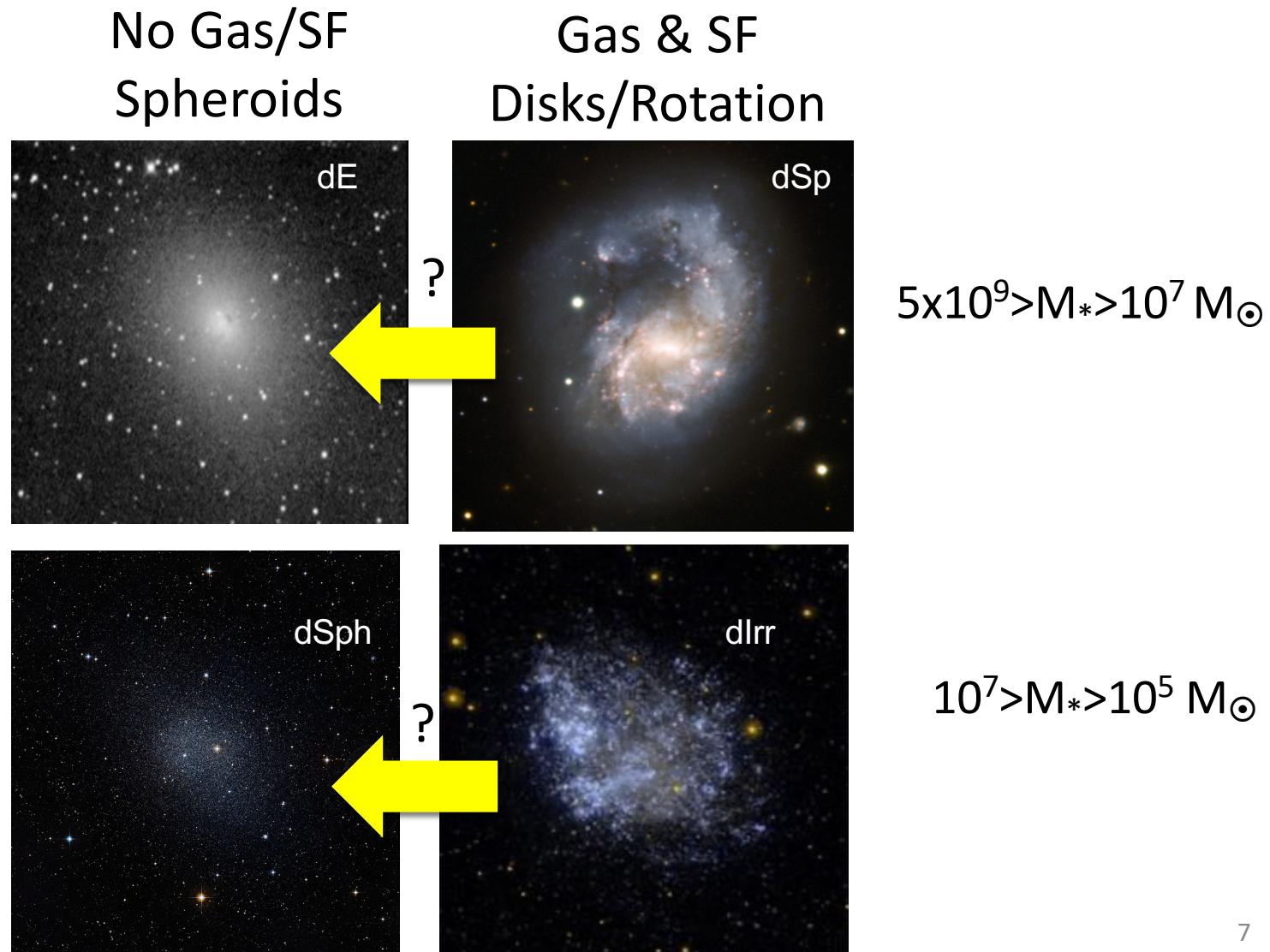
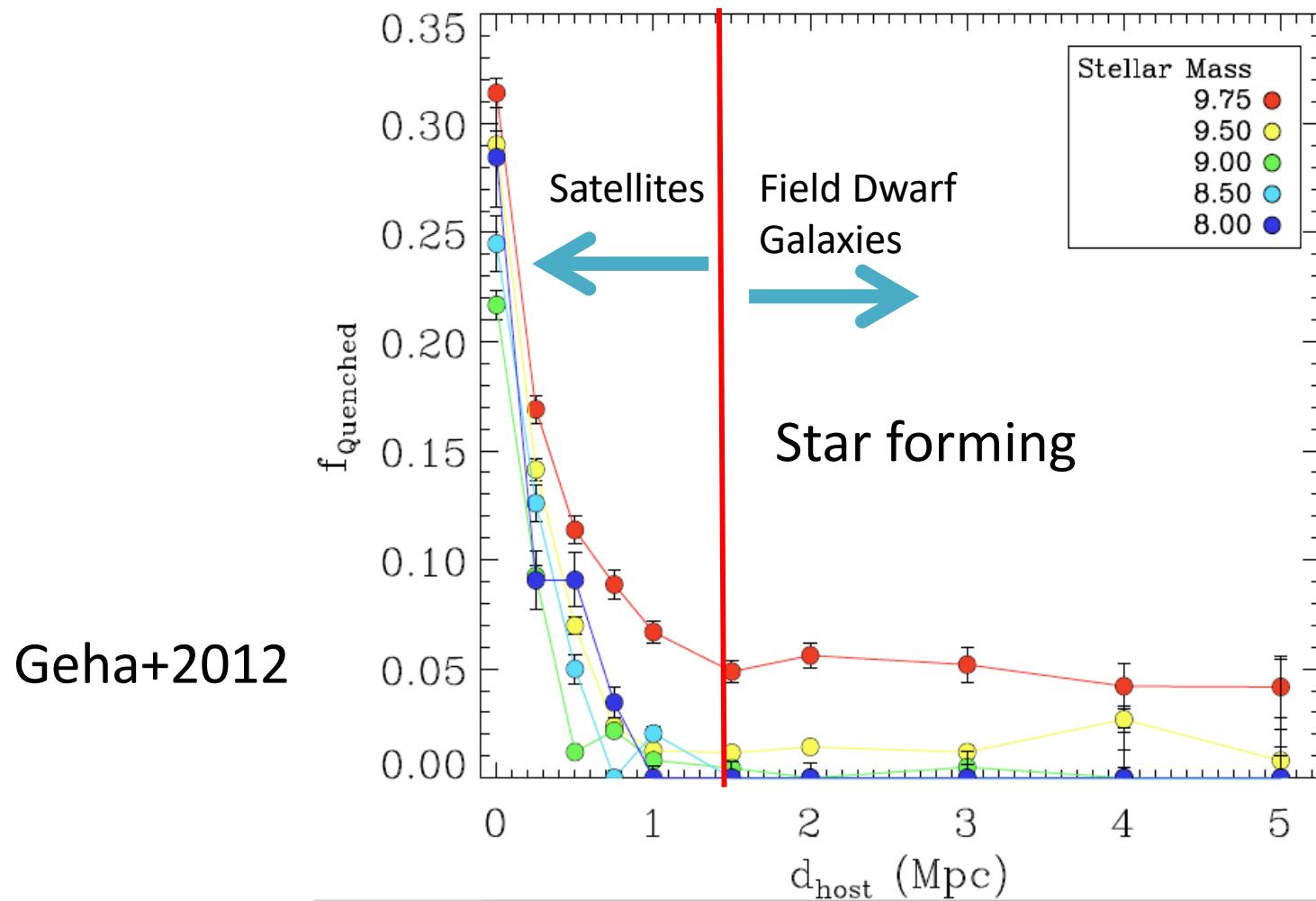


Recall: The Dwarf Galaxy Zoo: SO HOW DOES THIS TRANSITION HAPPEN?



