Stratiform precipitation, vertical heating profiles, and the Madden-Julian Oscillation

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August, 2002

ABSTRACT

The contribution of stratiform precipitation to the vertical profile of heating in the Madden-Julian Oscillation (MJO) is examined. Heating profiles in a strong MJO event in 1992-3 are calculated from the Tropical Ocean Global Atmosphere (TOGA) Coupled Ocean Atmosphere Response Experiment (COARE) sounding array data. Long-term heating estimates, derived from vorticity budgets in NCEP reanalysis, support the notion that this well-observed event was representative. Convective/stratiform precipitation estimates are obtained from TOGA COARE shipborne radars and the Tropical Rainfall Measurement Mission (TRMM) satellite. All datasets are filtered using a 30-70 day bandpass filter. The long-term MJO composites are constructed using linear correlation and linear regression with respect to time series of surface precipitation.

The observed MJO anomalous vertical heating profile is very top-heavy at the time of maximum precipitation. TRMM data and TOGA COARE radar data show that the anomalous stratiform precipitation contributes more than 50% to the anomalous total precipitation in the MJO. This large fraction of stratiform precipitation significantly increases the anomalous heating in the upper troposphere and decreases the anomalous heating in the lower troposphere. This helps to make the heating profile top-heavy.

The MJO anomalous vertical heating profiles in several GCMs and other models are shown to differ from the observations. The model heating profiles are generally middle-heavy.

Based on the above results, it is hypothesed that the weak and fast MJO in the current GCMs may be partly caused by their *middle-heavy anomalous vertical heating profiles*.

1. Introduction

Discovered by Madden and Julian (1971, 1972), the Madden-Julian Oscillation (MJO) is the dominant intraseasonal mode in tropical convection and circulation (e.g. Weickmann et al. 1985, Lau and Chan 1985, Salby and Hendon 1994, Wheeler and Kiladis 1998). It affects a wide range of tropical weather such as the onset and breaks of the Indian and Australian summer monsoons (e.g. Yasunari 1979, Hendon and Liebmann 1990), and the formation of tropical cyclones (e.g. Nakazawa 1986, Liebmann et al. 1994). It also drives teleconnections to the extratropics (e.g., Lau and Phillips 1986, Winkler et al. 2001) and impacts some important extratropical weather. On a longer timescale, the MJO is observed to trigger or terminate some El Nino events (e.g. Kessler et al. 1995, Takayabu et al. 1999, Bergman et al. 2001). Therefore, the MJO is important for both extended-range weather forecasting and long-term climate prediction.

Fig. 1a shows the standard deviation of the 30-70 day bandpass filtered anomaly of CPC Merged Analysis of precipitation (CMAP, see section 2 and 3 for details of the data and filtering). This intraseasonal variability has two heating centers, one in the Indian ocean, the other in the western Pacific. Fig. 1b shows the lag-correlation between the 30-70 day CMAP precipitation anomaly at 0N155E and the same quantity everywhere along the equator. The propagating part of intraseasonal variability (hereafter called the MJO) moves eastward from the Indian Ocean to the western Pacific. The phase speed in these two heating centers is about 5 m/s (or 4 degrees/day).

The motivation of this study is that eastward-propagating signals in general circulation models (GCMs) are generally too weak and propagate too fast compared with observations (e.g. Hayashi and Sumi 1986, Hayashi and Golder 1986, 1988, Lau et al. 1988, Slingo et al. 1996). This shortcoming is detrimental to both numerical weather prediction and climate prediction. Theoretical studies may offer useful

guidance on how to improve MJO simulations by GCMs, in particular how to make tropical variations stronger and propagate slower.

Observations suggest that the MJO in the eastern hemisphere is maintained by diabatic heating (Krishnamurti et al. 1985, Murakami and Nakazawa 1985, Yanai et al. 2000). The heating is strongly associated with a propagating large-scale circulation anomaly (e.g. Madden and Julian 1972, Weickmann et al. 1985, Salby and Hendon 1994), and this propagating circulation anomaly (hereafter called a "wave") also feeds back onto the heating (Hendon and Salby 1994, Zhang 1996), suggesting a wave-heating feedback theoretical framework for the MJO. In this view, the diabatic heating forces a certain kind of wave or large-scale circulation. This kind of large-scale circulation in turn influences the diabatic heating (somehow) in such a way that the whole disturbance may amplify, propagate, and decay. The heating may be divided into five components, namely the latent heatings associated with free troposphere moisture convergence, boundary layer moisture convergence, surface heat flux (latent plus sensible), and local moisture storage, as well as radiative heating. These five components have been emphasized in different types of theories historically, including the wave-CISK (Convective Instability of the Second Kind) mechanism (e.g. Hayashi 1970, Lindzen 1974, Lau and Peng 1987, Zhang and Geller 1994), the frictional wave-CISK mechanism (e.g. Hayashi 1971, Wang and Rui 1990, Salby et al. 1994), the WISHE (Wave Induced Surface Heat Exchange) mechanism (e.g. Emanuel 1987, Neelin et al. 1987), the charge-discharge mechanism (e.g. Blade and Hartmann 1993, Wang and Schlesinger 1999), and the cloud-radiation interaction mechanism (Hu and Randall 1994, Raymond 2001, Lee et al. 2001), respectively.

Because diabatic heating plays a key role in the MJO, this study examines observed heating, and compares heating profiles in observations and models. The purpose is to look for possible ways to improve model simulations of the MJO.

The seasonal mean vertical heating profile over the tropical oceans has been de-

rived from heat budget analysis of sounding arrays in several field experiments including Mashall Island (Yanai et al. 1973), GATE (e.g. Thompson et al. 1979), AMEX (e.g. Frank and McBride 1989), and TOGA COARE (Lin and Johnson 1996, Frank et al. 1996, Yanai et al. 2000, and Zhang and Lin 1999). The anomalous vertical heating profile for the MJO, which is the perturbation around the seasonal mean profile, has not been published, to our knowledge.

Stratiform precipitation is important to the vertical heating profile, because it is associated with heating in the upper troposphere and cooling in the lower troposphere, and thus shifts the heating peak upward (Houze 1982, 1989, 1997, Johnson 1984, Mapes and Houze 1995). The contribution of stratiform precipitation to the seasonal mean precipitation has been studied by many authors, but the contribution of anomalous stratiform precipitation to the anomalous precipitation in the MJO has not been reported, to our knowledge. In this study, we get this value from TOGA COARE shipborne radar data and the Tropical Rainfall Measuring Mission (TRMM) satellite data.

