

EPOCHS V1 Catalogue Information

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1 Introduction

This document provides information on the available columns in the EPOCHS v1 catalogue release. The EPOCHS v1 catalogue consists of 1165 galaxy candidates at $z > 6.5$ taken from 214 arcmin² of deep JWST NIRCам observations.

The EPOCHS paper series is as follows:

1. **EPOCHS I: The Discovery and Properties of Distant Galaxies in the Epoch of Reionization at $6.5 < z < 20$ with Pearls, JADES GTO, and Public JWST data** *Conselice et al. 2024, ApJ submitted*
2. **EPOCHS II: The Ultraviolet Luminosity Function from $7.5 < z < 13.5$ using 180 square arcminutes of deep, blank-fields from the PEARLS Survey and Public JWST data** *Adams et al. 2024, ApJ*
3. **EPOCHS III: Unbiased UV continuum slopes at $6.5 < z < 13$ from combined PEARLS GTO and public JWST NIRCам imaging** *Austin et al. 2024, ApJ submitted*
4. **EPOCHS IV: SED Modelling Assumptions and their impact on the Stellar Mass Function at $6.5 \leq z \leq 13.5$ using PEARLS and public JWST observations** *Harvey et al. 2024, ApJ submitted*
5. **EPOCHS V: The dependence of galaxy formation on galaxy structure at $z < 7$ from JWST observations** *Conselice et al. 2024, MNRAS*
6. **EPOCHS VI: the size and shape evolution of galaxies since $z \sim 8$ with JWST Observations** *Ormerod et al. 2024, MNRAS*
7. **EPOCHS VII: Discovery of high-redshift ($6.5 < z < 12$) AGN candidates in JWST ERO and PEARLS data** *Juodžbalis et al. 2023, MNRAS*
8. **EPOCHS VIII: An Insight into MIRI-selected Galaxies in SMACS-0723 and the Benefits of Deep MIRI Photometry in Revealing AGN and the Dusty Universe** *Li et al. 2023, MNRAS*
9. **EPOCHS IX. When cosmic dawn breaks: Evidence for evolved stellar populations in $7 < z < 12$ galaxies from PEARLS GTO and public NIRCам imaging** *Trussler et al. 2024, MNRAS*
10. **EPOCHS X: Environmental effects on Galaxy Formation and Protocluster Galaxy candidates at $4.5 < z < 10$ from JWST observations** *Li et al. 2024, MNRAS submitted*
11. More to come!

1.1 Notes

BAND refers to any of the ACS_WFC or NIRCам bands listed in Table 1. If a band is listed as available in Table 1, but is not given for a galaxy in that field, then that galaxy fall within our image mask in that filter.

Table 1: Table of the available HST/ACS_WFC and JWST/NIRCam data used in each field in the EPOCHS v1 catalogue. A “*” means the imaging only partially covers the NIRCam wideband footprint.

| Field | HST/ACS_WFC | | JWST/NIRCam | | | | | | | | |
|---------------|-------------|-------|-------------|-------|-------|-------|-------|-------|-------|-------|-------|
| | F606W | F814W | F090W | F115W | F150W | F200W | F277W | F335M | F356W | F410M | F444W |
| NEP | Y | N | Y | Y | Y | Y | Y | N | Y | Y | Y |
| El Gordo | N | N | Y | Y | Y | Y | Y | N | Y | Y | Y |
| MACS-0416 | N | N | Y | Y | Y | Y | Y | N | Y | Y | Y |
| CLIO | N | N | Y | N | Y | Y | Y | N | Y | N | Y |
| CEERS | Y | Y | N | Y | Y | Y | Y | N | Y | Y | Y |
| SMACS0723 | N | N | Y | N | Y | Y | Y | N | Y | N | Y |
| GLASS | N | N | Y | Y | Y | Y | Y | N | Y | N | Y |
| NGDEEP | Y* | Y* | N | Y | Y | Y | Y | N | Y | N | Y |
| JADES Deep GS | Y | N | Y | Y | Y | Y | Y | Y | Y | Y | Y |

Table 2: Table of EPOCHS v1 catalog column names, units, descriptions and column shape. A “*” indicates that the column has been corrected for any flux associated with the galaxy which falls outside the extraction aperture.

| Column Name | Unit | Description |
|--|---------------------------|--|
| IDs, Positions, Fluxes and local depths | | |
| ID | | Unique catalogue ID, consisting of number and fieldname |
| ALPHA_J2000 | degree | Right ascension |
| DELTA_J2000 | degree | Declination |
| FIELDNAME | | Field/pointing in Table 1 the galaxy is in |
| FLUX_APER_ BAND | nJy | Aperture corrected flux in 0.16 arcsec radius apertures |
| FLUXERR_APER_ BAND | nJy | Local-depth derived flux error from NMAD of 200 nearby empty apertures |
| sigma_ BAND | | SNR of detection in 0.16 arcsec aperture |
| local_depth_ BAND | AB Mag | 5σ local depth from NMAD of flux in 200 nearby empty apertures |
| unmasked_ BAND | Boolean | Whether galaxy is masked in BAND |
| auto_corr_factor_ BAND | | Correction factor in BAND for flux outside 0.16 arcsec aperture |
| Photometric Redshifts and Selection | | |
| zbest | | Photometric redshift using EAZY-py |
| zbest_l1 | | -1σ photometric redshift uncertainty using EAZY-py |
| zbest_u1 | | $+1\sigma$ photometric redshift uncertainty using EAZY-py |
| chi2_best | | χ^2 of EAZY-py fit |
| PDF_integral_eazy | | $\int_{0.94 \times z_{\text{best}}}^{1.06 \times z_{\text{best}}} \text{PDF}(z) dz$ - Integral of EAZY-py posterior redshift PDF |
| zbest_lowz | | Photometric redshift using EAZY-py, with $z_{\text{max}} = 6$ |
| chi2_best_lowz | | χ^2 of EAZY-py fit, with $z_{\text{max}} = 6$ |
| UV Properties | | |
| M_UV* | AB Mag | Absolute UV mag in 100Å tophat at 1500Å rest-frame flux at redshift zbest |
| M_UV_u1 | AB Mag | |
| M_UV_l1 | AB Mag | |
| BETA_UV | | UV slope $f \propto \lambda^\beta$ |
| BETA_UV_l1 | | |
| BETA_UV_u1 | | |
| SFR_UV* | $M_\odot \text{ yr}^{-1}$ | |
| SFR_UV_l1 | $M_\odot \text{ yr}^{-1}$ | |
| SFR_UV_u1 | $M_\odot \text{ yr}^{-1}$ | |
| Sample identifiers | | |
| certain_by_eye | Boolean | Visual inspection of cutout and SED boolean |
| EPOCHS_II | Boolean | Used in EPOCHS II (UV LF) |
| EPOCHS_III | Boolean | Used in EPOCHS III (UV and SF properties) |
| EPOCHS_IV | Boolean | Used in EPOCHS IV (SMF) |

Table 3: Table of EPOCHS v1 catalog column names, units, descriptions and column shape, specifically for the stellar population parameters calculated using **Bagpipes**. **EXT** indicates that the column name appears multiple time with different extensions, and in this case “**EXT**” can take the value of **zfix** or **zgauss**, depending on whether the redshift is fixed to the **EAZY-py** maximum likelihood result given by “zbest”, or allowed to vary within a Gaussian centered on “zbest”. A “*” indicates that the column has been corrected for any flux associated with the galaxy which falls outside the extraction aperture - for masses this is done by correcting the mass by the ratio of MAG_AUTO to MAG_APER in the longest wavelength F444W band, where this exceeds unity. For star formation rates the band covering the rest-frame 1500Å wavelength is used instead.

| Column Name | Unit | Description |
|---|-------------------------------------|--|
| Fiducial Bagpipes Results (z_{fix} or z_{gauss}) | | |
| redshift_pipes_zgauss | | Fitted redshift (zgauss only) |
| redshift_pipes_l1_zgauss | | Lower uncertainty (50th - 16th percentile) |
| redshift_pipes_u1_zgauss | | Upper uncertainty (84th - 50th percentile) |
| stellar_mass_pipes_**EXT* | log ₁₀ (M _⊙) | Total surviving stellar mass |
| stellar_mass_pipes_l1_**EXT | log ₁₀ (M _⊙) | Lower uncertainty (50th - 16th percentile) |
| stellar_mass_pipes_u1_**EXT | log ₁₀ (M _⊙) | Upper uncertainty (84th - 50th percentile) |
| SFR_10Myr_pipes_**EXT* | M _⊙ yr ⁻¹ | Average total star formation rate over a 10 Myr timescale |
| SFR_10Myr_pipes_l1_**EXT | M _⊙ yr ⁻¹ | Lower uncertainty (50th - 16th percentile) |
| SFR_10Myr_pipes_u1_**EXT | M _⊙ yr ⁻¹ | Upper uncertainty (84th - 50th percentile) |
| SFR_100Myr_pipes_**EXT* | M _⊙ yr ⁻¹ | Average total star formation rate over a 100 Myr timescale |
| SFR_100Myr_pipes_l1_**EXT | M _⊙ yr ⁻¹ | Lower uncertainty (50th - 16th percentile) |
| SFR_100Myr_pipes_u1_**EXT | M _⊙ yr ⁻¹ | Upper uncertainty (84th - 50th percentile) |
| mass_weighted_age_pipes_**EXT | Myr | Mass-weighted age of galaxy |
| mass_weighted_age_pipes_l1_**EXT | Myr | Lower uncertainty (50th - 16th percentile) |
| mass_weighted_age_pipes_u1_**EXT | Myr | Upper uncertainty (84th - 50th percentile) |
| beta_pipes_**EXT | | UV β slope of best-fitting Bagpipes spectra in Calzetti filters |
| beta_pipes_l1_**EXT | | Lower uncertainty (50th - 16th percentile) |
| beta_pipes_u1_**EXT | | Upper uncertainty (84th - 50th percentile) |
| Z_star_pipes_**EXT | Z _⊙ | Stellar metallicity |
| Z_star_pipes_l1_**EXT | Z _⊙ | Lower uncertainty (50th - 16th percentile) |
| Z_star_pipes_u1_**EXT | Z _⊙ | Upper uncertainty (84th - 50th percentile) |
| A_V_pipes_**EXT | AB mag | Dust extinction in V band |
| A_V_pipes_l1_**EXT | AB mag | Lower uncertainty (50th - 16th percentile) |
| A_V_pipes_u1_**EXT | AB mag | Upper uncertainty (84th - 50th percentile) |
| U-V_pipes_**EXT | AB mag | U-V colour |
| U-V_pipes_l1_**EXT | AB mag | Lower uncertainty (50th - 16th percentile) |
| U-V_pipes_u1_**EXT | AB mag | Upper uncertainty (84th - 50th percentile) |
| M_UV_pipes_**EXT | AB mag | Absolute UV Magnitude |
| M_UV_pipes_l1_**EXT | AB mag | Lower uncertainty (50th - 16th percentile) |
| M_UV_pipes_u1_**EXT | AB mag | Upper uncertainty (84th - 50th percentile) |
| chisq_phot_pipes_**EXT | | χ ² of fit |