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# Computer Vision Project

# Low Light Imagery

Team - Wanda Vision
TA Mentor – Pulkit Gera

# Inspiration

# Low Light Imaging

We derive inspiration from the CVPR 2018 paper "Learning to See in the Dark" which proposes a pipeline for processing low-light images, based on end-to-end training of a fully-convolutional network. The pipeline proposed can process raw sensor data as opposed to physical changes such as extending exposure time and using flash which can introduce blur and requires expensive hardware.







- A) Raw sensor data (ISO 8000)
- B) Image produced by camera using greater exposure (ISO 409,600)
- C) Image produced from the network proposed (example from paper)





**Top**: Image from the Sony α7S II sensor with short exposure **Bottom**: Output from the model as proposed in the paper



# DATASET: SID SONY



## See-in-the-Dark dataset (SID)

**Camera:** Sony α7S II

**Dataset Size:** 5094 pairs of short-exposure and long-exposure images

Image shape: 4240×2832

**Exposure for Input Images: 0.03 to 0.1 sec** 

**Exposure for Ground Truth Images: 10 sec** 

**Ground Truth (long exposure)** 

**Input Image (short exposure)** 





## **Dataset link:**

https://drive.google.com/file/d/1G6VruemZtpOyHjOC5N8Ww3ftVXOydSXx/view

# INSTRUCTIONS TO RUN

The Instructions to Run this repository locally have been written in the README file, a screenshot of which can be seen here.

#### Run locally:

· Clone the repository:

```
git clone https://github.com/Computer-Vision-IIITH-2021/project-wandavision.git
```

Navigate to the code folder:

```
cd see_in_dark
```

- Download and extract the dataset folders /long and /short in the Sony\_test/ folder.
- Create two new folders: mkdir saved\_model/

```
mkdir test result new/
```

#### Training from scratch:

- · Change the directory locations on the files as per your directory structure.
- Train the model from scratch: python train\_Sony.py
- Test the model: python test\_Sony.py

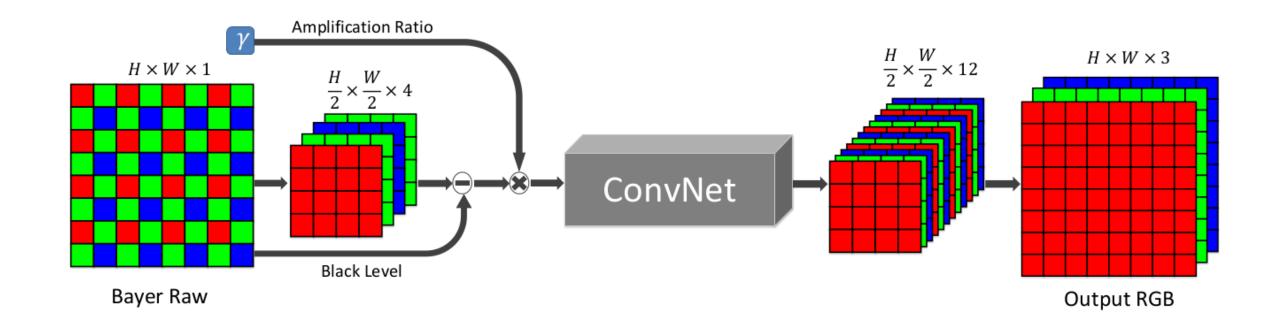
#### Using the pre-trained model:

- Pretrained model is saved as checkpoint\_sony\_e4000.pth under Sony\_test/saved\_model
- To test the pretrained model: python test\_Sony.py
- Test images are stored in test\_results\_Sony/ directory.

#### Get quantitative results:

- Run all cells of the notebook Quantitative\_results.ipynb changing the test directory path to where the test results are stored.
- The notebook generates PSNR and SSIM values between pairs of Ground Truth and output images.