The datasets used in this study are described in section 2. The methods are described in section 3. The results are reported in section 4. A summary is given in section 5.

2. Data

The datasets used include TOGA COARE data and long-term data.

The TOGA COARE datasets include:

(1) 120 day (November 1, 1992 to Feburary 28, 1993) of the 6-hourly IFA diabatic heating profiles calculated by Ciesilski et al. (2002).

(2) 120 day (November 1, 1992 to Feburary 28, 1993) of the daily OSA diabatic heating profiles calculated by Zhang and Lin (1999) using the method of Zhang and Lin (1997).

(3) 120 day (November 1, 1992 to Feburary 28, 1993) of the hourly IFA radiative heating profiles calculated by Qian and Cess (2002).

(4) The radar convective/stratiform precipitation every ten minutes during the three Intensive Observation Periods (IOPs) calculated by Short et al. (1997). We use the sum of the rainfall of the whole region sampled by radar (220 Km by 220 Km area).

The long term datasets include:

(1) 21 years (1979-1999) of the pentad "chi-corrected" diabatic heating profiles calculated by Winkler et al. (2001). The horizontal resolution is 2.5 degree longitude by 2.5 degree latitude. We average the data along the equator (between 5N and 5S) with a zonal resolution of 10 degree longitude. The diabatic heating profiles are determined from an improved iterative solution of the "chi problem" (Sardesmukh 1993). This iterative procedure is applied to twice-daily NCEP reanalysis wind fields to minimize the nonlinear vorticity budget imbalance at 28 atmospheric levels, and the modified divergent wind circulation is further constrained to satisfy the large-scale mass budget. Diabatic heating rates are finally determined as a balance requirement in the heat budget, using the modified wind circulation to compute the other terms. See Sardesmukh (1993) for details of the technique.

(2) 21 years (1979-1999) of the pentad CMAP precipitation calculated by Xie and Arkin (1997). The horizontal resolution is 2.5 degree longitude by 2.5 degree latitude. We average the data along the equator (between 5N and 5S) with a zonal resolution of 5 degree longitude.

(3) 4 years (1998-2001) of the half-hourly TRMM gridded convective/stratiform precipitation (product number 3G68) calculated by Kummerow et al. (2000) and Stocker et al. (2001). The dataset contains convective/stratiform precipitation values from three algorithms: precipitation radar (PR), TRMM Microwave Imager (TMI), and combined algorithm. The horizontal resolution is 0.5 degree longitude by 0.5 degree latitude. We average the data along the equator (between 5N and 5S) to pentad data

with a zonal resolution of 5 degree longitude.

3. Method

The MJO is a broadband phenomena with an averaged period of 45 days and a wide spread from 20 to 80 days (see review by Madden and Julian 1994). To emphasize the broadband nature of the oscillation, previous studies used wide frequency bands such as 30-60 days (Weickmann et al. 1985), 35-95 days (Salby and Hendon 1994), or 30-96 days (Wheeler and Kiladis 1998). In this study we also use a wide frequency band of 30-70 days. All datasets are filtered using a 30-70 day Murakami (1976) filter, whose response function is shown in Fig. 2. The central frequency correspond to a period of 45 days. The half amplitude is at periods of 30 days and 70 days. We have also tested the Lanczos filter (Duchan 1979), and the results are not sensitive to the different choice of filters.

For the TOGA COARE data, there are two MJO events (Fig. 3). We focus on the December 1992 event. This is a strong MJO event with its amplitude at the TOGA COARE location larger than two standard deviation of the 21 year data (Fig. 1a). It moves eastward with a phase speed of 5 m/s, which is similar to the phase speed of the 21 year composite (Fig. 1b). The maximum of the 30-70 day bandpass filtered precipitation anomaly at the TOGA COARE location occurs on December 19, 1992.

For the long term data, an MJO composite is constructed using linear regression with respect to an MJO index. In this study, we use filtered CMAP precipitation as our MJO index. Because the MJO is a large-scale phenomena dominated by wavenumber 0-6 (Wheeler and Kiladis 1998), the CMAP precipitation has been zonally filtered to keep only wavenumber 0-6. The confidence level of linear correlation is estimated following Oort and Yienger (1996).

4. Results

4.1 Observed vertical heating profile in the MJO

Fig. 4 shows the vertical structure of the diabatic heating anomaly in the MJO for (a) TOGA COARE IFA, (b) TOGA COARE OSA, and (c) 21 years (1979-1999) of chi-corrected data at 0N155E. In (c) only the anomaly with linear correlation above the 95% confidence level is plotted. The time lag is with respect to the time of maximum precipitation. The time evolution is from the right to the left, showing the local evolution of heating as the eastward-moving MJO passes the measurement longitude (Fig. 3, Fig. 1b). We can see that the heating has a slight westward phase tilt with height. It develops first in the lower troposphere and then shifts upward as it intensifies. This westward phase tilt with height may be associated with more shallow convection in the earlier stage, and more stratiform precipitation in the later stage.

Fig. 5 shows the vertical heating profile at the time of maximum precipitation for the above three different datasets. The three heating profiles look quite similar with each other. They are very top-heavy, i.e., with strong heating in the upper troposphere and weak heating in the lower troposphere.

What makes the anomalous heating profile in the MJO so top-heavy? From the work of Houze (1982, 1989, 1997) and Johnson (1984), we know that the stratiform precipitation associated with organized deep convective systems is characterized by heating in the upper troposphere associated with the mesoscale updraft, and cooling in the lower troposphere associated with the mesoscale downdraft. This has the effect of shifting the altitude of the heating peak of deep convective cloud systems upward. Therefore, next we look at how much the stratiform precipitation contributes to the anomalous precipitation in the MJO.

4.2 Contribution of stratiform precipitation to the heating profile

Fig. 6 shows the anomalous precipitation during the life cycle of the MJO. The thick solid line is total precipitation, the thin solid line is convective precipitation, and

the thin dashed line is stratiform precipitation. We can see that for the MJO event during TOGA COARE (Fig. 6a), the stratiform precipitation contributes about 50% of the anomalous total precipitation. For the four year (1998-2001) MJO composite of TRMM precipitation radar(PR), TRMM Microwave Imager (TMI) and combined data (Fig. 6b-d), the stratiform precipitation contributes more than 50% of the anomalous total precipitation.

