

Theory and Phenomenology of Superconductivity Homework 1

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Problem 1. • Consider a system consisting of N spin- $\frac{1}{2}$ particles, each of which can be in one of two quantum states, namely \uparrow and \downarrow . In presence of a magnetic field B , the energy of a spin in a \uparrow / \downarrow state is $\epsilon = \pm \mu_B B / 2$ where μ_B is the magnetic moment. Show that the partition function is

$$Z = 2^N \cosh^N \left(\frac{\mu_B B}{2k_B T} \right),$$

with $1/\beta = k_B T$ in the canonical ensemble. Find the average energy E and entropy S . Compute both quantities at zero temperature and $T \rightarrow \infty$.

Proof. The partition function of a single particle is

$$\begin{aligned} Z_1 &= \sum_{s \in \{\uparrow, \downarrow\}} e^{-\beta H(s)} \\ &= e^{-\beta \epsilon_{\uparrow}} + e^{-\beta \epsilon_{\downarrow}} \\ &= e^{\frac{\beta \mu_B B}{2}} + e^{-\frac{\beta \mu_B B}{2}} \\ &= 2 \cosh \frac{\mu_B B}{2k_B T}. \end{aligned}$$

The partition function of N particles is the product of their partition functions; in this case, it is simply

$$Z = Z_1^N = 2^N \cosh^N \left(\frac{\mu_B B}{2k_B T} \right).$$

The average energy is given by

$$\begin{aligned} \langle E \rangle &= - \frac{\partial \ln Z}{\partial \beta} \\ &= - \frac{\partial}{\partial \beta} N \ln \left[2 \cosh \left(\frac{\beta \mu_B B}{2} \right) \right] \\ &= -N \frac{\partial}{\partial \beta} \ln \cosh \frac{\beta \mu_B B}{2} \end{aligned}$$

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$$\begin{aligned}
&= -\frac{N}{\cosh \frac{\beta\mu_B B}{2}} \left(\sinh \frac{\beta\mu_B B}{2} \right) \left(\frac{\mu_B B}{2} \right) \\
&= -\frac{N\mu_B B}{2} \tanh \frac{\beta\mu_B B}{2} \\
&= -\frac{N\mu_B B}{2} \tanh \left(\frac{\mu_B B}{2k_B T} \right)
\end{aligned}$$

The Helmholtz free energy is given by

$$F(T, V, N) = -k_B T \ln Z.$$

The entropy can be computed from the Helmholtz free energy as

$$S = -\frac{\partial F}{\partial T}.$$

This leads to the expression

$$\begin{aligned}
\frac{\partial \ln Z}{\partial T} &= N \frac{\partial}{\partial T} \left(\ln \left[2 \cosh \left(\frac{\mu_B B}{2k_B T} \right) \right] \right) \\
&= N \frac{\partial}{\partial T} \left(\ln \cosh \left(\frac{\mu_B B}{2k_B T} \right) \right) \\
&= N \frac{1}{\cosh \frac{\mu_B B}{2k_B T}} \left(\sinh \frac{\mu_B B}{2k_B T} \right) \left(-\frac{\mu_B B}{2k_B T^2} \right) \\
&= -N \frac{\mu_B B}{2k_B T^2} \tanh \frac{\mu_B B}{2k_B T}
\end{aligned}$$

Then, we expand

$$\begin{aligned}
S &= -\frac{\partial F}{\partial T} \\
&= k_B \ln Z + k_B T \frac{\partial}{\partial T} \ln Z \\
&= k_B N \ln \left[2 \cosh \left(\frac{\mu_B B}{2k_B T} \right) \right] - \frac{N\mu_B B}{2T} \tanh \frac{\mu_B B}{2k_B T}
\end{aligned}$$

We note that

$$\lim_{x \rightarrow \infty} \tanh(x) = 1, \quad \lim_{x \rightarrow -\infty} \tanh(x) = -1, \quad \tanh(0) = 0.$$

Thus, the average energy $\langle E \rangle$ is, in the limits,

$$\begin{aligned}
\langle E \rangle &\xrightarrow{T \rightarrow 0} -\frac{N\mu_B B}{2}, \\
\langle E \rangle &\xrightarrow{T \rightarrow \infty} 0.
\end{aligned}$$