# IMAGE PROCESSING PIPELINE



# CHOICE OF ARCHITECTURE & LOSS FUNCTION

## Why FCN?

End-to-end image processing pipeline that runs fast even on large raw images, does not require a fixed size input output like Fully Connected Layers

## Why UNet?

- less memory consumption
- doesn't use fully connected layers to work on small image patches

For other model and training parameters, we performed quantitative experiments and decided on the best choice of parameters, results for which will be shown in subsequent slides.

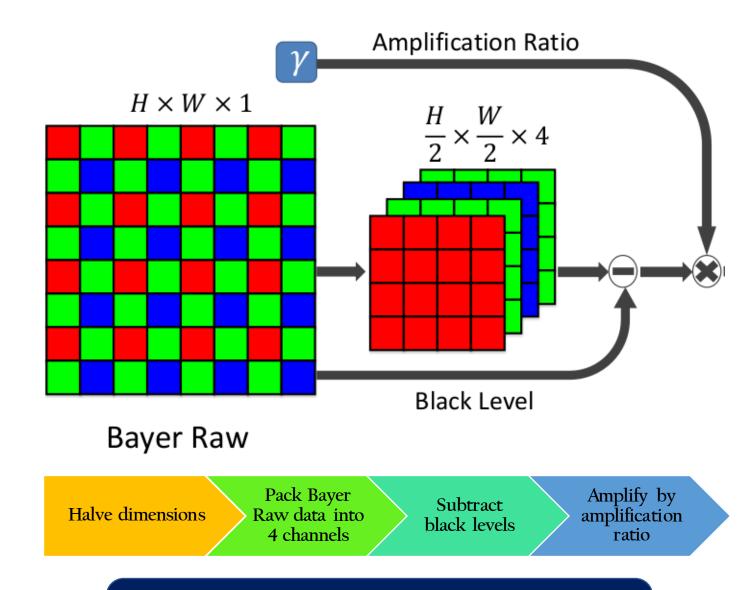


# PRE-PROCESSING

**Libraries Used** 

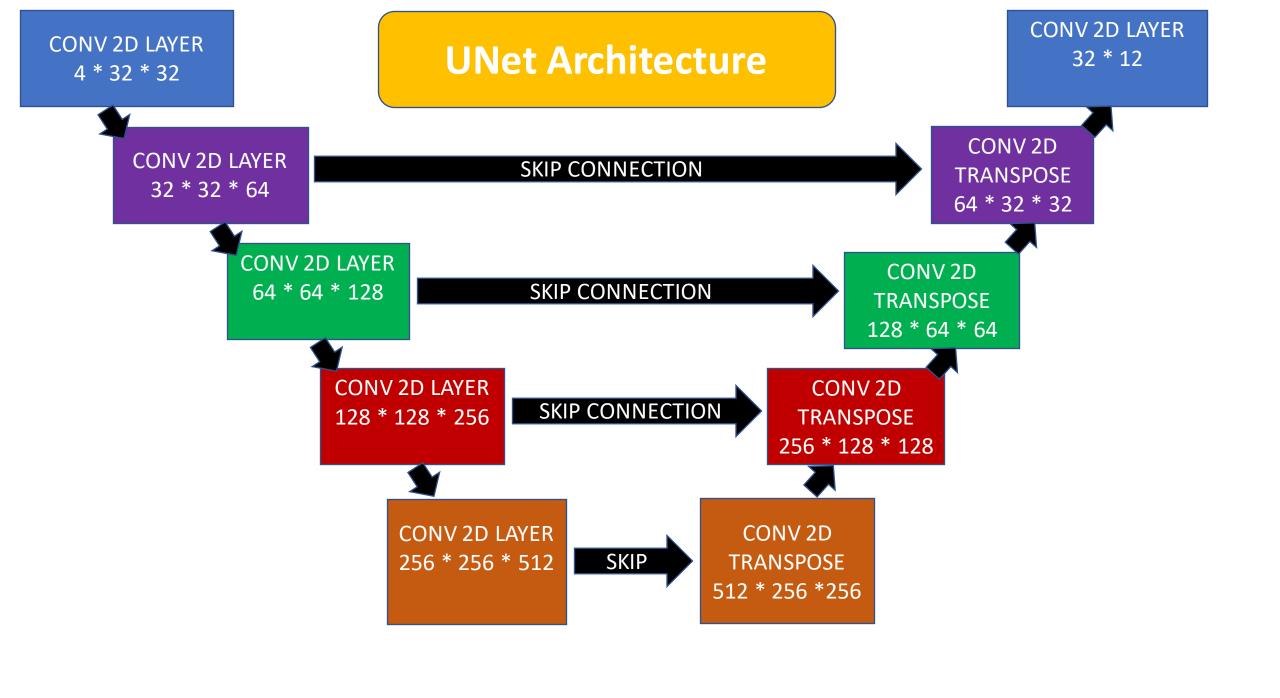
RawPy

**NumPy** 



## **Amplification Ratio**

= Ratio of exposure times between input and reference images



## MODEL CODE

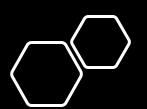
## **U-Net architecture:**

The first five convolution nets define the encoder architecture of the U-Net, such that we move from (4\*32) to (512\*512) 2D Convolutional Kernel.

The next 5 layers define the decoder architecture, along with the conserved skip connections

The size of convolution kernel is reduced from (512\*512) to (32\*12) using Transpose Convolutions(which halves the number of layers in the next stage), which finally gives us a 12-Layered Output.

```
def init (self, num classes = 10):
    super(SeeInDark, self). init ()
    self.conv1 1 = nn.Conv2d(4, 32, kernel size=3, stride=1, padding=1)
    self.conv1_2 = nn.Conv2d(32, 32, kernel_size=3, stride=1, padding=1)
    self.pool1 = nn.MaxPool2d(kernel size=2)
