NUEN 647 Final Project

Uncertainty quantification of depletion calculations for specific isotopes using ORIGEN2.

I Introduction

Determing composition of irradiated fuel is of importance for a myriad of reasons. Whether for flux calculations, reprocessing, or irradiation history verification, calculating fuel composition requires a Bateman solver, and a means for building a sparse matrix.

Applications using these compositions rarely report the uncertainty associated with results, even when inputs, such as flux shape, fission yield, cross sections, and half-lives have varying degrees of uncertainty. Further sources of error in this calculation are due to the multi-group approximation, and single point approximation, but will not be explored here.

Several isotope concentrations, shown in Table 1, were calculated as a function of burnup with the depletion code ORIGEN2 for a PWR system with 3 Wt% enriched uranium. ORIGEN2 solves the bateman equations with the matrix expoential method and requires a library with decay and cross section information. Cross sections and fission product yields are reduced to single group through flux averaging before execution of the code, with the assumption that the flux has the same shape as a typical PWR.

The uncertainty of concentrations were determined by varying the absorption and fission cross sections for ²³⁵U, ²³⁸U ²³⁹Pu, ²⁴⁰Pu, and ²⁴¹Pu. Originally, absorption cross sections and yields for the fission products were to be varied, but variance information for the lighter nuclides is either difficult to acquire or not available.

The uncertainties on cross sections were determined by calculating the range of the single group cross section, taking the mid point as a mean, the range as a standard deviation, and assuming a Gamma distribution.

Table 1: Isotope solve list.

$^{133}\mathrm{Cs}$	$^{136}\mathrm{Ba}$	$^{153}\mathrm{Eu}$
$^{134}\mathrm{Cs}$	$^{138}\mathrm{Ba}$	$^{154}\mathrm{Eu}$
$^{135}\mathrm{Cs}$	$^{149}\mathrm{Sm}$	$^{239}\mathrm{Pu}$
$^{137}\mathrm{Cs}$	$^{150}\mathrm{Sm}$	$^{242}\mathrm{Pu}$
$^{148}\mathrm{Nd}$	$^{106}\mathrm{Rh}$	$^{125}\mathrm{Sb}$

II Objectives

☑ Build ORIGEN2 model for thermal system which calculates concentrations of isotopes shown in Table 1.

Listing 1: PWR Input Deck

-1 -1 -1

```
RDA
          Irradiation of 1 MT of PWR fuel
     RDA
          Fuel enrichment is 3.0 w/o U-235
     RDA
          LPU 922350 922380 942390 942400 942410 -1
     RDA
                    601 602 603
             1 2 3
                                     8 0 1 38
     LIB
             101 102 103
                           10
     INP
                 -1
                     -1
                           1
10
     BUP
     IRP 100.0 37.5 1 2 4 2 BURNUP=3,750 MWD/MT
        200.0 37.5 2 3 4 0 BURNUP=7,500 MWD/MT
        300.0 37.5 3 4 4 0 BURNUP=11,250 MWD/MT
15
        400.0 37.5 4 5 4 0 BURNUP=15,000 MWD/MT
     DEC 500.0
                    5 6 4 0 DECAY FOR 100.0 DAYS
     DEC 4150.0
                    6 7 4 0 DECAY FOR 10 YEARS
                    7 8 4 0 DECAY FOR 200.0 YEARS
     DEC 73500.0
     BUP
     OPTL 24*8
     OPTA 4*8 5 19*8
     OPTF 4*8 5 19*8
             8
                 1 -1
     OUT
     END
  2 922340 270. 922350 30000. 922380 969730. 0 0.0
   0
```

The model irradiates 1 metric ton of US PWR fuel for a single cycle (15,000 MWd/Mt). The calculations use a constant power assumption of 37.5 W/g. The model does not include the oxygen because we are not interested in the activation of oxygen. Cross section modification throughout the calculation use the changing flux associated with a US PWR.

Initial verification of the model analyzed the end concentration of 137 Cs and calculated the burn-up from that value. This calculation does not have an exact value for the yield of 137 Cs and is used qualitatively as a sanity check.

$$\frac{552.8 \text{ g}^{137}\text{Cs}}{Mt} \cdot \frac{6.022E23 \text{ atoms}}{137 \text{ g}^{137}\text{Cs}} \cdot \frac{\text{Fission}}{0.06 \text{ atoms}} \cdot \frac{200 \text{ MeV}}{\text{Fission}} \cdot \frac{1.602E - 19 \text{ MJ}}{1 \text{ MeV}} \cdot \frac{1 \text{ day}}{86400 \text{ s}} = 15,018 \frac{\text{MWd}}{Mt}$$

 \square Determine how to vary cross section and or flux spectrum inputs for calculation

ORIGEN2 reads in cross section information through a file named "TAPE9.inp", specified by the 8th input on the LIB card. "TAPE9.inp" needs at least 3 cross section libraries. These are specified by the 5th 6th and 7th inputs on the LIB card as 601, 602, and 603 for the activation products, actinides, and fission products, respectively. The input for ²³⁵U from library 601 is shown in the listing below, with a corresponding key shown in Table 2 [1].

