Classical Mechanics Tutorial II

Engineering Physics

Indian Institute of Information Technology, Allahabad

Constraints

Holonomic constraints – They can be expressed as an equation connecting the cooridnates of the particles.

Non-holonomic constraints - Constraints which are not expressible in the form of an equation.

Before moving on, recall that constraints can also be rheonomous (explicit time dependence) or scleronomous (not explicitly dependent of on time)

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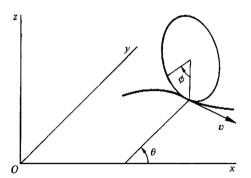
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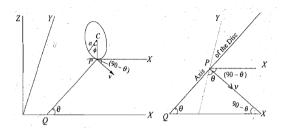
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Another example of a non-holonomic constraint

Consider a disc of radius R rolling (without slipping) on a horizontal plane x-y plane constrained to move so that the plane of the disc is always vertical. To describe the motion of the disc, we use the following coordinates: x,y coordinates of the center of the disc, the angle θ bewteen the axis of the disc and the x axis, and the angle of rotation ϕ about the axis of the disc.



Another example of a non-holonomic constraint



Since the disc remains vertical, the axis of rotation is perpendicular to the z axis. This tells us that the velocity of the center of the disc has a magnitude $|\mathbf{v}|=R\dot{\phi}$ and its direction is perpendicular to the axis of rotation $\Longrightarrow \dot{\mathbf{x}}=v\sin\theta$ and $\dot{\mathbf{y}}=-v\cos\theta$ which implies

$$dx - R\sin\theta d\phi = 0 \text{ and } dy + R\cos\theta d\phi = 0 \tag{1}$$

These constraints are not of the form $f(x, y, \theta, \phi) = 0$ and are hence non-holonomic. Actually neither of the equations can be integrated without solving the problem first, that is, we cannot first find the integrating factor $f(x, y, \theta, \phi) = 0$ that will covert them into exact differentials.

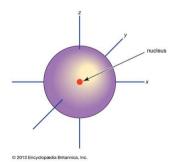
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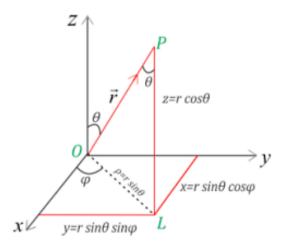
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- We will be using spherical polar coordinates (r, θ, ϕ) to describe this system.
- ▶ The constraint is the fixed radius r = I, where I is an arbitrary constant indicating the fixed length.



Spherical Polar Coordinates



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$$x = I \sin \theta \cos \phi \tag{2}$$

$$y = I\sin\theta\sin\phi\tag{3}$$

$$z = I\cos\theta\tag{4}$$

where r = I from the constraint equation. The expression for kinetic energy T is simply given by

$$T = \frac{1}{2}m(\dot{x}^2 + \dot{y}^2 + \dot{z}^2) \tag{5}$$

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$$\dot{x} = I(\cos\theta\cos\phi\dot{\theta} + \sin\theta(-\sin\phi)\dot{\phi}) \tag{6}$$

$$\dot{y} = I(\cos\theta\sin\phi\dot{\theta} + \sin\theta(+\cos\phi)\dot{\phi}) \text{ and } \dot{z} = -I\sin\theta\dot{\theta}$$

So using these equations

$$\dot{x} = I(\cos\theta\cos\phi\dot{\theta} + \sin\theta(-\sin\phi)\dot{\phi}) \tag{8}$$

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$$\dot{z} = -I\sin\theta\dot{\theta}\tag{10}$$

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$$T = \frac{1}{2}ml^2(\dot{\theta}^2 + \sin\theta^2\dot{\phi}^2) \tag{12}$$

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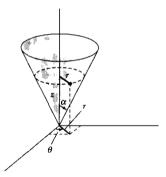
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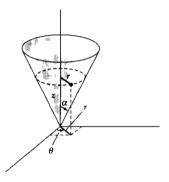
$$L = \frac{1}{2}ml^{2}(\dot{\theta}^{2} + \sin\theta^{2}\dot{\phi}^{2}) - V(\theta, \phi)$$
 (13)

Let us consider a different problem: A particle of mass m is constrained to move on the inside surface of a smooth cone of half angle α . The particle is subject to a gravitational force. First determine a set of generalized coordinates and the constrains, and then find Lagrange's equations of motions.