    self.conv2 1 = nn.Conv2d(32, 64, kernel size=3, stride=1, padding=1)
   self.conv2 2 = nn.Conv2d(64, 64, kernel size=3, stride=1, padding=1)
    self.pool2 = nn.MaxPool2d(kernel size=2)
    self.conv3 1 = nn.Conv2d(64, 128, kernel size=3, stride=1, padding=1)
    self.conv3 2 = nn.Conv2d(128, 128, kernel size=3, stride=1, padding=1)
    self.pool3 = nn.MaxPool2d(kernel size=2)
   self.conv4 1 = nn.Conv2d(128, 256, kernel size=3, stride=1, padding=1)
    self.conv4 2 = nn.Conv2d(256, 256, kernel size=3, stride=1, padding=1)
    self.pool4 = nn.MaxPool2d(kernel size=2)
   self.conv5 1 = nn.Conv2d(256, 512, kernel size=3, stride=1, padding=1)
    self.conv5 2 = nn.Conv2d(512, 512, kernel size=3, stride=1, padding=1)
    self.upv6 = nn.ConvTranspose2d(512, 256, 2, stride=2)
   self.conv6 1 = nn.Conv2d(512, 256, kernel size=3, stride=1, padding=1)
    self.conv6 2 = nn.Conv2d(256, 256, kernel size=3, stride=1, padding=1)
    self.upv7 = nn.ConvTranspose2d(256, 128, 2, stride=2)
   self.conv7_1 = nn.Conv2d(256, 128, kernel_size=3, stride=1, padding=1)
    self.conv7 2 = nn.Conv2d(128, 128, kernel size=3, stride=1, padding=1)
    self.upv8 = nn.ConvTranspose2d(128, 64, 2, stride=2)
   self.conv8 1 = nn.Conv2d(128, 64, kernel size=3, stride=1, padding=1)
   self.conv8 2 = nn.Conv2d(64, 64, kernel size=3, stride=1, padding=1)
    self.upv9 = nn.ConvTranspose2d(64, 32, 2, stride=2)
   self.conv9 1 = nn.Conv2d(64, 32, kernel size=3, stride=1, padding=1)
    self.conv9 2 = nn.Conv2d(32, 32, kernel size=3, stride=1, padding=1)
    self.conv10 1 = nn.Conv2d(32, 12, kernel size=1, stride=1)
```