The climatological mean fraction of stratiform precipitation for the four datasets are 26%, 52%, 32%, and 47%, respectively. Although the values differ among the different datasets, the four datasets consistently show that the fraction of anomalous stratiform precipitation in the MJO is much larger than the climatological mean fraction of stratiform precipitation. Since the MJO is associated with more organized convective systems, such as the super cloud clusters (SCCs), the 2-day waves, and the mesoscale convective systems (e.g. Nakazawa 1988, Mapes and Houze 1993, Chen et al. 1996), it is possible that these organized convective systems are responsible for the large fraction of stratiform precipitation.

How much does this stratiform precipitation affect the vertical heating profile? Following the method of Johnson (1984), we partition the observed heating profile into three components: stratiform, radiative and convective (Fig. 7). The stratiform heating is assumed to have the profile observed in Johnson (1984). The 30-70 day anomalous radiative profile at the time of maximum precipitation is calculated from the hourly IFA radiative heating profiles by Qian and Cess (2002). The convective profile is derived as the residual from the total. The decomposition in Fig. 7 assumes the fraction of stratiform precipitation to be 0.3, which is the lower bound of the observed values. We can see that the stratiform precipitation significantly increases the heating in the upper troposphere and decreases the heating in the lower troposphere, which helps to make the heating profile top-heavy.

4.3 Comparison with the model heating profiles

Now we compare the observed heating profile with several model heating profiles (Fig. 8a). The models include the GFDL GCM (Lau et al. 1988), the University of Illinois GCM (Wang and Schlesinger 1999), the Seoul National University GCM (Lee et al. 2001), and two theorectical models (Sui and Lau 1989, Hendon and Salby 1996). The model heating profiles are generally middle-heavy. For clarity, we plot the difference between the observed profile and the model profiles in Fig. 8b. The observed heating profile has stronger heating in the upper troposphere and weaker heating in the lower troposphere. This is similar to the stratiform heating profile, suggesting that the models may lack some equivalent of the observed anomalous stratiform precipitation.

The above results lead us to the following hypothesis: the weak and fast MJO in the current GCMs may be partly caused by their *middle-heavy anomalous vertical heating profiles*.

This hypothesis can be tested in the GCMs by adding some equivalent of the observed stratiform precipitation to make the diabatic heating profile top-heavy. There are two possible ways to do this. The first is to revise the convection schemes to include the mesoscale effects (with the observed magnitude). Some convection schemes (e.g. Molinari and Corsetti 1985, Sud and Walker 1993) have included the effect of mesoscale downdraft, and some schemes (e.g. Donner 1993, Alexander and Cotton 1998, Gray 2000, Donner et al. 2001) have included the effects of both mesoscale updraft and mesoscale downdraft. Donner et al. (2001) apply the Donner (1993) convection scheme in the GFDL GCM. The simulated seasonal mean fraction of stratiform precipitation agrees well with the TRMM observation. The MJO simulation is reported to be improved (Donner 2002, presentation in the Seventh Annual CCSM workshop), but further diagnostics need to be done to determine whether the model MJO heating profile becomes more top-heavy. The second way of adding equivalent of stratiform precipitation is to encourage more precipitation to occur via the large-

scale (resolved) condensation and microphysics schemes, by assuming some fraction of the convective condensate falls through air with large-scale mean humidity. Li et al. (2002) apply this modification to a GCM and find that it improves the mean state of the model. But the effect on the MJO simulation has not been studied. Some cautions need to be taken when applying the above modifications to a GCM: (1) These modifications may change the mean state of the GCM, which may also affect the amplitude and phase speed of the model MJO. We need to seperate the effect of heating profile from that of the mean state. (2) These modifications may induce the response of some other schemes, such as the shallow convection scheme, which may cancel the intended change in the final diabatic heating profile. Therefore we need to plot the final diabatic heating profile to see if it is really top-heavy.

This hypothesis can also be examined in theoretical models of the effect of vertical heating profiles on wave-heating feedbacks. Within the context of wave-CISK theory (Yamasaki 1968a, 1968b, 1969, Chang and Lim 1988, Sui and Lau 1989, Cho and Pendlebury 1997), the heating profile is specified outright, allowing direct study of the influence of profile shape. Cho and Pendlebury (1997) decomposed the specified heating profile into vertical modes, and found that instability required a sufficient amplitude of the full-wave mode, with heating above cooling, corresponding directly to the stratiform contribution to heating profiles. Yamasaki (1968a, 1968b, 1969) suggested in his numerical modeling studies that some tropical systems are unstable only if the cumulus heating profile has a maximum in the upper troposphere. Sui and Lau (1989) found that a "deep convection" heating profile (lighter dash-dot line in Fig. 8) increases the propagation speed of disturbances, compared to the case with a shallower heating profile. At first glance, this suggests that more top-heavy heating will make faster, not slower, waves, but in fact their "deep convection" heating profile is middle-heavy with strong heating in the lower troposphere, although its mathematical maximum is in the upper troposphere. Therefore their comparison was

between bottom-heavy and middle-heavy heating profiles. It might more generally be true that theories which attatch importance to higher vertical modes - whether by top-heavy or bottom-heavy excitation - will tend to generate slower-propagating modes.

5. Summary

Diabatic heating profiles calculated from the Tropical Ocean Global Atmosphere (TOGA) Coupled Ocean Atmosphere Response Experiment (COARE) sounding array and the long-term NCEP reanalysis, supplemented by the convective/stratiform precipitation from TOGA COARE and Tropical Rainfall Measurement Mission (TRMM) satellite, are used to study the contribution of the stratiform precipitation to the vertical heating profile in the MJO. All datasets are filtered using a 30-70 day bandpass filter. The long-term MJO composites are constructed using linear correlation and linear regression with respect to precipitation.

The observed MJO anomalous vertical heating profile is very top-heavy at the time of maximum precipitation. TRMM data and TOGA COARE radar data show that the anomalous stratiform precipitation contributes more than 50% to the anomalous total precipitation in the MJO. This large fraction of stratiform precipitation significantly increases the anomalous heating in the upper troposphere and decrease the anomalous heating in the lower troposphere. This helps to make the heating profile top-heavy.