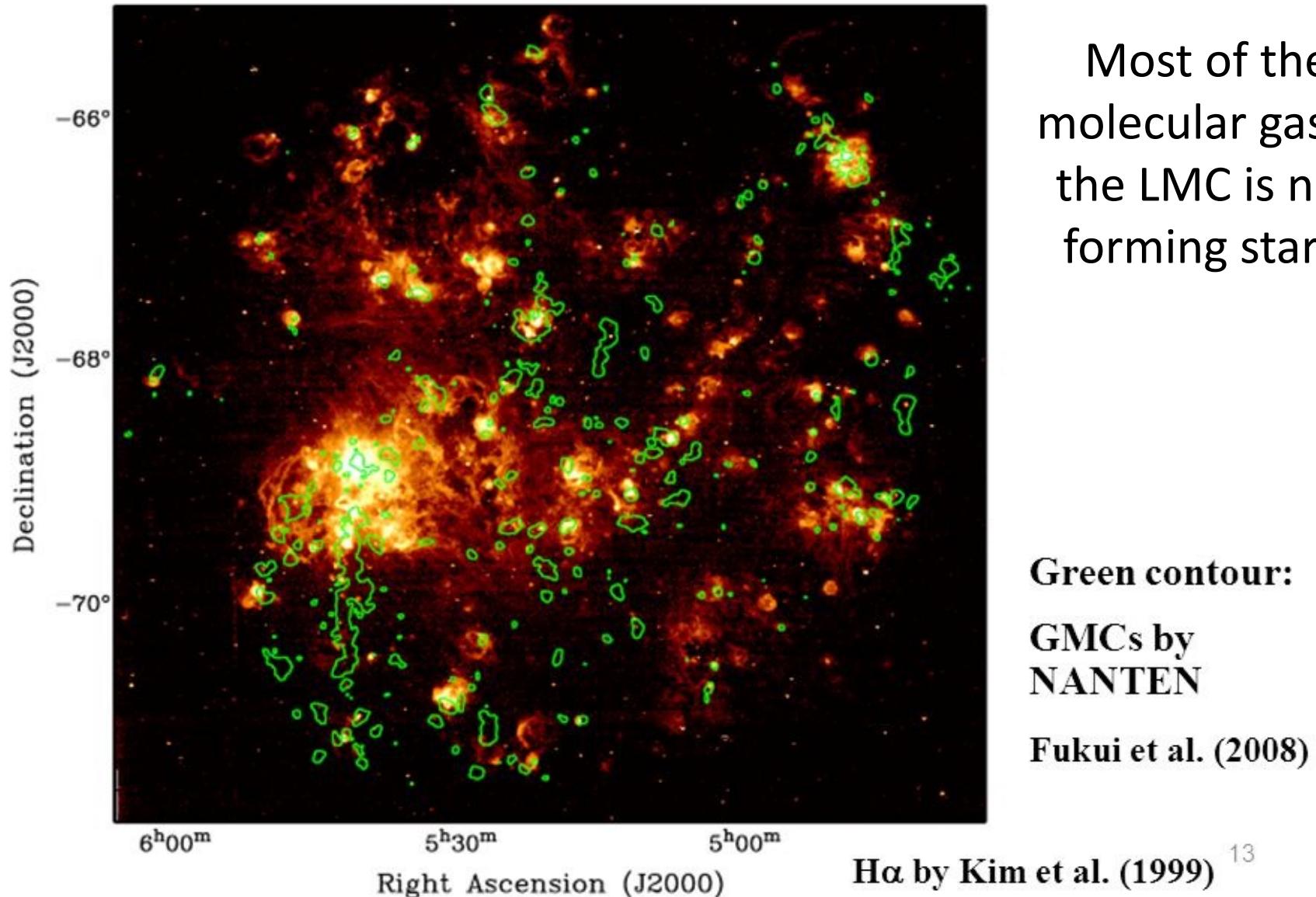
Recall: Dwarf galaxies are gas rich & forming stars
UNLESS they are close to a massive galaxy

Dwarfs ($10^8) DON'T quench via secular processes$



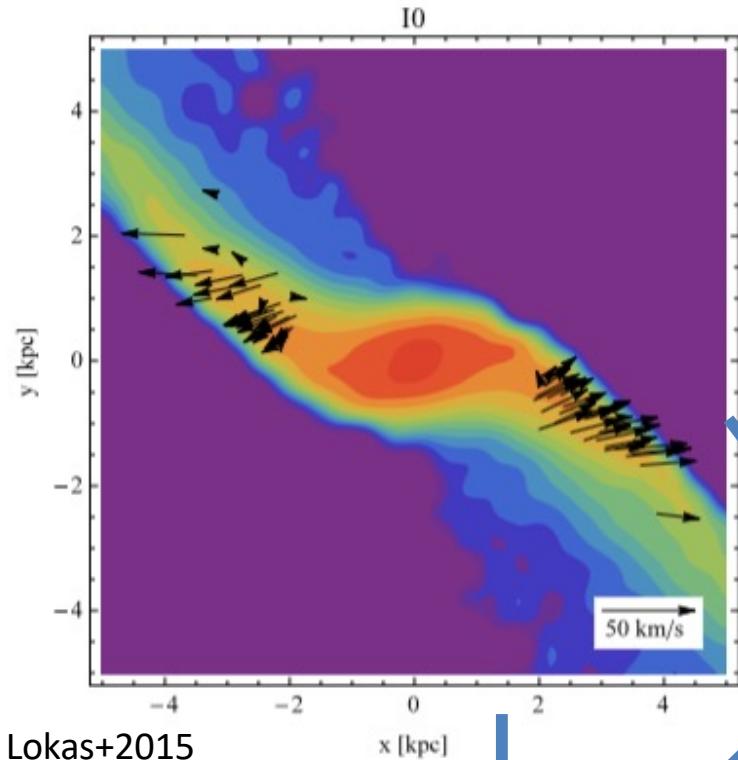
Star formation is inefficient at low metallicity

LMC: CO and H α

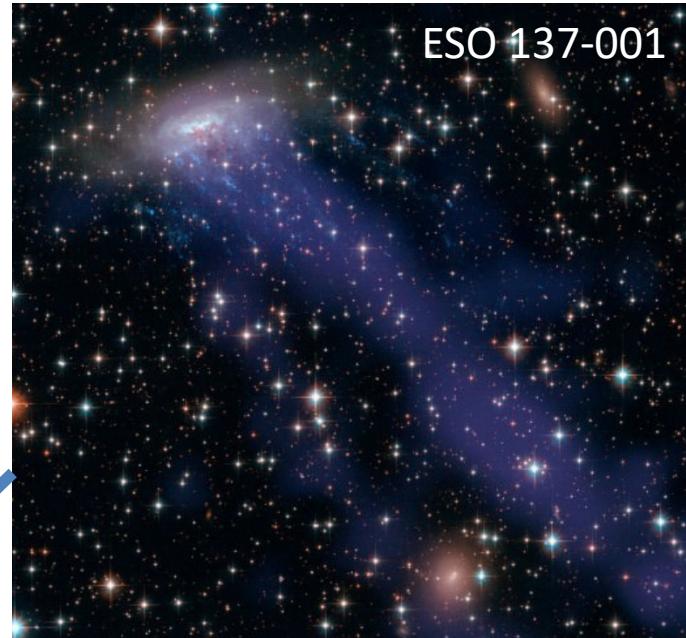


So why are the majority of
dwarf galaxies near MW type
hosts quenched?