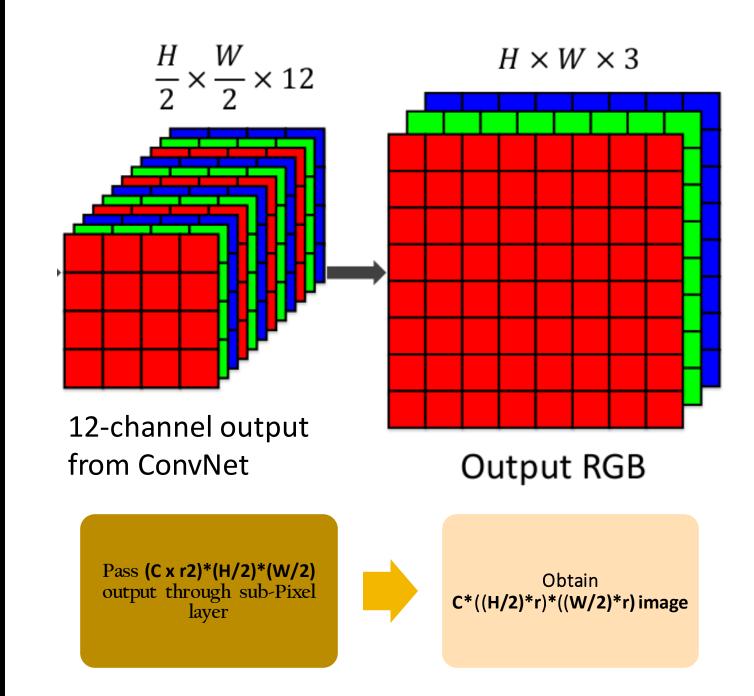


# POST-PROCESSING

**Libraries Used** 

**PyTorch** 

NumPy



TRAINING THE NETWORK

**Input:** short exposure images

Loss Function Used: L1 loss

Optimizer: ADAM

Batch: 512\*512 window from original image Data
Augmentation: Random flipping and rotation

Learning Rate: 10<sup>-4</sup> after 2000 epochs, 10<sup>-5</sup>

**Iterations:** 4000 epochs

**Ground truth:** longexposure images





# RESULTS OBTAINED

## TESTING AND EXPERIMENTS

Train & Validation

- We trained the model on 60 distinct image pairs of Ground Truth exposure 10s and input exposures of 0.1s, 0.04s and 0.033s.
- We validated our results on 10 more image pairs from the training set

Testing

• We tested our model on 51 distinct unseen test image pairs.

## Experiments

- Change ADAM optimizer to Adagrad
- Change L1 loss to L2 loss
- Change Leaky ReLU activation to ReLU activation

## **Relevant Links**

## Dataset Link:

• https://drive.google.com/file/d/1G6VruemZtpOyHjOC5N8Ww3ftVXOydSXx/view

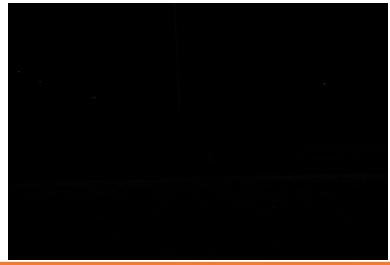
### Code Link:

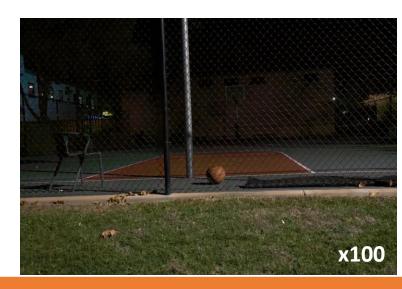
• https://github.com/Computer-Vision-IIITH-2021/project-wandavision

#### Results Link:

- Test images
- Note: there are images named \*\_100, \*\_250 and \*\_300 denoting the amplification ratio used.







