# Comparing clouds and their seasonal variations in ten atmospheric

#### general circulation models with satellite measurements 3

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- [1] To assess the current status of climate models in simulating clouds, basic cloud
- climatologies from ten atmospheric general circulation models are compared with satellite 10
- measurements from the International Satellite Cloud Climatology Project (ISCCP) and the 11
- Clouds and Earth's Radiant Energy System (CERES) program. An ISCCP simulator is 12 employed in all models to facilitate the comparison. Models simulated a four-fold
- difference in high-top clouds. There are also, however, large uncertainties in satellite high 14
- thin clouds to effectively constrain the models. The majority of models only simulated 15
- 30-40% of middle-top clouds in the ISCCP and CERES data sets. Half of the models 16
- underestimated low clouds, while none overestimated them at a statistically significant 17
- level. When stratified in the optical thickness ranges, the majority of the models simulated 18 optically thick clouds more than twice the satellite observations. Most models, however, 19
- underestimated optically intermediate and thin clouds. Compensations of these clouds 20
- biases are used to explain the simulated longwave and shortwave cloud radiative forcing at 21
- the top of the atmosphere. Seasonal sensitivities of clouds are also analyzed to compare 22
- with observations. Models are shown to simulate seasonal variations better for high clouds 23
- 24 than for low clouds. Latitudinal distribution of the seasonal variations correlate with
- satellite measurements at >0.9, 0.6-0.9, and -0.2-0.7 levels for high, middle, and low 25
- clouds, respectively. The seasonal sensitivities of cloud types are found to strongly depend 26
- on the basic cloud climatology in the models. Models that systematically underestimate 27
- middle clouds also underestimate seasonal variations, while those that overestimate 28
- optically thick clouds also overestimate their seasonal sensitivities. Possible causes of the 29
- systematic cloud biases in the models are discussed. 30
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- with satellite measurements, J. Geophys. Res., 110, XXXXXX, doi:10.1029/2004JD005021. 32

### Introduction

[2] Clouds are intrinsically coupled with the chaotic moist atmospheric circulations. Aside from directly interacting with air motions through latent heating, clouds also produce a net energy loss or gain to the atmosphere-Earth system through their radiative effects. Variations of clouds

thus have the potential to either amplify or reduce a climate 40 change. It has been known that the sensitivity of a climate 41 model strongly depends on its clouds [Cess et al., 1990; 42 Senior and Mitchell, 1993, 1996; Le Treut and Li, 1991; 43 Roeckner et al., 1987], and models simulate different cloud 44 feedbacks [Cubasch et al., 2001; Bony et al., 2004]. In the 45 last 10 years, several research programs have been orga- 46

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nized to narrow the range of uncertainty in cloud-climate feedbacks.

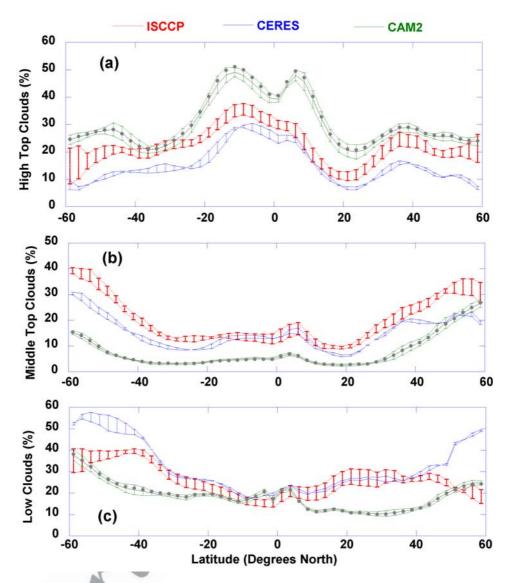
- [3] The Cloud Parameterization and Modeling Working Group (CPM) within the Atmospheric Radiation Measurement (ARM) program organized a model intercomparison project to compare cloud climatologies from general circulations models with satellite measurements. The first purpose of the project is to assess the current status of climate models in simulating clouds so that future progress can be more objectively measured. The second purpose is to reveal serious deficiencies in these models so as to guide future measurement and single-column modeling/cloud system resolving modeling (SCM/CSRM) activities [Randall et al., 1996; Ghan et al., 2000; Xu et al., 2002; Xie et al., 2002]. Weare and AMIP Modeling Groups [1996] compared zonally averaged model clouds from the Atmospheric Model Intercomparison Projects (AMIP I) with surface and satellite measurements. Weare [2004] also evaluated clouds in six AMIP II models against observations. The AMIP I study showed that global means of model high cloudiness are about two to five times larger than satellite measurements, but low clouds are 10% to 20% less than satellite and surface observations. The AMIP II study showed that the models simulated moderately well cloud albedo, but not the cloud water path.
- [4] The effort to more quantitatively characterize cloud errors in models has been hampered by the considerable amount of uncertainties of available cloud data. Each cloud data set, whether satellite based or surface based, has its specific viewing geometry that blocks some clouds. Past efforts have tried to empirically adjust model clouds into diagnostics that are comparable to observations. Assumptions made in these adjustments introduce additional uncertainties in the comparison. Another uncertainty is about the exact definition of cloudiness. Within a model, one can either use a threshold of hydrometeor concentration or an optical depth of the cloud condensate to define clouds. In observations, however, the threshold is dependent on the cloud detection algorithm, which is further related with satellite pixel sizes. In this regard, cloud radiative forcing (CRF) is a more objective measure. However, CRF only measures the accumulative effects of clouds and therefore it does not necessarily provide the physical insight on particular model biases.
- [5] The present study of evaluating clouds in climate models is aided by simulating the results of the International Satellite Cloud Climatology Project (ISCCP; see Rossow and Schiffer [1999]) through an ISCCP simulator developed in Klein and Jakob [1999] and Webb et al. [2001]. The ISCCP simulator not only minimizes the difference of sampling geometry between the models and data, but also allows the comparison of model cloud types with measurements that are stratified into both altitude ranges and optical properties. This paper reports results from this intercomparison project. Several studies have used the same approach to evaluate model clouds and their associated cloud radiative forcing at the top of the atmosphere [Webb et al., 2001; Tselioudis and Jakob, 2002; Williams et al., 2003; Lin and Zhang, 2004]. This study employs a wide range of models with different physical parameterization components to analyze common model biases. Furthermore, seasonal sensitivities of clouds in the models are evaluated and their

relationships with the basic climatology are examined. The 109 paper is organized as follows. Section 2 gives a brief 110 description of the data and the models used. Section 3 111 analyzes common model biases. Section 4 studies the 112 seasonal sensitivity of cloud types in the models. The last 113 section contains a summary and discussion of the results.