TIDAL STRIPPING

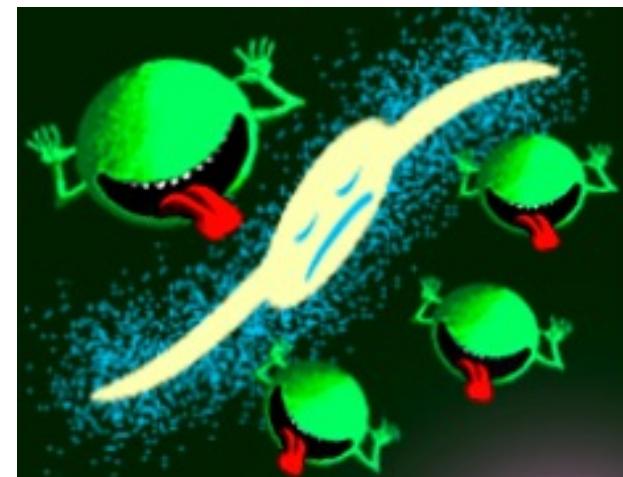


RAM PRESSURE STRIPPING



+

HARRASSMENT



STRANGULATION/STARVATION
“out of fuel”
(gas supply cut off, stars form passively)



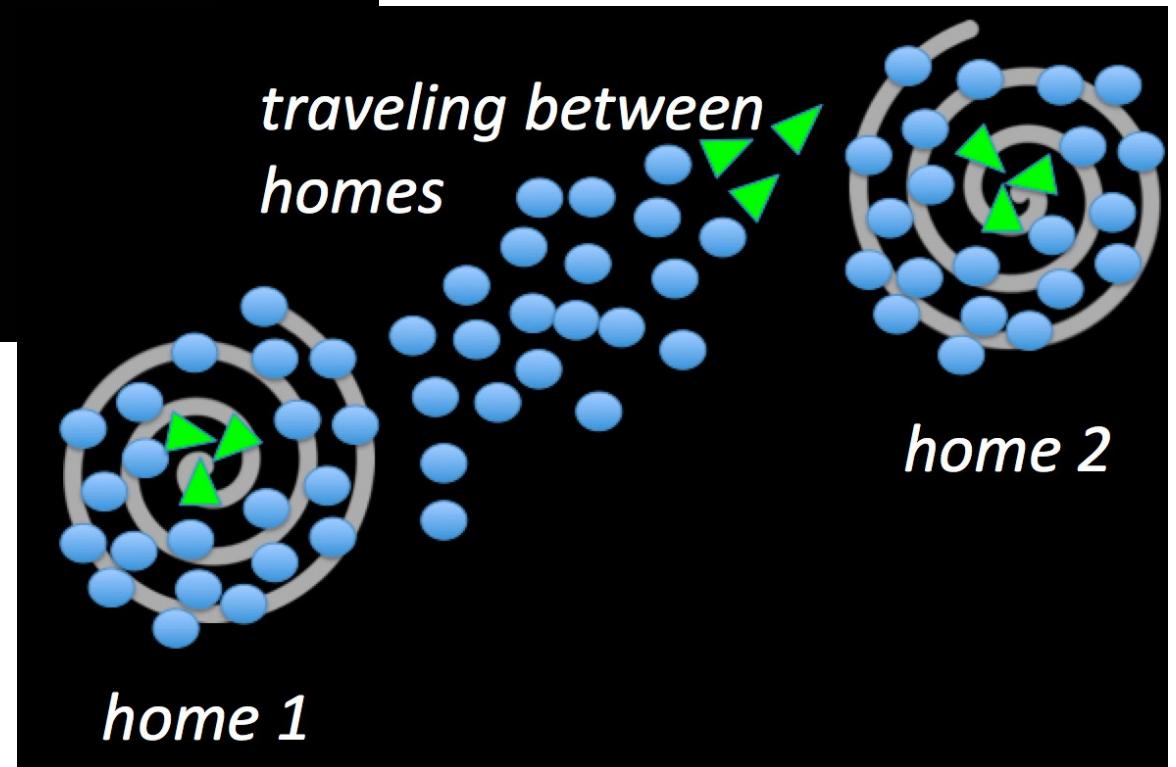
images: Dr. Aeree Chung



Simulation of interacting spiral galaxies.
Image adapted from Moreno + 2018



Dr. Jorge Moreno



A diagram of dancing that occurs during Hatun Puncha, where groups of dancers and musicians travel from home to home, dancing in spirals that change directions. Smaller groups merge to form larger dance troupes throughout the festival.

Diagram by Jessie M. Vallejo

Jorge Moreno's Suggested Change of Language

hierarchical scenario

ancestral scenario

devouring

including

cannibalism

collectivism

harassment

communion

stripping

sharing

strangulation

collaboration

starvation

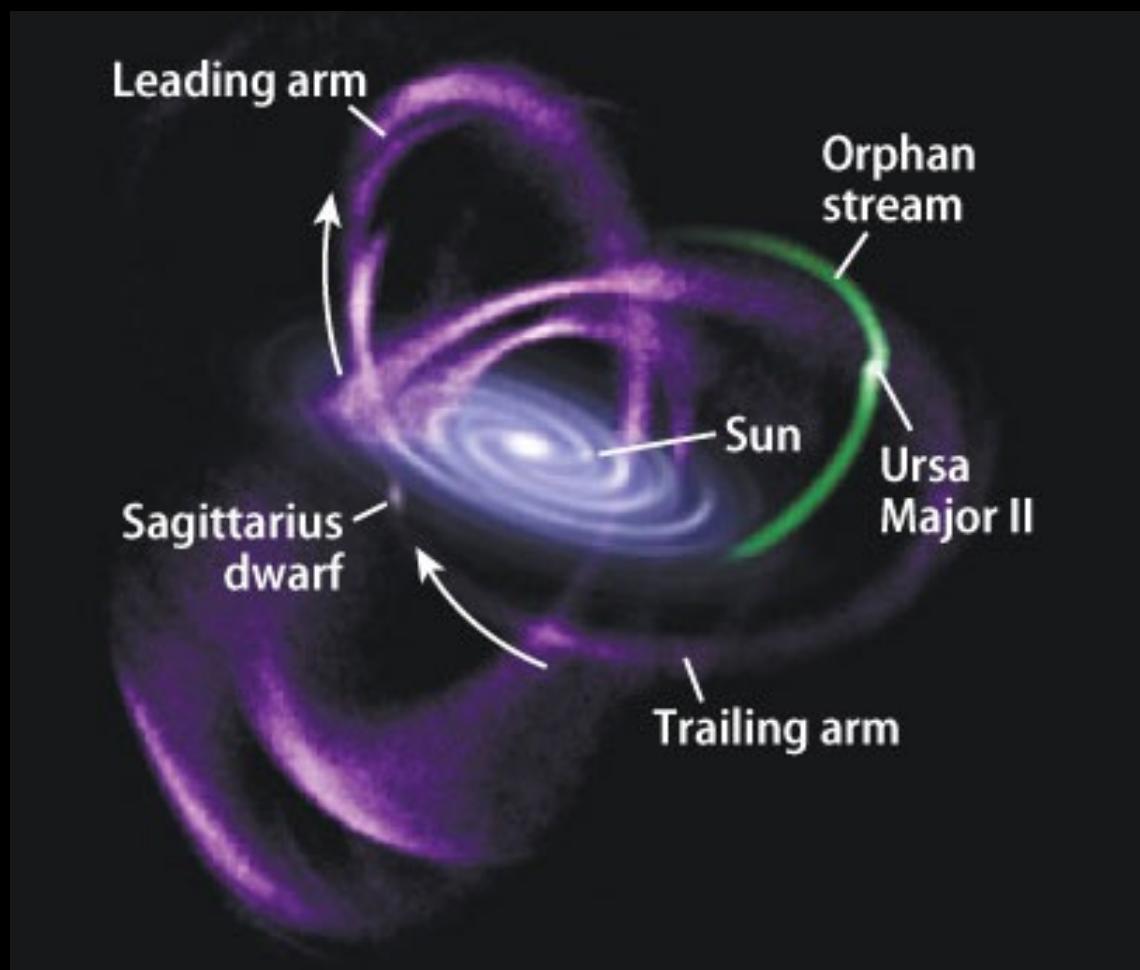
preservation

disruption

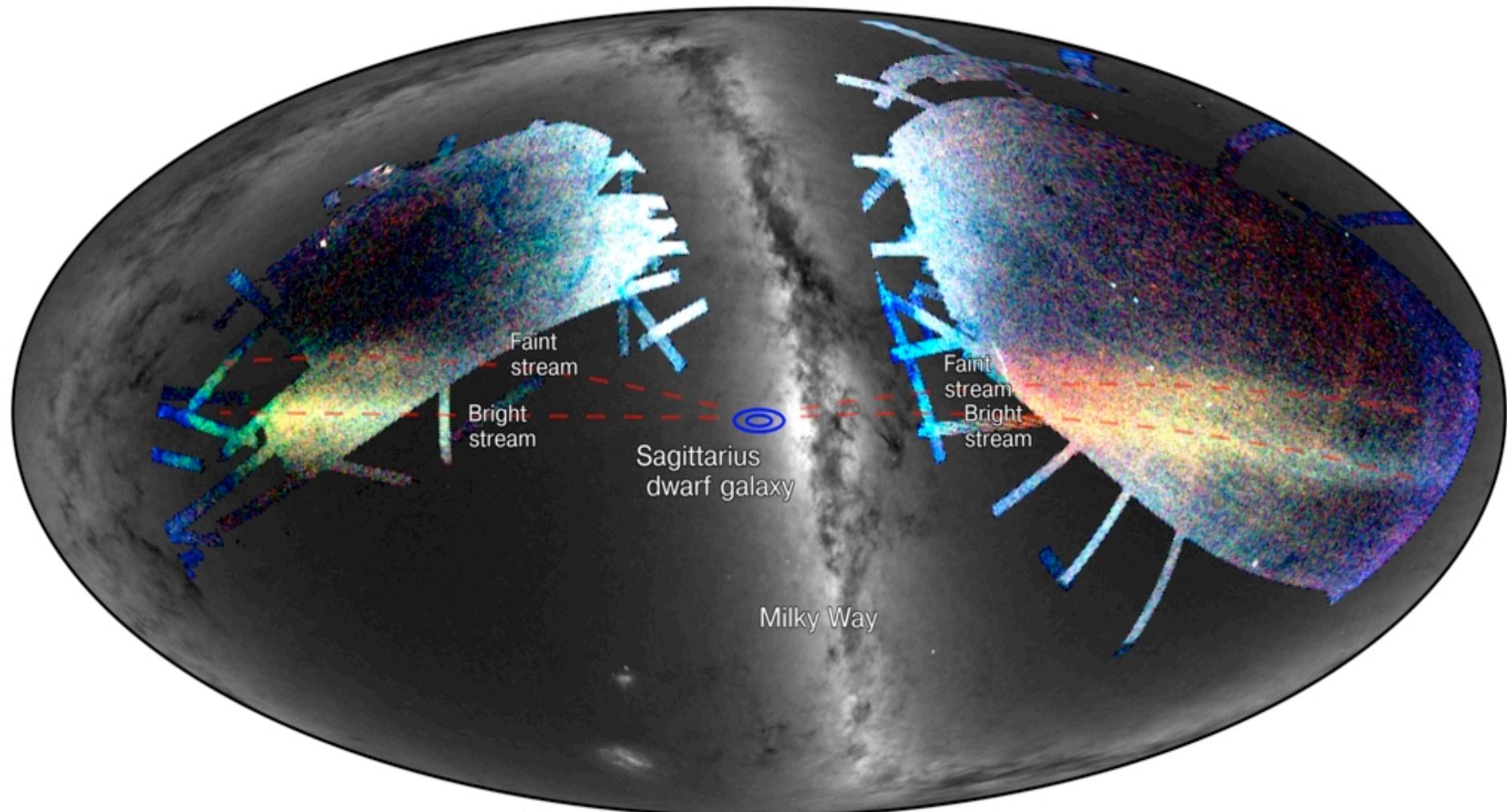
incorporation

<https://folklife.si.edu/magazine/intergalactic-pachamama-kichwa-cosmology-vs-western-astrophysics>

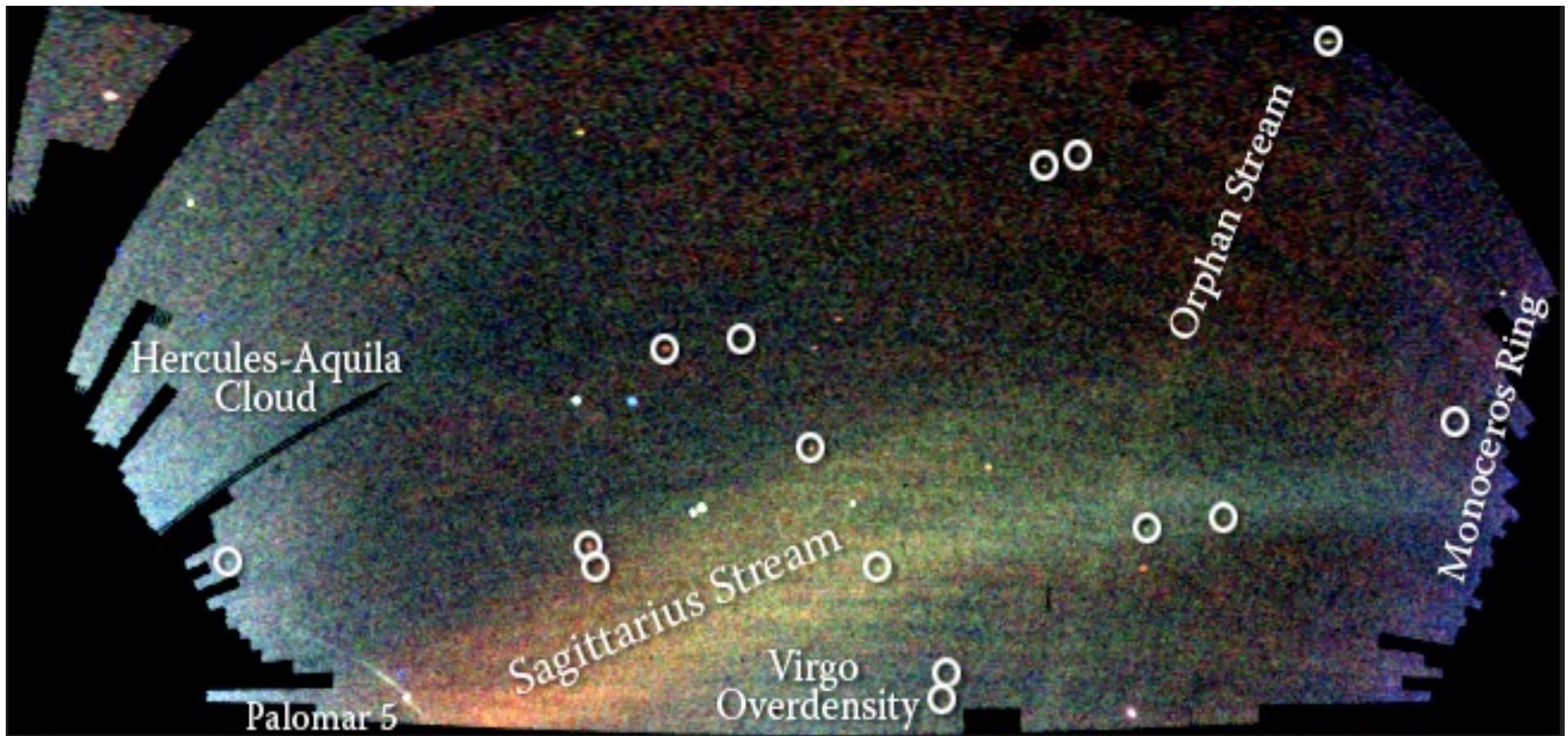
The MW Stellar Halo contains stellar streams,
remnants of the tidal sharing process



SDSS: Field Of Streams

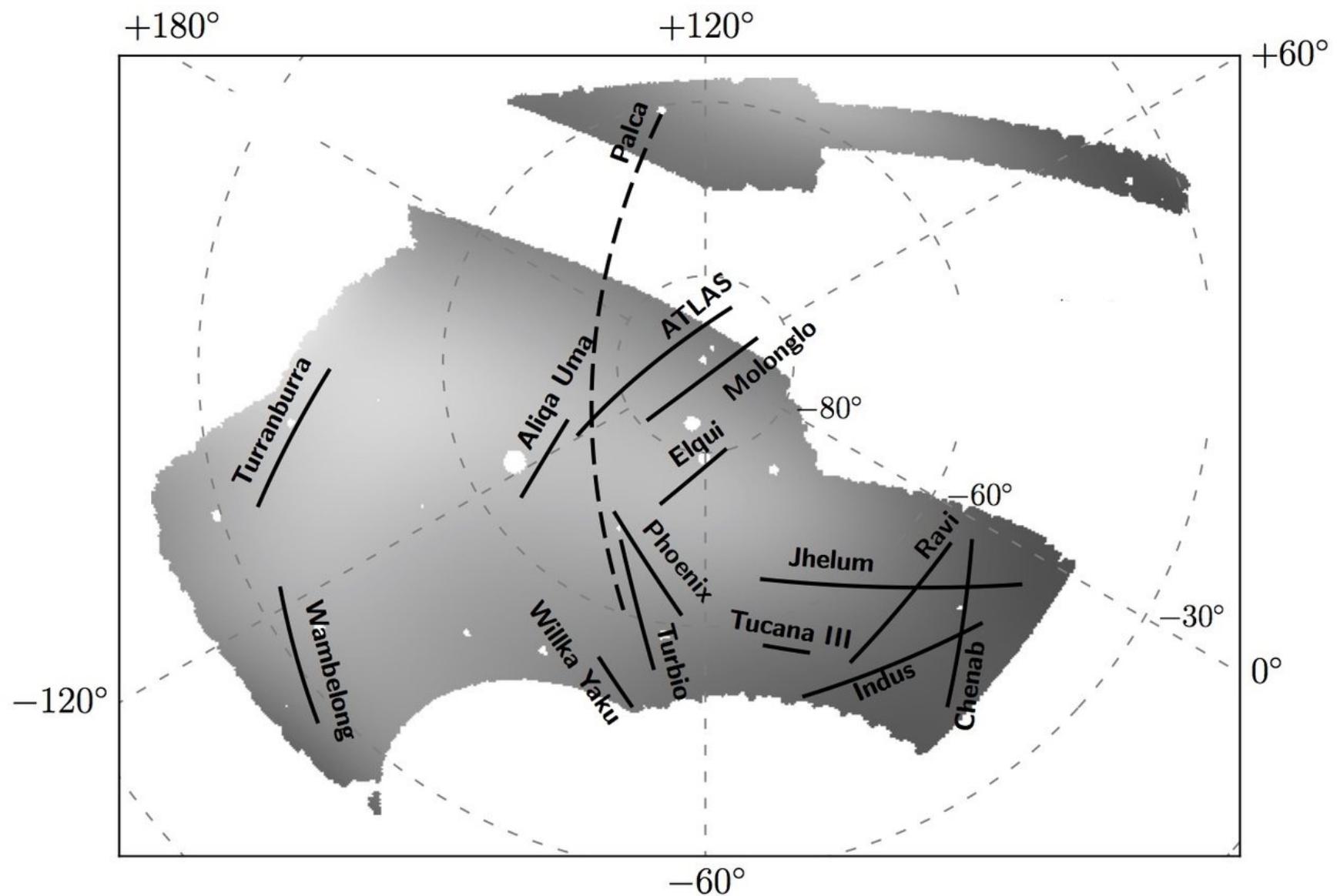


SDSS Field of Streams: Northern Hemisphere

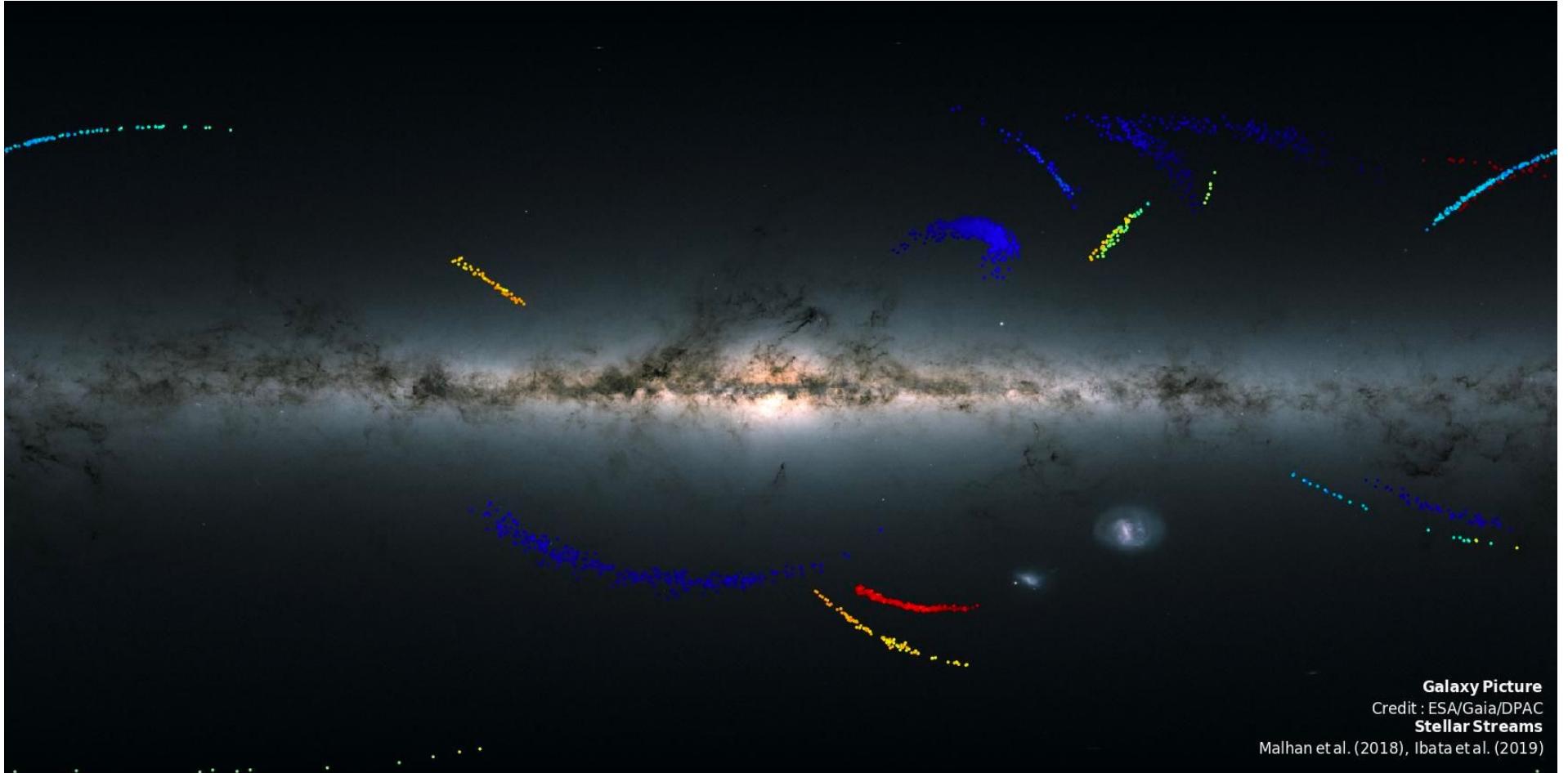


Color coding indicates radial velocity

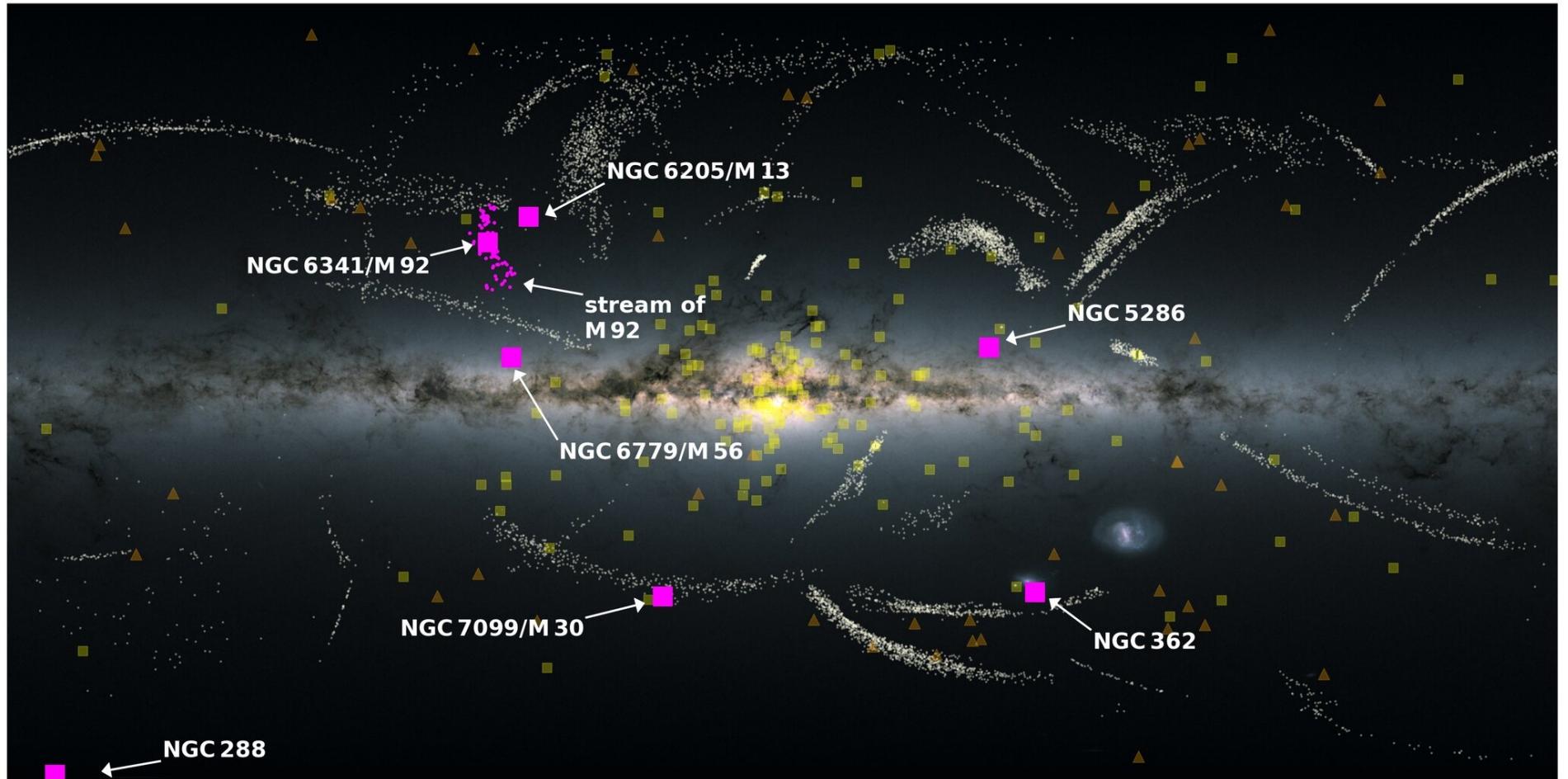
Dark Energy Survey



Gaia Streams: EDR2



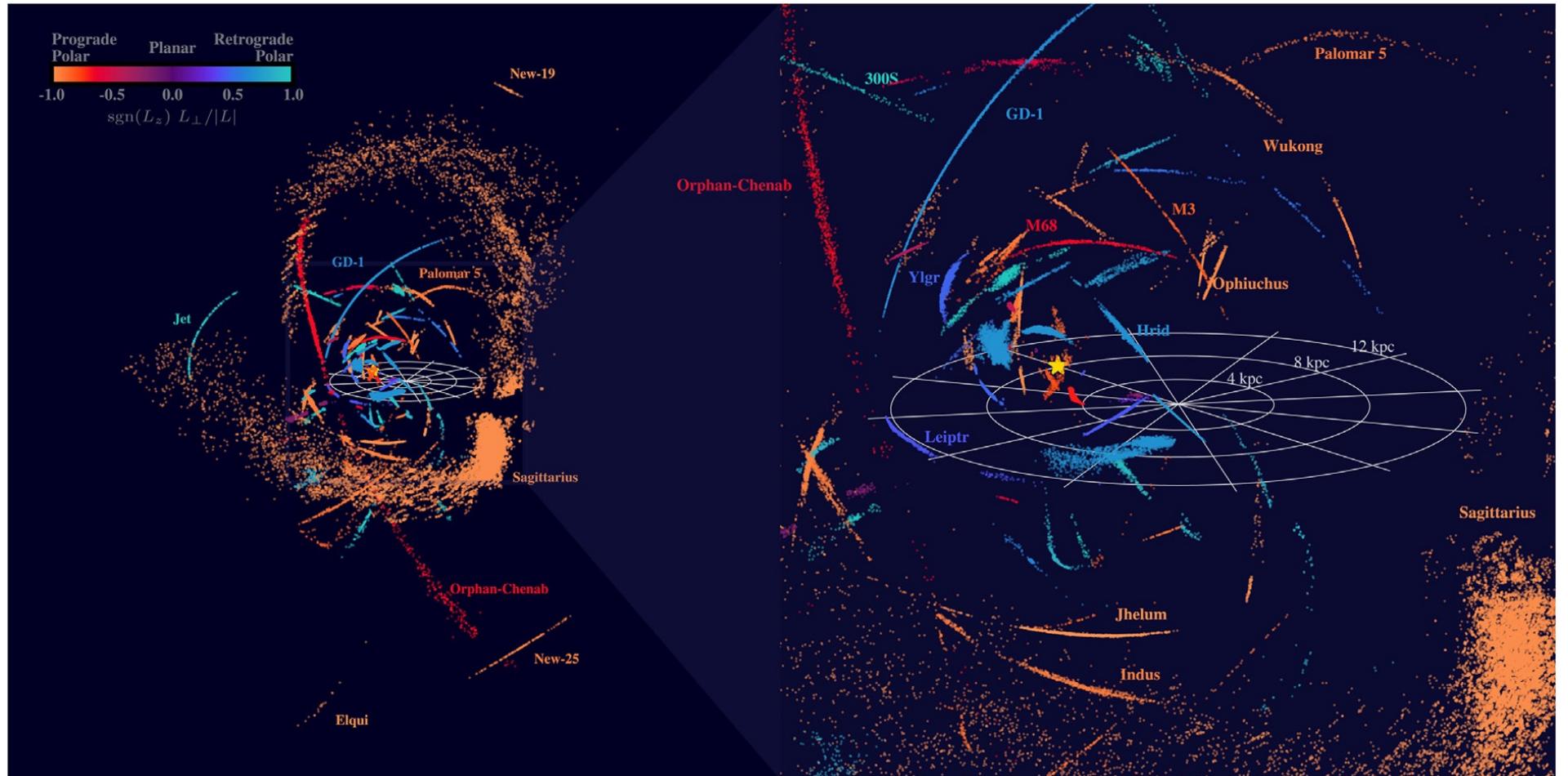
Gaia Stream : EDR3

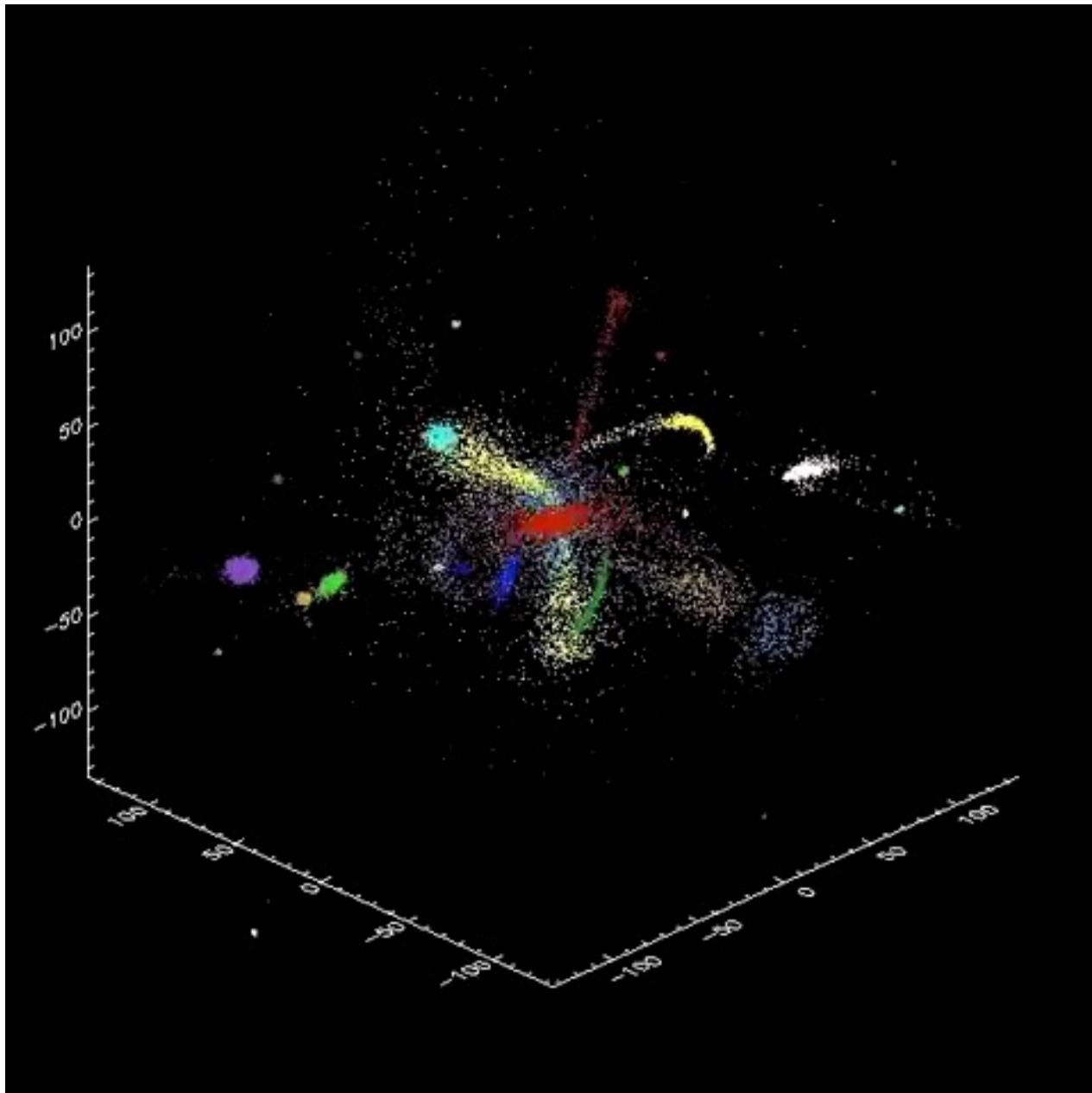


This image shows the Milky Way as seen by Gaia. The squares represent the location of globular clusters, the triangles the location of satellite galaxies, and the