Listing 2: ²³⁵U cross section library 602 input

602	922350	1.068E+01	2.338E-03	8.049E-07	4.752E+01	0.0	0.0	-1.0
00 -	022000	1.0002,01	2.0002 00	0.0102 0.	111022101	0.0	0.0	1.0

Table 2: Key to parameters in cross section library

LIB	NUCLID	(n,γ)	(n,2n)	(n,3n)	(n,f)	(n,γ^*)	(n,2n*)	YYN
-----	--------	--------------	--------	--------	-------	----------------	---------	-----

The cross sections, σ_{γ} and σ_{f} will be modified based on the locations in the cross section libraries shown above. Half-life information is contained in the decay libraries 1, 2, and 3 for activation products, actinides, and fission products, respectively. These values will not be modified to reduce the scope of the project.

Modifying the cross section libraries seems to have no bearing on the output of the code. This will be shown in the results section.

A program was writen to modify σ_{γ} , yield, or half-life based on lines from a text file. The project was simplified to only modify cross section input for the actnides, and therefore a modified version of is provided below. The code below is used to modify σ_{γ} and σ_{f} .

Listing 3: Script for modifying ORIGEN2 input.

```
#!/usr/bin/env python3
 #Please note, might have to run command
 # sed -i 's/E /E\+/g' TAPE9_BANK.inp
 # on file to make sure there are no spaces after E's
 ######################### Import Packages ####################
 import time
 start_time = time.time()
 import numpy as np
 import Functions as Fun
 from subprocess import call
 #As opposed to high class
 class ToPull:
    def init (self):
      self.Isotope = ""
25
      self.Type
 ListToPull=[]
 with open ('LISTTOPULL') as f:
    Lines=f.readlines()
 for i in Lines:
    i=i.split()
    ListToPull.append(ToPull())
    ListToPull[-1].Isotope =i[0]
    ListToPull[-1].Type
 ######################### Load all samples ##################
 with open ('SAMPLES/94Pu239a') as f:
    Pu239a=f.readlines()
```

```
Pu239a=Fun.StripNL(Pu239a)
 with open ('SAMPLES/94Pu240a') as f:
    Pu240a=f.readlines()
 Pu240a=Fun.StripNL(Pu240a)
 with open ('SAMPLES/94Pu241a') as f:
    Pu241a=f.readlines()
 Pu241a=Fun.StripNL(Pu241a)
 with open ('SAMPLES/94Pu239f') as f:
    Pu239f=f.readlines()
 Pu239f=Fun.StripNL(Pu239f)
 with open ('SAMPLES/94Pu240f') as f:
    Pu240f=f.readlines()
 Pu240f=Fun.StripNL(Pu240f)
 with open ('SAMPLES/94Pu241f') as f:
    Pu241f=f.readlines()
 Pu241f=Fun.StripNL(Pu241f)
60
 with open ('SAMPLES/92U235a') as f:
    U235a=f.readlines()
 U235a=Fun.StripNL(U235a)
 with open ('SAMPLES/92U238a') as f:
    U238a=f.readlines()
 U238a=Fun.StripNL(U238a)
 with open ('SAMPLES/92U235f') as f:
   U235f=f.readlines()
 U235f=Fun.StripNL(U235f)
 with open ('SAMPLES/92U238f') as f:
    U238f=f.readlines()
 U238f=Fun.StripNL(U238f)
 with open ('ORIGENBACKUP/TAPE9 BANK.inp') as f:
    TAPE9content=f.readlines()
 f.close()
 for k in range (0, len(U238f)):
 #for k in range(0,1):
    \#k=2
    #Open output file
95
    output=open("TAPE8.inp", "w")
    #Loop through the TAPE9 template file, make changes,
```

```
#and write to new TAPE9 (output)
100
      for i in TAPE9content:
         hold=i.split()
         toprint=False
         if '602' in hold[0] and "-" not in hold[0]:
             if "942390" in hold[1]:
                #Replace the gamma x-section
105
                i=i.replace(hold[2],Pu239a[k])
                #Replace the fission x-section
                i=i.replace(hold[5],Pu239f[k])
                toprint=True
             if "942400" in hold[1]:
110
                #Replace the gamma x-section
                i=i.replace(hold[2],Pu240a[k])
                #Replace the fission x-section
                i=i.replace(hold[5],Pu240f[k])
                toprint=True
115
             if "942410" in hold[1]:
                #Replace the gamma x-section
                i=i.replace(hold[2],Pu241a[k])
                #Replace the fission x-section
                i=i.replace(hold[5],Pu241f[k])
120
                toprint=True
             if "922350" in hold[1]:
                #Replace the gamma x-section
                i=i.replace(hold[2],U235a[k])
                #Replace the fission x-section
125
                i=i.replace(hold[5],U235f[k])
                toprint=True
             if "922380" in hold[1]:
                #Replace the gamma x-section
                i=i.replace(hold[2],U238a[k])
130
                #Replace the fission x-section
                i=i.replace(hold[5],U238f[k])
                toprint=True
             if "ACTINIDE+AND" in i:
                toprint=True
135
             if toprint:
                i=i.replace("\n","")
                print(i, file=output)
      print(" -1", file=output)
140
      output.close() #important or ORIGEN2 wont run
      #endfile record being detected otherwise
      145
      #Run with new TAPE9, also deletes old files
      print("Run number :"+str(k+1))
      call(["./r"])
150
      ################## COLLECT OUTPUT #############################
```

```
155
      #Open output file and save contents
      with open ('TAPE6.OUT') as f:
          content=f.readlines()
      #Work on Page by Page Basis for output
160
      #Also grab time steps
      Pages=[];Page=""
      TimeDays=[]; AD="ACTINIDES+DAUGHTERS" #For space
      for i in content:
          if "PAGE" in i:
165
              #Save the Page if it has useful information
              if AD in Page or "FISSION PRODUCTS" in Page:
                 Pages.append(Page)
             Page=""
          else:
170
             Page=Page+i
          #If we have located the time steps pull the info
          if "TIME, SEC" in i:
             i=i.replace("TIME, SEC","")
             i=i.split()
175
             for j in i:
                 TimeDays.append(str(round(float(j)/86400,2)))
      #Pages has a list of strings, each with a page on it
      #from the output
180
      #The only pages saved in Pages are ones with text:
      #"ACTINIDES+DAUGHTERS" or "FISSION PRODUCTS" the rest
      #of the output is trashed
      #Loop through all isotopes and grab info
185
      for item in ListToPull:
          Isotope=item.Isotope
          Type=item.Type
          Info=Fun.Graboutput(Isotope, Pages, Type)
190
             SaveFile=open("OUTPUTDATA/"+Isotope+'.out','w')
             print('D,'.join(TimeDays)+'D',file=SaveFile)
             print(','.join(Info),file=SaveFile)
             SaveFile.close()
          else:
195
             SaveFile=open("OUTPUTDATA/"+Isotope+'.out','a')
             print(','.join(Info),file=SaveFile)
             SaveFile.close()
   print("--- %s seconds ---" % (time.time() - start time))
```