The MJO anomalous vertical heating profiles in several current GCMs and other models are shown to differ from the observations. The model heating profiles are generally middle-heavy.

Based on the above results, we have the following hypothesis: the weak and fast MJO in the current GCMs maybe partly caused by their middle-heavy anomalous vertical heating profiles.

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FIGURE CAPTIONS

Fig. 1 (a) Standard deviation of the 30-70 day bandpass filtered anomaly of the CMAP precipitation from 1979-1999. The unit is mm/day. The thick solid polygons are the sounding arrays during TOGA COARE. The inner one is the IFA. The outer one is the OSA.

(b) The lag-correlation between the 30-70 day CMAP precipitation anomaly and itself at 0N155E.

Fig. 2 The response function of the Murakami filter used in this study.

Fig. 3 The 30-70 day bandpass filtered anomaly of the CMAP precipitation during TOGA COARE (November 1, 1992 to Febuary 28, 1993). The unit is mm/day. The thick dashed line indicates the location of the TOGA COARE sounding arrays.

Fig. 4 The vertical structure of the diabatic heating anomaly in the MJO for (a) TOGA COARE IFA. (b) TOGA COARE OSA. (c) 21 years (1979-1999) of chi-corrected data at 0N155E. The time lag is with respect to the time of maximum precipitation. The heating anomaly has been normalized by its column-integration at the time of maximum precipitation. The unit is (K/day)/(mm/day). The first contour is 0.03, and the contour interval is 0.15. Negative contours are shaded. In (c) only the anomaly with linear correlation above the 95% confidence level is plotted.

Fig. 5 The MJO anomalous vertical heating profile at the time of maximum precipitation for TOGA COARE IFA (thick solid line), TOGA COARE OSA (thick dashed line), and 21 years (1979-1999) of chi-corrected data at 0N155E (thin solid line).

Fig. 6 The precipitation anomaly during the life cycle of the MJO for (a) TOGA COARE radar. (b) 4 years (1998-2001) of TRMM PR data at 0N147E. (c) 4 years (1998-2001) of TRMM TMI data at 0N147E. and (d) 4 years (1998-2001) of TRMM combined data at 0N147E. The thick solid line is the total precipitation. The thin solid line is the convective precipitation. The thin dashed line is the stratiform pre-

cipitation.

Fig. 7 Partition of the observed heating profile (thick solid line) into three components: the stratiform component (thin solid line), the radiative component (thin dotted line), and the convective component (thin dashed line).

Fig. 8 (a) Comparison of the observed heating profile with the model heating profiles. The thick solid line is the observed heating profile. All other lines are the model heating profiles. The models include the GFDL GCM used in Lau et al. 1988 (thick dashed line), the University of Illinois GCM used in Wang and Schlesinger (1999)'s A3 experiment (thin solid line) and M2 experiment (thin dotted line), the theoretical models of Sui and Lau (1989)'s deep convection case (thin dashed line), and Hendon and Salby (1996, thin dot-dashed line). (b) The difference between the observed heating profile and the model heating profiles.

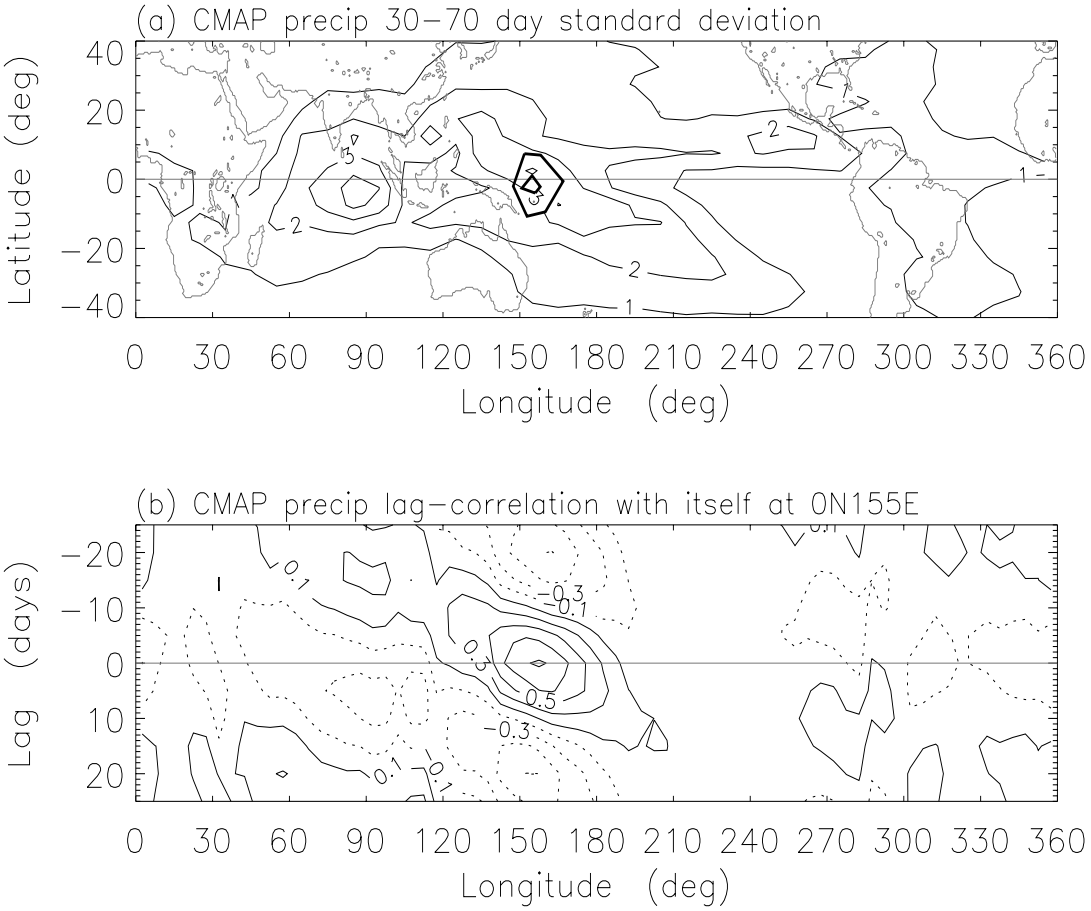
(a) CMAP precip 30−70 day standard deviation
Latitude (deg)
40
20
0
−20
−40
0 30 60 90 120 150 180 210 240 270 300 330 360
Longitude (deg)
(b) CMAP precip lag−correlation with itself at 0N155E
Lag (days)
−20
−10
0
10
20
0 30 60 90 120 150 180 210 240 270 300 330 360
Longitude (deg)

