

Theory and Phenomenology of Superconductivity Homework 1

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Problem 1. Bose-Einstein condensation was obtained experimentally for the first time in 1995 by confining particles in a harmonic potential. Suppose you have particles in a potential $V(\mathbf{r}) = \frac{1}{2}m\omega^2 r^2$. Then, the particle states are characterized by three quantum numbers, n_x, n_y, n_z , all integers and greater or equal to zero. The energy of each state is given by

$$\epsilon_{n_x, n_y, n_z} = \hbar\omega(n_x + n_y + n_z).$$

Notice that the zero energy level has been redefined so that the lowest energy state (the ground state) has energy 0.

- Show that the grand canonical partition function can be written as

$$\ln Z_G = - \sum_{n_x, n_y, n_z} \ln \left(1 - ze^{-\beta\hbar\omega(n_x + n_y + n_z)} \right),$$

where $z = e^{\mu\beta}$ is the fugacity.

- Introduce $n = n_x + n_y + n_z$ and rewrite the sum over states as the sum over n with multiplicity $\Omega(n)$. Compute the number of states of a particle $\Omega(n)$ at a given energy $E_n = \hbar\omega n$ (notice that it is equivalent as, for instance, to distribute n equal books in 3 indistinguishable boxes).
- Which condition must be satisfied in order to approximate the sum by an integral? Approximate the sum in the grand canonical partition function by an integration assuming that the dominant contributions are from terms with $n \gg 1$. $\ln Z_G$ should be written in terms of $\text{Li}_s(z)$ defined as

$$\text{Li}_s(z) = \frac{1}{\Gamma(s)} \int_0^\infty \frac{x^{s-1}}{e^x/z - 1} dx,$$

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with Γ is the Gamma function. $\text{Li}_s(z)$ is the polylogarithm of order s . It will become useful that $\text{Li}_s(1) = \zeta(s)$, where ζ is the Riemann Zeta function and $z\text{Li}'_s(z) = \text{Li}_{s-1}(z)$.

- Find the expected number of particles $N = \langle N \rangle$ expressed in terms of the fugacity. (Notice that the condensate is a particular limit when $z \rightarrow 1$.) Identify the occupation of the ground state and the occupation of excited states.
- Introduce $0 < f \leq 1$ to represent how many particles are in the lowest energy state, such that $z = 1 - 1/fN$. Discuss the thermodynamic limit and find the critical temperature T_c . The critical temperature is where the ground state starts to become occupied when lowering the temperature.
- Calculate the internal energy and specific heat of the gas of temperature below the critical temperature.

Proof. (a) The grand canonical partition function is given by

$$\begin{aligned}
 Z &= \sum e^{-\beta(H-\mu N)} \\
 &= \sum_{n_x, n_y, n_z=0}^{\infty} \sum_{N(n_x, n_y, n_z)=0}^{\infty} e^{-\beta N(n_x, n_y, n_z)(\hbar\omega - \mu)(n_x + n_y + n_z)} \\
 &= \sum_{n_x, n_y, n_z=0}^{\infty} \sum_{N(n_x, n_y, n_z)=0}^{\infty} (e^{-\beta(\hbar\omega - \mu)(n_x + n_y + n_z)})^{N(n_x, n_y, n_z)} \\
 &= \sum_{n_x, n_y, n_z=0}^{\infty} \frac{1}{1 - e^{-\beta(\hbar\omega - \mu)(n_x + n_y + n_z)}}
 \end{aligned}$$

where $N(n_x, n_y, n_z)$ is the number of particles in the state (n_x, n_y, n_z) . Notably, n_x, n_y and n_z represent just energy states, rather than the number of bosons in a state[1]. Taking the logarithm yields

$$\ln Z = - \sum_{n_x, n_y, n_z=0}^{\infty} \ln \left[1 - ze^{-\beta\hbar\omega(n_x + n_y + n_z)} \right].$$

(b) Clearly, the partition function is given by

$$\ln Z_G = \sum_{n=0}^{\infty} \sum_{n_x + n_y + n_z = n} \ln \left(1 - ze^{-\beta\hbar\omega(n_x + n_y + n_z)} \right).$$

Since the expression inside the sum only depends on n , the inner sum reduces to simply counting $\Omega(n)$. We begin by choosing any number between 0 and n for n_x . In general, after choosing n_x , we will then have $n - n_x + 1$ possibilities for n_y . This fixes n_z . Thus,

$$\Omega(n) = \sum_{n_x=0}^n (n - n_x + 1) = \frac{(n+1)(n+2)}{2}.$$

Thus the partition function is

$$\ln Z_G = \sum_{n=0}^{\infty} \Omega(n) \ln(1 - ze^{-\beta\hbar\omega n}) = \sum_{n=0}^{\infty} \frac{(n+1)(n+2)}{2} \ln(1 - ze^{-\beta\hbar\omega n}).$$

- (c) In general, a sum can be approximated by an integral when it does not change much over an interval of length 1. If we ignore the prefactor, this occurs when the energy spacing between the levels is small, i.e. $\beta\hbar\omega \leq 1$.
- (d) For high n , the prefactor simplifies to $\frac{n^2}{2}$. Thus, we approximate the partition function using an integral

$$\begin{aligned} \ln Z_G &\approx \int_0^{\infty} \frac{n^2}{2} \ln(1 - ze^{-\beta\hbar\omega n}) \, dn \\ &= \left[\frac{n^3}{6} \ln(1 - ze^{-\beta\hbar\omega n}) \right]_0^{\infty} - \int_0^{\infty} \frac{n^3}{6} \frac{1}{1 - ze^{-\beta\hbar\omega n}} (-ze^{-\beta\hbar\omega n}) (-\beta\hbar\omega) \, dn \end{aligned}$$

We must show that the constant part vanishes. Clearly the lower limit vanishes. The upper limit is a bit more subtle: We expand

$$\ln(1 - x) \approx -x$$

and note that the exponential prefactor now standing in front of the polynomial forces the term to vanish. Thus, we have

$$\begin{aligned} \ln Z_G &= -\frac{\beta\hbar\omega}{6} \int_0^{\infty} \frac{n^3}{\frac{1}{z}e^{\beta\hbar\omega n} - 1} \, dn \\ &= -\frac{\beta\hbar\omega}{6} \int_0^{\infty} \frac{1}{(\beta\hbar\omega)^4} \frac{(\beta\hbar\omega n)^3}{\frac{1}{z}e^{\beta\hbar\omega n} - 1} \, d(\beta\hbar\omega n) \\ &= \frac{1}{6(\beta\hbar\omega)^3} (\Gamma(4) \text{Li}_4(z)) \\ &= \frac{1}{(\beta\hbar\omega)^3} \text{Li}_4(z) \end{aligned}$$

(e) The grand canonical potential is given by

$$J = -\frac{1}{\beta} \ln Z_G.$$

We can then determine the expectation value of the particle number as

$$N = - \left(\frac{\partial J}{\partial \mu} \right)_T.$$

This gives us

$$\begin{aligned} \langle N \rangle &= \frac{1}{\beta} \frac{\partial}{\partial \mu} \ln Z_G \\ &= \frac{1}{\beta} \frac{1}{\beta \hbar \omega} \frac{\partial}{\partial \mu} \text{Li}_4(z) \\ &= \frac{1}{\beta} \frac{1}{\beta \hbar \omega} \left(\frac{\partial}{\partial z} \text{Li}_4(z) \right) \frac{\partial z}{\partial \mu} \end{aligned}$$

The first derivative is clear:

$$\frac{\partial z}{\partial \mu} = z\beta.$$

For the second, we use the hint to get

$$\begin{aligned} \langle N \rangle &= \frac{1}{\beta(\beta \hbar \omega)^3} \beta z \text{Li}_4'(z) \\ &= \frac{1}{(\beta \hbar \omega)^3} \text{Li}_3(z) \end{aligned}$$

Anticipating a transition when lowering temperature, we analyse what happens when we lower the temperature while keeping the number of particles constant. As temperature decreases, the polylogarithmic function must increase to keep the number of particles constant. Hence, z increases and thus μ must increase. However, the maximum value of μ is 0 - at this value of μ , we have $z = 1$ and $\text{Li}_3(1) = \zeta(3) \approx 1.2$.

Since we can decrease T arbitrarily, but cannot increase $\text{Li}_3(z)$ arbitrarily to compensate, we must conclude that we have missed out something in our analysis. This is the assumption that there are no macroscopically occupied states. Thus, above a transition temperature T_c , we have a macroscopic occupation of the ground state.

We compute this occupation number by bounding $\langle N \rangle$ using the maximum value

$$\langle N \rangle \leq \frac{\zeta(3)}{(\beta \hbar \omega)^3} = \zeta(3) \left(\frac{k_B T}{\hbar \omega} \right)^3.$$

Thus, for a fixed total particle number N_{tot} , we find that the critical temperature is

$$T_c = \left(\frac{\hbar\omega}{k_B} \right) \left(\frac{N_{\text{tot}}}{\zeta(3)} \right)^{1/3}.$$

Above this temperature, the occupation is microscopic and follows the distribution derived above. Below this temperature, the ground state is occupied microscopically; the number of particles in the excited states is limited to our bounding value, and all particles not there are in the ground state.

- (f) Below the transition temperature, the number of particles in excited states remains at

$$N = \frac{\zeta(3)}{(\beta\hbar\omega)^3}.$$

The number of particles in the ground state is hence

$$N_0 = N_{\text{tot}} - \frac{\zeta(3)}{(\beta\hbar\omega)^3}.$$

If we define $f = N_0/N_{\text{tot}}$, we get

$$f = 1 - \frac{\zeta(3)}{N_{\text{tot}}(\beta\hbar\omega)^3}.$$

- (g) Below the transition temperature, the particles in the ground state do not contribute any energy, and hence the energy is just

$$E = .$$

□

Problem 2. Prove that the entropy per photon in black body radiation is independent of temperature and in d spatial dimensions is given by

$$s = \frac{S}{N} = k_B(d+1) \frac{\sum_{n=1}^{\infty} n^{-d-1}}{\sum_{n=1}^{\infty} n^{-d}} = k_B(d+1) \frac{\zeta(d+1)}{\zeta(d)},$$

where $\zeta(s)$ is the Riemann Zeta function.

Proof. We

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