✓ Determine variance of Cross-sections

Both σ_{γ} and σ_{f} were determined by flux averaging via:

$$\sigma = \frac{\int \sigma(E)\phi(E)dE}{\int \phi(E)dE}$$

where:

$$\phi(E) = C_1 \cdot \frac{E}{E_0^2} \cdot exp\left(-\frac{E}{E_0}\right) \qquad E < E_{max,th}$$

$$= \frac{C_2}{E} \qquad E_{max,th} < E < E_{max,epi}$$

$$= C_3 \cdot \frac{\sqrt{\frac{E}{E_f}}}{E_f} \cdot exp\left(-\frac{E}{E_f}\right) \qquad E > E_{max,epi}$$

and:

$$C_1 = \frac{E_0^2}{E_{max,th}^2} e^{E_{max,th}/E_0}$$

$$C_2 = 1$$

$$C_3 = \frac{E_f}{E_{max,epi}} \cdot e^{\frac{E_{max,epi}}{E_f}} \frac{1}{\sqrt{\frac{E_{max,epi}}{E_f}}}$$

Where: $E_{max,th} = 0.50$ eV, $E_{max,epi} = 1E5$ eV, $\theta_{th} = 0.09$ eV (764 K), and $\theta_{fis} = 1.35E6$ eV. These values were picked because they minimized the difference between the cross-sections in the TAPE9 file, and those calculated with ENDF-VII and the above method. This is highlighted in figure 1. Most of this error is from 238 U and 240 Pu.

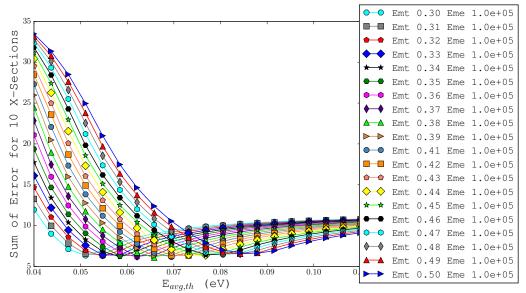


Figure 1: Minimized error for 10 cross section calculations

The flux spectrum is shown as below:

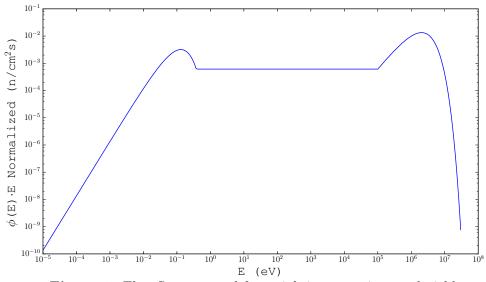


Figure 2: Flux Spectra used for weighting x-sections and yields

Table 3 shows the difference between my calculated cross section and the cross sections provided in ORIGEN2. The ratios will be used as a correction factor for the cross sections so that the input to ORIGEN2 will be constant. Differences have been attributed to the fact that the flux spectrum used does not capture reasonances. The cross sections were calculated with ENDF VII, and ORIGEN2 could have been processed with ENDF V, which would further account for differences, but has not been verified. The reason ENDF VII is being used for the current analysis is because the variance information is included in ENDF VII and not ENDF V.

Table 3: Comparison of one-group cross sections

$Isotope^{Rxn}$	ENDF VII	ORIGEN2	Ratio
$^{239}\mathrm{Pu}^{\gamma}$	6.544e + 01	6.909E+01	1.06
$^{240}\mathrm{Pu}^{\gamma}$	1.521e + 02	2.228E+02	1.46
$^{241}\mathrm{Pu}^{\gamma}$	4.518e + 01	4.202E+01	0.93
$^{235}\mathrm{U}^{\gamma}$	9.387e + 00	1.068E+01	1.14
$^{238}\mathrm{U}^{\gamma}$	4.098e+00	8.872 E-01	0.22
$^{239}\mathrm{Pu^f}$	1.179e + 02	1.211E+02	1.03
$^{240}\mathrm{Pu^f}$	9.609 e-01	5.787E-01	0.60
$^{241}\mathrm{Pu^f}$	1.253e + 02	1.259E+02	1.01
$^{235}\mathrm{U^f}$	4.621e+01	4.752E + 01	1.03
$^{238}\mathrm{U^f}$	2.091e-01	9.281E-02	0.44

Both σ_{γ}^{error} and σ_{f}^{error} were determined by calculating the single group cross section with error subtracted, and then added. These values will constitue a mean with error. This was done with the following code:

Listing 4: Script calculating average x-sections

```
#!/usr/bin/env python3
  11 11 11
  This program will compute 1-group cross sections with a weighted
 flux. Parameters for the flux were
5 determined in a subdirectory called Reduce Err.
  import time
 start time = time.time()
 import Functions as f
 from scipy import interpolate
 from scipy import integrate
 from scipy.integrate import trapz
  #To fix X-section data to ORIGEN values
 Ratios=[1.05578868993,1.46437937788,0.929974069639,1.13769098926,
       0.216470218472,1.02697467277,0.602248684356,1.0049132997,
       1.02836592353,0.4437673342571
  ############### Import X-Section Data #####################
 #Get list of csv files with X-section information
 Names=f.GETcsvFiles("X Sections")
 Names=["Pu 239 94 a.csv",
      "Pu 240 94 a.csv",
35
      "Pu 241 94 a.csv",
      "U 235 92 a.csv",
      "U 238 92 a.csv",
      "Pu_239_94_f.csv",
      "Pu_240_94_f.csv",
      "Pu_241_94_f.csv",
      "U_235_92_f.csv",
      "U_238_92_f.csv"]
45 | #Flux Parameters
 Emt = 0.38
          #Max thermal energy in ev
 Eme=1e5
           #Max epithermal energy in ev
 E0=0.0658
            #Thermal average in ev (1045 K)
           #Fission average in ev
 Ef=1.35e6
  #Loop through all the X-sections I got
  index=0
  for Name in Names:
    Element=Name.split('_')[0]
```

```
Isotope=Name.split('_')[1]
       Protons=Name.split('_')[2]
       Reaction=Name.split('_')[3].split('.')[0]
       #Do not do Averaging of variances
60
       if 'V' in Reaction:
           continue
       Xsec = f.np.genfromtxt('X_Sections/'+Name, delimiter=',')
       #Modify Xsections to match with ORIGEN2
65
       Xsec[:,1]=Xsec[:,1]*Ratios[index]
       index=index+1
       #Set energy, and convert from MeV to ev
       E=f.copy.copy(Xsec[:,0])*10**6
70
       #Gather Variance and make function for it
       VarName=Name.split(".")[0]+"V.csv"
       Var=f.np.genfromtxt('X_Sections/'+VarName, delimiter=',')
       Var_int=interpolate.interp1d(Var[:,0],Var[:,1],
75
                                     fill_value=0,bounds_error=False)
       #Determine the absolute err from the variance
       ErrAb = (Var_int(E)/100) *Xsec[:,1]
       #Find minimum X-section
80
       Xmin=Xsec[:,1]-ErrAb
       #Find Max X-section
       Xmax=Xsec[:,1]+ErrAb
       #Calculate flux (yes we need E)
85
       F=f.flux(E,Emt,Eme,E0,Ef)
       #Make function for Min-X-Section(E) * Flux(E)
       X_phimin=interpolate.interp1d(E,F*Xmin,
                                   fill_value=0, bounds_error=False)
90
       #Make function for Max-X-Section(E) * Flux(E)
       X_phimax=interpolate.interp1d(E,F*Xmax,
                                   fill_value=0,bounds_error=False)
       #Perform the integral for Max-X-Section(E) * Flux(E)
       X_int_max=integrate.trapz(X_phimax(E),E)
       #Perform the integral for Min-X-Section(E) * Flux(E)
       X_int_min=integrate.trapz(X_phimin(E),E)
       #Perform the integral for Flux(E)
       Phi_int=integrate.trapz(F,E)
       #Average X-section value
100
       Avgmin=X int min/Phi int
       Avgmax=X_int_max/Phi_int
       Avg= (Avgmin+Avgmax) /2
105
       print (Protons+Element+Isotope+Reaction+' '+str(Avg)+
              '+/-'+str(Avg-Avgmin))
       #With the ratio fixes, Ratio should be one
       #Find TAPE9's X-section value for comparison
110
       #TAPE9_X=f.LoopTAPE(Protons, Isotope, Reaction)
```

With results in Table 4

Table 4: Errors in single group cross sections

$Isotope^{Rxn}$	σ with 1STD Error
$^{239}\mathrm{Pu}^{\gamma}$	69.09 ± 8.15
$^{240}\mathrm{Pu}^{\gamma}$	222.8 ± 50.9
$^{241}\mathrm{Pu}^{\gamma}$	42.02 ± 10.92
$^{235}\mathrm{U}^{\gamma}$	10.68 ± 3.23
$^{238}\mathrm{U}^{\gamma}$	0.887 ± 0.175
$^{239}\mathrm{Pu^f}$	121.1 ± 1.2
$^{240}\mathrm{Pu^f}$	0.579 ± 0.003
$^{241}\mathrm{Pu^f}$	125.9 ± 2.3
$^{235}\mathrm{U^f}$	47.52 ± 0.71
$^{238}\mathrm{U^f}$	$0.093 \pm 8.2 e-7$

Cross section versus flux for several isotopes are shown in the following figures.

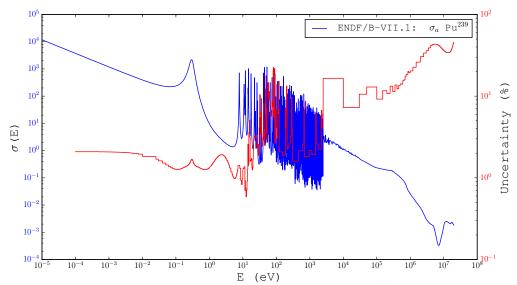


Figure 3: σ_a with corresponding error for ²³⁹Pu

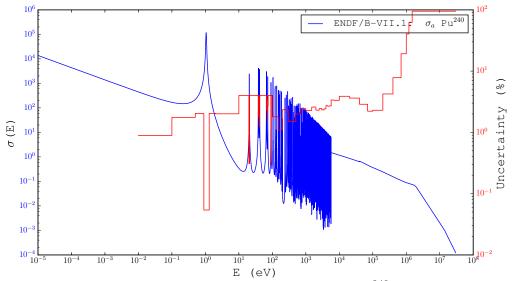
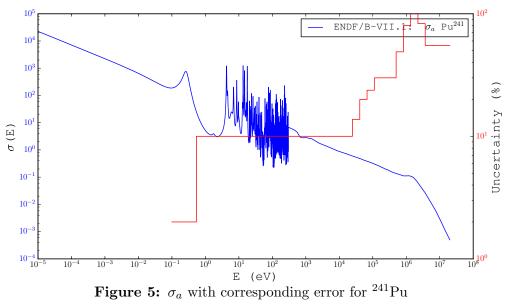
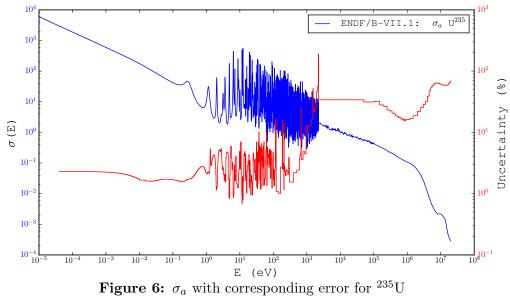
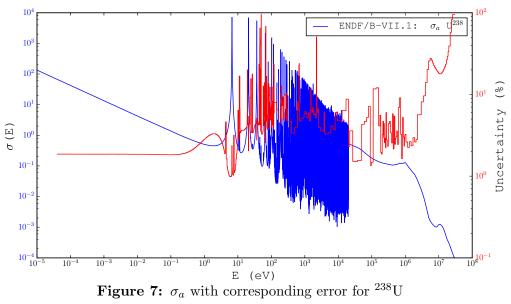
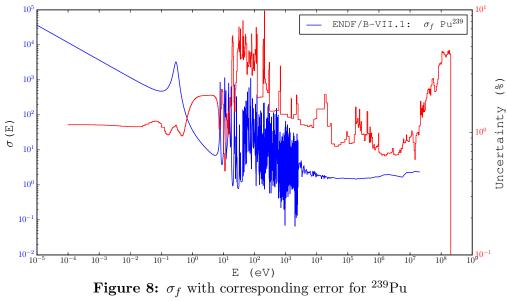


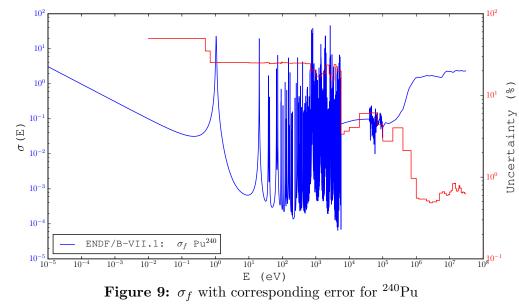
Figure 4: σ_a with corresponding error for $^{240}\mathrm{Pu}$

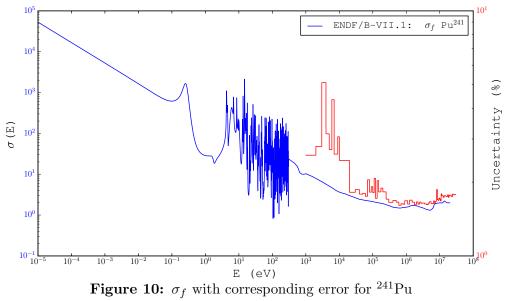


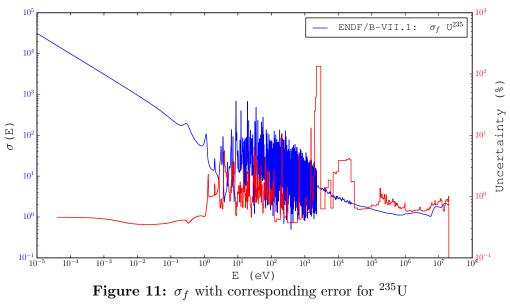












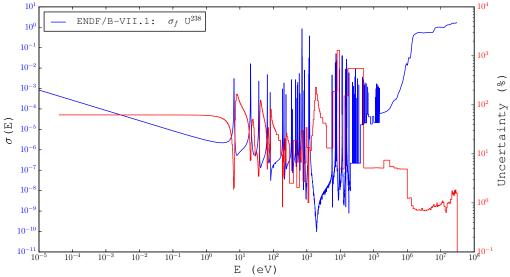


Figure 12: σ_f with corresponding error for $^{238}\mathrm{U}$

The distribution for all 10 of the above cross sections are assumed as Gamma distributions to avoid negative cross section values.

$$\pi(\theta) = \frac{\theta^{\alpha - 1} e^{-\theta/\beta}}{\Gamma(\alpha)\beta^{\alpha}}, \qquad \theta, \alpha, \beta > 0.$$

Therefore, we say that $\theta \sim G(\alpha, \beta)$. Where the mean and errors determined above fit into α and β via

$$\alpha = \frac{\mathrm{Mean}^2}{\mathrm{Error}^2}$$

and

$$\beta = \frac{\text{Error}^2}{\text{Moon}}$$

With the following codes the sampling space was determined.

Listing 5: Sample Generation code

```
import Functions as Fun
15
 class OneGroupwError:
   def __init__(self):
      self.Element = ""
      self.XSec
              = 0
                   #X Section
      self.Err
              = 0
                   #Error
 N=Fun.N
 Nbins=Fun.Nbins
 0=[]
 with open ('OneGroupXSections') as f:
   Lines=f.readlines()
 for i in Lines:
   i=i.split()
   O.append(OneGroupwError())
40
   O[-1].Element = i[0]
   O[-1].XSec = float(i[1].split('+/-')[0])
   O[-1].Err = float(i[1].split('+/-')[1])
45
 for i in O:
   #Make Samples
   alpha=(i.XSec**2)/(i.Err**2)
   beta=(i.Err**2)/(i.XSec)
   Sample=np.random.gamma(alpha, beta, size=N)
50
   #Make histogram plot
   Fun.PlotHistSave (Sample, N, i. Element, Nbins)
   #Convert Data to string
   Sample=[str(x) for x in Sample]
   #open the output file and save
   output=open("SAMPLES/"+i.Element,"w")
   print("\n".join(Sample),file=output)
   output.close()
 print("--- %s seconds ---" % (time.time() - start time))
```

Below are histograms of the sampling space, most of the others look the same, with samples surrounding the mean.

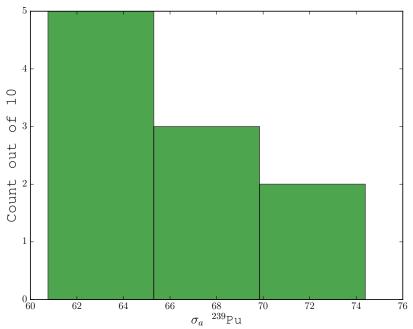
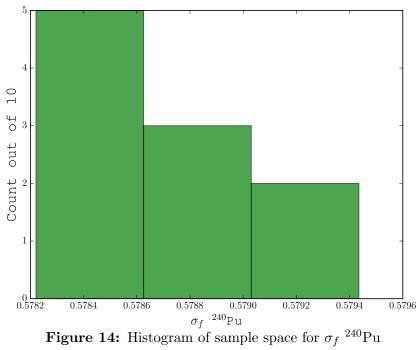


Figure 13: Histogram of sample space for σ_a ²³⁹Pu



The sampling spaces were run through ORIGEN with the code shown in the first section.

☑ Plot Results

The function for plotting results are shown below.

Listing 6: Code for plots

```
#!/usr/bin/env python3
 11 11 11
5
 import time
 start time = time.time()
 import Functions as Fun
 from scipy import interpolate
 from scipy import integrate
 from scipy.integrate import trapz
 import sys
 import os
 import numpy as np
 import copy
 #For histogram plots
 TimeIndex=3
 #Below should be the same as in 'Sample_Gen' file (I know its lame)
 N=Fun.N
Nbins=Fun.Nbins
 #Which Time Steps for range...
 T1 = 0
 Tf=5
 #15 Elements, cant do all at once Do 8 at a time (lame again)
 \#E1 = 9
 \#Ef = 1.5
 E1 = 0
 |Ef=9
 #As opposed to high class
 class ToPull:
   def __init__(self):
     self.File = ""
 #Gather all output file names
 ListToPull=[]
 with open ('LISTTOPULL') as f:
   Lines=f.readlines()
 for i in Lines:
```

```
i=i.split()
55
     ListToPull.append(ToPull())
     ListToPull[-1].File =i[0]+'.out'
  #Loop through output file names
  for index in range(E1,Ef):
     fileparse=ListToPull[index].File
65
     #Save output file information
     with open ("OUTPUTDATA/"+fileparse) as f:
        Lines=f.readlines()
70
     #Get time information from first row
     TimeList=Lines[0].replace('\n','').replace('D','').split(',')
     Time=[]
     for times in TimeList:
        Time.append(float(times))
75
     #Make matrix of data for rest
     Data=[]
     del Data
     for i in range(1,len(Lines)):
80
         line=Lines[i].replace('\n','')
        line=line.split(',')
        line=[float(i) for i in line]
        try:
            Data=np.vstack((Data, line))
85
         except NameError:
            Data=copy.copy(line)
     #Plot Time vs grams for all different runs
     HISTSamples=[]
90
     fig=Fun.plt.figure(figsize=Fun.FigureSize)
     ax=fig.add_subplot(111)
     label=''
     for run in Data:
        HISTSamples.append(run[TimeIndex])
95
         #print(run) Plot
         (fig, ax) =Fun.plot(Time[T1:Tf], run[T1:Tf],
                       ax,'black',label,fig,fileparse)
     Element=fileparse.split('.')[0]
     Fun.plt.savefig("PLOTS/"+Element+'Post XY.pdf')
     #Make histogram for specified time for all different runs
     Fun.PlotHistSave2(HISTSamples, N, fileparse, Nbins)
  ############################ Time to Execute ###################################
  print("--- %s seconds ---" % (time.time() - start_time))
```

IIIResults

Quantities of interest are shown in Table 1 above. The uncertain parameters were the absorption and fission cross-sections for ²³⁵U, ²³⁸U, ²³⁹Pu, ²⁴⁰Pu, and ²⁴¹Pu.

Results are graphically represented below.

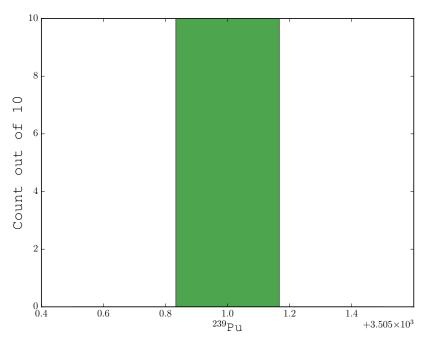
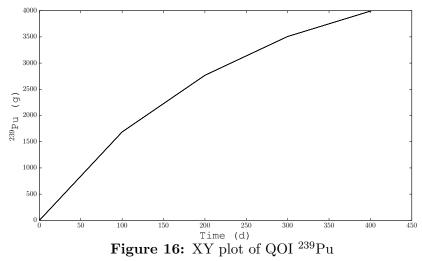
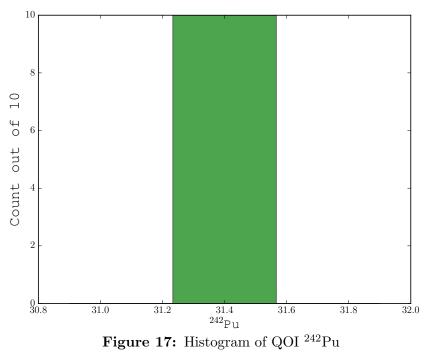
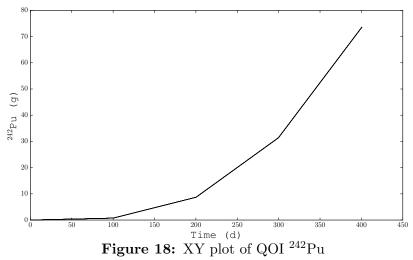


Figure 15: Histogram of QOI 239 Pu







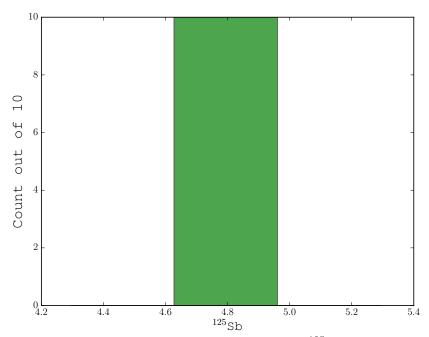
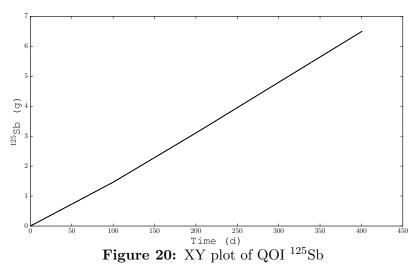


Figure 19: Histogram of QOI $^{125}\mathrm{Sb}$



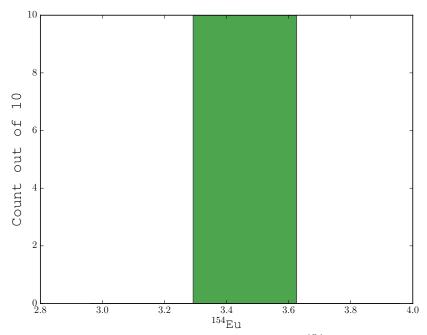
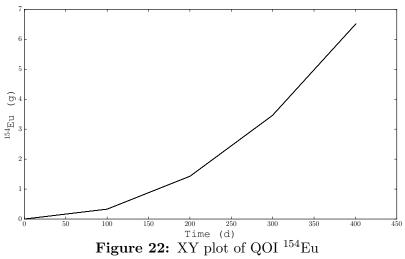


Figure 21: Histogram of QOI 154 Eu



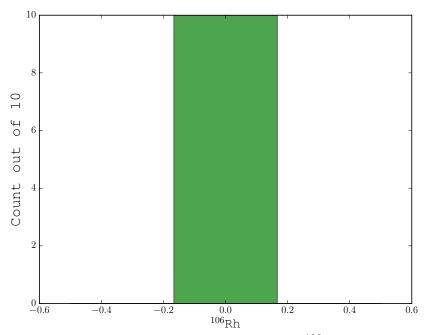
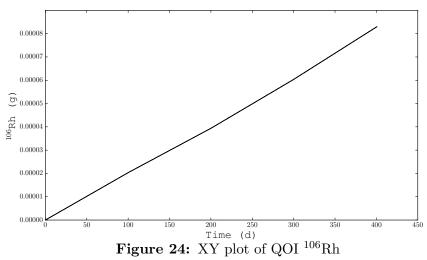


Figure 23: Histogram of QOI $^{106}\mathrm{Rh}$



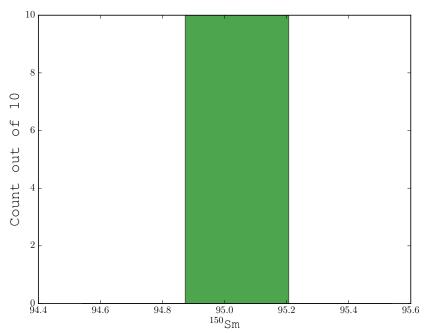
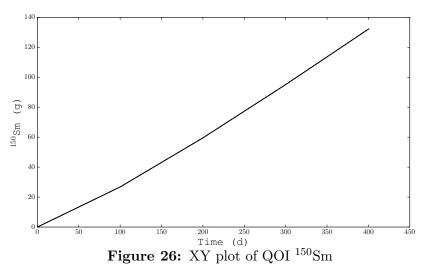


Figure 25: Histogram of QOI $^{150}\mathrm{Sm}$



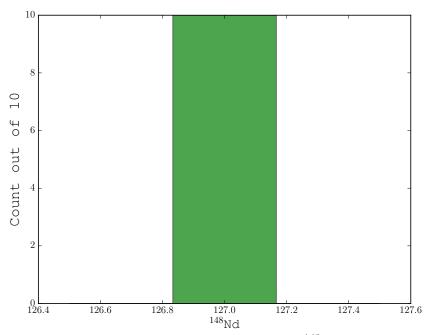
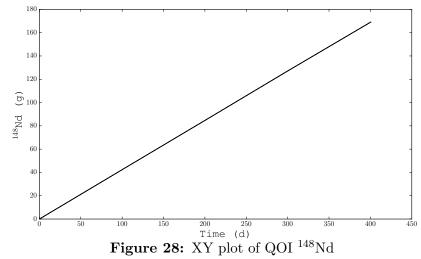


Figure 27: Histogram of QOI $^{148}\mathrm{Nd}$



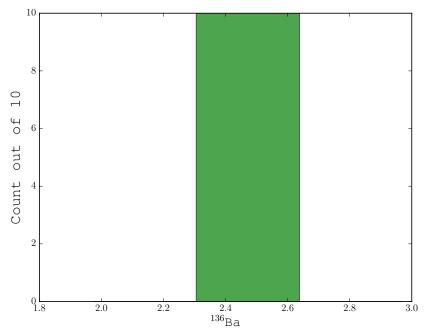
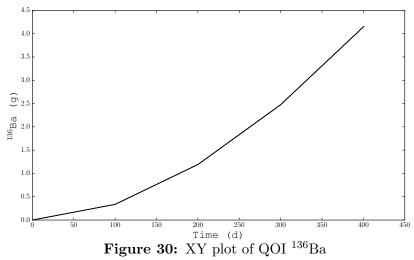


Figure 29: Histogram of QOI 136 Ba



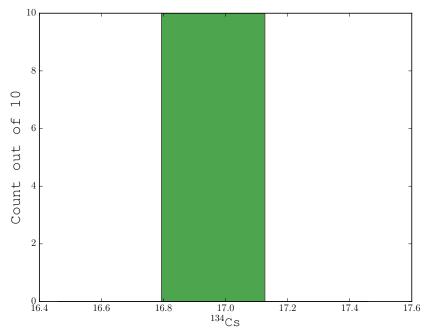
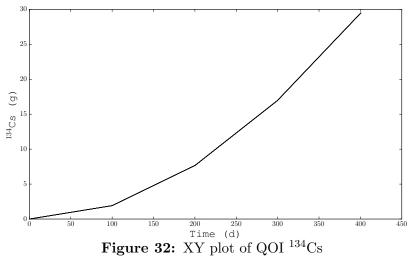


Figure 31: Histogram of QOI ¹³⁴Cs



I apologize for the wall of graphs. I left some out.

Conclusions IV

There was an unexpectedly large amount of error in the one group cross sections, but currently there is no variation in the output, even when I change the cross section by a factor of over 1000, I don't think ORIGEN is taking up what I am laying down.

References

[1] Allen G Croff. User's manual for the origen2 computer code. Technical report, Oak Ridge National Lab., 1980.