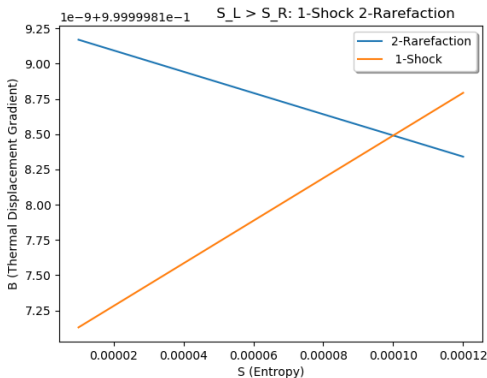
## Case I : $S_L > S_R$

- The non linear scalar equation whose root is to be solved by given left and right data and obtain star state.
- The different test case for right shock and left rarefaction with given left and right data,

Test Case	$S_L$	$S_R$	$\beta_L$	$\beta_R$	$S_*$	$\beta_*$
1.	0.009	0.006	1.0	1.0	0.00101038	0.99954
2.	0.0104	0.0100	0.0	0.0	$7.441401e^{-06}$	$1.5416999e^{-07}$
3.	0.0118	0.0114	0.0	-1.0	0.00107168	-0.00163836
4.	0.0121	0.0115	-1.0	-1.0	-0.00074492	-0.99973143
5.	0.0121	0.01	1.0	1.0	0.0001	1.0002685

# Case I : $S_L > S_R$

solution for left rarefaction and right shock



- The intersection between the two curves is the star state  $S_*$  and  $\beta_*$

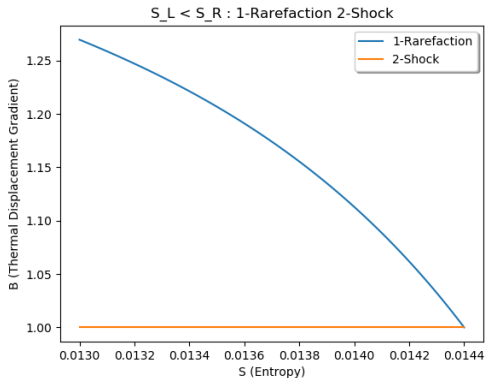


## case II : $S_L < S_R$

- similarly ,the root is to be solved by given left and right data and obtain star state.
- The different test case for left shock and right rarefaction

Test Case	$S_L$	$S_R$	$\beta_L$	$\beta_R$	$S_*$	$\beta_*$
1.	0.0182	0.0184	2.0	1.0	0.018197	5.0303712
2.	0.0192	0.0198	1.0	2.0	0.019196	45.4961
3.	0.0201	0.0206	0.0	0.0	0.020096	$-5.844 e^{-5}$
4.	0.0311	0.0318	-1.0	1.0	0.03109047	$-1.44158 e^{-5}$
5.	0.0142	0.0144	1.0	1.0	0.0141999	1.0017792

## Case II : $S_L < S_R$



- The intersection between the two curves is the star state  $S_*$  and  $\beta_*$

# Conclusion

- The Riemann solution proposed to non linear thermal response Green Naghdi-II thermal equations.
- The shock and rarefaction equations derived analytically.
- The solution is combination of a shock and a rarefaction(i.e. 1-shock 2-rarefaction or 1-rarefaction 2-shock).
- The solution of the star state requires to find numerically the root of a non linear scalar equation.

# Future scope

- The Riemann solution can be used for coupled problem (i.e. Thermoelasticity).
- Numerical simulation for FVM (Godunov method or even higher order).

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Thank You