In the limit of $T \rightarrow \infty$, the second term in the entropy vanishes due to the $\frac{1}{T}$ factor. Since $\cosh(0) = 1$, we are left with

$$S \xrightarrow{T \rightarrow \infty} = k_B N \ln 2.$$

The limit $T \rightarrow 0$ is the most subtle. We massage the entropy into a desired form by

$$\begin{aligned} S &= k_B N \left\{ \ln \left[2 \cosh \left(\frac{\mu_B B}{2k_B T} \right) \right] - \frac{\mu_B B}{2k_B T} \tanh \frac{\mu_B B}{2k_B T} \right\} \\ &= k_B N \left\{ \ln \left[2 \cosh \left(\frac{\mu_B B}{2k_B T} \right) \right] - \ln \exp \left[\frac{\mu_B B}{2k_B T} \tanh \frac{\mu_B B}{2k_B T} \right] \right\} \\ &= k_B N \ln \frac{e^{\frac{\mu_B B}{2k_B T}} + e^{-\frac{\mu_B B}{2k_B T}}}{\left(e^{\frac{\mu_B B}{2k_B T}} \right)^{\tanh \frac{\mu_B B}{2k_B T}}} \\ &= k_B N \ln \frac{1 + e^{-\frac{\mu_B B}{k_B T}}}{\left(e^{\frac{\mu_B B}{2k_B T}} \right)^{\tanh \left(\frac{\mu_B B}{2k_B T} \right) - 1}} \end{aligned}$$

Now all we need to do is to simplify

$$\begin{aligned} \exp \left(\frac{\mu_B B}{2k_B T} \left(\tanh \left(\frac{\mu_B B}{2k_B T} \right) - 1 \right) \right) &= \exp \left(\frac{\mu_B B}{2k_B T} \left(\frac{e^{\frac{\mu_B B}{k_B T}} - 1}{e^{\frac{\mu_B B}{k_B T}} + 1} - 1 \right) \right) \\ &= \exp \left(\frac{\mu_B B}{2k_B T} \frac{-2}{e^{\frac{\mu_B B}{k_B T}} + 1} \right) \\ &\xrightarrow{T \rightarrow 0} 1 \end{aligned}$$

and hence

$$S \xrightarrow{T \rightarrow 0} k_B N \ln 1 = 0. \quad \square$$

Problem 2. • Compute the partition function of a quantum harmonic oscillator at frequency ω in the canonical ensemble. *Hint:* The energy levels are given by

$$E_n = \hbar \omega \left(n + \frac{1}{2} \right),$$

with $n \in \mathbb{Z}$.

- A simple model of a solid can be made considering N atoms that vibrate all of them at the same frequency ω . Consider these vibrations as a harmonic oscillator. Show that at high temperatures, $k_B T \gg \hbar \omega$, one has a heat capacity

$$C_V = Nk_B.$$

- Derive the limit also for low temperatures.

Proof. The partition function is given by

$$\begin{aligned}
 Z &= \sum_n e^{-\beta E_n} \\
 &= \sum_{n=0}^{\infty} e^{-\beta \hbar \omega (n + \frac{1}{2})} \\
 &= e^{-\frac{\beta \hbar \omega}{2}} \sum_{n=0}^{\infty} e^{-\beta \hbar \omega n} \\
 &= e^{-\frac{\beta \hbar \omega}{2}} \frac{e^{\beta \hbar \omega}}{e^{\beta \hbar \omega} - 1} \\
 &= \frac{1}{e^{\frac{\beta \hbar \omega}{2}} + e^{-\frac{\beta \hbar \omega}{2}}} \\
 &= \frac{1}{2 \sinh\left(\frac{\beta \hbar \omega}{2}\right)}
 \end{aligned}$$

We then determine the average energy of a single particle as in the previous problem

$$\begin{aligned}
 U := \langle E \rangle &= -\frac{\partial \ln Z}{\partial \beta} \\
 &= \frac{\partial}{\partial \beta} \ln \left[2 \sinh\left(\frac{\beta \hbar \omega}{2}\right) \right] \\
 &= \frac{\partial}{\partial \beta} \ln \left[\sinh\left(\frac{\beta \hbar \omega}{2}\right) \right] \\
 &= \frac{1}{\sinh\left(\frac{\beta \hbar \omega}{2}\right)} \left[\cosh\left(\frac{\beta \hbar \omega}{2}\right) \right] \frac{\hbar \omega}{2} \\
 &= \frac{\hbar \omega}{2} \coth\left(\frac{\beta \hbar \omega}{2}\right)
 \end{aligned}$$