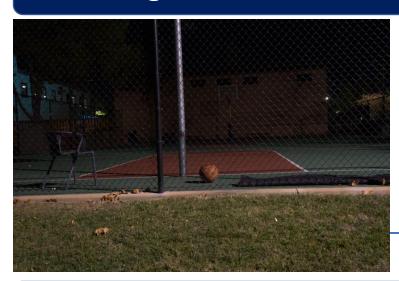
**Ground Truth** 

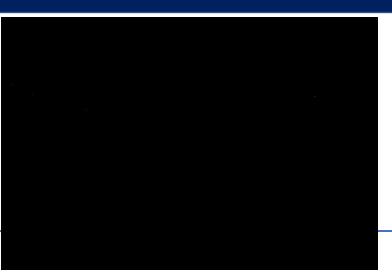
Input Image

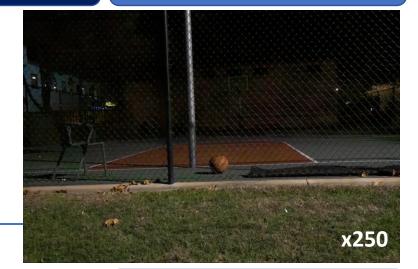
Output Image

**Test Image 1: Ground truth Exposure = 10s,Input Exposure = 0.1s** 

**PSNR = 31.02209** 



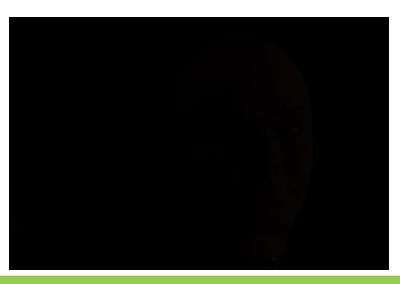




**Test Image 1: Ground truth Exposure = 10s,Input Exposure = 0.04s** 

**PSNR = 29.6558** 





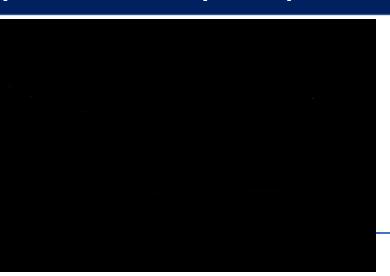


Ground Truth Input Image Output Image

**Test Image 2: Ground truth Exposure = 10s,Input Exposure = 0.1s** 

**PSNR = 33.34167** 



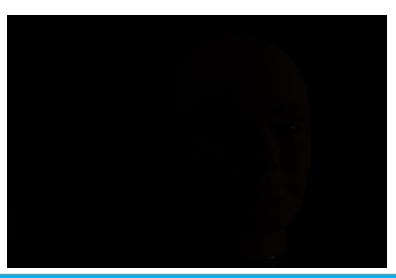




**Test Image 2: Ground truth Exposure = 10s,Input Exposure = 0.04s** 

**PSNR = 31.39138** 







Ground Truth Input Image Output Image

**Test Image 3: Ground truth Exposure = 10s,Input Exposure = 0.1s** 

**PSNR = 21.84601** 



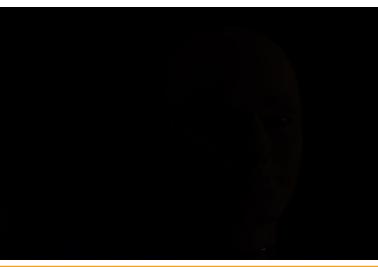




**Test Image 3: Ground truth Exposure = 10s,Input Exposure = 0.04s** 

**PSNR = 20.68907** 





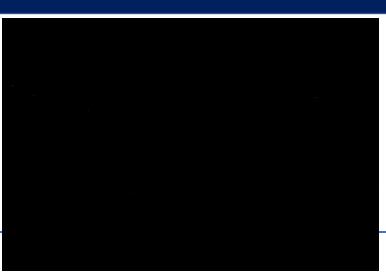


Ground Truth Input Image Output Image

**Test Image 4: Ground truth Exposure = 10s,Input Exposure = 0.1s** 

**PSNR = 32.50815** 



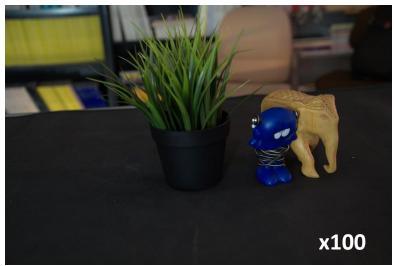


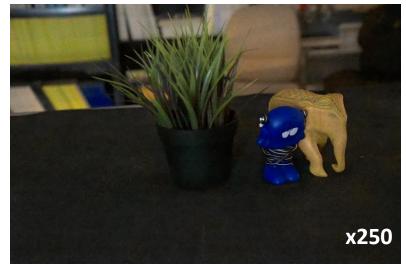


**Test Image 4: Ground truth Exposure = 10s,Input Exposure = 0.04s** 

**PSNR = 30.23718** 









With input exposure = 0.04s

With input exposure = 0.033s

PSNR = 30.882503

PSNR = 24.537966

PSNR = 23.445992