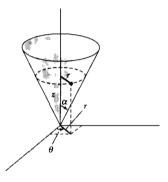
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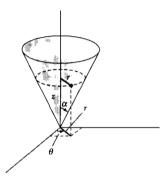


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- ▶ The constraint is the fixed radius $z = r \cot \alpha$, where we can see that r is the "height" and z is the "base" of the triangle formed by cutting the cone vertically. So $r/z = \tan \alpha$

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$$V = mgz = mgr \cot \alpha \tag{18}$$

So we now have,

$$T = \frac{1}{2}m(\dot{r}^2\csc^2\alpha + r^2\dot{\theta}^2)$$

and,

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So the Lagrangian can be written as L = T - U

$$L = \frac{1}{2}m(\dot{r}^2\csc^2\alpha + r^2\dot{\theta}^2) - mgr\cot\alpha$$
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and Lagrange's equations of motions are given by

$$\frac{d}{dt}\left(\frac{\partial L}{\partial \dot{q}}\right) - \left(\frac{\partial L}{\partial q}\right) = 0 \tag{20}$$

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 θ is a cyclic coordinate in this Lagrangian.

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But $mr^2\dot{\theta}=mr^2\omega$ is just the angular momentum about z-axis. This equation simply gives us the conservation of angular momentum. Similarly, we can calculate Lagrange's equation for the r coordinate

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$$\ddot{r} - r\dot{\theta}^2 \sin^2 \alpha + g \sin \alpha \cos \alpha = 0 \tag{24}$$

Lagrange's equation of motion for dissipative systems

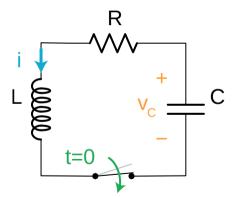
The Lagrange equation for a system with dissipation is given by

$$\frac{d}{dt}\left(\frac{\partial L}{\partial \dot{q}_j}\right) - \frac{\partial L}{\partial q_j} + \frac{\partial \mathcal{F}}{\partial \dot{q}_j} = 0. \tag{25}$$

Here \mathcal{F} is a scalar function known as Rayleigh's dissipation function and must be specified along with L to obtain the equations of motion.

One of the advantages of the Lagrangian formulation is that it can be easily extended to systems that are not studied in classical mechanics such as electrical circuits!

We consider a physical system which has a "key" in series with an inductance L, A capacitance C and a resistance R. We first charge the capacitor and then close the key The capacitor will now begin to discharge. What is equation that captures this situation?



We choose the electric charge q as our dynamical variable (generalised coordinate).

The inductor will give rise to a kinetic energy term since the energy stored in an inductor contains current I, the time derivative of q. The inductance L is the electrical analogue of mass. So $=\frac{1}{2}L\dot{q}^2$.

The capacitor gives us the potential energy. The capacitance term 1/C is analogous to the spring constant k. So $U=\frac{1}{2}\frac{q^2}{C}$

If we now introduce a dissipation function

$$\mathcal{F} = \frac{1}{2}R\dot{q}_j^2\tag{26}$$

We can use

$$\frac{d}{dt}\left(\frac{\partial L}{\partial \dot{q}_j}\right) - \frac{\partial L}{\partial q_j} + \frac{\partial \mathcal{F}}{\partial \dot{q}_j} = 0. \tag{27}$$

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to get

$$L\frac{d^2q}{dt^2} + R\frac{dq}{dt} + \frac{q}{C} = 0 (29)$$

This is exactly what we would get from Kirchhoff's law!

Taylor Series

Taylor's theorem provides a way of expressing a function as a power series in x, known as a Taylor series.

It can be applied only to those functions that are continuous and differentiable within the x-range of interest.

To express f(x) as a power series in x-a about the point x=a. We shall assume that, in a given x-range, f(x) is a continuous, single-valued function of x having continuous derivatives. Then

$$f(x) = f(a) + (x - a)f'(a) + \frac{(x - a)^2}{2!}f''(a) + \cdots + \frac{(x - a)^n}{n!}f^{(n)}(a) + \cdots$$

alternatively setting x = a + h,

$$f(a+h) = f(a) + hf'(a) + \frac{h^2}{2!}f''(a) + \cdots + \frac{h^n}{n!}f^{(n)}(a) + \cdots$$

Example:
$$\sin(x) = x - \frac{x^3}{3!} + \frac{x^5}{5!} - \frac{x^7}{7!} + \dots = \sum_{n=0}^{\infty} (-1)^n \frac{x^{(2n+1)}}{(2n+1)!}$$
.

Taylor Series: More than one variable: The Taylor series work the same way for functions of two variables. (There are just more of each derivative!) For an analytic function f(x, y, z),

$$f(x,y) \approx f(a,b) + f_x(a,b)(x-a) + f_y(a,b)(y-b)$$

$$\frac{f_{xx}(a,b)}{2} + (x-a)^2 + f_{xy}(a,b,z)(x-a)(y-b) + \frac{f_{yy}(a,b)}{2} \quad (y-b)^2 + \cdots \quad (30)$$

Or, with $x = a + \delta x$ and $y = b + \delta y$

$$f(a + \delta x, b + \delta y) \approx f(a, b) + f_{x}(a, b)\delta x + f_{y}(a, b)\delta y + \frac{f_{xx}(a, b)}{2} (\delta x)^{2} + f_{xy}(a, b)\delta x\delta y + \frac{f_{yy}(a, b)}{2} (\delta y)^{2} + \cdots$$
(31)