At high temperatures, β is low and we can make use of the well known expansion in small β

$$U = \frac{\hbar \omega}{2} \frac{2}{\beta \hbar \omega} + O(\beta) = k_B T + O\left(\frac{1}{T}\right).$$

By inspection, we see that the heat capacity $\frac{\partial U}{\partial T} = k_B$, and for N noninteracting harmonic oscillators, we simply multiply by N to get

$$C_V = N k_B.$$

In the low temperature limit, β is very high. We approximate

$$\cosh(x) \approx \frac{e^x}{2}$$

$$\sinh(x) \approx \frac{e^x}{2}$$

to get

$$U \approx \frac{\hbar\omega}{2}.$$

Since U does not depend on T in this limit, the heat capacity $\frac{\partial U}{\partial T}$ vanishes - without putting in any heat, it is possible to change the temperature. \square

Problem 3. • Consider the Gibbs entropy for a probability distribution $p(n)$,

$$S = -k_B \sum_n p(n) \ln p(n).$$

- Through the use of a Lagrange multiplier, show that when restricted to states of fixed energy E , the entropy is maximized by the microcanonical ensemble, in which all such states are equally likely. Further show that in this case, the Gibbs entropy coincides with the Boltzmann entropy. *Hint:* Recall that probabilities are positive and constrained to sum up to 1.
- Show that at fixed average energy, i.e.: $\langle E \rangle = \sum_n p(n) E_n$, the entropy is maximized by the canonical ensemble. Moreover, show that the Lagrange multiplier imposing the constraint is proportional to the inverse of temperature, β . Check that maximizing the entropy is equivalent to minimizing the free energy.

Proof. (a) The constraint equation is given by

$$\sum_n p(n) = 1.$$

Hence, the (constrained) maximization of the entropy can be replaced by an unconstrained maximization over the functional

$$S' = -k_B \sum_n p(n) \ln p(n) + \lambda \left(\sum_n p(n) - 1 \right).$$

We extremize this by taking the partial derivatives. The partial derivative with respect to an occupation $p(k)$ is given by

$$\frac{\partial S'}{\partial p(k)} = -k_B (\ln p(k) + 1) + \lambda = 0$$

and the partial derivative with respect to λ returns the constraint equation

$$\sum_n p(n) = 1$$

as expected. Rearranging the partial derivative with respect to $p(k)$ yields

$$\ln p(k) = \frac{\lambda}{k_B} - 1.$$

Since this must be true for all k , it tells us that $\ln p(k)$ is the same for all k . Since \ln is monotonic, then $p(k)$ must be the same for all k . In this case, the normalisation condition yields, for a total of N microstates,

$$p(k) = \frac{1}{N} \forall k.$$

The Gibbs entropy is given by

$$\begin{aligned} S &= -k_B \sum_{n=1}^N \frac{1}{N} \ln \frac{1}{N} \\ &= k_B \ln N \end{aligned}$$

This is, by definition, the Boltzmann entropy.

(b) We follow the same procedure, defining instead the functional

$$S' = -k_B \sum_n p(n) \ln p(n) + \lambda \left(\sum_n p(n) E_n - U \right) + \eta \left(\sum_n p(n) - 1 \right)$$

where we denote the average energy by $\langle E \rangle =: U$. Then, the derivatives yield

$$\begin{aligned} \frac{\partial S'}{\partial p(k)} &= -k_B (\ln p(k) + 1) + \lambda E_k + \eta = 0 \\ \frac{\partial S'}{\partial \lambda} &= \sum_n p(n) E_n - U = 0 \\ \frac{\partial S'}{\partial \eta} &= \sum_n p(n) - 1 = 0 \end{aligned}$$

We proceed in a manner analogous to the previous part: By solving the equation for $p(k)$. We have

$$\ln p(k) + 1 = \frac{1}{k_B} (\eta + \lambda E_k)$$

and

$$p(k) = e^{-1} e^{\frac{1}{k_B} (\eta + \lambda E_k)}$$

$$= e^{-1} e^{\eta/k_B} e^{\lambda E_k}$$

From this, we see by inspection that λ is proportional to the inverse temperature β .

(c) The Helmholtz free energy is given by

$$F = U - TS.$$

Since U and T are fixed in the canonical ensemble, minimizing S is equivalent to maximizing F . □