Figure 1:

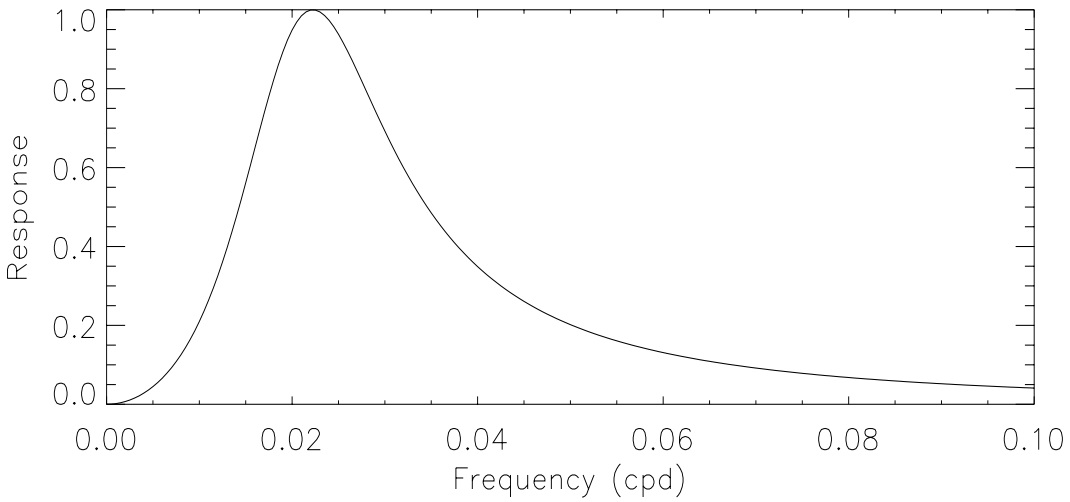
1.0
0.8
0.6
0.4
0.2
0.0
Response
0.00
0.02
0.04
0.06
0.08
0.10
Frequency (cpd)

Figure 2:

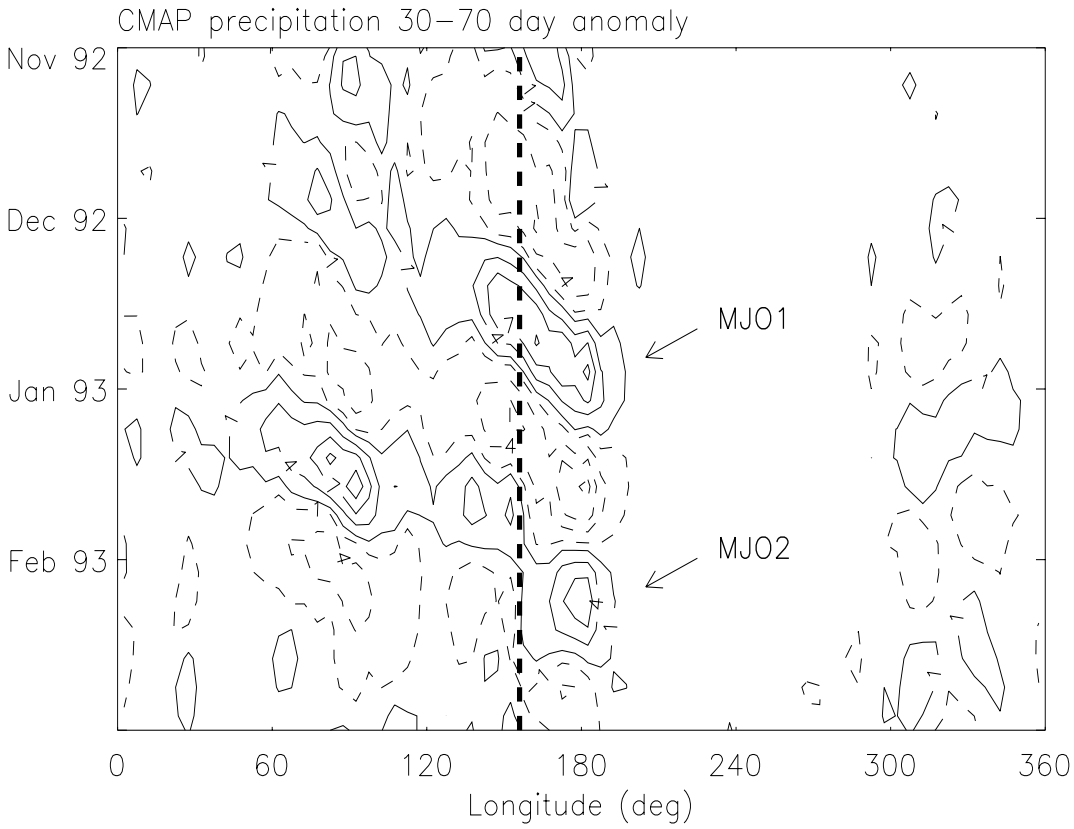
CMAP precipitation 30−70 day anomaly
Nov 92
Dec 92
Jan 93
Feb 93
MJO1
MJO2
0
60
120
180
240
300
360
Longitude (deg)

Figure 3:

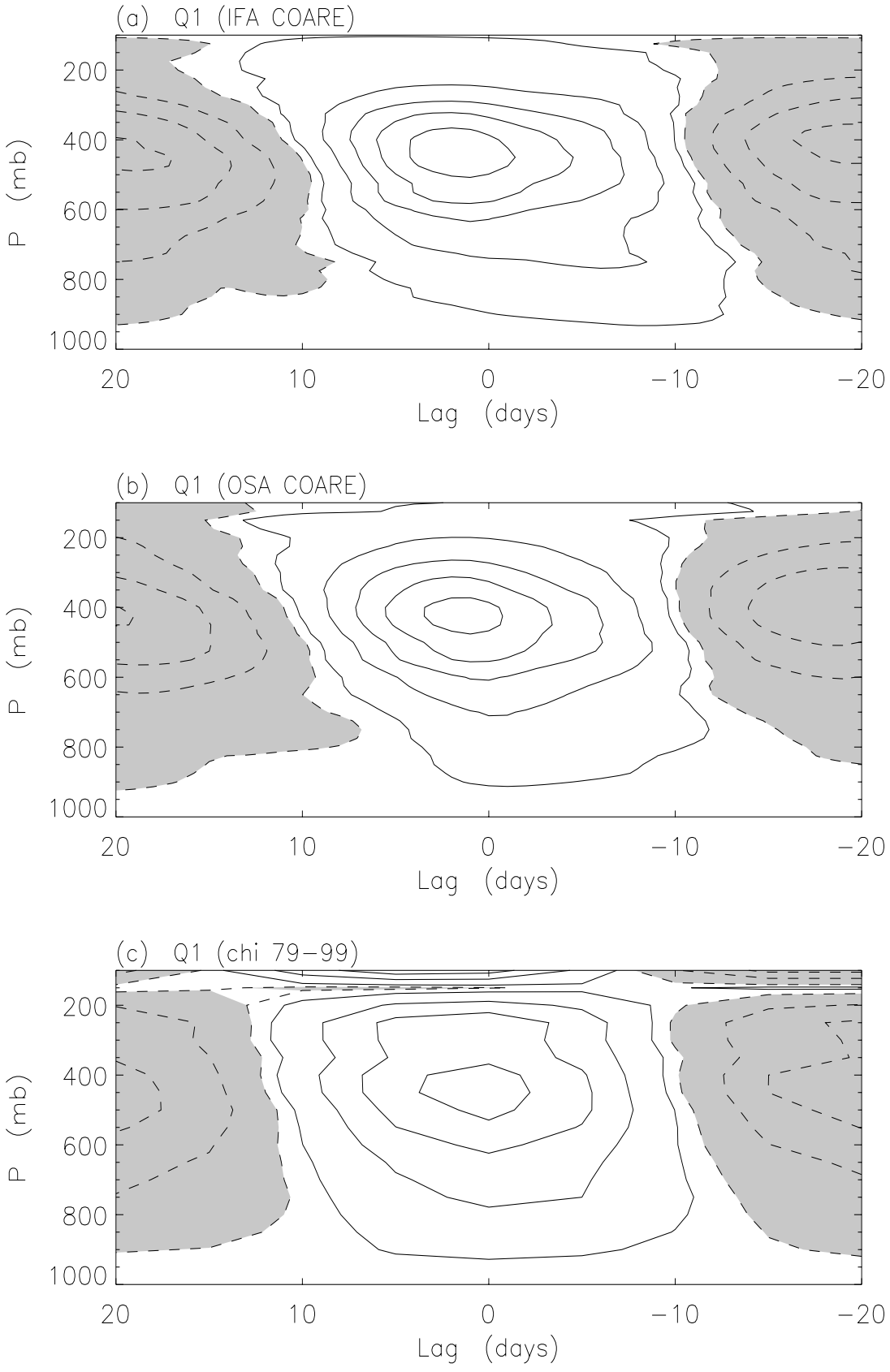
(a) Q1 (IFA COARE)
P (mb)
200
400
600
800
1000
20
10
0
−10
−20
Lag (days)
(b) Q1 (OSA COARE)
P (mb)
200
400
600
800
1000
20
10
0
−10
−20
Lag (days)
(c) Q1 (chi 79−99)
P (mb)
200
400
600
800
1000
20
10
0
−10
−20
Lag (days)

Figure 4:

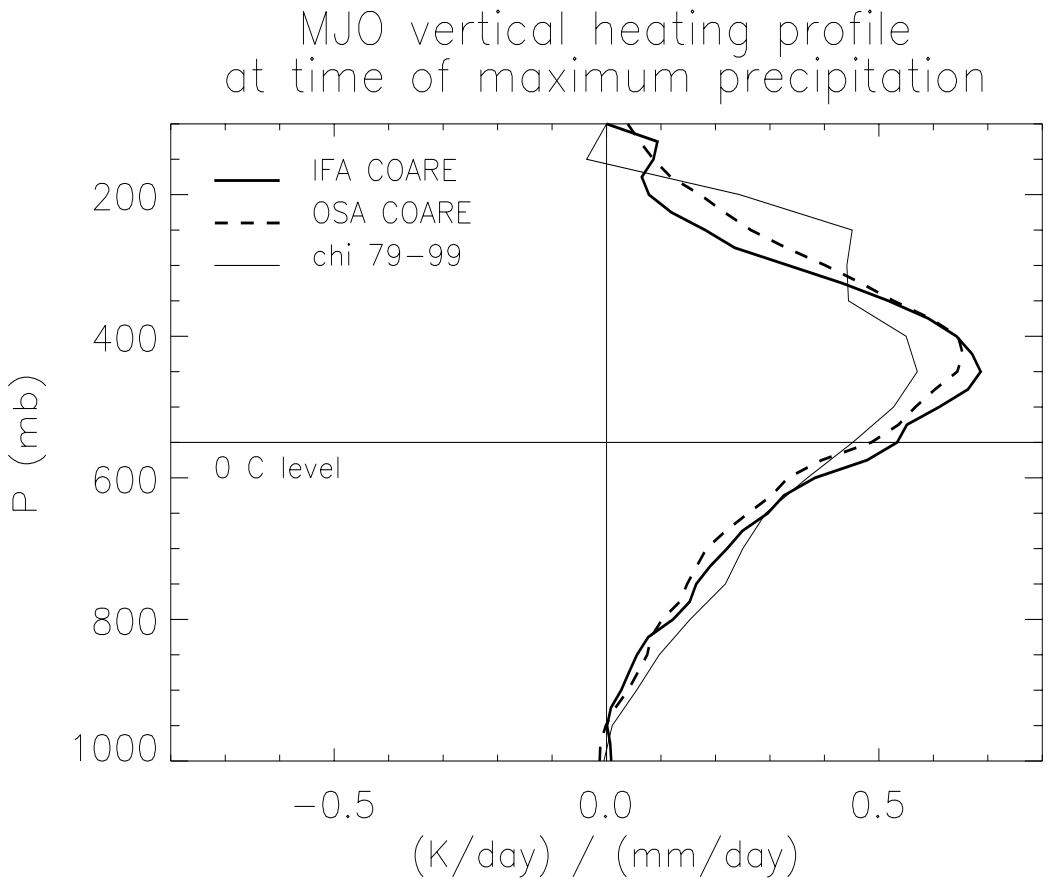
MJO vertical heating profile
at time of maximum precipitation
IFA COARE
OSA COARE
chi 79–99
0 C level
P (mb)
200
400
600
800
1000
−0.5
0.0
0.5
(K/day) / (mm/day)