# **QUANTIFYING RESULTS**

## **Peak Signal-to-Noise Ratio(PSNR)**

Peak Signal to Noise Ratio is a ratio between the maximum signal power and the power of corrupting noise that affects the fidelity of the representation.

Since the signals have a very dynamic and wide range, PSNR is usually expressed as a logarithmic quality using the decibel scale.

If we define two images as I and K, then MSE between two images is defined as-

$$MSE = \frac{1}{m \times n} \sum_{i=0}^{m-1} \sum_{j=0}^{n-1} [I(i,j) - K(i,j)]^{2}$$

MAXI represents the maximum pixel value in the image. Then finally, the PSNR can be defined as-

$$PSNR = 10 \times \log_{10}(\frac{MAX_I^2}{MSE})$$

# Structural Similarity Index Measure (SSIM)

## What is SSIM?

SSIM is a method for predicting the perceived quality of images and considers the structural similarity of two images, as opposed to PSNR, which is an absolute metric.

A Higher SSIM value indicates better structural similarity between two images.

## SSIM formula for images x (Img1) and y (Img2)

$$SSIM(x,y) = \frac{(2\mu_x \mu_y + c_1)(2\sigma_{xy} + c_2)}{(\mu_x^2 + \mu_y^2 + c_1)(\sigma_x^2 + \sigma_y^2 + c_2)}$$

## **Notation used above:**

 $\mu_x$  - The mean of pixels of image 1

 $\mu_y$  - The mean of pixels of image 2

 $\sigma x^2$  - The variance of pixels of image 1

 $\sigma y^2$  - The variance of pixels of image 2

 $\sigma xy$  - The co-variance between pixels of image 1 and image 2

 $c_1 = (k_1 L)^2$  and  $c_2 = (k_2 L)^2$  - The constants to stabilise a weak denominator

L - The entire range of values of pixels, [0, 255]