## 2. Data and Models

## 2.1. ISCCP and CERES Cloud Data

- [6] The ISCCP D2 monthly cloud frequencies are used in 117 the comparison. ISCCP combines infrared and visible 118 radiances from geostationary satellites with the TIROS 119 Operational Vertical Sounder (TOVS) atmospheric temper- 120 ature and humidity as well as correlative surface ice/snow 121 data to obtain cloud information. It first identifies whether 122 an image pixel of size 4-10 km is cloudy or clear, and then 123 retrieves the optical thickness. Optical depth is used to 124 estimate cloud emissivity, which is then used to determine 125 the cloud-top pressure. The cloudy pixels are sorted into 126 different bins defined by ranges in cloud-top pressure 127 (height) and cloud optical depth [Rossow et al., 1996] 128 sampled at a nominal resolution of 30 km and 3 hours. 129 The frequencies of each cloud bin are accumulated to an 130 equal area map with 280 km resolution. They are further 131 averaged to monthly means. Three bins of cloud-top 132 pressures (high: <440 mb, middle: 440-680 mb, and 133 low: >680 mb) and three bins of cloud optical thickness 134 (thin:  $\langle 3.6, \text{ medium: } 3.6-23, \text{ and thick } \rangle 23$ ) define nine 135 ISCCP cloud types. They will be referred to as high thin 136 or middle thick clouds, and so on.
- [7] Four seasons of CERES clouds binned into the same 138 ISCCP optical depth and cloud altitude ranges are also 139 used in this study. This product is based on pixel-level 140 radiance data in multiple channels from the Terra Moder- 141 ate Resolution Imaging Spectrometer (MODIS) [King et 142 al., 1992]. The CERES cloud detection algorithm differs 143 from ISCCP in that it uses four instead of two wave- 144 lengths to decide whether a given pixel is clear or cloudy 145 [Trepte et al., 1999]. Radiance measurements at these four 146 wavelengths are also used to estimate cloud phase, optical 147 depth, particle size, and temperature [Minnis et al., 1995, 148 1998]. Cloud-top height is determined from cloud temper- 149 ature using a lapse rate method [Minnis et al.,. 1992; 150 Garreaud et al., 2001] for low clouds over ocean areas 151 and using the vertical temperature and water vapor profiles 152 from the NASA Global Modeling and Assimilation 153 (GMAO)'s Goddard Earth Observing System DAS product 154 (GEOS-DAS V4.0.3) (see http://dao.gsfc.nasa.gov/sub- 155 pages/atbd.html) for other regions. In addition, the CERES 156 cloud analysis is performed using 1-km MODIS pixels 157 sampled to a 4-km resolution instead of taking one 4–10 km 158 pixel in each 30 km by 30 km box as in ISCCP. Regional 159 cloud type frequencies derived for 1° × 1° latitude- 160 longitude equal area regions are used to determine cloud 161 type frequencies of occurrence over the entire  $60^{\circ}\text{S} - 60^{\circ}\text{N}$  162 domain.
- [8] Caveats in both the ISCCP and CERES data include 164 uncertainties in cloud detection, partial cloudy pixels, and 165 emissivity correction of cloud top altitudes. These mostly 166 affect optically thin clouds. Multilayer clouds with thin high 167 clouds above low clouds pose a special difficulty for ISCCP 168



**Figure 1.** Interannual ranges of clouds DJF seasons from ISCCP (red), CERES (blue), and CAM2 (green) for: (a) high top clouds, (b) middle top clouds, and (c) low clouds. El Nino years have been excluded in the ISCCP data and CAM simulations. The CERES data refer to the two DJF seasons of 2001 and 2002. The difference between the black line and dashed line with solid circle for the CAM2 represent whole day sampling versus daytime sampling of calculating the ISCCP clouds in the simulator.

and CERES to correctly retrieve cloud tops since all these methods are based on single-layer clouds. This factor however is considered in the ISCCP simulator, which uses a simulated infrared brightness temperature to determine cloud pressure under a single-layer cloud assumption.

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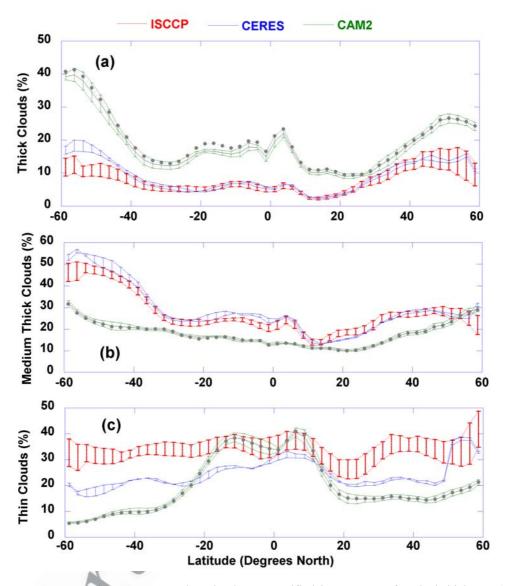
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[9] The ISCCP D2 data are available from 1983 to 2001. The CERES cloud data are from the DJF seasons of 2001 and 2002, and JJA seasons of 2000 and 2001. Each DJF season refers to the months of January and February of the year and December of the previous year. The red lines in Figure 1a show the range of maximum and minimum high cloud frequencies in ISCCP from ten DJF seasons after excluding the El Nino years of 1983, 1987, 1990–1995, and 1998. The blue lines show the range of high cloud frequencies from the 2 CERES DJF seasons of 2001 and 2002. While the patterns of the latitudinal distributions are similar

in the two data sets, high clouds are systematically less 185 frequent in the CERES, especially at middle latitudes. 186 Interannual variability of high cloud frequencies is generally 187 smaller than the difference between the two satellite data 188 sets. For the overlapping season of JJA in 2001 (not shown), 189 the difference in the two data sets is similar to that in 190 Figure 1a. This suggests that the main source of differ- 191 ence is from the cloud retrieving algorithms. This dis- 192 crepancy is primarily due to differences in optically thin 193 clouds. We have also examined high clouds from the 194 High Resolution Infrared Radiometer Sounder (HIRS) 195 [Jin et al., 1996; Wylie and Menzel, 1999] and the adjusted 196 surface-based observations [Norris, 1999], and found similar 197 differences among all these data sets. While it is beyond the 198 scope of this paper to explain the difference between the 199 satellite data sets, we believe that the cloud detection algo- 200



**Figure 2.** Same as Figure 1 except that clouds are stratified into ranges of optical thickness (a) for optical thick clouds, (b) for optically intermediate clouds, and (c) for optically thin clouds.

rithms, which are further related with the satellite pixel sizes, contributed significantly to these differences.

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[10] Middle and low clouds from ISCCP and CERES agree better except at high latitudes (Figures 1b and 1c). Both data sets show increasing amount of middle and low clouds toward high latitudes, especially the southern high latitudes. This latitudinal variation is also seen in surface-based observations. It is therefore unlikely due to the masking of high clouds in the satellite product. Poleward of about 40° CERES reported fewer middle clouds, but more low clouds than ISCCP. These differences in middle and low clouds at high latitudes may be caused by several factors, including the specifications of radiative and microphysical properties of liquid and ice particles, the temperature profiles, and the surface properties.

[11] When clouds are stratified against the optical thickness ranges, Figure 2 shows that optically thick and intermediate clouds in ISCCP and CERES agree well. The discrepancy between the two data sets is mainly in optically

thin clouds (Figure 2c) with ISCCP reporting about 32% 220 thin clouds versus 25% in CERES.

[12] Given the differences between the two data sets, in 222 our comparison, if a model is statistically different from 223 both data sets toward one direction, it is judged as biased. 224 As an example, Figures 1 and 2 also included simulated 225 clouds from the CAM2 as reported in the work of *Lin and* 226 *Zhang* [2004]. For each cloud type, if the model result 227 satisfies the following

$$C_m > \max(C_{\mathit{ISCCP}}, C_{\mathit{CERES}}) + \delta_{\max}$$
 or

$$C_m < \min(C_{ISCCP}, C_{CERES}) - \delta_{\min}$$

where  $C_m$ ,  $C_{ISCCP}$ ,  $C_{CERES}$  represent cloud frequencies from 232 the model, ISCCP, and CERES, the model is judged to 233 overestimate or underestimate clouds.  $\delta_{max}$  and  $\delta_{min}$  are the 234

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235 95% confidence intervals based on a student *t* test. Standard 236 deviations for the satellite data are from ten seasons of 237 ISCCP data.

### 2.2. Models and Simulations

- [13] Table 1 lists the participating models, along with their physical parameterization components related to clouds. Cloud fraction schemes can be categorized broadly into three groups: relative humidity-based schemes, statistical total water schemes, and prognostic cloud fraction schemes. In principle, the relative humidity cloud scheme can be also considered as a special case of the statistical scheme in which the probability distribution is implicitly assumed rather than diagnosed. All models used the maximum-random cloud overlapping assumptions in the vertical except for the GFDL model that will be described later.
- [14] Two versions of the UKMO GCM are used. These are the HadAM3 and HadAM4. Multiple enhancements are made in HadAM4 relative to HadAM3, including the boundary layer scheme, precipitation scheme and the statistical cloud scheme. These are described in the work of Webb et al. [2001] and Williams et al. [2003]. Three versions of the NCAR CAM are used (CAM2, CAM2c, and CAM2x). CAM2c differs from CAM2 with only one change: the triggering condition of convection in the model It is known that convection in the CAM2 is too frequent. An empirical convection triggering condition was imposed based on measurements of the large-scale atmospheric dynamics and SCM simulations at the ARM SGP site [Xie and Zhang, 2000]. This modification was implemented in the CAM2c by Xie et al. [2004] and was found to improve many aspects of the model climate. CAM2x is a developmental version of the CAM3. Main differences include the parameterization of clouds from shallow and deep convections, advection of condensates, and separation of cloud ice and cloud water in the CAM2x (W. D. Collins et al., The formulation and atmospheric simulation of the Community Atmosphere Model: CAM3, submitted to J. Climate, 2005, hereinafter referred to as Collins et al., submitted manuscript, 2005).
- [15] All models used the ISCCP simulator described by *Klein and Jakob* [1999] and *Webb et al.* [2001] and updated at http://gcss-dime.giss.nasa.gov/simulator.html. Since model grids are much coarser than the ISCCP or CERES pixels, clouds at each model layer are first downscaled to cloudy and cloud-free subcolumns, along with liquid and ice water contents. The columns are then vertically aligned with overlapping assumptions consistent with the radiation routines in the models. Optical depth of each subcolumn is calculated based on the model radiation code. The optical depths are then used to calculate the emissivity, brightness temperature, and emissivity-adjusted cloud top pressure. The frequencies in different bins of cloud optical depths and cloud top pressures are used to compare with satellite measurements.
- [16] The ISCCP simulator incorporates the satellite view of clouds and adjusts the physical cloud top to mimic satellite measurements. There are still issues in comparing its output with measurements. Chief among them is the cutoff value of the optical depth to define cloudiness. The satellite algorithms used a cutoff value of 0.02 as the lowest detectable limit. Because the satellites may not actually

report clouds at this level, a cutoff optical depth value of 296 0.3 is used in the ISCCP simulator as a proxy (B. Rossow, 297 personal communication, 2003). This choice is somewhat 298 arbitrary and it affects the model cloud amount. To under- 299 stand the sensitivity of ISCCP clouds with this cutoff value, 300 Figure 3 shows the frequency differences of high thin 301 clouds from the ISCCP simulator relative to a cutoff optical 302 thickness value of 0.4 in the CAM2. High thin clouds are 303 almost linear to the cutoff value in the optical thickness 304 range of 0.1 to 0.4 in this model, with about 6% relative 305 increase of cloud frequency per 0.1 optical thickness de- 306 crease. Below 0.1, the sensitivity is larger. The HIRS and 307 ISCCP comparison by Jin et al. [1996] also implied large 308 sensitivity of high thin clouds to the cutoff value of optical 309 thickness in real observations. Figure 3 also includes the 310 differences of middle and low thin clouds between the 311 optical thickness of 0.01 and 0.4. There is very little 312 sensitivity in these cloud types. We therefore have less 313 confidence in optically thin clouds than optically thick to 314 intermediate clouds in inferring model biases.

- [17] Other issues about the ISCCP simulator include the 316 possibility of subpixel clouds that the satellite sees as 317 optically thin clouds, assumptions on the vertical overlap- 318 ping of the subcolumns, and the distribution of condensates 319 among the subcolumns. These are expected to be on the 320 secondary order. Two versions of the GFDL model are used 321 to highlight the impact of subgrid-scale cloud structure on 322 the ISCCP simulator results. One version (GFDL0) uses the 323 random overlap assumption, and a constant value of con- 324 densate in each layer for every subcolumn in the ISCCP 325 simulator. The second version (GFDL) infers internal inho- 326 mogeneity in each layer by fitting a symmetric beta distri- 327 bution of total water with fixed shape parameters to the 328 model values of cloud fraction and mean condensate. Cloud 329 overlapping varies smoothly between maximum and ran- 330 dom as the distance between adjacent layer increases 331 [Hogan and Illingworth, 2000; R. Pincus et al., Overlap 332 assumptions for assumed-PDF cloud schemes in large-scale 333 models, submitted to Journal of Geophysical Research, 334 2004]. Since this structure is used in both the ISCCP 335 simulator and the radiation calculations [Pincus et al., 336 2003], the model produces slightly different clouds in the 337 two simulations. The second version is more physically 338 based in terms of both its overlapping assumption and its 339 subgrid-scale inhomogeneity. 340
- [18] Figure 4 shows the cloud frequencies produced by 341 the different implementations of the ISCCP simulator in 342 these two simulations relative to ISCCP and CERES data, 343 averaged equatorward of 60°, for the nine ISCCP cloud 344 types in DJF. The second version simulated slightly more 345 optically thin clouds but less optically thick clouds. This is 346 likely due to the impact of the consideration of subgrid- 347 scale inhomogeneity. This version also simulated less low 348 clouds. This is consistent with the impact of the new 349 overlapping scheme relative to the random assumption. 350 None of these differences in the two calculations modifies 351 the nature of the model biases relative to the two satellite 352 measurements, and so the second version is used in the rest 353 of this paper. For all other models, including those 354 that employ statistical distribution of cloud liquid, the 355 ISCCP simulator only assigns constant condensates to the 356 subcolumns.

t1.1		Table 1. Description of Physical Parameterization Schemes in the	rization Schemes in the Models				
t1.2		Stratiform Clouds	Convective Clouds	Convection	PBL	Cloud Microphysical and Precipitation	Resolution
t1.3	CAM2	RH-based; Kiehl et al. [1996]	diagnostic; Rasch and Kristjánsson [1998]	Hack [1994]; mass flux; Zhang and McFarlane [1995]	first-order nonlocal;  Holtslag and Boville [1993]	Rasch and Kristjánsson [1998]	T42L26
t1.4	CAM2c	as above	as above	as above, modified with <i>Xie et al.</i> [2004]	as above	as above	as above
t1.5	t1.5 CAM2x	as above with modifications in the work of Collins et al.	diagnostic amount; Collins et al. (submitted manuscript, 2005)	as above	as above	as above with modifications; Collins et al. (submitted manuscript, 2005)	as above
9. 6 of	GFDL	prognostic; Teathe [1993]; Geophysical Fluid Dynamics Laboratory Global Atmospheric Model Development Team (GFDL	prognostic; <i>Tiedike</i> [1993]; <i>GFDL GAMDT</i> [2004]	RAS; Moorth and Suarez [1992]	cloud entrainments;  Lock et al. [2000];  GFDL GAMDT [2004]	Rotstayn [1997]; GFDL GAMDT [2004]	$2.5 \times 2.0 \text{ L}24$
185	GISS	GAMDT) [2004] RH based, Sundqvist type;	diagnostic; $Del$	mass flux; Del Genio	second-order;	Del Genio et al. [2005]	$4 \times 5 L12$
t1.8	GSFC	Det Genio et al. [2002] RH based, Sundqvist type; Del Genio et al. [1996]	diagnostic; Del Genio et al [1996]	and Ido [1993] RAS; Moorthi and Suarez [1992]	2.5 order; Helfand and Lahraca [1988]	Del Genio et al. [1996]; Sud and Walter [1999]	$2.5 \times 2 \text{ L40}$
t1.9	HadAM3	statistical; Smith [1990]	diagnostic; Gregory and Rowntree [1990]	mas flux, Gregory and Rowntree [1990]; Gregory, and Allen [1901]	first-order, Smith [1990]; Pope et al. [2000]	Smith [1990]	$3.75 \times 2.5L19$
t1.1	t1.10 HadAM4	as above with modifications of Webb et al. [2001], Cusack et al. [1999]	as above with modifications Gregory [1999]	mass flux; Gregory and Rowntree [1990]; Gregory and Allen [1991]	first-order with cloud entrainment; Lock et al. [2000]; Martin et al. [2000]	Wilson and Ballard [1999]	$3.75\times2.5L30$
t1.1	t1.11 ECHAM5	prognostic; TownLine [2002]	diagnostic; Roeckner	mass flux; <i>Tiedtke</i> [1989];	first-order, Brinkop	Lohmann and Roeckner [1996]	T42L19
t1.1;	t1.12 LMD	statistical; [2002] Le Treut and Li [1991]	statistical; Bony and Emanuel [2001]	Emanuel [1991]	first-order Li [1999]	Le Treut and Li [1991]	$3.75 \times 2.5L19$

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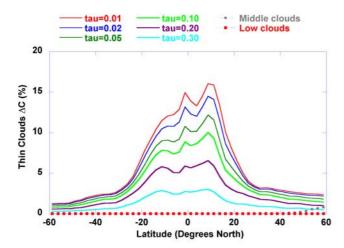
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**Figure 3.** Frequency differences of high thin clouds from the ISCCP simulator with different optical cutoff values of optical thickness relative to the cutoff optical thickness value of 0.4 in the CAM2. For middle and low thin clouds, the differences are between the optical thickness of 0.01 and 0.4.

[19] Since the calculation of optical depth in the models does not explicitly involve sunlight, the ISCCP simulator can calculate cloud types during nighttime, while the ISCCP cloud data are for daytime only. In Figures 1 and 2, simulated clouds from the whole day and daytime only calculations in the CAM2 are shown as the black line and the dashed line with solid circles. While diurnal variations of clouds are seen, their magnitudes are much smaller than the signatures sought in this study. Clouds are from daytime-only simulations for most models.

[20] Some participating models carried out AMIP long simulations. Others only carried out one-year simulations. Ranges of interannual variabilities of model clouds from 10 DJF seasons in CAM2 are shown in Figures 1 and 2 as the green lines. The interannual variations in the model are smaller than those in ISCCP. When statistical tests are carried out, we use the ISCCP standard deviations as an upper bound estimates for the models. All model results are from 1-year integrations spun off from an AMIP type simulation with prescribed monthly sea surface temperature (SST). Most models used SST for the year 2000. The ISCCP cloud data are also from year 2000. The CERES data are averages for the two winter and summer seasons of 2001 and 2002, respectively.

### 3. Results

[21] Frequencies of high clouds from ISCCP, CERES and the models for the DJF season are shown in Figure 5a. Most models are able to simulate the observed cloud maximum in the tropics and minimum in the northern subtropics. The simulated frequencies however differ greatly among the models even though the same cutoff values of the optical thickness (0.3) are used in the ISCCP simulators. There is a four-fold difference in the simulated high clouds among the models, ranging from 11% in the GISS GCM to 44% in the GSFC when averaged from 60°S to 60°N. The ISCCP and CERES measurements are 23% and 15% respectively with

an estimate of the interannual standard deviation of 1.5%. In 395 the tropics from 30°N to 30°S, the GSFC model and the 396 CAM series overestimated high clouds above the 95% 397 confidence level; the GISS and HadAM3 underestimated 398 high clouds. In middle latitudes poleward of 30°, GSFC, 399 CAM2, CAM2c, and ECHAM5 overestimated high clouds, 400 while GISS underestimated high clouds. High clouds from 401 half of the models are within the range of the two observational data sets. This is due to the large spread in the two 403 data sets. Along with uncertainties related to the cutoff 404 value of optical thickness for high thin clouds, the accuracy 405 of currently available satellite high clouds is therefore still 406 not sufficient to effectively constrain the models.

[22] While half of the models simulated high clouds 408 within the range of the two satellite data sets, the situation 409 is worse for middle and low clouds. Figure 5b shows that 410 most models substantially underestimated middle clouds. 411 Nine out of the ten models significantly underestimated 412 middle clouds in the tropics, and eight models did so in 413 middle latitudes. The grand mean of middle clouds from all 414 models is only one third of the satellite measurements in the 415 tropics and one half in middle latitudes. The GFDL model is 416 the only exception to simulate middle clouds in close 417 agreement with the ISCCP and CERES data at all latitudes, 418 while the GISS model did well at middle latitudes. Because 419 of the satellite view, middle clouds may be affected by 420 shielding from high clouds. We have examined the middle 421 cloud distribution in the CAM2 directly from the model 422 output without this shielding effect and they are still 423 significantly less than the satellite observation.

[23] Figure 5c compares the low clouds with satellite 425 measurements. In the tropics from 30°S to 30°N, eight out 426 of the ten models underestimated low clouds at the 95% 427 significant level. Poleward of 30°, half of the models underestimated low clouds at the statistically significant level. 428 None of the models overestimated low clouds. The grand 430 mean of model low clouds is about 70 and 80% of satellite 431 observations in the tropics and middle latitudes respectively. 432 This underestimation is probably a lower limit because there 433 is less shielding of low clouds by middle clouds in the 434 models. We have also examined the surface-based low 435 clouds adjusted to satellite view, courtesy of Joel Norris. 436 In low latitudes, the surface-based low clouds are more than 437 the two satellite data sets. In middle latitudes, they are less, 438 but are still more than those in half of the models.

[24] Compensations of model clouds at different heights 440 are seen in some models. The GISS model simulated the 441 smallest amount of high clouds, but it had the largest 442 amount of low and middle clouds that are in close agree-443 ment with the satellite data. The GSFC model simulated the 444 largest amount of high clouds, but it was among the models 445 simulating the smallest amount of middle clouds. The 446 HadAM3 and the LMD, on the other hand, were at the 447 lower end of the spectrum of simulating clouds at all 448 altitudes. Because of the use of the ISCCP simulator, a 449 model with excessive high cloud cover tends to have fewer 450 low clouds because of masking by high clouds. Examina-451 tion of the geographical distribution of different cloud types 452 however does not suggest this shielding effect to be the 453 leading cause of these compensations in the models.

[25] In summary, models simulated a four-fold difference 455 in high clouds. The available observations however also 456

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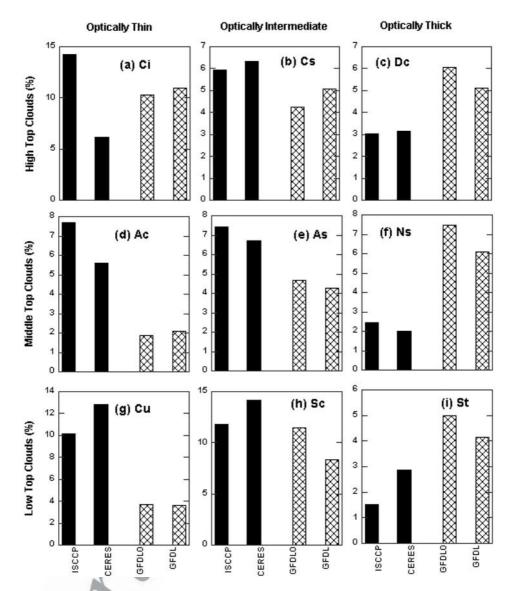
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**Figure 4.** Cloud frequency averaged from 60°N to 60°S in the DJF season for the nine ISCCP cloud types in ISCCP, CERES, and two versions of the ISCCP simulator in the GFDL model. See text for details.

show large spread. Models are found to significantly underestimate middle and low clouds. Averaged over 60°S to 60°N, the models simulated a grand mean of 6.9% middle clouds, in contrast to 17.5% in ISCCP and 14.3% in CERES. This is only 40% of the smaller satellite value from CERES. The models simulated a grand mean of 20.2% low clouds, in contrast to 23.4% and 29.9% in ISCCP and CERES measurements.

[26] We next examine the collective impact of these clouds on the cloud radiative forcing (CRF) at the top of the atmosphere (TOA). Figure 6 shows the longwave, shortwave, and net CRFs at the TOA, averaged from 60°S to 60°N for the DJF season from the ISCCP FD product [Zhang et al., 1995], ERBE, CERES, and all models. Standard deviations of interannual variation in the ISCCP FD, ERBE, the CAM2 for non-El Nino years, and the range of the two seasons in CERES, are also shown in Figure 6 as error bars. Mean values from ERBE and CERES are also

drawn as dashed horizontal lines. The ERBE data are from 475 the monthly S-4 product that combined measurements from 476 ERBS, NOAA-9 and NOAA-10 [Harrison et al., 1990]. 477 The CERES data are from its ERBE-like Monthly Regional 478 Averages (ES-9) product [Wielicki et al., 1996]. Both data 479 sets were acquired from the NASA Langley Distributed 480 Active Archive Center. The magnitudes of cloud forcing in 481 CERES are smaller than in ERBE. Similar to the cloud 482 products, we do not know the exact reasons of this differ- 483 ence. We thus use the same statistical procedure to judge the 484 model biases by using the interannual variations from the 485 ERBE in the t test. Only the GSFC model significantly 486 overestimated longwave cloud forcing, which is consistent 487 with the overestimation of high clouds in this model 488 (Figure 6a). Seven models underestimated the longwave 489 cloud forcing. This is consistent with the underestimation 490 of middle and low clouds in the models. CAM2 and 491 CAM2x simulated longwave cloud forcing that falls 492

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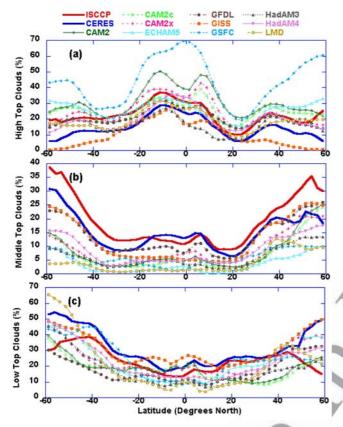
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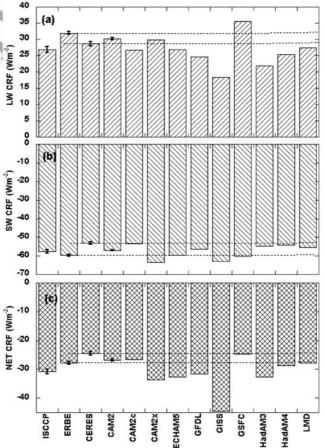
**Figure 5.** (a) High top clouds, (b) middle top clouds, and (c) low top clouds in the DJF from satellite measurements and from the models. ISCCP data are from year 2000, CERES data are from two seasons of 2001 and 2002. Model results are from one year simulations with most of them forced with prescribed monthly sea-surface temperature of year 2000.

within the range of the ERBE and CERES measurements since these models overestimated high clouds but with deficient middle and low clouds. The GISS GCM simulated the smallest longwave cloud forcing, consistent with underestimation of high clouds in this model. Overall, the underestimations of middle and low clouds in the models showed signatures in the longwave cloud forcing except for those that have overestimated high clouds.

[27] Since middle and low clouds should have a large impact on the shortwave CRF at the TOA, one would expect most models to also underestimate the magnitude of shortwave cloud forcing. This is, however, not seen in Figure 6b. Eight models simulated shortwave cloud forcing that is within the range of the two observational data sets. Two models (CAM2x and GISS) even overestimated the magnitude of shortwave cloud forcing. As a result, half of the models simulated the net cooling of clouds within the observational range, and the other half overestimated the net cooling effect (Figure 6c). This is consistent with *Potter and Cess* [2004] who demonstrated negative biases of the net CRF in many GCMs.

[28] The biases in the shortwave cloud forcing can only be explained by compensatory changes in cloud types. The ISCCP simulator allows us to compare model clouds with observation in optical thickness ranges. For optically thick 517 clouds (Figure 7a), eight models are found to overestimate 518 them at all latitudes. The GSFC and LMD models are the 519 exceptions. The grand mean of all models in the tropics is 520 12.4%, more than double the satellite measurements of 521 4.7% in ISCCP and 5.1% in CERES. In middle latitudes, 522 the simulated mean of optically thick cloud amount of 523 19.5% is also about twice of the observed values of 10% 524 in ISCCP and 12% in CERES. The HadAM3 simulated the 525 largest amount of optically thick clouds. This compensates 526 for the small amount of middle and low clouds to explain 527 its shortwave cloud forcing that was close to HadAM4 528 simulations.

[29] Figure 7b shows the comparison of optically inter-530 mediate clouds. The GSFC model is an outlier that simu-531 lated significantly more than observations. Except for the 532 GISS GCM, all eight other models underestimated this type 533 of clouds at the 95% statistically significance level. The 534 grand average of simulated cloud amount in the tropics, 535 including the GSFC values, is 16.2% in contrast to obser-536 vations of 20.0% in ISCCP and 21.8% in CERES. In middle 537



**Figure 6.** DJF cloud radiative forcing at the TOA averaged from 60°N to 60°S from measurements and from the models: (a) longwave CRF, (b) shortwave CRF, and (c) net CRF. The error bars for ISCCP FD, ERBE, CAM2 are their interannual standard deviations excluding the El Nino years. The error bar for CERES is from the two DJF seasons of 2001 and 2002. The ERBE and CERES values are also drawn as horizontal lines.

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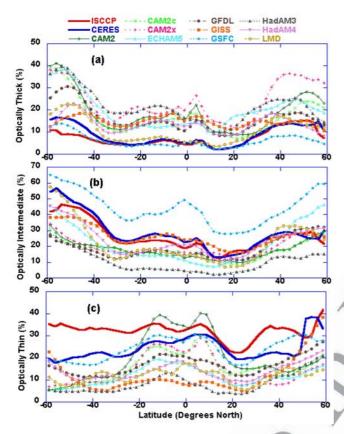


Figure 7. (a) Optically thick clouds, (b) optically intermediate clouds, and (c) optically thin clouds in the DJF from satellite measurements and from the models.

latitudes, all models except the GSFC and GISS models underestimated this cloud type, even though some models were able to simulate better in one of the hemispheres. The grand average of simulated amount in middle latitudes is 27.0% versus 32.2% in ISCCP and 34.7% in CERES. The HadAM3 simulated about one third of the satellite observations.

[30] Comparison of the optically thin clouds is shown in Figure 7c. Except for the CAM2 and the GSFC models, all other models significantly underestimated thin clouds. The models simulated about 60% of the smaller satellite values from CERES. These underestimations are mostly contributed by middle and low thin clouds, which are not very sensitive to the cutoff optical depth in the ISCCP simulators. The HadAM3 simulated one third of the CERES value.

[31] We have therefore seen systematic cloud biases in the models as follow: In the altitude ranges, models systematically underestimate middle and low clouds. In the optical thickness range, they overestimate optically thick clouds and underestimate optically thin and intermediate clouds. These biases can be more clearly seen in Figure 8 where the nine ISCCP cloud types are averaged from 60°N to 60°S. The same statistical procedure with error bars from ISCCP is used to judge the models against both satellite data sets.

[32] For high thin clouds (Figure 8a), the two satellite data sets show the largest differences, with CERES giving less than half of the ISCCP value. Eight models simulated high thin clouds within the two data sets. The two exceptions are the CAM2 that simulated excessive high thin 566 clouds and the GISS GCM that simulated too few of them. 567 The data sets therefore cannot effectively constrain the 568 models. Moreover, as noted in the previous section, the 569 cutoff value of the optical thickness used in the ISCCP 570 simulator has the largest impact on this cloud type. It is 571 difficult to draw firm conclusions from the comparison of 572 this cloud type with observations.

[33] For all other cloud types, the spread between the two 574 data sets is smaller. The impact of the optical thickness 575 cutoff value is also small or absent. For high clouds with 576 intermediate thickness, the GSFC model had significantly 577 more clouds than those in the other models and in the two 578 observations. The majority of models only slightly under- 579 estimated this cloud type, but the HadAM3 simulated one 580 third of the observed values. For high thick clouds, eight of 581 the ten models had significantly positive biases. The grand 582 average of high thick clouds from the models is 6.1%, about 583 twice the ISCCP value of 3.0% and the CERES value of 584 3.2%. The two exceptions are the GISS and the GSFC 585 models that simulated the correct amount of this cloud type. 586 The GISS model simulated small amount of high clouds in 587 all optical ranges, while the GSFC simulated more than 588 twice high clouds with intermediate optical depth.

[34] Middle clouds with thin and intermediate optical 590 depths are both significantly underestimated in all models. 591 The grand average of simulated middle thin clouds is only 592 15% of ISCCP value and 20% of the CERES value. The 593 grand average of middle intermediate clouds is only 40% of 594 the two satellite measurements. For middle thick clouds, 595 seven models overestimated this cloud type, with the GFDL 596 simulated three times more than observations. Three models 597 underestimated this cloud type, with the GSFC model 598 simulated about one quarter of the observed values.

[35] All models also underestimated low thin clouds, with 600 the grand mean only about 30% of the satellite measure- 601 ments. Because low thin clouds may be subpixel to satellite 602 measurements, they are better combined with low interme- 603 diate clouds as a single cloud type to minimize the mis- 604 match between observation and the ISCCP simulator 605 results. The grand mean of this combined type in all models 606 is only 55% of the ISCCP value of 21.8% and 65% of the 607 CERES value of 27.1%. On the other hand, eight models 608 significantly overestimated low top optically thick clouds, 609 with the grand mean two to three times of the satellite 610 measurements. The LMD model simulated this cloud type 611 comparable to measurements, while the GSFC model is the 612 only one that significantly underestimated this cloud type. 613

[36] To summarize the above results for systematic model 614 biases, we can categorize the nine ISCCP cloud types into 615 three groups. In one group, we combine the four ISCCP 616 cloud types with middle and low tops, thin and intermediate 617 optical depths. All models significantly underestimated 618 clouds in this group. The grand mean of all models is about 619 half of both the ISCCP and CERES measurements of 41% 620 and 43% respectively. In the second group, we combine the 621 three optical thick clouds of all altitudes. The majority of the 622 models significantly overestimated this group of clouds. 623 The grand mean of these cloud types in all models is 15.4%, 624 while the two satellite values are 6.9% and 8.1% respec- 625 tively. The only exception is the GSFC model that simulated 626 less than half of the low and middle top optically thick 627

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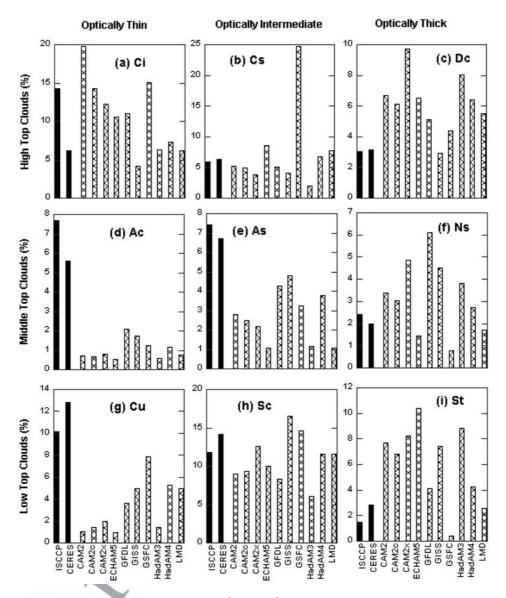
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**Figure 8.** Cloud frequency averaged from 60°N to 60°S in the DJF season for the nine ISCCP cloud types in satellite measurements and in the models.

clouds. The systematic overestimation of optically thick clouds is more pronounced for high and low top thick clouds than for middle top thick clouds, suggesting possibly multiple sources of model errors. The third group of clouds consists of high thin clouds. The spread in the two satellite data sets is too large to constrain the models.

[37] It is difficult to completely rule out the possibility that problems in the observations and the ISCCP simulator caused the systematic differences between models and data. In particular, it has been argued that middle clouds in ISCCP may be exaggerated by the misrepresentation of cirrus over low clouds, which ISCCP simulator may not completely capture because of possible sampling differences. We believe however that the above results reveal true physical deficiencies of clouds in the models. As an example, Figures 9a and 9b show the visible and infrared cloud images, respectively, from GOES east associated with an Atlantic cyclone on 19 February 2004 at 1500 UTC. Even though the optical properties of these clouds are not

available, the classic high-altitude comma-shaped frontal 647 cloud band and the low shallow cumulus and stratocumulus 648 after the cold front can be clearly identified from the 649 images. Figure 9c shows the 15-hour forecast simulation 650 of visible clouds as measured by TOA albedo from the 651 CAM2, which was initialized with the NCEP operational 652 analysis. It is seen that, consistent with the systematic 653 models biases discussed above, the CAM2 missed a con-654 siderable amount of low and middle clouds behind the front. 655 Figure 9d shows a forecast simulation of clouds from using 656 the Weather Research and Forecasts (WRF) mesoscale 657 model. Although the frontal cloud structure is improved in 658 the mesoscale mode, it still missed the low clouds behind 659 the cold front.

### 4. Seasonal Variations of Clouds

[38] Figure 10 shows the seasonal variations of clouds 662 from DJF to JJA for high, middle and low clouds in ISCCP, 663

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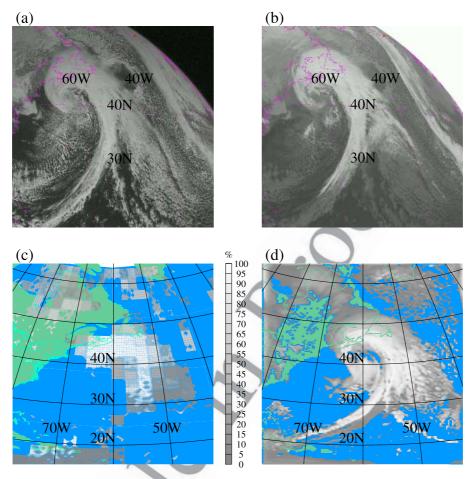
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**Figure 9.** (a) Visible and (b) infrared cloud images from GOES east at 1500 UTC on 19 February 2004. Simulated TOA albedo from 15-hour forecasts with NCEP operational initial conditions in (c) CAM2 and (d) WRF.

CERES and the models. Seasonal variation is defined as the difference of cloud frequencies between the JJA and the DJF seasons. In observations, the dominant pattern of seasonal high cloud variation is the movement of the ITCZ. There is little seasonal variation of high clouds poleward of 30°. The seasonal variation of middle clouds has a similar ITCZ pattern in the tropics, but it shows summertime reductions at middle latitudes, indicating the impact of wintertime middle latitude storm track clouds. This seasonal variability is more pronounced in the Northern Hemisphere than in the Southern Hemisphere. North of 50°, the two satellite data sets diverge significantly from each other for middle clouds, with the CERES data indicating summertime increase. Top view middle clouds from HIRS also show summertime reduction that is more like ISCCP. For low clouds, the two satellite data sets show reduction of clouds in the summer. This could be partly related with shielding of low clouds by high and middle clouds that maximize in the summer. The two satellite data sets again diverge from each other north of 50°, and the ISCCP seasonal variation is more consistent with both surface-based and HIRS measurements at these latitudes (figures not shown). We therefore use the ISCCP seasonal variation as a benchmark when quantitative evaluations are carried out.

[39] Most models were able to simulate the seasonal variation of high clouds (Figure 10a). The correlations with

ISCCP and CERES are all above 0.9. Weare and AMIP 690 Modeling Groups [1996] showed that nearly all AMIP I 691 models had tropical peaks in seasonal variability that were 692 poleward of observations. Figure 10a shows that only the 693 CAM2, CAM2x and the GSFC model still have this 694 tendency. The GISS model simulated a tropics-like variation 695 at middle latitudes. To facilitate the discussion, we define 696 seasonal amplitude as the area weighted root mean square of 697 the seasonal variation. The GISS GCM simulated the small-698 est amplitude, 81% of the ISCCP measurement. The GSFC 699 model simulated the largest seasonal amplitude, 180% of 700 the satellite data.

[40] For middle clouds, most models simulated summer-702 time reduction in middle latitudes, but they tend to simulate 703 little seasonal variability in the tropics (Figure 10b). The 704 correlations of the seasonal changes with observations are in 705 the range of 0.6 to 0.9. The ECHAM5 simulated the 706 smallest seasonal variations, with amplitudes about 50% 707 the ISCCP measurement, while the GISS model had the 708 largest seasonal amplitude, about 140% of the ISCCP value. 709 [41] The ability of the models in simulating low clouds is 710

[41] The ability of the models in simulating low clouds is 710 poorer (Figure 10c). Correlations of simulated low cloud 711 variations with ISCCP values are from -0.2 in the CAM2 712 to 0.7 in ECHAM5. The majority of the models had 713 correlations from 0.2 to 0.5. The RMS differences between 714 the models and ISCCP are as large as the observed seasonal 715

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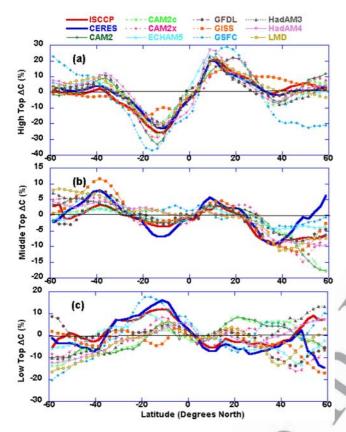


Figure 10. Seasonal variations (JJA minus DJF) of clouds from measurements and the models: (a) high top clouds, (b) middle top clouds, and (c) low top clouds.

amplitudes. These results are similar to what were found in the work of Weare and AMIP Modeling Groups [1996] about 10 years ago. While satellite-view low clouds are necessarily affected by middle and high clouds and may contain larger uncertainties than high and middle clouds, after examining the geographical distributions of seasonal variations of low clouds, we tend to conclude that the poorer quality of low cloud seasonal variation in the models is not mainly caused by the shielding effects of middle and high clouds.

[42] We next examine the individual ISCCP cloud types to search for the first-order controlling factors of the seasonal cloud variations in the models. Figures 11a and 11b show the seasonal variations of high clouds with intermediate and thick optical depth. The two satellite data sets agree well with each other. The seasonal amplitudes of high intermediate clouds in the models differ by several folds, with the HadAM3 showing the smallest amplitude of 30% of the ISCCP value, and the GSFC model showing the largest amplitude of four times the ISCCP measurement. Large differences are also seen for high thick clouds. The GSFC model had the smallest amplitude of 70% of the satellite data, while the two HadAM models showed the largest seasonal amplitudes of about twice the ISCCP and CERES values. Figures 11c and 11d relate the magnitudes of the seasonal amplitudes to the mean annual cloud frequencies for these two cloud types. The solid circle inside a square denotes the ISCCP data, and

that inside a triangle represents the CERES data. Models that 744 simulated large annual frequencies also had greater seasonal 745 variation, and vice versa. The linear correlation between the 746 seasonal amplitudes and the basic cloud amounts is 0.85 for 747 high intermediate clouds and 0.7 for high thick clouds. For 748 high thin clouds, the relationship (not shown) is similar to 749 those shown in Figures 11c and 11d among the models.

[43] A different type of relationships between the seasonal 751 variation and the basic distribution of clouds can be seen in 752 Figures 12a and 12c for middle thin clouds. Figure 12a shows 753 that the models produced very little seasonal changes in this 754 cloud type in the tropics and subtropics. This can be explained 755 by the basic annual cloud frequency in the models shown in 756 Figure 12c. All models simulated very little middle thin 757 clouds between 40°N and 40°S. Therefore there is also little 758 seasonal change and intermodel variability. A contrasting 759 case is shown in Figures 12b and 12d for low thick clouds. 760 The seasonal variations of low thick clouds in many models 761 are substantially larger than in the observations. With the 762 exception of the GSFC and the LMD models, all models 763 simulated more than twice the seasonal amplitudes of the 764 ISCCP value. The ECHAM5 amplitude is six times the 765 satellite data, followed by HadAM3 and CAM2x with 766 amplitudes 5 and 4 times of the ISCCP data. This exaggerated 767 variation in most models is clearly related to the basic cloud 768 distributions shown in Figure 12d since most models sub- 769 stantially overestimated the mean frequency of this cloud 770 type. The GSFC model had the smallest amount of low thick 771 clouds, and thus the smallest seasonal variation. The 772 ECHAM5 generated the greatest amount of this cloud type 773 and had the largest amount of seasonal variation. The LMD 774 model produced the best climatology of low thick clouds and 775 it simulated the best seasonal cycle of this cloud type relative 776 to ISCCP and CERES.

[44] The cloud biases illustrated in the previous section 778 thus have direct relevance to the sensitivities of clouds in 779 models. Even though our results do not invalidate cloud 780 feedbacks and climate sensitivity results from the models, 781 the fact that cloud feedback uncertainties have not been 782 reduced appreciably in the last 15 years [Cubasch et al., 783 2001; Bony et al., 2004] suggests the need to improve the 784 model clouds beyond what have been done.

## **Summary and Discussion**

[45] We have used ISCCP simulators in ten GCMs to 787 compare with satellite cloud analysis from ISCCP and 788 CERES. We have shown a four-fold difference in high 789 clouds among the models. There is also a large difference in 790 high thin clouds between the satellite data sets and a large 791 sensitivity of high thin clouds to the cutoff value of optical 792 thickness. The available satellite data are therefore not 793 accurate enough to constrain high thin clouds in most 794 models. We have also shown that the majority of models 795 only simulated thirty to 40% of the observed middle clouds. 796 Some models only simulated less than a quarter of observed 797 middle clouds. For low clouds, half of the models under- 798 estimated them while none overestimated them at the 799 statistically significant level. The grand mean of low clouds 800 from all models is about 70-80% of observations. When 801 stratified in the optical thickness ranges, the majority of the 802 models simulated optically thick clouds more than twice the 803

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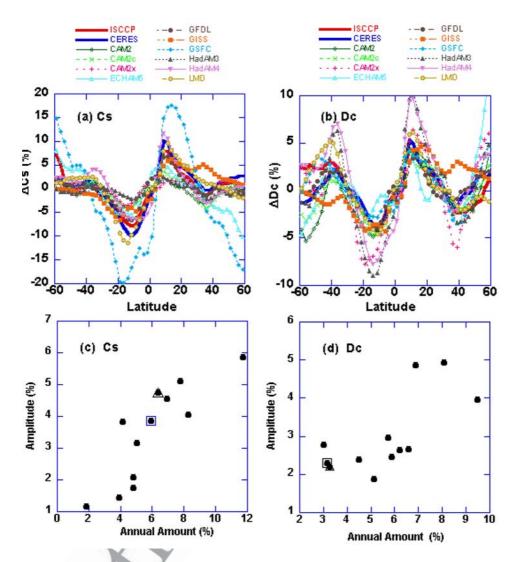
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**Figure 11.** Seasonal variations of high top clouds with (a) intermediate and (b) thick optical depths. Relationships between the seasonal amplitudes averaged from 60°N to 60°S with the mean cloud amount for high-top clouds with (c) intermediate and (d) thick. To fit the GSFC model in Figure 11c that is represented by the data point in the upper right corner, the magnitudes of the amplitude and basic frequency are all scaled by half.

satellite observations. Most models however underestimated optically intermediate and optically thin clouds. The grand mean of all models simulated about 80% of optical intermediate clouds and 60% of optically thin clouds. The underestimations of middle and low clouds are related with the negative biases of TOA longwave cloud forcing in some models, while the overestimations of optically thick clouds explains the reasonable or excessive shortwave cloud forcing in the models. These results further quantify the model cloud biases reported in the work of *Weare and AMIP Modeling Groups* [1996]. They also explain the *Weare* [2004] result that models tend to simulate the albedo moderately well, but not the cloud water path.

[46] We have categorized the nine ISCCP cloud types into three groups to describe the systematic model biases. The first group consists of middle and low clouds with intermediate and thin optical thickness. Models underestimate this group of clouds. The grand mean of all model results is about half of both the ISCCP and CERES measurements of 822 41% and 43% respectively. The second group consists of the 823 three optical thick clouds of all altitudes. The majority of the 824 models significantly overestimated this group of clouds. 825 The grand mean of these cloud types in all models is twice the 826 two satellite measurements of 6.9% and 8.1% respectively. 827 The third group consists of thin cirrus for which the models 828 show a several fold difference but they cannot be accurately 829 constrained by the available satellite data. 830

[47] We also presented seasonal sensitivities of clouds. 831 Models are shown to simulate latitudinal distributions of 832 seasonal variations that correlate with satellite measure- 833 ments at >0.9, 0.6–0.9, and -0.2 to 0.7 respectively for 834 high, middle and low clouds. For individual ISCCP cloud 835 types, the differences of seasonal amplitudes among the 836 models and satellite measurements can reach several hun- 837 dred percent. The dominant factor that determines the 838 seasonal amplitude of a particular cloud type in the models 839

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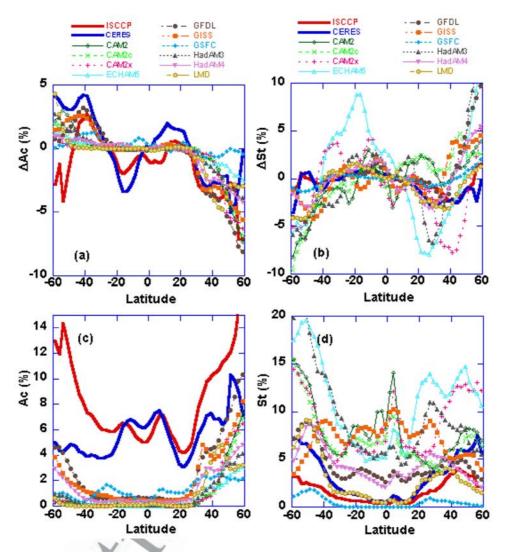
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**Figure 12.** Seasonal variations of (a) middle top thin clouds and (b) low top thick clouds. Basic annual cloud distributions for (c) middle top thin clouds and (d) low-top thick clouds.

is the simulated magnitude of the basic cloud frequencies. Models that systematically underestimate middle clouds also underestimate seasonal variations, while models that overestimate optically thick clouds also overestimate the seasonal sensitivity of these clouds.

[48] Even though the systematic cloud biases are common to most models, some models simulate certain cloud types better than others. It is highly desirable if the positive and negative attributes of model clouds can be associated with specific physical parameterizations. As pointed out in the work of Webb et al. [2001], in assessing clouds in models, many model components can be as important as the cloud and precipitation schemes. Without carrying out controlled experiments by isolating individual physical parameterization components, it is difficult to pinpoint the source of the model differences. Nevertheless, certain relationships with physical parameterizations can be observed. Both the HadAM4 and the GFDL models used the Lock et al. [2000] PBL scheme that contained additional turbulent mixing due to cloud top entrainment, and these two models simulated relatively better low clouds in all optical thickness ranges. On the other hand, the overestimation of optically

thick clouds is common to models that used very different 862 cloud schemes. For example, CAM used a relative humidity- 863 based cloud scheme, the HadAM used a statistical cloud 864 scheme, while the GFDL model used a prognostic cloud 865 scheme. They have similar biases in optically thick high top 866 clouds, in both the tropics and middle latitudes.

[49] Williams et al. [2003] and Lin and Zhang [2004] 868 showed that optically thick clouds occur with strong vertical 869 ascent associated with either large-scale convective systems 870 or middle latitude frontal systems in models. Cloud schemes 871 in current GCMs are designed to account for partial cloud 872 cover, but not for subgrid distribution of temperature and 873 moisture tendencies as a result of the subgrid circulations. 874 These subgrid structures are sometimes part of the large- 875 scale systems, sometimes they are self organized and main- 876 tained within the grid box [Katzfey and Ryan, 2000; Ryan et 877 al., 2000]. When a grid box has only a fractional area of 878 strong upward motion, the mean vertical motion for the 879 GCM grid box is upward. In a relative humidity-based 880 fractional cloud scheme, the fractional cloud cover merely 881 allows water to condense before the whole grid is saturated, 882 but it does not prevent the humidity in the clear-sky portion 883

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of the grid from rising as a result of the mean vertical motion. In traditional statistical schemes, the grid-scale upward motion typically shifts the total water distribution upward and the saturation threshold downward, rather than changing the subgrid-scale distributions of total water. In the prognostic cloud amount scheme of *Tiedtke* [1993], stratiform cloud tendency was formulated based on the difference of the grid box mean mixing ratio, which is controlled by the mean vertical motion, and its saturation value. One possible cause of the overestimation of high thick clouds in most models may be therefore due to the discretization of advective tendencies for the grid boxes, which has not been adequately accounted for in current fractional cloud schemes. Diagnosing the subgrid-scale distribution of total water as a function of subgrid-scale processes (e.g., convective or boundary-layer processes) may improve that situation. For instance, the LMD model, whose cloud scheme is physically coupled to the convection scheme and uses both large-scale and subgrid-scale predictors to diagnose the total water distribution [Bonv and Emanuel, 2001], does not overestimate the occurrence of optically thick clouds as much as most other models (Figures 7 and 8). Other possible causes of the model biases include vertical resolution and cloud microphysical properties. The overestimation of optically thick high clouds in the models can be also confirmed if daily TOA radiative fluxes are used to compare with model results. This has been reported in the work of Norris and Weaver [2001].

[50] The overestimation of low optically thick clouds in many models could be due to completely different reasons. Again, regardless of which type of cloud schemes is used, the mean relative humidity is a major control variable of clouds in the models. In the planetary boundary layer (PBL), the dynamic range of the relative humidity variation is small, which makes it a poor predictor of clouds. None of the models had a PBL scheme that directly predicts clouds. Instead, clouds are predicted by the model's stratiform or convective cloud scheme with its parameters modified by the PBL processes.

[51] With respect to the underestimation of low and middle optically thin clouds, it is well known that even mesoscale models cannot simulate clouds from shallow convections (see Figure 9 and Bretherton et al. [2004]). The simulation of cumulus and stratocumulus in the tropics has been a challenge for the modeling community for a long time. In middle latitudes, especially over the oceans, these same types of clouds frequently occur after a cold front due to the temperature contrast between warm water and cold air. Insufficient vertical resolution in the models is also a possible cause. In addition, observations also show that surface heterogeneity and topography can often generate shallow mesoscale and synoptic-scale circulation systems that cannot be adequately resolved by coarse resolution models.

[52] Much more needs to be done to fully understand the physical causes of model cloud biases presented here and to improve the models. Process oriented study, with observations of vertical cloud distributions and cloud optical properties, will be most useful to associate clouds with transient atmospheric dynamical circulations on different scales. Some of these challenges are discussed in several papers in this volume.

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