Figure 5:

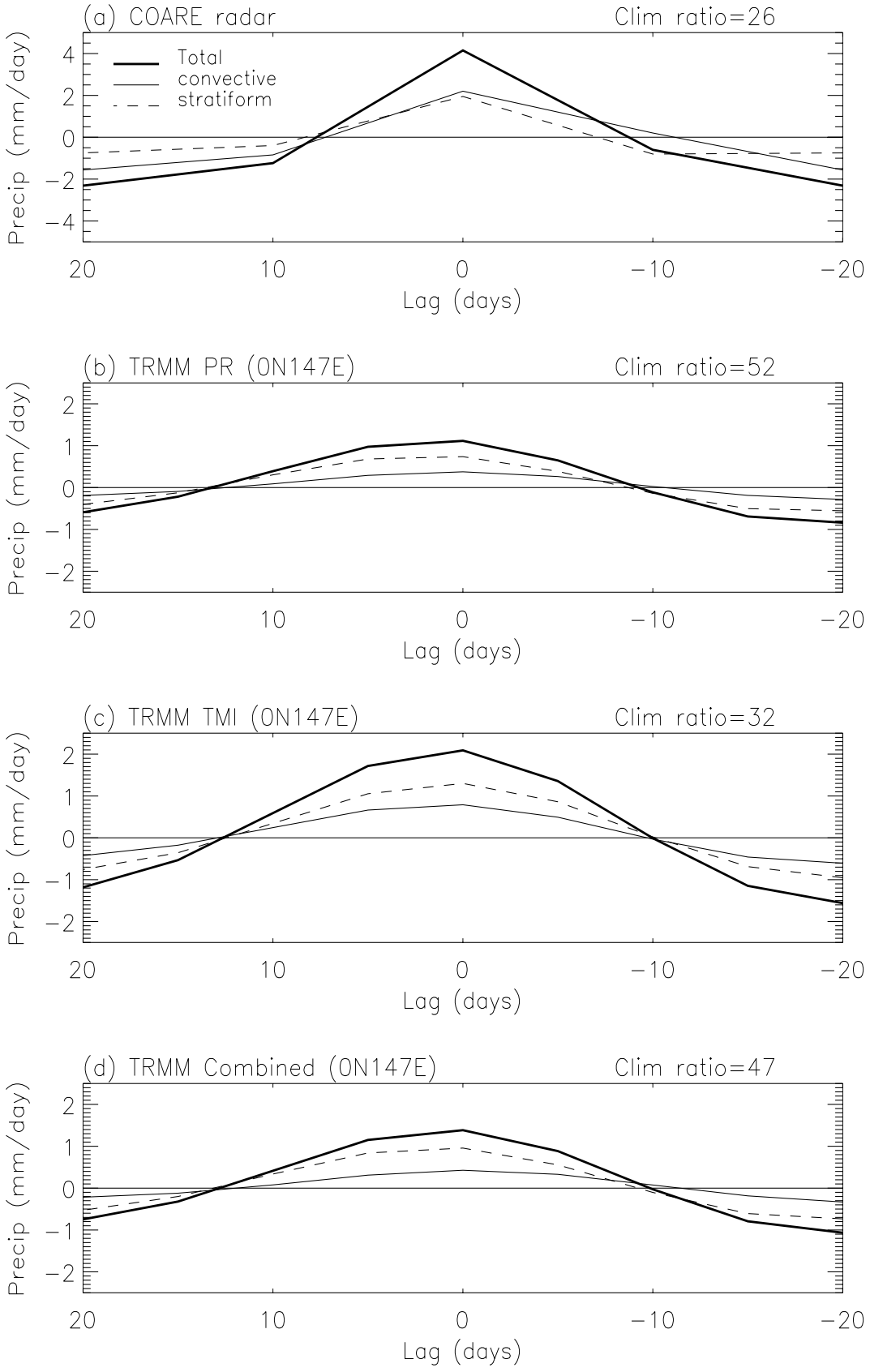
(a) COARE radar
Clim ratio=26
Total
convective
stratiform
Precip (mm/day)
Lag (days)
(b) TRMM PR (0N147E)
Clim ratio=52
Precip (mm/day)
Lag (days)
(c) TRMM TMI (0N147E)
Clim ratio=32
Precip (mm/day)
Lag (days)
(d) TRMM Combined (0N147E)
Clim ratio=47
Precip (mm/day)
Lag (days)

Figure 6:

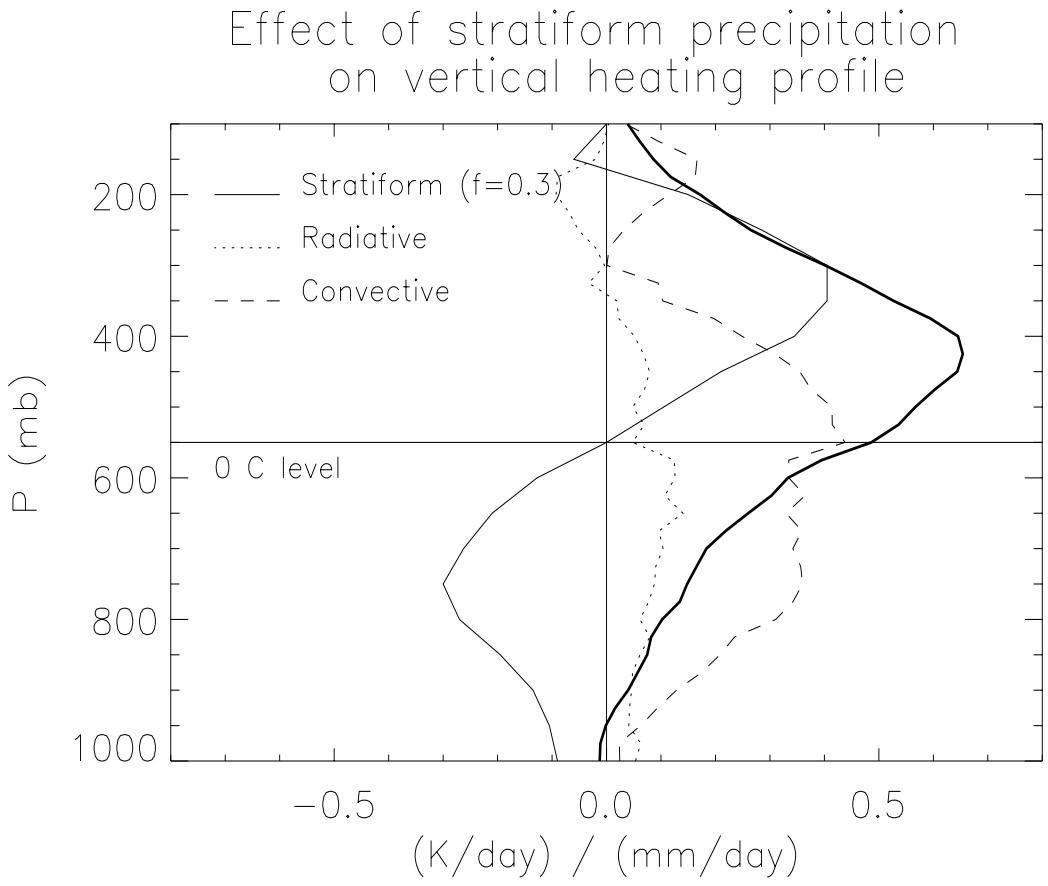
Effect of stratiform precipitation
on vertical heating profile
Stratiform (f=0.3)
Radiative
Convective
0 C level
P (mb)
200
400
600
800
1000
−0.5
0.0
0.5
(K/day) / (mm/day)

Figure 7:

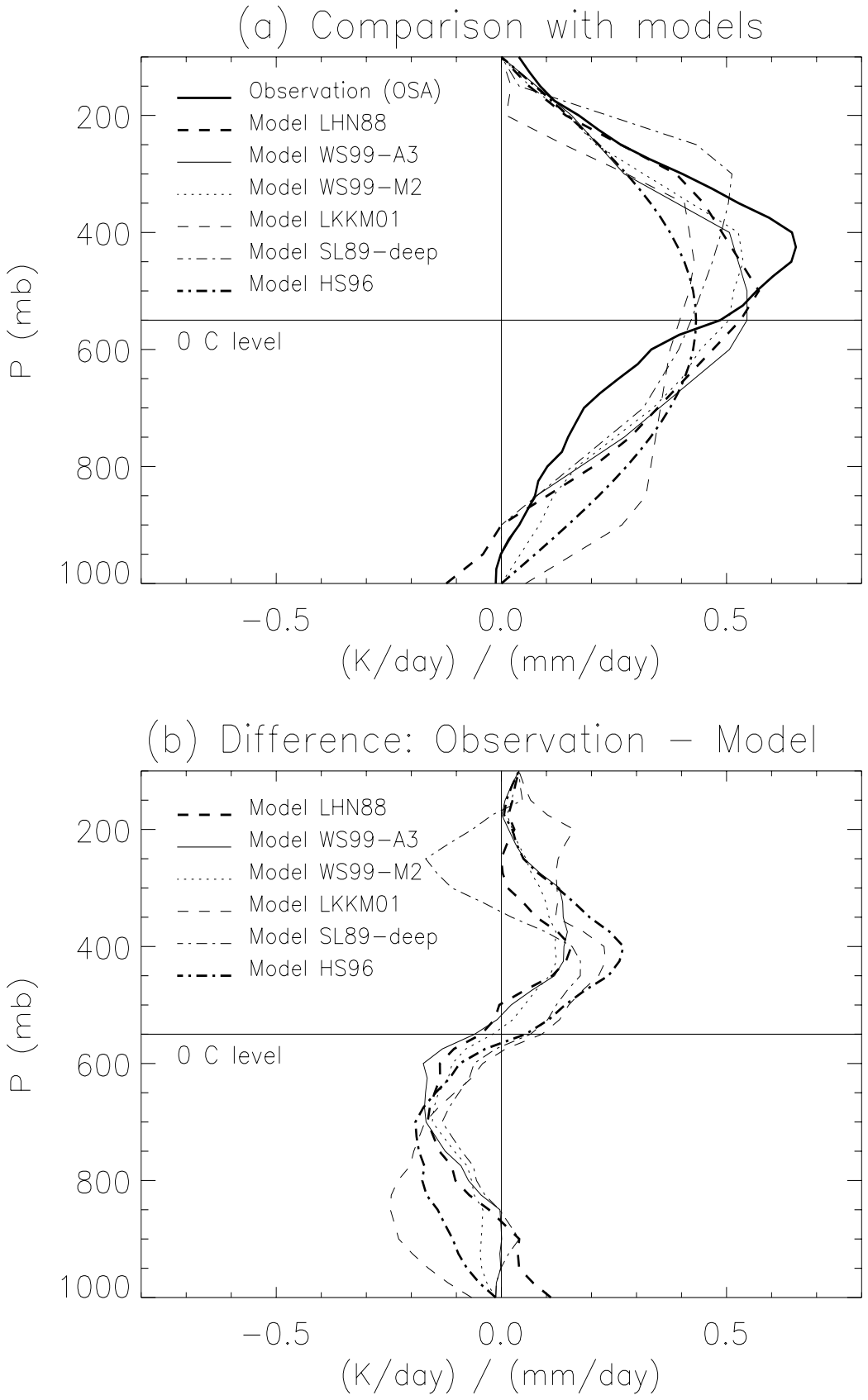
(a) Comparison with models
Observation (OSA)
Model LHN88
Model WS99−A3
Model WS99−M2
Model LKKM01
Model SL89−deep
Model HS96
0 C level
P (mb)
200
400
600
800
1000
−0.5
0.0
0.5
(K/day) / (mm/day)
(b) Difference: Observation − Model
Model LHN88
Model WS99−A3
Model WS99−M2
Model LKKM01
Model SL89−deep
Model HS96
0 C level
P (mb)
200
400
600
800
1000
−0.5
0.0
0.5
(K/day) / (mm/day)

Figure 8: