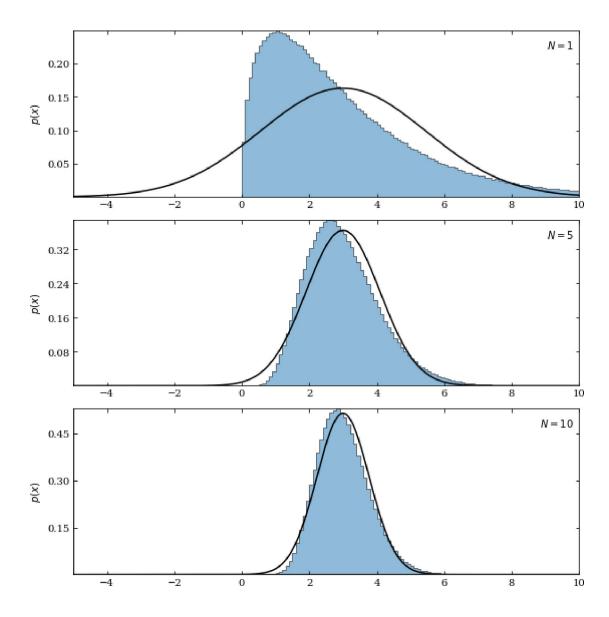
Data Science Analysis Assignment - 2

Q1. In the class, we demonstrated the Central Limit Theorem for a sample drawn from a uniform distribution. Reproduce a similar plot for a sample drawn the from chi-square distribution with degrees of freedom equal to 3, for samples drawn once, 5 times, and 10 times. Either plot all of these on one multipanel figure similar to AstroML figure 3.20.

```
In [67]:
          #importing required libraries
          import numpy as np
          from matplotlib import pyplot as plt
          from scipy.stats import norm
          from astroML.plotting import setup_text_plots
          setup text plots(fontsize = 10, usetex = False)
          # Generate the uniform samples
          N = [1, 5, 10]
          x = np.random.chisquare(3, (10, 1000000))
          # Plot the results
          fig = plt.figure(figsize = (8, 8))
          fig.subplots_adjust(hspace = 0.1)
          for i in range(len(N)):
              ax = fig.add_subplot(3, 1, i + 1)
              # take the mean of the first N[i] samples
              x_i = x[:N[i], :].mean(0)
              # histogram the data
              ax.hist(x i, bins = np.linspace(0, 10, 101),
                      histtype='stepfilled', alpha=0.5, density=True)
              # plotting the expected chi-square pdf
              mu = 3 #degrees of freedom
              sigma = np.sqrt(6)/np.sqrt(N[i])
              dist = norm(mu, sigma)
              x pdf = np.linspace(-5, 10, 1000)
              #plotting the pdf and setting the limits for the axes
              ax.plot(x_pdf, dist.pdf(x_pdf), '-k')
              ax.set_xlim(-5, 10)
              ax.set_ylim(0.001, None)
              ax.yaxis.set_major_locator(plt.MaxNLocator(5))
              ax.text(0.99, 0.95, r"$N = %i$" % N[i],
                      ha='right', va='top', transform=ax.transAxes)
              ax.set ylabel('$p(x)$')
          plt.tight_layout()
          plt.show()
```

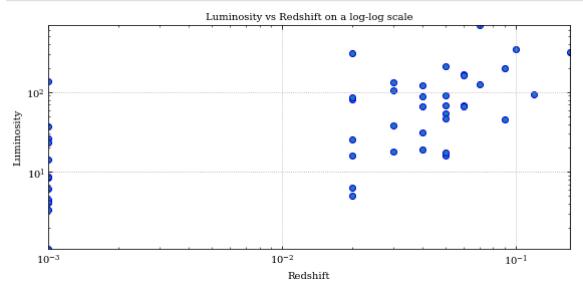


Q2. The luminosity and redshift of galaxy clusters from XMM-BCS survey (details available at arXiv:1512.01244) can be downloaded http://www.iith.ac.in/~shantanud/test.dat. Plot the luminosity as a function of redshift on a log-log scale. By eye, do you think the datasets are correlated? Calculate the Spearman, Pearson and Kendall-tau correlation coefficients and the p-value for the null hypothesis.

```
#scattering the dataset
ax.scatter(x,y)

#adding title and lables
plt.title('Luminosity vs Redshift on a log-log scale')
plt.xlabel('Redshift')
plt.ylabel('Luminosity')

plt.grid()
plt.show()
```