small dots are stellar streams. (Malhan+2022)



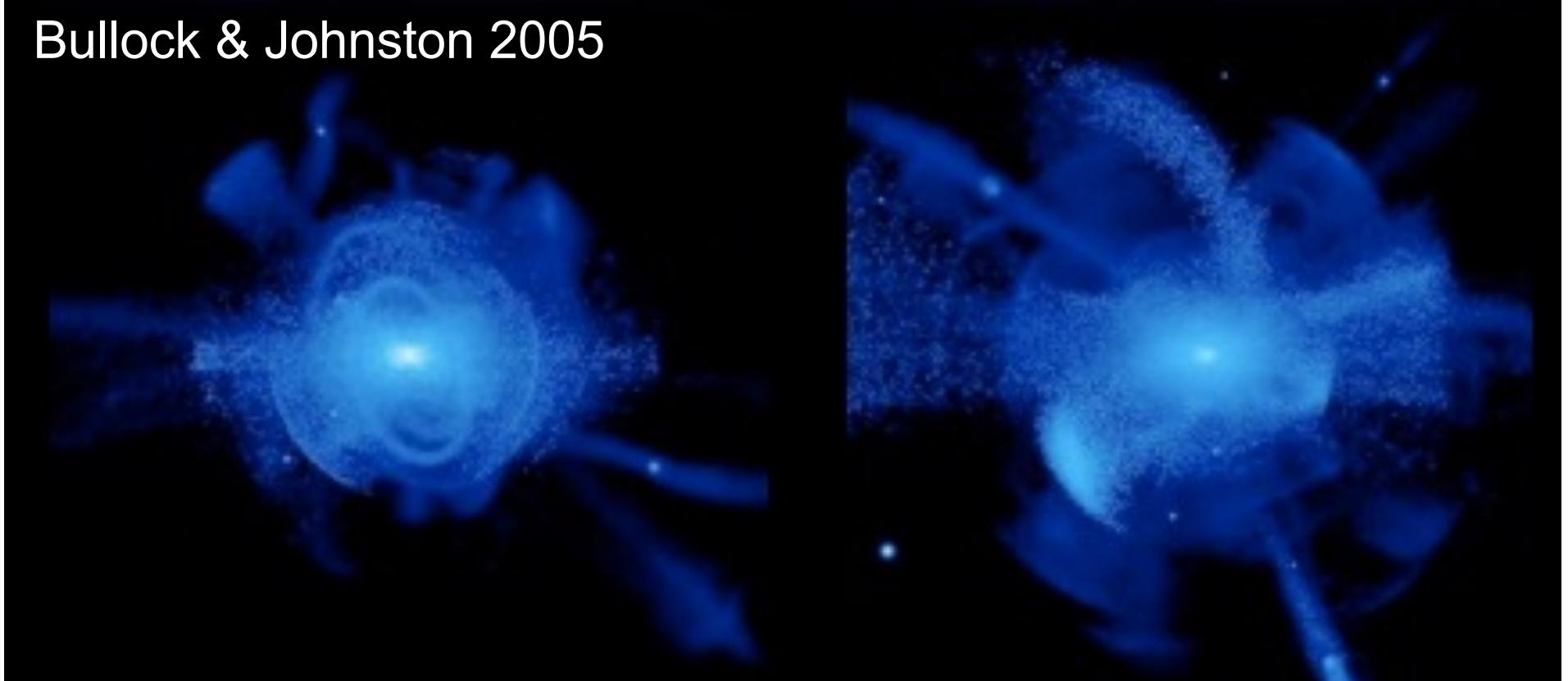
The Milky Way galaxy and the stellar streams (colored dots), globular clusters (star symbols) and dwarf galaxies (small cubes)
© S. Payne-Wardenar / Khyati Malhan, MPIA. Malhan+2022





Bullock & Johnston 2005

Bullock & Johnston 2005





Elliptical Galaxy
NGC 474

P-A Duc
CFHT

Tidal Radius: Point mass host, no rotation

$$r_t \sim (m/4M)^{1/3} D.$$

Accounting for centrifugal forces:

$$\mathbf{v}' \equiv \frac{d\mathbf{x}'}{dt'} = \mathbf{v} - \Omega \times \mathbf{x}.$$

↑
Angular speed

Jacobi Constant:

$$E_J = \frac{1}{2}\mathbf{v}'^2 + \Phi_{\text{eff}}(\mathbf{x}'), \quad \text{where } \Phi_{\text{eff}}(\mathbf{x}') \equiv \Phi(\mathbf{x}') - \frac{1}{2}(\vec{\Omega} \times \mathbf{x}')^2$$

For a satellite (m) orbiting a host (M):

$$\Phi_{\text{eff}}(x) = -\frac{GM}{|D-x|} - \frac{Gm}{|x|} - \frac{\Omega^2}{2} \left(x - \frac{DM}{M+m} \right)^2.$$

$$\frac{\partial \Phi_{\text{eff}}}{\partial x} = 0 = -\frac{GM}{(D-x)^2} \pm \frac{Gm}{x^2} - \Omega^2 \left(x - \frac{DM}{M+m} \right)$$

Solutions: Lagrange points L1 and L2 are

$$x = \pm r_J, \quad \text{where } r_J = D \left(\frac{m}{3M+m} \right)^{1/3}$$

For an extended host (Isothermal Sphere):

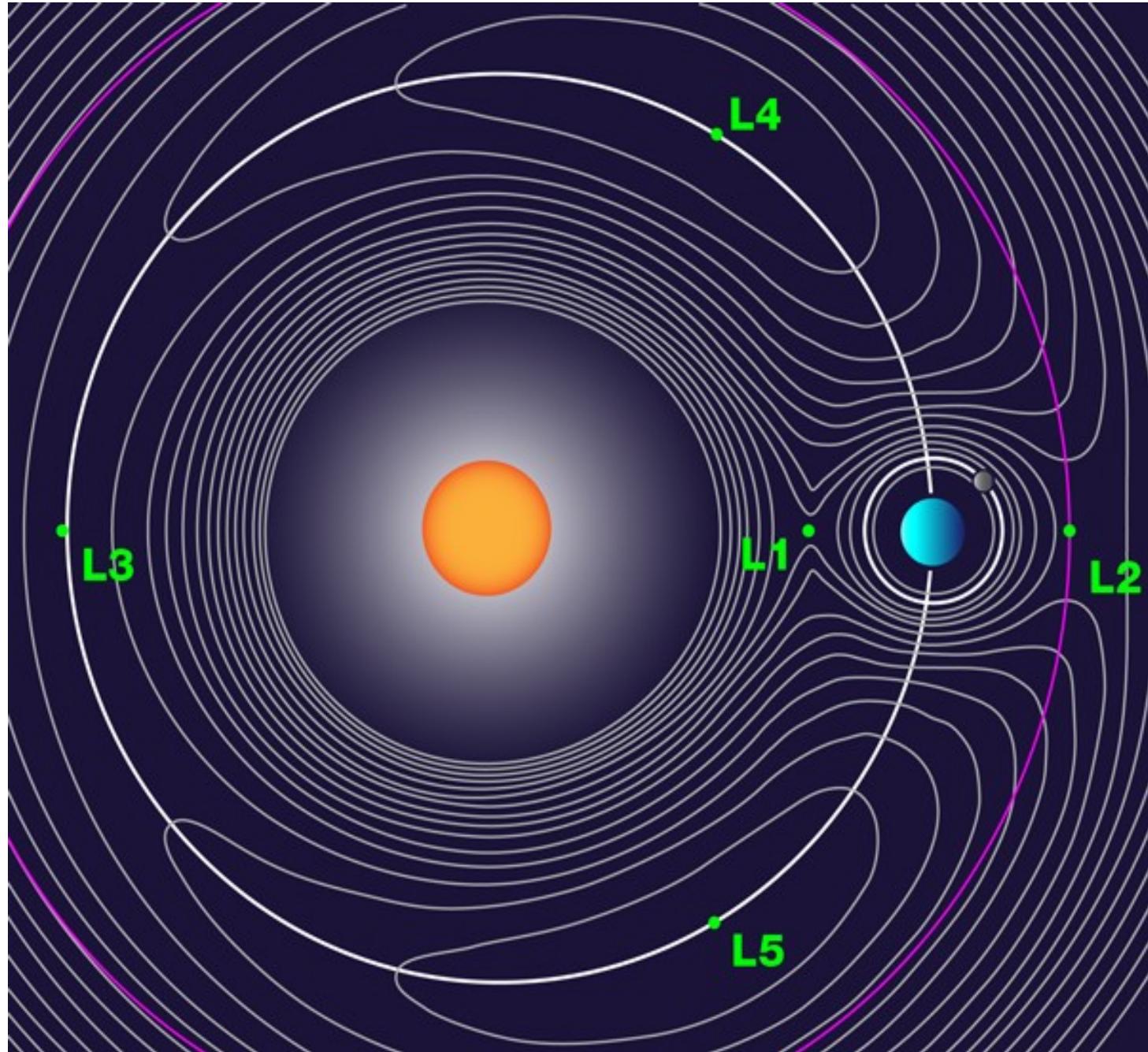
Szebehely 1967,
Valtonen & Karttunen 2006

$$r_J = D \left[\frac{m}{2M(< D)} \right]^{1/3}$$

For an extended host (in general):

$$r_J = \left(\frac{GM_{\text{sat}}}{\Omega^2 - d^2 \Phi / d^2 r} \right)^{1/3}$$

Chapter 4 of Sparke & Gallagher



Choi + 2009