 $k_1 = 0.01$  and  $k_2 = 0.03$ 

## **Code snippets for PSNR & SSIM**

## SSIM

```
def ssim(img1, img2):
    C1 = (0.01 * 255)**2
    C2 = (0.03 * 255)**2
    img1 = img1.astype(np.float64)
    img2 = img2.astype(np.float64)
    kernel = cv2.getGaussianKernel(11, 1.5)
    window = np.outer(kernel, kernel.transpose())
    mu1 = cv2.filter2D(img1, -1, window)[5:-5, 5:-5]
    mu2 = cv2.filter2D(img2, -1, window)[5:-5, 5:-5]
    mu1 sq = mu1**2
    mu2 sq = mu2**2
    mu1 mu2 = mu1 * mu2
    sigmal sq = cv2.filter2D(imq1**2, -1, window)[5:-5, 5:-5] - mu1 sq
    sigma2 \ sq = cv2.filter2D(img2**2, -1, window)[5:-5, 5:-5] - mu2 \ sq
    sigma12 = cv2.filter2D(imq1 * imq2, -1, window)[5:-5, 5:-5] - mu1 mu2
    ssim map = ((2 * mu1 mu2 + C1) * (2 * sigma12 + C2)) / ((mu1 sq + mu2 sq + C1) *
                                                             (sigmal sq + sigma2 sq + C2))
    return ssim map.mean()
def calc ssim(img1, img2):
    if not imgl.shape == img2.shape:
        raise ValueError('Input images must have the same dimensions.')
    if imql.ndim == 2:
        return ssim(img1, img2)
    elif imal.ndim == 3:
        if img1.shape[2] == 3:
            ssims = []
            for i in range(3):
                ssims.append(ssim(img1, img2))
            return np.array(ssims).mean()
        elif img1.shape[2] == 1:
            return ssim(np.squeeze(img1), np.squeeze(img2))
    else:
        raise ValueError('Wrong input image dimensions.')
```

## **PSNR**

```
def calc_psnr(img1, img2):
    img1 = img1.astype(np.float64)
    img2 = img2.astype(np.float64)
    if img1.ndim == 2:
        mse = np.mean((img1 - img2)**2)
    elif img1.ndim == 3:
        mse1 = np.mean((img1[:,:,0] - img2[:,:,0])**2)
        mse2 = np.mean((img1[:,:,1] - img2[:,:,1])**2)
        mse3 = np.mean((img1[:,:,2] - img2[:,:,2])**2)
        mse = (mse1 + mse2 + mse3)/3
    if mse == 0:
        return float('inf')
    psnr = 20 * math.log10(255.0 / math.sqrt(mse))
    return psnr
```

# QUANTITATIVE RESULTS OF EXPERIMENTS

EXP. NO.	OPTIMIZER	LOSS FUNCTION	ACTIVATION FUNCTION	PSNR	SSIM
1	ADAM	L1 LOSS	LEAKY RELU	24.481	0.6489
2	ADAGRAD	L1 LOSS	LEAKY RELU	21.587	0.6251
3	ADAM	L2 LOSS	LEAKY RELU	23.698	0.6045
4	ADAM	L1 LOSS	RELU	22.668	0.6102
Pretrained best model: GitHub model					

## LIMITATIONS

Loss of some hues in the output images with very low exposure

Amplification ratio is chosen externally based on ratio of exposure times.

Currently the model is trained on only images from the Sony  $\alpha$ 7S II camera

Not suited for realtime applications at full resolutions The SID dataset is limited in that it does not contain humans and dynamic objects.

## REFERENCES

- 1. C. Chen, Q. Chen, J. Xu and V. Koltun, "Learning to See in the Dark," 2018 IEEE/CVF Conference on Computer Vision and Pattern Recognition, 2018, pp. 3291-3300, doi: 10.1109/CVPR.2018.00347.
- 2. Original Project Website-

http://cchen156.github.io/SID.html

3. Demo YouTube Video for the original Project-

https://youtu.be/qWKUFK7MWvg

**4. PSNR:** 

https://en.wikipedia.org/wiki/Peak signal-to-noise ratio

**5. SSIM:** 

https://en.wikipedia.org/wiki/Structural similarity

