

Solar-sail perturbed cr3bp

Previous videos:

- Recap of the cr3bp
- Equilibria in the cr3bp
- Solar-sail acceleration model

cr3bp – recap from video “recap of the cr3bp”

cr3bp – the equations of motion **in dimensionless form**

- Equations of motion in rotating frame – **dimensionless**

$$\ddot{\mathbf{r}} = -\left(\frac{1-\mu}{r_1^3}\mathbf{r}_1 + \frac{\mu}{r_2^3}\mathbf{r}_2\right) - 2\boldsymbol{\omega} \times \dot{\mathbf{r}} - \boldsymbol{\omega} \times (\boldsymbol{\omega} \times \mathbf{r})$$

- $\mathbf{r} = [x \ y \ z]^T$, $\dot{\mathbf{r}} = [\dot{x} \ \dot{y} \ \dot{z}]^T$, $\ddot{\mathbf{r}} = [\ddot{x} \ \ddot{y} \ \ddot{z}]^T$

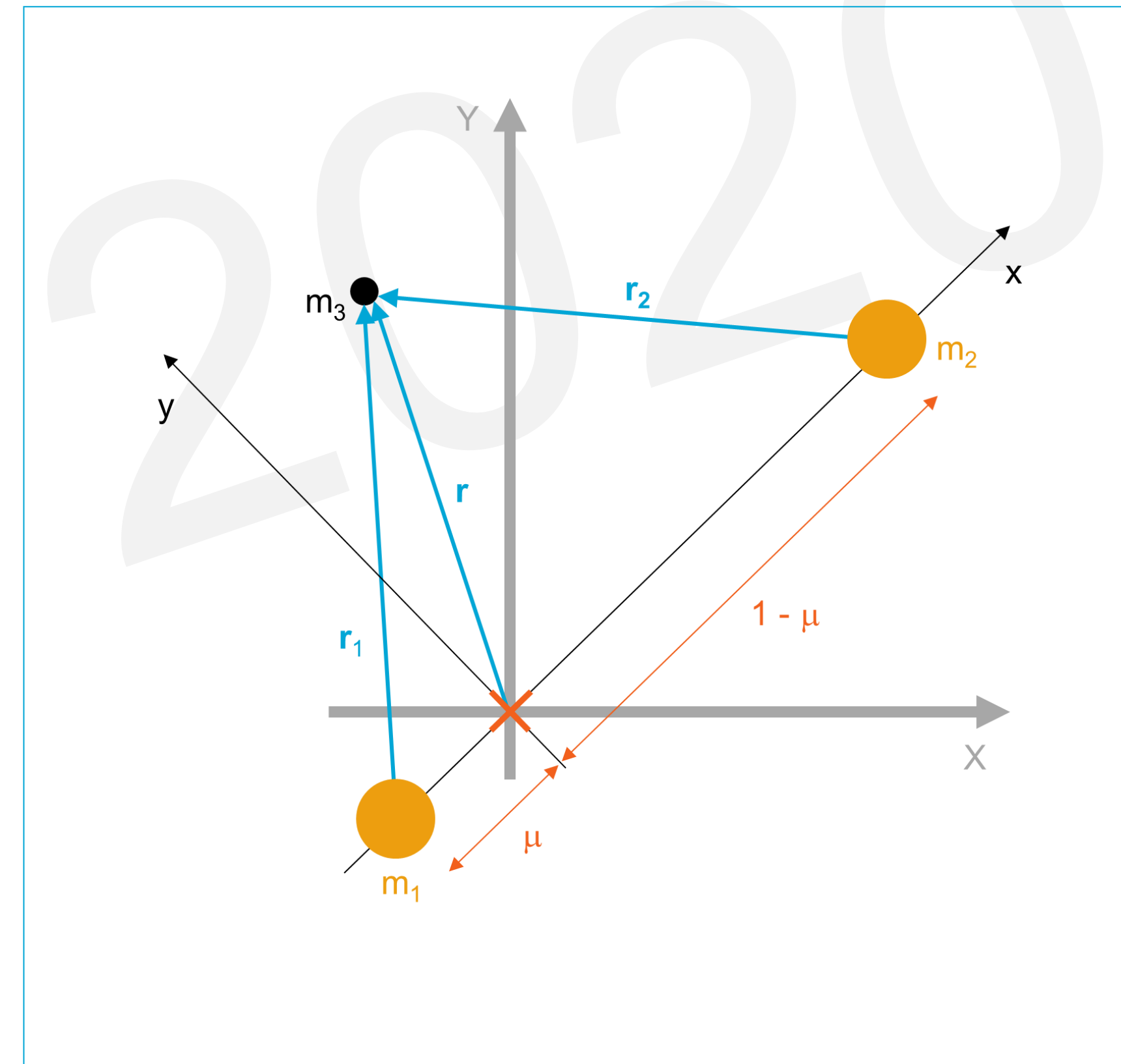
- $\mu = \frac{m_1}{m_1 + m_2}$ Erratum: this should be $\mu = \frac{m_2}{m_1 + m_2}$

- $\mathbf{r}_1 = [x + \mu \ y \ z]^T$

- $\mathbf{r}_2 = [x - (1 - \mu) \ y \ z]^T$

- $\boldsymbol{\omega} = [0 \ 0 \ \omega]^T = [0 \ 0 \ 1]^T$

- Equations of motion are fully parameterized by μ



cr3bp – recap from video “recap of the cr3bp”

cr3bp – the equations of motion in dimensionless form + potential formulation

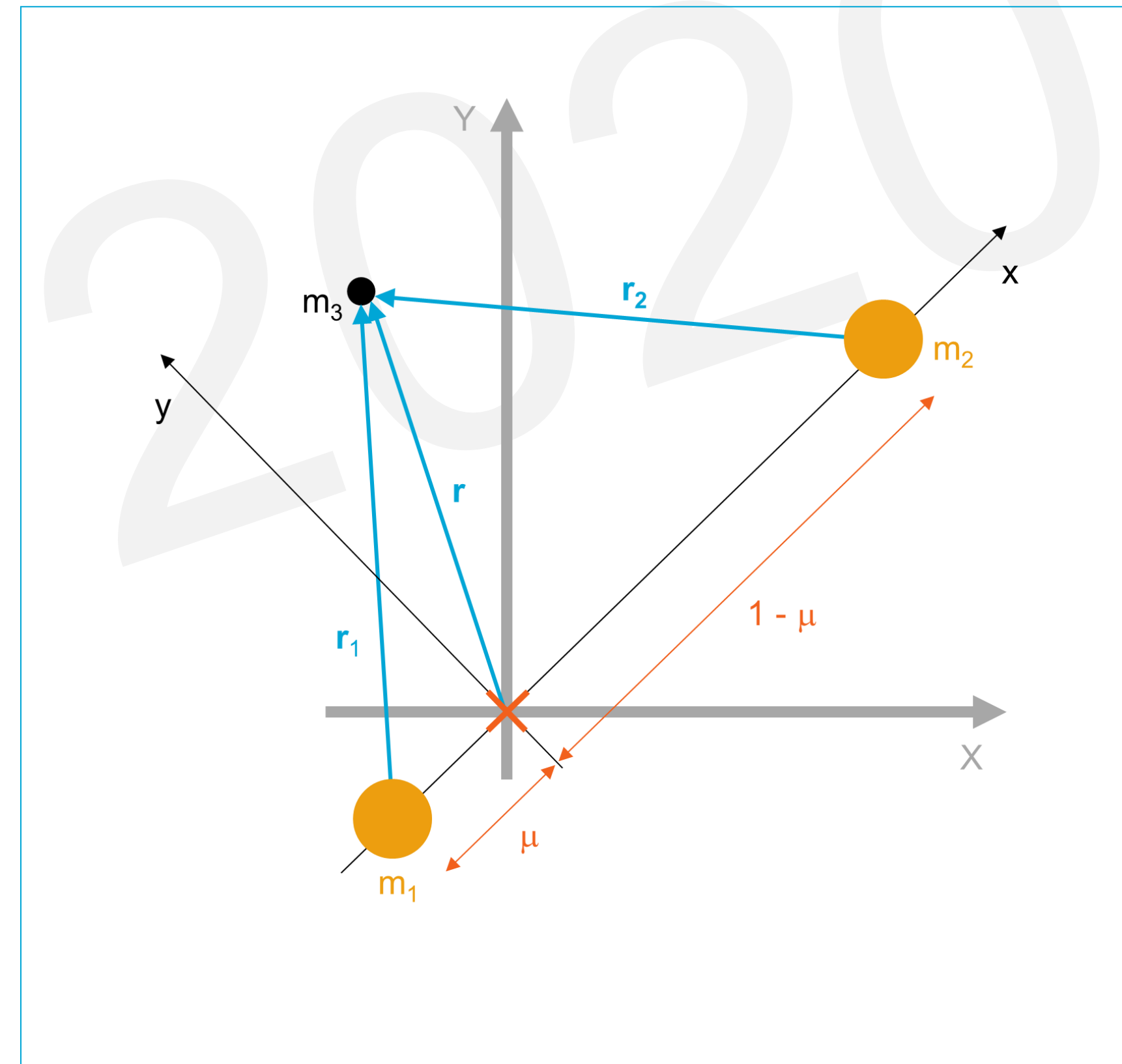
- Equations of motion in rotating frame – **dimensionless**

$$\ddot{\mathbf{r}} = - \underbrace{\left(\frac{1-\mu}{r_1^3} \mathbf{r}_1 + \frac{\mu}{r_2^3} \mathbf{r}_2 \right)}_{\text{Gravitational terms}} - \underbrace{2\boldsymbol{\omega} \times \dot{\mathbf{r}} - \boldsymbol{\omega} \times (\boldsymbol{\omega} \times \mathbf{r})}_{\text{Coriolis and centrifugal terms}}$$

These terms can be written in potential form

$$U = - \left(\frac{1-\mu}{r_1} + \frac{\mu}{r_2} \right) - \frac{1}{2} (x^2 + y^2) \quad \text{Effective potential}$$

$$\ddot{\mathbf{r}} = -2\boldsymbol{\omega} \times \dot{\mathbf{r}} - \nabla U$$



cr3bp – recap from video “equilibria in the cr3bp”

Equilibria – computation

- Also called “Lagrange points”
- Locations where, if
 - The body with negligible mass, m_3 , has zero velocity
 - No net acceleration acts on m_3

It will stay stationary *w.r.t.* the rotating frame $R(x,y,z)$

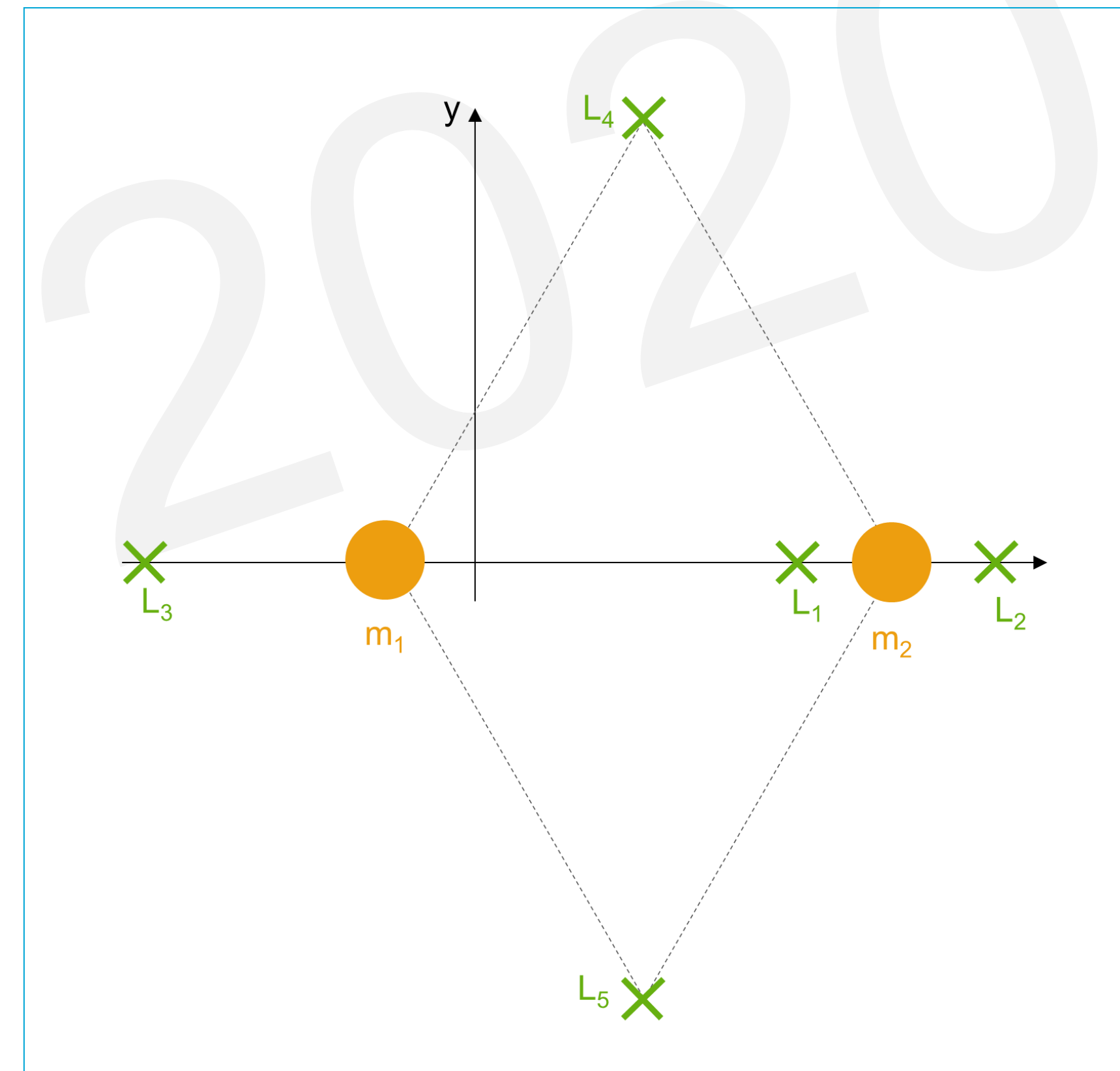
Trace out an orbit *w.r.t.* the inertial frame $I(X,Y,Z)$

$$\begin{matrix} \ddot{\mathbf{x}} \\ =0 \end{matrix} = -2\boldsymbol{\omega} \times \begin{matrix} \dot{\mathbf{x}} \\ =0 \end{matrix} - \nabla U$$

- Lagrange points are located where

$$\nabla U = \mathbf{0}$$

$$U = -\left(\frac{1-\mu}{r_1} + \frac{\mu}{r_2}\right) - \frac{1}{2}(x^2 + y^2)$$



Solar-sail perturbed cr3bp

- Equations of motion

$$\ddot{\mathbf{r}} = -2\boldsymbol{\omega} \times \dot{\mathbf{r}} - \nabla U$$

$$\circ \mathbf{r} = [x \ y \ z]^T, \dot{\mathbf{r}} = [\dot{x} \ \dot{y} \ \dot{z}]^T, \ddot{\mathbf{r}} = [\ddot{x} \ \ddot{y} \ \ddot{z}]^T$$

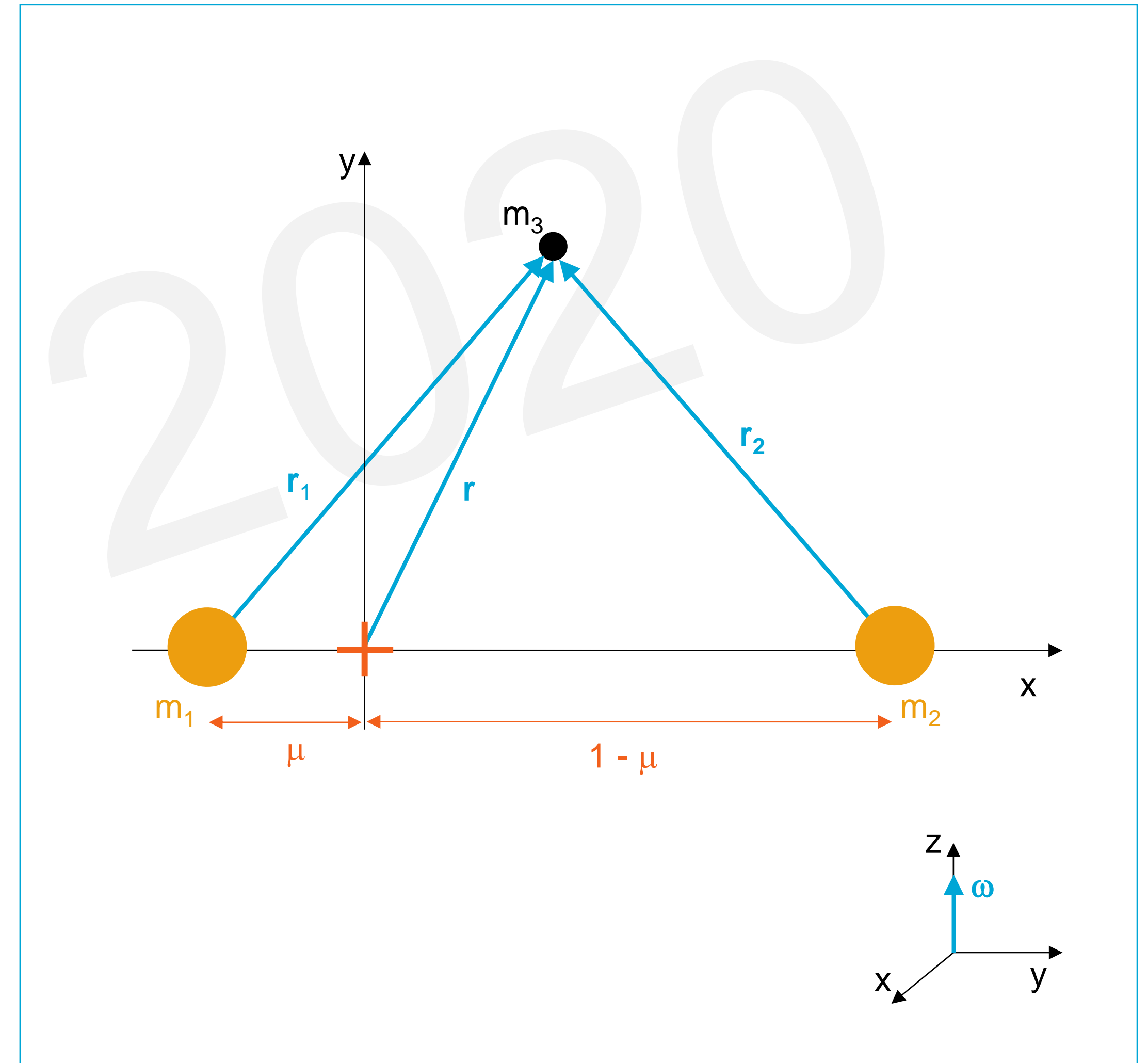
$$\circ U = -\left(\frac{1-\mu}{r_1} + \frac{\mu}{r_2}\right) - \frac{1}{2}(x^2 + y^2)$$

$$\circ \mu = \frac{m_2}{m_1 + m_2}$$

$$\circ \mathbf{r}_1 = [x + \mu \ y \ z]^T$$

$$\circ \mathbf{r}_2 = [x - (1 - \mu) \ y \ z]^T$$

$$\circ \boldsymbol{\omega} = [0 \ 0 \ \omega]^T = [0 \ 0 \ 1]^T$$



Solar-sail perturbed cr3bp

- Equations of motion

$$\ddot{\mathbf{r}} = -2\boldsymbol{\omega} \times \dot{\mathbf{r}} - \nabla U + \mathbf{a}_s$$

$$\circ \mathbf{r} = [x \ y \ z]^T, \dot{\mathbf{r}} = [\dot{x} \ \dot{y} \ \dot{z}]^T, \ddot{\mathbf{r}} = [\ddot{x} \ \ddot{y} \ \ddot{z}]^T$$

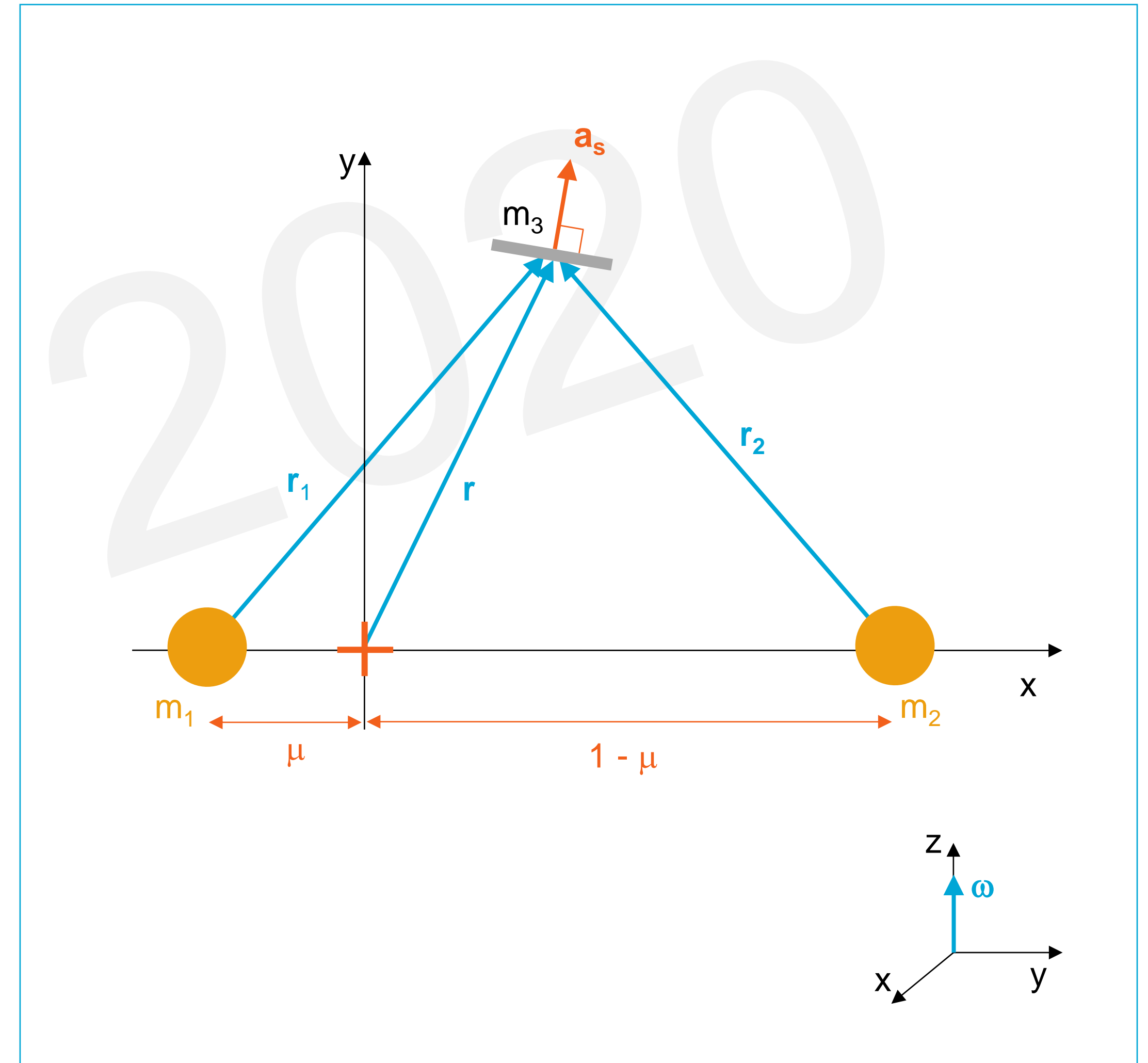
$$\circ U = -\left(\frac{1-\mu}{r_1} + \frac{\mu}{r_2}\right) - \frac{1}{2}(x^2 + y^2)$$

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$$\circ \mathbf{r}_1 = [x + \mu \ y \ z]^T$$

$$\circ \mathbf{r}_2 = [x - (1-\mu) \ y \ z]^T$$

$$\circ \boldsymbol{\omega} = [0 \ 0 \ \omega]^T = [0 \ 0 \ 1]^T$$

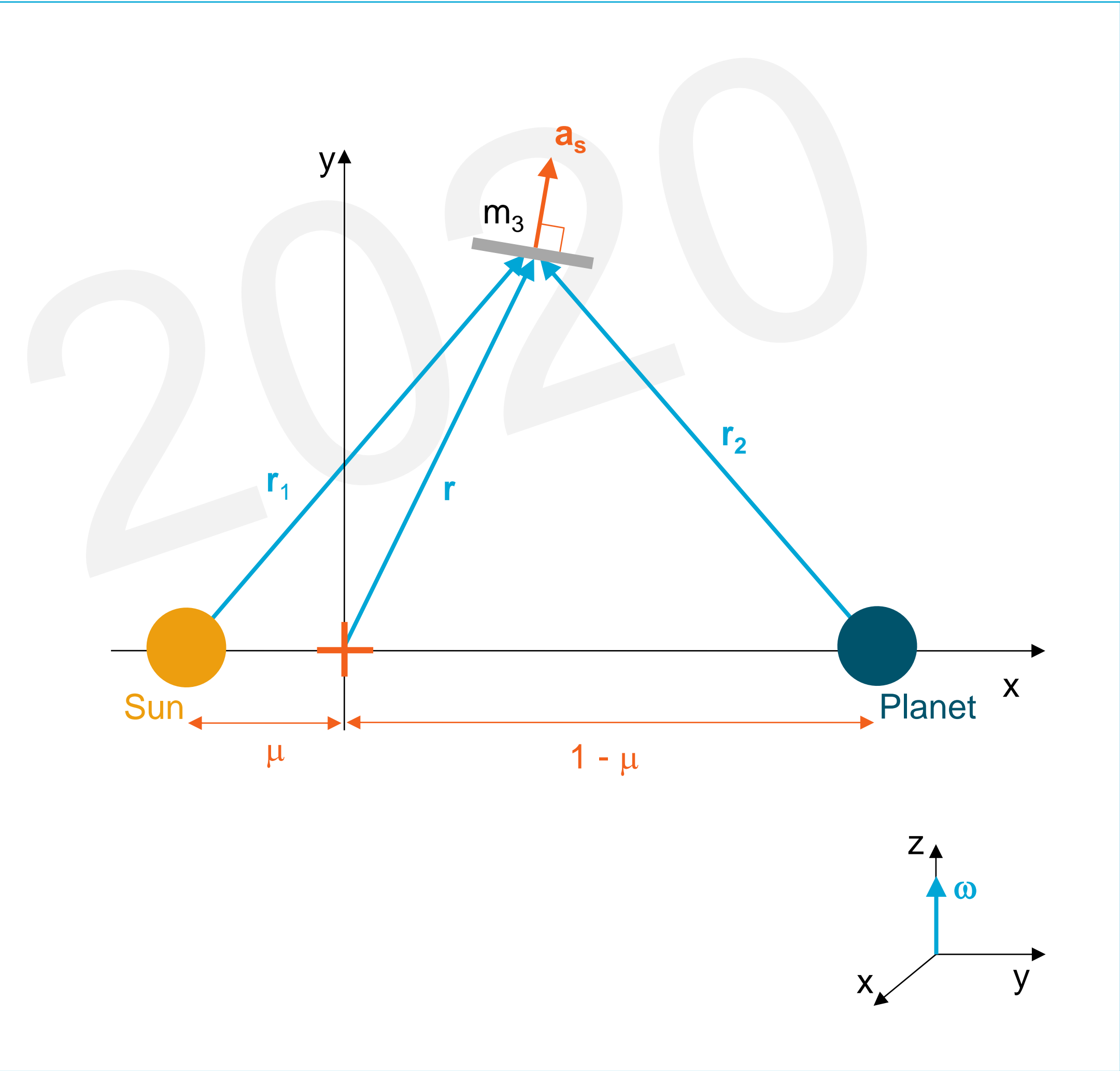
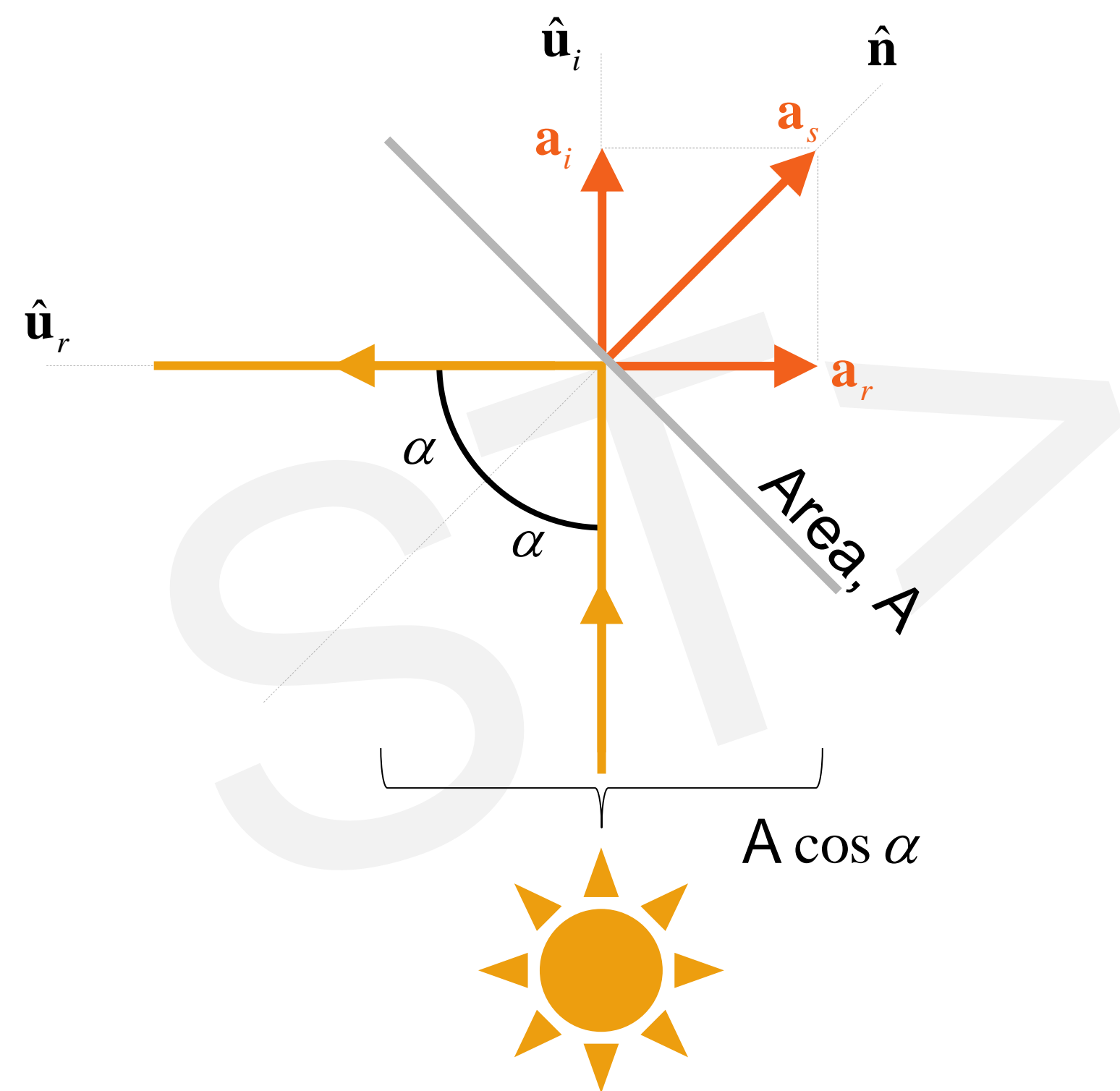


Solar-sail acceleration – recap from video “solar-sail acceleration model”

- Dimensional solar-sail acceleration

$$\mathbf{a}_s = \boxed{\beta} \frac{\mu_\odot}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

Solar-sail lightness number

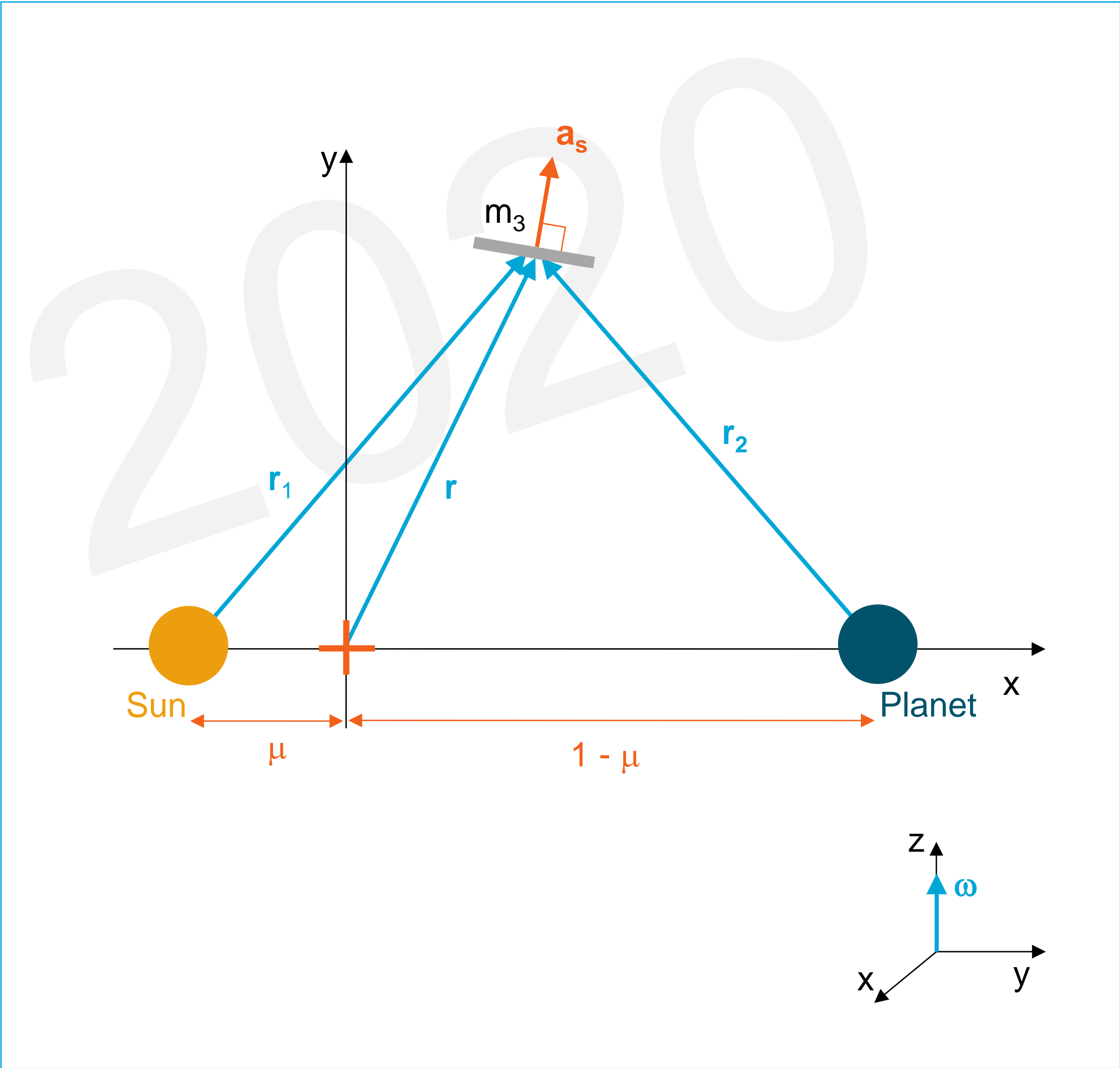
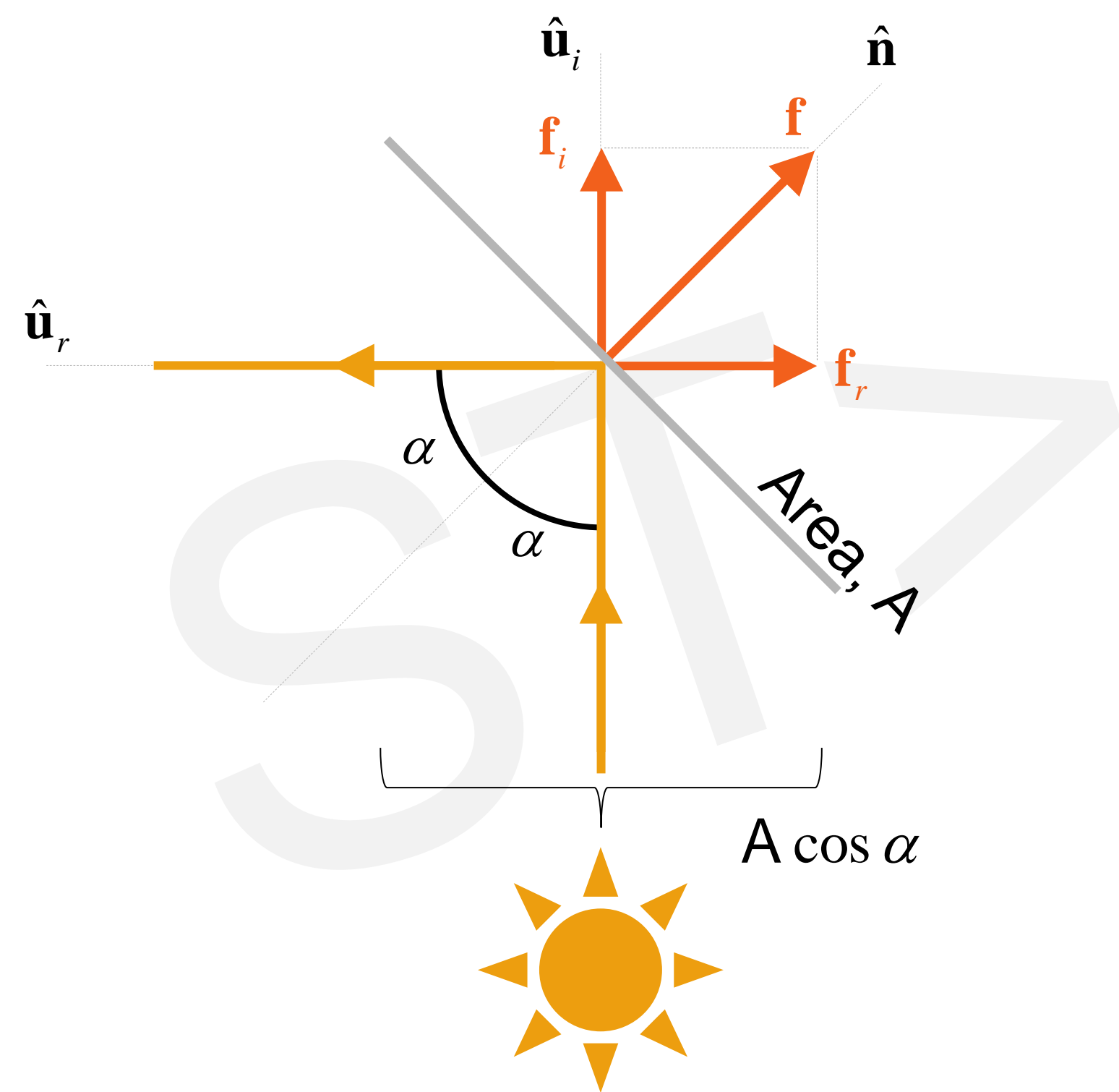


Solar-sail acceleration – recap from video “solar-sail acceleration model”

- Dimensional solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{\mu_{\odot}}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

Solar gravitational acceleration

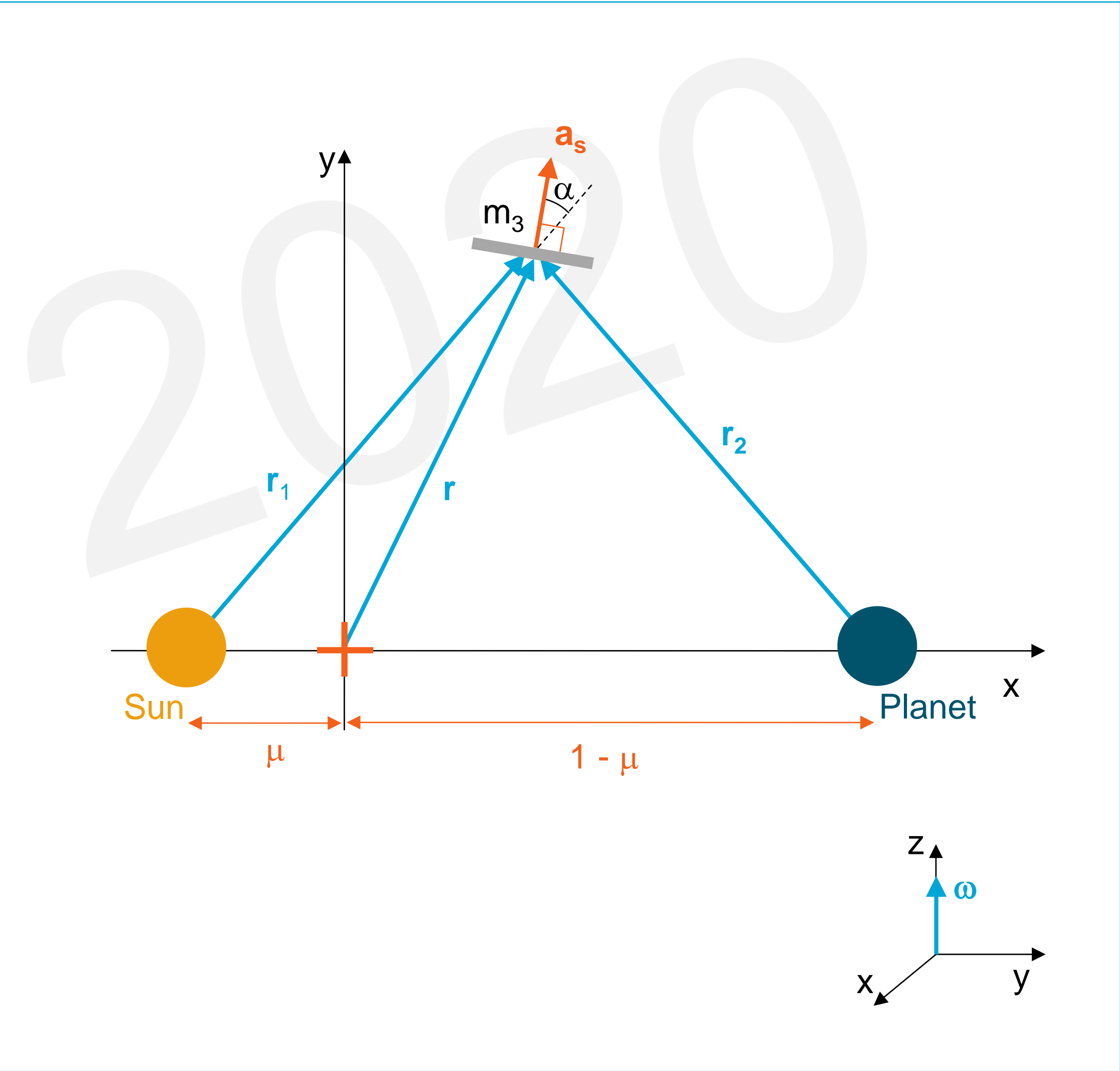
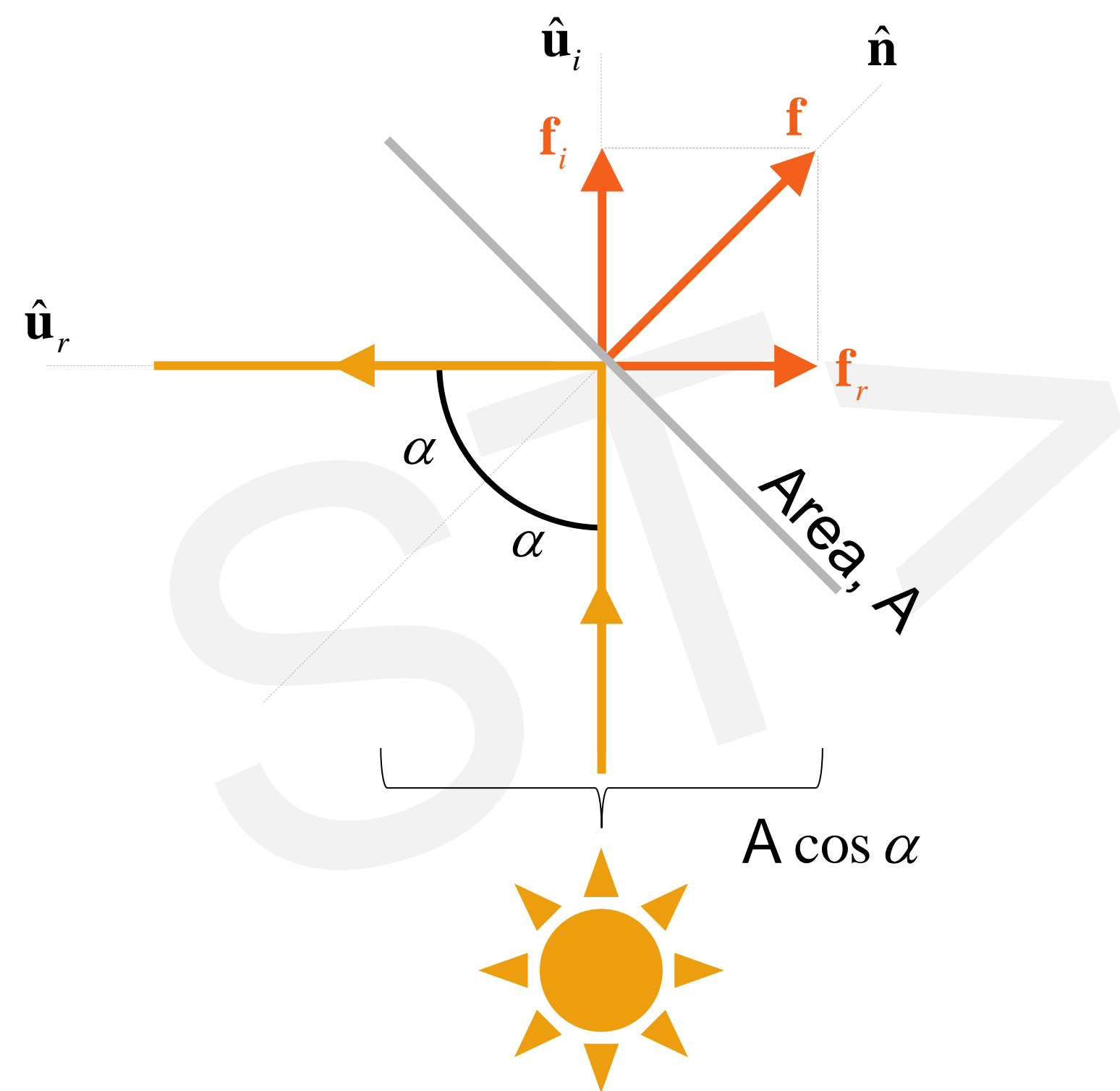


Solar-sail acceleration – recap from video “solar-sail acceleration model”

- Dimensional solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{\mu_{\odot}}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

Accounting for sail attitude

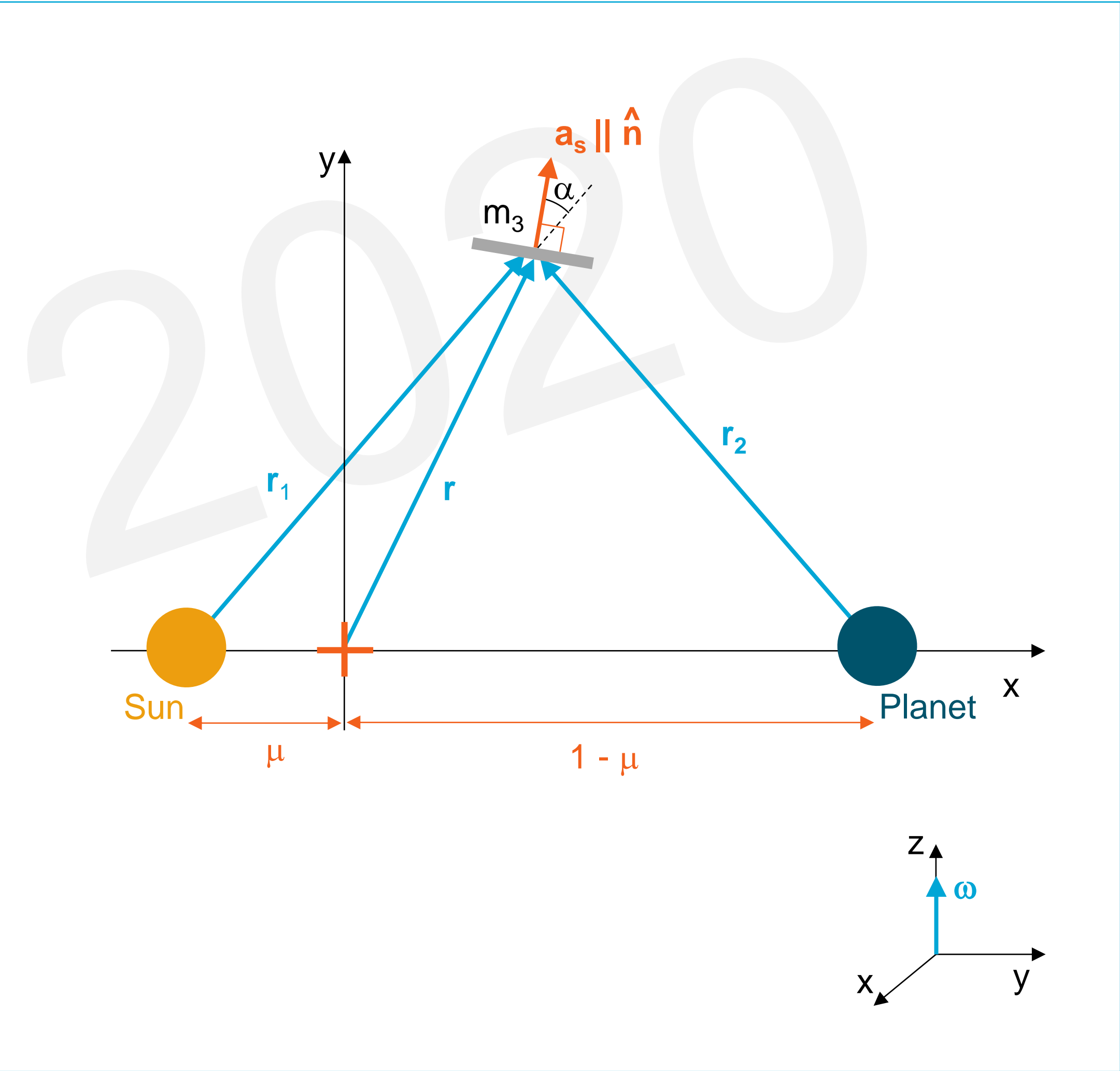
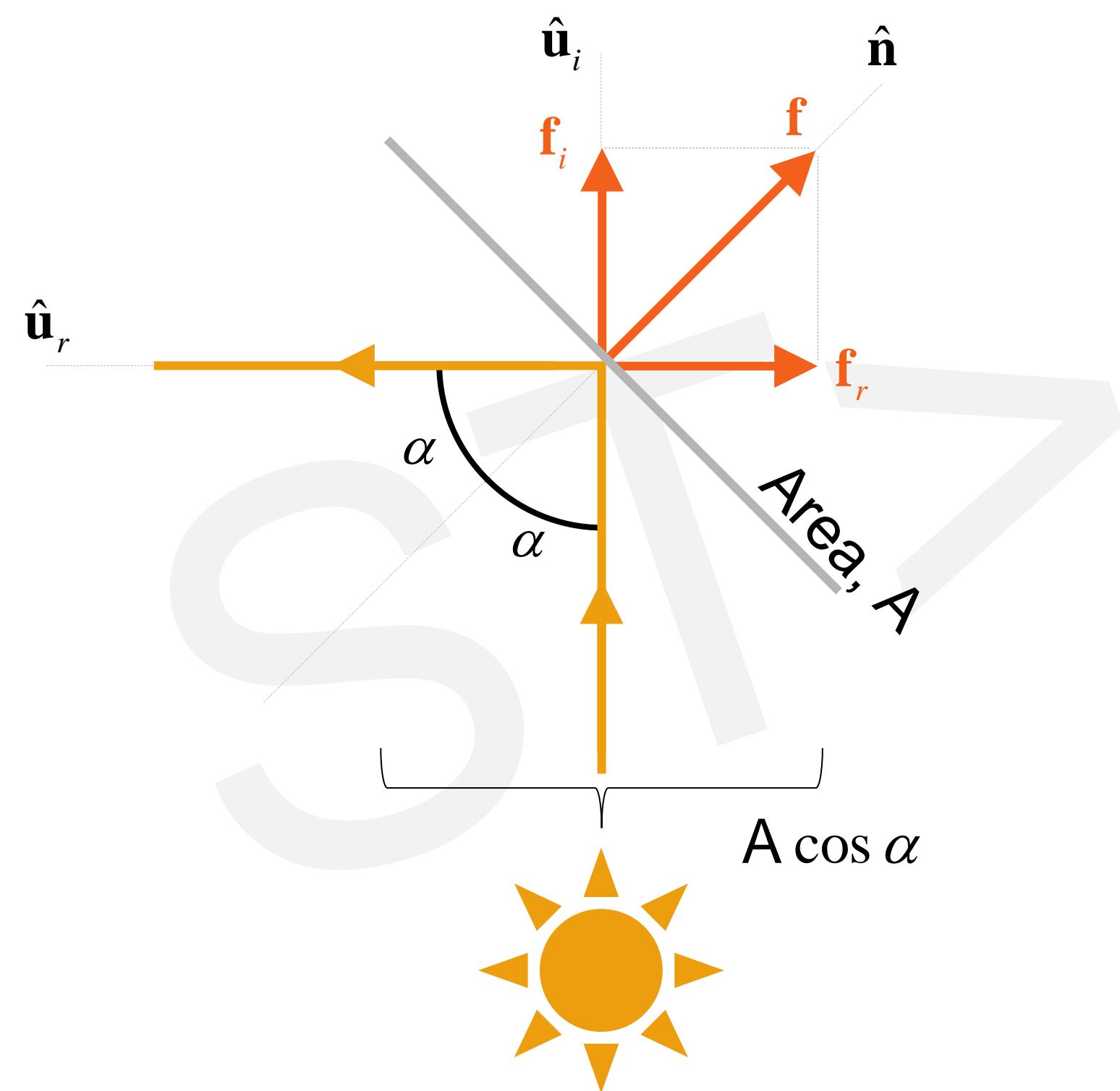


Solar-sail acceleration – recap from video “solar-sail acceleration model”

- Dimensional solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{\mu_\odot}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

Unit vector normal to the sail



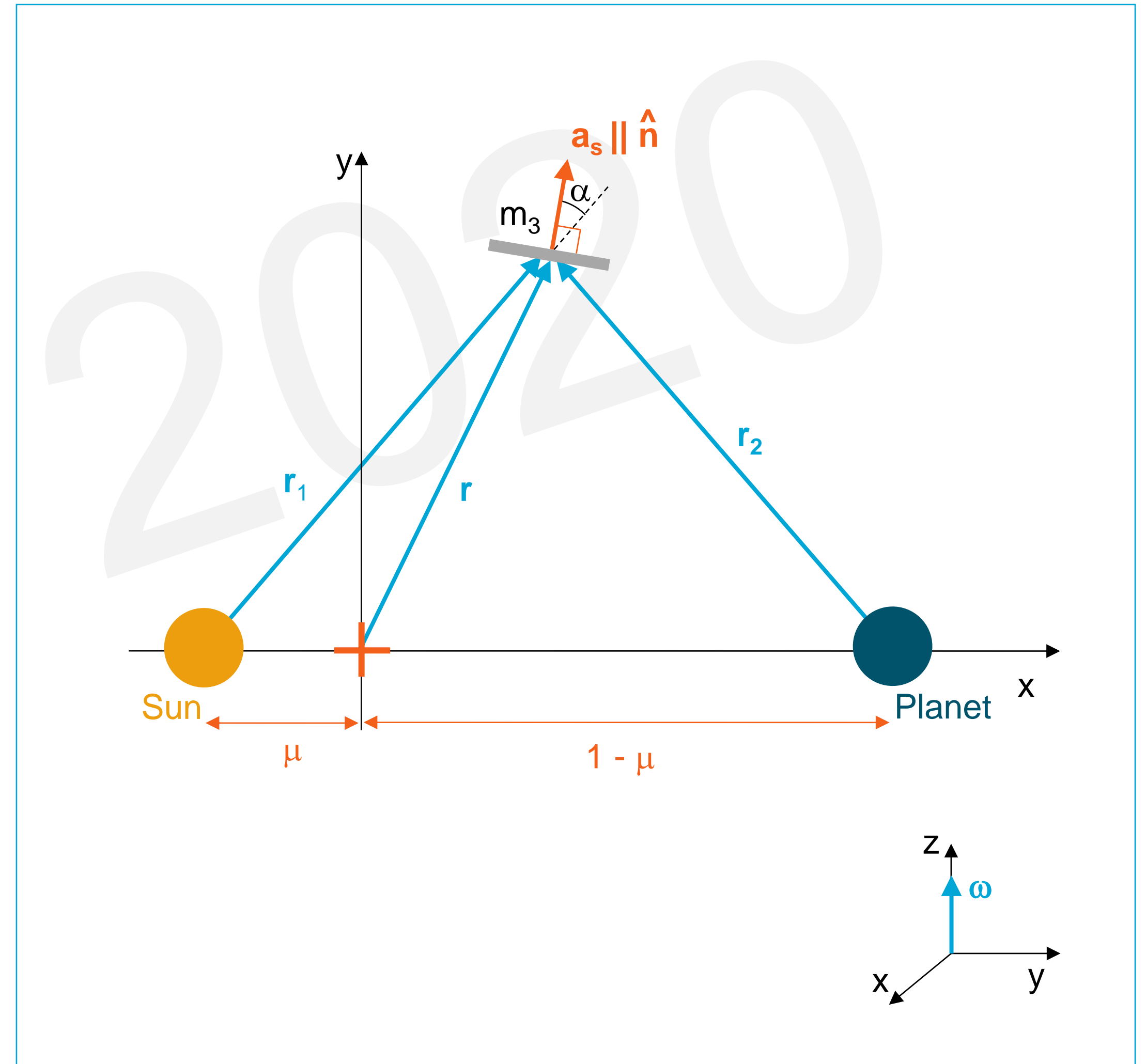
Solar-sail acceleration in the dimensionless cr3bp

- Dimensional solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{\mu_{\odot}}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

- **Dimensionless** solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$



Solar-sail acceleration in the dimensionless cr3bp

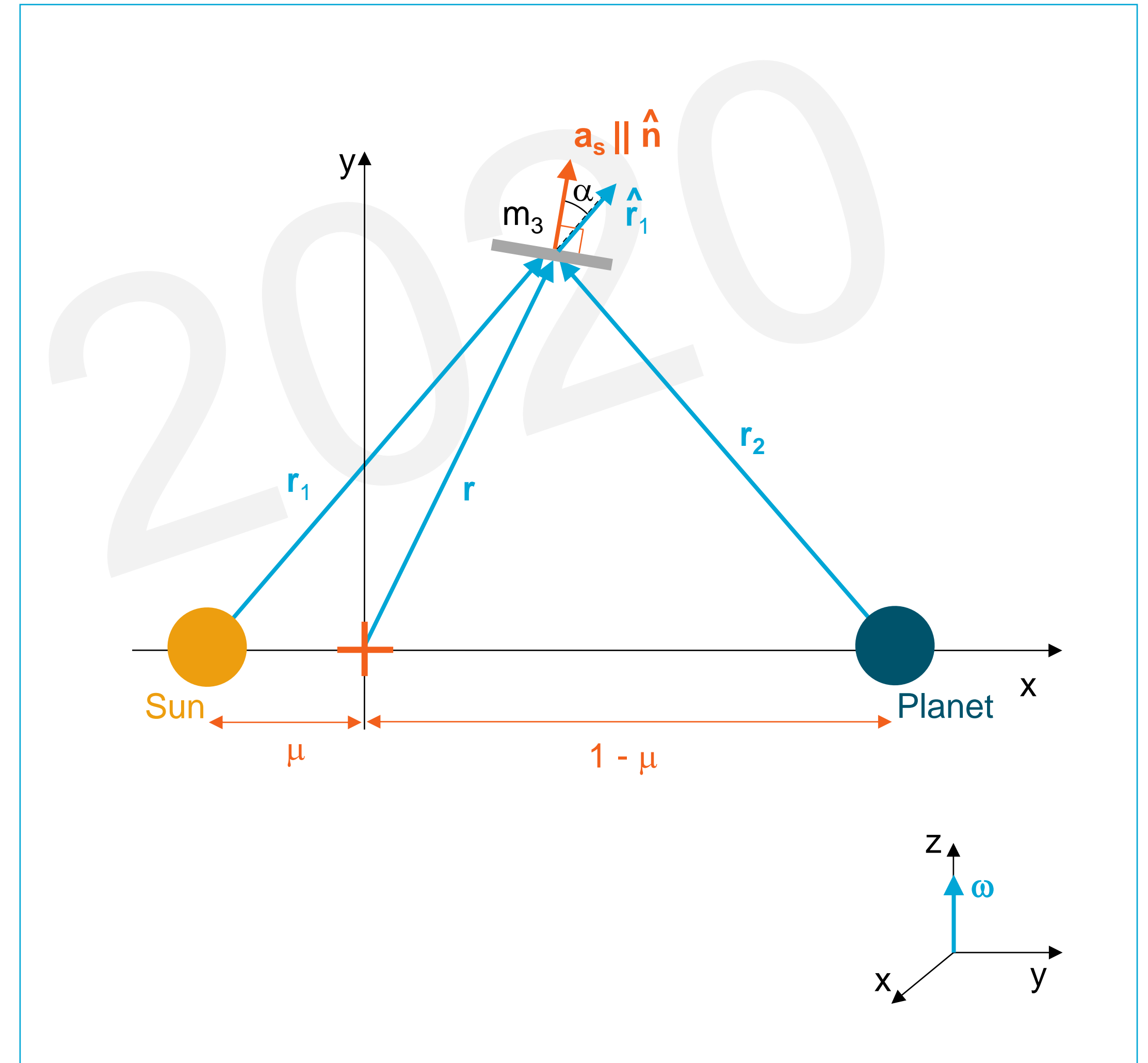
- Dimensional solar-sail acceleration in

$$\mathbf{a}_s = \beta \frac{\mu_{\odot}}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

- Dimensionless solar-sail acceleration in

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

- More **general form** of solar-sail acceleration



Solar-sail acceleration in the dimensionless cr3bp

- Dimensional solar-sail acceleration

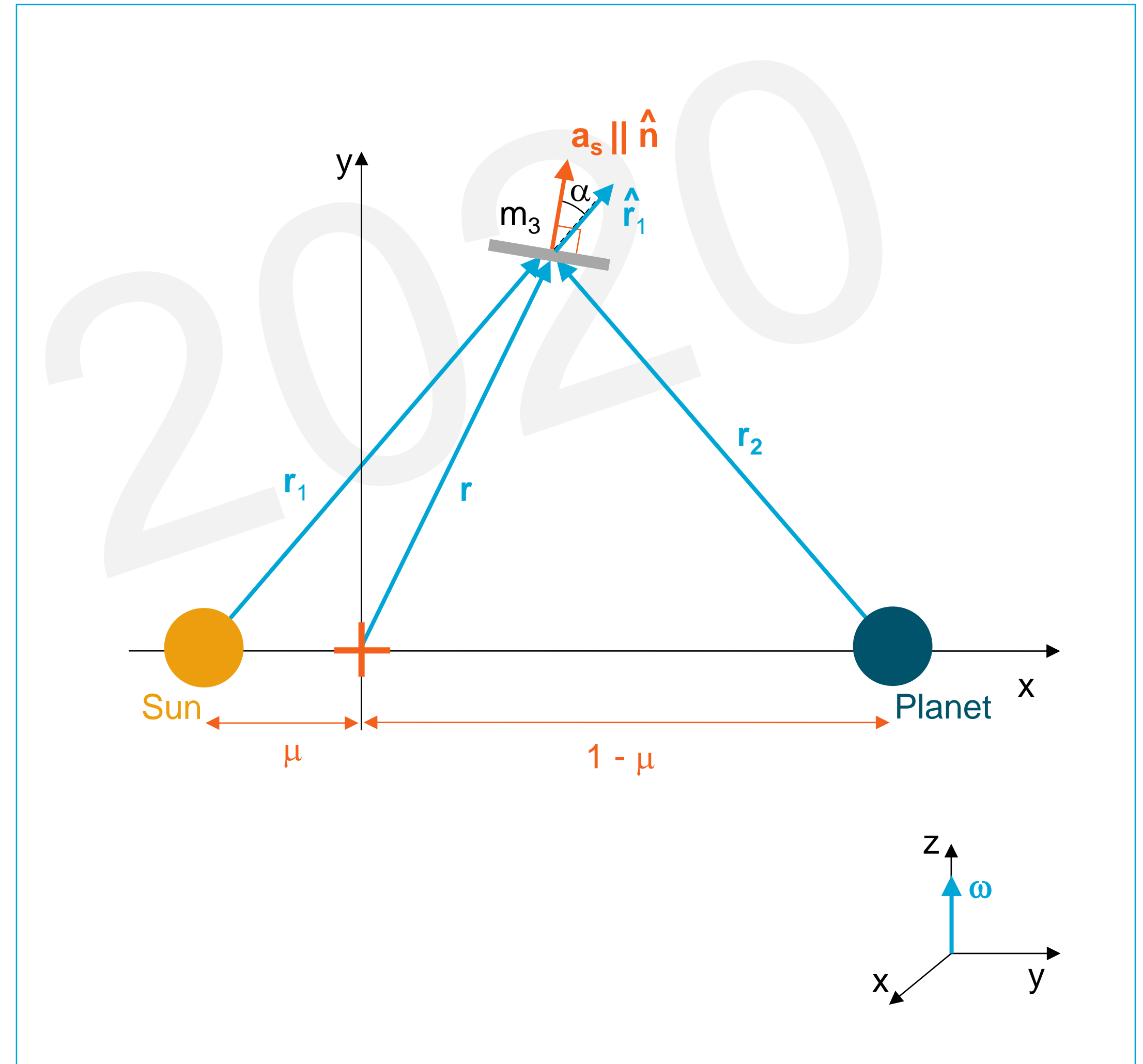
$$\mathbf{a}_s = \beta \frac{\mu_{\odot}}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

- Dimensionless solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

- More **general form** of solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}}$$



Solar-sail acceleration in the dimensionless cr3bp

- Dimensional solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{\mu_{\odot}}{r^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

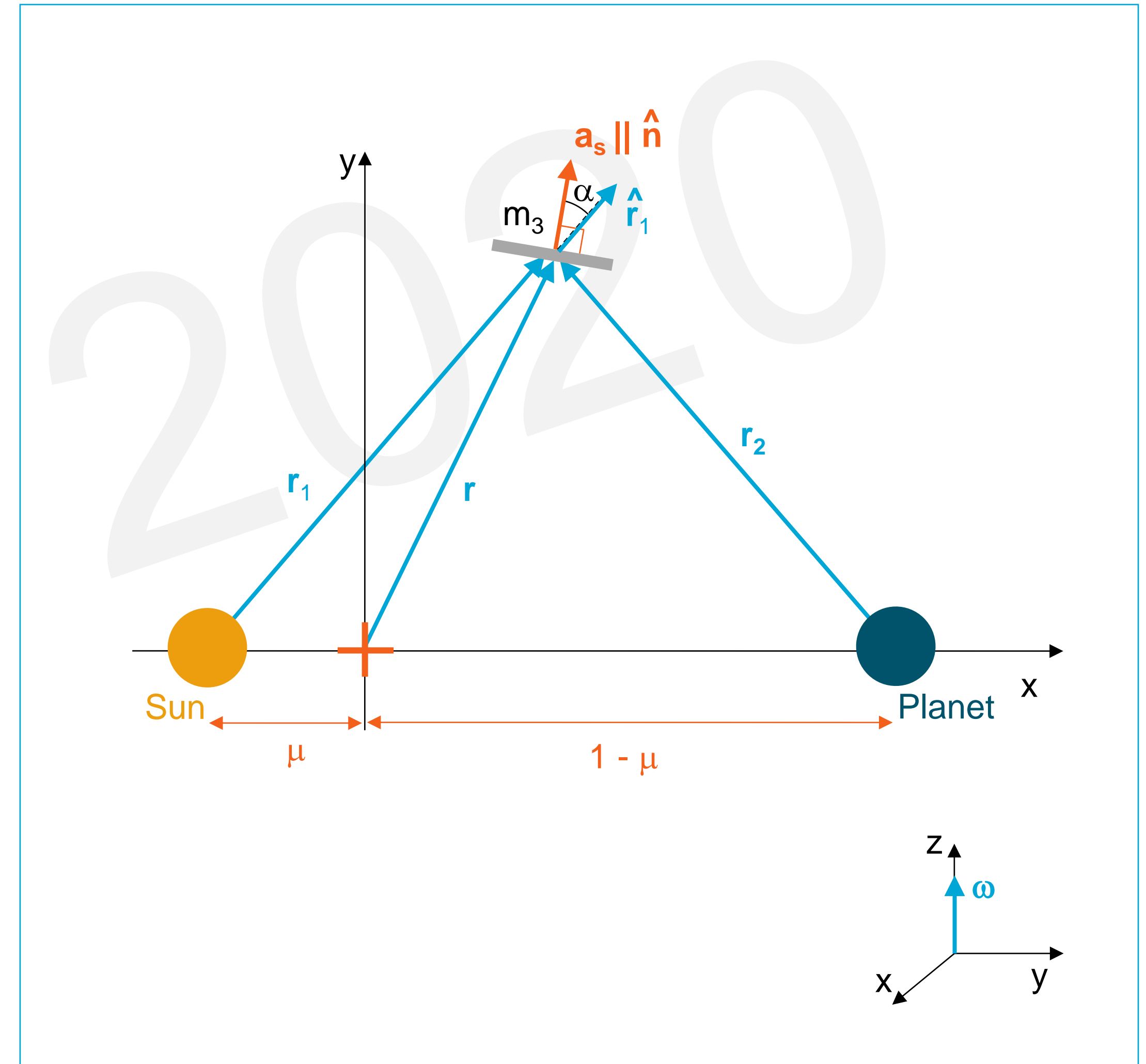
- Dimensionless solar-sail acceleration in

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\cos \alpha)^2 \hat{\mathbf{n}}$$

- More **general form** of solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}}$$

$\cos \alpha = \hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}}$

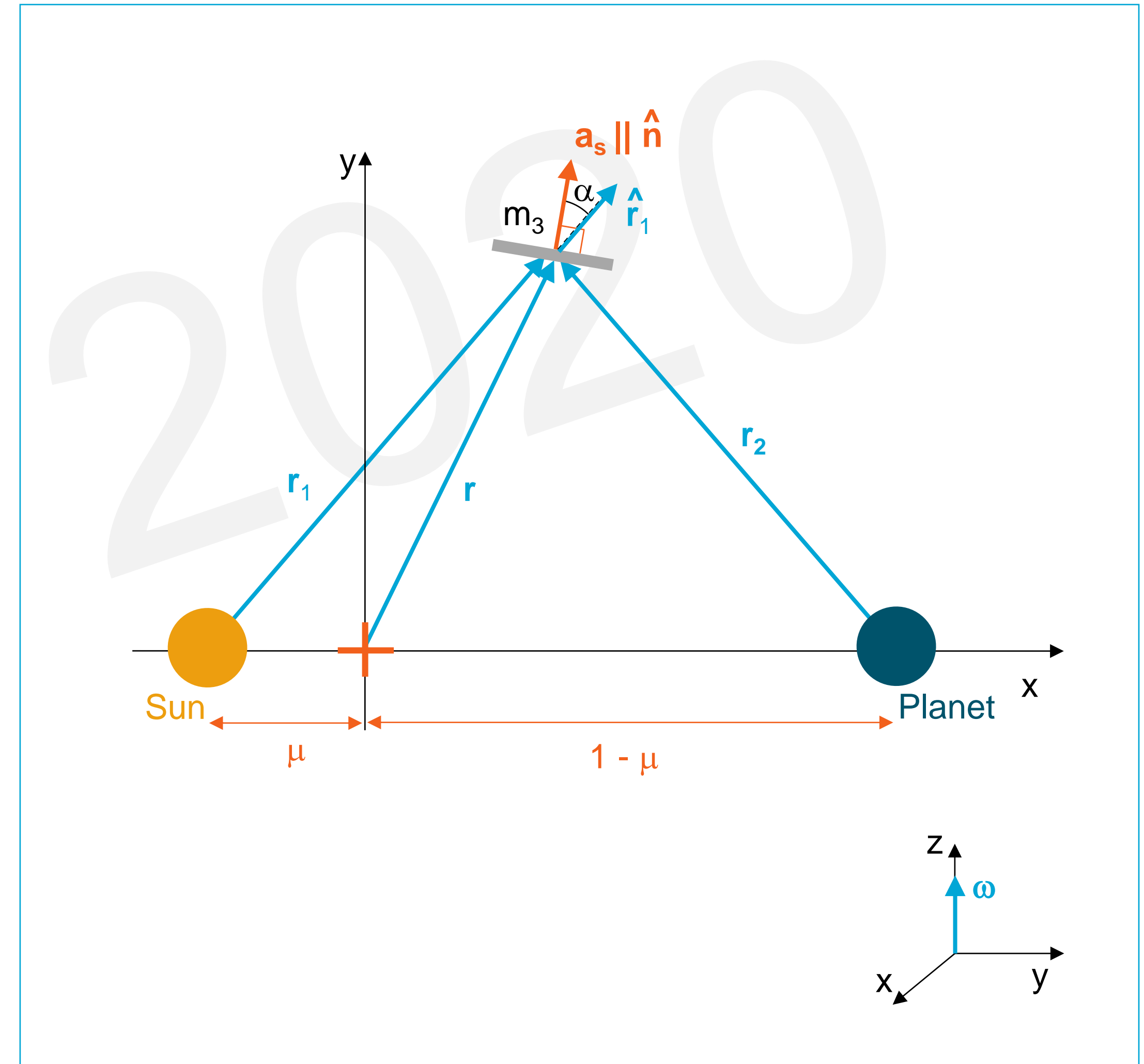


Solar-sail acceleration in the dimensionless cr3bp

- Dimensionless solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}}$$

- Define the sail orientation in **3D**



Solar-sail acceleration in the dimensionless cr3bp

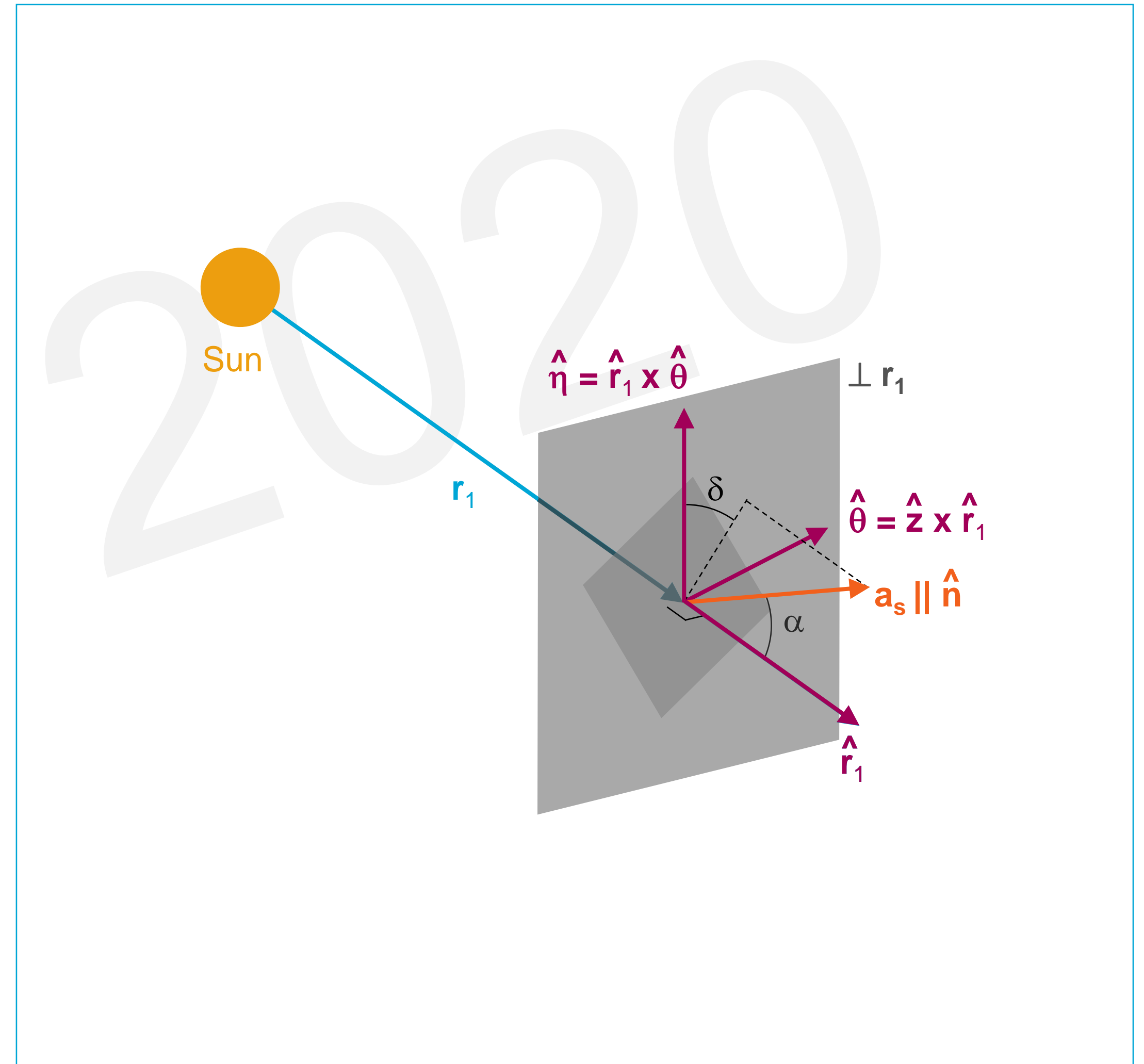
- Dimensionless solar-sail acceleration

$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}}$$

- Define the sail orientation in **3D**

$$\hat{\mathbf{n}}|_{(\hat{\mathbf{r}}_1, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\eta}})} = \begin{bmatrix} \cos \alpha \\ \sin \alpha \sin \delta \\ \sin \alpha \cos \delta \end{bmatrix}$$

α = cone angle
 δ = clock angle



Solar-sail acceleration in the dimensionless cr3bp

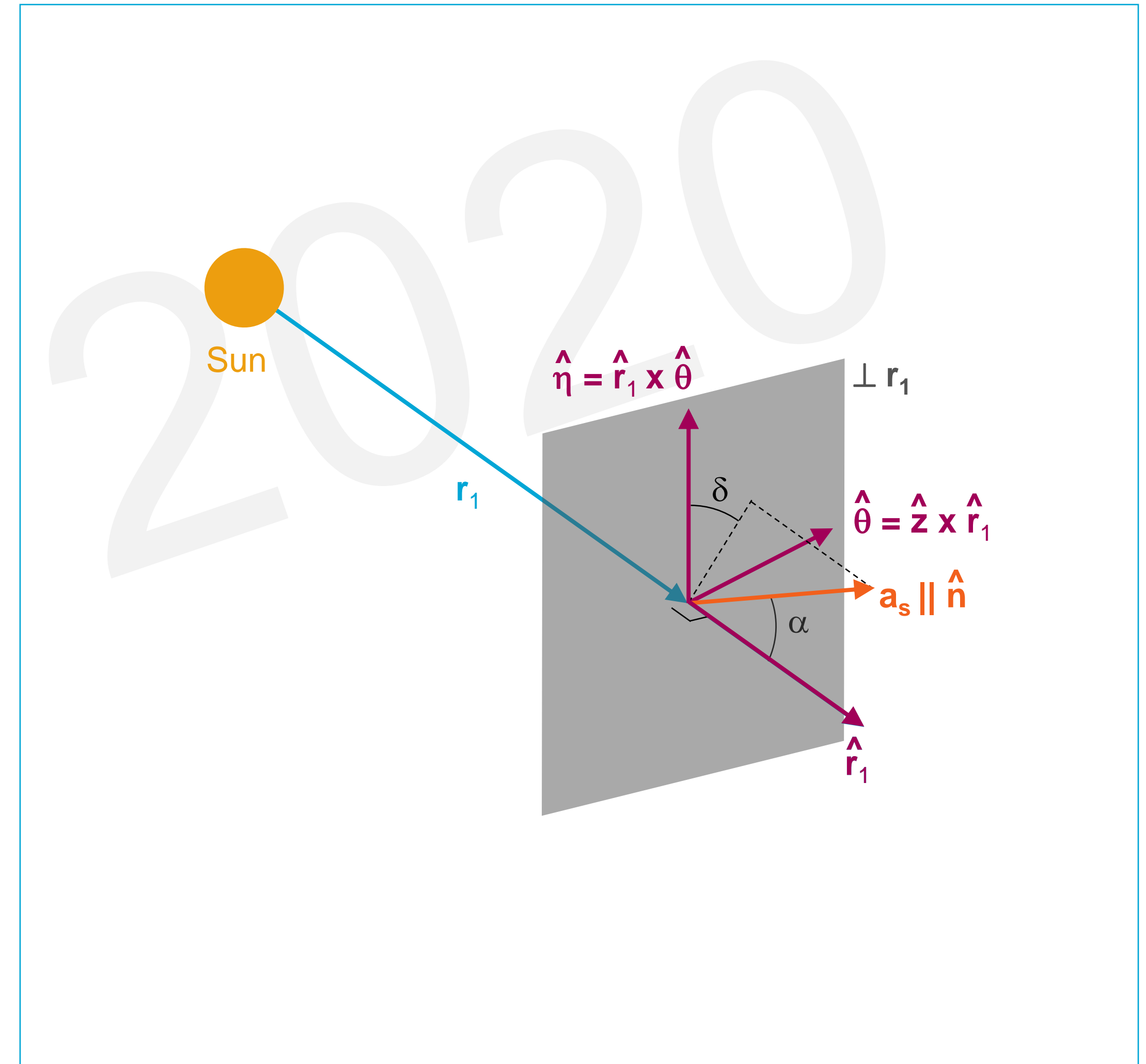
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Reference frame transformation!



Solar-sail acceleration in the dimensionless cr3bp

- Dimensionless solar-sail acceleration

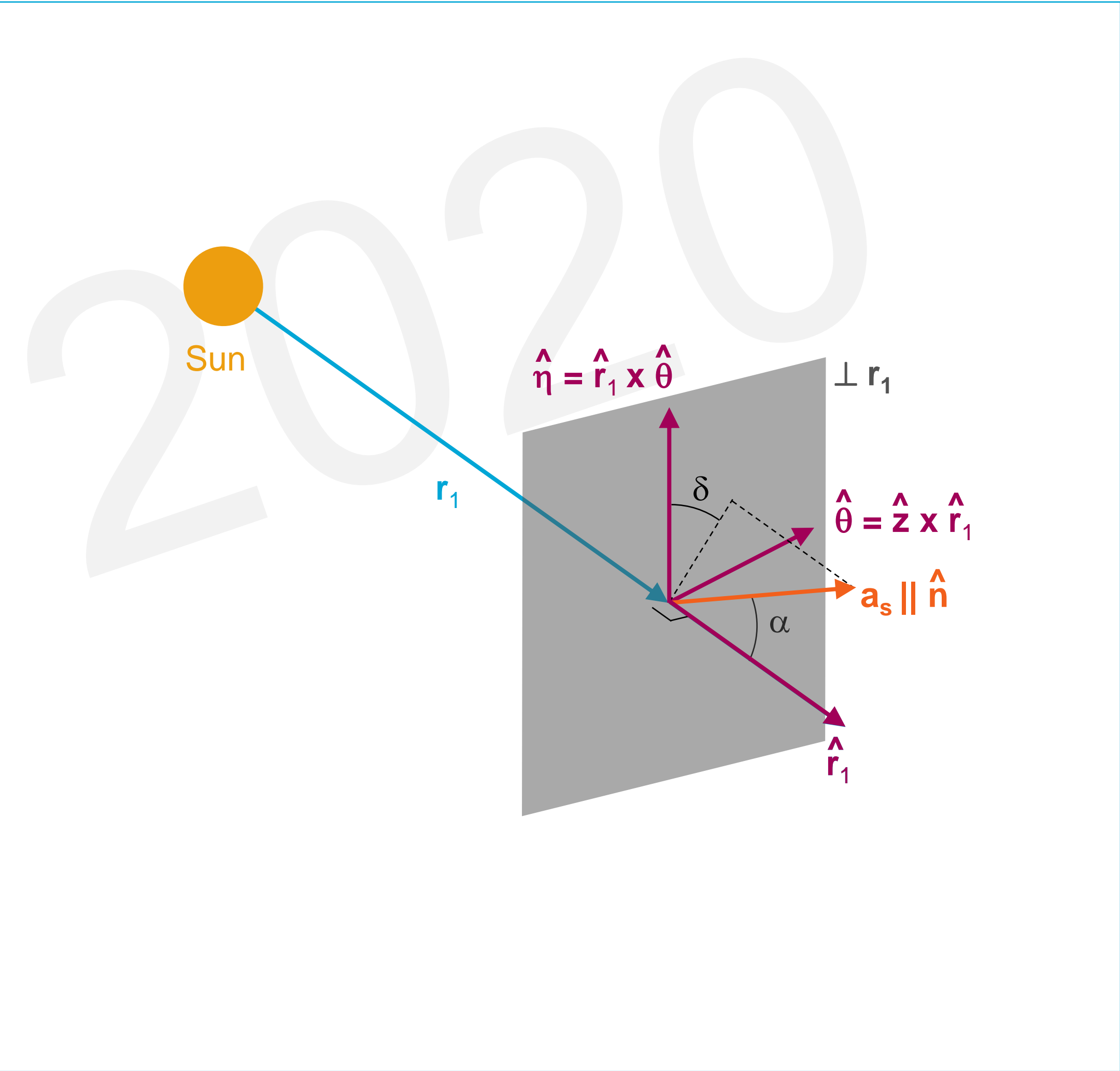
$$\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}}$$

- Define the sail orientation in **3D**

$$\hat{\mathbf{n}}|_{(\hat{\mathbf{r}}_1, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\eta}})} = \begin{bmatrix} \cos \alpha \\ \sin \alpha \sin \delta \\ \sin \alpha \cos \delta \end{bmatrix}$$

$$\hat{\mathbf{n}}|_{(\hat{\mathbf{x}}, \hat{\mathbf{y}}, \hat{\mathbf{z}})} = \hat{\mathbf{n}} = \begin{bmatrix} \hat{\mathbf{r}}_1 & \hat{\boldsymbol{\theta}} & \hat{\boldsymbol{\eta}} \end{bmatrix} \hat{\mathbf{n}}|_{(\hat{\mathbf{r}}_1, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\eta}})}$$

Reference frame transformation!



Solar-sail perturbed cr3bp

- Equations of motion

$$\ddot{\mathbf{r}} = -2\boldsymbol{\omega} \times \dot{\mathbf{r}} - \nabla U + \mathbf{a}_s$$

- $\mathbf{r} = [x \ y \ z]^T, \dot{\mathbf{r}} = [\dot{x} \ \dot{y} \ \dot{z}]^T, \ddot{\mathbf{r}} = [\ddot{x} \ \ddot{y} \ \ddot{z}]^T$

- $U = -\left(\frac{1-\mu}{r_1} + \frac{\mu}{r_2}\right) - \frac{1}{2}(x^2 + y^2)$

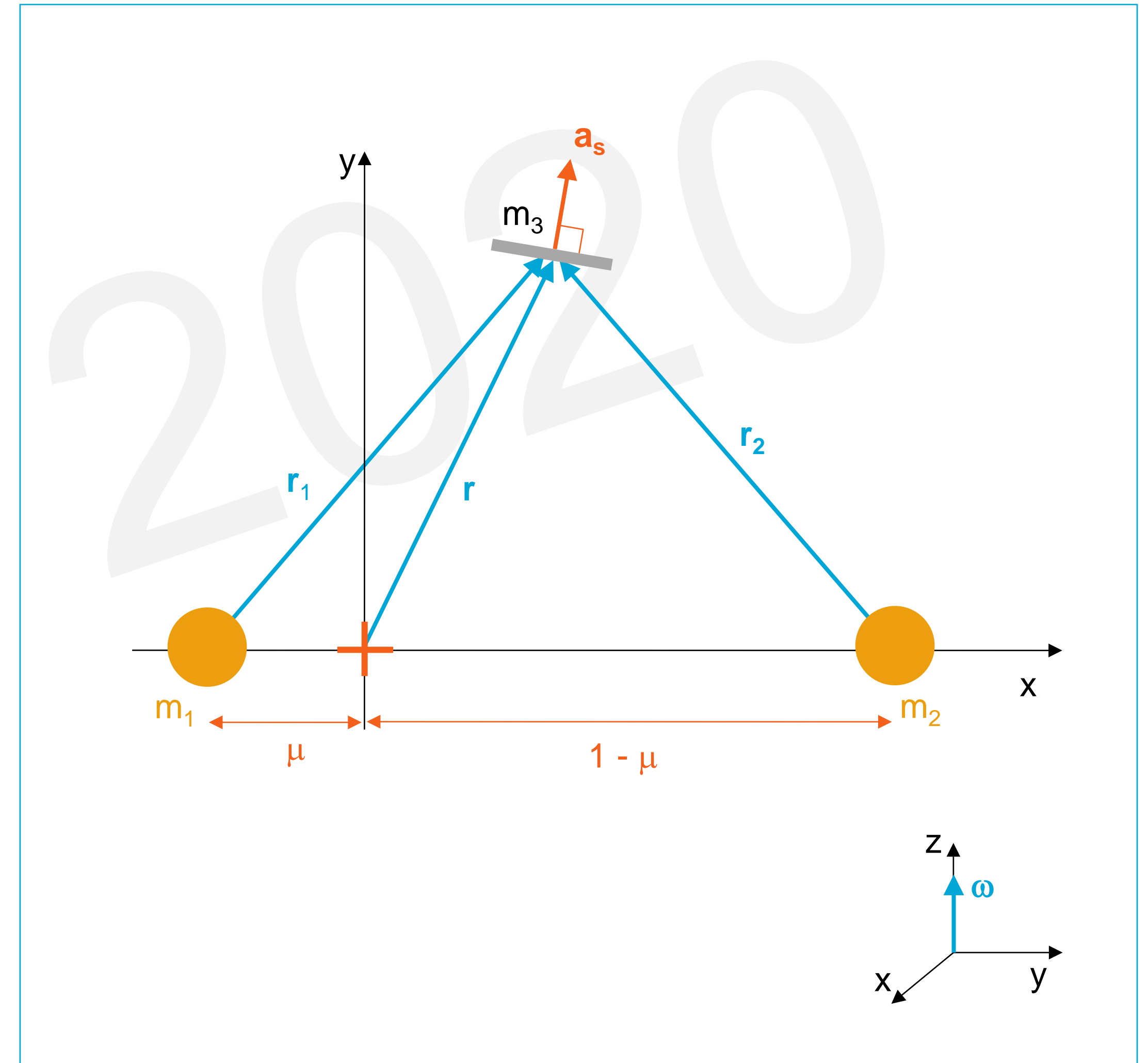
- $\mu = \frac{m_2}{m_1 + m_2}$

- $\mathbf{r}_1 = [x + \mu \ y \ z]^T$

- $\mathbf{r}_2 = [x - (1 - \mu) \ y \ z]^T$

- $\boldsymbol{\omega} = [0 \ 0 \ \omega]^T = [0 \ 0 \ 1]^T$

- $\mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}}$



Solar-sail perturbed cr3bp

- Equations of motion

$$\ddot{\mathbf{r}} = -2\boldsymbol{\omega} \times \dot{\mathbf{r}} - \nabla U + \mathbf{a}_s$$

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$$\circ U = -\left(\frac{1-\mu}{r_1} + \frac{\mu}{r_2}\right) - \frac{1}{2}(x^2 + y^2)$$

$$\circ \mu = \frac{m_2}{m_1 + m_2}$$

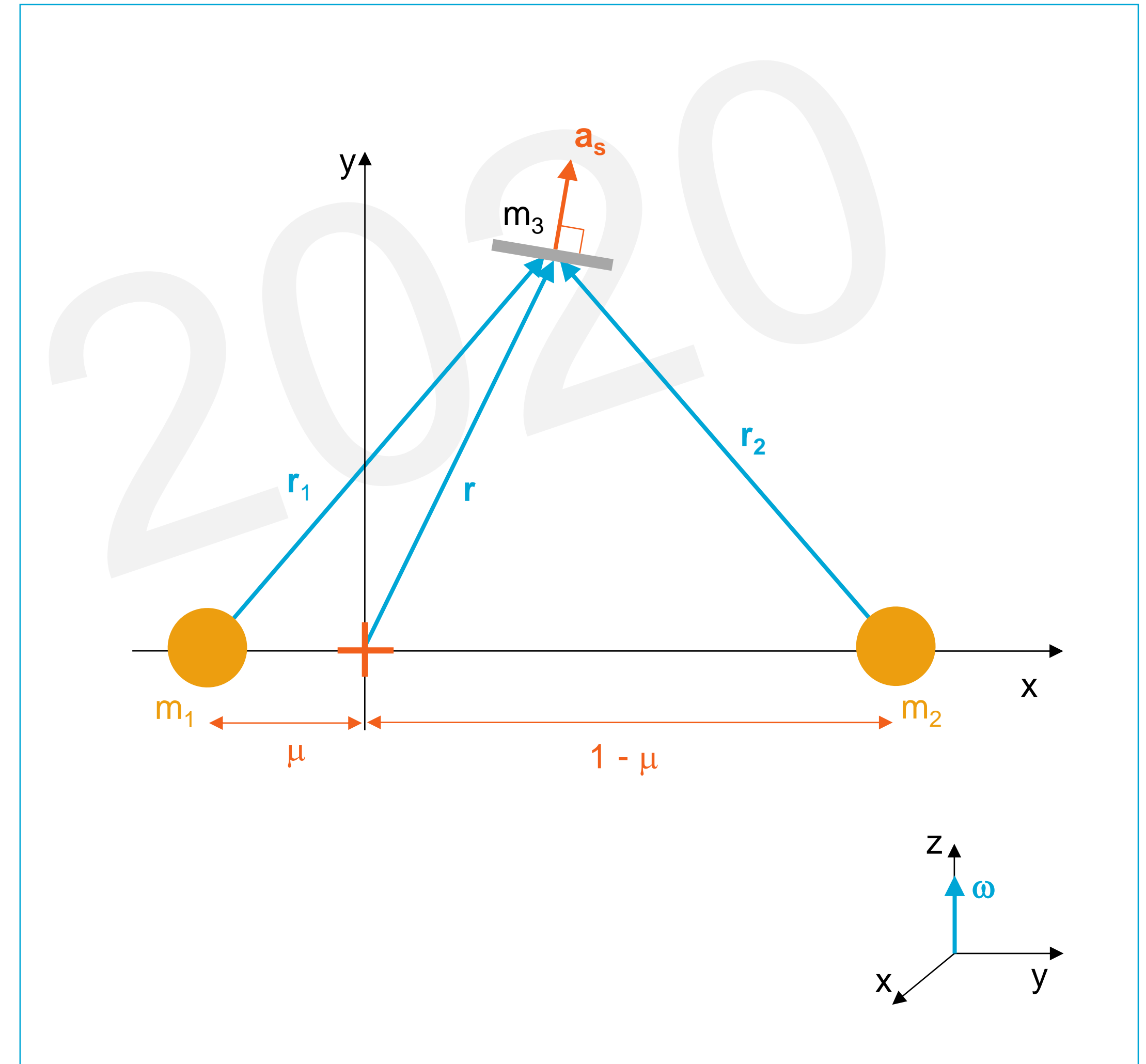
$$\circ \mathbf{r}_1 = [x + \mu \ y \ z]^T$$

$$\circ \mathbf{r}_2 = [x - (1-\mu) \ y \ z]^T$$

$$\circ \boldsymbol{\omega} = [0 \ 0 \ \omega]^T = [0 \ 0 \ 1]^T$$

$$\circ \mathbf{a}_s = \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}}$$

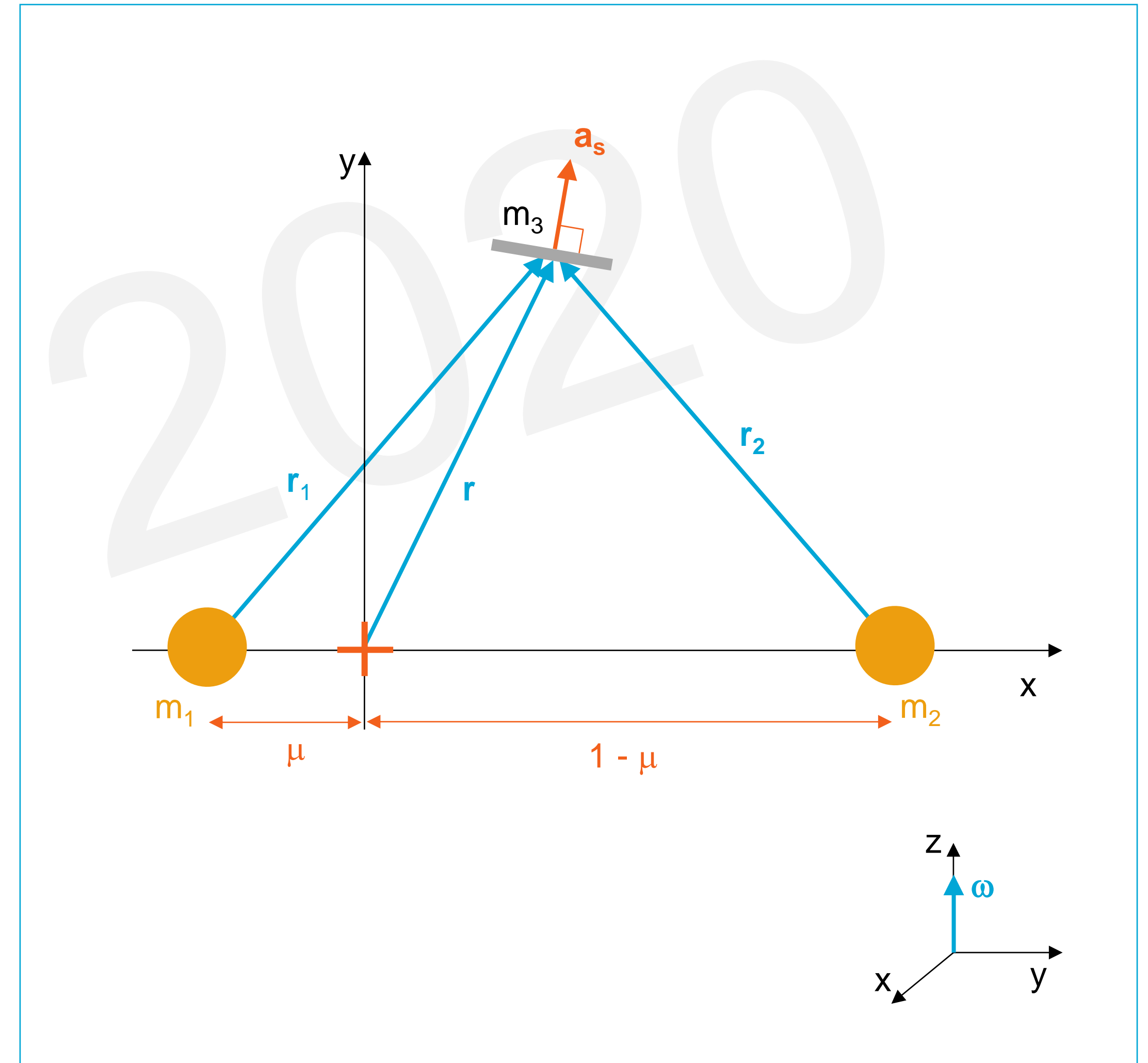
$$\circ \hat{\mathbf{n}} = [\hat{\mathbf{r}}_1 \ \hat{\boldsymbol{\theta}} \ \hat{\boldsymbol{\eta}}] \hat{\mathbf{n}}|_{(\hat{\mathbf{r}}_1, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\eta}})} \quad \hat{\mathbf{n}}|_{(\hat{\mathbf{r}}_1, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\eta}})} = \begin{bmatrix} \cos \alpha \\ \sin \alpha \sin \delta \\ \sin \alpha \cos \delta \end{bmatrix}$$



Solar-sail Lagrange points

- Deriving equilibria in the solar-sail perturbed cr3bp

$$\begin{matrix} \ddot{\mathbf{x}} \\ =0 \end{matrix} = -2\boldsymbol{\omega} \times \begin{matrix} \dot{\mathbf{x}} \\ =0 \end{matrix} - \nabla U + \mathbf{a}_s$$



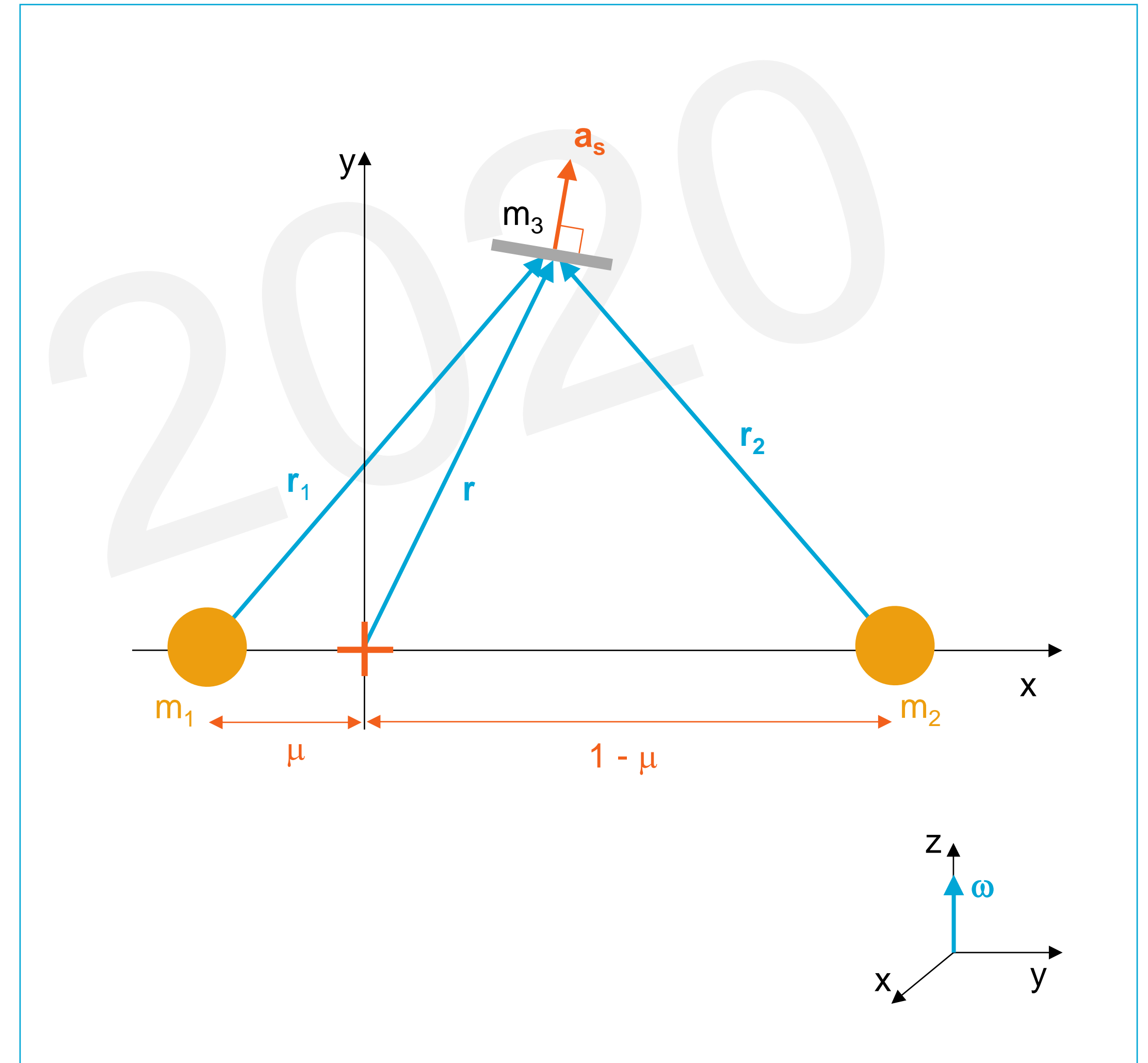
Solar-sail Lagrange points

- Deriving equilibria in the solar-sail perturbed cr3bp

$$\begin{aligned} \ddot{\mathbf{x}} &= -2\boldsymbol{\omega} \times \dot{\mathbf{x}} - \nabla U + \mathbf{a}_s \\ &= 0 \quad \quad = 0 \end{aligned}$$

- Solar-sail Lagrange points (aka **artificial Lagrange points**)

$$\nabla U = \mathbf{a}_s$$



Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration

$$\mathbf{a}_s = \nabla U$$

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration

$$\mathbf{a}_s = \nabla U$$
$$\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} = \nabla U$$

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration
- Take the cross product of both sides

$$\mathbf{a}_s = \nabla U$$
$$\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} = \nabla U$$
$$\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \times \hat{\mathbf{n}} = \nabla U \times \hat{\mathbf{n}}$$

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration
- Take the cross product of both sides
- Rewrite

$$\begin{aligned}\mathbf{a}_s &= \nabla U \\ \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} &= \nabla U \\ \left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \times \hat{\mathbf{n}} &= \nabla U \times \hat{\mathbf{n}} \\ \underbrace{\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \right)}_{\text{Scalar term}} \underbrace{(\hat{\mathbf{n}} \times \hat{\mathbf{n}})}_{=0} &= \nabla U \times \hat{\mathbf{n}}\end{aligned}$$

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration
- Take the cross product of both sides
- Rewrite
- To obtain the required sail orientation

$$\begin{aligned}\mathbf{a}_s &= \nabla U \\ \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} &= \nabla U \\ \left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \times \hat{\mathbf{n}} &= \nabla U \times \hat{\mathbf{n}} \\ \left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \right) (\hat{\mathbf{n}} \times \hat{\mathbf{n}}) &= \nabla U \times \hat{\mathbf{n}}\end{aligned}$$

$$\mathbf{0} = \nabla U \times \hat{\mathbf{n}} \Rightarrow \hat{\mathbf{n}} = \frac{\nabla U}{|\nabla U|}$$

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration
- Take the cross product of both sides
- Rewrite
- To obtain the required sail orientation
- Take the dot product of both sides

$$\mathbf{a}_s = \nabla U$$

$$\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} = \nabla U$$

$$\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \times \hat{\mathbf{n}} = \nabla U \times \hat{\mathbf{n}}$$

$$\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \right) (\hat{\mathbf{n}} \times \hat{\mathbf{n}}) = \nabla U \times \hat{\mathbf{n}}$$

$$\mathbf{0} = \nabla U \times \hat{\mathbf{n}} \Rightarrow \hat{\mathbf{n}} = \frac{\nabla U}{|\nabla U|}$$

$$\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \cdot \hat{\mathbf{n}} = \nabla U \cdot \hat{\mathbf{n}}$$

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration
- Take the cross product of both sides
- Rewrite
- To obtain the required sail orientation
- Take the dot product of both sides
- Rewrite

$$\begin{aligned} \mathbf{a}_s &= \nabla U \\ \beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} &= \nabla U \\ \left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \times \hat{\mathbf{n}} &= \nabla U \times \hat{\mathbf{n}} \\ \left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \right) (\hat{\mathbf{n}} \times \hat{\mathbf{n}}) &= \nabla U \times \hat{\mathbf{n}} \\ \mathbf{0} = \nabla U \times \hat{\mathbf{n}} &\Rightarrow \hat{\mathbf{n}} = \frac{\nabla U}{|\nabla U|} \\ \left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \cdot \hat{\mathbf{n}} &= \nabla U \cdot \hat{\mathbf{n}} \\ \underbrace{\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \right)}_{\text{Scalar term}} \underbrace{(\hat{\mathbf{n}} \cdot \hat{\mathbf{n}})}_{=1} &= \nabla U \cdot \hat{\mathbf{n}} \end{aligned}$$

Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration
- Take the cross product of both sides
- Rewrite
- To obtain the required sail orientation
- Take the dot product of both sides
- Rewrite
- To obtain the required sail performance (i.e., the solar-sail lightness number)

$$\mathbf{a}_s = \nabla U$$

$$\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} = \nabla U$$

$$\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} \right) \times \hat{\mathbf{n}} = \nabla U \times \hat{\mathbf{n}}$$

$$\left(\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \right) (\hat{\mathbf{n}} \times \hat{\mathbf{n}}) = \nabla U \times \hat{\mathbf{n}}$$

$$\mathbf{0} = \nabla U \times \hat{\mathbf{n}} \Rightarrow \hat{\mathbf{n}} = \frac{\nabla U}{|\nabla U|}$$

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Solar-sail Lagrange points – required sail orientation + performance

- Deriving equilibria in the solar-sail perturbed cr3bp
- Required sail acceleration
- Substitute the sail acceleration
- Take the cross product of both sides
- Rewrite
- To obtain the required sail orientation
- Take the dot product of both sides
- Rewrite
- To obtain the required sail performance (i.e., the solar-sail lightness number)

$$\mathbf{a}_s = \nabla U$$

$$\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 \hat{\mathbf{n}} = \nabla U$$

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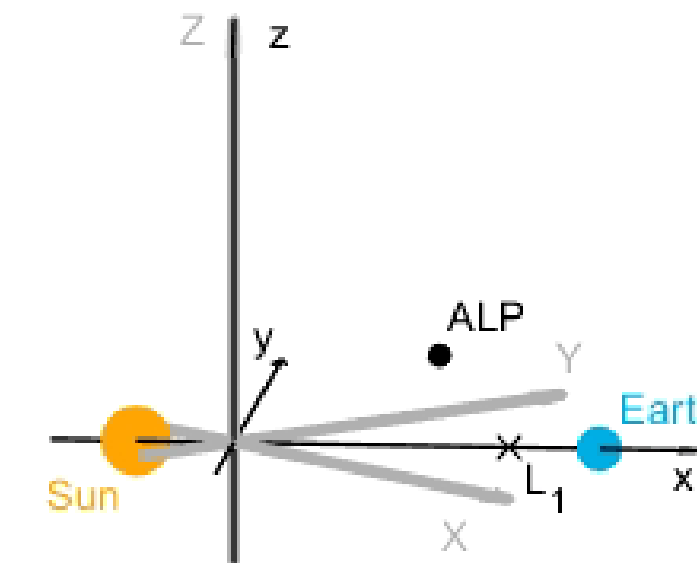
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Only a function
of the position
coordinates (x,y,z)

$$\beta \frac{1-\mu}{r_1^2} (\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2 = \nabla U \cdot \hat{\mathbf{n}} \Rightarrow \beta = \frac{r_1^2}{1-\mu} \frac{\nabla U \cdot \hat{\mathbf{n}}}{(\hat{\mathbf{r}}_1 \cdot \hat{\mathbf{n}})^2}$$

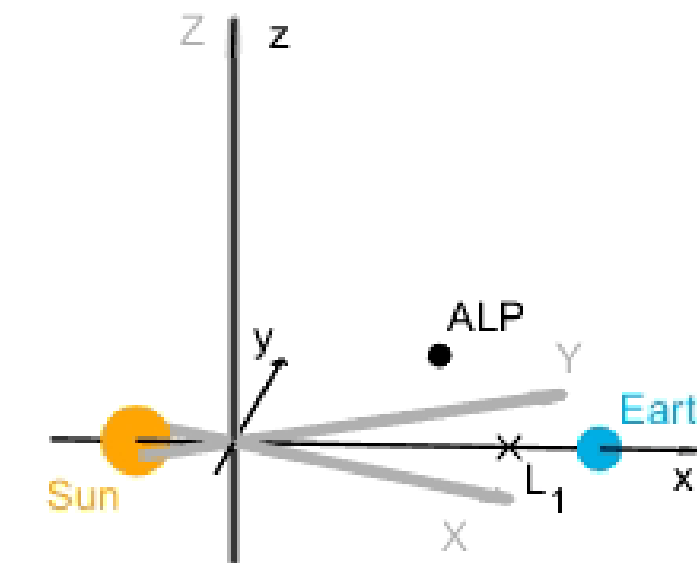
Example

- Let's say we want to create an artificial Lagrange point (ALP)
 - in Sun-Earth cr3bp
 - above the ecliptic
 - Sunward of the L_1 point



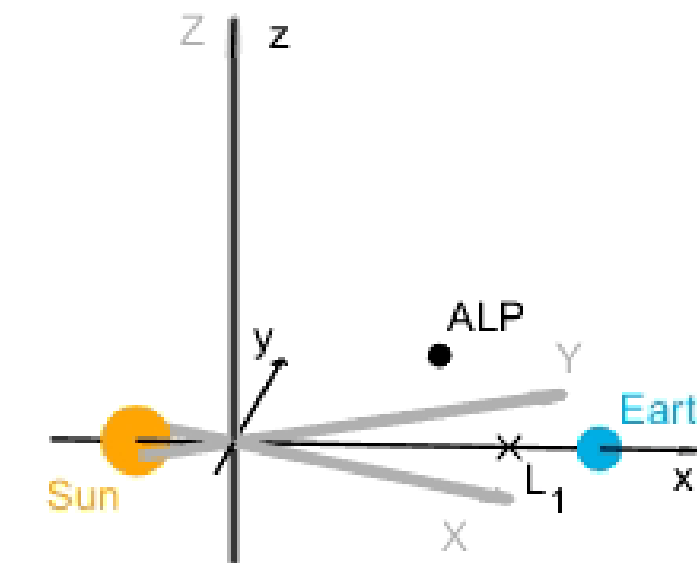
Example

- Let's say we want to create an artificial Lagrange point (ALP)
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- What is the required sail orientation and performance?



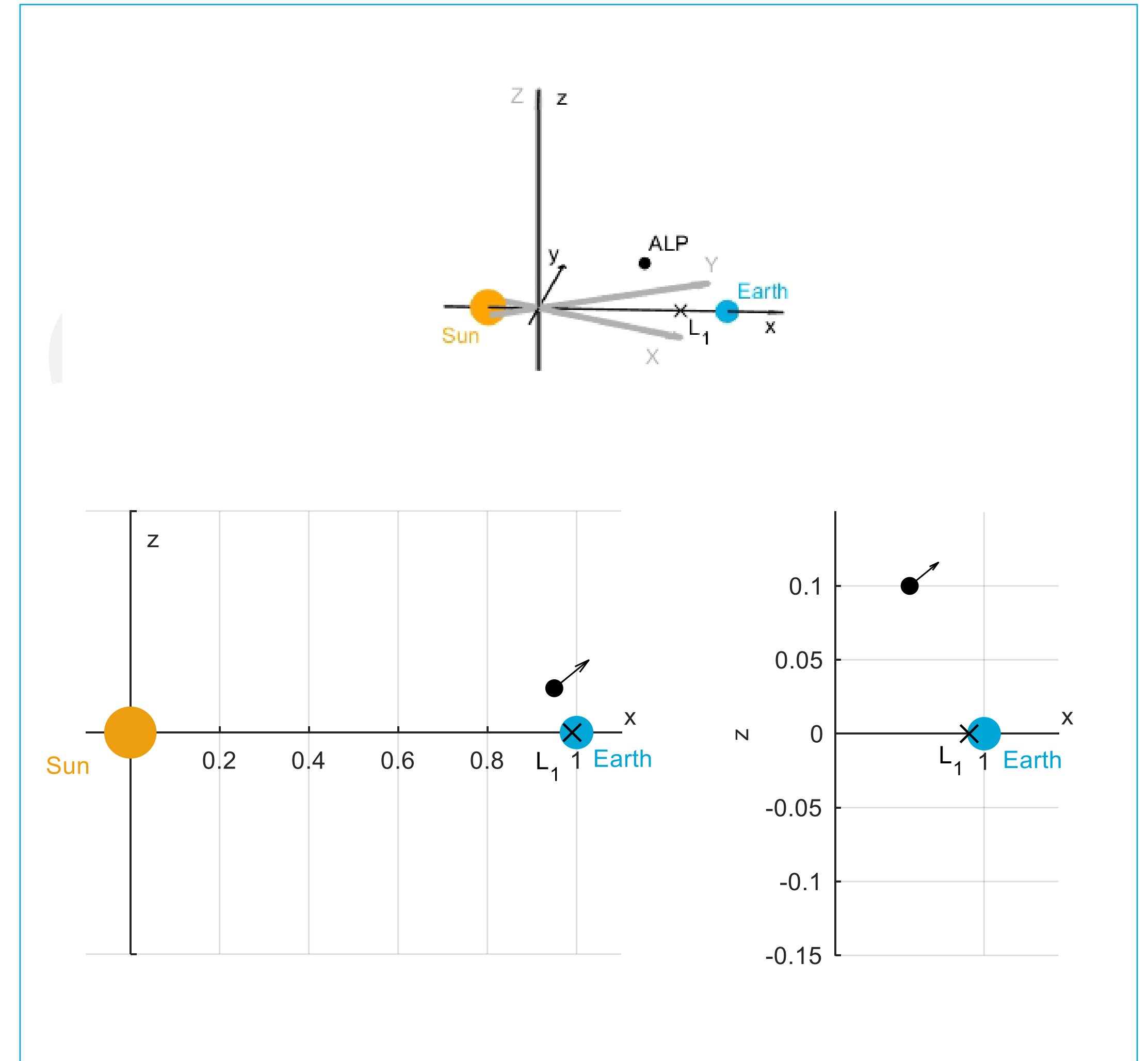
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- Example
 - $(x, y, z) = (0.95, 0, 0.1)$



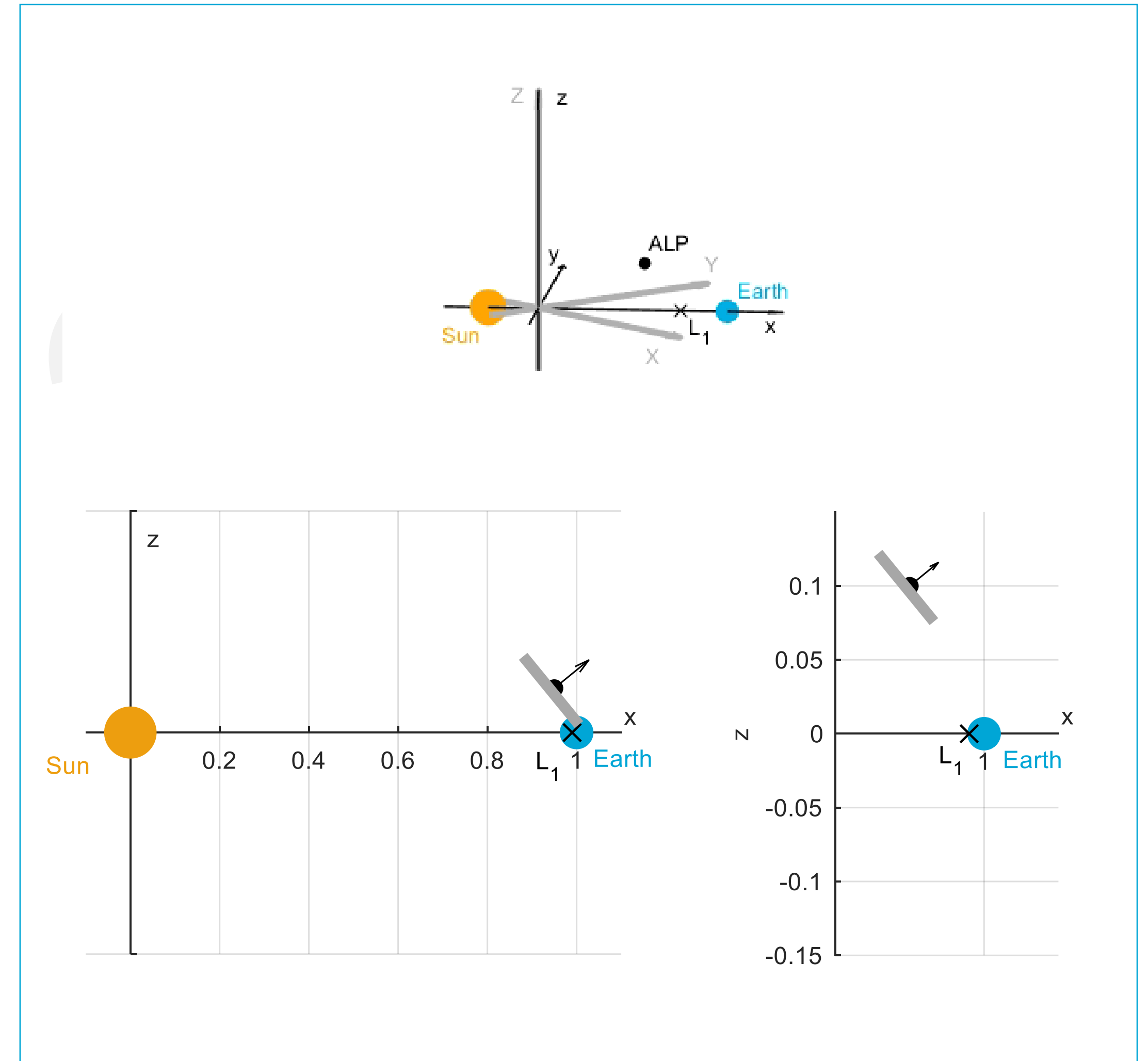
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 - $(x, y, z) = (0.95, 0, 0.1)$
 - $\hat{\mathbf{n}} = \begin{bmatrix} 0.7723 \\ 0 \\ 0.6352 \end{bmatrix}$



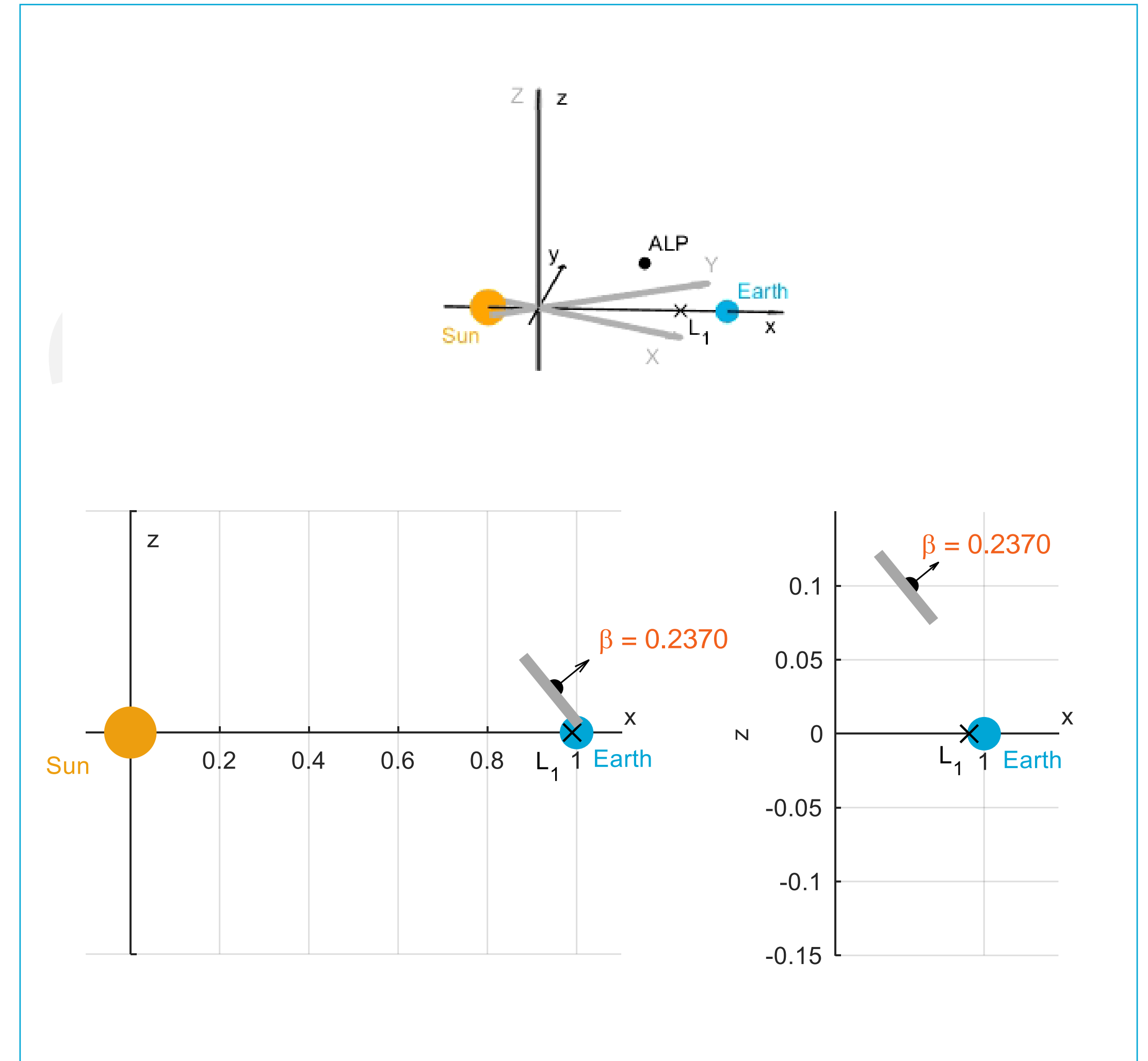
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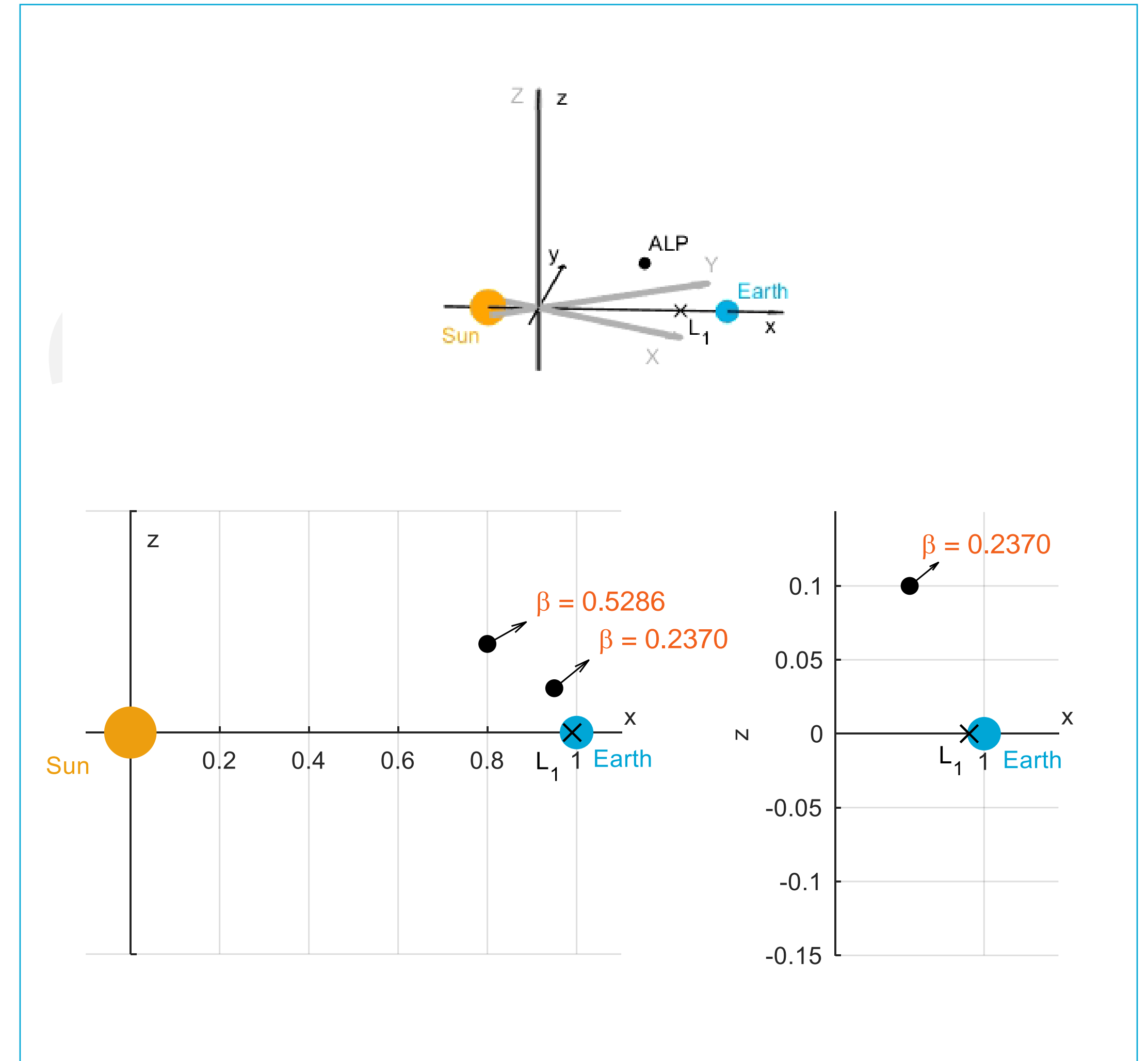
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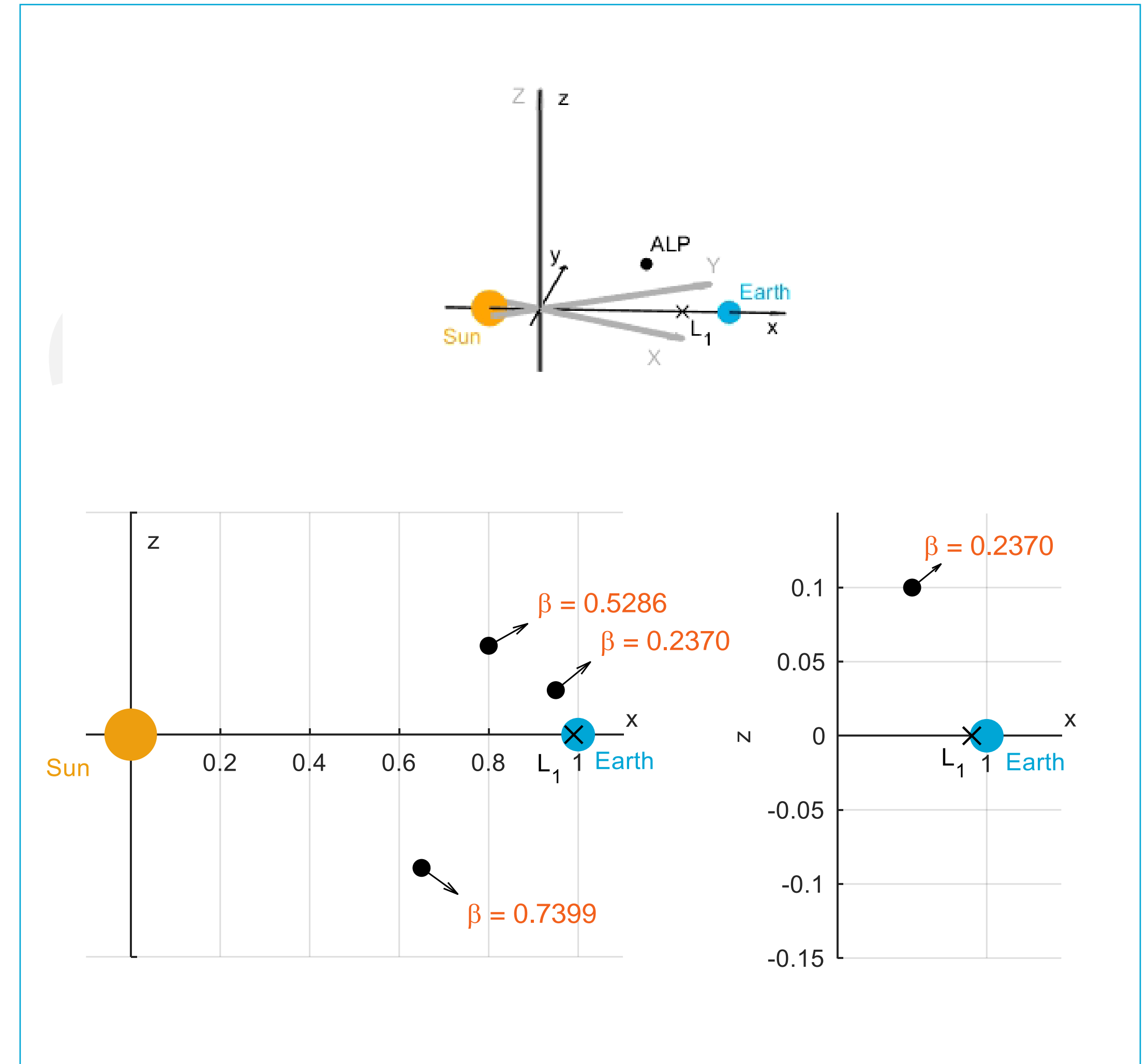
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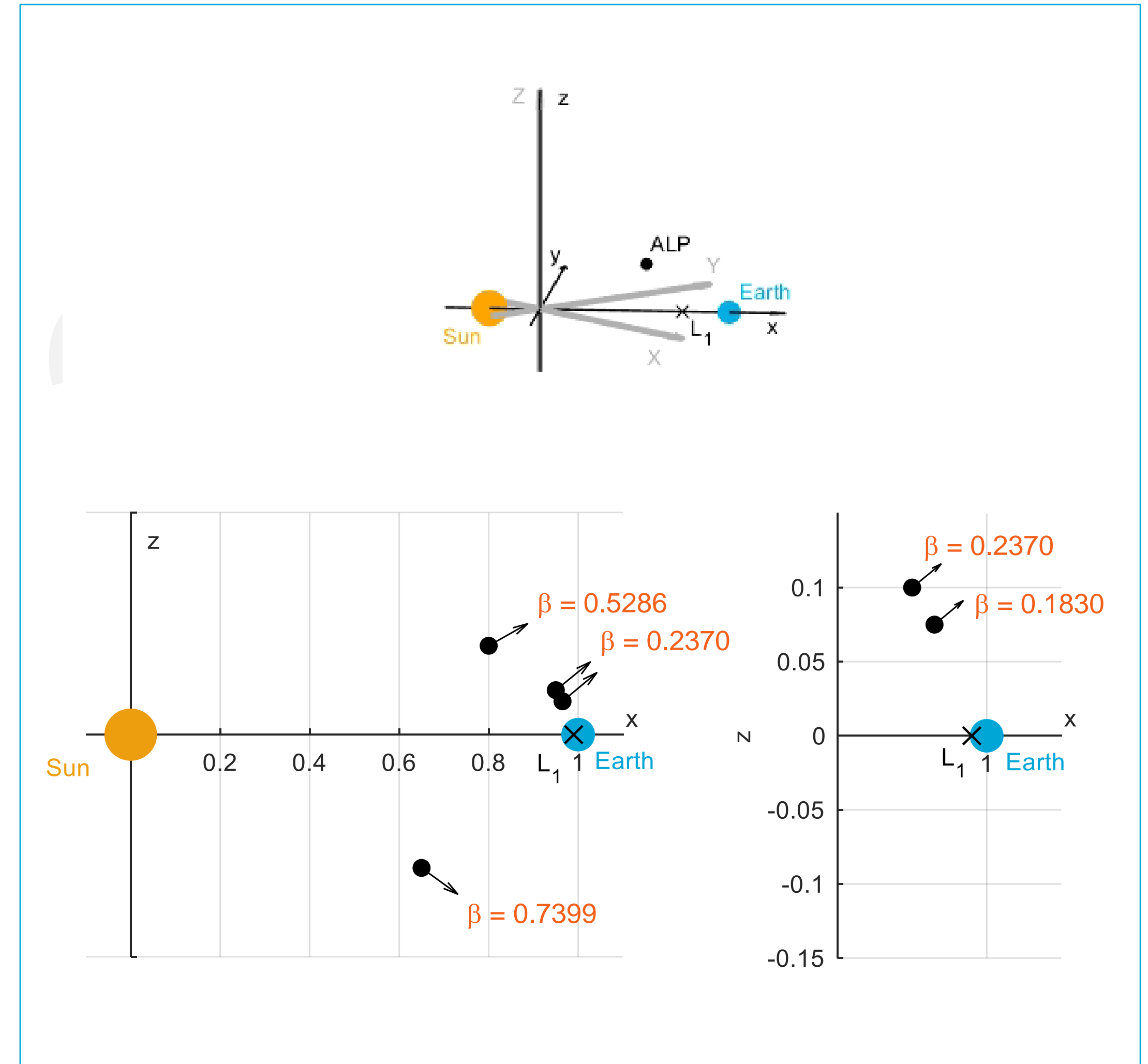
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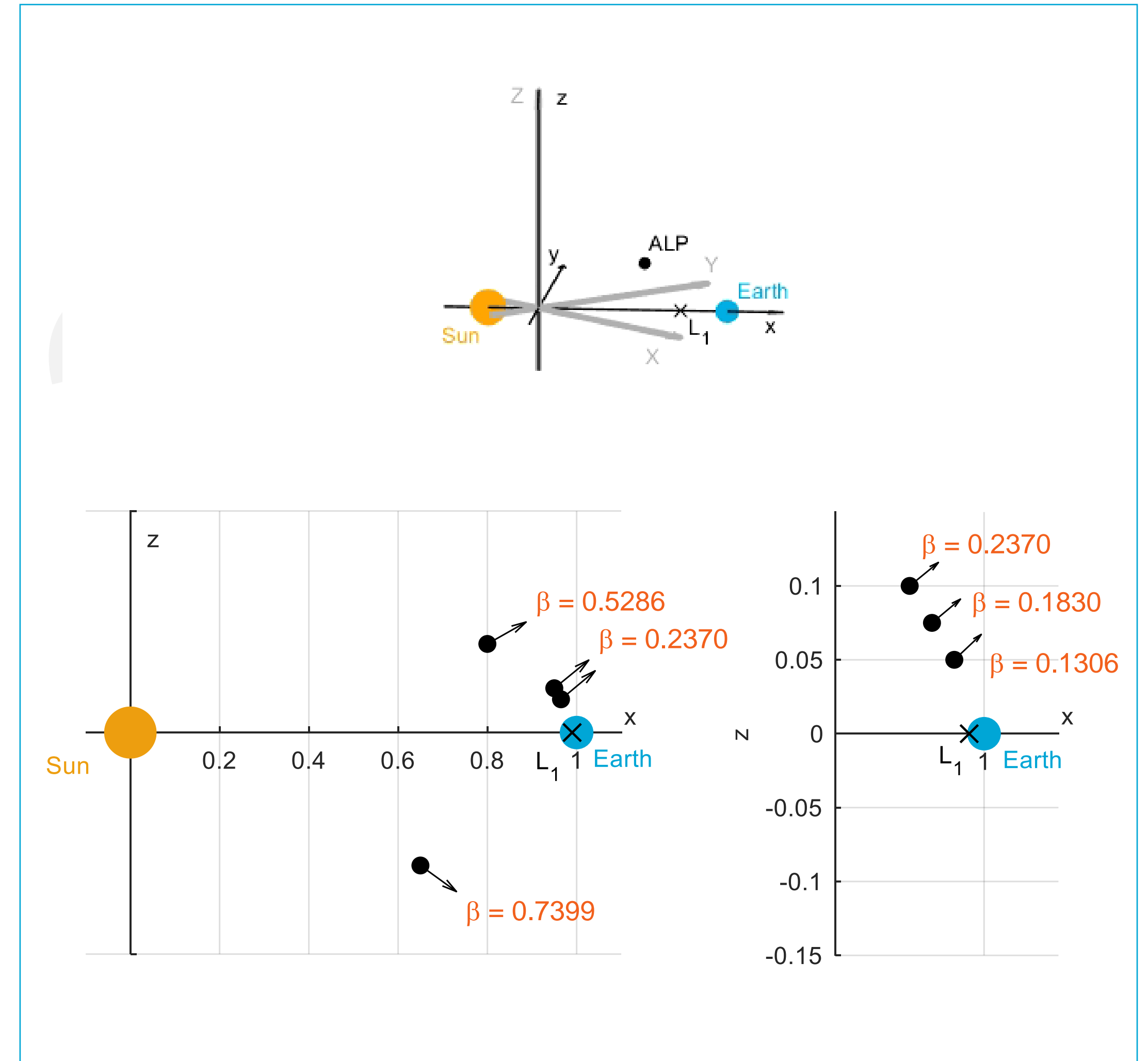
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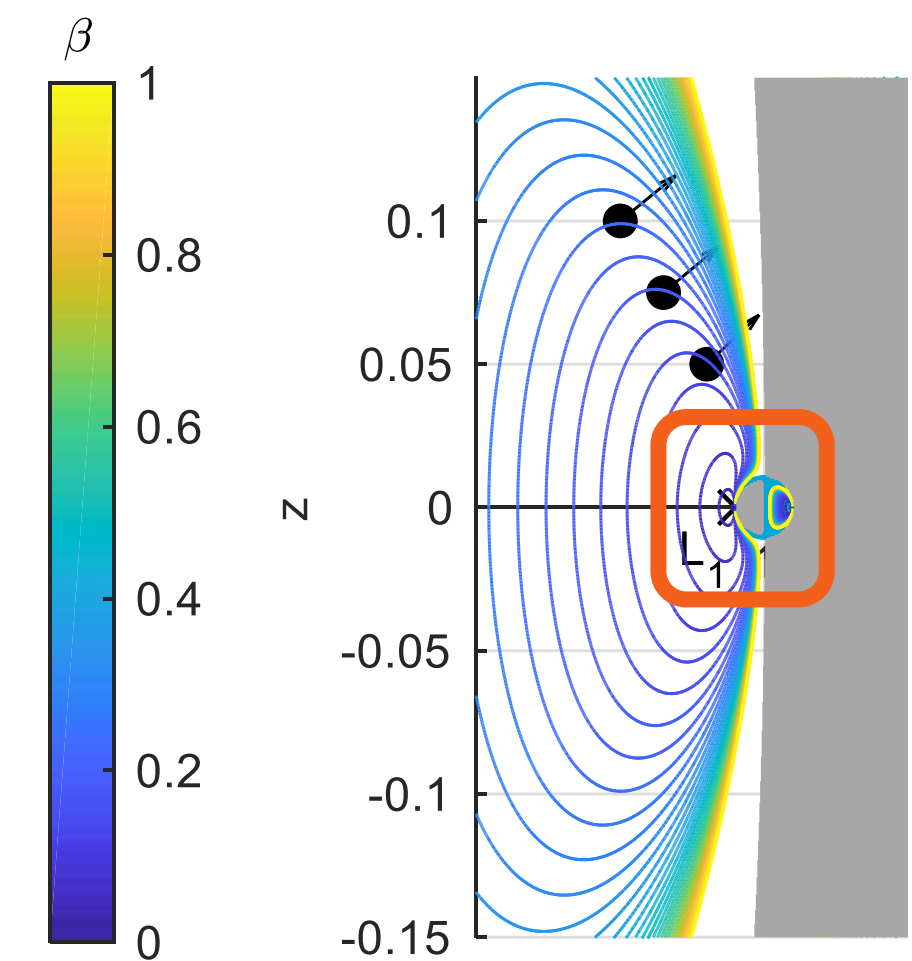
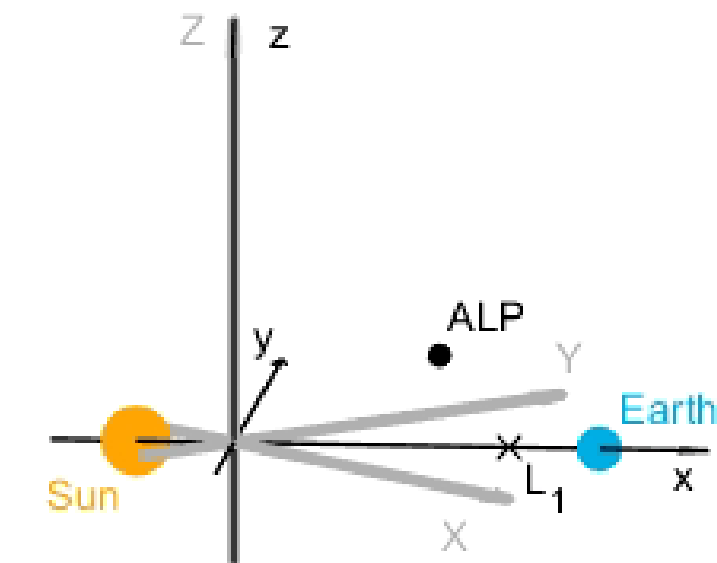
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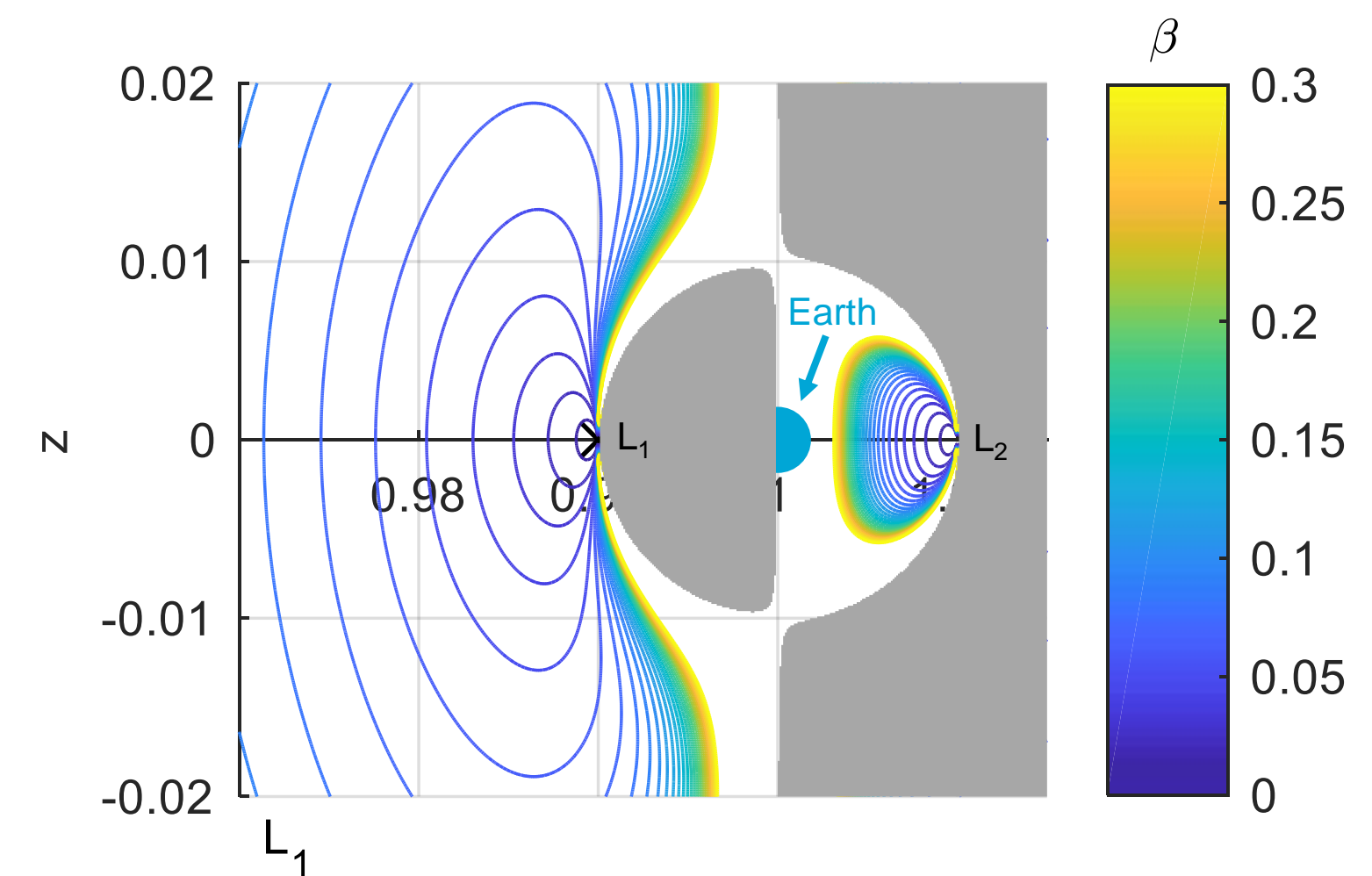
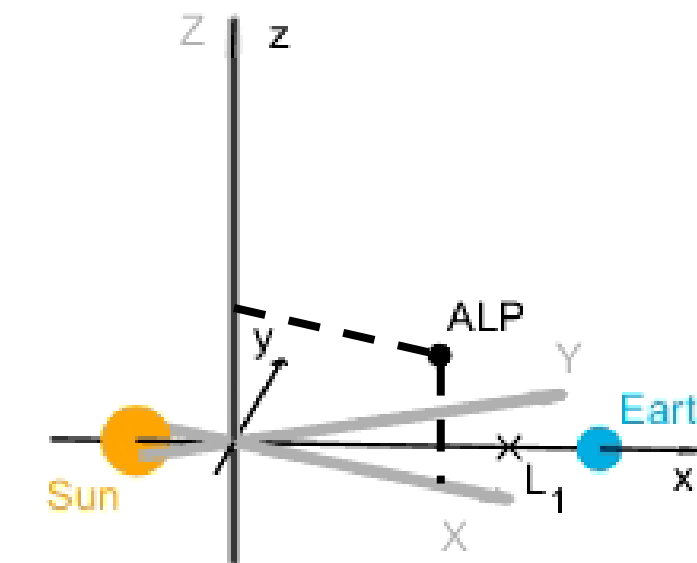
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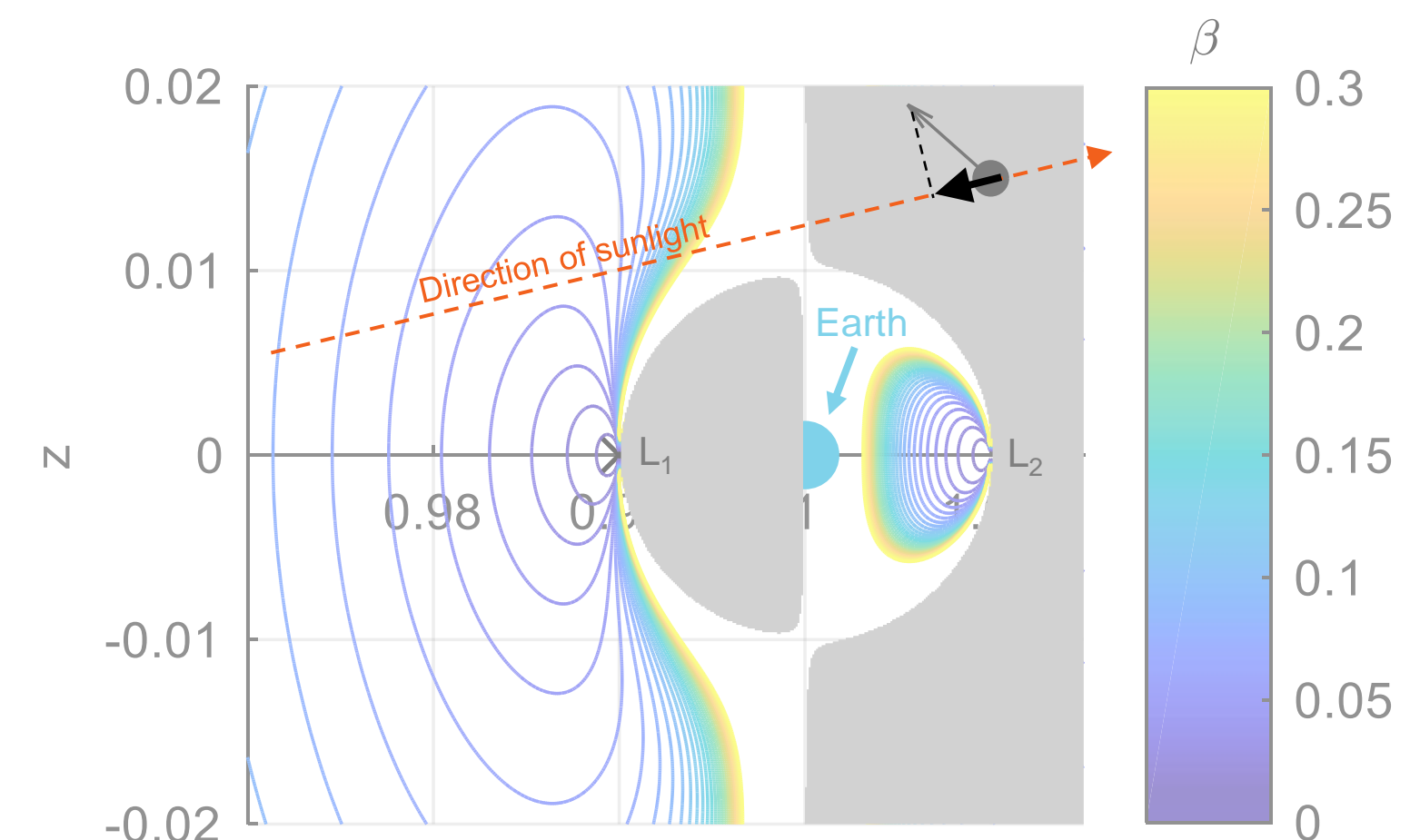
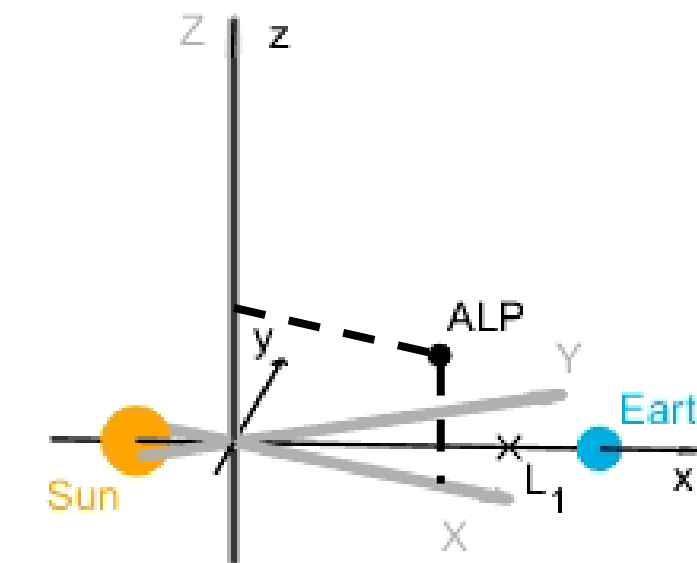
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Artificial Lagrange points - utility

- Previously proposed Sunjammer mission
 - Mass: 45 kg
 - Sail size: 38 x 38 m²
 - Lightness number ~0.04



Credit: NASA

Artificial Lagrange points - utility

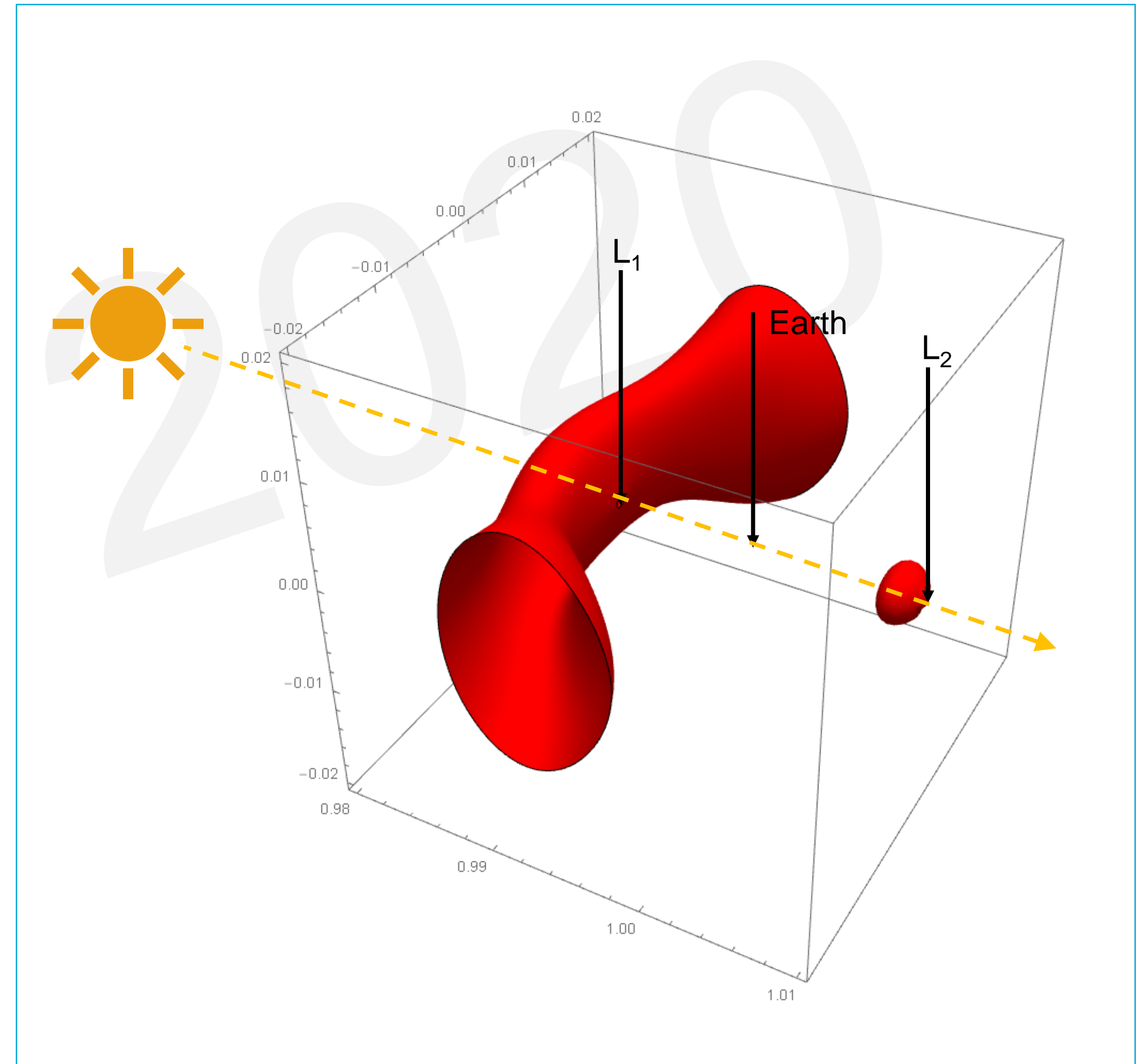
- Previously proposed Sunjammer mission
 - Mass: 45 kg
 - Sail size: 38 x 38 m²
 - Lightness number ~0.04
- **Mission objective** – increase the warning time for solar storms



Credit: NASA

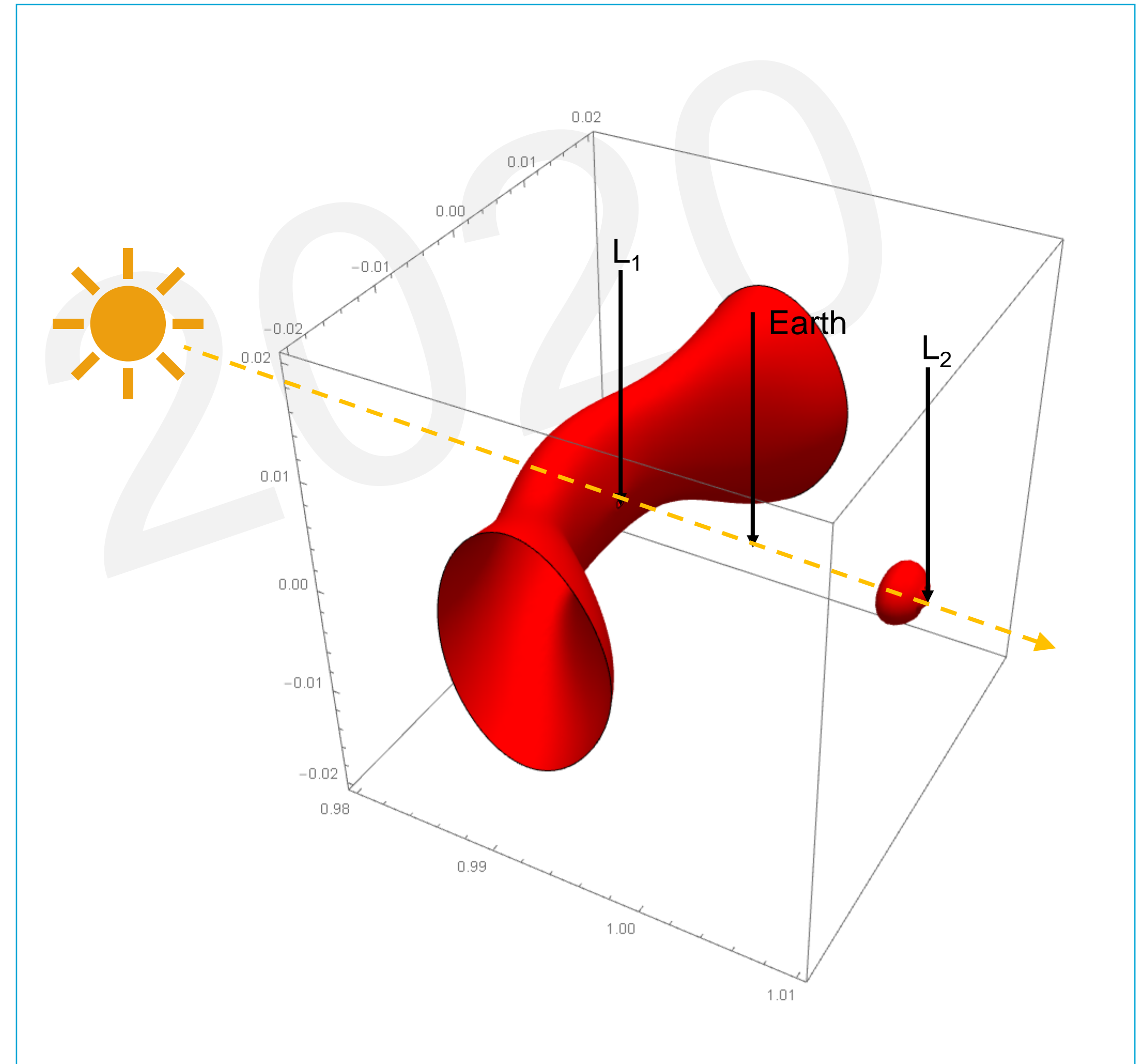
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- Mission objective – increase the warning time for solar storms
- **Mission destination** – sub- L_1 point



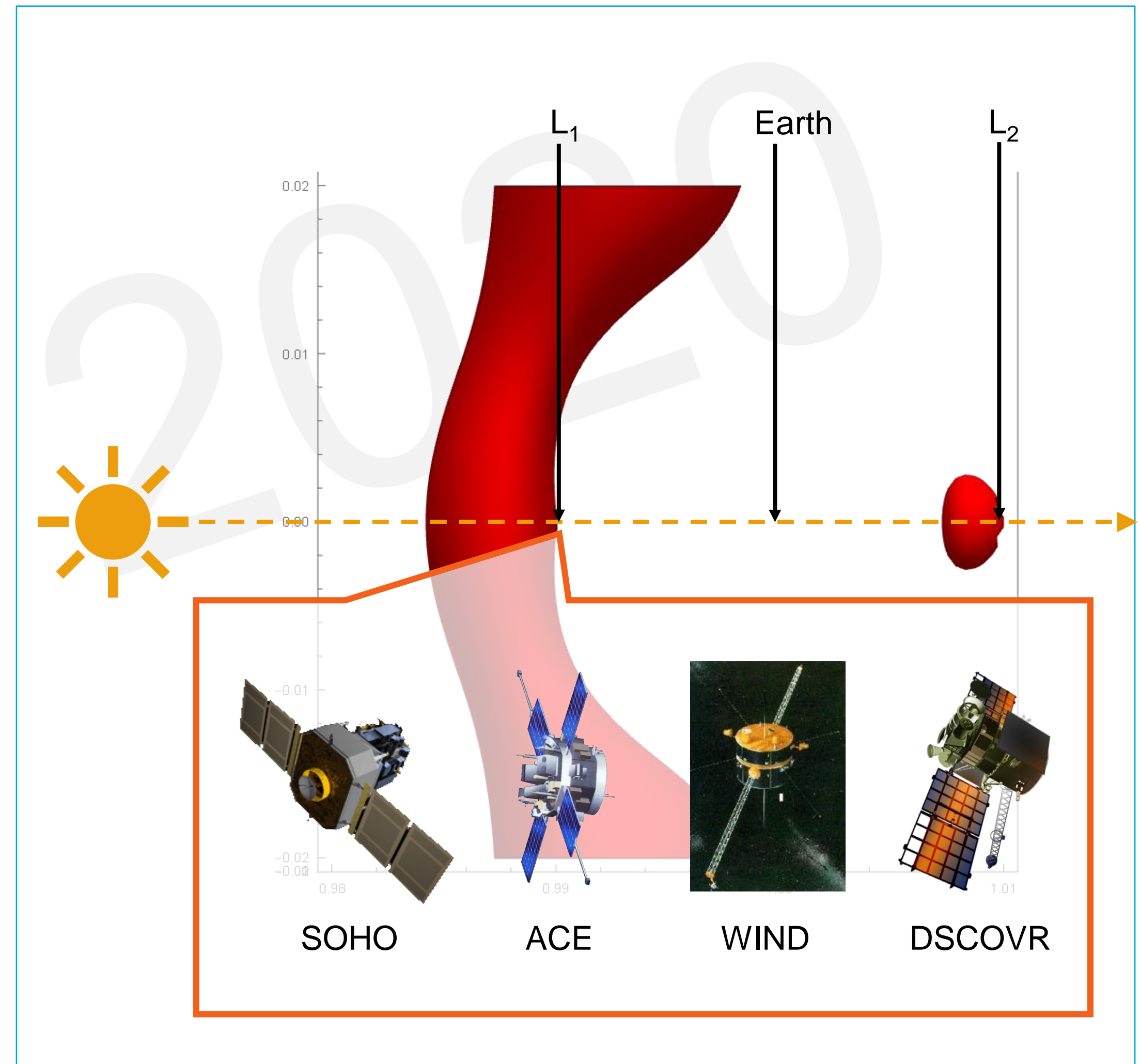
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- Mission destination – sub- L_1 point
- Satellites with similar objective are currently stationed at the L_1 point



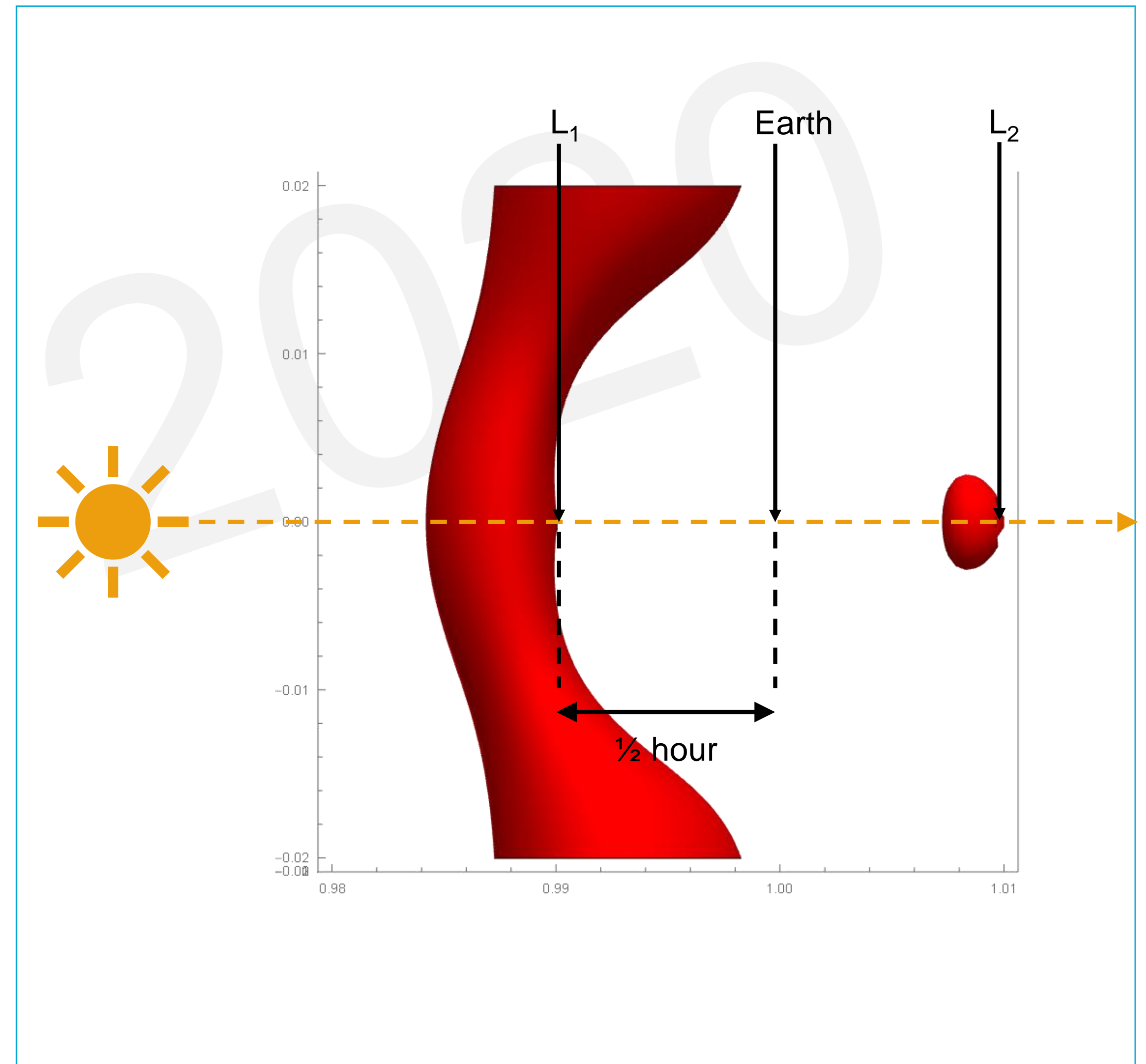
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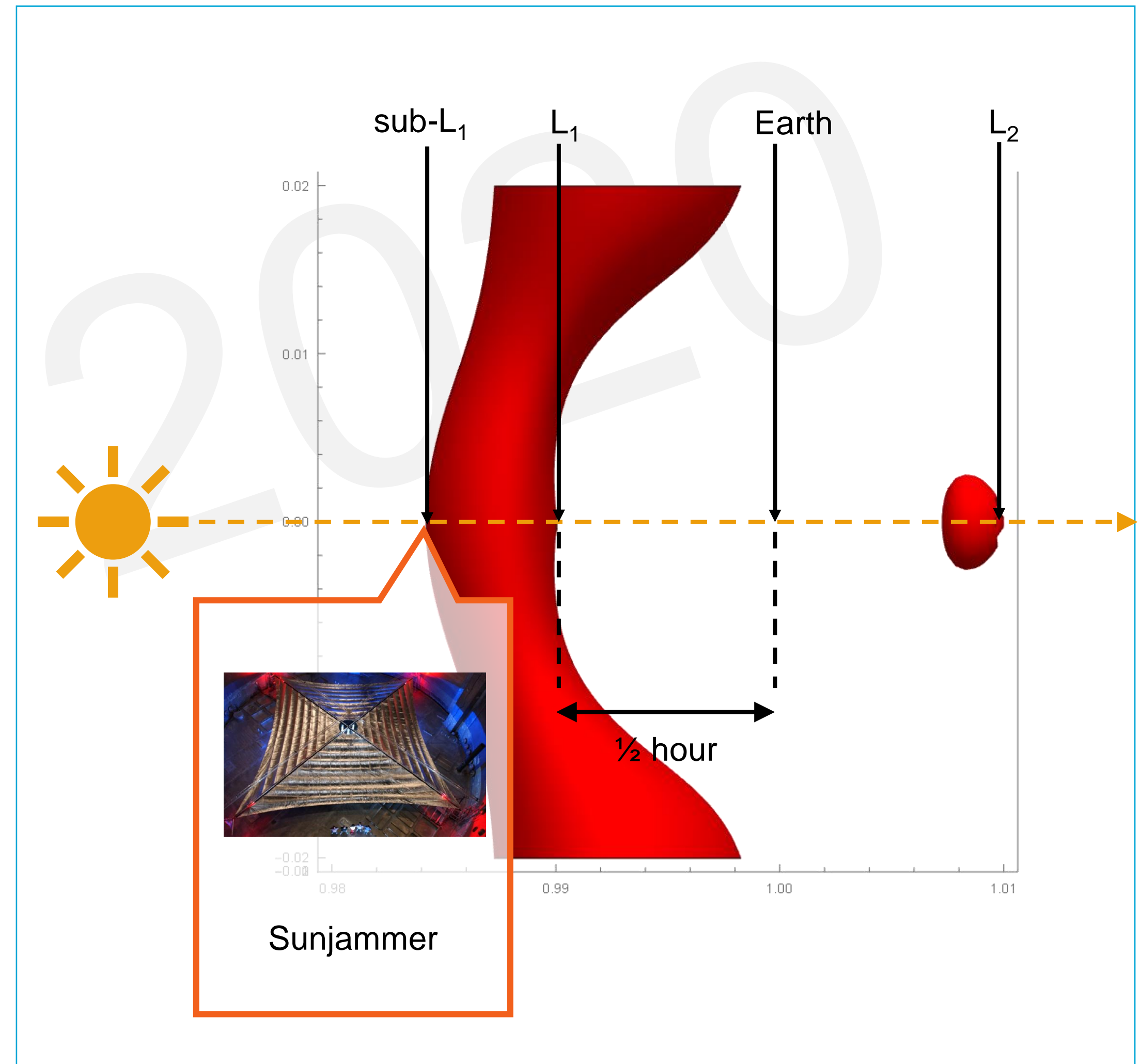
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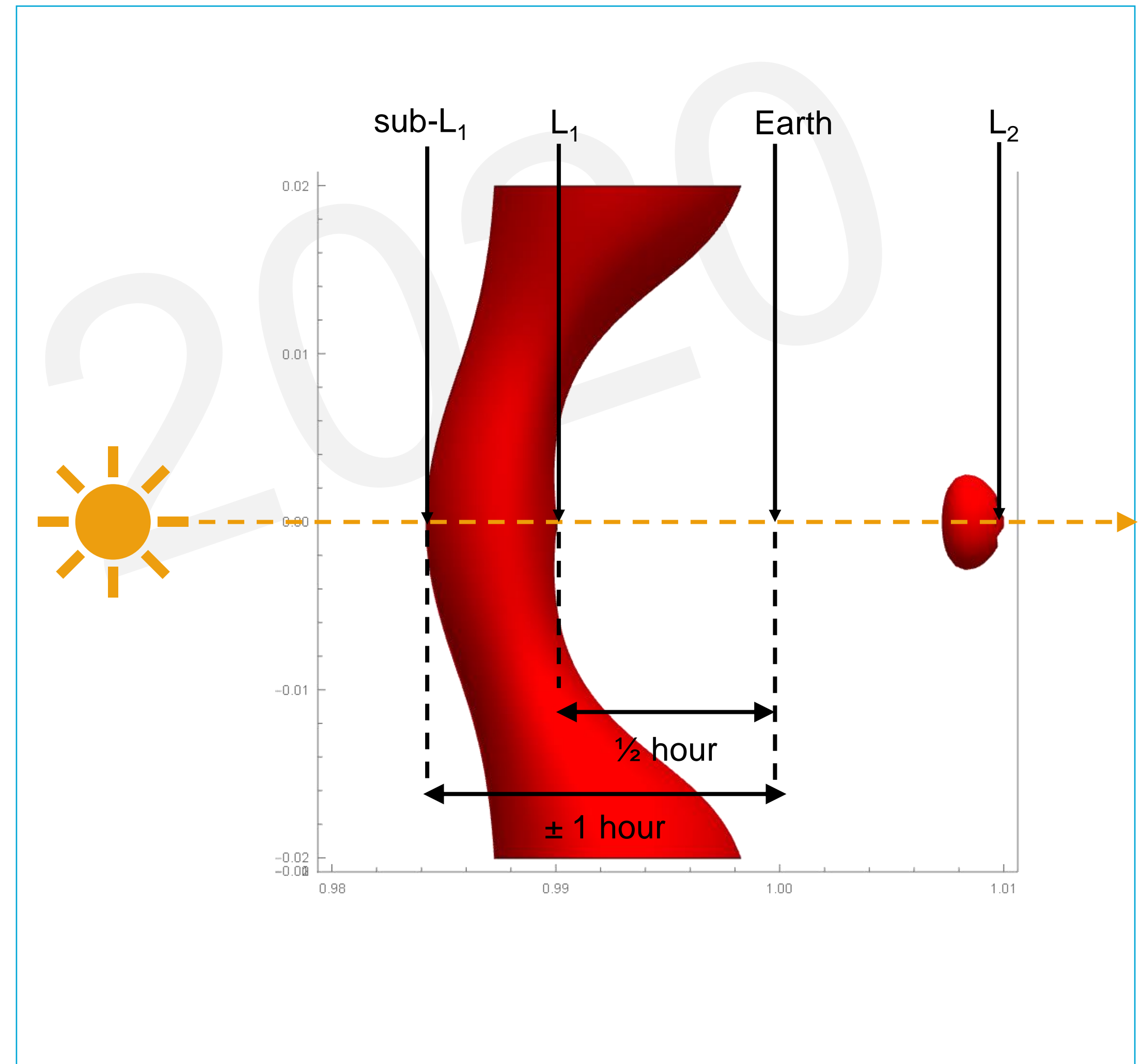
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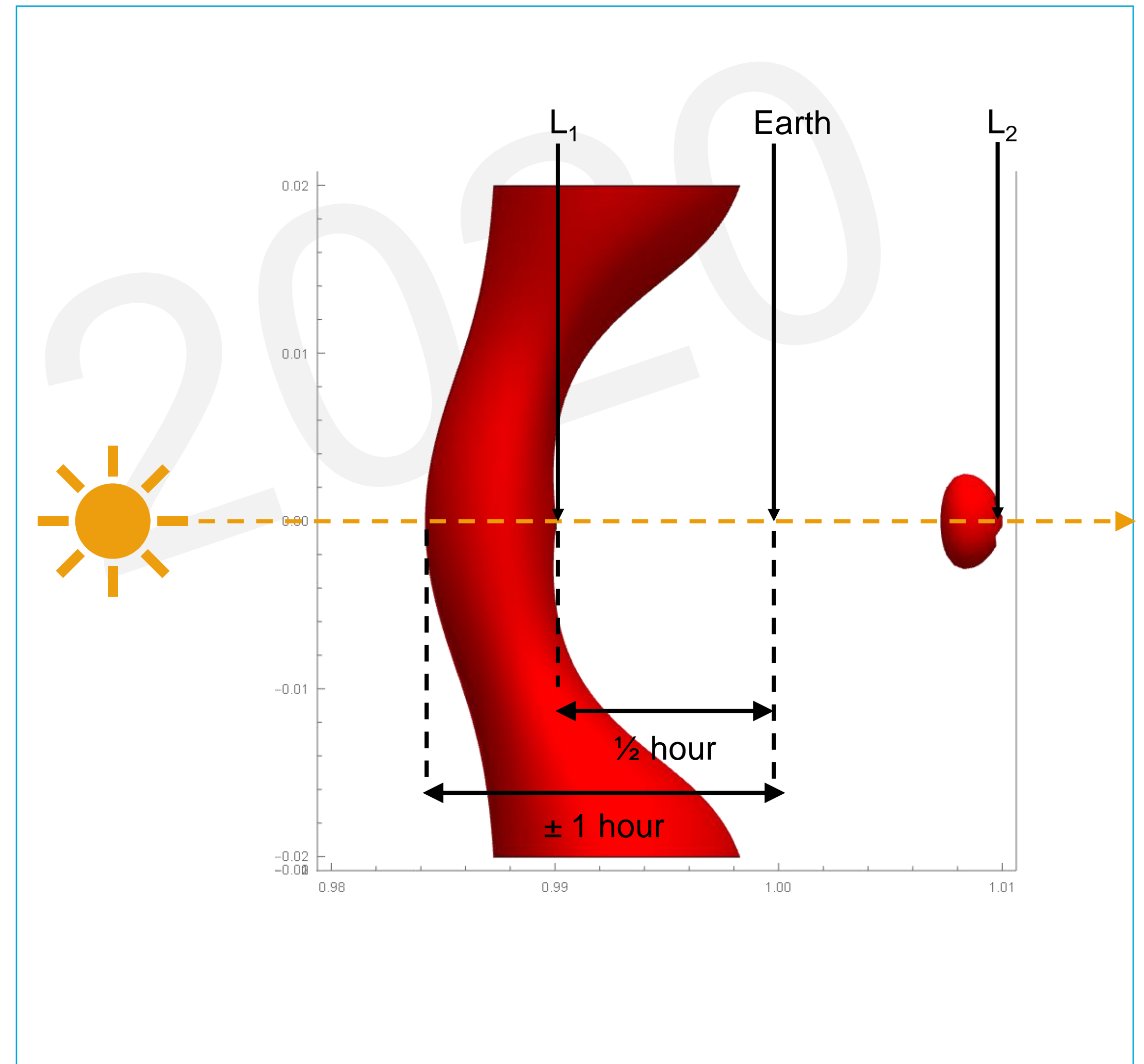
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- Satellites with similar objective are currently stationed at the L_1 point
- Sunjammer could have **doubled the warning time** for solar storms

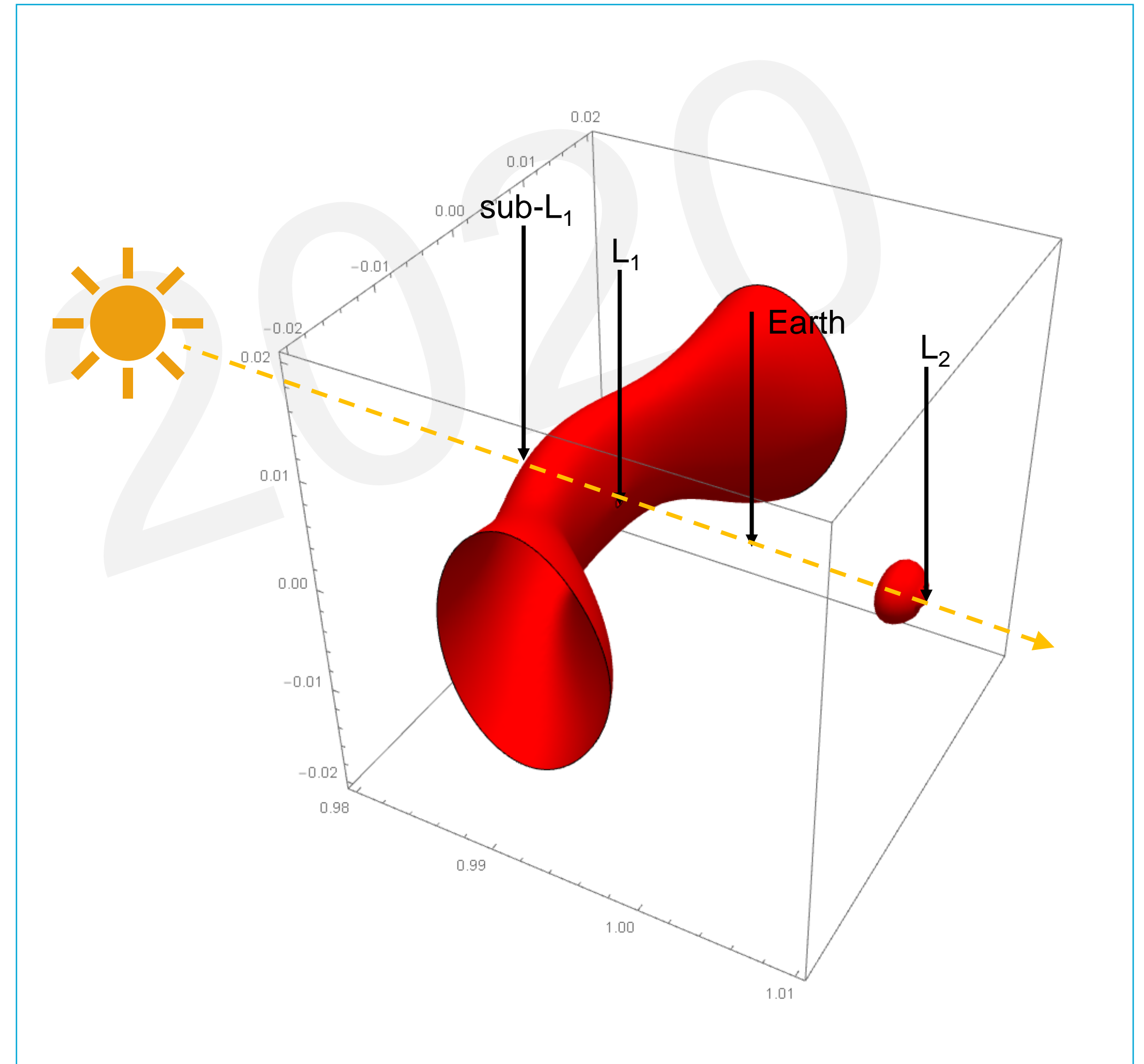


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- Satellites with similar objective are currently stationed at the L_1 point
- Sunjammer could have doubled the warning time for solar storms
- Due to budgetary reasons, Sunjammer was cancelled in 2015

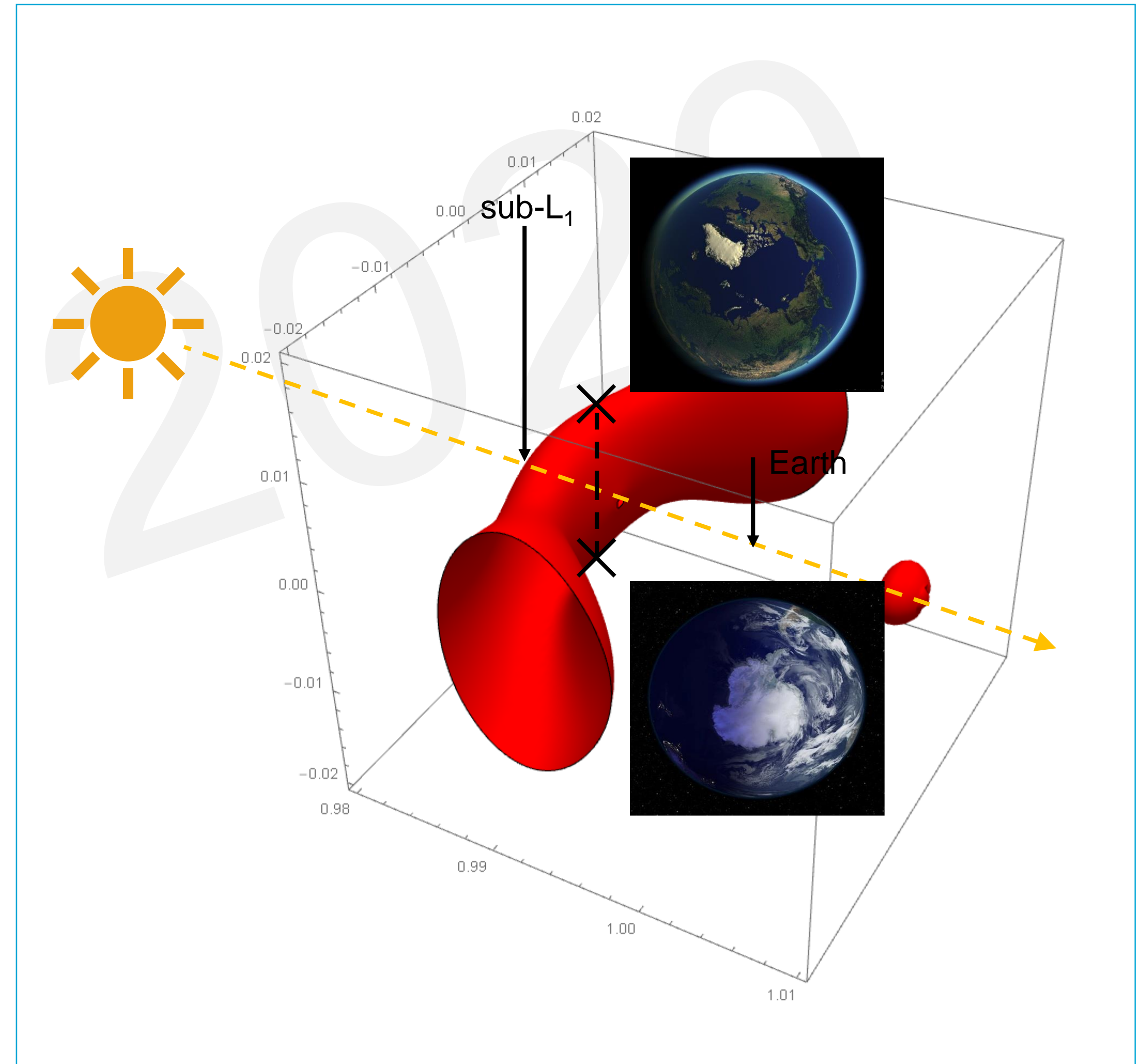


Artificial Lagrange points - utility



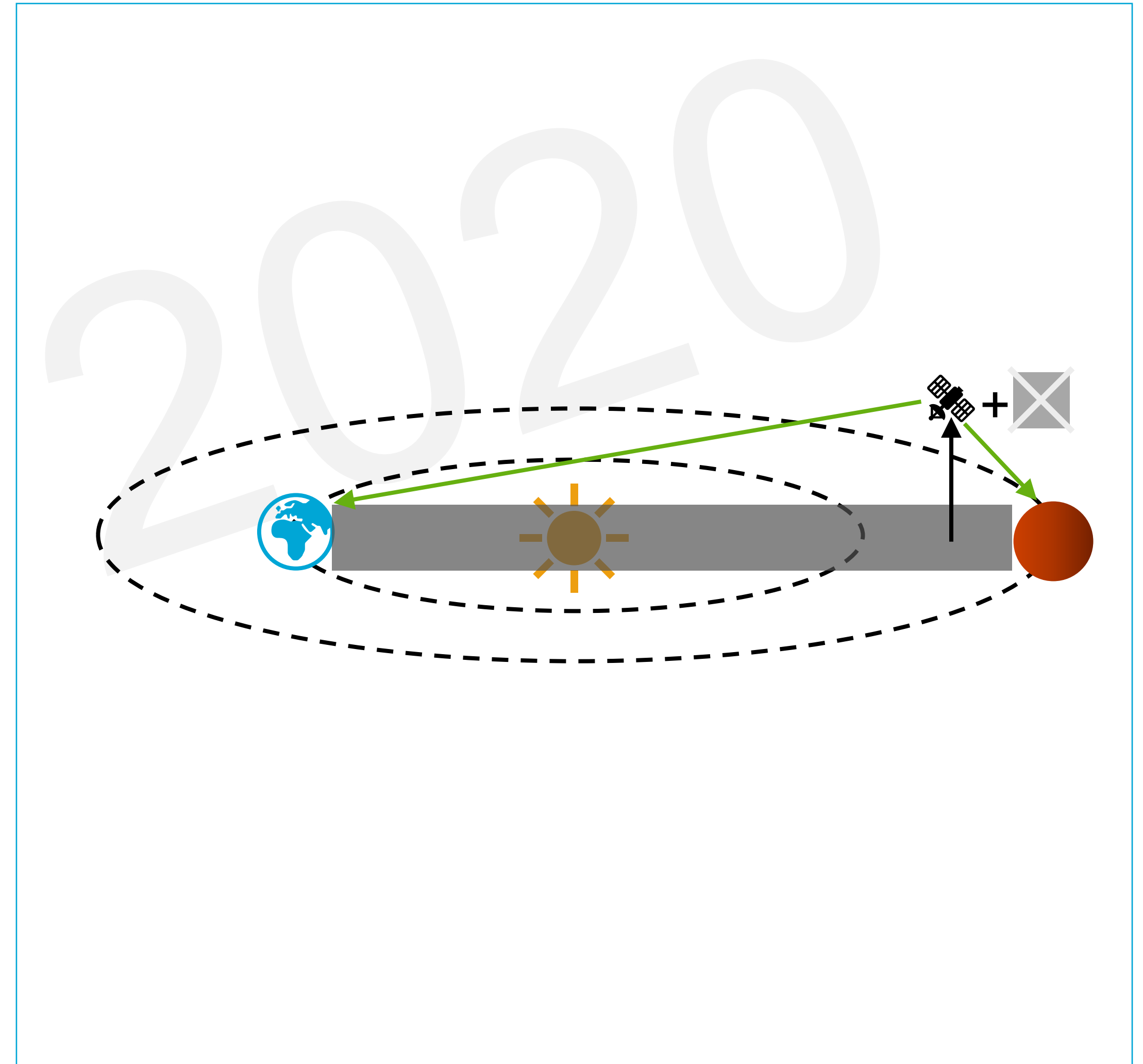
Artificial Lagrange points - utility

- High-latitude observation of the Earth



Artificial Lagrange points - utility

- High-latitude observation of the Earth
- Earth-Mars communication during conjunction



ST>A – Learning objectives today – revisited

- Summarize a short history on solar sailing
- Explain the basic workings, advantages and disadvantages of solar-sail propulsion
- Derive a solar-sail acceleration model
- Compute artificial Lagrange points in the solar-sail perturbed cr3bp
- **Identify mission applications of solar-sail propulsion in the cr3bp**



End of video