Pearson: 0.5144497852670243 Spearman: 0.6596325957535454 Kendall: 0.5029584682704178

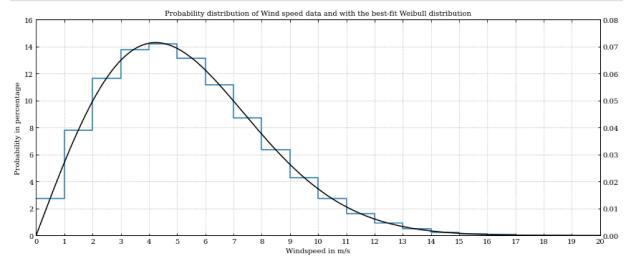
Correlation coefficients

P-values

Pearson: 0.00025464716576124137 Spearman: 6.166489759081011e-07 Kendall: 2.969686227473415e-06

Q3. Wind speed data from the Swiss Wind Power data website can be found at http://wind-data.ch/tools/weibull.php. Using the data provided on the website, plot the probability distribution and overlay the best-fit Weibull distribution (with the parameters shown on the website). (20 points) (Hint: A on the website is same as λ , which was used in class to parameterize the Weibull distribution.)

```
In [139... | #reading data from csv file
          data2 = pd.read_csv("D:\CLASSES\SEM 4\Data Science Analysis\A2\data2.csv",
                              skipinitialspace=True, usecols = [0,1])
          #making the plot more presentable
          fig, ax = plt.subplots(1, 1, figsize =(12, 5), tight_layout = True)
          plt.grid()
          #plotting the step histogram and setting limits
          ax.step(np.arange(0, 20), data2['Probability'], where = 'post')
          ax.set_xlim(0, 20)
          ax.set_ylim(0, 16)
          ax.set xticks(np.arange(0, 21))
          #adding labels
          ax.set ylabel('Probability in percentage')
          ax.set xlabel('Windspeed in m/s')
          #importing library required to plot Weibull dist.
          from scipy.stats import dweibull
          #making an axis on RHS for the dist.
          ax1 = ax.twinx()
          #making the Weibull dist.
          \#k=2, loc=0, lambda=6
          dist = dweibull(2, 0, 6)
          #generating x values for the pdf
          x_pdf = np.linspace(0, 20, 1000)
          #plotting
          ax1.plot(x_pdf, dist.pdf(x_pdf),'-k')
          ax1.set_ylim(0,0.08)
          plt.title('Probability distribution of Wind speed data and with the best-fit Weibull
          plt.show()
```



Q4. Generate two arrays of size 1000 drawn from a Gaussian distribution of mean of zero and standard deviation of one. Calculate Pearson correlation coefficient and its p-value using scipy module. Also check if the p- value agrees with that calculated using the Student-t distibution.

```
from scipy.stats import pearsonr

#creating gaussian dist with mean 0 and std 1
#mu, sigma = 0, 1
dist = norm(0, 1)

#creating two arrays of size 1000 drawn from the dist.
data1 = dist.rvs(1000)
data2 = dist.rvs(1000)

#data1 = np.random.default_rng().normal(mu, sigma, 1000)
#data2 = np.random.default_rng().normal(mu, sigma, 1000)

corr, p_v = pearsonr(data1, data2)

print('Pearson Correlation Coefficient is %s \np value is %s' %(corr, p_v))
```

Pearson Correlation Coefficient is 0.021066530903662433 p value is 0.5057800818556575

```
In [122...
```

```
from scipy import stats

#creating t distribution
dist2 = stats.t(998)

r = corr
t = r*np.sqrt(998/(1-r**2))
p = 2*(1-dist.cdf(t))

print(f"P value obtained from students-t distribution : {p}")
```

P value obtained from students-t distribution : 0.505626265053148